

Supersymmetric Dark Matter or Not

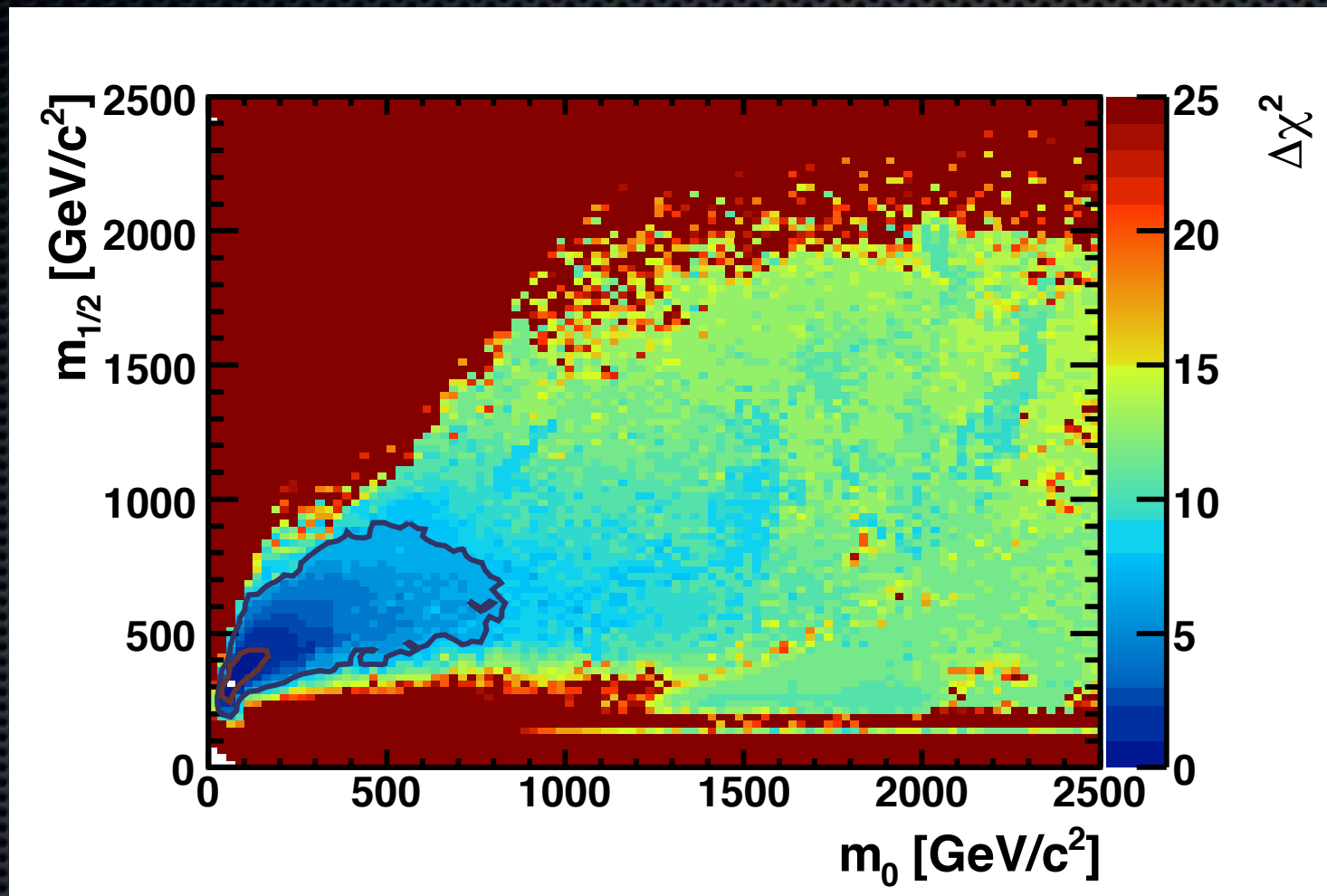
1) With CMSSM-like models pushed to high mass scales, can we still 'guarantee' Supersymmetry's discovery at the LHC. Viable dark matter models in the CMSSM tend to lie in strips (co-annihilation, funnel, focus point), how far up in energy do these strips extend?

2) Non-Supersymmetric SO(10) - gauge coupling unification; neutrino masses; AND DARK MATTER.

$\Delta\chi^2$ map of $m_0 - m_{1/2}$ plane

Mastercode

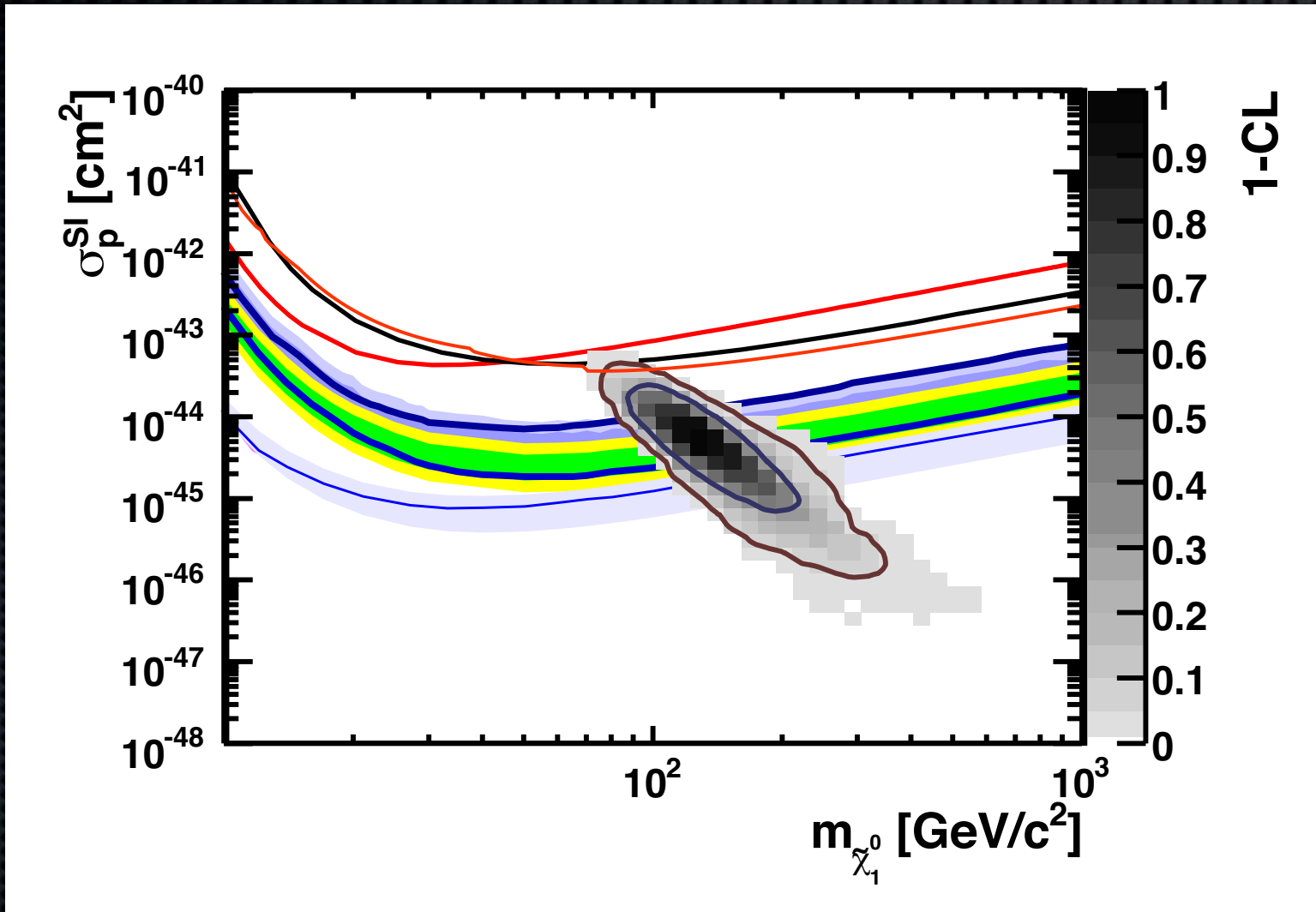
2009



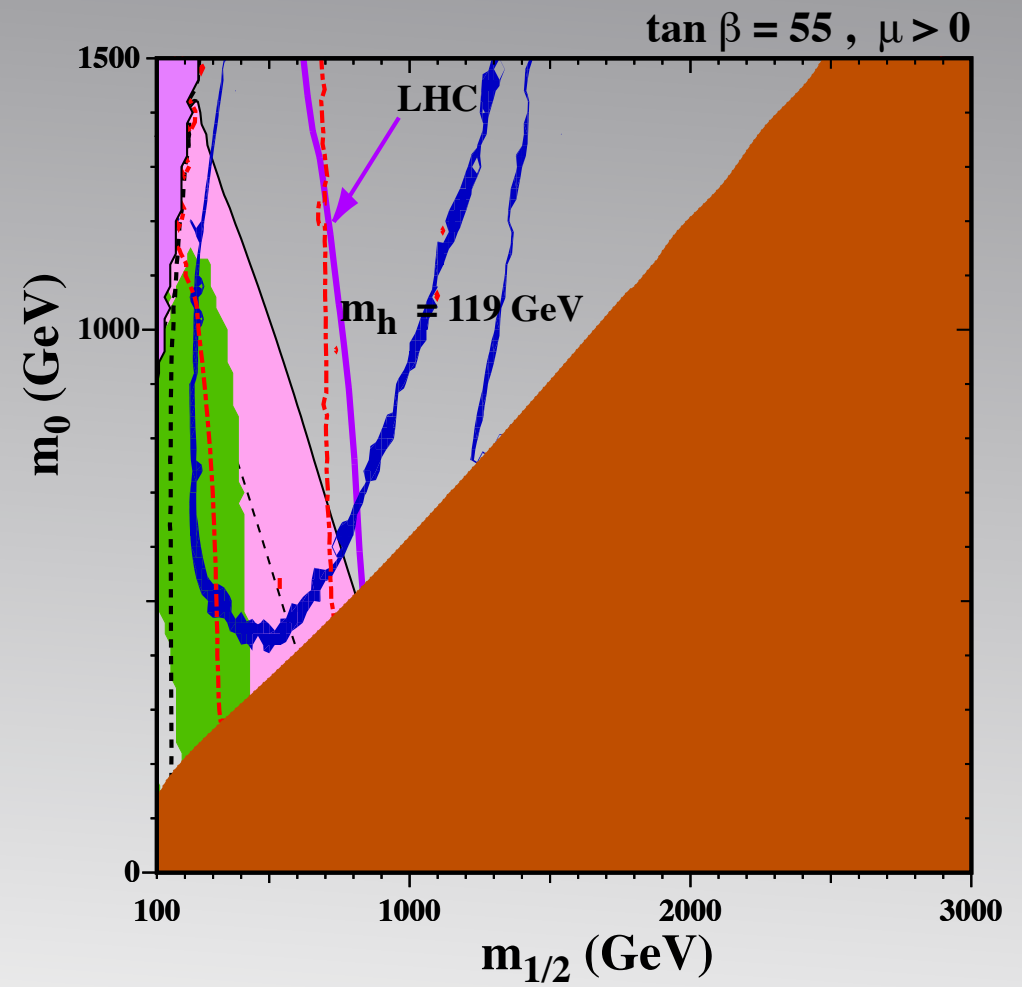
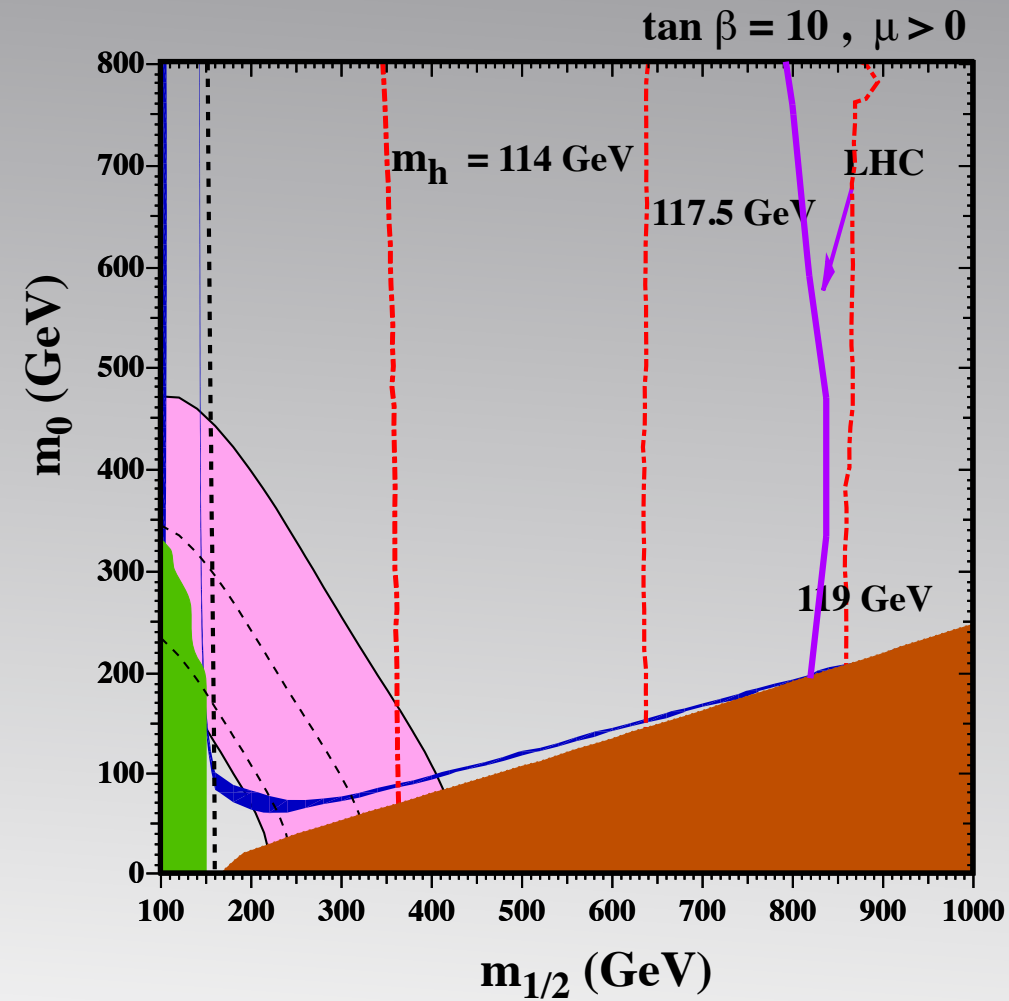
Elastic scattering cross-section

