

# Post-Newtonian analysis of bobbing effects in spinning binary systems

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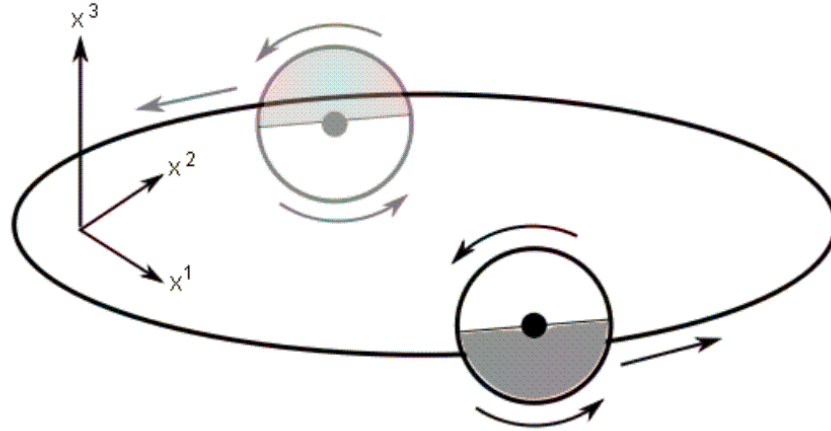
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# Outline

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(Newtonian; circular orbits)
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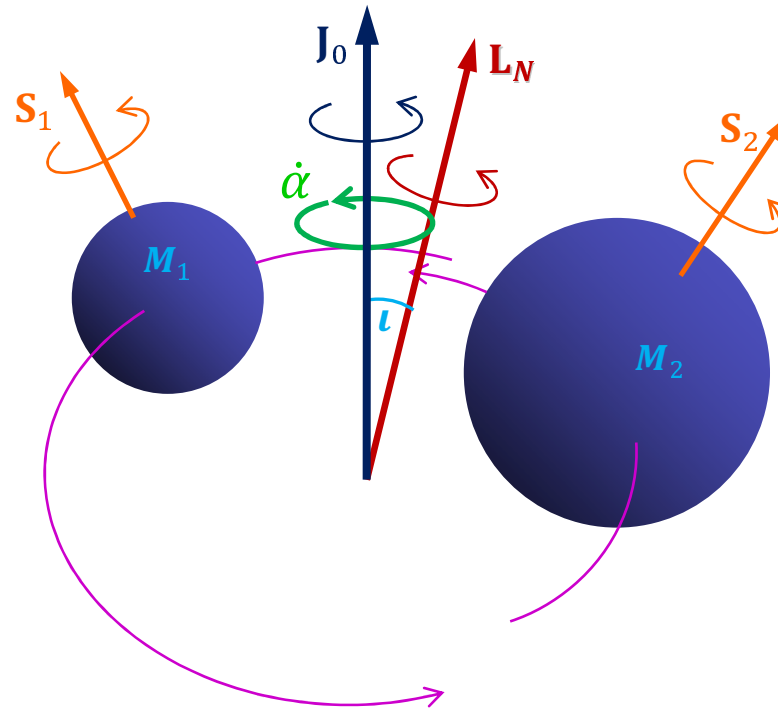
# 1. What is bobbing? How is it observed in spinning binary systems?



<Excerpted from *Gralla et. al* [Phys. Rev. D 81, 104012 (2010)]>

The lab frame centroid of a body (solid line) is confined to the  $x^1$ - $x^2$  plane. However, since the body has a (constant) spin vector lying in the plane, its center of mass (solid dot) is displaced in the  $x^3$  direction by a velocity dependent factor, giving rise to bobbing:

$$m\vec{z}_{\text{CM}} = m\vec{z}_{\text{Lab}} + \vec{S} \times \frac{d\vec{z}_{\text{L}}}{dt} + \mathcal{O}\left(\left|\frac{d\vec{z}_{\text{L}}}{dt}\right|^2\right)$$

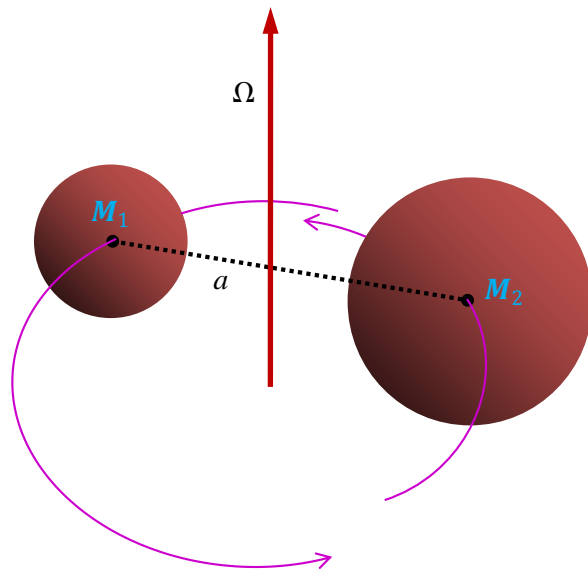


$$\mathbf{J}_0 = \mathbf{L}_N + \mathbf{S}_1 + \mathbf{S}_2$$

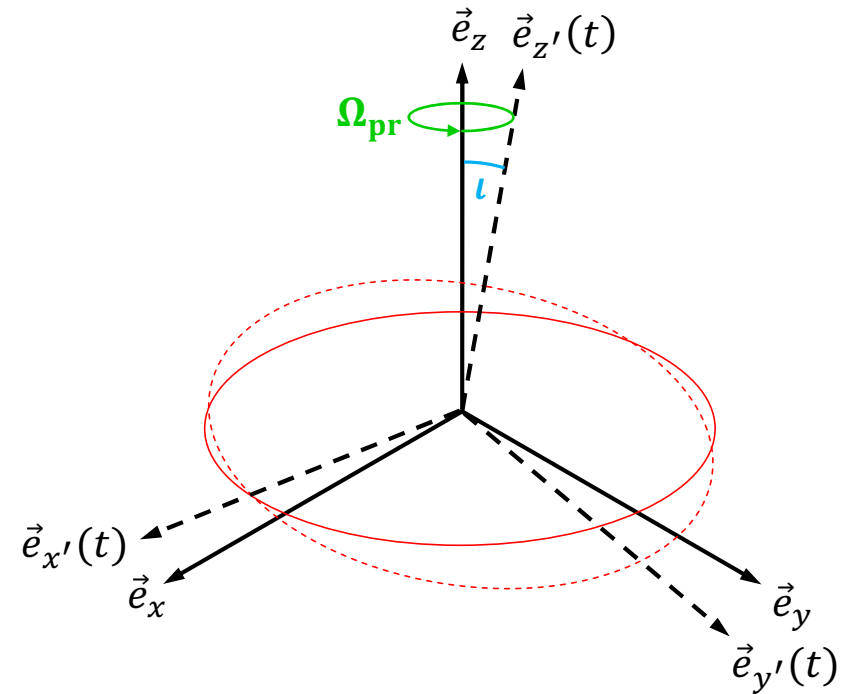
Spinning binary inspiral system

In spinning binary systems, this kinematics can be represented by the interactions between the orbital angular momentum ( $\mathbf{L}_N$ ) and spins of two bodies ( $\mathbf{S}_1, \mathbf{S}_2$ ), i.e. **spin-orbit** interactions. The net bobbing can be observed from the inclination angle ( $\iota$ ).

