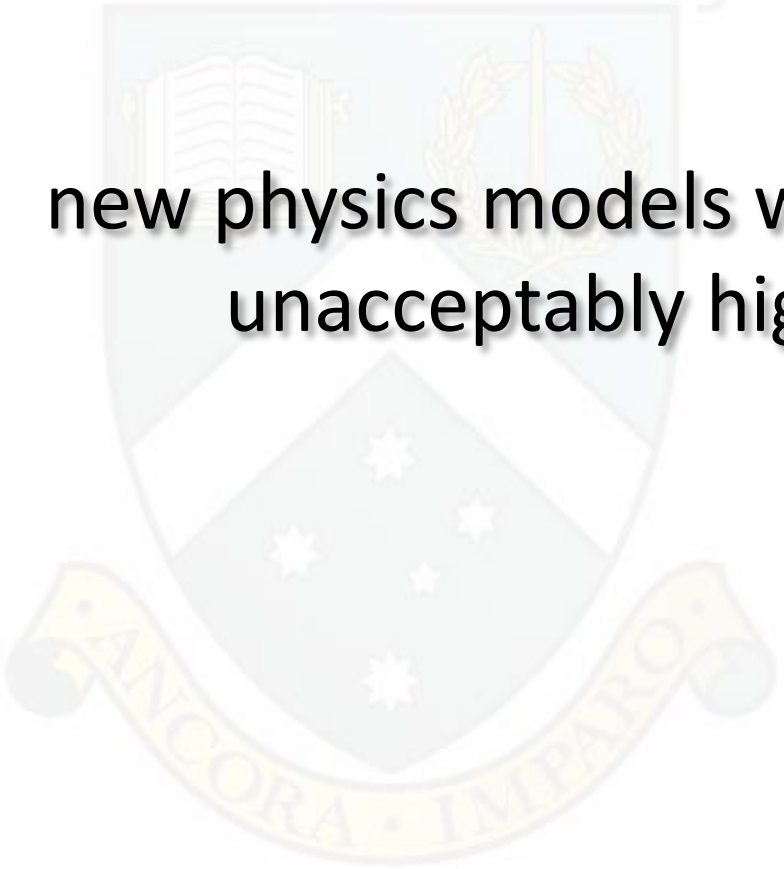
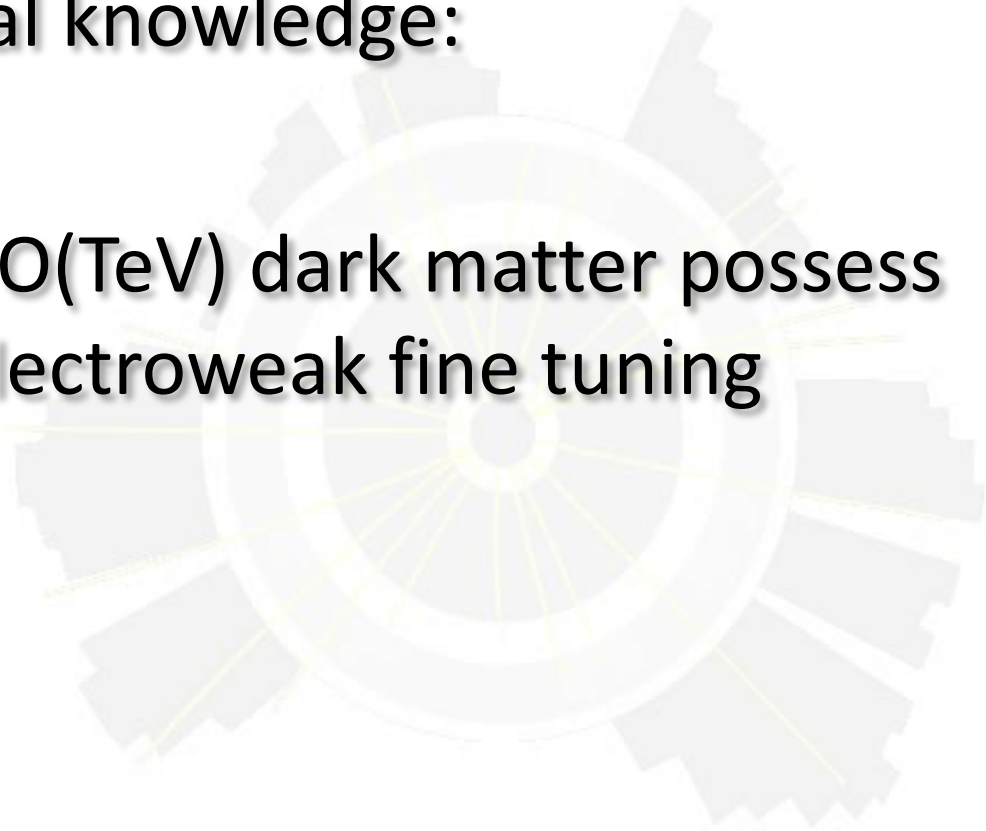