Mastercode

2009



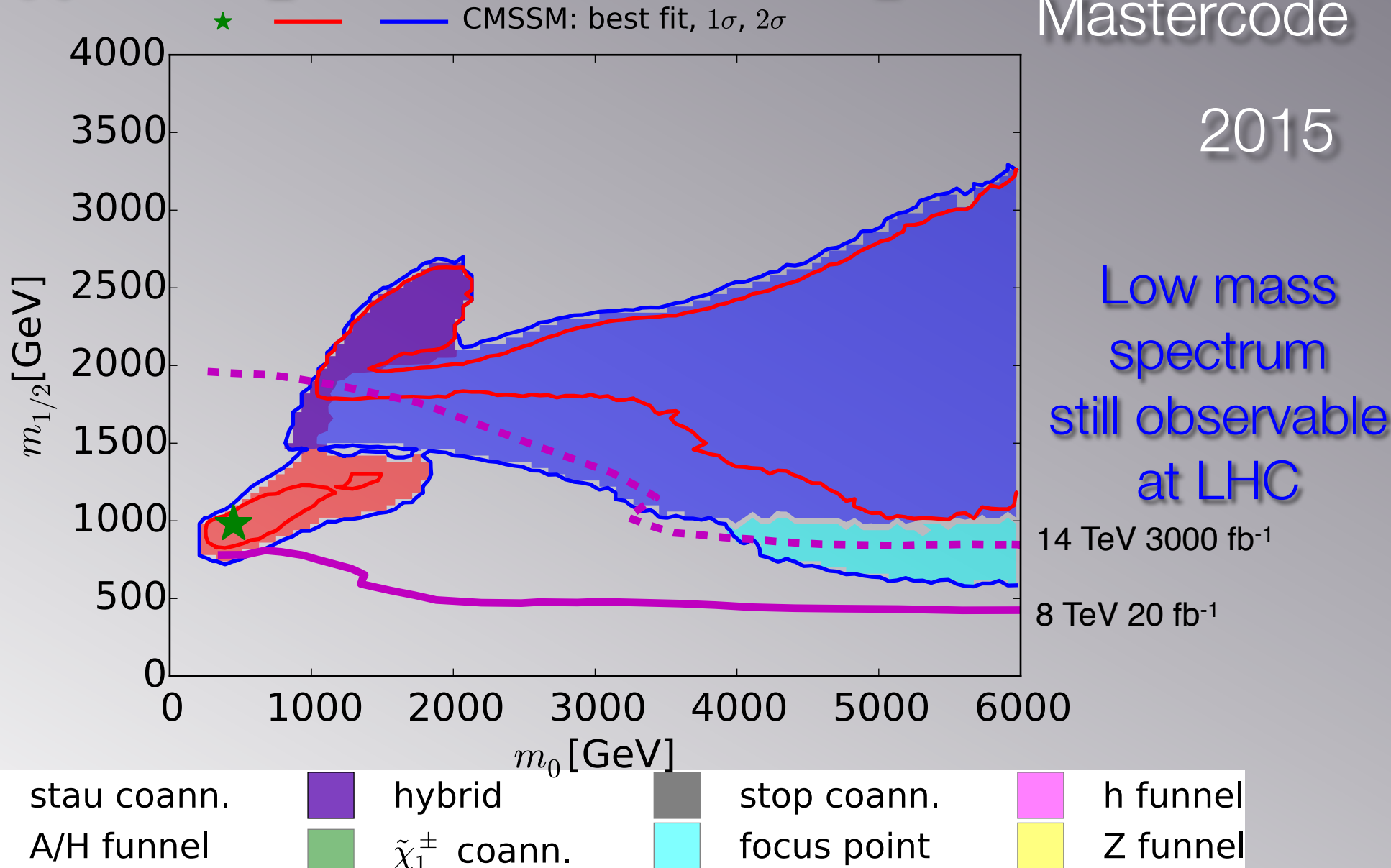
$m_{1/2} - m_0$ planes incl. LHC



$\Delta\chi^2$ map of $m_0 - m_{1/2}$ plane

Mastercode

2015



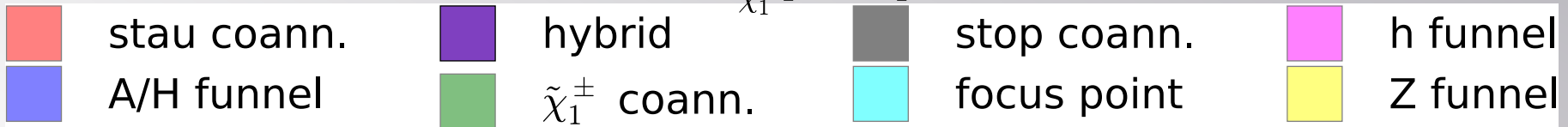
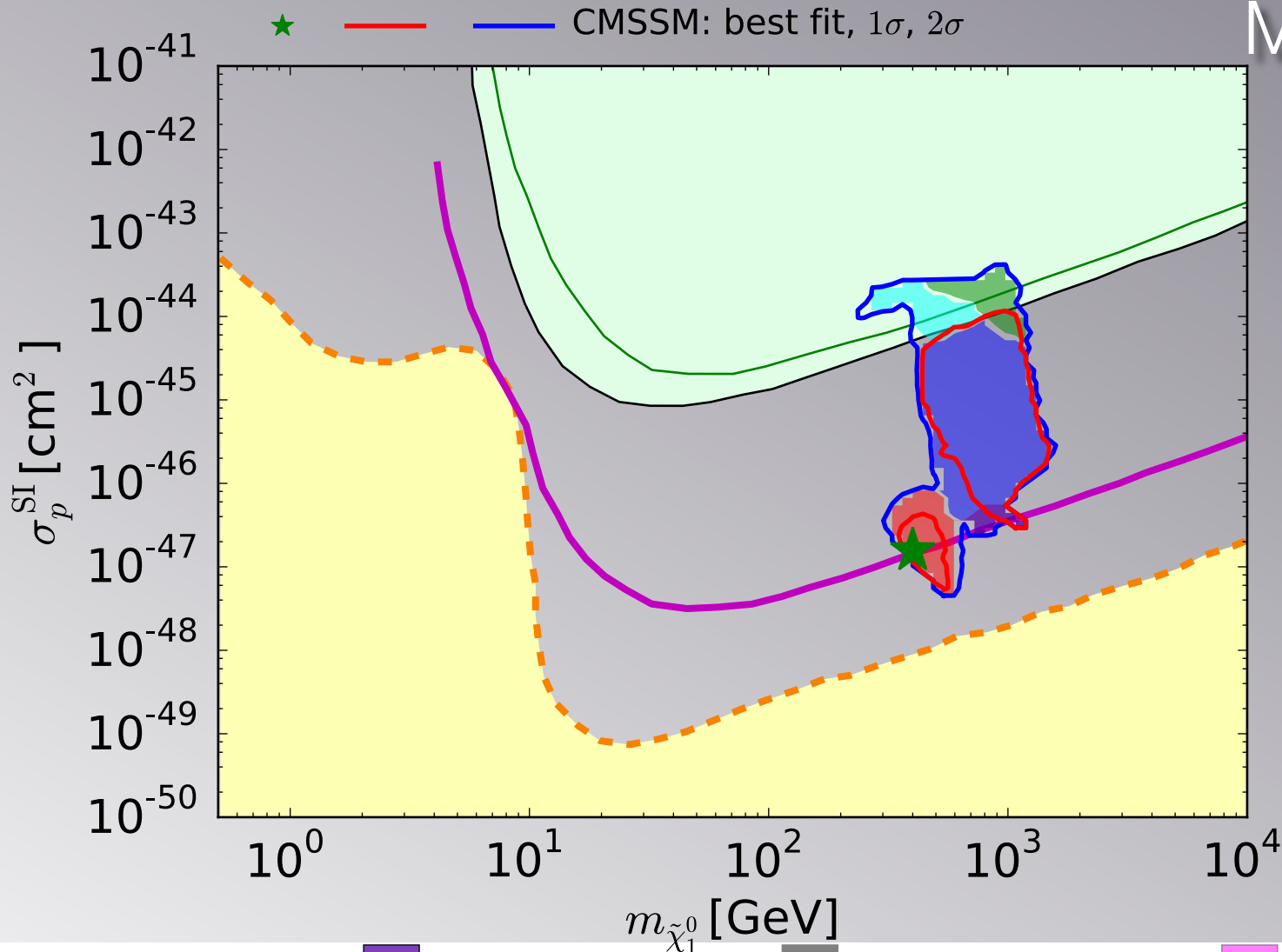
CMSSM

Bagnaschi, Buchmueller, Cavanaugh, Citron, De Roeck, Dolan,
 Ellis, Flacher, Heinemeyer, Isidori, Malik, Martinez Santos,
 Olive, Sakurai, de Vries, Weiglein

