

CR



linac

SC

31

1932 e^+

n

d

Cockcroft-Walton

Cyclotron

Cosmic Ray \rightarrow Particle Physics

not C.R.

e^+

$\mu^\pm \rightarrow e^\pm + \nu + \bar{\nu}$

$\pi^+ \rightarrow \mu + \nu$

$\pi^- + AZ \rightarrow$ star

K^\pm K^0

Σ^\pm Λ^0

γ

m

π^0 η^0

ρ w

Σ^0

1948 Part. Phys. Cosmic rays \rightarrow Accelerators
pion physics

1952 \rightarrow BNL Cosmotron 3GeV

Strange Particles

$K, \Sigma \Lambda$

Ξ

* visual (emulsion, Cl. chamber)
decay mode

θ $K_2\pi$

τ $K_3\pi$

κ $K_3\pi$

$K\gamma$ pair creation

Gell-Mann - Nishijima

$\theta - \tau$

Parity violation

Berkeley

Berkeley
6 GeV

$\bar{p} \bar{n}$

Target

(H target)

$C_n H_{2n} - C$

H target \rightarrow 1960's

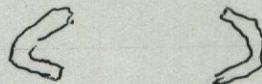
detectors + mag fd
space/time resol.

1932 cyclotron

1945 synchrotron

AG PS

1952 AG (Strong Focusing.)



BM + Focus.

Combined

BM + FM + A

Sep. function

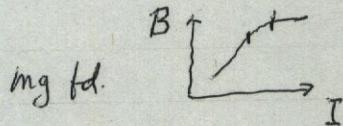
CERN ES 2 GeV

INS ES

CERN PS 25/27 GeV

1959

BNL AGS 30/33 1960
Gev



exp.

USA FNAL

SU₃

ν exp.

B C.

1953

Glaser

propane

aloxretz

H

liq H, liq D

1960 - 70's

HBC

DBC

{ BEBC 30 m^3
BNL, ANL

(1970's pol H target
→ エコナウス

Xe BC

Fren BC.

Lagrange

HBC dominate age

Heavy liquid B.C.

(Gargamelle)

sensitive
γ, e

France (Orsay) + CERN (^{Exptl}~~Apparatus~~ Dept)

1972 ν-e scattering

N.C.

bkgd → N.C.

1972 ν-e, NC

1973 Renom. gauge theory)

WE complete,

USA

1953 → CERN

1968 \sum_{WE} (H.E. phys Budget) > (USA)

1953 → CERN | dom. | cosmic ray HEP WE
| Accel. collid. HEP USA

1980's Z, W, gluon

3:0

USA

SSC

collider

colliding beam machine

MURA

Taj et al

internal report

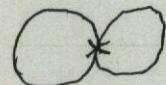
early

1950's

Princeton

O'Neil Richter

e^-e^- collider



QED

Frascati



e^+e^- in the same ring

QED + hadron physics

$e^+e^- \rightarrow$ all charged pair, etc.

Frascati- Orsay

$e^+e^- \not\rightarrow l^+l^-$

e^- machine

1960's SLAC 2 mile e^- linac

1960's

CEA

G GENES

\rightarrow

DESY

6 GeV
ES

Jentzke

QED B,D.

