

Direct Reactions with Rare Isotopes

Claudio Conti (Itapeva, Brazil)

Mesut Karakoc (Commerce & Akdeniz-Turkey)

Wenhui Long (Lanzhou, China)

Kazu Ogata (Osaka, Japan)



Low Energy problems

Theories: e.g. radiative capture

Schrödinger equation:

e.g., Woods-Saxon

$$-\frac{\hbar^2}{2\mu} \frac{d^2}{dr^2} \psi^l(r) + \left[V(r) + \frac{\hbar^2}{2\mu} \frac{l(l+1)}{r^2} \right] \psi^l(r) = E \psi^l(r)$$

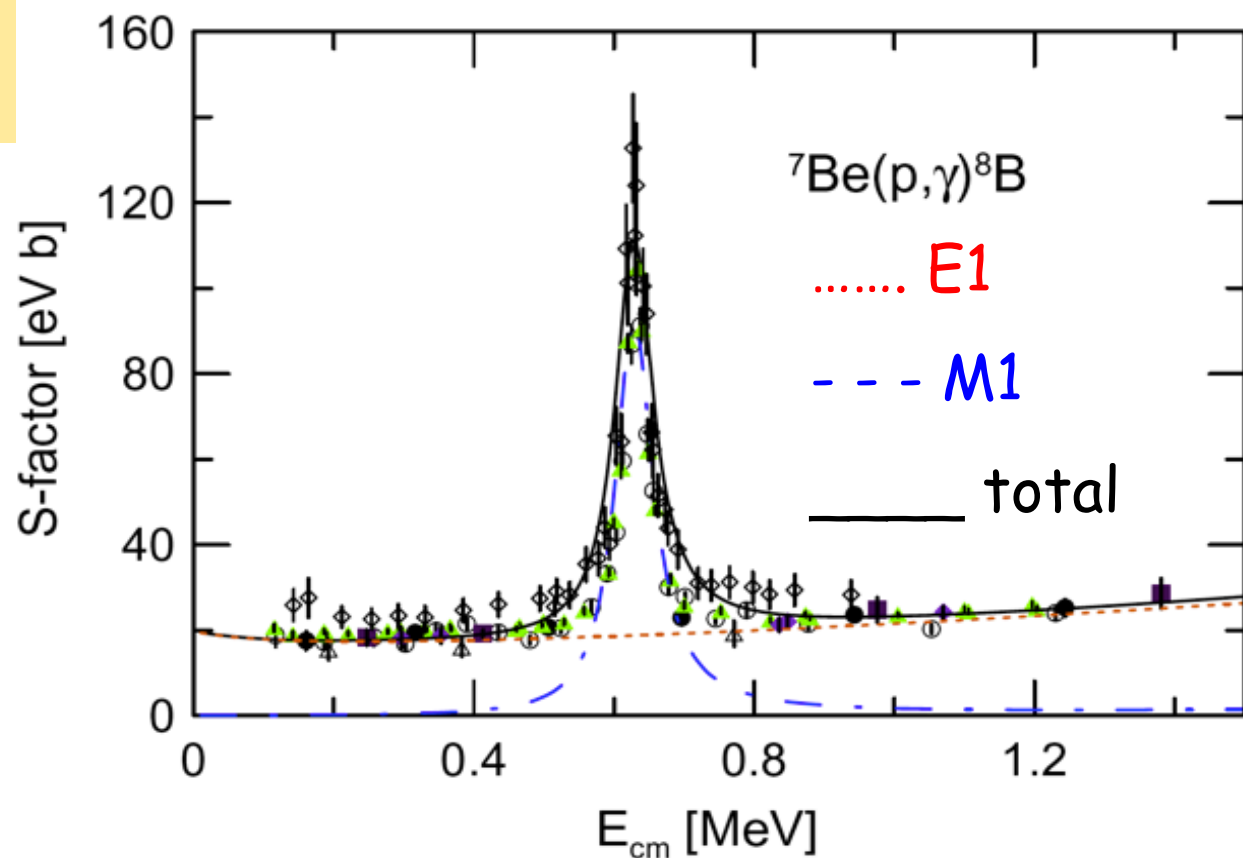
Simple to solve numerically for $E < 0$ or $E > 0$

$$\int_0^\infty r^\lambda \psi^l(E_i, r) \psi^l(E_f, r) dr$$



S-factors, X-sections

Bertulani
Z. Phys. A356, 293 (1996)



Cluster models and reactions

What one needs

$$g(r) = \left\langle \chi^{(A)} \left| \hat{A} \Phi^{(A-a)} \Phi^{(a)} \delta(r - r_{A-a,a}) \right. \right\rangle$$

$$g_{bound}(r) \rightarrow C_{lj} \frac{W_{-\eta, l+1/2}(r)}{r}$$

$$g_{scat}(r \rightarrow \infty) \sim I_l(r) - S_l O_l(r)$$

What one does

$$\int dr' [H(r, r') - EN(r, r')] g(r') = 0 \quad \text{all } r \quad \text{Hill-Wheeler, 1955}$$

$$\begin{Bmatrix} H \\ N \end{Bmatrix}(r, r') = \left\langle \hat{A} \Phi^{(A-a)} \Phi^{(a)}(r') \left| \begin{Bmatrix} H \\ N \end{Bmatrix} \right| \hat{A} \Phi^{(A-a)} \Phi^{(a)}(r) \right\rangle$$

ab initio
model

Problem:

intercluster effective interactions
in continuum

$$H = T + \sum v_{ij} + \sum v_{ijk}$$

Ab initio models: e.g., no core shell model

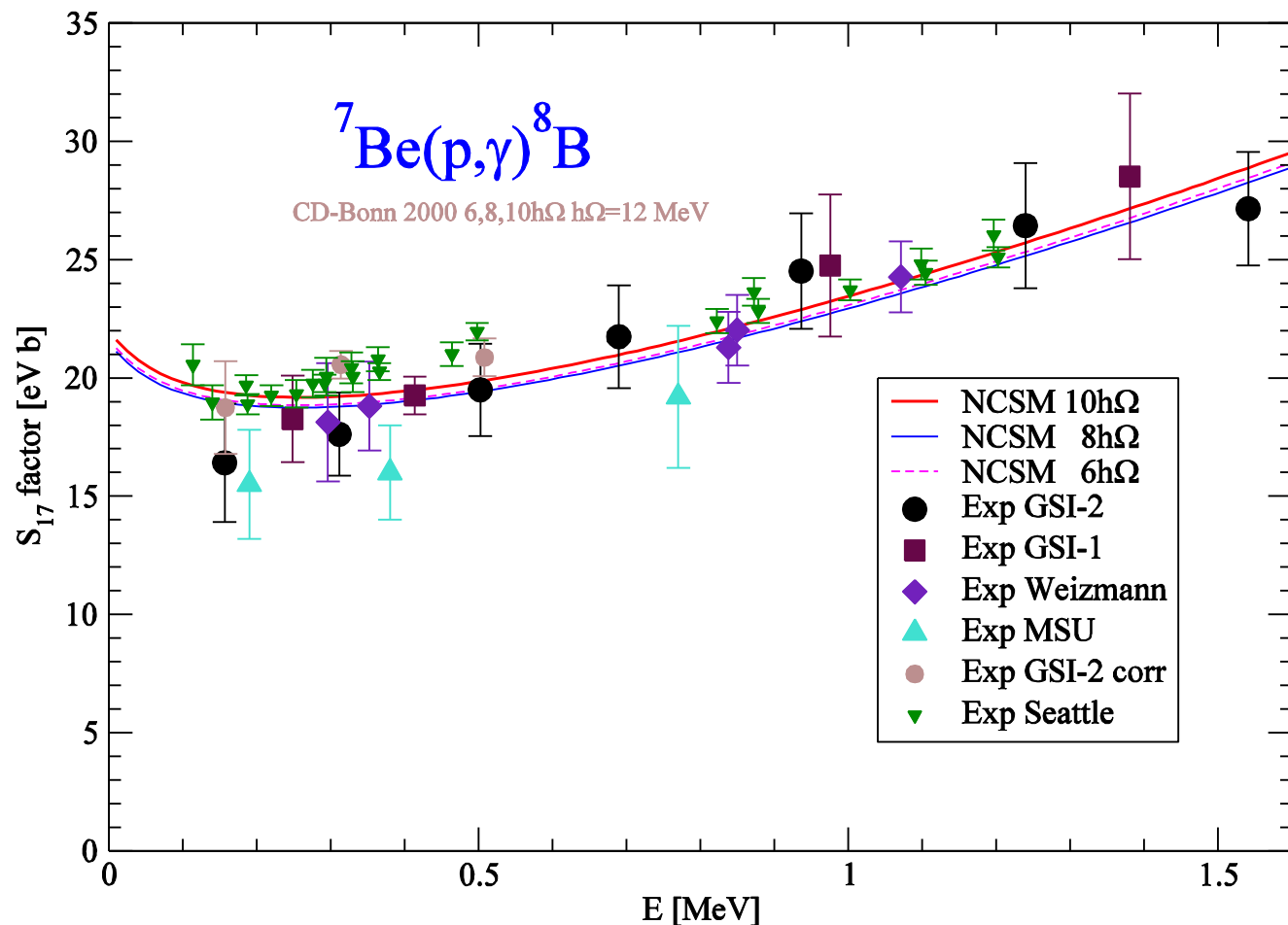
- Realistic interactions: CD-Bonn, INOY
- Accurate wave functions of ${}^7\text{Be}$, ${}^8\text{B}$
- Deduce the ANC of ${}^8\text{B}$