Elastic scattering cross-section

Mastercode

2015



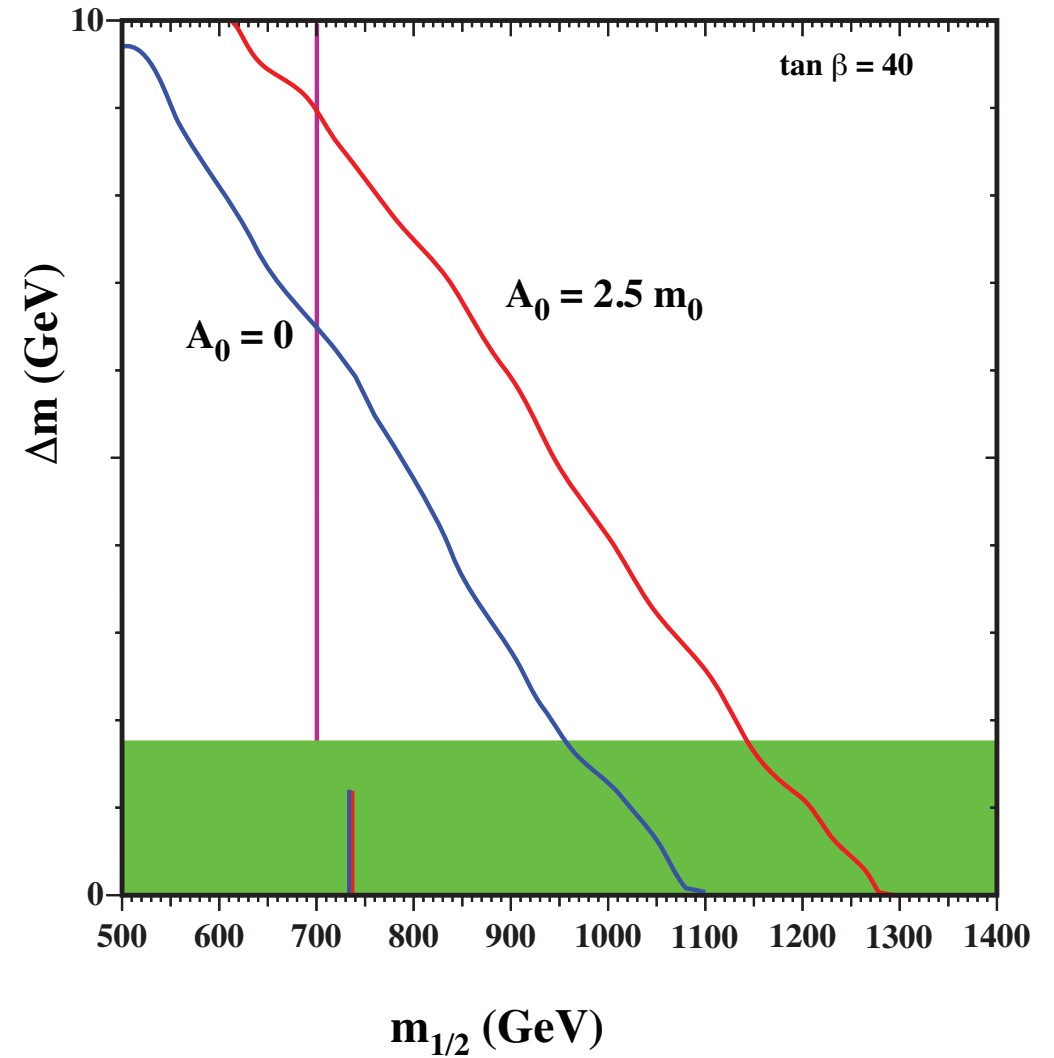
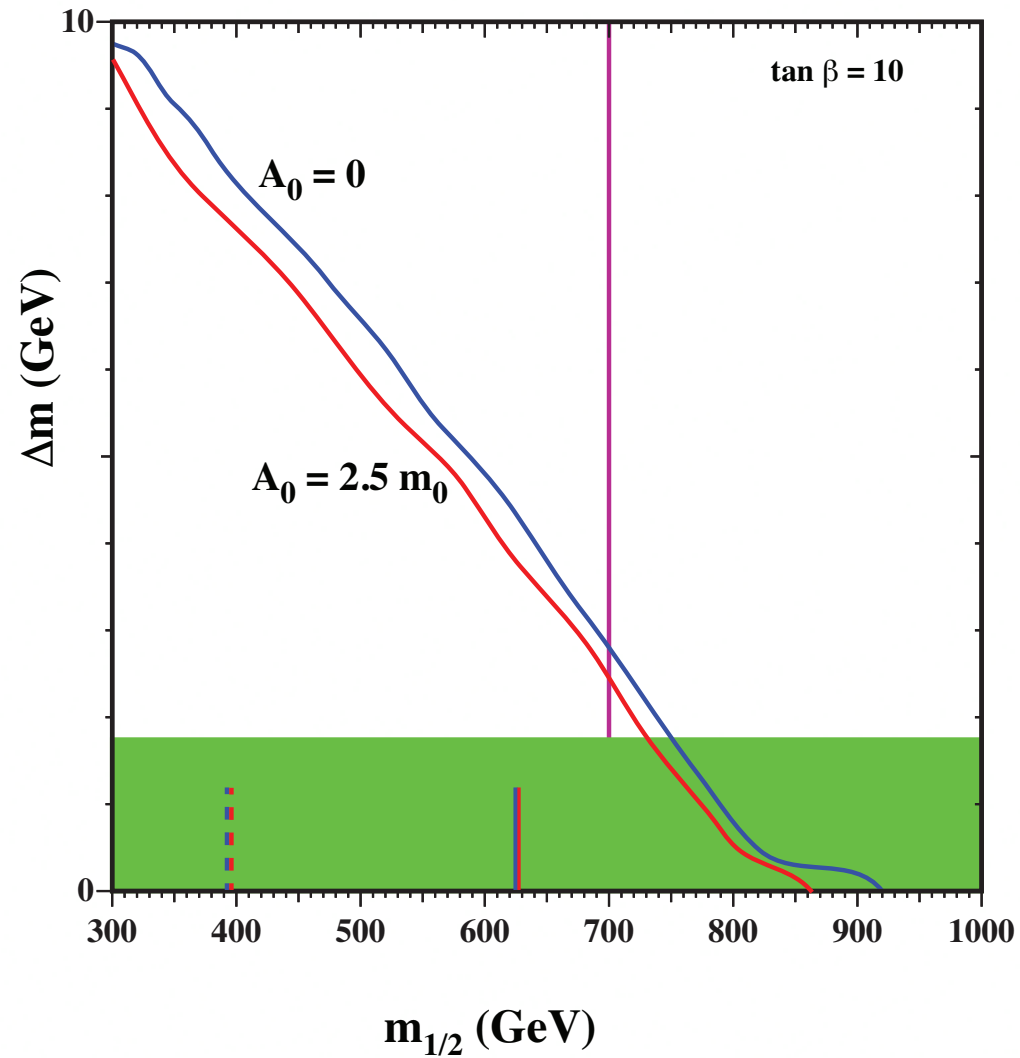
CMSSM

Bagnaschi, Buchmueller, Cavanaugh, Citron, De Roeck, Dolan, Ellis, Flacher, Heinemeyer, Isidori, Malik, Martinez Santos, Olive, Sakurai, de Vries, Weiglein

The Strips:

- ✦ Stau-coannihilation Strip
 - ✦ extends only out to ~ 1 TeV

Stau strip (end points)

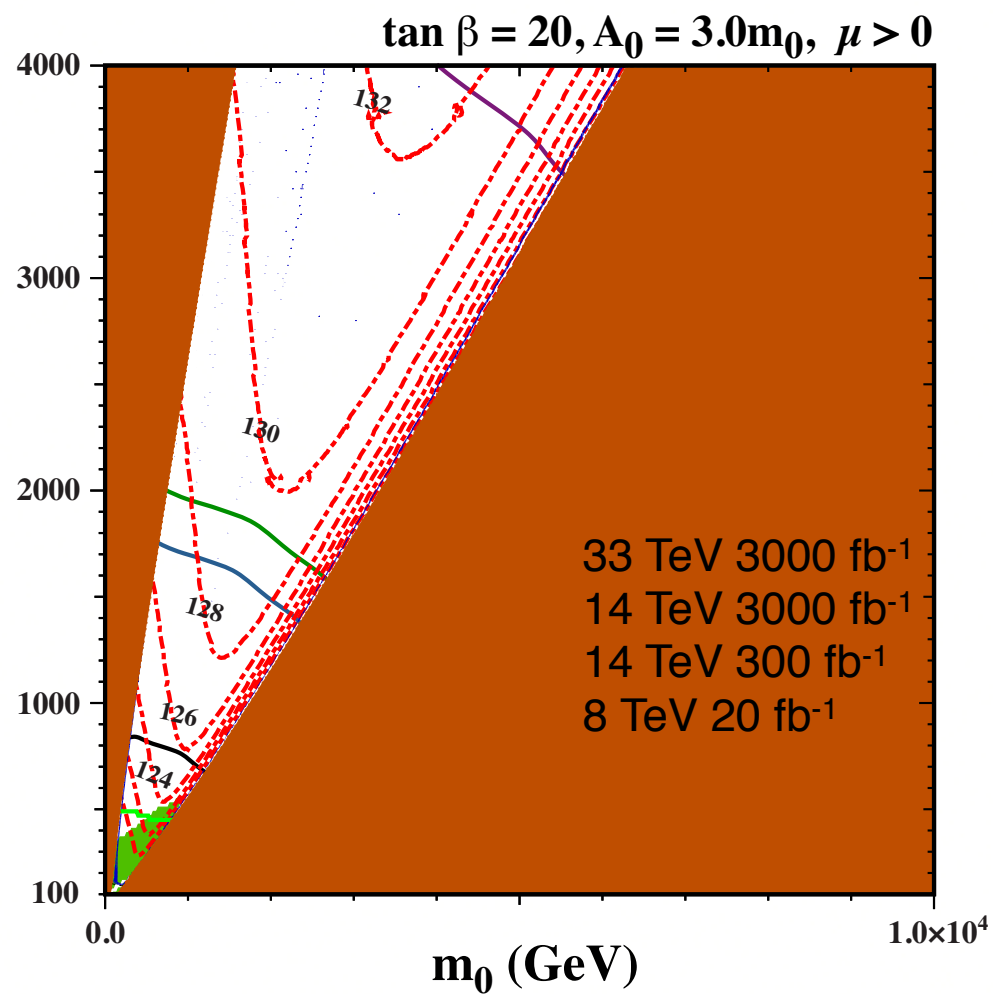
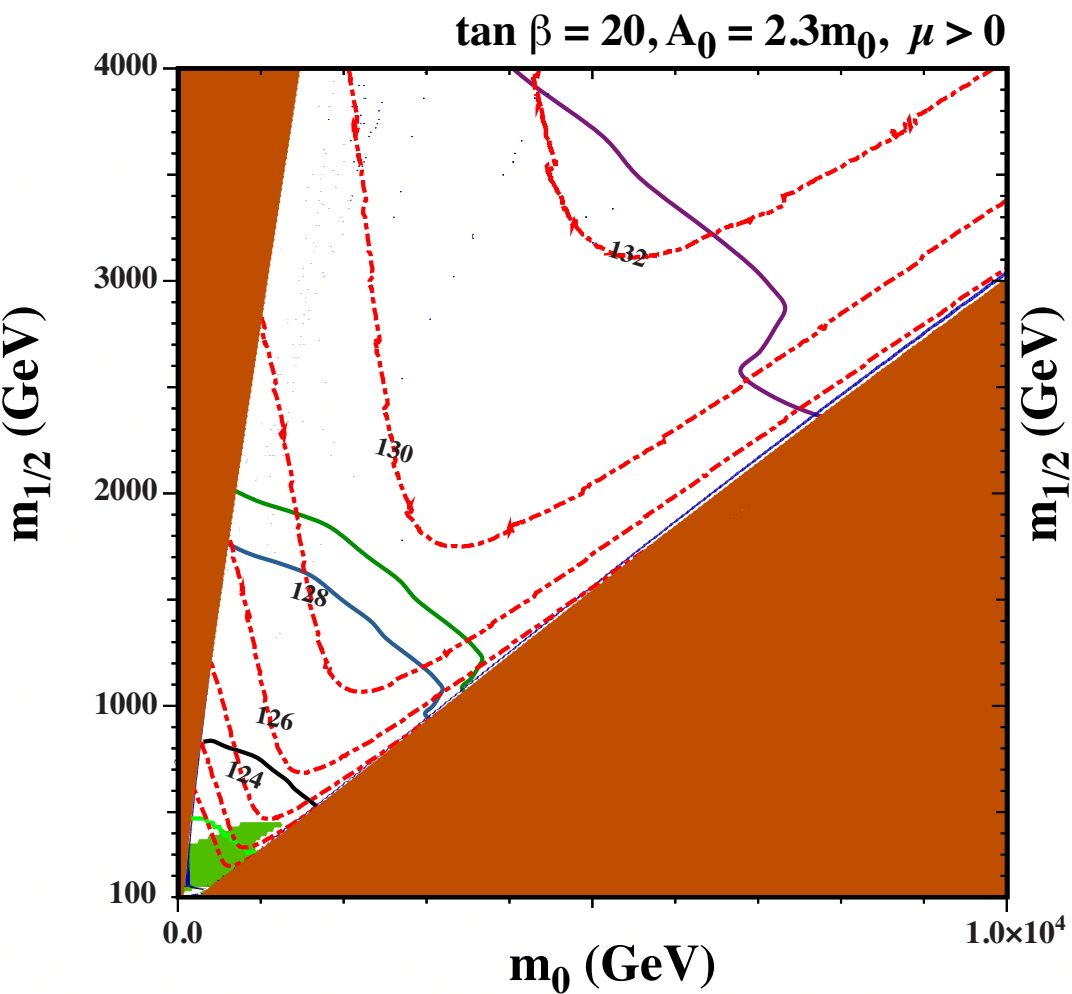


Citron, Ellis, Luo, Marrouche,
Olive, de Vries

The Strips:

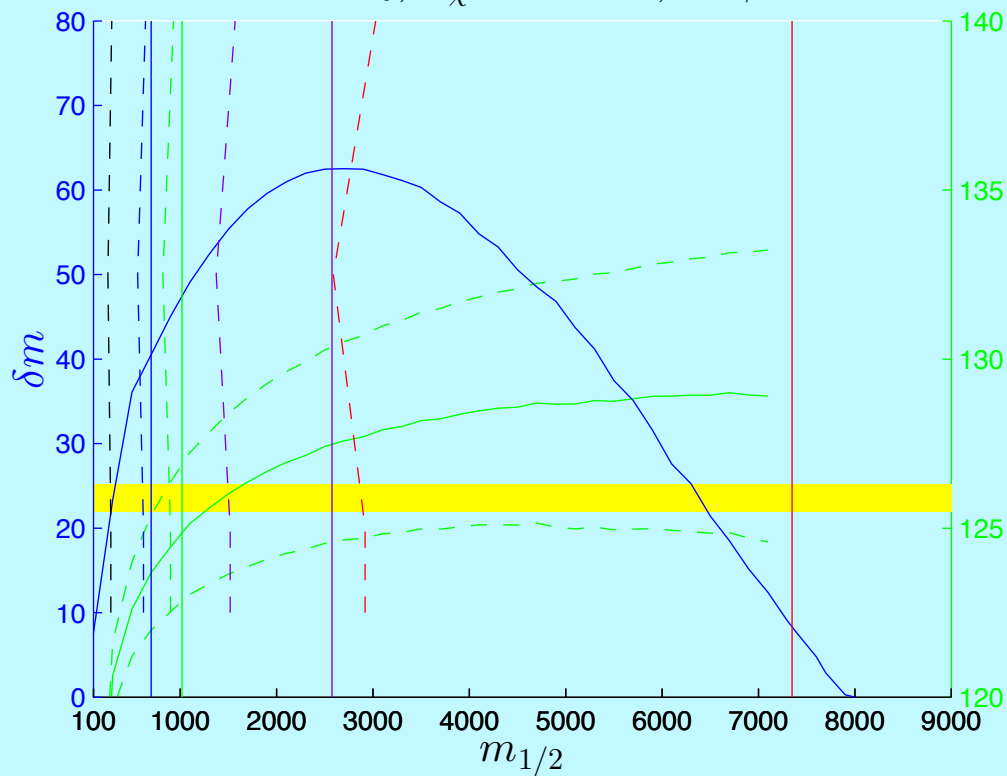
- ✦ Stau-coannihilation Strip
 - ✦ extends only out to ~ 1 TeV
- ✦ Stop-coannihilation Strip