$\gamma \rightarrow e^+e^-$

S. Ting

QED OK

1970's

SLAC

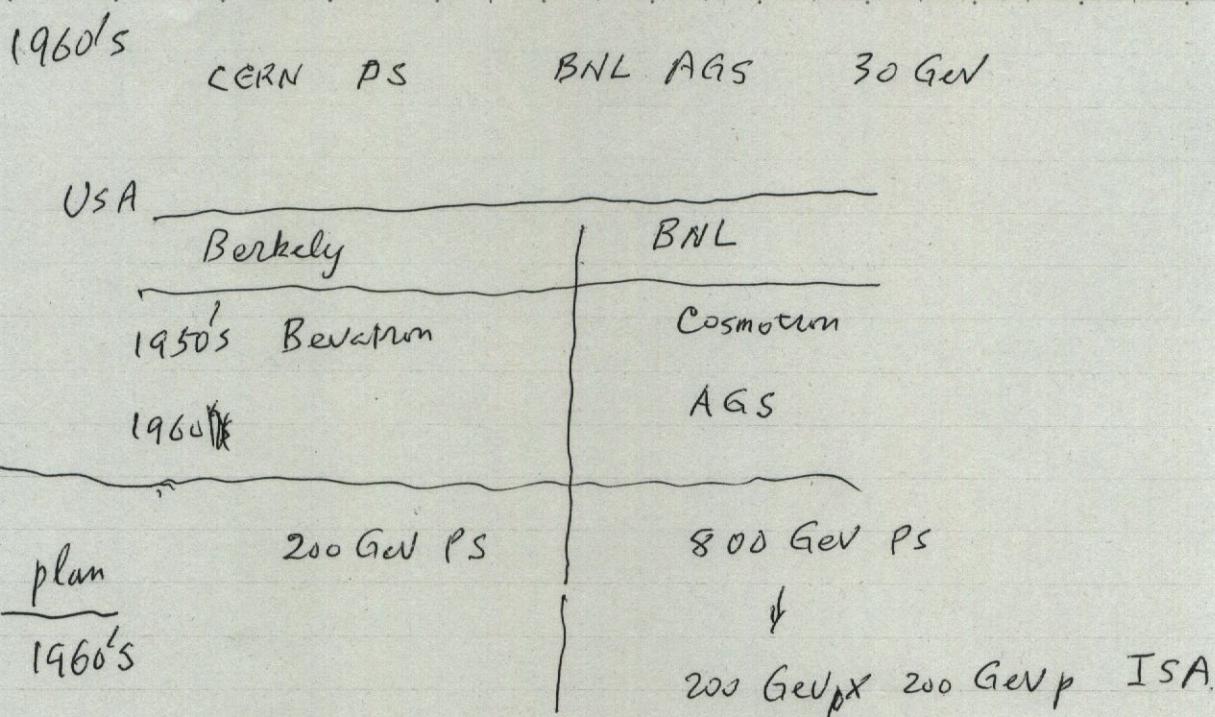
SPEAR

J/ψ

e^+e^- physics

USA (SLAC)

dominance



cost up.

Sup C. Magnet

$\left\{ \begin{array}{l} \text{mess mod } X \\ \text{management Non Exist} \end{array} \right.$

Budget 1/4 Tunnel Refrig.

$\underbrace{\text{RHIC}}$

B Wilson 200 GeV (Berkeley) \rightarrow 500 GeV PS

X Berkeley, BNL Acc. Expert

Fermilab 500 GeV PS

1970's

CERN ISR $30 \text{ GeV } p \times 30 \text{ GeV } p$

Fermilab 500 GeV PS $\cancel{\times}$ 1st gen. exp || γ

CERN 450 GeV SPS OK "

Next Projects in

Fermilab

\star $500 \text{ GeV} \rightarrow 1000 \text{ GeV } p$ S.C. Mag.
 $\left. \begin{array}{l} e - p \\ p - \bar{p} \end{array} \right\} \rightarrow \text{DESY}$
 $\rightarrow \text{CERN}$
 (Rubbia)

C. Rubbia late 1970's

SPS $\rightarrow S p \bar{p} S$ $350 \text{ GeV } p \times \bar{p}$

early 1980's

$\rightarrow Z^0, W^+$

early
1980's

Z, W CERN

gluon DESY (gluon jet)

WE : USA = 3 : 0

USA

SSC

Higgs?

LEP $Z^0 \rightarrow \text{SM cut}$

LHC

dusty like Nach.Target

20th century Energy $\times 10$ per 7 yrs

21st century
Accel. $\frac{\text{GeV}}{m} \rightarrow \frac{\text{TeV}}{m}$

Laser?

detectors space/time resolution

↓

like emulsion 10^{-15} sec

small but efficient accelerators / colliders

↓

detectors

(ILC) → ?

ILC etc

study of dark matter particles
and their exc. states

New Sym.

SUSY, ...

~~Ex 10 / 7 yrs~~

Galaxy

rotational velocity

Dark Matter

$$1/3 \text{ GeV/cm}^3$$

Gal.