## 2. Precessing binary inspirals and GWs (Newtonian; circular orbits)



Binary inspiral system



$$\vec{e}_{x'}(t) = -\sin(\Omega_{\text{pr}}t) \vec{e}_x + \cos(\Omega_{\text{pr}}t) \vec{e}_y$$

$$\vec{e}_{y'}(t) = -\cos l \cos(\Omega_{\text{pr}}t) \vec{e}_x - \cos l \sin(\Omega_{\text{pr}}t) \vec{e}_y + \sin l \vec{e}_z$$

$$\vec{e}_{z'}(t) = \sin l \cos(\Omega_{\text{pr}}t) \vec{e}_x + \sin l \sin(\Omega_{\text{pr}}t) \vec{e}_y + \cos l \vec{e}_z$$

Time-varying frame for precessing binary:

$$\begin{bmatrix} \vec{e}_{x'}(t) \\ \vec{e}_{y'}(t) \\ \vec{e}_{z'}(t) \end{bmatrix} = \begin{bmatrix} -\sin(\Omega_{\text{pr}}t) & \cos(\Omega_{\text{pr}}t) & 0 \\ -\cos\iota \cos(\Omega_{\text{pr}}t) & -\cos\iota \sin(\Omega_{\text{pr}}t) & \sin\iota \\ \sin\iota \cos(\Omega_{\text{pr}}t) & \sin\iota \sin(\Omega_{\text{pr}}t) & \cos\iota \end{bmatrix} \begin{bmatrix} \vec{e}_x \\ \vec{e}_y \\ \vec{e}_z \end{bmatrix}$$

Thus,

$$\begin{bmatrix} \vec{e}_{\hat{r}}(t) \\ \vec{e}_{\hat{\theta}}(t) \\ \vec{e}_{\hat{\phi}}(t) \end{bmatrix} = \begin{bmatrix} \cos\psi & 0 & -\sin\psi \\ 0 & 1 & 0 \\ \sin\psi & 0 & \cos\psi \end{bmatrix} \begin{bmatrix} \sin\theta & 0 & \cos\theta \\ \cos\theta & 0 & -\sin\theta \\ 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} \cos\phi & \sin\phi & 0 \\ -\sin\phi & \cos\phi & 0 \\ 0 & 0 & 1 \end{bmatrix} \\ \times \begin{bmatrix} -\sin(\Omega_{\text{pr}}t) & \cos(\Omega_{\text{pr}}t) & 0 \\ -\cos\iota \cos(\Omega_{\text{pr}}t) & -\cos\iota \sin(\Omega_{\text{pr}}t) & \sin\iota \\ \sin\iota \cos(\Omega_{\text{pr}}t) & \sin\iota \sin(\Omega_{\text{pr}}t) & \cos\iota \end{bmatrix}^{-1} \begin{bmatrix} \vec{e}_{x'}(t) \\ \vec{e}_{y'}(t) \\ \vec{e}_{z'}(t) \end{bmatrix}$$

Now, from

$$I_{x'(t)x'(t)} = \mu a^2 \cos^2(\Omega t), \quad I_{y'(t)y'(t)} = \mu a^2 \sin^2(\Omega t), \quad I_{x'(t)y'(t)} = I_{y'(t)x'(t)} = \mu a^2 \cos(\Omega t) \sin(\Omega t)$$

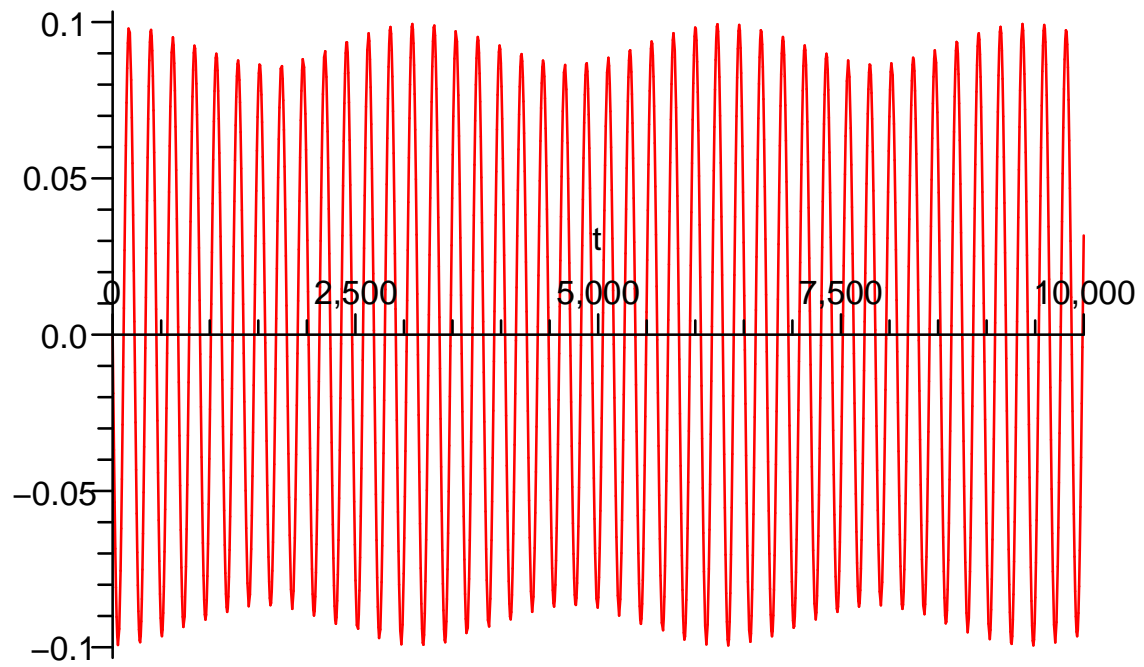
gravitational waves are computed (with polarization angle  $\psi = 0$ ):

$$\mathbf{h}^{\text{TT}} = h_+ \mathbf{e}^+(t) + h_\times \mathbf{e}^\times(t); \quad (\text{TT: Transverse and Trace-free})$$

$$\mathbf{e}^+(t) = (\vec{e}_{\hat{\theta}}(t) \otimes \vec{e}_{\hat{\theta}}(t) - \vec{e}_{\hat{\phi}}(t) \otimes \vec{e}_{\hat{\phi}}(t)), \quad \mathbf{e}^\times = (\vec{e}_{\hat{\theta}}(t) \otimes \vec{e}_{\hat{\phi}}(t) + \vec{e}_{\hat{\phi}}(t) \otimes \vec{e}_{\hat{\theta}}(t))$$

$$\begin{aligned}
\leftarrow h_+ &= h_{\hat{\theta}(t)\hat{\theta}(t)}^{\text{TT}} = \frac{2}{r} \left[ \ddot{I}_{\hat{\theta}(t)\hat{\theta}(t)} (t-r) \right]^{\text{TT}} \\
&= -\frac{\mu(M\Omega)^{2/3}}{r} \left[ \frac{(1+\cos\iota)^2(1+\cos^2\theta)}{2} \cos((2\Omega+2\Omega_{\text{pr}})t-2\Omega r-2\phi) \right. \\
&\quad + \frac{(1-\cos\iota)^2(1+\cos^2\theta)}{2} \cos((2\Omega-2\Omega_{\text{pr}})t-2\Omega r+2\phi) \\
&\quad + \sin\iota(1+\cos\iota)\sin 2\theta \cos((2\Omega+\Omega_{\text{pr}})t-2\Omega r-\phi) \\
&\quad - \sin\iota(1-\cos\iota)\sin 2\theta \cos((2\Omega-\Omega_{\text{pr}})t-2\Omega r+\phi) \\
&\quad \left. + 3\sin^2\iota\sin^2\theta \cos(2\Omega t) \right]
\end{aligned}$$