Stop strip

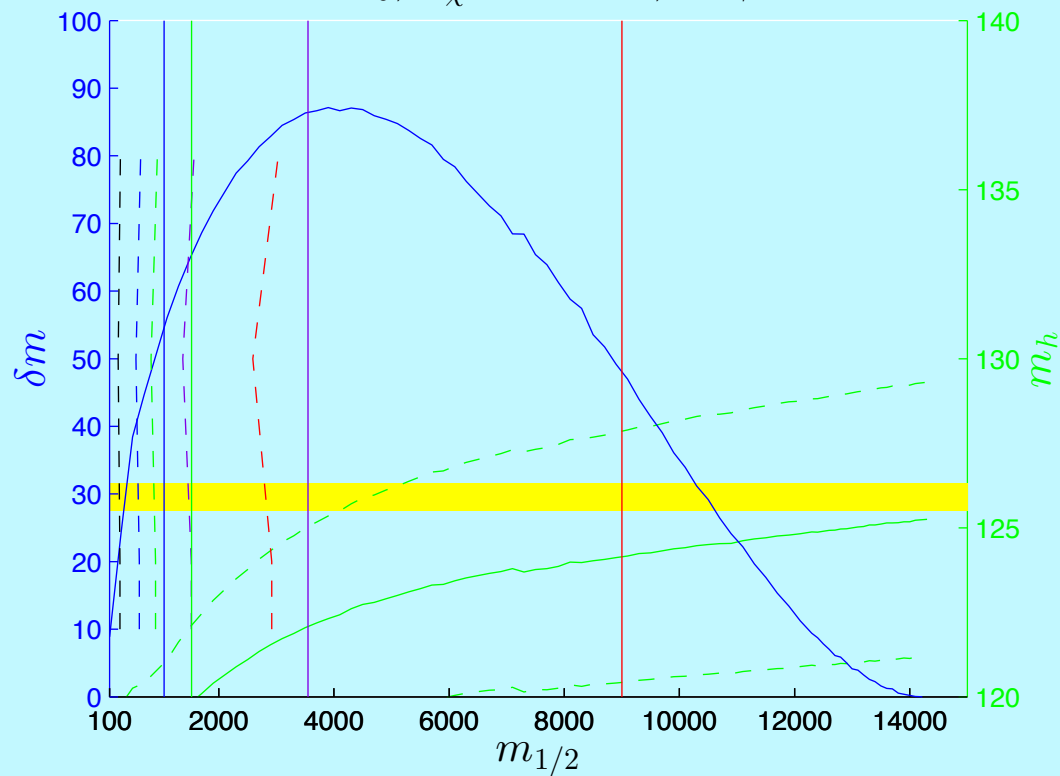


Stop strip

$A = 2.3m_0, \Omega_\chi h^2 = 0.12, \tan\beta = 20$



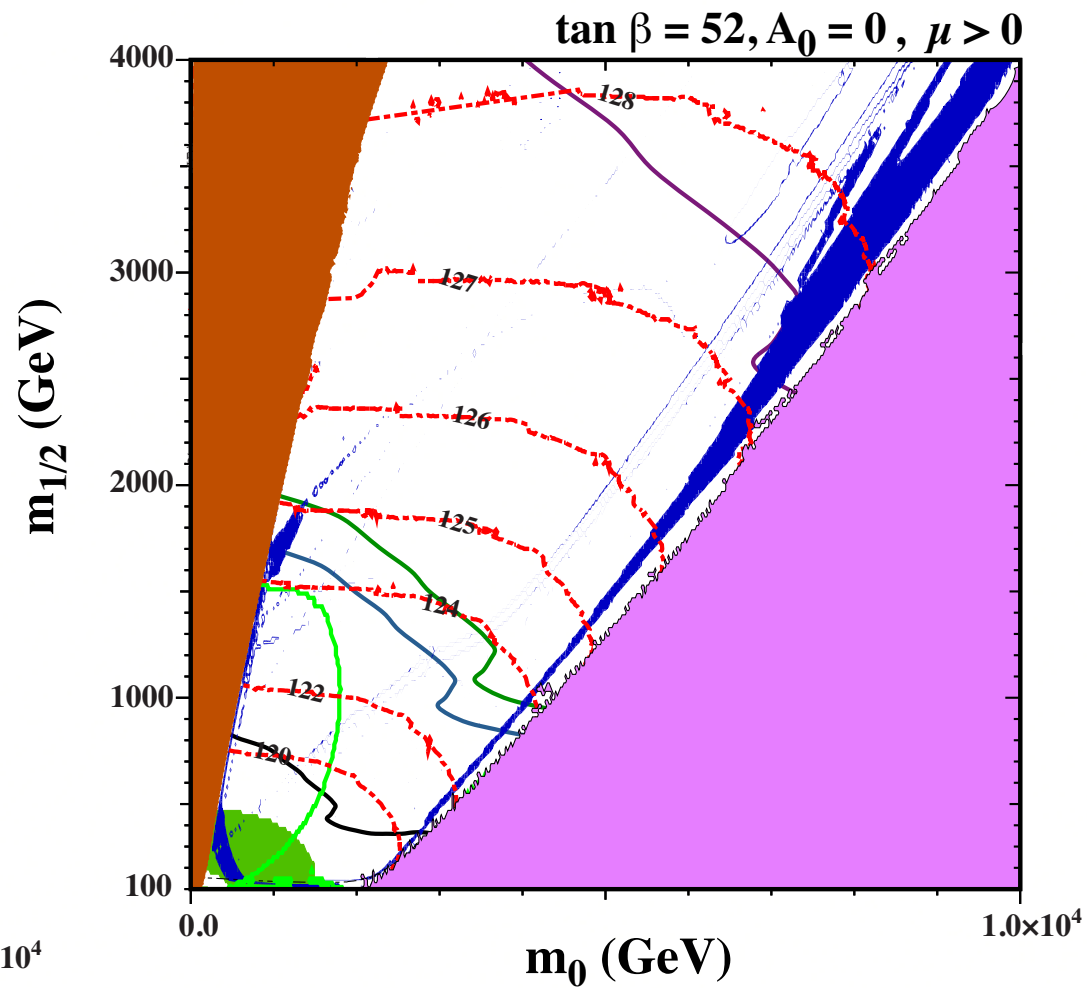
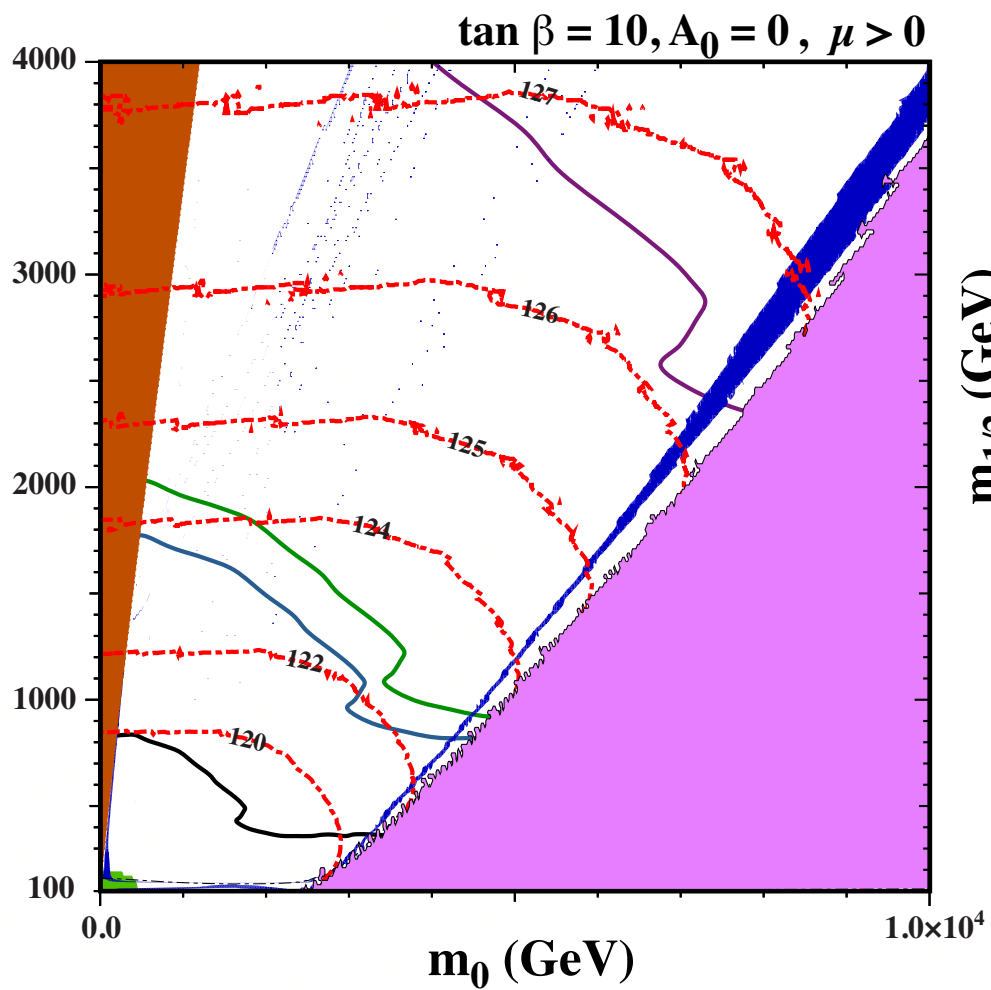
$A = 3m_0, \Omega_\chi h^2 = 0.12, \tan\beta = 20$



The Strips:

- ✦ Stau-coannihilation Strip
 - ✦ extends only out to ~ 1 TeV
- ✦ Stop-coannihilation Strip
- ✦ Funnel
 - ✦ associated with high $\tan \beta$, problems with $B \rightarrow \mu\mu$
- ✦ Focus Point