Naturalness, supersymmetry and dark matter

Peter Athron, Csaba Balázs, Ben Farmer, Doyoun Kim
PRD90 5 055008 (2014)

conventional knowledge:

new physics models with $O(\text{TeV})$ dark matter possess
unacceptably high electroweak fine tuning



example:

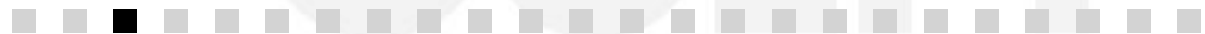
the Minimal Supersymmetric Standard Model is pushed by the LHC to parameter regions which are fine tuned at worse than percent level



this talk:

electroweak fine tuning is overestimated by the
conventional measures

$O(\text{TeV})$ dark matter can be natural





outline





NATURALNESS
???

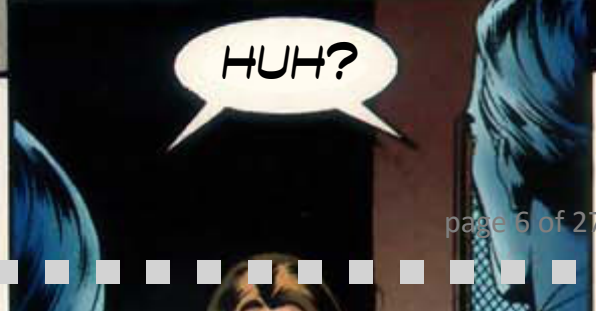
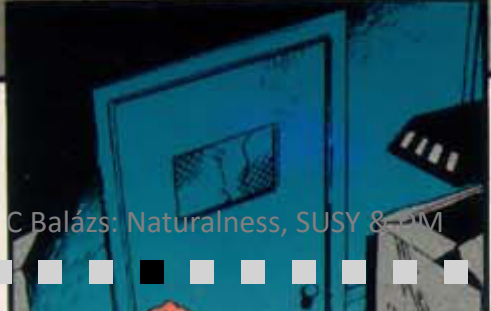
NOW THEY
TELL ME!

GOSH!!!

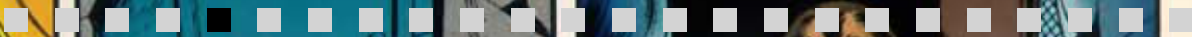
HEY!



HMMM...



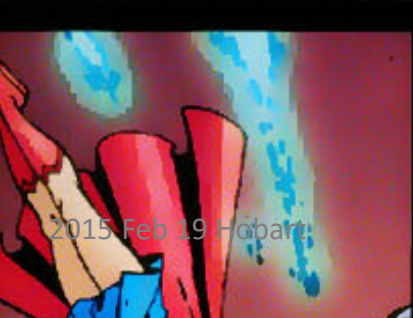
HUH?



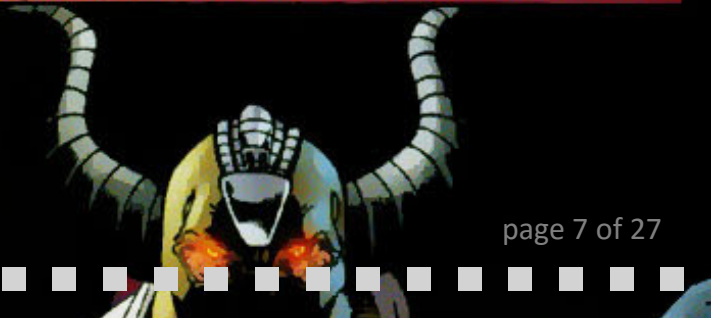


LETS ...

