

Emergence of an expanding (3+1)-dimensional spacetime in the type IIB matrix model

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1. Introduction

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Type IIB matrix model (a.k.a. IKKT model)

⇒ Promising candidate for nonperturbative string theory

[N. Ishibashi, H. Kawai, Y. Kitazawa and A. Tsuchiya, hep-th/9612115]

$$S = \underbrace{\frac{-N}{4} \text{tr}[A_\mu, A_\nu]^2}_{=S_b} + \underbrace{\frac{-N}{2} \text{tr}\bar{\psi}_\alpha(\Gamma^\mu)_{\alpha\beta}[A_\mu, \psi_\beta]}_{=S_f}$$

- Dimensional reduction of the D=10 super-Yang-Mills theory to 0 dimension
- A_μ ($\mu=0, 1, \dots, 9$), Ψ_α ($\alpha=1, 2, \dots, 16$ after Weyl projection)
⇒ $N \times N$ Hermitian traceless matrices.
- $N=2$ supersymmetry ⇒ eigenvalues of A_μ are interpreted as the spacetime coordinates.

How does our (3+1)-dim spacetime emerge dynamically?

2. Lorentzian type IIB matrix model

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Sign problem of the type IIB matrix model

Euclidean version : Wick rotation $A_{10} = -iA_0$, $\Gamma_{10} = i\Gamma^0$
 contracted by the Euclidean metric $\delta_{\mu\nu} = \text{diag}(1, 1, 1, \dots, 1)$

SSB $\text{SO}(10) \rightarrow \text{SO}(3)$
 (3dim space time)

$$Z = \int dA \left(e^{-S_b} \right) \underbrace{\int d\psi e^{-S_f}}_{\text{complex}}$$

[K.N. Anagnostopoulos, T. Azuma, Y. Ito J. Nishimura, T. Okubo and S.K. Papadoudis arXiv:2002.07410]

\mathcal{M} is a $16(N^2-1) \times 16(N^2-1)$ sparse matrix
 $= \text{Pf } \mathcal{M}$

Lorentzian version : no Wick rotation

contracted by the **Lorentzian metric** $\eta_{\mu\nu} = \text{diag}(-1, 1, 1, \dots, 1)$

$$Z = \int dA \left(e^{iS_b} \right) \underbrace{\int d\psi e^{iS_f}}_{\text{real}}$$

$= \text{Pf } \mathcal{M}$

[J. Nishimura and A. Tsuchiya,
 arXiv:1904.05919]

We employ the **Complex Langevin Method (CLM)**

2. Lorentzian type IIB matrix model

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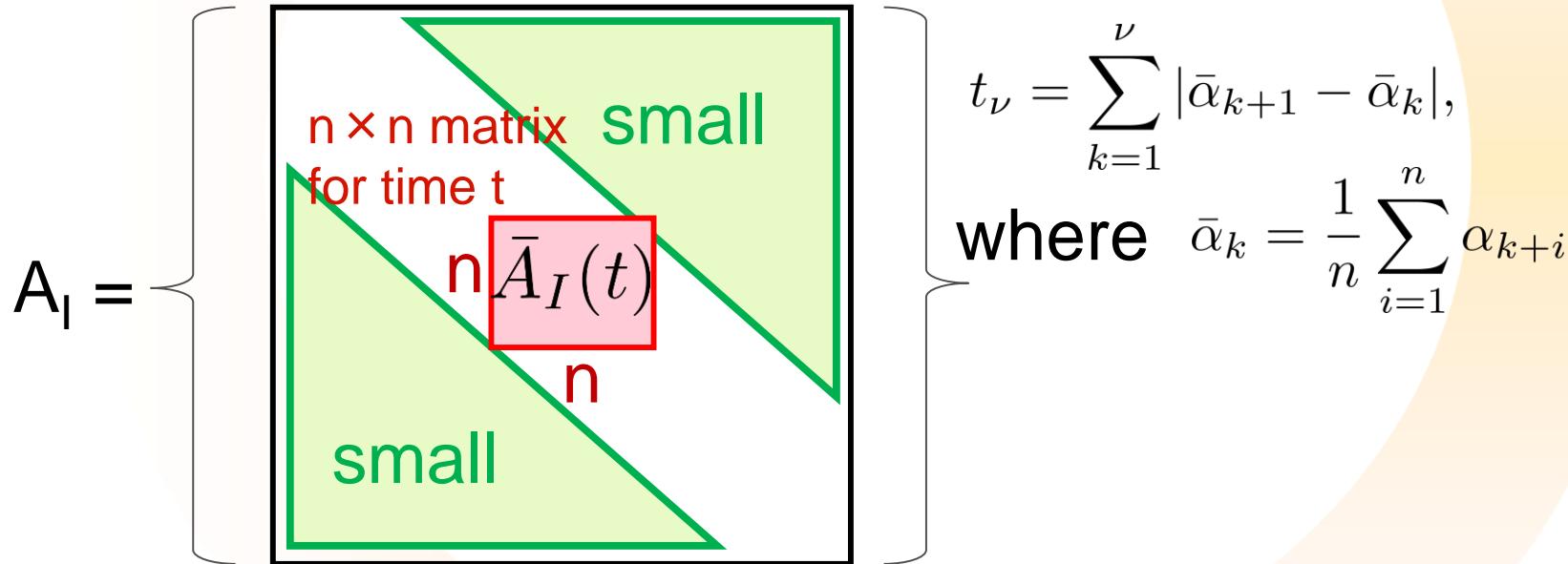