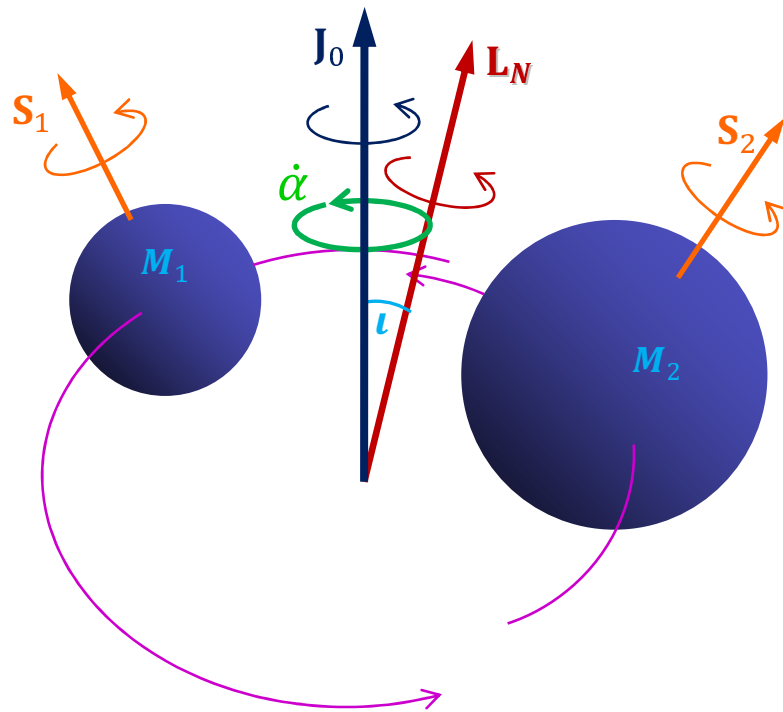
$$\begin{aligned}
h_\times &= h_{\hat{\theta}(t)\hat{\phi}(t)}^{\text{TT}} = \frac{2}{r} \left[ \ddot{I}_{\hat{\theta}(t)\hat{\phi}(t)} (t-r) \right]^{\text{TT}} \\
&= -\frac{\mu(M\Omega)^{2/3}}{r} \left[ (1+\cos\iota)^2 \cos\theta \sin((2\Omega+2\Omega_{\text{pr}})t-2\Omega r-2\phi) \right. \\
&\quad - (1-\cos\iota)^2 \cos\theta \sin((2\Omega-2\Omega_{\text{pr}})t-2\Omega r+2\phi) \\
&\quad + 2\sin\iota(1+\cos\iota)\sin\theta \sin((2\Omega+\Omega_{\text{pr}})t-2\Omega r-\phi) \\
&\quad \left. + 2\sin\iota(1-\cos\iota)\sin\theta \sin((2\Omega-\Omega_{\text{pr}})t-2\Omega r+\phi) \right]
\end{aligned}$$



Plot of  $h_x$  with  $\Omega = 0.012$ ,  $\Omega_{pr} = 0.002$ ,  $\iota = \pi/15$ ,  $\theta = \pi/10$ ,  $\phi = 0$

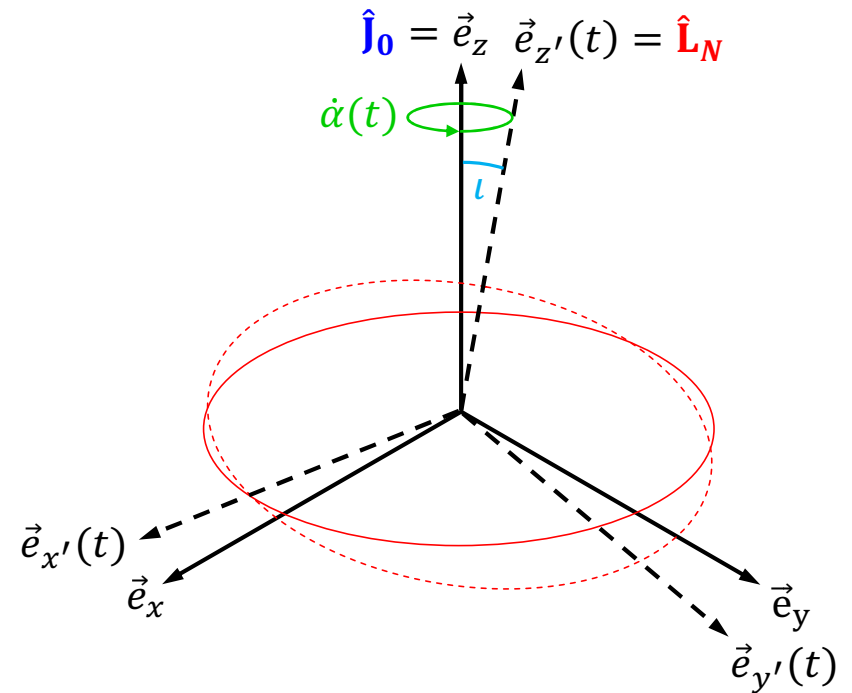


### 3. Spinning binary inspirals and GWs (Post-Newtonian; circular orbits)



$$\mathbf{J}_0 = \mathbf{L}_N + \mathbf{S}_1 + \mathbf{S}_2$$

Spinning binary inspiral system



$$\vec{e}_{x'}(t) = -\sin \alpha \vec{e}_x + \cos \alpha \vec{e}_y$$

$$\vec{e}_{y'}(t) = -\cos \iota \cos \alpha \vec{e}_x - \cos \iota \sin \alpha \vec{e}_y + \sin \iota \vec{e}_z$$

$$\vec{e}_{z'}(t) = \sin \iota \cos \alpha \vec{e}_x + \sin \iota \sin \alpha \vec{e}_y + \cos \iota \vec{e}_z$$

Formally, following *Kidder* [Phys. Rev. D 52, 821 (1995)], *Will et. al* [Phys. Rev. D 54, 4813 (1996)] and *Arun et. al* [Phys. Rev. D 79, 104023 (2009)], one can perform the multipole expansion

$$h^{ij}(t, \mathbf{x}) = \frac{2}{D_L} \frac{d^2}{dt^2} \sum_{m=0}^{\infty} \hat{N}_{k_1} \cdots \hat{N}_{k_m} I_{\text{EW}}^{ijk_1 \cdots k_m}(t - r),$$

where  $\hat{\mathbf{N}} \equiv \mathbf{x}/D_L$ ,  $I_{\text{EW}}^{ijk_1 \cdots k_m}$ : Epstein-Wagoner (EW) moments. Through quite involved analysis, one can determine

$$h_{+,\times} = \frac{2M\nu v^2}{D_L} \left[ H_{+,\times}^{(0)} + H_{+,\times}^{(1/2)} + H_{+,\times}^{(1/2,\text{SO})} + H_{+,\times}^{(1)} + H_{+,\times}^{(1,\text{SO})} + H_{+,\times}^{(3/2)} + H_{+,\times}^{(3/2,\text{SO})} \right],$$

where  $\nu = \frac{M_1 M_2}{M^2}$ , and  $H_{+,\times}^{(0)}$ ,  $H_{+,\times}^{(1/2)}$ ,  $H_{+,\times}^{(1/2,\text{SO})}$ ,  $H_{+,\times}^{(1)}$ ,  $H_{+,\times}^{(1,\text{SO})}$ ,  $H_{+,\times}^{(3/2)}$ ,  $H_{+,\times}^{(3/2,\text{SO})}$  are given as follows ( $\nu$ -expanded polarizations with the assumption  $S \ll L$ ):

<Excerpted from *Arun et. al* [Phys. Rev. D 79, 104023 (2009)]>

$$H_+^{(0)} = -(c_\theta^2 + 1)\cos 2(\alpha + \Psi), \quad (3.16a)$$

$$H_+^{(1/2)} = \nu \delta s_\theta \left[ \left( \frac{c_\theta^2}{8} + \frac{5}{8} \right) \cos(\alpha + \Psi) - \frac{9}{8}(c_\theta^2 + 1)\cos 3(\alpha + \Psi) \right], \quad (3.16b)$$

$$H_+^{(1)} = \nu^2 \left[ \left( -\frac{c_\theta^4}{3} + \frac{3c_\theta^2}{2} + \frac{19}{6} + \left( c_\theta^4 + \frac{11c_\theta^2}{6} - \frac{19}{6} \right) \nu \right) \cos 2(\alpha + \Psi) + \frac{4}{3}(1 - c_\theta^4)(3\nu - 1)\cos 4(\alpha + \Psi) \right], \quad (3.16c)$$