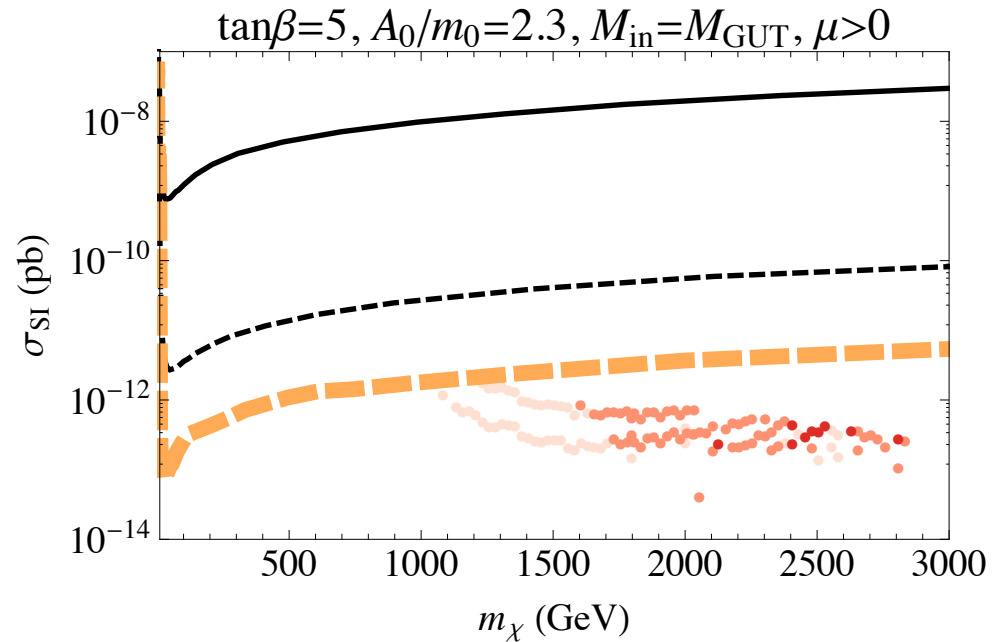
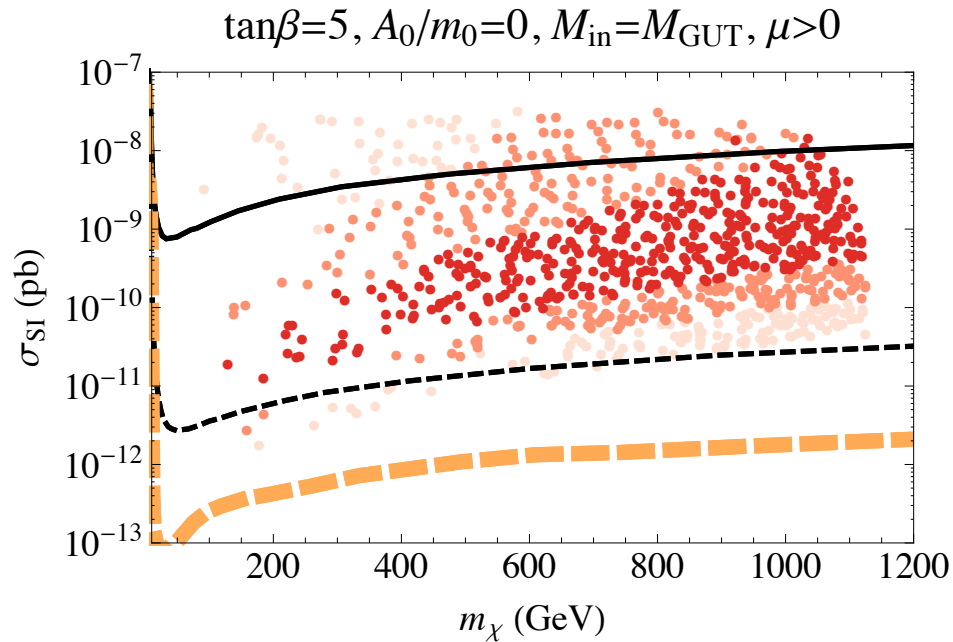
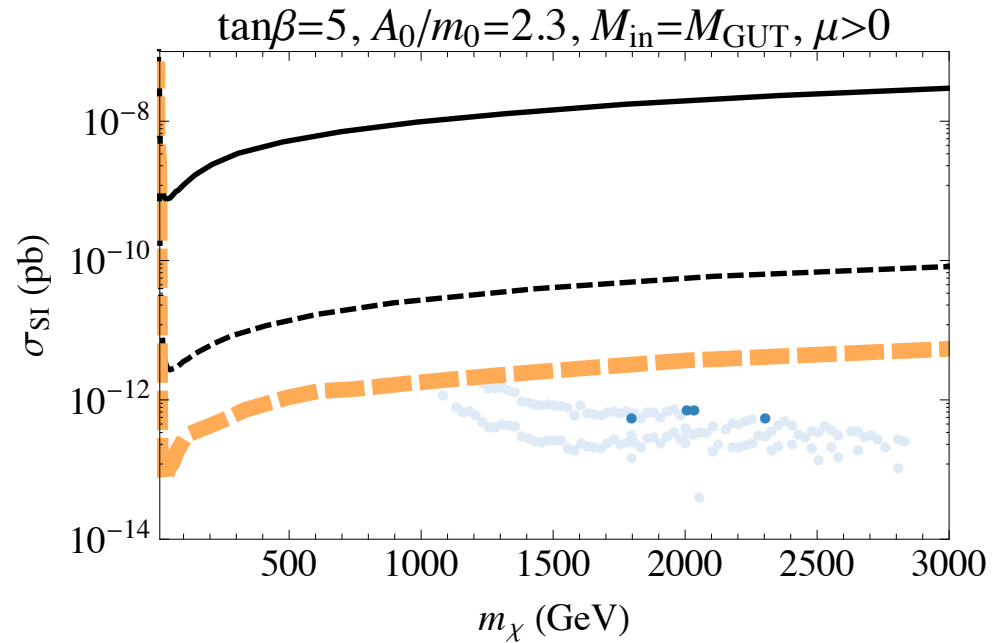
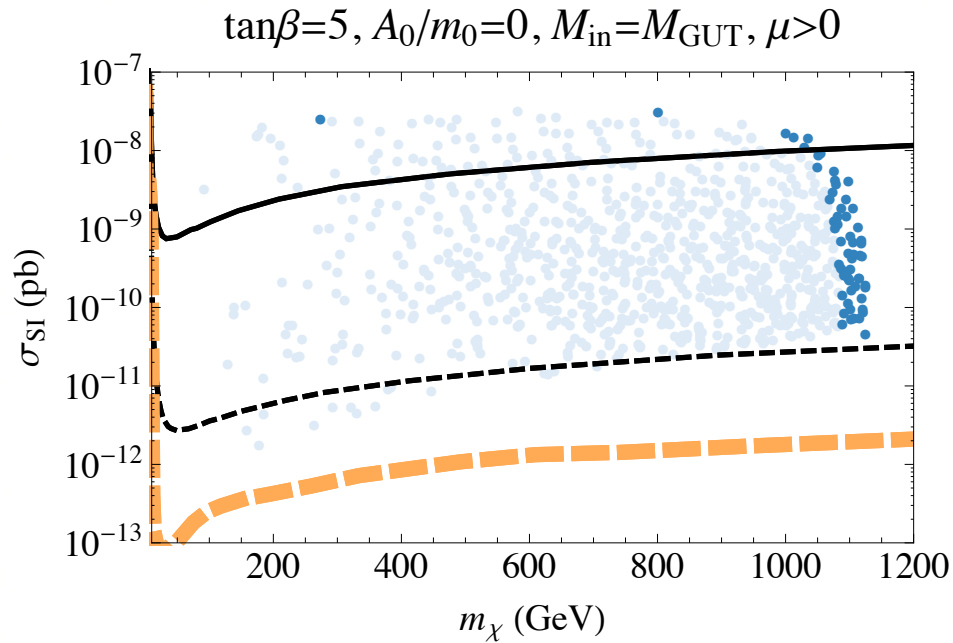
Focus Point



Buchmueller, Citron, Ellis, Guha, Marrouche, Olive, de Vries, Zheng

Ellis, Olive, Zheng

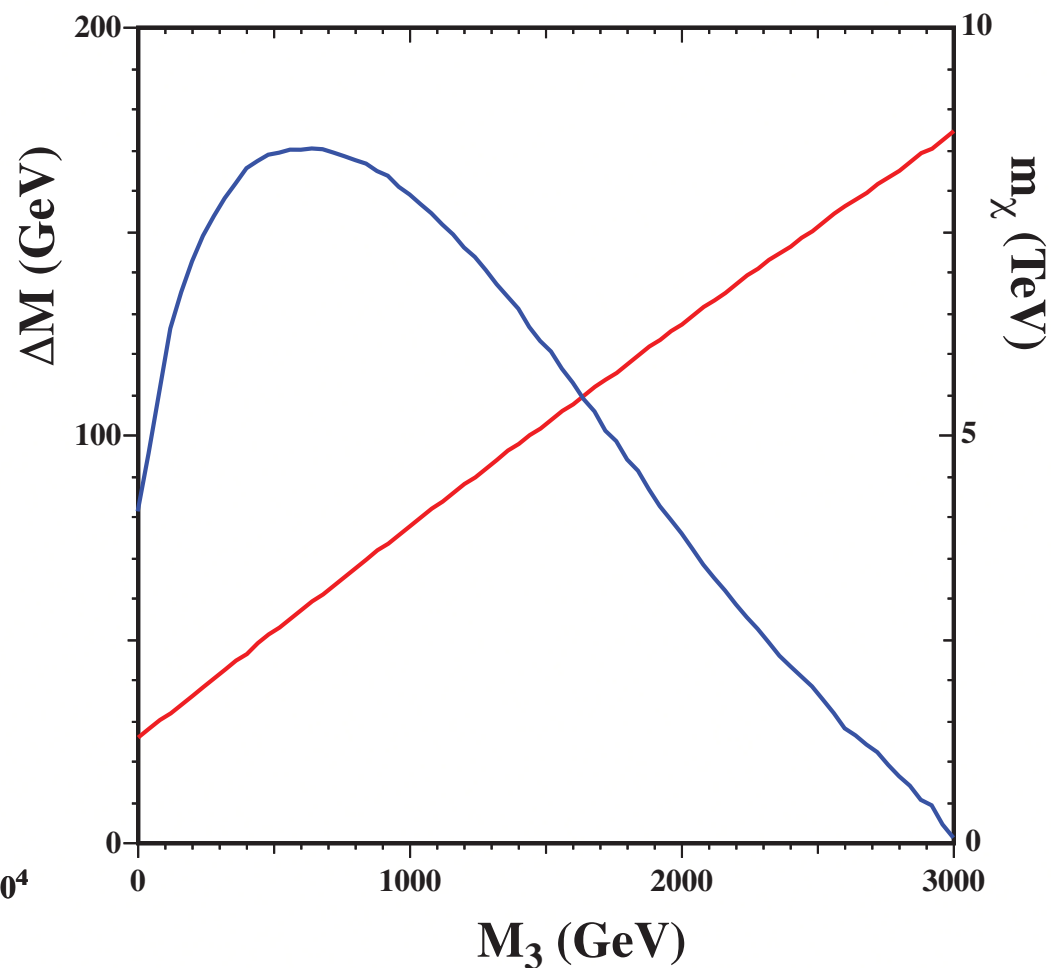
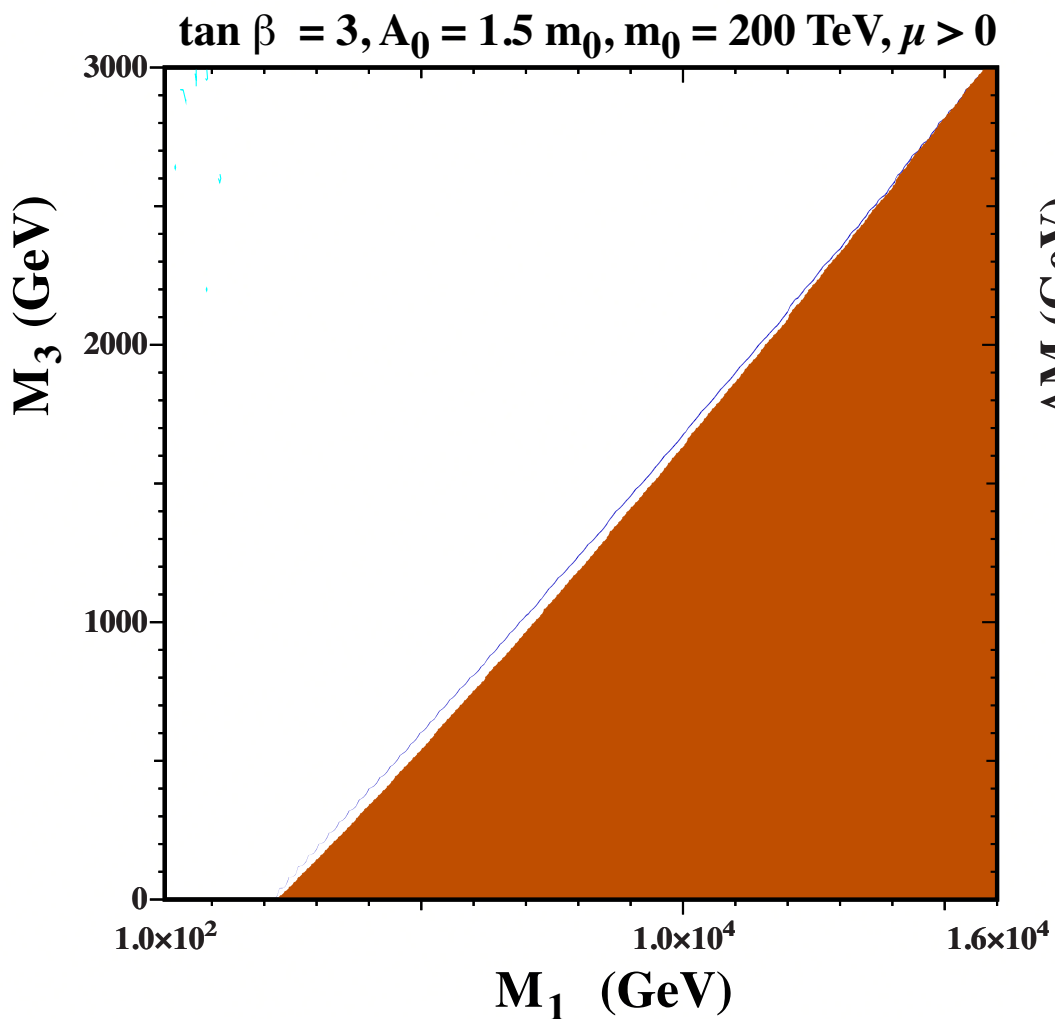
Direct detectability