Time development of the Lorentzian version:
⇒ gauge fixing to diagonalize A_0

[S.W. Kim, J. Nishimura and A. Tsuchiya, arXiv:1108.1540]

$$A_0 = \text{diag}(\alpha_1, \alpha_2, \dots, \alpha_N), \text{ where } \alpha_1 < \alpha_2 < \dots < \alpha_N.$$

Band-diagonal structure of A_I



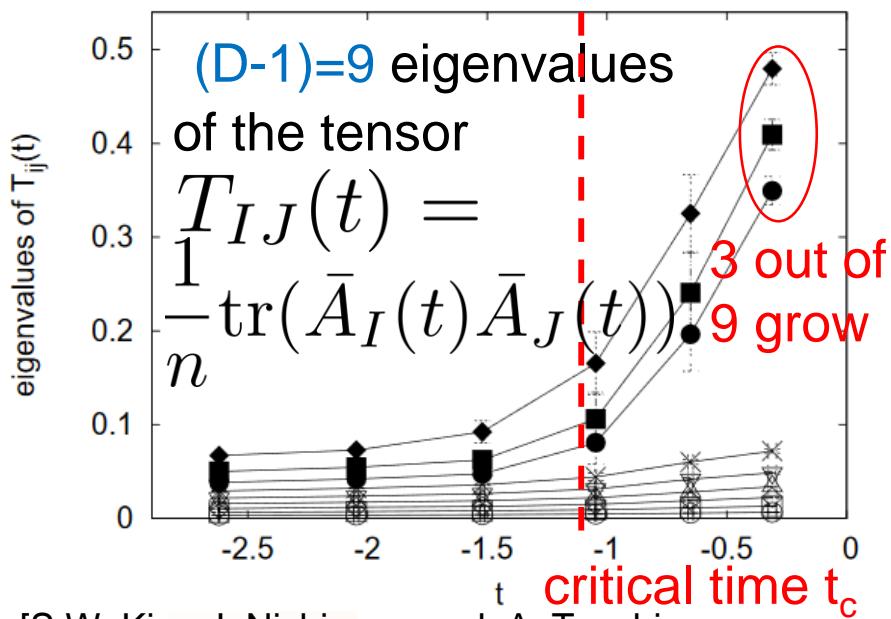
2. Lorentzian type IIB matrix model

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Results with the approximation to avoid the sign problem

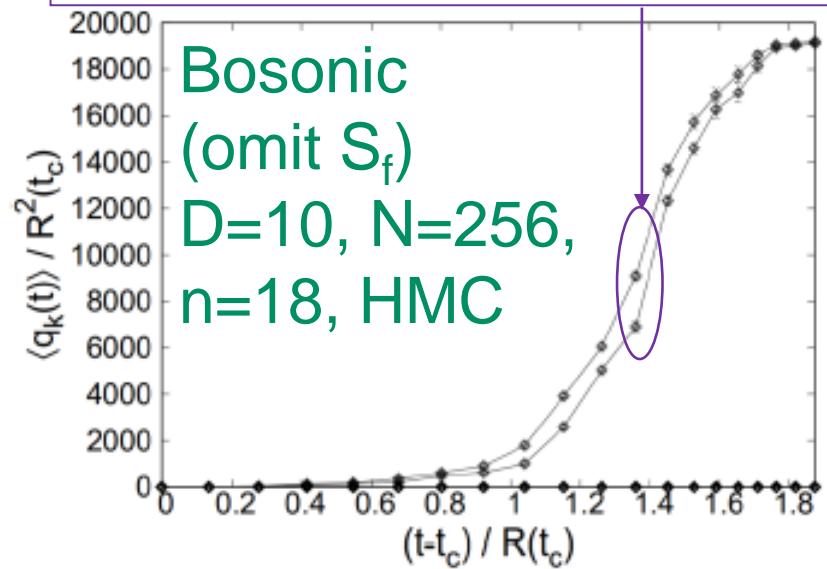
SUSY (include S_f) D=10,
N=16, n=4, HMC (not CLM)



[S.W. Kim, J. Nishimura and A. Tsuchiya, arXiv:1108.1540]

Dynamical emergence
of (3+1)-dim spacetime.