$$\begin{aligned} H_+^{(3/2)} = & \nu^3 \left[ \delta s_\theta \left( \frac{c_\theta^4}{192} - \frac{5c_\theta^2}{16} - \frac{19}{64} + \left( -\frac{c_\theta^4}{96} - \frac{c_\theta^2}{8} + \frac{49}{96} \right) \nu \right) \cos(\alpha + \Psi) - 2\pi(c_\theta^2 + 1)\cos 2(\alpha + \Psi) \right. \\ & + \delta s_\theta \left( \left( -\frac{81c_\theta^4}{128} + \frac{45c_\theta^2}{16} + \frac{657}{128} \right) + \left( \frac{81c_\theta^4}{64} + \frac{9c_\theta^2}{8} - \frac{225}{64} \right) \nu \right) \cos 3(\alpha + \Psi) + \delta s_\theta \frac{625}{384}(1 - c_\theta^4)(2\nu - 1) \\ & \left. \times \cos 5(\alpha + \Psi) \right], \end{aligned} \quad (3.16d)$$

$$H_+^{(1/2,SO)} = -2\iota c_\theta s_\theta \cos(\alpha + 2\Psi), \quad (3.16e)$$

$$\begin{aligned} H_+^{(1,SO)} = & \nu^2 [(c_\theta(\chi_a^x + \delta\chi_s^x) - s_\theta(\chi_a^z + \delta\chi_s^z))\cos(\alpha + \Psi) - c_\theta(\chi_a^y + \delta\chi_s^y)\sin(\alpha + \Psi)] \\ & + \nu \iota \delta c_\theta \left[ \frac{1}{4}s_\theta^2 \cos \Psi - \left( \frac{c_\theta^2}{8} + \frac{5}{8} \right) \cos(2\alpha + \Psi) + \left( -\frac{9}{8} + \frac{27c_\theta^2}{8} \right) \cos(2\alpha + 3\Psi) \right] \\ & + \iota^2 \left[ -\frac{3}{2}s_\theta^2 \cos 2\Psi + \frac{1}{2}(c_\theta^2 + 1)\cos 2(\alpha + \Psi) \right], \end{aligned} \quad (3.16f)$$

$$\begin{aligned} H_+^{(3/2,SO)} = & \nu^3 \left[ s_\theta c_\theta (2\delta\chi_a^x + (2 - \nu)\chi_s^x) + \left( \frac{4}{3}(1 + c_\theta^2)\delta\chi_a^z + \frac{4}{3}((1 + c_\theta^2) + \nu(1 - 5c_\theta^2))\chi_s^z - s_\theta c_\theta (2\delta\chi_a^x + (2 + 7\nu)\chi_s^x) \right) \right. \\ & \left. \times \cos 2(\alpha + \Psi) - s_\theta c_\theta (2\delta\chi_a^y + (2 - \nu)\chi_s^y) \sin 2(\alpha + \Psi) \right] + \nu^2 \iota s_\theta \left[ c_\theta \left( -c_\theta^2 + 4 + \left( 3c_\theta^2 + \frac{2}{3} \right) \nu \right) \cos(\alpha + 2\Psi) \right. \\ & + c_\theta \left( -\frac{c_\theta^2}{3} - 1 + (c_\theta^2 + 3)\nu \right) \cos(3\alpha + 2\Psi) + c_\theta^3 \left( \frac{16}{3} - 16\nu \right) \cos(3\alpha + 4\Psi) - (\chi_a^y + \delta\chi_s^y) \sin \Psi \left. \right] \\ & + \nu \iota^2 s_\theta \delta \left[ \left( -\frac{3c_\theta^2}{16} + \frac{9}{16} \right) \cos(\alpha - \Psi) - \left( \frac{11c_\theta^2}{32} + \frac{23}{32} \right) \cos(\alpha + \Psi) + \frac{27}{32}(c_\theta^2 + 1)\cos 3(\alpha + \Psi) - \frac{1}{32}(c_\theta^2 + 1) \right. \\ & \left. \times \cos(3\alpha + \Psi) + \left( -\frac{45}{32} + \frac{135c_\theta^2}{32} \right) \cos(\alpha + 3\Psi) \right] + \iota^3 c_\theta s_\theta \left[ \frac{1}{2} \cos(\alpha - 2\Psi) + \frac{5}{6} \cos(\alpha + 2\Psi) \right] \end{aligned} \quad (3.16g)$$

$$H_{\times}^{(0)} = -2c_{\theta} \sin 2(\alpha + \Psi), \quad (3.17a)$$

$$H_{\times}^{(1/2)} = \nu \delta c_{\theta} s_{\theta} \left[ -\frac{9}{4} \sin 3(\alpha + \Psi) + \frac{3}{4} \sin(\alpha + \Psi) \right], \quad (3.17b)$$

$$H_{\times}^{(1)} = \nu^2 c_{\theta} \left[ \left( -\frac{4c_{\theta}^2}{3} + \frac{17}{3} + \left( -\frac{13}{3} + 4c_{\theta}^2 \right) \nu \right) \sin 2(\alpha + \Psi) + s_{\theta}^2 \left( -\frac{8}{3} + 8\nu \right) \sin 4(\alpha + \Psi) \right], \quad (3.17c)$$

$$H_{\times}^{(3/2)} = \nu^3 c_{\theta} \left[ \delta s_{\theta} \left( \left( -\frac{21}{32} + \frac{5c_{\theta}^2}{96} \right) + \left( -\frac{5c_{\theta}^2}{48} + \frac{23}{48} \right) \nu \right) \sin(\alpha + \Psi) - 4\pi \sin 2(\alpha + \Psi) \right. \\ \left. + \delta s_{\theta} \left( \left( -\frac{135c_{\theta}^2}{64} + \frac{603}{64} \right) + \left( -\frac{171}{32} + \frac{135c_{\theta}^2}{32} \right) \nu \right) \sin 3(\alpha + \Psi) + \delta s_{\theta} \left( \frac{625}{192} (2\nu - 1) s_{\theta}^2 \right) \sin 5(\alpha + \Psi) \right], \quad (3.17d)$$

$$H_{\times}^{(1/2, \text{SO})} = -2\iota s_{\theta} \sin(\alpha + 2\Psi), \quad (3.17e)$$

$$H_{\times}^{(1, \text{SO})} = \nu^2 [( \chi_a^y + \delta \chi_s^y ) \cos(\alpha + \Psi) + c_{\theta} (c_{\theta} (\chi_a^x + \delta \chi_s^x) - s_{\theta} (\chi_a^z + \delta \chi_s^z)) \sin(\alpha + \Psi)] \\ + \iota \nu \delta \left[ s_{\theta}^2 \sin \Psi - \left( \frac{c_{\theta}^2}{2} + \frac{1}{4} \right) \sin(2\alpha + \Psi) + \left( -\frac{9}{4} + \frac{9c_{\theta}^2}{2} \right) \sin(2\alpha + 3\Psi) \right] + \iota^2 c_{\theta} \sin 2(\alpha + \Psi), \quad (3.17f)$$

$$H_{\times}^{(3/2, \text{SO})} = \nu^3 \left[ s_{\theta} (2\delta \chi_a^y + (2 - \nu) \chi_s^y) (1 + \cos 2(\alpha + \Psi)) + \left( \frac{8}{3} c_{\theta} \delta \chi_a^z + c_{\theta} \left( \frac{8}{3} - \left( \frac{4}{3} + 4c_{\theta}^2 \right) \nu \right) \chi_s^z \right. \right. \\ \left. \left. - s_{\theta} (2\delta \chi_a^x + (2 + (3 + 4c_{\theta}^2) \nu) \chi_s^x) \right) \sin 2(\alpha + \Psi) \right] + \iota \nu^2 s_{\theta} [(c_{\theta} (\chi_a^x + \delta \chi_s^x) - s_{\theta} (\chi_a^z + \delta \chi_s^z)) \sin \Psi \\ + \left( (-3c_{\theta}^2 + 6) + \left( -\frac{16}{3} + 9c_{\theta}^2 \right) \nu \right) \sin(\alpha + 2\Psi) + \left( -\left( c_{\theta}^2 + \frac{1}{3} \right) + (3c_{\theta}^2 + 1) \nu \right) \sin(3\alpha + 2\Psi) \\ + \left( \left( -\frac{8}{3} + 8c_{\theta}^2 \right) + (-24c_{\theta}^2 + 8) \nu \right) \sin(3\alpha + 4\Psi) ] + \iota^2 \nu \delta c_{\theta} s_{\theta} \left[ \frac{3}{8} \sin(\alpha - \Psi) - \frac{17}{16} \sin(\alpha + \Psi) \right. \\ \left. + \frac{27}{16} \sin 3(\alpha + \Psi) - \frac{1}{16} \sin(3\alpha + \Psi) + \frac{45}{16} \sin(\alpha + 3\Psi) \right] + \iota^3 \left[ \frac{1}{2} s_{\theta} \sin(\alpha - 2\Psi) + \frac{5}{6} s_{\theta} \sin(\alpha + 2\Psi) \right], \quad (3.17g)$$