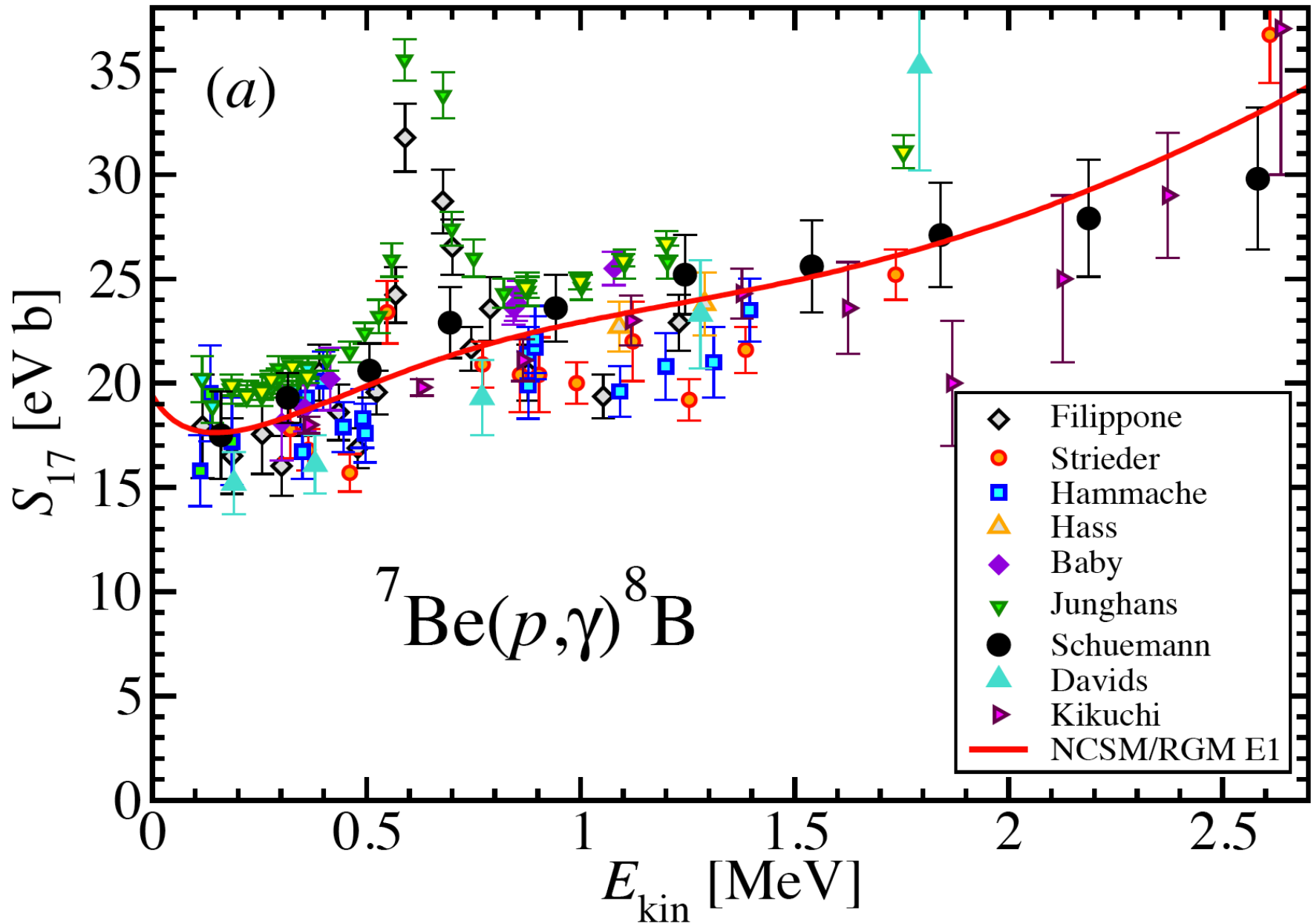
But: ${}^7\text{Be}+p$ scattering states defined in a potential model
 "Ab initio" only in the final state

Convergence test with #
 of oscillator shells

$$S_{17} = 22 \pm 1 \text{ eV}\cdot\text{b}$$

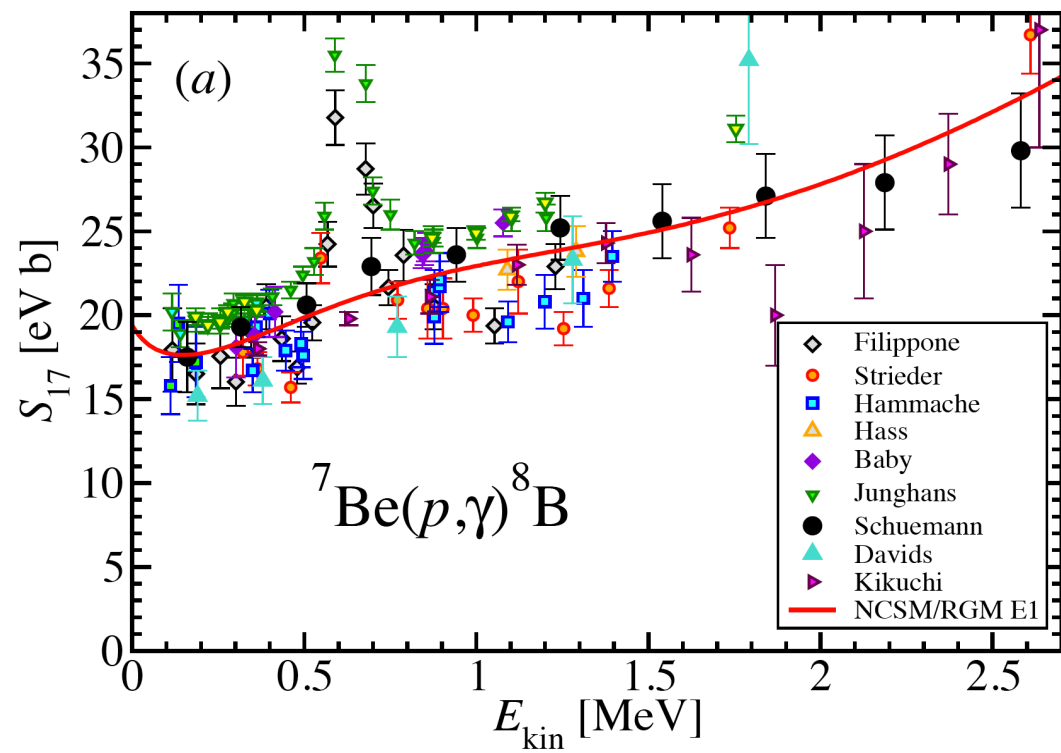
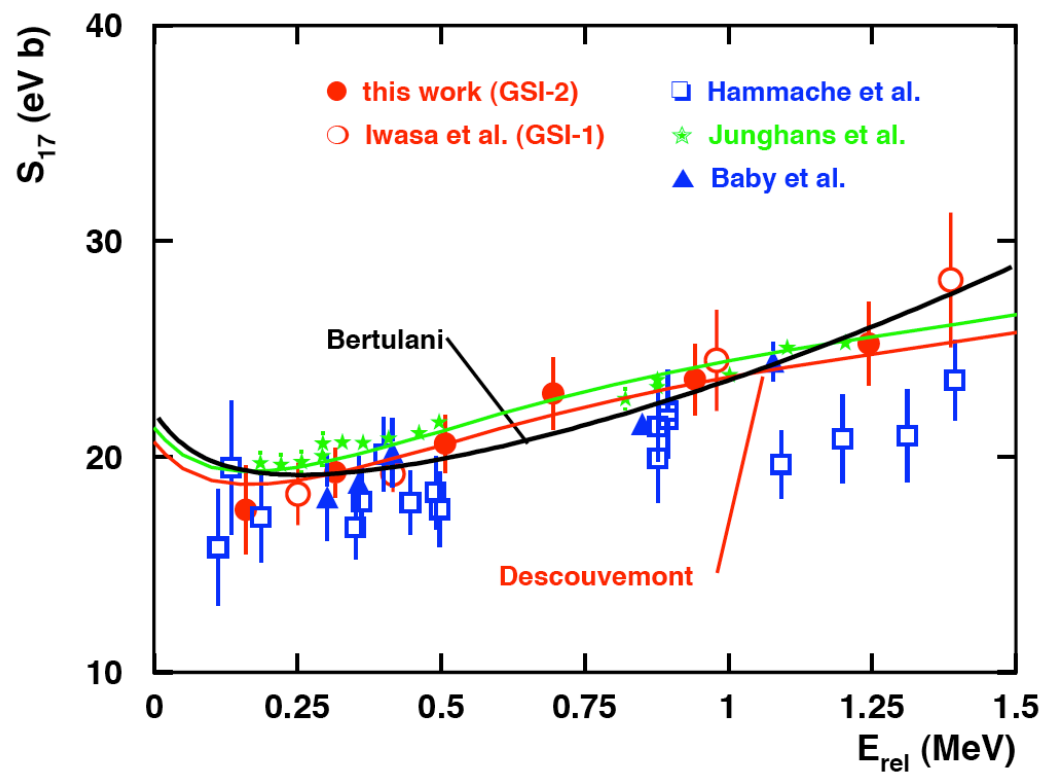
Navratil, Bertulani, Caurier
 PLB 634, 191 (2006)
 PRC 73, 065801 (2006)





$S_{17} = 19.4 \pm 0.7 \text{ eV.b}$

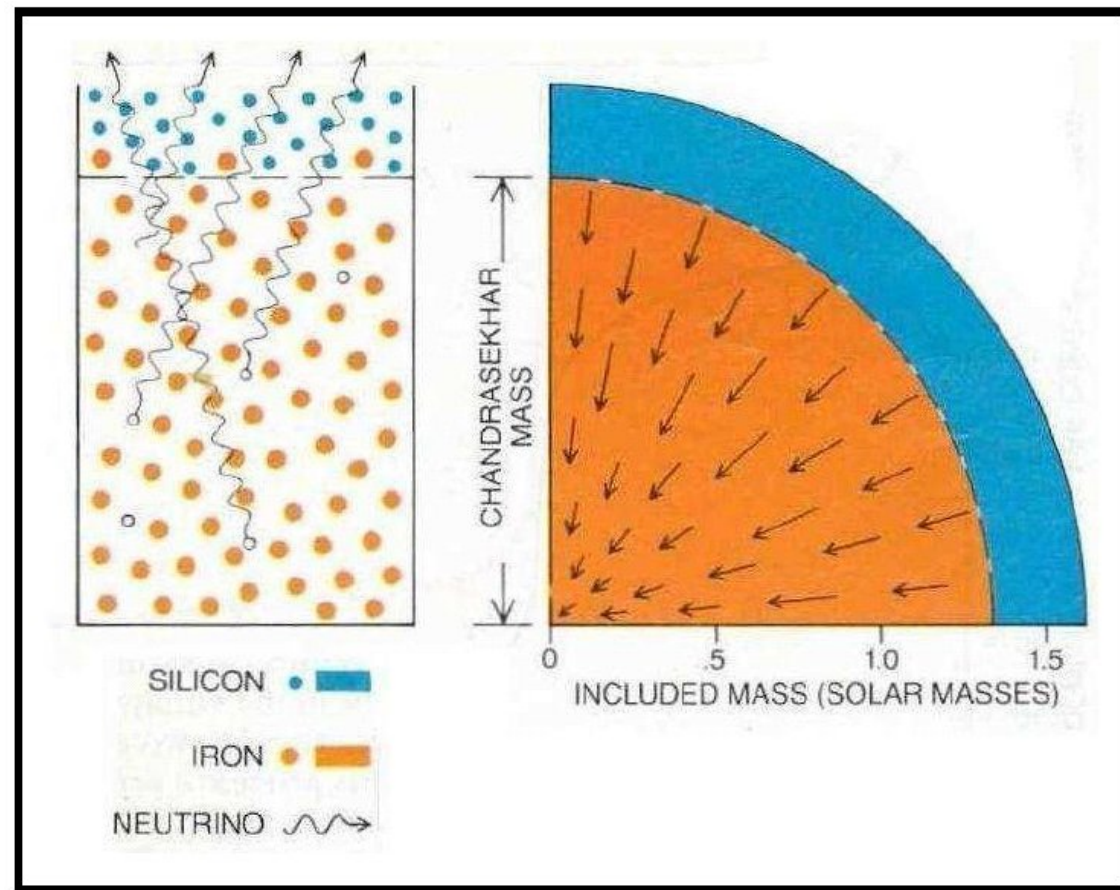
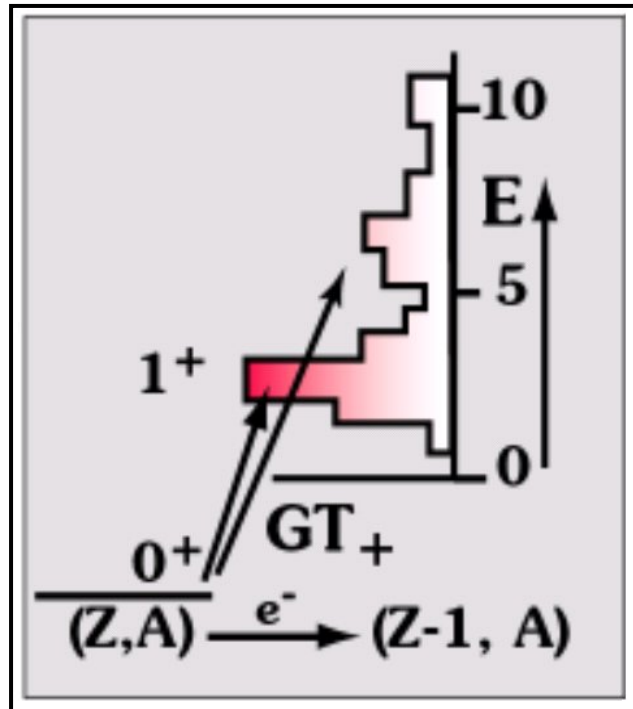
Quaglioni, Navratil, Roth
PLB 704 (2011) 379