2 of the n eigenvalues of $Q(t) = \sum_{I=1}^9 (\bar{A}_I(t))^2$ grow.



[T. Aoki, M. Hirasawa, Y. Ito, J. Nishimura and A. Tsuchiya, arXiv:1904.05914]

$\bar{A}_I(t) \propto \sigma_I \oplus 0_{n-2}$ ($I = 1, 2, 3$)
Pauli-matrix space structure

2. Lorentzian type IIB matrix model

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Equivalence between the Lorentzian and Euclidean version

[Y. Asano, private communication]

Contour deformation $A_0 = e^{-3i\pi u/8} \tilde{A}_0, A_I = e^{i\pi u/8} \tilde{A}_I$

$$S_b \rightarrow \tilde{S}_b = e^{i\pi u/2} \left\{ \frac{-N}{4} \text{tr}[\tilde{A}_I, \tilde{A}_J]^2 + \frac{N}{2} e^{-i\pi u} \text{tr}[\tilde{A}_0, \tilde{A}_I]^2 \right\}$$

w.r.t. worldsheet

w.r.t. target space

u=0: Lorentzian, **u=1:** Euclidean

$$e^{iS_b(A)} = e^{-S(\tilde{A})}, \quad S(\tilde{A}) = -\frac{N}{4} e^{-i\pi(1-u)/2} \text{tr}[\tilde{A}_I, \tilde{A}_J]^2 - \frac{N}{2} e^{i\pi(1-u)/2} \text{tr}[\tilde{A}_0, \tilde{A}_I]^2$$

real part is positive ($0 < u \leq 1$).

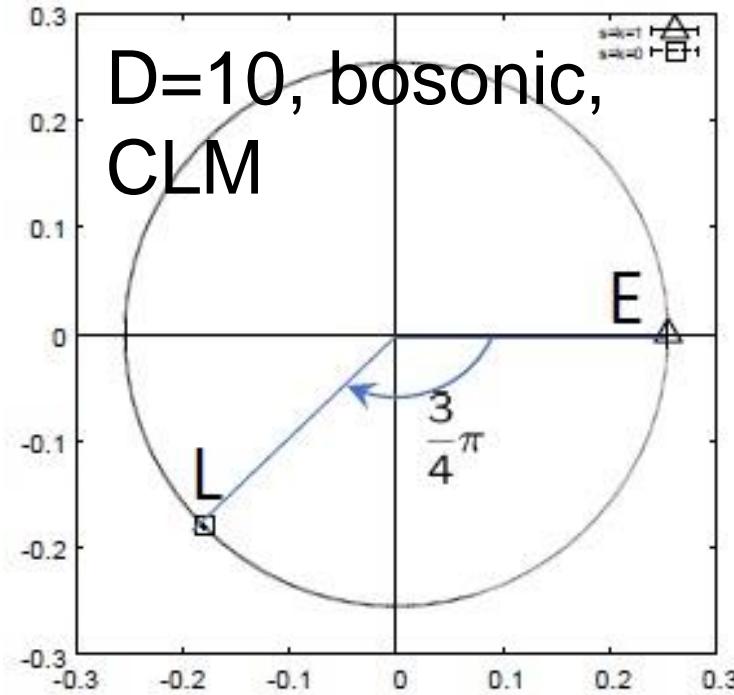
Cauchy's theorem:

$\langle \mathcal{O}(e^{-3i\pi u/8} \tilde{A}_0, e^{i\pi u/8} \tilde{A}_I) \rangle_u$ is independent of u .