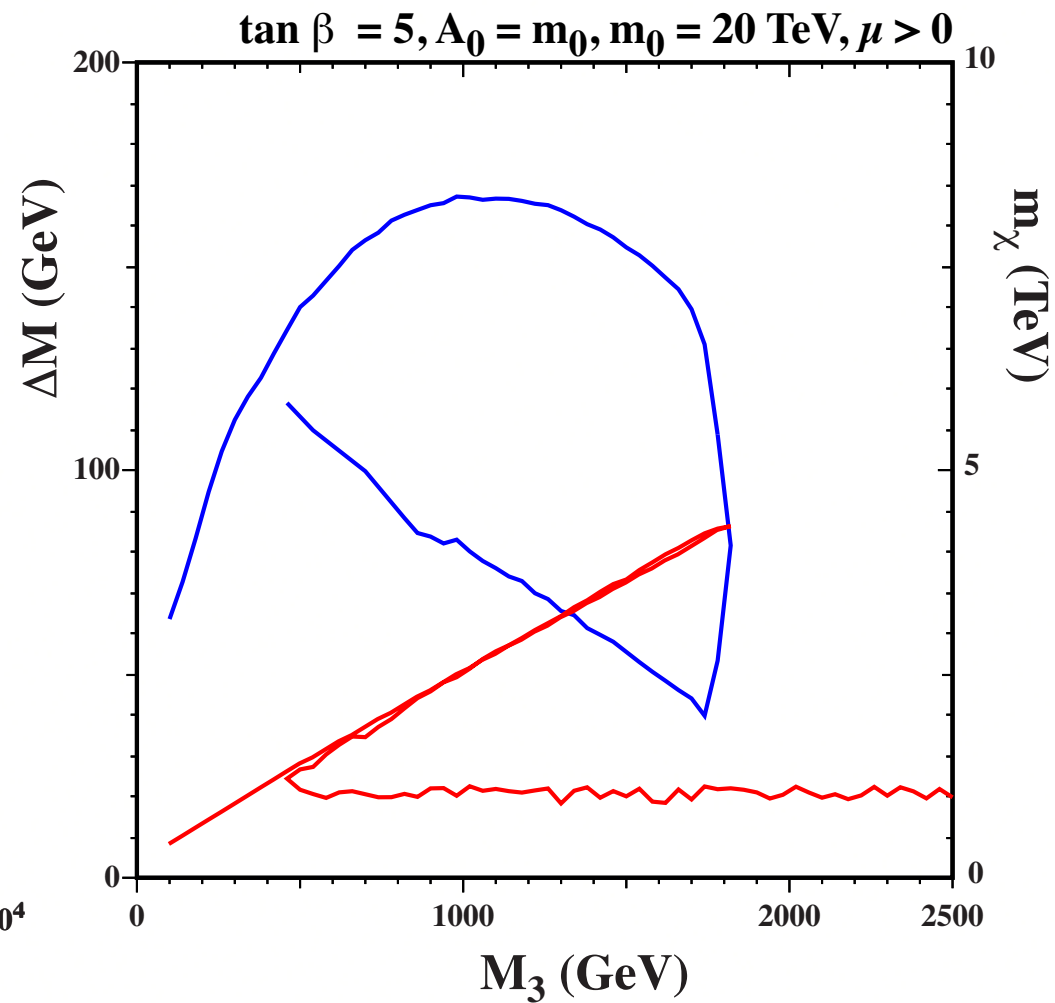
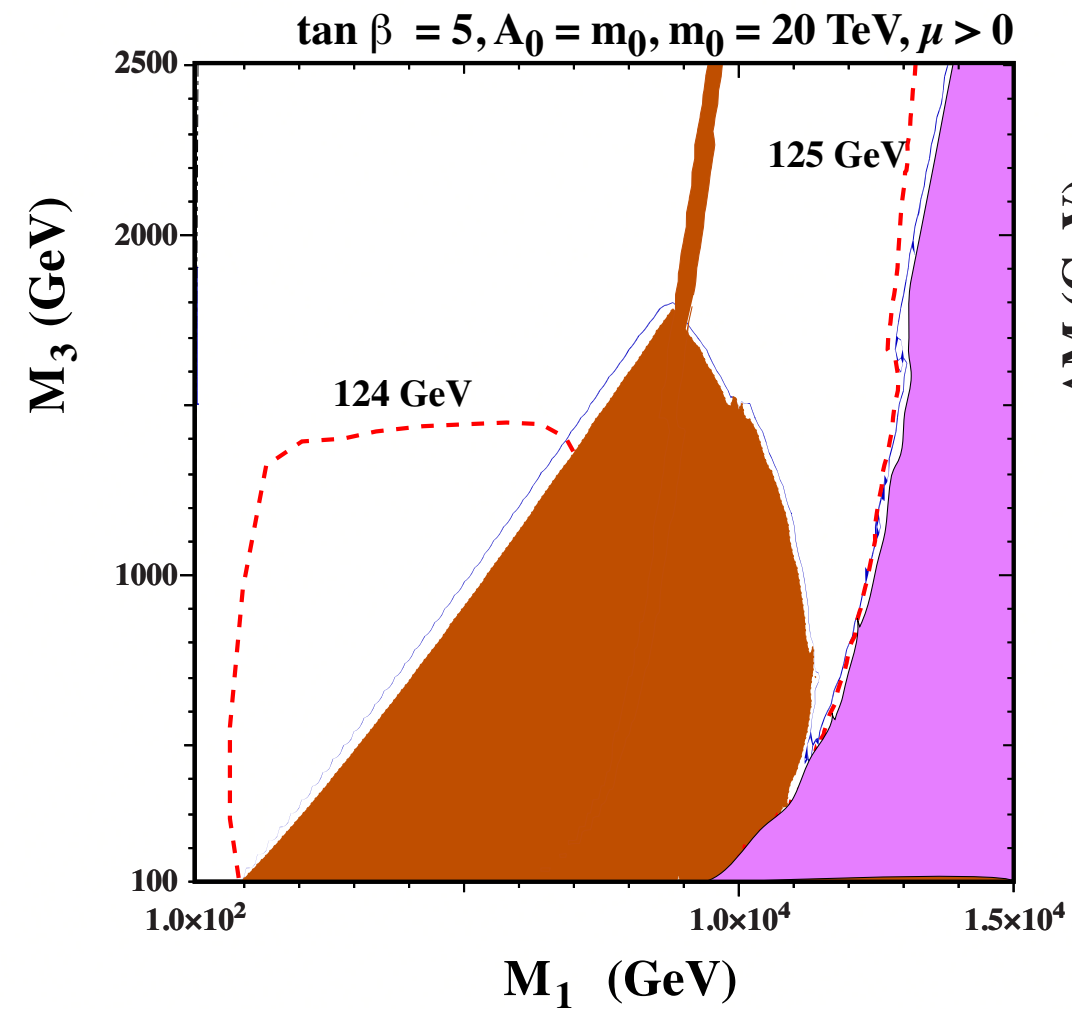
The Strips:

- ✦ Stau-coannihilation Strip
 - ✦ extends only out to ~ 1 TeV
- ✦ Stop-coannihilation Strip
- ✦ Funnel
 - ✦ associated with high $\tan \beta$, problems with $B \rightarrow \mu\mu$
- ✦ Focus Point
- ✦ Gluino-coannihilation Strip

Gluino Strips ($M_3 \neq M_1 = M_2$ @ M_{GUT})



Gluino Strips ($M_3 \neq M_1 = M_2$ @ M_{GUT})



May require more general models
which are concordant with LHC MET;
Higgs; and $B_s \rightarrow \mu^+\mu^-$; and Dark Matter

Other Possibilities

- ✦ NUHM1,2: $m_1^2 = m_2^2 \neq m_0^2$, $m_1^2 \neq m_2^2 \neq m_0^2$
 - ✦ μ and/or m_A free
- ✦ subGUT models: $M_{in} < M_{GUT}$
 - ✦ with or without mSUGRA

Why Supersymmetry (still)?

- ✦ Gauge Coupling Unification
- ✦ Gauge Hierarchy Problem
- ✦ Stabilization of the Electroweak Vacuum
- ✦ Radiative Electroweak Symmetry Breaking
- ✦ Dark Matter
- ✦ Improvement to low energy phenomenology?

but, $m_h \sim 126$ GeV, and no SUSY?

SO(10) GUT?

- ✦ Gauge Coupling Unification
- ✦
- ✦ Stabilization of the Electroweak Vacuum?
- ✦
- ✦ Dark Matter
- ✦ Improvement to low energy phenomenology?

Neutrino masses...

Recipe for constructing an SO(10) DM model

0. Get a copy of Slansky's review (Phys Rep 79 (1981) 1)
(or something equivalent)

Table 42
SO₁₀ tensor products

$$\begin{aligned}
 10 \times 10 &= 1_s + 45_a + 54_s \\
 \overline{16} \times 10 &= 16 + 144 \\
 16 \times 16 &= 10_s + 120_a + 126_s \\
 \overline{16} \times 16 &= 1 + 45 + 210 \\
 45 \times 10 &= 10 + 120 + 320 \\
 45 \times 16 &= 16 + 144 + 560 \\
 45 \times 45 &= 1_s + 45_a + 54_s + 210_s + 770_s + 945_a \\
 54 \times 10 &= 10 + 210' + 320 \\
 54 \times 16 &= 144 + 720 \\
 54 \times 45 &= 45 + 54 + 945 + 1386 \\
 54 \times 54 &= 1_s + 45_a + 54_s + 660_s + 770_s + 1386_a \\
 120 \times 10 &= 45 + 210 + 945 \\
 120 \times \overline{16} &= 16 + 144 + 560 + 1200 \\
 120 \times 45 &= 10 + 120 + 126 + 126 + 320 + 1728 + 2970 \\
 120 \times 54 &= 120 + 320 + 1728 + 4312 \\
 120 \times 120 &= 1_s + 45_a + 54_s + 210_s + 210_s + 770_s + 945_a + 1050_s + \overline{1050}_s + 4125_s + 5940_s \\
 126 \times 10 &= 210 + 1050 \\
 \overline{126} \times \overline{16} &= 144 + 672 + 1200 \\
 126 \times 16 &= 16 + 560 + 1440 \\
 126 \times 45 &= 120 + 126 + 1728 + 3696 \\
 126 \times 54 &= \overline{126} + 1728 + 4950 \\
 126 \times 120 &= 45 + 210 + 945 + 1050 + 5940 + 6930 \\
 126 \times 126 &= 54_s + 945_a + 1050_s + 2772_s + 4125_s + 6930_a \\
 \overline{126} \times 126 &= 1 + 45 + 210 + 770 + 5940 + 8910 \\
 \overline{144} \times 10 &= 16 + 144 + 560 + 720 \\
 \overline{144} \times 16 &= 45 + 54 + 210 + 945 + 1050 \\
 \overline{144} \times \overline{16} &= 10 + 120 + 126 + 320 + 1728 \\
 \overline{144} \times 45 &= 16 + 144_s + 144_s + 560 + 720 + 1200 + 3696' \\
 \overline{144} \times 54 &= 16 + 144 + 560 + 720 + 2640 + 3696' \\
 \overline{144} \times 120 &= 16 + 144_s + 144_s + 560_s + 560_s + 720 + 1200 + 1440 + 3696' + 8800 \\
 \overline{144} \times 126 &= 144 + 560 + 720 + 1200 + 1440 + 5280 + 8800 \\
 \overline{144} \times 126 &= 16 + 144 + 560 + 1200 + 1440 + 3696' + 11088 \\
 144 \times 144 &= 10_s + 120_{a_1} + 120_{a_2} + 126_s + \overline{126}_s + 210'_s + 320_s + 320_s + 1728_s + 1728_s + 2970_s + 3696_a + 4312_a + 4950_s \\
 \overline{144} \times 144 &= 1 + 45_s + 45_s + 54_s + 210_s + 210_s + 770_s + 945_s + 945_s + 1050_s + \overline{1050}_s + 1386_s + 5940 + 8085 \\
 210 \times 10 &= 120 + 126 + 126 + 1728 \\
 210 \times 16 &= 16 + 144 + 560 + 1200 + 1440 \\
 210 \times 45 &= 45 + 210_s + 210_s + 945 + 1050 + \overline{1050} + 5940 \\
 210 \times 54 &= 210 + 945 + \overline{1050} + 1050 + 8085 \\
 210 \times 120 &= 10 + 120_s + 120_s + 126 + \overline{126} + 320 + 1728_1 + 1728_2 + 2970 + 3696 + \overline{3696} + 10560 \\
 210 \times 126 &= 10 + 120 + 126 + 320 + 1728 + 2970 + 3696 + 6930' + 10560' + \overline{3696} + \overline{3696} + 10560 \\
 210 \times 144 &= 16 + 144_s + 144_s + 560_s + 560_s + 672 + 720 + 1200_1 + 1200_2 + 1440 + 3696' + 8800 + 11088 \\
 210 \times 210 &= 1_s + 45_s + 45_s + 54_s + 210_s + 210_s + 770_s + 945_{a_1} + 945_{a_2} + \overline{1050}_s + 1050_s + 4125_s + 5940_s + 5940_s + 6930_a + \overline{6930}_a + 8910_s \\
 210 \times 210 &= 1_s + 45_s + 45_s + 54_s + 210_s + 210_s + 770_s + 945_s + 945_s + 1050_s + 1050_s + 4125_s + 5940_s + 5940_s + 6930_s + \overline{6930}_s + 8910_s
 \end{aligned}$$

GROUP THEORY FOR UNIFIED MODEL BUILDING

R. SLANSKY

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Georgi, Nanopoulos; Vayonakis;
Masiero; Shafi, Sondermann, Wetterich;
del Aguila, Ibanez;
Mohapatra, Senjanovic;
Mambrini, Nagata,
Olive, Quevillon, Zheng;
Nagata, Olive, Zheng

Recipe for constructing an SO(10) DM model

0. Get a copy of Slansky's review (Phys Rep 79 (1981) 1)
(or something equivalent)

1. Pick an Intermediate Scale Gauge Group

$$SO(10) \xrightarrow{R_1} G_{\text{int}}$$

G_{int}	R_1
$SU(4)_C \otimes SU(2)_L \otimes SU(2)_R$	210
$SU(4)_C \otimes SU(2)_L \otimes SU(2)_R \otimes D$	54
$SU(4)_C \otimes SU(2)_L \otimes U(1)_R$	45
$SU(3)_C \otimes SU(2)_L \otimes SU(2)_R \otimes U(1)_{B-L}$	45
$SU(3)_C \otimes SU(2)_L \otimes SU(2)_R \otimes U(1)_{B-L} \otimes D$	210
$SU(3)_C \otimes SU(2)_L \otimes U(1)_R \otimes U(1)_{B-L}$	45, 210
$SU(5) \otimes U(1)$	45, 210
Flipped $SU(5) \otimes U(1)$	45, 210