To complete the waveforms, one should determine the following quantities,

- Invariant velocity:  $v(t)$
- Total phase:  $\Psi(t)$
- Precession phase:  $\alpha(t)$
- Inclination angle:  $\iota(t)$
- Spin vectors:  $\chi_i = \mathbf{S}_i/M_i^2$  ( $i = 1, 2$ )
- Amplitude factors:  $F^{(n)} = Mv^{n+2}/2D_L$  ( $n = 0, \frac{1}{2}, 1, \frac{3}{2}, \dots$ : PN-order),

by solving the following equations simultaneously,

- Spin-precession eqns:  $\dot{\mathbf{S}}_1 = \boldsymbol{\Omega}_1 \times \mathbf{S}_1$ ,  $\dot{\mathbf{S}}_2 = \boldsymbol{\Omega}_2 \times \mathbf{S}_2$  ;  
 $\boldsymbol{\Omega}_{1,2} = M^{2/3}\omega_{\text{orb}}^{5/3} \left(\frac{3}{4} + \frac{\nu}{2} \mp \frac{3}{4}\delta\right) \hat{\mathbf{L}}_N$ ,  $\nu = \frac{M_1M_2}{M^2}$ ,  $\delta = \frac{M_1-M_2}{M}$ ,  $M = M_1 + M_2$
- Spin/angular momentum evolution eqn:  $\dot{\hat{\mathbf{L}}}_N = -\frac{v}{\nu M^2} (\dot{\mathbf{S}}_1 + \dot{\mathbf{S}}_2)$  ;  $v \equiv (M\omega_{\text{orb}})^{1/3}$
- Evolution eqn for orbital frequency:

$$\frac{\dot{\omega}_{\text{orb}}}{\omega_{\text{orb}}^2} = \frac{96}{5}\nu v^5 \left\{ 1 - \left(\frac{743}{336} + \frac{11}{4}\nu\right) v^2 + \left[\left(\frac{19}{3}\nu - \frac{113}{12}\right) \chi_s \cdot \hat{\mathbf{L}}_N - \frac{113}{12}\delta\chi_a \cdot \hat{\mathbf{L}}_N\right] v^3 + 4\pi v^3 \right\}; \quad \chi_{s/a} = \frac{1}{2}(\chi_1 \pm \chi_2)$$

## A. Equal-mass binaries

For binaries with  $M_1 = M_2 = \frac{M}{2}$ ,  $\delta = \frac{M_1 - M_2}{M} = 0$ ,  $\nu = \frac{M_1 M_2}{M^2} = \frac{1}{4}$ , we solve the following equations simultaneously:

- Spin-precession eqns: with  $\mathbf{S} \equiv \mathbf{S}_1 + \mathbf{S}_2$  and  $\bar{\mathbf{S}} \equiv \mathbf{S}_1 - \mathbf{S}_2$

$$\dot{\mathbf{S}} = \boldsymbol{\Omega} \times \mathbf{S}, \quad \dot{\bar{\mathbf{S}}} = \boldsymbol{\Omega} \times \bar{\mathbf{S}} \quad ; \quad \boldsymbol{\Omega} \equiv \frac{7}{8} M^{2/3} \omega_{\text{orb}}^{5/3} \hat{\mathbf{L}}_N$$

- Spin/angular momentum evolution eqn:  $\dot{\hat{\mathbf{L}}}_N = -\frac{4v}{M^2} \dot{\mathbf{S}}$  ;  $v \equiv (M\omega_{\text{orb}})^{1/3}$

- Evolution eqn for orbital frequency:

$$\frac{\dot{\omega}_{\text{orb}}}{\omega_{\text{orb}}^2} = \frac{24}{5} v^5 \left[ 1 - \frac{487}{168} v^2 + \left( 4\pi - \frac{47}{6} \chi_s \cdot \hat{\mathbf{L}}_N \right) v^3 \right]; \quad \chi_s \equiv \frac{2\mathbf{S}}{M^2}$$

In the basis  $\{\vec{e}_{x'}(t), \vec{e}_{y'}(t), \vec{e}_{z'}(t) = \hat{\mathbf{L}}_N\}$  we obtain trivial solutions:

$$\mathbf{S} = \mathbf{S}_o = \text{const}, \quad \bar{\mathbf{S}} = \bar{\mathbf{S}}_o = \text{const} \quad \Rightarrow \quad \chi_s = \chi_{so} = 2\mathbf{S}_o/M^2, \quad \chi_a = \chi_{ao} = 2\bar{\mathbf{S}}_o/M^2$$

Then project the solutions into  $\{\vec{e}_x, \vec{e}_y, \vec{e}_z\}$  basis via:

$$\begin{bmatrix} \vec{e}_{x'}(t) \\ \vec{e}_{y'}(t) \\ \vec{e}_{z'}(t) \end{bmatrix} = \begin{bmatrix} -\sin \alpha & \cos \alpha & 0 \\ -\cos \iota \cos \alpha & -\cos \iota \sin \alpha & \sin \iota \\ \sin \iota \cos \alpha & \sin \iota \sin \alpha & \cos \iota \end{bmatrix} \begin{bmatrix} \vec{e}_x \\ \vec{e}_y \\ \vec{e}_z \end{bmatrix}$$

The following are determined:

- $v(t) = \frac{1}{2}\Theta^{-1/8} \left[ 1 + \frac{487}{4032}\Theta^{-1/4} + \left( -\frac{\pi}{10} + p\frac{47\chi_{so} \cos \beta_o}{240} \right) \Theta^{-3/8} + \mathcal{O}(\Theta^{-1/2}) \right]$  ;  $\Theta \equiv \frac{t_c - t}{20M}$   
( $t_c$ : coalescence time)
- $\Psi(t) = -4\Theta^{5/8} \left[ 1 + \left( \frac{2435}{4032} - p\frac{35}{96} \right) \Theta^{-1/4} + \left( -\frac{3\pi}{4} + p\frac{59\chi_{so} \cos \beta_o}{64} \right) \Theta^{-3/8} + \mathcal{O}(\Theta^{-1/2}) \right]$  ;
- $\alpha(t) = -p\frac{35}{24}\Theta^{3/8} \left[ 1 + \frac{3}{2}\chi_{so} \cos \beta_o \Theta^{-1/8} + \mathcal{O}(\Theta^{-1/4}) \right]$
- $\iota \approx p\chi_s \sin \beta_o \Theta^{-1/8} \left[ 1 + \frac{487}{4032}\Theta^{-1/4} + \left( -\frac{\pi}{10} + \frac{47}{240}\chi_{so} \cos \beta_o \right) \Theta^{-3/8} + \mathcal{O}(\Theta^{-1/2}) \right]$
- $\begin{bmatrix} \chi_s^x \\ \chi_s^y \\ \chi_s^z \end{bmatrix} = \begin{bmatrix} -\chi_{so} \sin \beta_o \cos(\alpha(t)) \\ -\chi_{so} \sin \beta_o \sin(\alpha(t)) \\ \chi_{so} \cos \beta_o \end{bmatrix}, \quad \begin{bmatrix} \chi_a^x \\ \chi_a^y \\ \chi_a^z \end{bmatrix} = \begin{bmatrix} -\chi_{ao} \sin \bar{\beta}_o \cos(\alpha(t)) \\ -\chi_{ao} \sin \bar{\beta}_o \sin(\alpha(t)) \\ \chi_{ao} \cos \bar{\beta}_o \end{bmatrix}$
- $F(0) = \frac{Mv^2}{2D_L} = \frac{M}{8D_L}\Theta^{-1/4}, \quad F(1/2) = \frac{Mv^3}{2D_L} = \frac{M}{16D_L}\Theta^{-3/8},$   
 $F(1) = \frac{Mv^4}{2D_L} = \frac{M}{32D_L}\Theta^{-1/2} \left( 1 + \frac{487}{1008}\Theta^{-1/4} \right),$   
 $F(3/2) = \frac{Mv^5}{2D_L} = \frac{M}{64D_L}\Theta^{-5/8} \left[ 1 + \frac{2435}{4032}\Theta^{-1/4} + \left( -\frac{\pi}{2} + p\frac{47\chi_{so} \cos \beta_o}{48} \right) \Theta^{-3/8} \right]$