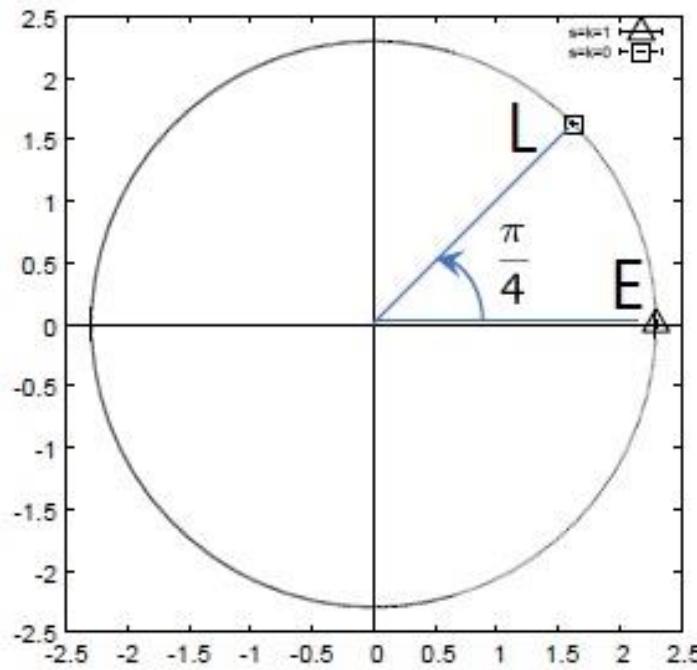
2. Lorentzian type IIB matrix model

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$$\left\langle \frac{1}{N} \text{tr}(A_0^{(L)})^2 \right\rangle = e^{-3i\pi/4} \left\langle \frac{1}{N} \text{tr}(A_0^{(E)})^2 \right\rangle$$



$$\left\langle \frac{1}{N} \text{tr}(A_I^{(L)})^2 \right\rangle = e^{i\pi/4} \left\langle \frac{1}{N} \text{tr}(A_I^{(E)})^2 \right\rangle$$



Equivalence of the Euclidean ($u=1$) and
Lorentzian ($u \rightarrow +0$) model.

→ The spacetime is Euclidean.

[K. Hatakeyama, et. al. , arXiv:2112.15368]

2. Lorentzian type IIB matrix model

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Adding the Lorentzian mass term

$$Z = \int dA d\psi e^{i(S+S_\gamma)}, S_\gamma = \frac{-N\gamma}{2} \text{tr}(A^\mu A_\mu) = \frac{N\gamma}{2} \{\text{tr}(A_0)^2 - \text{tr}(A_I)^2\}$$

$$e^{iS_\gamma(A)} = e^{-S_\gamma(\tilde{A})}, \quad S_\gamma(\tilde{A}) = \frac{N\gamma}{2} \{ e^{-i\pi(2+3u)/4} \text{tr}(\tilde{A}_0)^2 + e^{i\pi(2+u)/4} \text{tr}(\tilde{A}_I)^2 \}$$

real part is negative ($0 < u \leq 1$).

At $\gamma > 0$, we cannot define the model by contour deformation

⇒ Equivalence to the Euclidean model ($u=1$) is violated.

We consider the limit $N \rightarrow \infty \Rightarrow \gamma \rightarrow 0$.

3. Complex Langevin Method

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Complex Langevin Method (CLM)

⇒ Promising method to solve complex-action systems.

[Parisi, Phys.Lett. 131B (1983) 393, Klauder, Phys.Rev. A29 (1984) 2036]

- Introducing τ_a

⇒ time order $\alpha_1 < \alpha_2 < \dots < \alpha_N$ for complexified α_i ($k=2, \dots, N$)

$$\alpha_1 = 0, \quad \alpha_k = \sum_{i=1}^{k-1} e^{\tau_i}$$

$$A_0 = \text{diag}(\alpha_1, \alpha_2, \dots, \alpha_N)$$

- Mass term to avoid the near-zero modes of the

Dirac operator: $\text{Pf } \mathcal{M}_{m_f} = \int d\psi e^{iS_{m_f}}$,

$$S_{m_f} = \frac{-N}{2} \text{tr} \{ \bar{\psi}_\alpha (\Gamma^\mu)_{\alpha\beta} [A_\mu, \psi_\beta] + m_f \bar{\psi}_\alpha (\Gamma^7 \Gamma^{8\dagger} \Gamma^9)_{\alpha\beta} \psi_\beta \}$$

Effective action $Z = \int dA e^{-S_{\text{eff}}}$,

$$S_{\text{eff}} = -i(S_b + S_\gamma) - \log \text{Pf } \mathcal{M}_{m_f}$$

$$- \log \prod_{1 \leq a < b \leq N} (\alpha_a - \alpha_b)^2$$

$$- \sum_{a=1}^{N-1} \tau_a$$

3. Complex Langevin Method

$$\frac{d(A_I)_{ab}(\sigma)}{d\sigma} = - \frac{dS_{\text{eff}}}{d(A_I)_{ba}} + \eta_{I,ab}(\sigma), \quad \frac{d\tau_a(\sigma)}{d\sigma} = - \frac{dS_{\text{eff}}}{d\tau_a} + \eta_a(\sigma)$$

drift term Hermitian-matrix white noise drift term real-number white noise

fictitious Langevin time fictitious Langevin time

- A_I : Hermitian \rightarrow general complex traceless matrices.
- τ_a : Real number \rightarrow complex number.