... GET ...



... NATURAL



DESTROY
YOU!!!

NO!

HELP!

HA-HA-HA...





PREPARE
...

... TO ...

... **DIE!!!**

the electroweak naturalness problem

physical phenomena characterized by disparate
(energy or length) scales are separated:
their governing laws can be understood largely
independently from each other

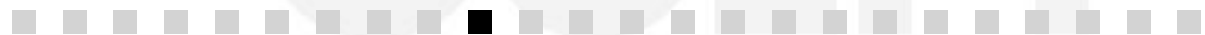
when viewed as an effective theory
the Standard Model Higgs mass receives
quantum corrections from the Planck scale

$m_H = 125 \text{ GeV} \Rightarrow$ the Standard Model is unnatural



What is natural?

naturalness is quantified by a fine-tuning measure



Barbieri-Ellis-Giudice measure

the most used, misused and abused electroweak fine tuning measure

$$\frac{\partial(\textit{electroweak observable})}{\partial(\textit{theory parameters})}$$

measures the “sensitivity” of the “weak scale” to the change in the “theory parameter(s)”

Barbieri-Ellis-Giudice measure: example

Constrained Minimal Supersymmetric Standard Model

theory parameters: $M_0, M_{1/2}, A_0, \tan\beta, \text{sign}(\mu)$

dominant fine tuning term: $\frac{\partial m_Z}{\partial \mu}$

Barbieri-Ellis-Giudice measure: example

electroweak symmetry breaking condition:

$$\frac{m_Z^2}{2} = \frac{M_{H_d}^2 - M_{H_u}^2 \tan^2 \beta}{\tan^2 \beta - 1} - \mu^2$$

fixes μ by connecting it to m_Z

this justifies $\frac{\partial m_Z}{\partial \mu}$

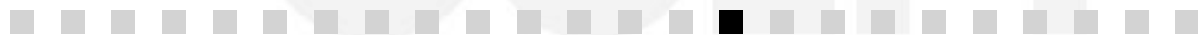
Bayesian evidence

$$\mathcal{E} = \int \mathcal{L}(\mu) \pi(\mu) d\mu$$

for a theory with a single parameter μ
 \mathcal{E} quantifies the plausibility of the theory

the theory predicts $m_Z(\mu)$
so $m_Z(\mu)$ is invertible

$$m_Z(\mu) \Rightarrow \mu(m_Z)$$



write \mathcal{E} as an integral over m_Z

$$\mathcal{E} = \int \mathcal{L}(m_Z) \pi(m_Z) \frac{d\mu}{dm_Z} dm_Z$$

$$\mathcal{E} = \int \mathcal{L}(m_Z) \pi(m_Z) \left(\frac{dm_Z}{d\mu} \right)^{-1} dm_Z$$

Allanach, Hooper JHEP 0810:071 (2008)

Bayesian evidence \sim inverse of fine-tuning measure!

$$\mathcal{E} = \int \mathcal{L}(m_Z) \pi(m_Z) \left(\frac{dm_Z}{d\mu} \right)^{-1} dm_Z$$

fixing μ using m_Z automatically induces $\frac{dm_Z}{d\mu}$ in \mathcal{E}

$\mathcal{E} \sim$ plausibility that the theory correctly predicts m_Z

this is Occam's razor at work!

MSSM

trade $\{m_Z, m_t, \tan\beta\}$
for $\{\mu, y_t, B_0\}$

$$\text{naturalness prior} = \begin{vmatrix} \frac{\partial m_Z}{\partial \mu} & \frac{\partial m_t}{\partial \mu} & \frac{\partial \tan\beta}{\partial \mu} \\ \frac{\partial m_Z}{\partial y_t} & \frac{\partial m_t}{\partial y_t} & \frac{\partial \tan\beta}{\partial y_t} \\ \frac{\partial m_Z}{\partial B_0} & \frac{\partial m_t}{\partial B_0} & \frac{\partial \tan\beta}{\partial B_0} \end{vmatrix}$$

Cabrera, Casas, deAustri JHEP 0903:075,2009

Barbieri-Ellis-Giudice EW FT measure:

$$\frac{\partial(\text{electroweak observable})}{\partial(\text{theory parameters})}$$

- evidence = integral over parameters
- observables can be used to eliminate parameters
- evidence ratios have clear normalization scale
- Observables: subjective choice!
- Parameters: subjective choice!
- Form: depends on prior!
- Expression: determinant!
- What does $\frac{\partial m_Z}{\partial \mu}$ have to do with $\Delta m_H, \mu, m_{\tilde{t}}, m_{\tilde{g}}, \dots$?
- Beyond MSSM, NMSSM, SUSY: evidence can be defined!