More energetic problems



Theories:
e.g. neutrinos

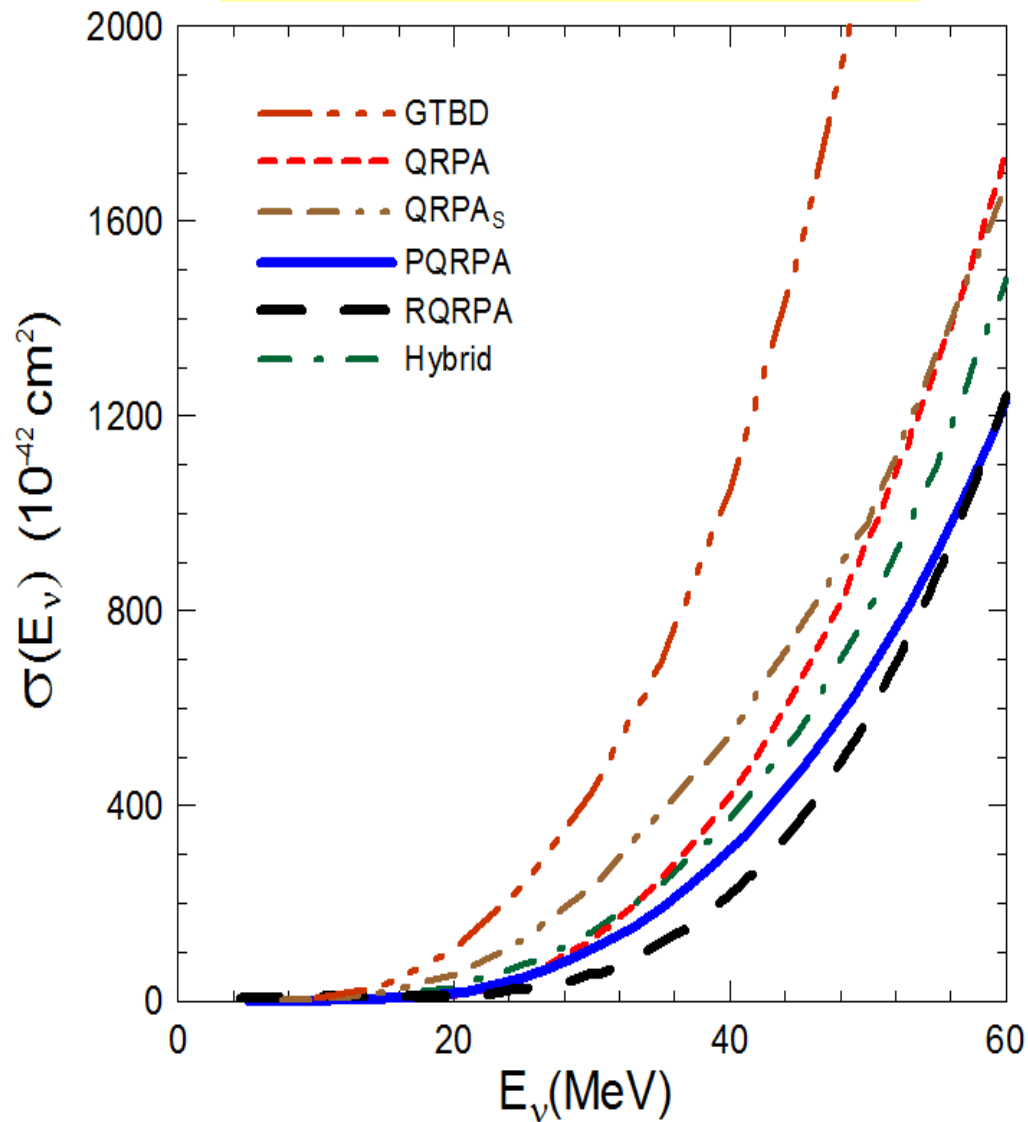
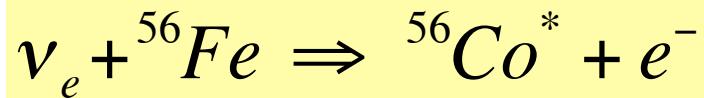


$e^- + (Z,A) \leftrightarrow (Z-1,A) + \nu_e$ Needs $\left| \langle B \| \sigma \tau \| A \rangle \right|^2$ for nuclei $A > 40$

Number of target nuclei Neutrino flux
Interaction cross section Efficiency

$$N_{ev} = N_t \int_0^{\infty} F(E_\nu) \cdot \sigma(E_\nu) \cdot \varepsilon(E_\nu) dE_\nu$$

Supernovae neutrinos



$$\langle \sigma_e \rangle = \int dE_\nu \sigma(E_\nu) n(E_\nu),$$

$$n(E_\nu) = \frac{96E_\nu^2}{M_\mu^4} (M_\mu - 2E_\nu),$$

Model	$\langle \sigma_e \rangle$
QRPA	264.6
PQRPA	197.3
Hybrid ^(a) [14]	228.9
Hybrid ^(b) [14]	238.1
TM [26]	214
RPA [27]	277
QRPA _s [15]	352
RQRPA [16]	140
Exp[5]	$256 \pm 108 \pm 43$

Samana, Bertulani, PRC 78, 024312 (2008)

Samana, Krmpotic, Bertulani, CPC 181, 1123 (2010)

Brilliant ideas



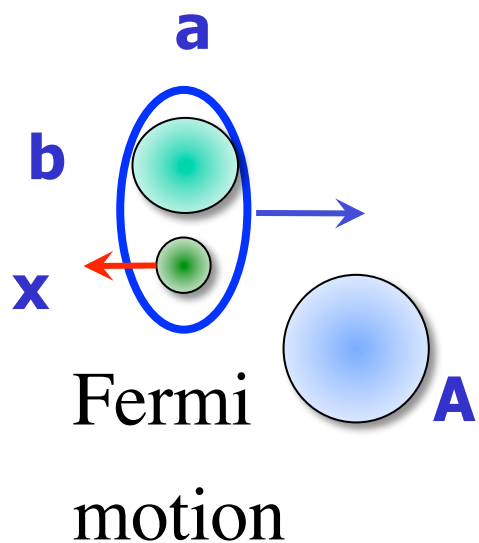
Trojan Horse Method

Astrophysically
relevant

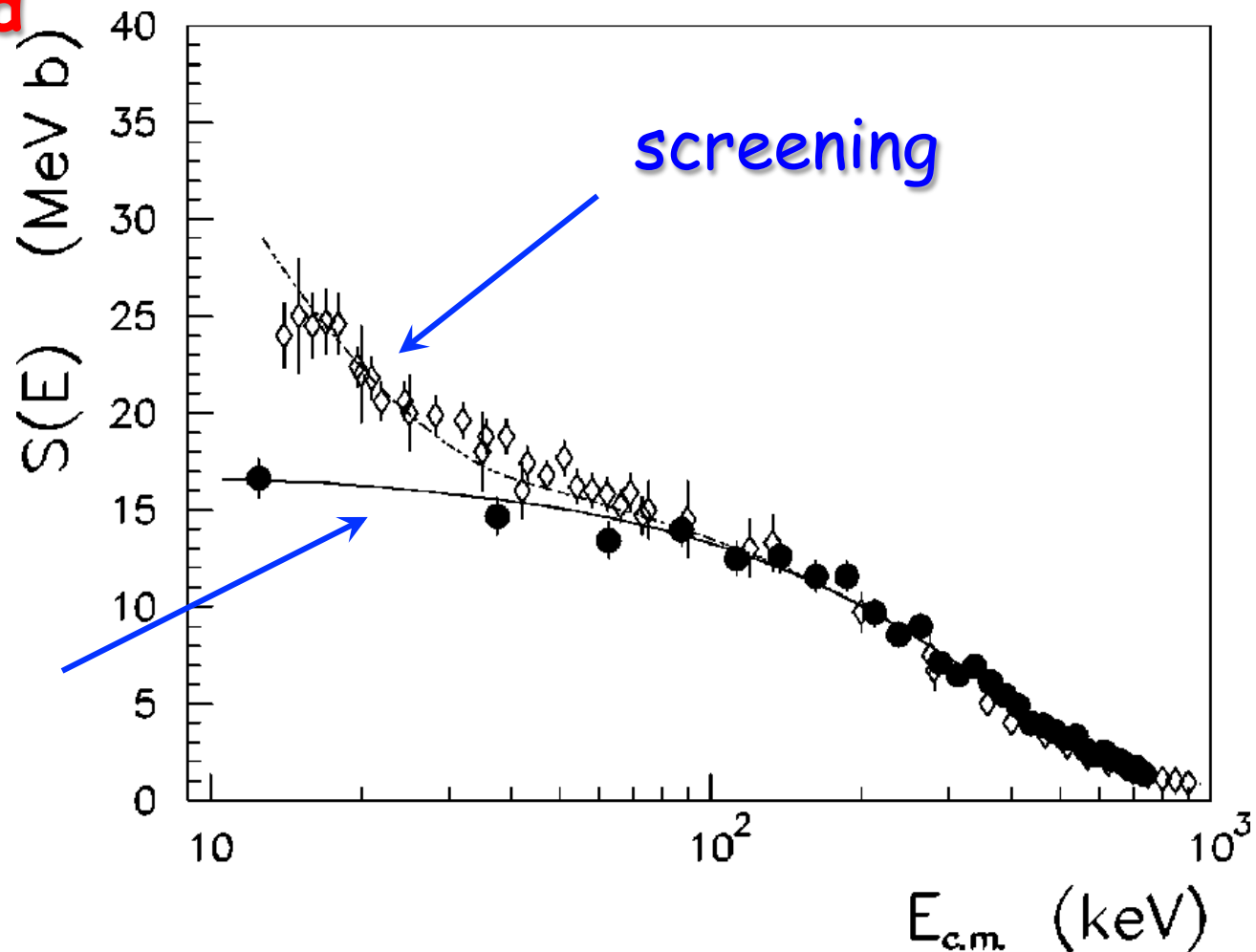
cross section $\sigma(E)$ for
 $B + x \rightarrow C + D$

from transfer

$A + B \rightarrow C + D + s$



THM



${}^2\text{H}({}^6\text{Li}, \alpha){}^4\text{He}$

from

${}^6\text{Li}({}^6\text{Li}, \alpha\alpha){}^4\text{He}$

Multi-nucleon transfer (e.g. Pollaro, Torino)