$p = 0$  for non-spinning case,  $p = 1$  for spinning case,  $\chi_{so} \equiv 2S_o/M^2 = \text{const}$ ,  
 $\chi_{ao} \equiv 2\bar{S}_o/M^2 = \text{const}$  ;  $\mathbf{S} \equiv \mathbf{S}_1 + \mathbf{S}_2 = \mathbf{S}_o = \text{const}$ ,  $\bar{\mathbf{S}} \equiv \mathbf{S}_1 - \mathbf{S}_2 = \bar{\mathbf{S}}_o = \text{const}$ ,  
 $\beta_o \equiv \cos^{-1} \left( \frac{\mathbf{L}_N \cdot \mathbf{S}}{L_N S} \right) = \text{const}$ ,  $\bar{\beta}_o \equiv \cos^{-1} \left( \frac{\mathbf{L}_N \cdot \bar{\mathbf{S}}}{L_N \bar{S}} \right) = \text{const} \Rightarrow$  all determined by **initial conditions**

## B. Unequal-mass binaries

For binaries with  $M_1 \neq M_2$ ,  $\delta = \frac{M_1 - M_2}{M}$ ,  $\nu = \frac{M_1 M_2}{M^2} = \frac{1 - \delta^2}{4}$ , we solve the following equations simultaneously:

- Spin-precession eqns: with  $\mathbf{S} \equiv \mathbf{S}_1 + \mathbf{S}_2$  and  $\bar{\mathbf{S}} \equiv \mathbf{S}_1 - \mathbf{S}_2$

$$\dot{\mathbf{S}} = \left(1 - \frac{\delta^2}{7}\right) \boldsymbol{\Omega} \times \mathbf{S} - \frac{6\delta}{7} \boldsymbol{\Omega} \times \bar{\mathbf{S}}, \quad \dot{\bar{\mathbf{S}}} = \left(1 - \frac{\delta^2}{7}\right) \boldsymbol{\Omega} \times \bar{\mathbf{S}} - \frac{6\delta}{7} \boldsymbol{\Omega} \times \mathbf{S} \quad ; \quad \boldsymbol{\Omega} \equiv \frac{7}{8} M^{2/3} \omega_{\text{orb}}^{5/3} \hat{\mathbf{L}}_N$$

- Spin/angular momentum evolution eqn:  $\dot{\hat{\mathbf{L}}}_N = -\frac{v}{\nu M^2} \dot{\mathbf{S}} \quad ; \quad v \equiv (M \omega_{\text{orb}})^{1/3}$

- Evolution eqn for orbital frequency:

$$\frac{\dot{\omega}_{\text{orb}}}{\omega_{\text{orb}}^2} = \frac{96}{5} \nu v^5 \left\{ 1 - \left( \frac{743}{336} + \frac{11}{4} \nu \right) v^2 + \left[ \left( \frac{19}{3} \nu - \frac{113}{12} \right) \chi_s \cdot \hat{\mathbf{L}}_N - \frac{113}{12} \delta \chi_a \cdot \hat{\mathbf{L}}_N \right] v^3 + 4\pi v^3 \right\}; \quad \chi_s \equiv \frac{2\mathbf{S}}{M^2}, \quad \chi_a \equiv \frac{2\bar{\mathbf{S}}}{M^2}$$

Solve the equations in the basis  $\{\vec{e}_{x'}(t), \vec{e}_{y'}(t), \vec{e}_{z'}(t) = \hat{\mathbf{L}}_N\}$  first and project the solutions into  $\{\vec{e}_x, \vec{e}_y, \vec{e}_z\}$  basis via:

$$\begin{bmatrix} \vec{e}_{x'}(t) \\ \vec{e}_{y'}(t) \\ \vec{e}_{z'}(t) \end{bmatrix} = \begin{bmatrix} -\sin \alpha & \cos \alpha & 0 \\ -\cos \iota \cos \alpha & -\cos \iota \sin \alpha & \sin \iota \\ \sin \iota \cos \alpha & \sin \iota \sin \alpha & \cos \iota \end{bmatrix} \begin{bmatrix} \vec{e}_x \\ \vec{e}_y \\ \vec{e}_z \end{bmatrix}$$



In the basis  $\{\vec{e}_{x'}(t), \vec{e}_{y'}(t), \vec{e}_{z'}(t) = \hat{\mathbf{L}}_N\}$  the equations reduce to

$$\begin{aligned} S^{x'} &= 0, & S^{z'} &= \mathbf{S} \cdot \hat{\mathbf{L}}_N = \text{const} = S_o \cos \beta_o, & \bar{S}^{z'} &= \bar{\mathbf{S}} \cdot \hat{\mathbf{L}}_N = \text{const} = \bar{S}_o \cos \bar{\beta}_o, \\ \dot{S}^{y'} &= -\frac{6\delta}{7}\Omega \bar{S}^{x'}, & S^{y'} \bar{S}^{y'} &= \text{const} = S_o^{y'} \bar{S}_o^{y'}, & \dot{\bar{S}}^{x'} &= -\frac{6\delta}{7}\Omega (S_o^{y'} \bar{S}_o^{y'})^2 (S^{y'})^{-3} + \frac{6\delta}{7}\Omega S^{y'}, \\ \dot{\alpha} &= \left(1 - \frac{\delta^2}{7}\right) \Omega - \frac{6\delta}{7}\Omega (\bar{S}^{y'}/S^{y'}); & \Omega &\equiv \frac{7}{8}M^{2/3}\omega_{\text{orb}}^{5/3} = \frac{7}{8}M^{-1}v^5 \end{aligned}$$

Solving for  $S^{y'}$ ,

$$\ddot{S}^{y'} - \frac{32\nu}{M}v^8 \dot{S}^{y'} - \left(\frac{6\delta}{7}\Omega\right)^2 (S_o^{y'} \bar{S}_o^{y'})^2 (S^{y'})^{-3} + \left(\frac{6\delta}{7}\Omega\right)^2 S^{y'} = 0 \Rightarrow \text{anharmonic oscillator}$$

Ignoring the damping term, and splitting  $S^{y'} = S_o^{y'} + \Delta^{y'}$ , and then taking  $(S^{y'})^{-3} = (S_o^{y'})^{-3} \left[1 - 3(S_o^{y'})^{-1} \Delta^{y'} + \mathcal{O}(\Delta^2)\right]$  (treating  $\mathcal{O}(\Delta^2)$  as higher order perturbation pieces),

$$\ddot{\Delta}^{y'} + \left(\frac{6\delta}{7} \sqrt{1 + 3 \left(\frac{\bar{S}_o^{y'}}{S_o^{y'}}\right)^2} \Omega\right)^2 \left(\Delta^{y'} + \frac{S_o^{y'} \left[(S_o^{y'})^2 - (\bar{S}_o^{y'})^2\right]}{(S_o^{y'})^2 + 3(\bar{S}_o^{y'})^2}\right) = 0 \Rightarrow \text{harmonic oscillator}$$