NMSSM

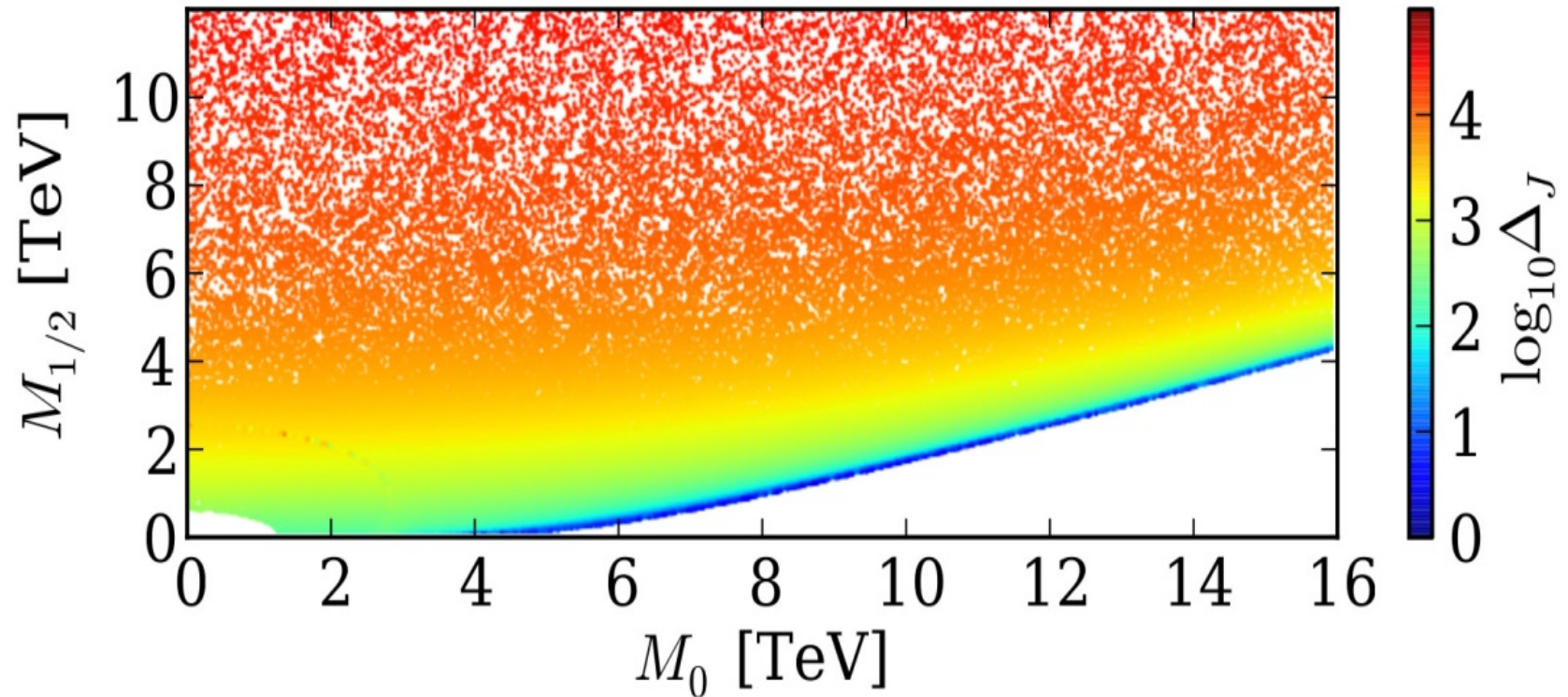
trade $\{\ln m_Z^2, \ln \lambda, \ln \tan \beta\}$
 for $\{\ln m_S^2, \ln \lambda_0, \ln \kappa_0\}$