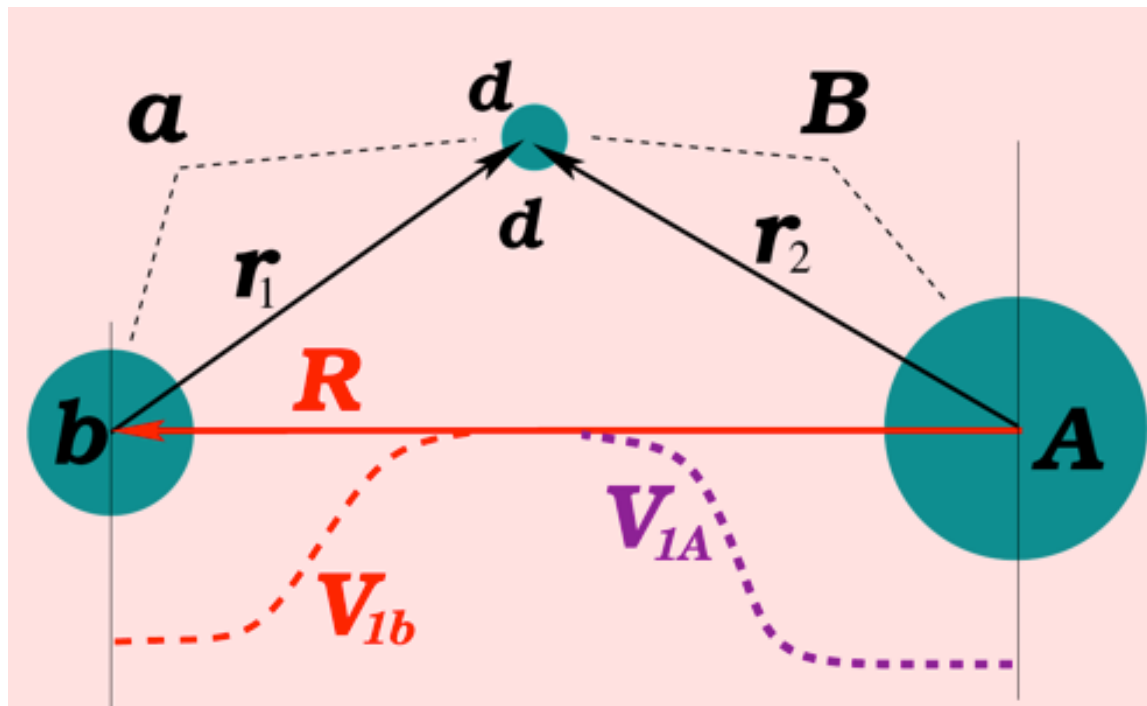
$$P_{\beta} = \left| \frac{i}{\hbar} \int_{-\infty}^{\infty} dt F_{\beta\alpha}(\mathbf{R}) e^{i(E_{\beta}-E_{\alpha})t/\hbar+\dots} \right|^2 \sim \tau_{coll} |F_{\beta\alpha}(D)|^2 g(Q_{\beta\alpha})$$

$$F_{\beta\alpha}(\mathbf{R}) \sim \int d^3\mathbf{r}_1 e^{i\mathbf{Q}\cdot\mathbf{r}_1} \phi_{a_n}^{(A)}(\mathbf{R} + \mathbf{r}_1) (V_{1A} - \langle U \rangle) \phi_{a_n}^{(b)}(\mathbf{r}_1)$$

\mathbf{Q} = momentum transfer

V_{1A} transfer interaction.

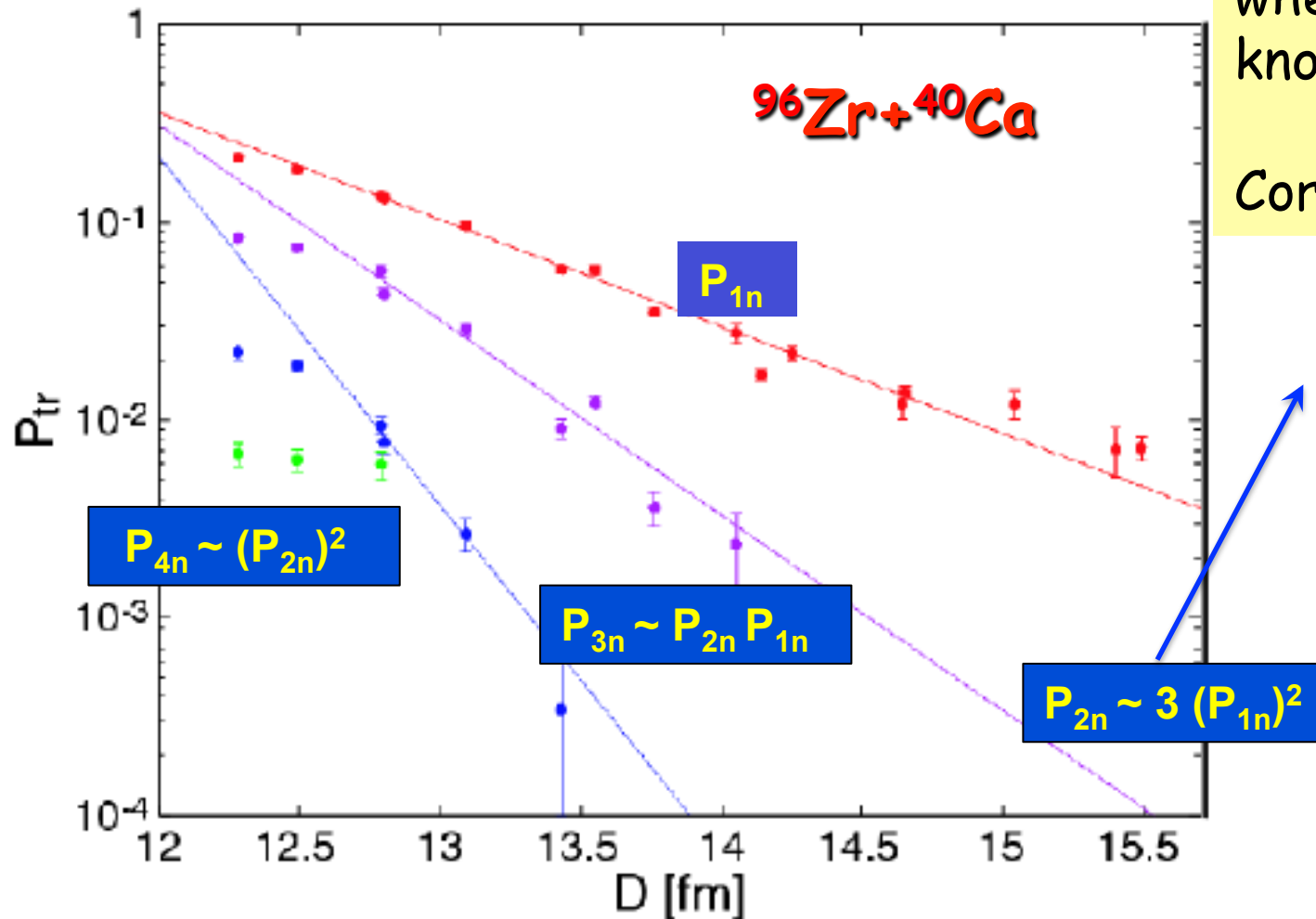
Why not V_{1b} ?? POST-PRIOR representation



$$\frac{P_{tr}}{\sin(\theta_{c.m.}/2)} \propto \exp(-2\alpha D)$$

“That is what happens when theorists do not know what to do”

Corradi, Legnaro, 2011



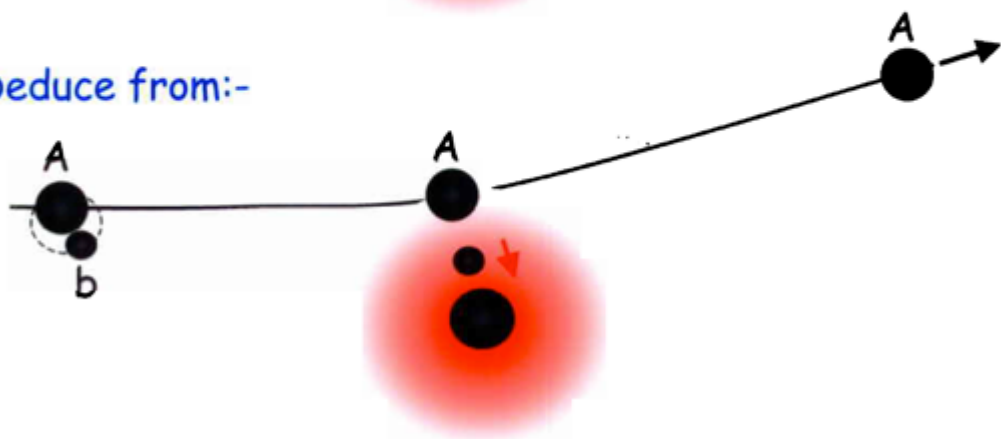
$$D = \frac{Z_1 Z_2 e^2}{2E_{c.m.}} \left(1 + \frac{1}{\sin(\theta_{c.m.}/2)} \right)$$