The drift term involves $\frac{d}{d(A_I)_{ba}} \{-\log \text{Pf } \mathcal{M}_{m_f}\} = -\frac{1}{2} \text{Tr} \left(\frac{d\mathcal{M}_{m_f}}{d(A_I)_{ba}} \mathcal{M}_{m_f}^{-1} \right)$

- \mathcal{M}_{m_f} 's near-zero modes \Rightarrow singular drift problem.
- We use conjugate gradient (CG) method and noisy estimator.

Large-scale numerical simulation using supercomputers.

3. Complex Langevin Method

The condition to justify the CLM: [K. Nagata, J. Nishimura and S. Shimasaki, arXiv:1606.07627]

The probability distribution of the drift norms falls exponentially or faster.

$$u_A = \sqrt{\frac{1}{9N^3} \sum_{I=1}^9 \sum_{a,b=1}^N \left| \frac{dS_{\text{eff}}}{d(A_I)_{ba}} \right|^2}, \quad u_\alpha = \sqrt{\frac{1}{N} \sum_{a=1}^{N-1} \left| \frac{dS_{\text{eff}}}{d\tau_a} \right|^2}$$

Look at the drift terms \Rightarrow Get the drift of the CLM.

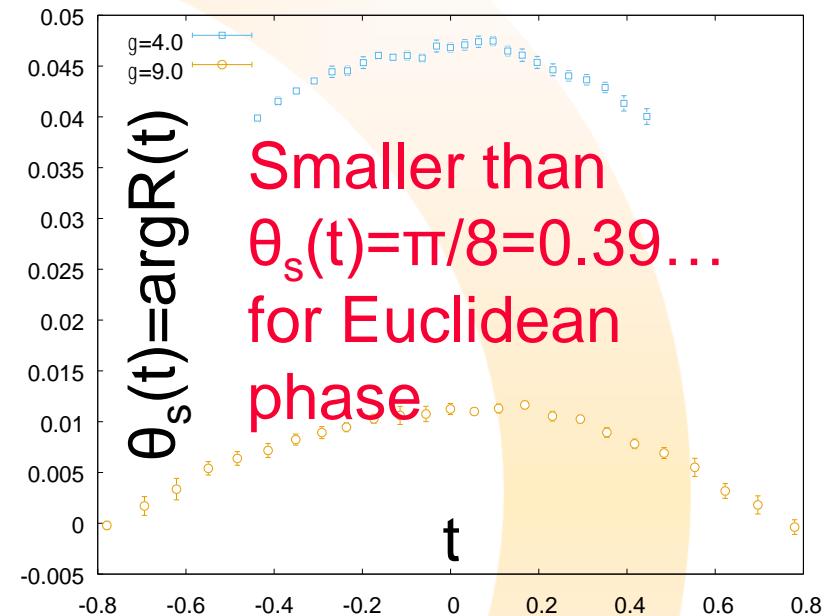
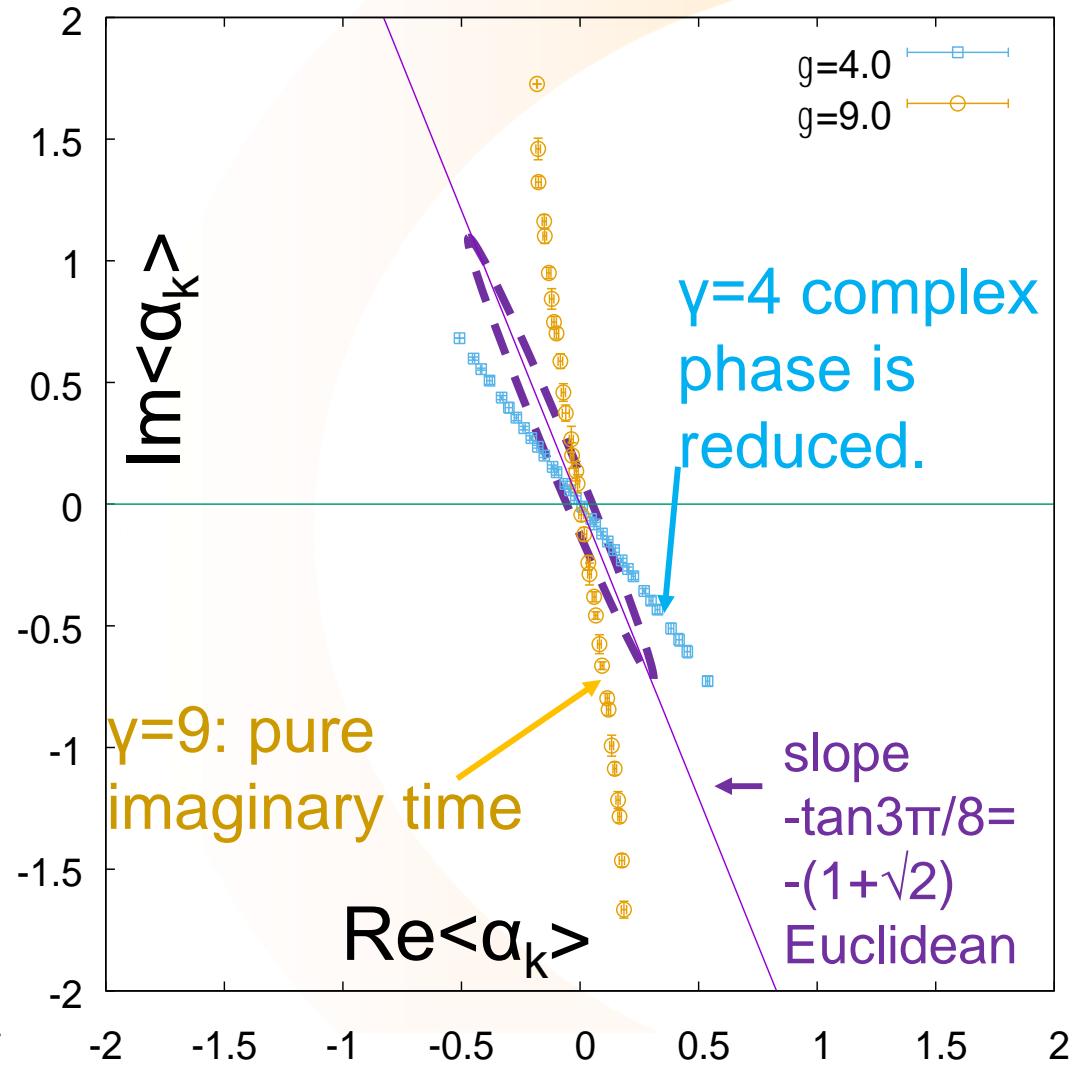
Dynamical stabilization: [F. Attanasio and B. Jäger arXiv:1808.04400]