Solutions:

$$\begin{aligned}
S^{x'} &= 0, \\
S^{y'} &= S_o^{y'} + \mathcal{A} \sin (f (v) + \varphi) + C, \\
S^{z'} &= S_o \cos \beta_o, \\
\bar{S}^{x'} &= \sqrt{1 + 3 (\bar{S}_o^{y'} / S_o^{y'})^2} \mathcal{A} \cos (f (v) + \varphi), \\
\bar{S}^{y'} &= - (\bar{S}_o^{y'} / S_o^{y'}) \mathcal{A} \sin (f (v) + \varphi) + \bar{S}_o^{y'} - C (\bar{S}_o^{y'} / S_o^{y'}), \\
\bar{S}^{z'} &= \bar{S}_o \cos \bar{\beta}_o,
\end{aligned}$$

where

$$\begin{aligned}
f (v) &\equiv -\frac{5\delta}{32(1-\delta^2)} \sqrt{1 + 3 \left(\frac{\bar{S}_o^{y'}}{S_o^{y'}}\right)^2} v^{-3}, \quad C \equiv \frac{S_o^{y'} \left[ (\bar{S}_o^{y'})^2 - (S_o^{y'})^2 \right]}{(S_o^{y'})^2 + 3(\bar{S}_o^{y'})^2}, \quad \mathcal{A} \equiv \sqrt{C^2 + \frac{(\bar{S}_o^{x'})^2 (S_o^{y'})^2}{(S_o^{y'})^2 + 3(\bar{S}_o^{y'})^2}}, \\
S_o &= \sqrt{S_1^2 + S_2^2 + 2S_1S_2 (\cos \theta_1 \cos \theta_2 + \sin \theta_1 \sin \theta_2 \cos (\phi_1 - \phi_2))}, \\
\bar{S}_o &= \sqrt{S_1^2 + S_2^2 - 2S_1S_2 (\cos \theta_1 \cos \theta_2 + \sin \theta_1 \sin \theta_2 \cos (\phi_1 - \phi_2))}, \\
\beta_o &= \arccos \left( \frac{S_1 \cos \theta_1 + S_2 \cos \theta_2}{S_o} \right), \quad \bar{\beta}_o = \arccos \left( \frac{S_1 \cos \theta_1 - S_2 \cos \theta_2}{\bar{S}_o} \right), \\
S_o^{y'} &= S_o \sin \beta_o, \quad \bar{S}_o^{x'} = -\frac{2S_1S_2 \csc \beta_o \sin \theta_1 \sin \theta_2 \sin(\phi_1 - \phi_2)}{S_o}, \quad \bar{S}_o^{y'} = \frac{(S_1^2 - S_2^2) \csc \beta_o}{S_o} - \bar{S}_o \cot \beta_o \cos \bar{\beta}_o
\end{aligned}$$

Determine the following quantities accordingly:

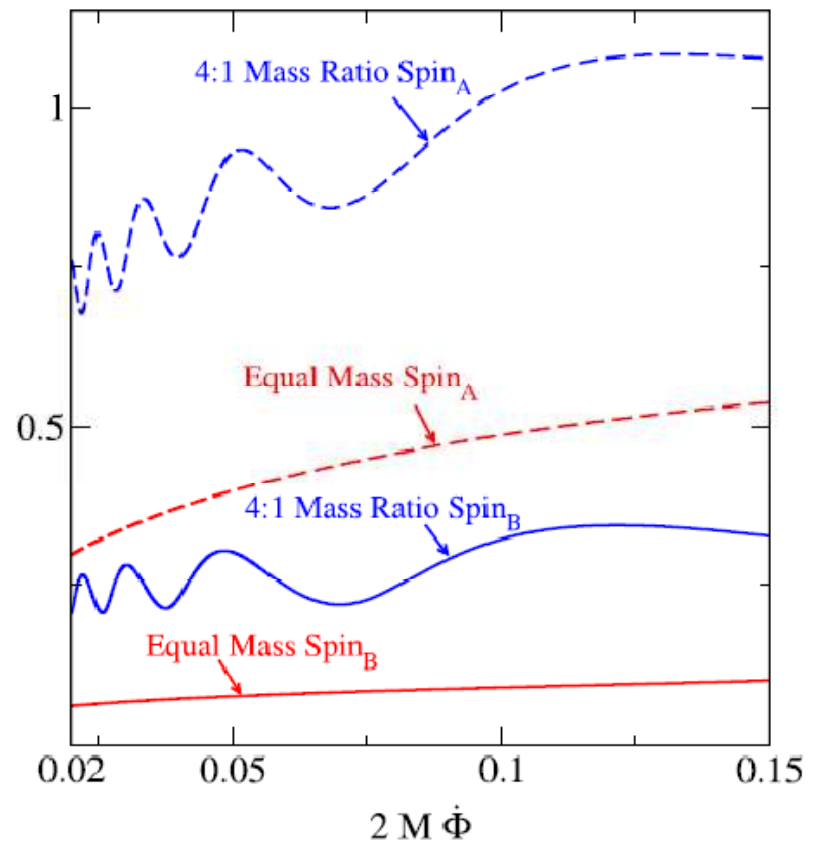
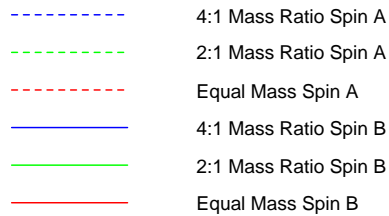
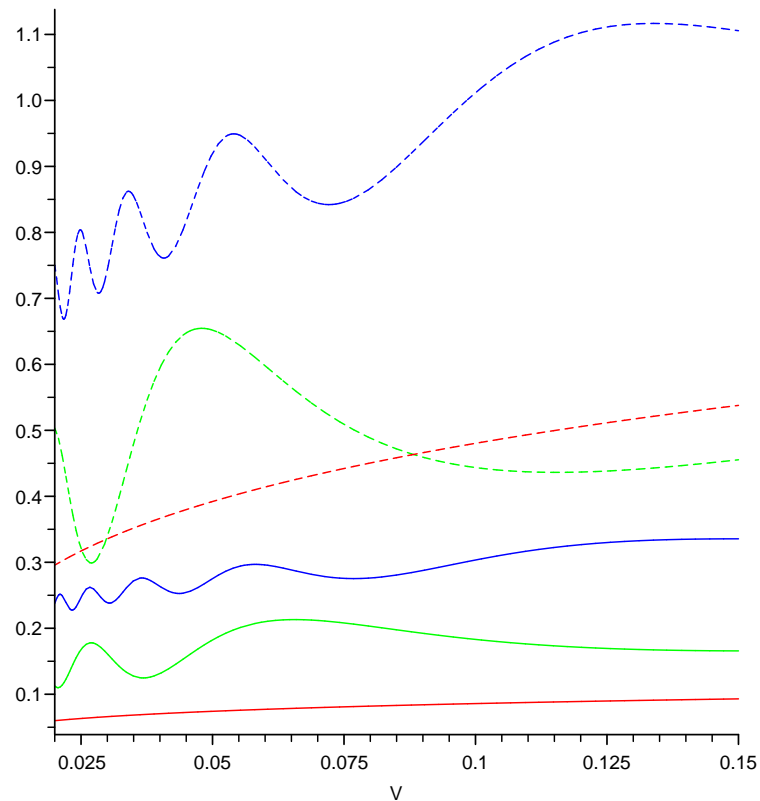
- Invariant velocity:  $v$
- Total phase:  $\Psi$
- Precession phase:  $\alpha$
- Inclination angle:  $\iota$
- Amplitude factors:  $F^{(n)} = Mv^{n+2}/2D_L$  ( $n = 0, \frac{1}{2}, 1, \frac{3}{2}, \dots$ : PN-order),

In particular, the **inclination angle**  $\iota$  undergoes **bobbing** (oscillation) unlike the equal-mass case:

$$\iota = \arcsin\left(\frac{S^{y'}}{J_0}\right) \approx \arcsin\left(\frac{S_0^{y'} + \mathcal{A} \sin(f(v) + \varphi) + C}{\sqrt{L_N^2 + 2L_N S_0^{z'} + (S_0^{x'})^2 + (S_0^{y'})^2 + (S_0^{z'})^2}}\right); \quad L_N = \frac{(1 - \delta^2) M^2}{4v}$$