Surrogate reactions

Reaction of interest



Deduce from:-



e.g., (n,f) from transfer reactions

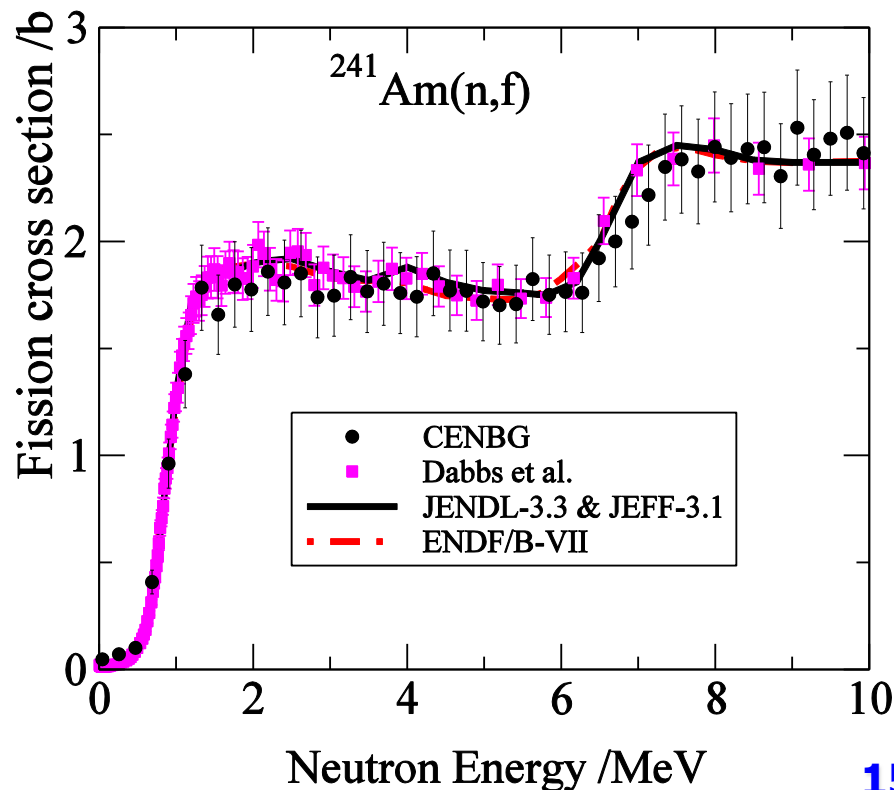
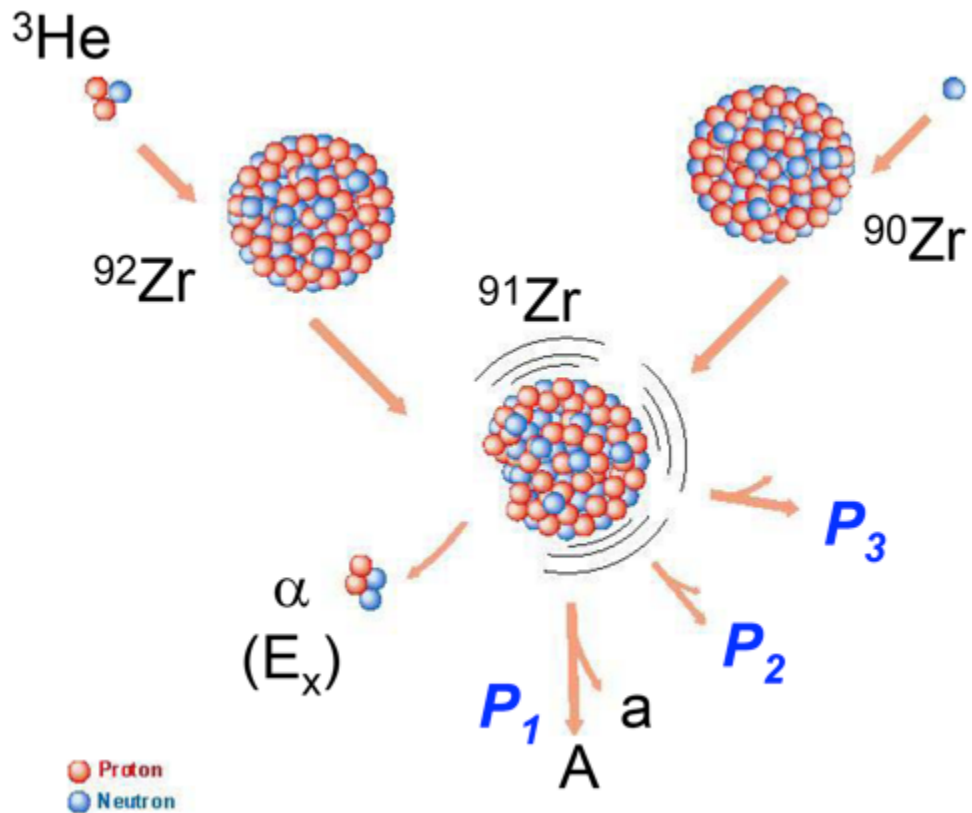
Kessedjian, et al., PLB 692, 297 (2010)

Fission cross sections not sensitive to differences J^π distributions!!!

→ Hauser-Feshbach = Ewing-Weisskopf

→ Surrogate transfer reactions work

BUT, unfortunately, most often it doesn't work this way.



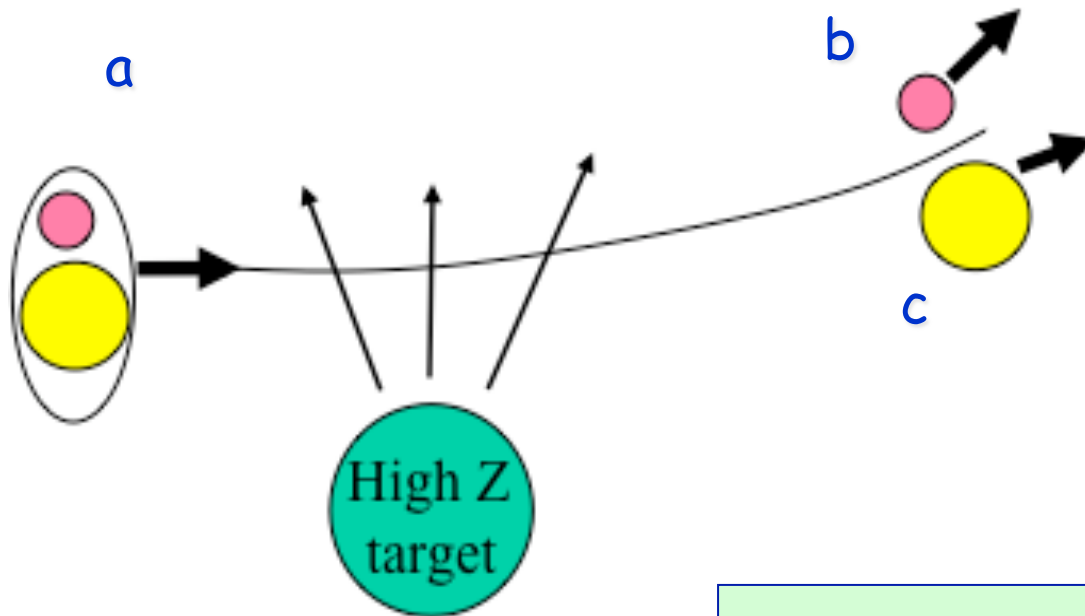


Other Beautiful ideas

Coulomb dissociation method

Baur, Bertulani, Rebel
NPA 458 (1986) 188

$$\frac{d\sigma}{dE_\gamma d\Omega} = \frac{1}{E_\gamma} \sum_l \frac{dn_l(E_\gamma, \Omega)}{dE_\gamma d\Omega} \sigma_{\gamma+a \rightarrow b+c}(E_\gamma)$$



Theory

detailed balance

$$\sigma_{b+c \rightarrow a+\gamma} = \frac{2(2j_a + 1)}{(2j_b + 1)(2j_c + 1)} \frac{k_{bc}^2}{k_\gamma^2} \sigma_{\gamma+a \rightarrow b+c}$$

Applications to radiative capture (n,γ) and (p,γ) reactions in nuclear astrophysics.

DWBA

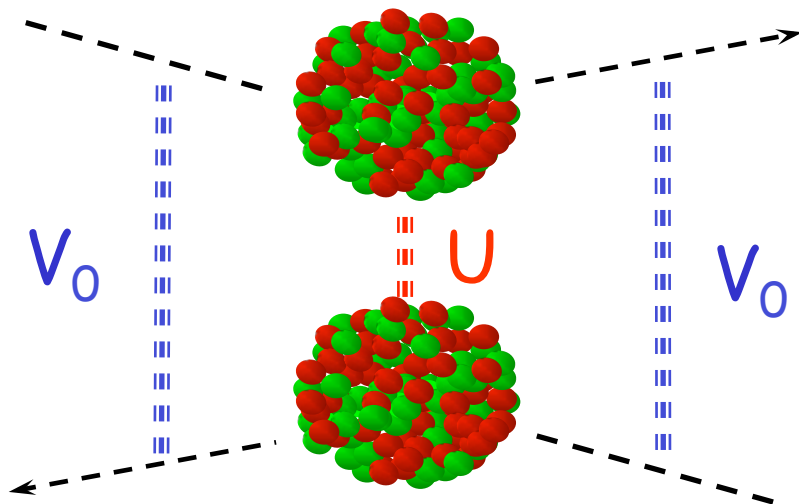
$$f_{inel}(\theta) = -\frac{4\pi^2\mu}{\hbar^2} \int d^3r \chi_{\mathbf{k}'}^{(-)*}(\mathbf{r}) U(\mathbf{r}) \Psi_{\mathbf{k}}^{(+)}(\mathbf{r})$$

$$\Psi^{\pm} \sim \chi^{\pm}$$



$$f_{DWBA}(\mathbf{k}', \mathbf{k}) = -\frac{4\pi^2\mu}{\hbar^2} \langle \chi_{\mathbf{k}'}^{(-)} | U | \chi_{\mathbf{k}}^{(+)} \rangle$$

$$T_{DWBA}(\mathbf{k}', \mathbf{k}) = \langle \chi_{\mathbf{k}'}^{(-)} | U | \chi_{\mathbf{k}}^{(+)} \rangle$$