After each Langevin step, $A_I \rightarrow \frac{A_I + \eta A_I^\dagger}{1 + \eta}$ Here, $\eta = 0.01$.

($\eta = 0$: do nothing, $\eta = 1$: Hermitize completely)

4. Result

SUSY (include S_f), $N=32$, $n=8$, $u=0$, $m_f=3.5$ (preliminary)

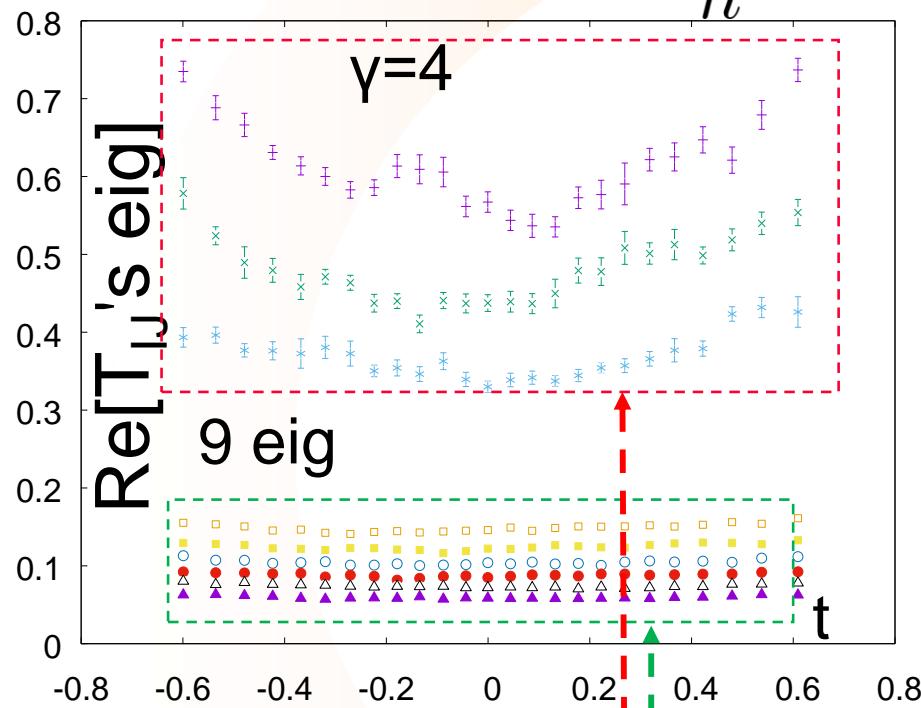


$$R^2(t) = \left\langle \frac{1}{n} \sum_{I=1}^9 \text{tr}(\bar{A}_I(t))^2 \right\rangle$$