500 km/sec

escape velocity

Star

50 km/sec

"

Dark Matter

Ideal gas

$$\sqrt{\langle v^2 \rangle} \sim 400 \text{ km/sec}$$

ILC,

SUSY particles

Dark matter — No container

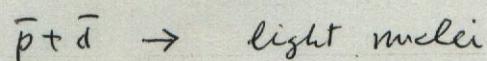
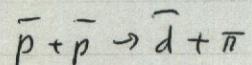
laser accel.

Space resolution — detector

(emulsion)

Anti-matter Factory

\bar{p} \bar{d}



New heavy nuclei $Z > 100$

2nd World War

govt → Funds to (Nuel.) Physicists

冷戰；原爆之後，¹⁹⁴⁵ ¹⁹⁴⁶ 航空

20^c

政府，軍 → Scientists
(AEC)

科系

H.E. Phys.

20^c ISLA, SSC

20^c 末 科學的技術

(HEP) → space science
Inform. "

Life "

21^c

Space research

聯合 mission

	WF	SF (AG)
	JINR	CERN
	Synchrophasotron	CPS
const. started		1955
1st beam	1957	24 Nov. 1959
Magnet units	W.F. 48	S.F. 100
Straight Sections	4	20 (3m)
Field { at injection	150 G	147 G
{ at max	13 kG	14 kG
Rise Time	3.3 sec	1.2 sec
Magnet Weight	(Fe 36 000 tons Cu 2 700 tons)	(Fe 3000 tons Al 130 tons)
Power Input { Av	4 MW	1.6 MW
peak	140 MW	32 MW
Mag. Sec. gap	5.3 x 7.5 m 40 x 200 cm	1.16 x 1.94 m 10.0 x 15.0 cm
Ring Diameter	72 m	200 m
Energy	10 GeV	28 / 24 GeV
Repetition Rate	0.08 sec ⁻¹	0.2 / 0.33 sec ⁻¹
Internal Beam	{ 1967 (1980) 10 ¹⁰ (10 ¹²) p/pulse 10 ⁹ (10 ¹¹) p/s	{ 1967 (1980) 10 ¹² (1.8 x 10 ¹³) (1 ~ $\frac{1}{4}$) x 10 ¹² (0.5 x 10 ¹³)
cost accel. only		120 MSF (1954 - 59)

Longitudinal polarization can be obtained by additional magnetic devices making spin rotate or by injecting longitudinally polarized beams, P_L being not affected during motion along the orbit.

e^\pm Storage Rings in the World

e^-e^- : Vepp -1 (Novosibirsk, USSR)	2 x 0.17 GeV	1963-1966
Princeton-Stanford (USA)	2 x 0.55 GeV	1961-1966
e^+e^- : Ada (Frascati, Italy; Orsay, France)	2 x 0.2 GeV	1961-1965
Vepp -2 (Novosibirsk, USSR)	2 x 0.7 GeV	1967-1974
Vepp -2 M (Novosibirsk, USSR)	2 x 0.7 GeV	1974-
ACO (Orsay, France)	2 x 0.55 GeV	1967-1976
Adone (Frascati, Italy)	2 x 1.5 GeV	1970-
By-Pass (Cambridge, USA)	2 x 2.5 GeV	1973-1974
Spear (Stanford, USA)	2 x 4.2 GeV	1973-
Doris (Hamburg, Germany)	2 x 5 GeV	1974-
Petra (Hamburg, Germany)	2 x 19 GeV	1978-
CESR (Cornell, USA)	2 x 8 GeV	1979-
PEP (Stanford, USA)	2 x 15 GeV	1980-
Vepp -4 (Novosibirsk, USSR)	2 x 7 GeV	1980(?)
$e^\pm e^\pm$: DCI (Orsay, France)	2 x 1.9 GeV	1976-

1.3 Detectors and Measurements

During an experiment one has to identify the nature of the particle produced, their kinematical characteristics and the rate of the reactions.

The intersection regions of the storage rings are surrounded by detectors whose accuracy is defined by:

- their solid angle Ω as close as possible to 4π in order that no particle escapes detection,
- their ability to identify particles, charged ones, neutral ones and unstable ones. Charged particles are easily detected by various kinds of chambers using ionisation effects; their identification is done using secondary reactions in various materials (showers for electrons, nuclear interactions for hadrons) and dE/dx informations (energy loss by ionisation). Neutral particles (photons or