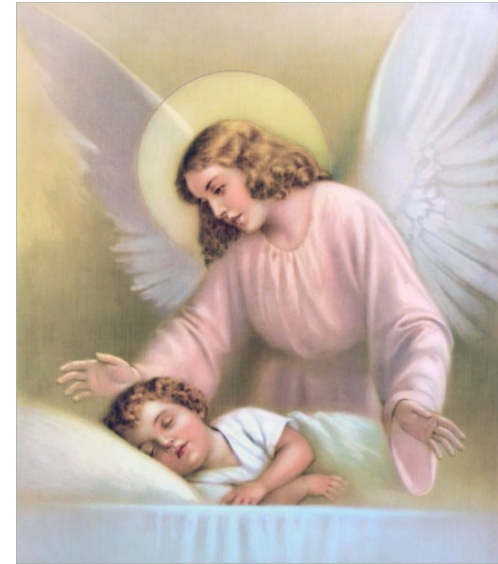


Distorted: all orders in V_0

Born: only first order in U

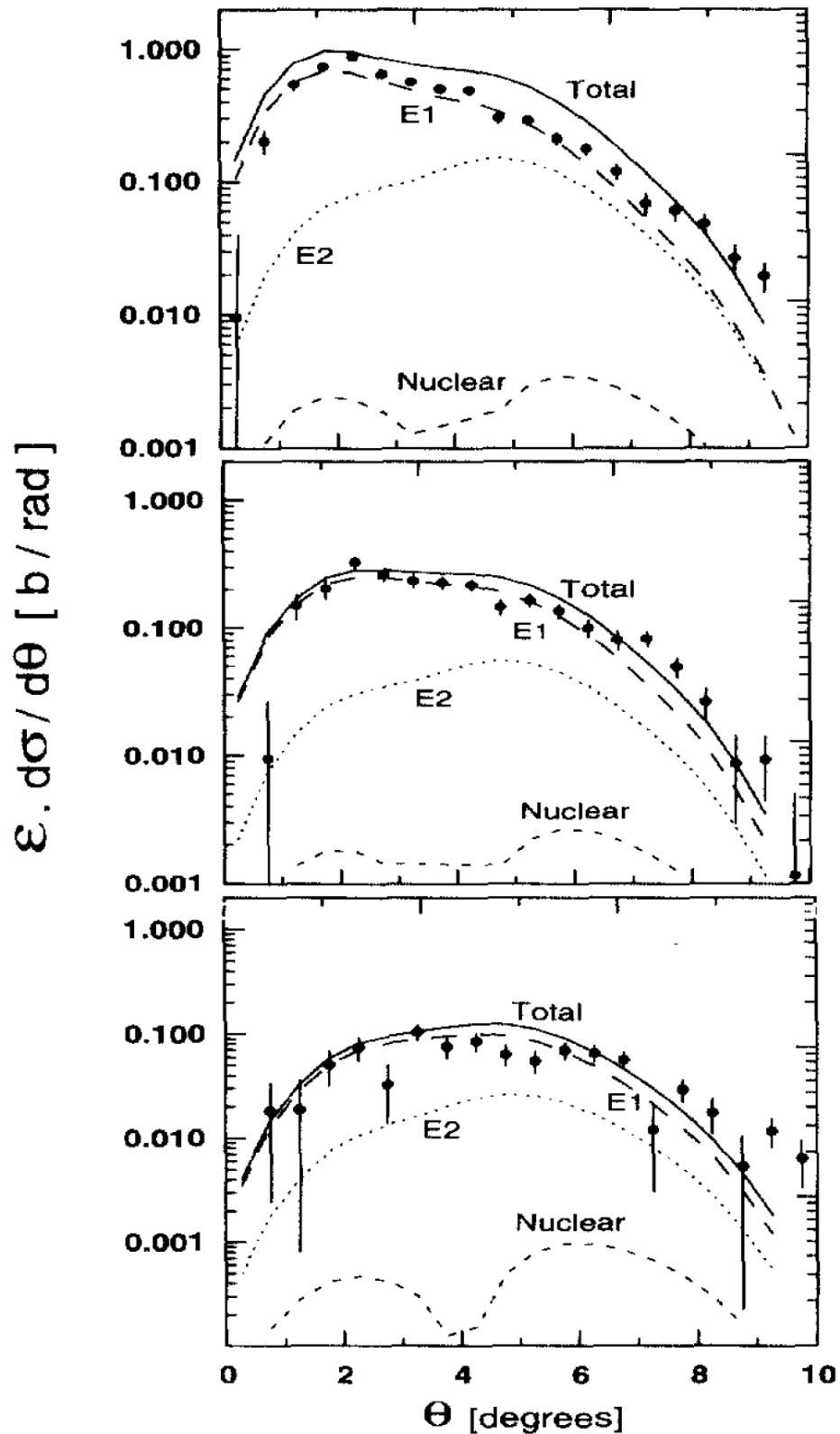
Coulomb excitation + Nuclear excitation

$$f_{inel}^C(\theta) \approx \int d^3r d^3r' \Psi_{\mathbf{k}'}^{(-)*}(\mathbf{r}) \varphi_f(\mathbf{r}') V_C(\mathbf{r}, \mathbf{r}') \Psi_{\mathbf{k}}^{(+)}(\mathbf{r}) \varphi_i(\mathbf{r}')$$



$$f_{inel}^N(\theta) \approx \int d^3r d^3r' \Psi_{\mathbf{k}'}^{(-)*}(\mathbf{r}) \varphi_f(\mathbf{r}') U_N(\mathbf{r}, \mathbf{r}') \Psi_{\mathbf{k}}^{(+)}(\mathbf{r}) \varphi_i(\mathbf{r}')$$

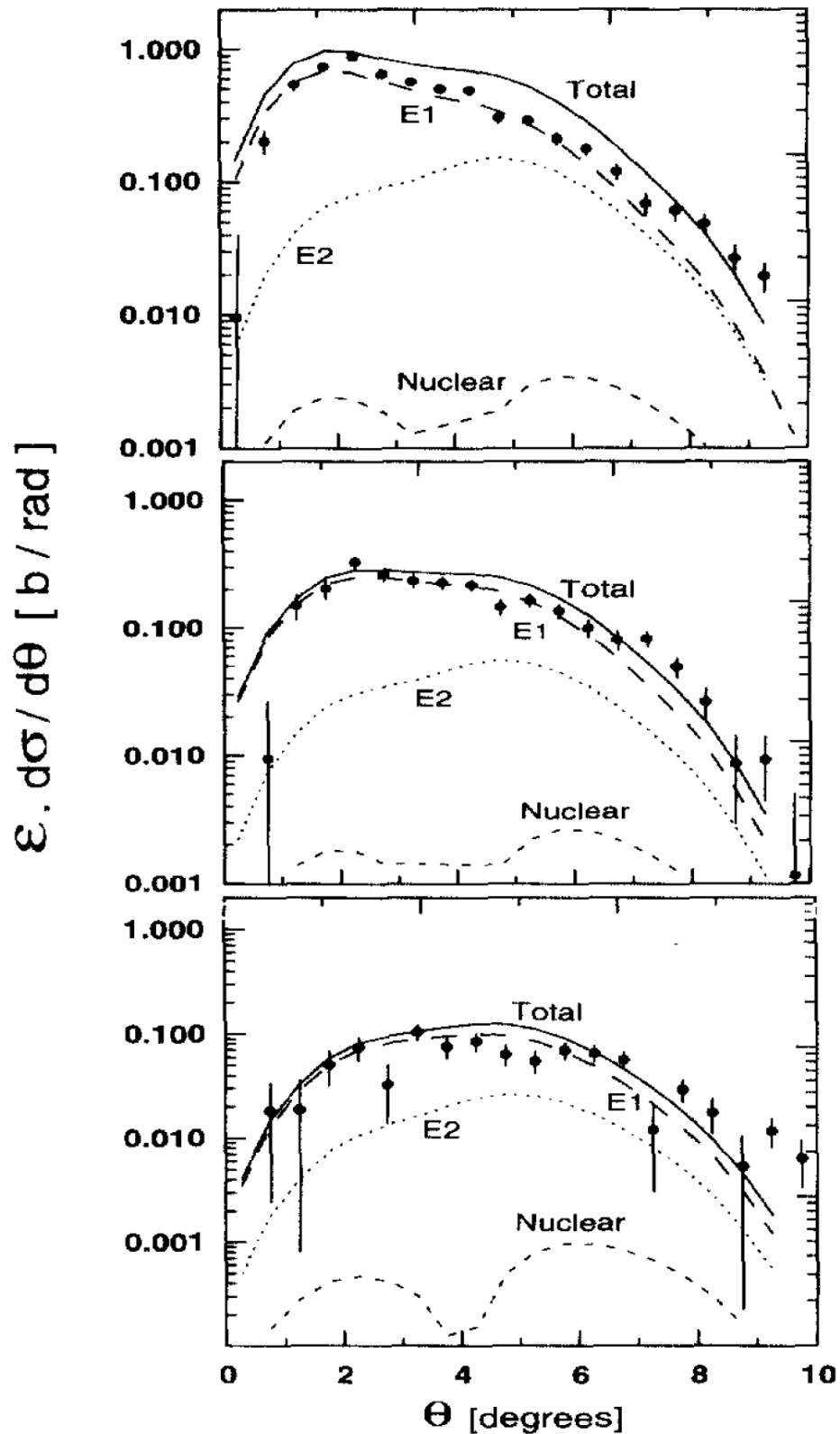
$$\frac{d\sigma}{d\Omega} = \left| f_{inel}^N(\theta) + f_{inel}^C(\theta) \right|^2$$



Relevant for ${}^7\text{Be}(p,\gamma){}^8\text{B}$ (Sun)

Data: Kikuchi et al, PLB 391, 261 (1997)

Calc: Bertulani, Gai, NPA 636, 227 (1998)