4. Result

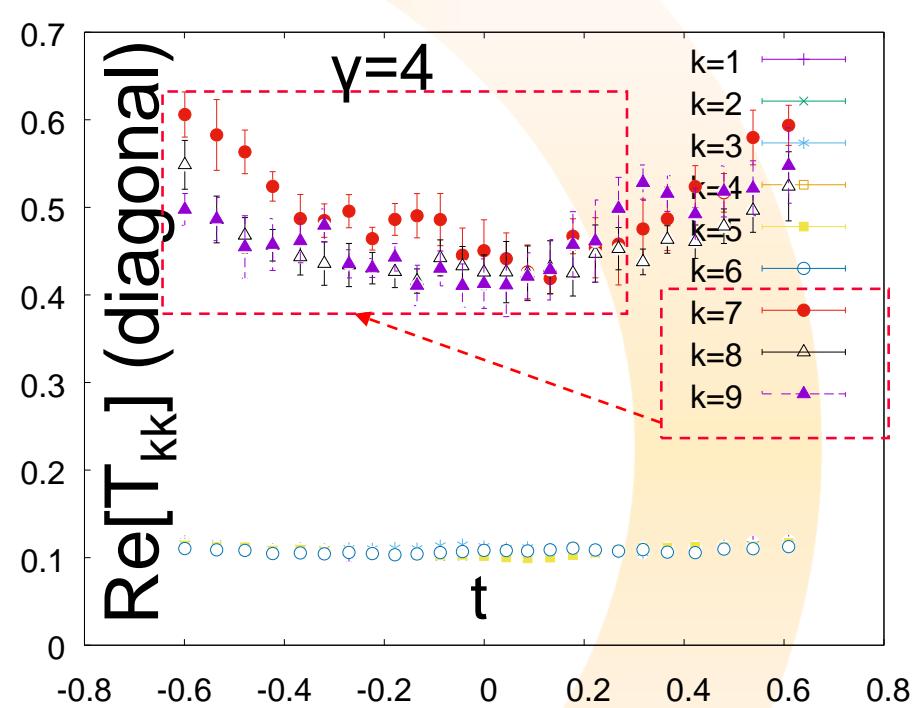
SUSY (include S_f), $N=32$, $n=8$, $u=0$, $m_f=3.5$ (preliminary)

$$9 \times 9 \text{ tensor } T_{IJ} = \frac{1}{n} \text{tr}(\bar{A}_I(t)\bar{A}_J(t))$$



expanding 3dim
space(similarly for $\gamma=9$)

shrunken 6dim space



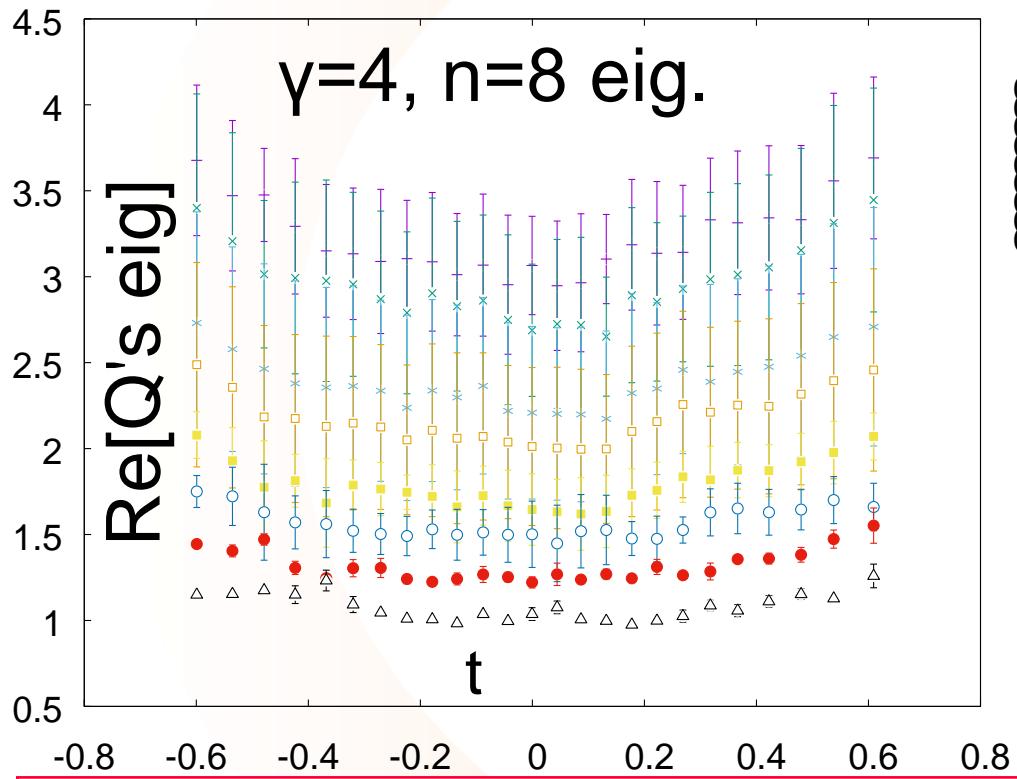
effect of $m_f \bar{\psi}_\alpha (\Gamma^7 \Gamma^{8\dagger} \Gamma^9)_{\alpha\beta} \psi_\beta$