Recipe for constructing an SO(10) DM model

0. Get a copy of Slansky's review (Phys Rep 79 (1981) 1) (or something equivalent)
1. Pick an Intermediate Scale Gauge Group
2. Use **126** to break G_{int} to SM

$$SO(10) \xrightarrow{R_1} G_{\text{int}} \xrightarrow{R_2} G_{\text{SM}} \otimes \mathbb{Z}_2$$

$$R_2 = \mathbf{126} + \dots$$

Neutrino see-saw: Majorana mass for ν_R from $16 \ 16 \ 126 \rightarrow m_{\nu R} \sim M_{\text{int}}$

Recipe for constructing an SO(10) DM model

0. Get a copy of Slansky's review (Phys Rep 79 (1981) 1) (or something equivalent)
1. Pick an Intermediate Scale Gauge Group
2. Use **126** to break G_{int} to SM
3. Pick DM representation and insure proper splitting within the multiplet, and pick low energy field content

Remnant Z_2 symmetry

Fermions from **10, 45, 54, 120, 126, or 210** representations;

Scalars from **16, 144**

Kadastik, Kannike, Raidal;
Frigerio, Hambye;
Mambrini, Nagata,
Olive, Quevillon, Zheng;
Nagata, Olive, Zheng

Model	$B - L$	$SU(2)_L$	Y	SO(10) representations
F_1^0	0	1	0	45, 54, 210
$F_2^{1/2}$		2	1/2	10, 120, 126, 210'
F_3^0		3	0	45, 54, 210
F_3^1		3	1	54
$F_4^{1/2}$		4	1/2	210'
$F_4^{3/2}$		4	3/2	210'
S_1^0	1	1	0	16, 144
$S_2^{1/2}$		2	1/2	16, 144
S_3^0		3	0	144
S_3^1		3	1	144
\widehat{F}_1^0	2	1	0	126
$\widehat{F}_2^{1/2}$		2	1/2	210
\widehat{F}_3^1		3	1	126

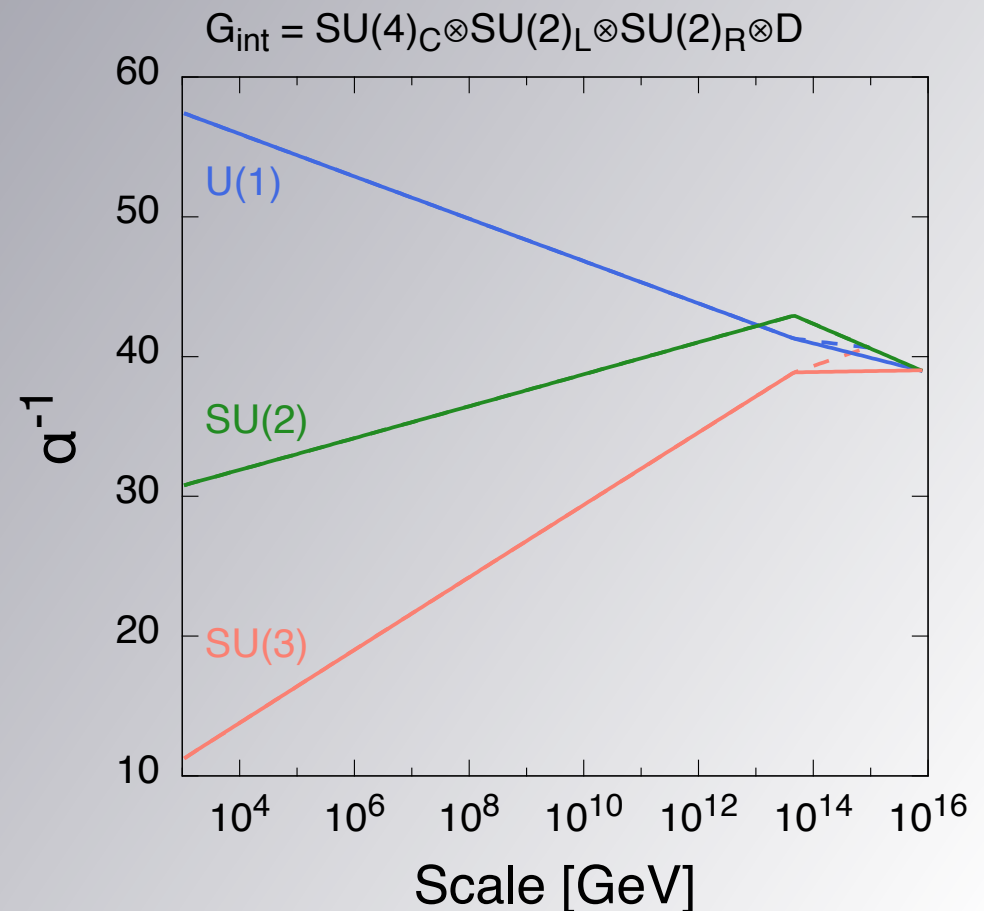
Recipe for constructing an SO(10) DM model

0. Get a copy of Slansky's review (Phys Rep 79 (1981) 1) (or something equivalent)
1. Pick an Intermediate Scale Gauge Group
2. Use **126** to break G_{int} to SM
3. Pick DM representation and insure proper splitting within the multiplet, and pick low energy field content
4. Use RGEs to obtain Gauge Coupling Unification

Recipe for constructing an SO(10) DM model

4. Use RGEs to obtain Gauge Coupling Unification

Fixes M_{GUT} , M_{int} , α_{GUT}



Examples:

Scalars

Model	R_{DM}	S_n^Y	SO(10) representation
$G_{\text{int}} = \text{SU}(4)_C \otimes \text{SU}(2)_L \otimes \text{SU}(2)_R (\otimes D)$			
SA _{422(D)}	4, 1, 2	S_1^0	16, 144
SB _{422(D)}	4, 2, 1	$S_2^{1/2}$	16, 144
SC _{422(D)}	4, 2, 3	$S_2^{1/2}$	144
SD _{422(D)}	4, 3, 2	S_3^1	144
SE _{422(D)}	4, 3, 2	S_3^0	144
$G_{\text{int}} = \text{SU}(4)_C \otimes \text{SU}(2)_L \otimes \text{U}(1)_R$			
SA ₄₂₁	4, 1, -1/2	S_1^0	16, 144
SB ₄₂₁	4, 2, 0	$S_2^{1/2}$	16, 144
SC ₄₂₁	4, 2, 1	$S_2^{1/2}$	144
SD ₄₂₁	4, 3, 1/2	S_3^1	144
SE ₄₂₁	4, 3, -1/2	S_3^0	144
$G_{\text{int}} = \text{SU}(3)_C \otimes \text{SU}(2)_L \otimes \text{SU}(2)_R \otimes \text{U}(1)_{B-L} (\otimes D)$			
SA _{3221(D)}	1, 1, 2, 1	S_1^0	16, 144
SB _{3221(D)}	1, 2, 1, -1	$S_2^{1/2}$	16, 144
SC _{3221(D)}	1, 2, 3, -1	$S_2^{1/2}$	144
SD _{3221(D)}	1, 3, 2, 1	S_3^1	144
SE _{3221(D)}	1, 3, 2, 1	S_3^0	144

Examples:

Scalars

Higgs portal models
Inert Higgs doublet models

Model	$\log_{10} M_{\text{GUT}}$	$\log_{10} M_{\text{int}}$	α_{GUT}	$\log_{10} \tau_p(p \rightarrow e^+ \pi^0)$
$G_{\text{int}} = \text{SU}(4)_C \otimes \text{SU}(2)_L \otimes \text{SU}(2)_R$				
SA ₄₂₂	16.33	11.08	0.0218	36.8 ± 1.2
SB ₄₂₂	15.62	12.38	0.0228	34.0 ± 1.2
$G_{\text{int}} = \text{SU}(3)_C \otimes \text{SU}(2)_L \otimes \text{SU}(2)_R \otimes \text{U}(1)_{B-L}$				
SA ₃₂₂₁	16.66	8.54	0.0217	38.1 ± 1.2
SB ₃₂₂₁	16.17	9.80	0.0223	36.2 ± 1.2
SC ₃₂₂₁	15.62	9.14	0.0230	34.0 ± 1.2
$G_{\text{int}} = \text{SU}(3)_C \otimes \text{SU}(2)_L \otimes \text{SU}(2)_R \otimes \text{U}(1)_{B-L} \otimes D$				
SA _{3221D}	15.58	10.08	0.0231	33.8 ± 1.2
SB _{3221D}	15.40	10.44	0.0233	33.1 ± 1.2

other models have M_{GUT} too low

mass splitting:

$$\begin{aligned}
 -\mathcal{L}_{\text{int}} = & M^2 |R_{\text{DM}}|^2 + \kappa_1 R_{\text{DM}}^* R_{\text{DM}} R_1 + \{\kappa_2 R_{\text{DM}} R_{\text{DM}} R_2^* + \text{h.c.}\} \\
 & + \lambda_1^{\mathbf{1}} |R_{\text{DM}}|^2 |R_1|^2 + \lambda_2^{\mathbf{1}} |R_{\text{DM}}|^2 |R_2|^2 + \{\lambda_{12}^{\mathbf{126}} (R_{\text{DM}} R_{\text{DM}})_{\mathbf{126}} (R_1 R_2^*)_{\overline{\mathbf{126}}} + \text{h.c.}\} \\
 & + \lambda_1^{\mathbf{45}} (R_{\text{DM}}^* R_{\text{DM}})_{\mathbf{45}} (R_1^* R_1)_{\mathbf{45}} + \lambda_1^{\mathbf{210}} (R_{\text{DM}}^* R_{\text{DM}})_{\mathbf{210}} (R_1^* R_1)_{\mathbf{210}} \\
 & + \lambda_2^{\mathbf{45}} (R_{\text{DM}}^* R_{\text{DM}})_{\mathbf{45}} (R_2^* R_2)_{\mathbf{45}} + \lambda_2^{\mathbf{210}} (R_{\text{DM}}^* R_{\text{DM}})_{\mathbf{210}} (R_2^* R_2)_{\mathbf{210}} ,
 \end{aligned}$$

Examples:

SM Fermion Singlets: Produced thermally out of equilibrium
 \Rightarrow Fermionic candidates (NETDM)

Mambrini, Olive, Quevillon, Zaldivar

To aid in gauge coupling unification, DM should be a SM singlet but a non-singlet under the Intermediate gauge group. Requires splitting multiplets

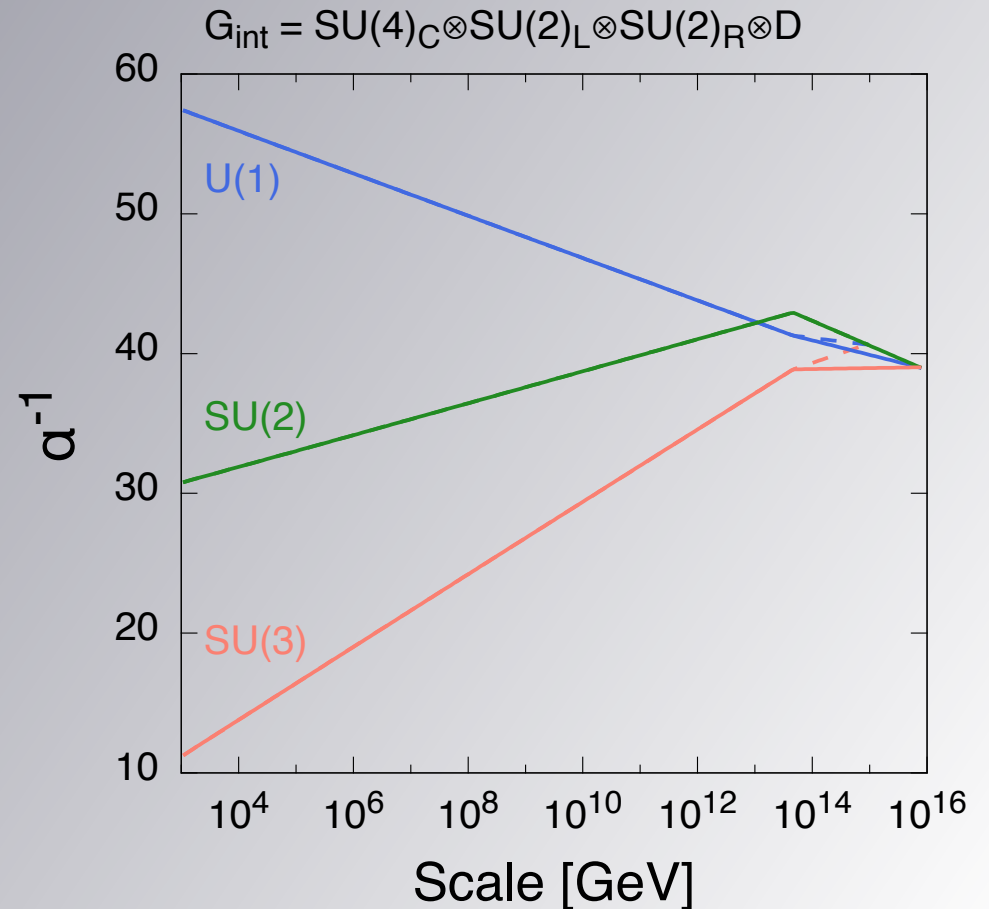
G_{int}	R_{DM}	SO(10)
$SU(4)_C \otimes SU(2)_L \otimes SU(2)_R$	(1, 1, 3)	45
	(15, 1, 1)	45, 210
	(10, 1, 3)	126
	(15, 1, 3)	210
$SU(4)_C \otimes SU(2)_L \otimes U(1)_R$	(15, 1, 0)	45, 210
	(10, 1, 1)	126
$SU(3)_C \otimes SU(2)_L \otimes SU(2)_R \otimes U(1)_{B-L}$	(1, 1, 3, 0)	45, 210
	(1, 1, 3, -2)	126
$SU(3)_C \otimes SU(2)_L \otimes U(1)_R \otimes U(1)_{B-L}$	(1, 1, 1, -2)	126