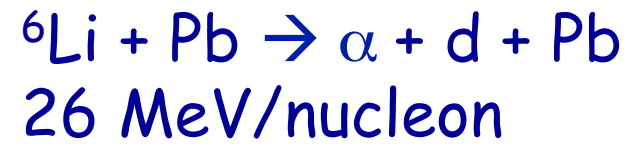
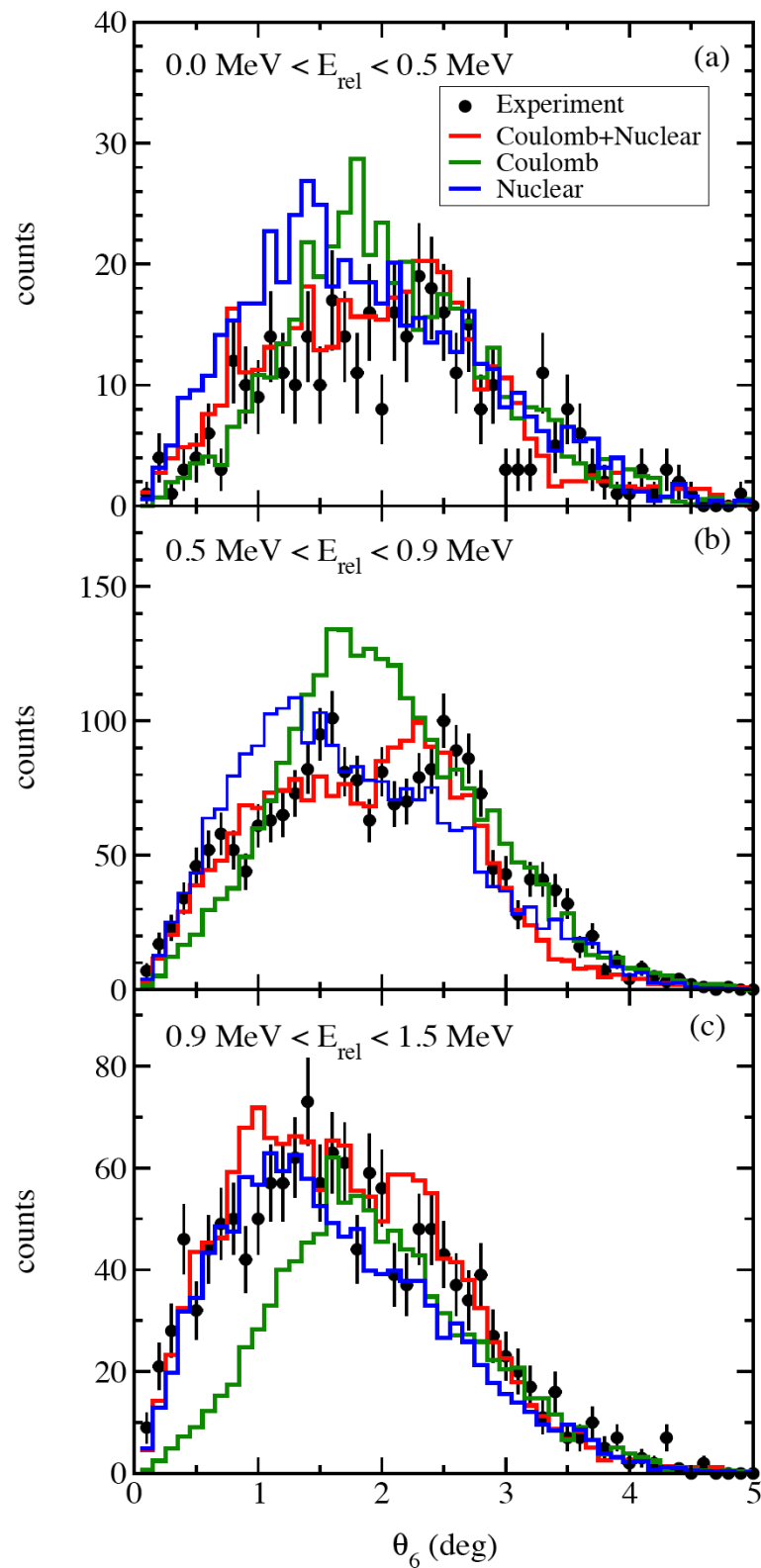
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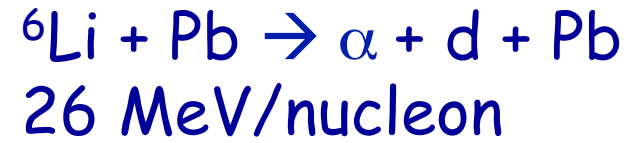
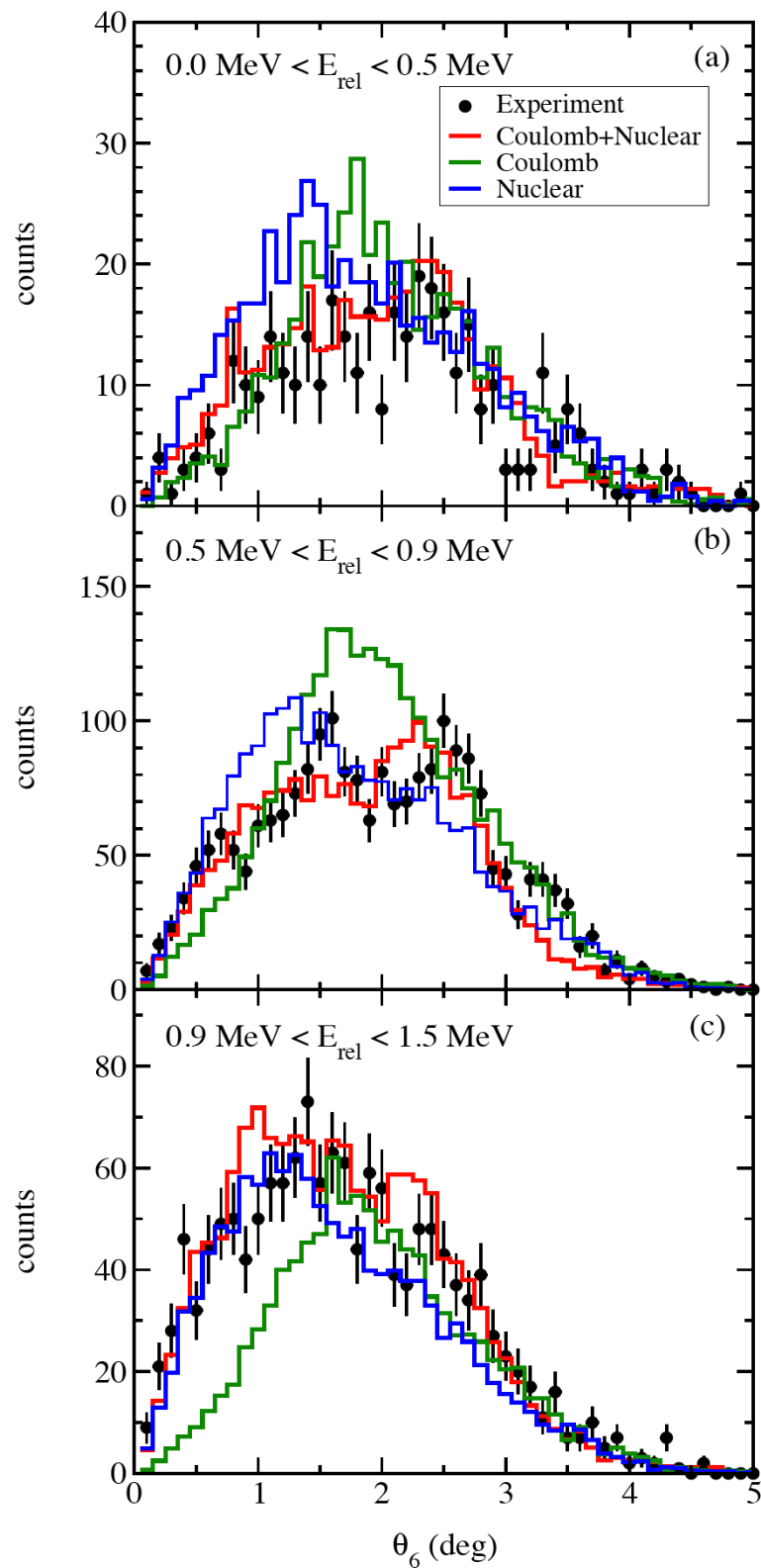


Good wins!



Relevant for BBN

Data: Hammache et al., PRC 82 (2010) 065803
 Calc: Stefan Typel



Relevant for BBN

Data: Hammache et al., PRC 82 (2010) 065803
Calc: Stefan Typel

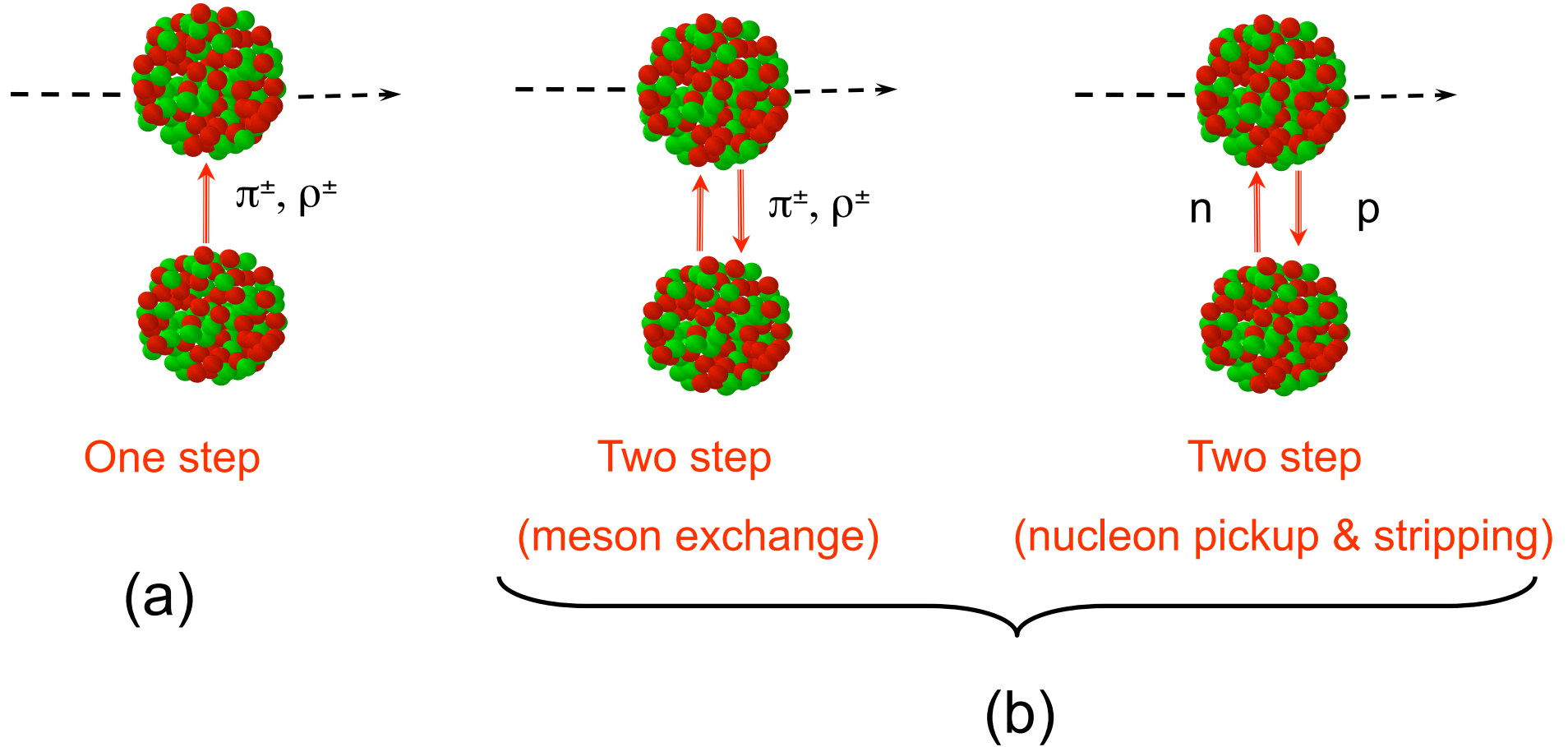


Evil wins!

More ways



Charge-exchange reactions (e.g., Fujita, RCNP)



$$(a) \quad T_{DWBA}(\mathbf{k}', \mathbf{k}) = \langle \chi_{\mathbf{k}'}^{(-)} | U | \chi_{\mathbf{k}}^{(+)} \rangle$$

$$(b) \quad T_{DWA}(\mathbf{k}', \mathbf{k}) = \sum_{\gamma=0} C_{\gamma} \langle \chi_{\mathbf{k}'}^{(-)} | U (G^{(+)} U)^{\gamma} | \chi_{\mathbf{k}}^{(+)} \rangle$$

Effective interaction V_{NN} (phenomenological)

$$V_{NN}(\mathbf{r}) = V^C(r) + V_{\sigma}^C(r)(\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2) + \left[V_{\tau}^C(r) + V_{\sigma\tau}^C(r)(\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2) \right] (\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2) \\ + \left[V^T(r) + V_{\tau}^T(r)(\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2) \right] S_{12}(\hat{\mathbf{r}}) + V^{LS}(r) \mathbf{l} \cdot (\boldsymbol{\sigma}_1 + \boldsymbol{\sigma}_2)$$

Antisimetrization: $V_{NN}(\mathbf{r}) = \left[1 - (-)^l P_x \right] V_{12}(\mathbf{r})$

$$P_x : \mathbf{r} \rightarrow -\mathbf{r}$$

$V^{LS}(r) \mathbf{l} \cdot (\boldsymbol{\sigma}_1 + \boldsymbol{\sigma}_2)$ small and usually neglected

Notation: $V^C(r) = V_{00}^0(r), \quad V_{\sigma}^C(r) = V_{10}^0(r), \quad V_{\tau}^C(r) = V_{01}^0(r)$
 $V_{\sigma\tau}^C(r) = V_{11}^0(r), \quad V^T(r) = V_{10}^2(r), \quad V_{\tau}^T(r) = V_{01}^2(r)$

$$V_{12}(\mathbf{r}) = \sum_{\substack{K=0,2 \\ ST}} V_{ST}^K(r) C_S^K Y_K(\hat{\mathbf{r}}) [\boldsymbol{\sigma}_1 \otimes \boldsymbol{\sigma}_2]^K [\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2]^T$$

$K = 0$: central force

$$\boldsymbol{\sigma}^{S=0} = 1, \quad \boldsymbol{\sigma}^{S=1} = \boldsymbol{\sigma}$$