4. Result

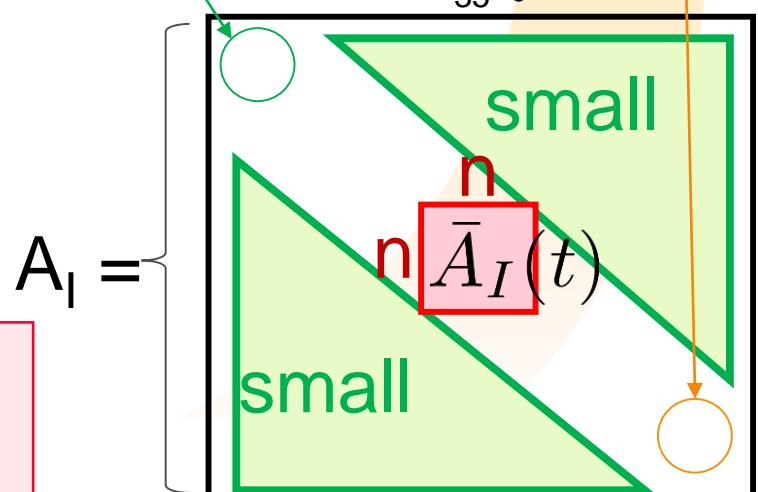
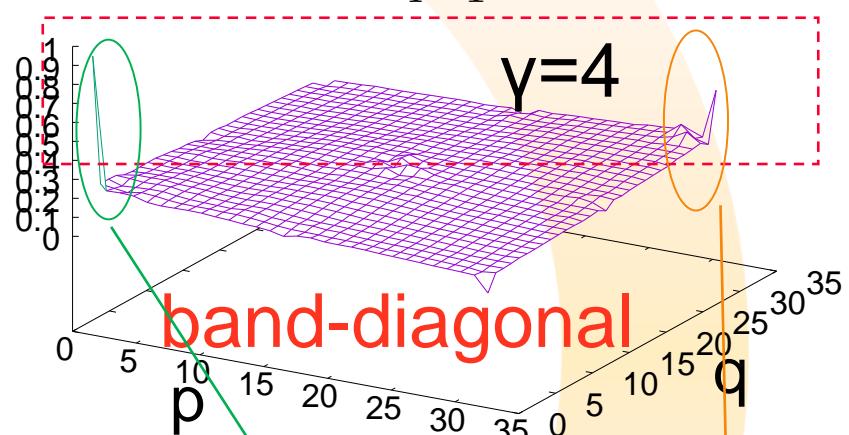
SUSY (include S_f), $N=32$, $n=8$, $u=0$, $m_f=3.5$ (preliminary)

$$n \times n \text{ matrix } Q(t) = \sum_{I=1}^9 (\bar{A}_I(t))^2$$

$$A_{pq} = \sum_{I=1}^9 |(A_I)_{pq}|^2$$



departure from Paulian
structure (similarly for $\gamma=9$)



5. Conclusion

Complex Langevin Method (CLM) for the type IIB matrix model.

Equivalence of the Euclidean and Lorentzian model without the Lorentzian mass term.

Introduce a Lorentzian mass term
⇒ emergence of **3dim space**.

At larger N , and smaller γ and m_f :

Can we observe the transition from
Euclidean to Lorentzian geometry?

Large-scale numerical simulation
at larger N is important.