$$\text{naturalness prior} = \begin{vmatrix} \frac{\partial \ln m_Z^2}{\partial \ln m_S^2} & \frac{\partial \ln \lambda}{\partial \ln m_S^2} & \frac{\partial \ln \tan \beta}{\partial \ln m_S^2} \\ \frac{\partial \ln m_Z^2}{\partial \ln \lambda_0} & \frac{\partial \ln \lambda}{\partial \ln \lambda_0} & \frac{\partial \ln \tan \beta}{\partial \ln \lambda_0} \\ \frac{\partial \ln m_Z^2}{\partial \ln \kappa_0} & \frac{\partial \ln \lambda}{\partial \ln \kappa_0} & \frac{\partial \ln \tan \beta}{\partial \ln \kappa_0} \end{vmatrix}$$

Ahron, Balazs, Farmer, Kim PRD90 5 055008 (2014)

naturalness prior map in the CMSSM

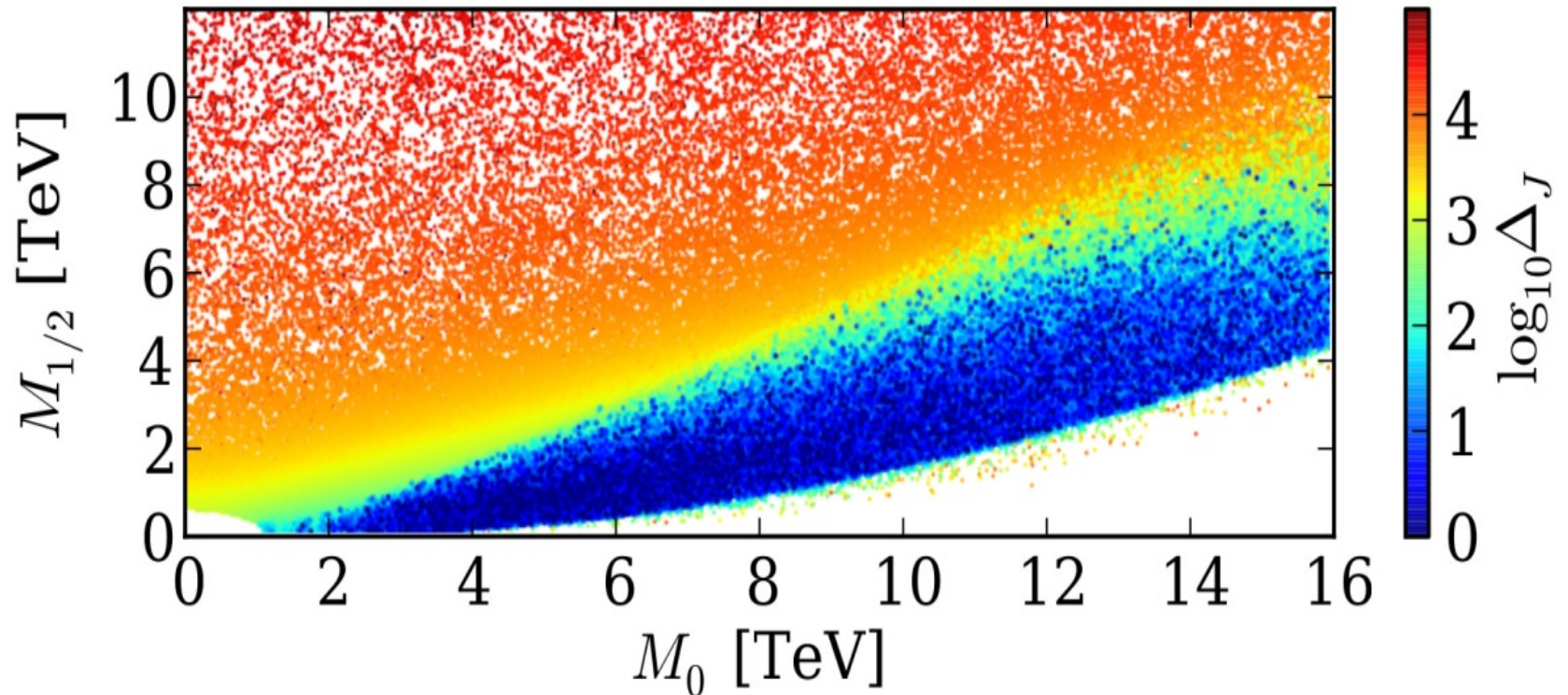
$$A_0 = -2.5 \text{ TeV}, \tan\beta = 10$$