Mambrini, Nagata,
 Olive, Quevillon, Zheng

Gauge Coupling Unification

Also fix particle content below M_{GUT}

$SU(4)_C \otimes SU(2)_L \otimes SU(2)_R$				
R_{DM}	R_2	$\log_{10}(M_{\text{int}})$	$\log_{10}(M_{\text{GUT}})$	g_{GUT}
$(1, 1, 3)_W$	$(10, 1, 3)_C$ $(1, 1, 3)_R$	10.8	15.9	0.53
$(1, 1, 3)_D$	$(10, 1, 3)_C$ $(1, 1, 3)_R$	9.8	15.7	0.53
$SU(4)_C \otimes SU(2)_L \otimes SU(2)_R \otimes D$				
R_{DM}	R_2	$\log_{10}(M_{\text{int}})$	$\log_{10}(M_{\text{GUT}})$	g_{GUT}
$(15, 1, 1)_W$	$(10, 1, 3)_C$ $(\overline{10}, 3, 1)_C$ $(15, 1, 1)_R$	13.7	16.2	0.56
$(15, 1, 1)_W$	$(10, 1, 3)_C$ $(\overline{10}, 3, 1)_C$ $(15, 1, 3)_R$ $(15, 3, 1)_R$	14.2	15.5	0.56
$(15, 1, 1)_D$	$(10, 1, 3)_C$ $(\overline{10}, 3, 1)_C$ $(15, 1, 3)_R$ $(15, 3, 1)_R$	14.4	16.3	0.58
$SU(3)_C \otimes SU(2)_L \otimes SU(2)_R \otimes U(1)_{B-L}$				
R_{DM}	R_2	$\log_{10}(M_{\text{int}})$	$\log_{10}(M_{\text{GUT}})$	g_{GUT}
$(1, 1, 3, 0)_W$	$(1, 1, 3, -2)_C$ $(1, 1, 3, 0)_R$	6.1	16.6	0.52



Examples:

SM Fermion Singlets: Produced thermally out of equilibrium
 \Rightarrow Fermionic candidates (NETDM)

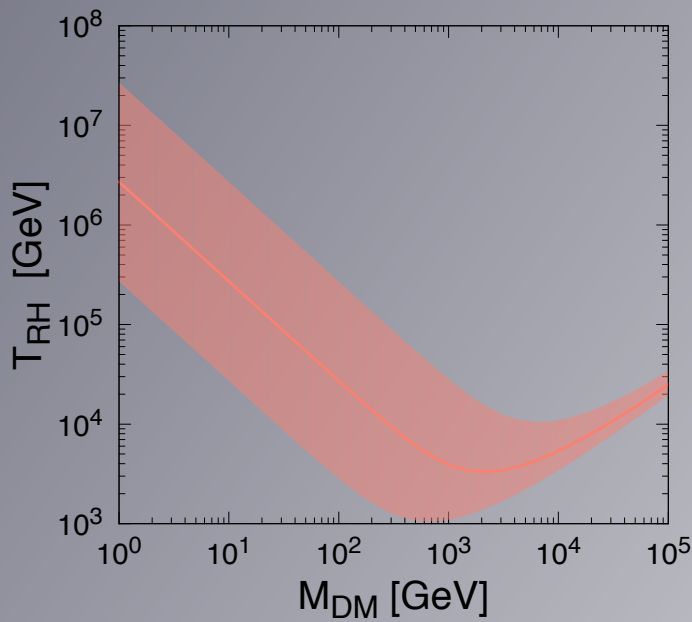
For mass splittings:

$$\mathcal{L}_{\text{int}} = -\frac{M_{45_W}}{2} 45_W 45_W - \frac{y_{54}}{2} 45_W 45_W 54_R - \frac{y_{210}}{2} 45_W 45_W 210_R + \text{h.c.} ,$$

$$15 = 8 + 3 + \bar{3} + 1$$

$$M_{15} \sim M_{45_W} - y_{54} v_{54} \sim M_{\text{int}};$$

$$M_1 \sim M_{15} - y_{210} v_{210}$$



G_{int}	$SU(4)_C \otimes SU(2)_L \otimes SU(2)_R \otimes D$
R_{DM}	$(15, 1, 1)_W$ in 45_W
R_1	54_R
R_2	$(10, 1, 3)_C \oplus (10, 3, 1)_C \oplus (15, 1, 1)_R$
$\log_{10}(M_{\text{int}})$	13.664(5)
$\log_{10}(M_{\text{GUT}})$	15.87(2)
g_{GUT}	0.5675(2)

$$(10, 1, 3)_C, (10, 3, 1)_C \in \mathbf{126}; \quad (15, 1, 1)_R \in \mathbf{210}$$

Examples:

Non-Singlets: Fermions

Model	$B - L$	$SU(2)_L$	Y	SO(10) representations
F_1^0		1	0	45, 54, 210
$F_2^{1/2}$		2	1/2	10, 120, 126, 210'
F_3^0	0	3	0	45, 54, 210
F_3^1		3	1	54
$F_4^{1/2}$		4	1/2	210'
$F_4^{3/2}$		4	3/2	210'

SO(10) representation	$SU(4)_C \otimes SU(2)_L \otimes SU(2)_R$
45	(1, 3, 1)
54	SM Triplets (Wino) (1, 3, 3)
210	(15, 3, 1)

SO(10) representation	$SU(4)_C \otimes SU(2)_L \otimes SU(2)_R$	$B - L$
10, 120, 210'	(1, 2, 2)	0
120, 126	(15, 2, 2)	0
210	SM Doublets (Higgsino) (10, 2, 2) \oplus ($\overline{10}$, 2, 2)	± 2
210'	(1, 4, 4)	0
54, 210	(1, 1, 1)	0
45	SM Singlets for mixing (1, 1, 3)	0
45, 210	(Bino) (15, 1, 1)	0
210	(15, 1, 3)	0
126	(10, 1, 3)	2

Examples:

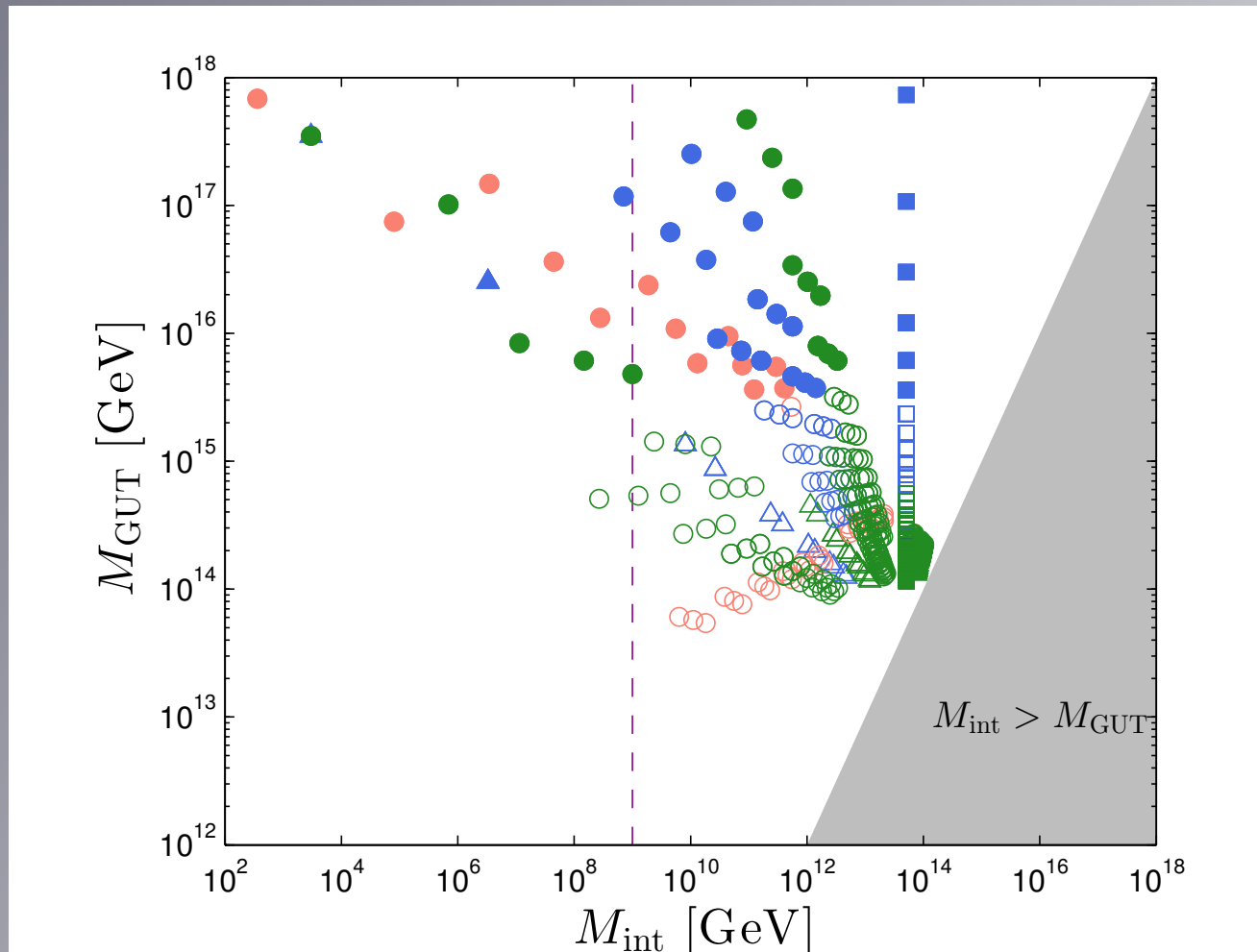
Non-Singlets: Fermions

R_{DM}	Additional Higgs in R_1	$\log_{10} M_{\text{int}}$	$\log_{10} M_{\text{GUT}}$	α_{GUT}	$\log_{10} \tau_p(p \rightarrow e^+ \pi^0)$		
$G_{\text{int}} = \text{SU}(4)_C \otimes \text{SU}(2)_L \otimes \text{SU}(2)_R$							
$(\mathbf{1}, \mathbf{3}, \mathbf{1})$	$(\mathbf{15}, \mathbf{1}, \mathbf{1})$ $(\mathbf{15}, \mathbf{1}, \mathbf{3})$	6.54	17.17	0.0252	39.8 ± 1.2		
Model	R_{DM}	R'_{DM}	Higgs	$\log_{10} M_{\text{int}}$	$\log_{10} M_{\text{GUT}}$	α_{GUT}	$\log_{10} \tau_p$
$G_{\text{int}} = \text{SU}(4)_C \otimes \text{SU}(2)_L \otimes \text{U}(1)_R$							
FA ₄₂₁	$(\mathbf{1}, \mathbf{2}, 1/2)_D$	$(\mathbf{15}, \mathbf{1}, 0)_W$	$(\mathbf{15}, \mathbf{1}, 0)_R$ $(\mathbf{15}, \mathbf{2}, 1/2)_C$	3.48	17.54	0.0320	40.9 ± 1.2
$G_{\text{int}} = \text{SU}(4)_C \otimes \text{SU}(2)_L \otimes \text{SU}(2)_R$							
FA ₄₂₂	$(\mathbf{1}, \mathbf{2}, \mathbf{2})_W$	$(\mathbf{1}, \mathbf{3}, \mathbf{1})_W$	$(\mathbf{15}, \mathbf{1}, \mathbf{1})_R$ $(\mathbf{15}, \mathbf{1}, \mathbf{3})_R$	9.00	15.68	0.0258	34.0 ± 1.2
FB ₄₂₂	$(\mathbf{1}, \mathbf{2}, \mathbf{2})_W$	$(\mathbf{1}, \mathbf{3}, \mathbf{1})_W$	$(\mathbf{15}, \mathbf{1}, \mathbf{1})_R$ $(\mathbf{15}, \mathbf{2}, \mathbf{2})_C$ $(\mathbf{15}, \mathbf{1}, \mathbf{3})_R$	5.84	17.01	0.0587	38.0 ± 1.2

Examples:

Non-Singlets: Fermions

Summary Plot



Summary

- ✦ LHC susy and Higgs searches have pushed CMSSM-like models to “corners”
- ✦ Though some phenomenological solutions are still viable typically along “strips” in parameter space
- ✦ NUHM models with “low” μ still promising as are subGUT models; PGM (with wino DM or Higgsino DM)
- ✦ Several possibilities in non-SUSY SO(10) models
- ✦ Challenge lies in detection strategies