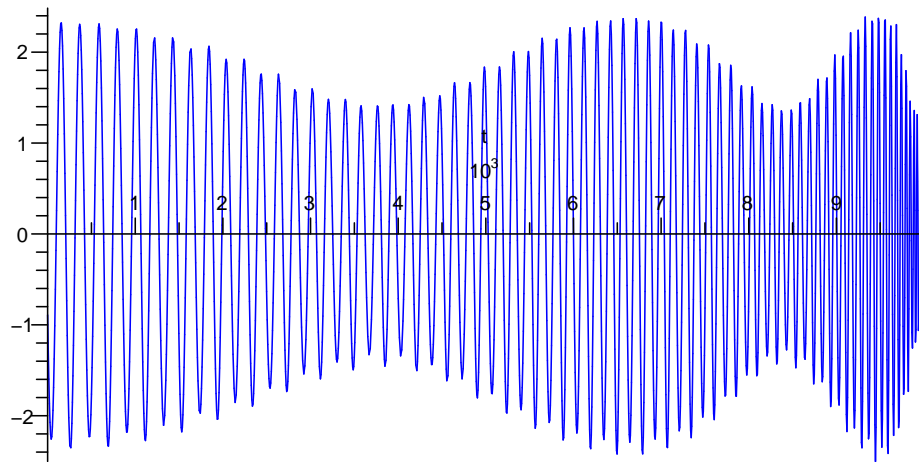
$$f(v) \equiv -\frac{5\delta}{32(1-\delta^2)} \sqrt{1 + 3 \left(\frac{\bar{S}_0^{y'}}{S_0^{y'}}\right)^2} v^{-3}, \quad C \equiv \frac{S_0^{y'} \left[ (\bar{S}_0^{y'})^2 - (S_0^{y'})^2 \right]}{(S_0^{y'})^2 + 3(\bar{S}_0^{y'})^2}, \quad \mathcal{A} \equiv \sqrt{C^2 + \frac{(\bar{S}_0^{x'})^2 (S_0^{y'})^2}{(S_0^{y'})^2 + 3(\bar{S}_0^{y'})^2}} \quad \text{will}$$

characterize the bobbing patterns: **mass-ratios** and **initial spin configurations** are main factors.

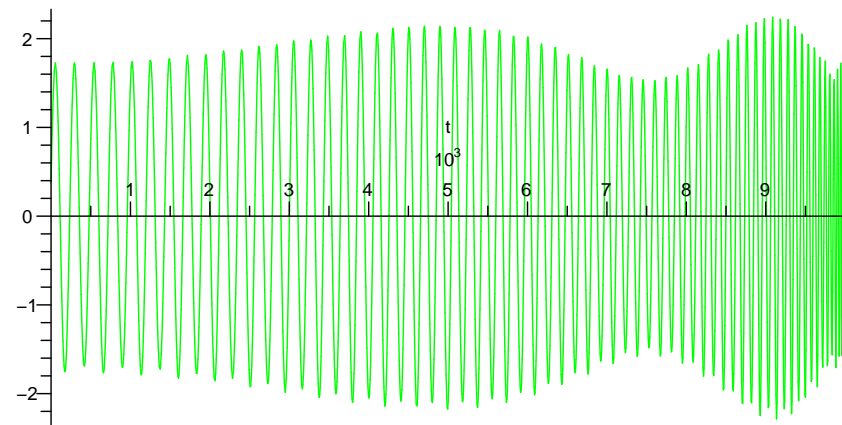


<LEFT> Plots of inclination angle  $\iota$  vs.  $V \equiv 2M\dot{\Phi} \approx 2v^3$  for **4:1 mass-ratio** binary ( $\delta = \frac{3}{5}$ ), **2:1 mass-ratio** binary ( $\delta = \frac{1}{3}$ ) and **equal-mass** binary ( $\delta = 0$ ) constructed from analytic solutions (to 1st order perturbation).

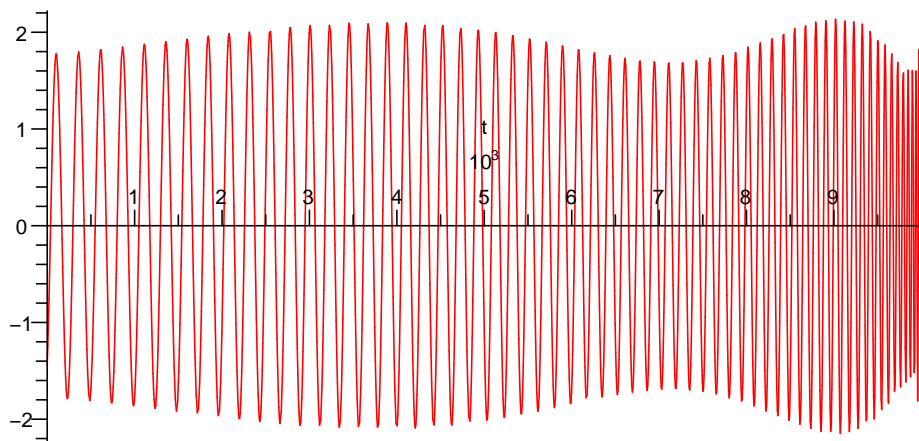
<RIGHT> Numerical plots of  $\iota$  vs.  $V \equiv 2M\dot{\Phi} \approx 2v^3$  for **4:1 mass-ratio** binary ( $\delta = \frac{3}{5}$ ) and **equal-mass** binary ( $\delta = 0$ ) from *Arun et. al*, PRD79-104023 (2009). Initial spin orientations relative to orbital angular momentum are chosen to be  $\text{Spin}_A = \{\theta_1 = \pi/2, \phi_1 = 0, \theta_2 = \pi/2, \phi_2 = \pi/2\}$ ,  $\text{Spin}_B = \{\theta_1 = \pi/6, \phi_1 = \pi/4, \theta_2 = \pi/6, \phi_2 = \pi\}$ .



— 4:1 Mass Ratio Binary



— 2:1 Mass Ratio Binary



— Equal Mass Binary

Plots of  $h_+(t)$  for **4:1 mass-ratio** binary ( $\delta = \frac{3}{5}$ ), **2:1 mass-ratio** binary ( $\delta = \frac{1}{3}$ ) and **equal-mass** binary ( $\delta = 0$ ) through 0.5 PN order: with initial spin orientation relative to orbital angular momentum,  $\{\theta_1 = \pi/2, \phi_1 = 0, \theta_2 = \pi/2, \phi_2 = \pi/2\}$ .

## 4. Conclusions and discussions

- $h_+(t)$  and  $h_\times(t)$  for spinning binaries with **unequal masses** can be determined **analytically** via **perturbation** - through 1.5 PN order (**spin-orbit** interactions) and the spins exhibit **bobbing effects**, which depend on the **mass-ratios** and **initial spin orientations** relative to the orbital angular momentum.
- The bobbing is caused by a purely **kinematical effect of spin**, and in fact ubiquitous in **relativistic mechanics**, occurring independently of the type of force holding two spinning bodies in orbit: an **electromagnetic analog** studied by *Gralla et. al* [Phys. Rev. D 81, 104012 (2010)].
- When **spin-spin** interactions (2 PN) and **radiation reaction** (2.5 PN) are involved, spinning binaries with **general mass ratios** assume much more complicated dynamic evolutions. The spin evolutions can be described by **non-linear oscillators** and **damping** (*Racine* [Phys. Rev. D 78, 044021 (2008)]), which makes the analysis of higher-order PN effects much more difficult.
- Recently, some new techniques (e.g. effective field theory approach) have provided more systematic computational schemes for higher-order PN terms (beyond 2.5 PN). Collaborations with Y. Chen (Caltech), B. L. Hu (Maryland) and C. Galley (JPL) are currently on-going.