Ahron, Balazs, Farmer, Kim PRD90 5 055008 (2014)

naturalness prior map in the CNMSSM

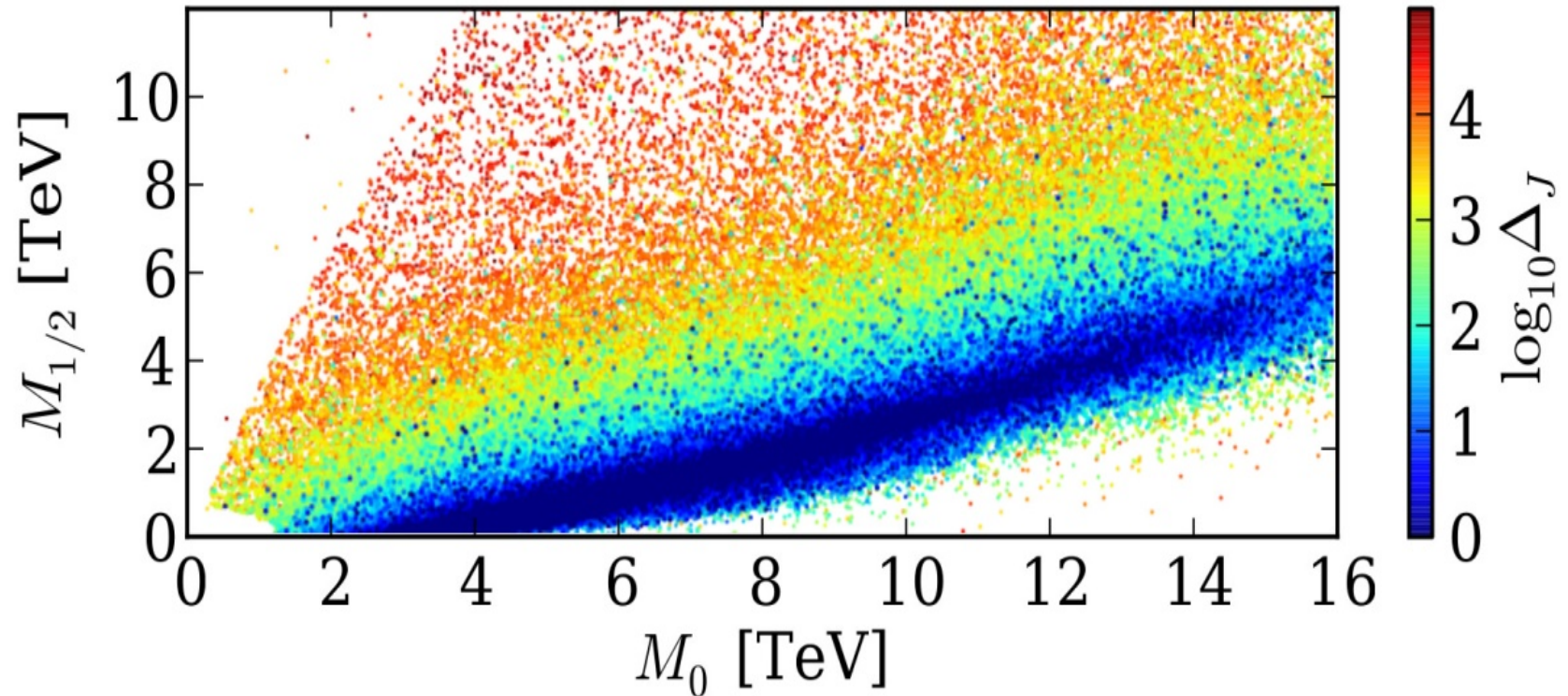
$$A_0 = -2.5 \text{ TeV}, \tan\beta = 10$$



Ahron, Balazs, Farmer, Kim PRD90 5 055008 (2014)

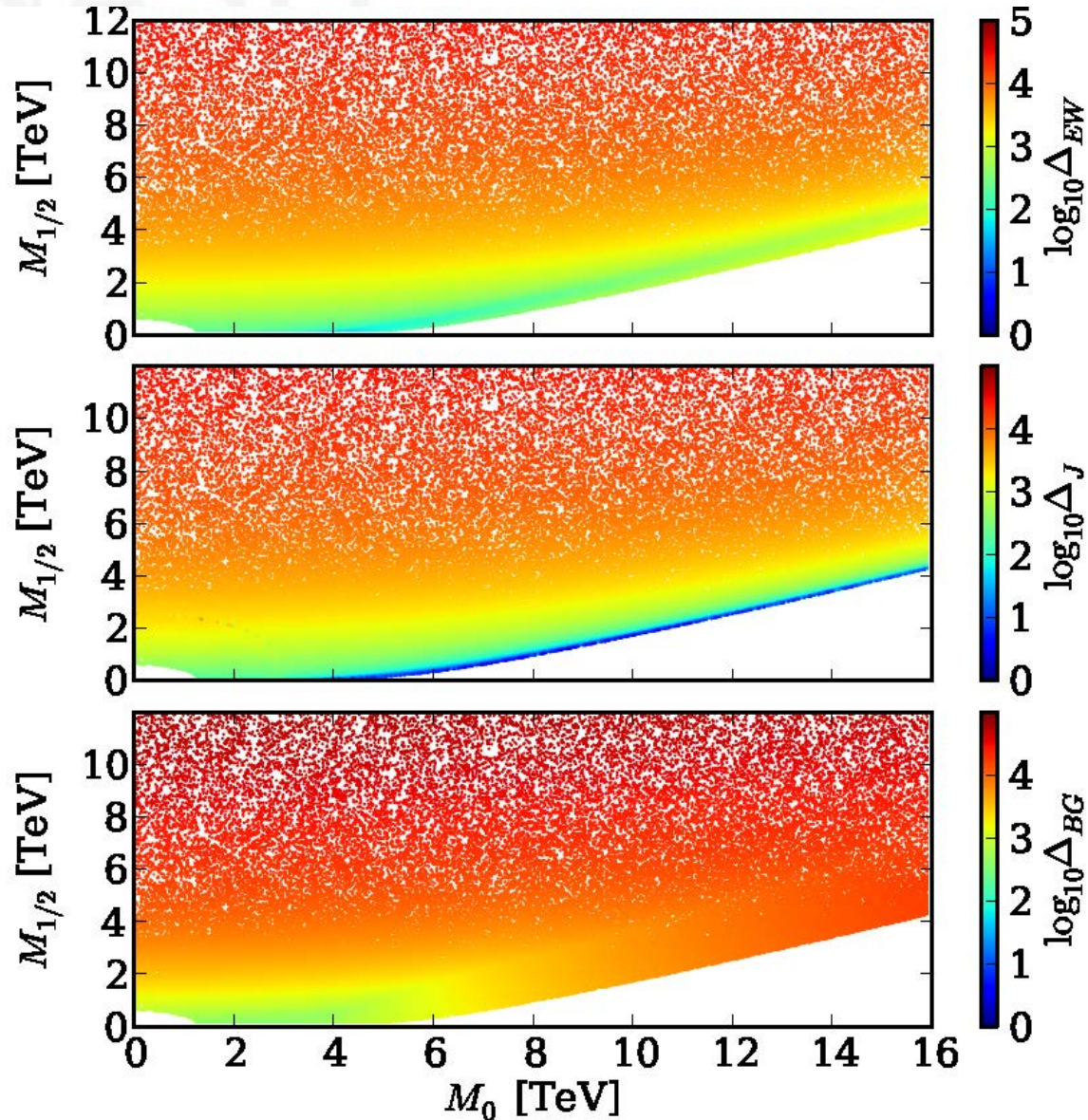
naturalness prior map in NMSSM-11

$$A_0 = -2.5 \text{ TeV}, \tan\beta = 10$$



Ahron, Balazs, Farmer, Kim PRD90 5 055008 (2014)

naturalness prior agrees with EW FT measure



conclusions

naturalness is a robust principle

it can be quantified within the Bayesian framework

fine tuning is measured by the naturalness prior

according to the naturalness prior $O(\text{TeV})$ dark matter
can be perfectly natural