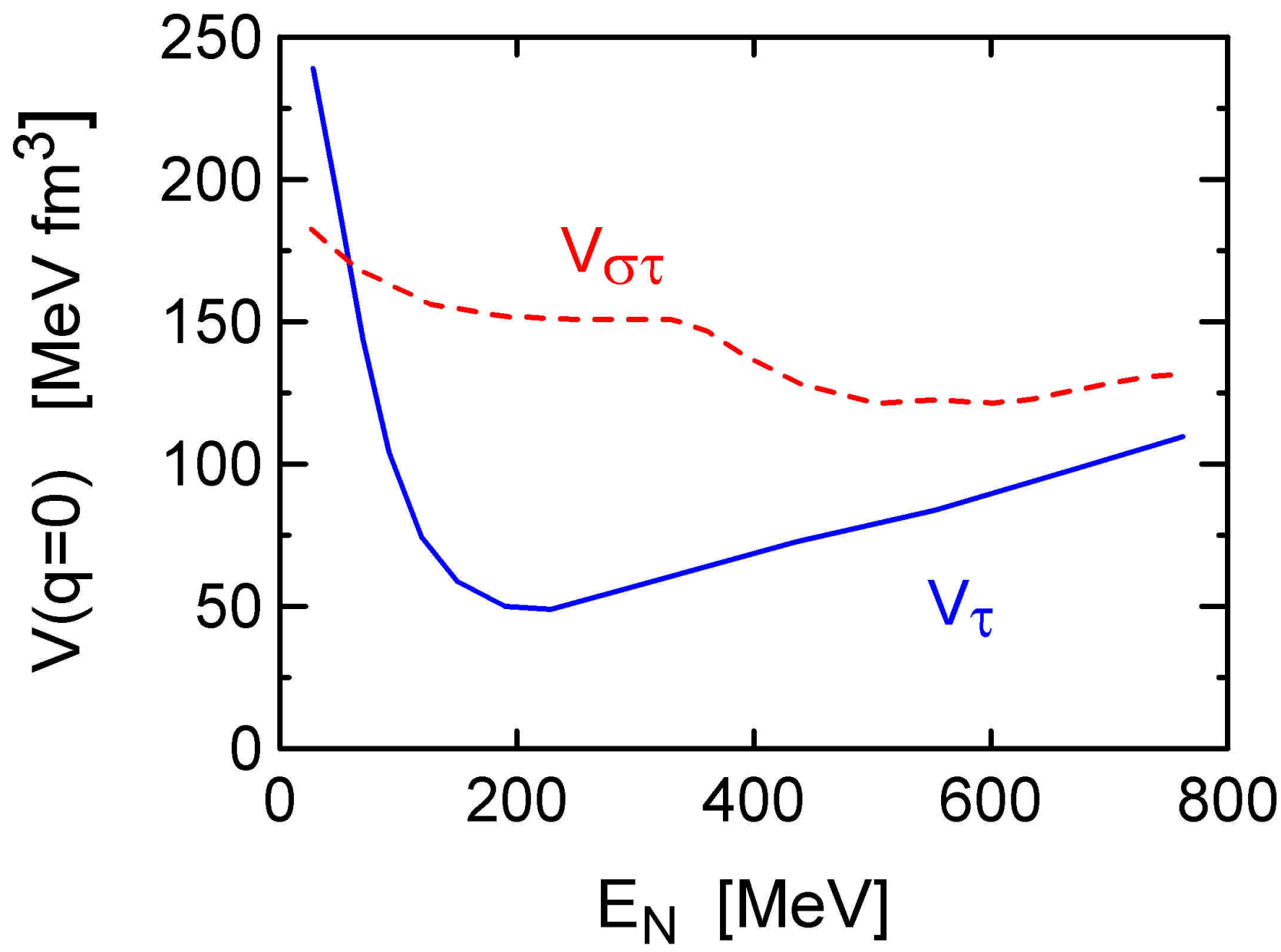
$$C_0^0 = \sqrt{4\pi}, \quad C_1^0 = -\sqrt{12\pi}$$

$K = 2$: tensor force

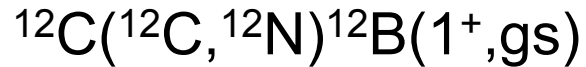
$$\boldsymbol{\tau}^{T=0} = 1, \quad \boldsymbol{\tau}^{T=1} = \boldsymbol{\tau}$$

$$C_0^2 = 0, \quad C_1^2 = \sqrt{25\pi/5}$$

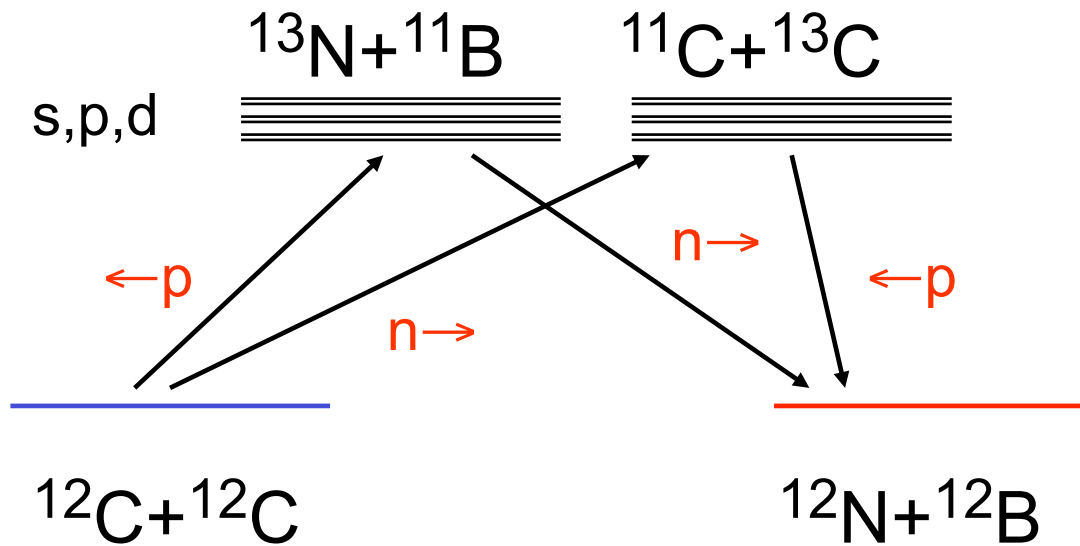
Love, Franey, NPA 1981, 1985



Lenske, Wolter, Bohlen, PRL 62, 1457 (1989)



Two step (proton pickup & neutron-stripping)

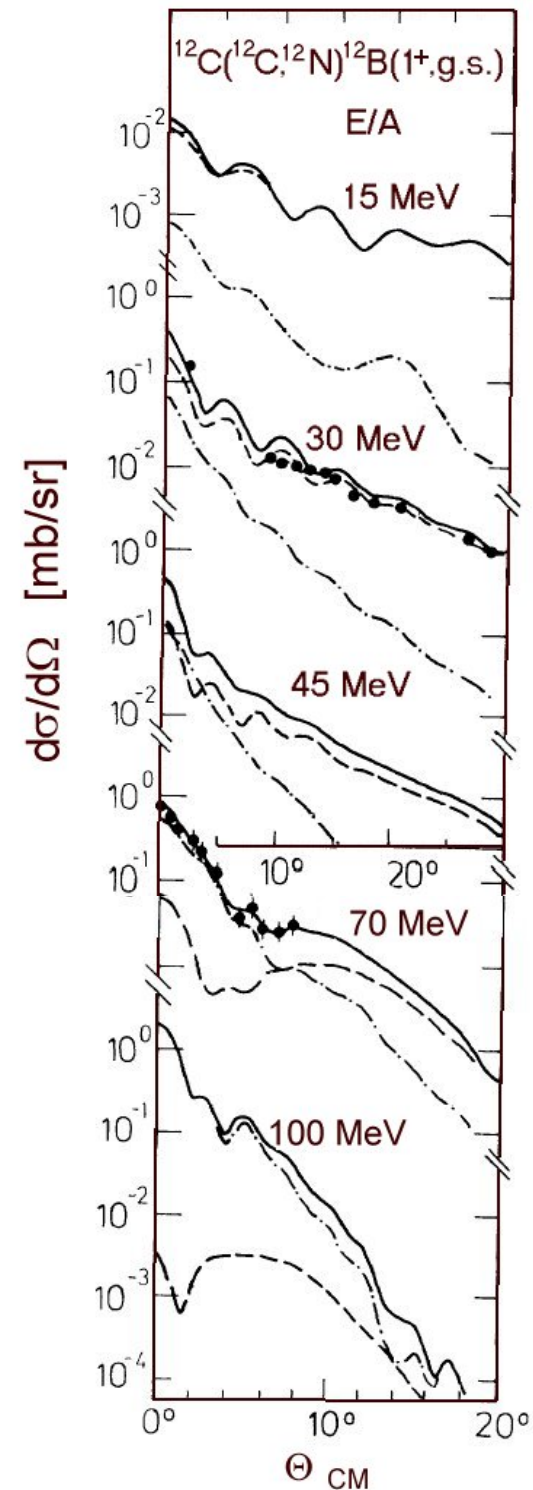


Bertulani, NPA 554, 493 (1993)



Two step (double $\pi+\rho$ exchange)

$$\sigma_{2\text{nd}} \sim 10^{-4} \times \sigma_{1\text{st}}$$



DBWA

$$T_{ch.exch.}(\mathbf{k}', \mathbf{k}) = \int d^3r S(b) \exp[i\mathbf{q} \cdot \mathbf{r}] \langle bB | U(\mathbf{r}) | aA \rangle$$

$$|aA\rangle = |aA; J_a M_a T_a N_a; J_A M_A T_A N_A\rangle$$

eikonal + few pages of algebra

Bertulani, NPA 554, 493 (1993)

$$T_{ch.exch.}(\mathbf{k}', \mathbf{k}) = \sum_{\substack{K=0,2 \\ ST}} \sum_{\substack{LL'JJ' \\ MM'\mu}} C(KS; LL' JJ' MM' \mu) \int db b S(b) J_0(qb) \\ \times \int dp p J_{M'-M-\mu}(pb) \tilde{V}_{ST}^K(p) \tilde{\rho}_{LJST}^{aA}(p) \tilde{\rho}_{L'J'ST}^{bB}(p)$$

$$\tilde{\rho}_{LJST}^{aA}(p) = \int dr r^2 j_L(pr) \left\langle J_a T_a \left\| \sum_i \frac{\delta(r-r_i)}{r_i^2} \mathfrak{S}_M^{LSJ} \tau^T \right\| J_b T_b \right\rangle$$

STRUCTURE INPUT
beautifully factorized

$$\mathfrak{S}_M^{LSJ} = \sum_{\mu M_L} \langle LM_L S\mu | JM \rangle i^L Y_{LM_L}(\hat{\mathbf{r}}) \sigma^{S\mu}$$

$$T_{aA \rightarrow bB}(\mathbf{k}', \mathbf{k}) = \sum_{\dots} \sum_{\dots} \dots \int db b S(b) J_0(qb) \int dp p J_{\dots}(pb) \tilde{\rho}_{\dots}^{aA}(p) \tilde{\rho}_{\dots}^{bB}(p)$$

$$S(b) \sim 1 \implies p \sim q$$

• $S(b) \neq 1$ but largest value of $T_{aA \rightarrow bB}$ occurs when $J_0(qb)$ oscillates in phase with $J_{\dots}(pb)$

$$\implies p \sim q$$

Forward scattering: $q \sim 0$

Bertulani, NPA 554, 493 (1993)

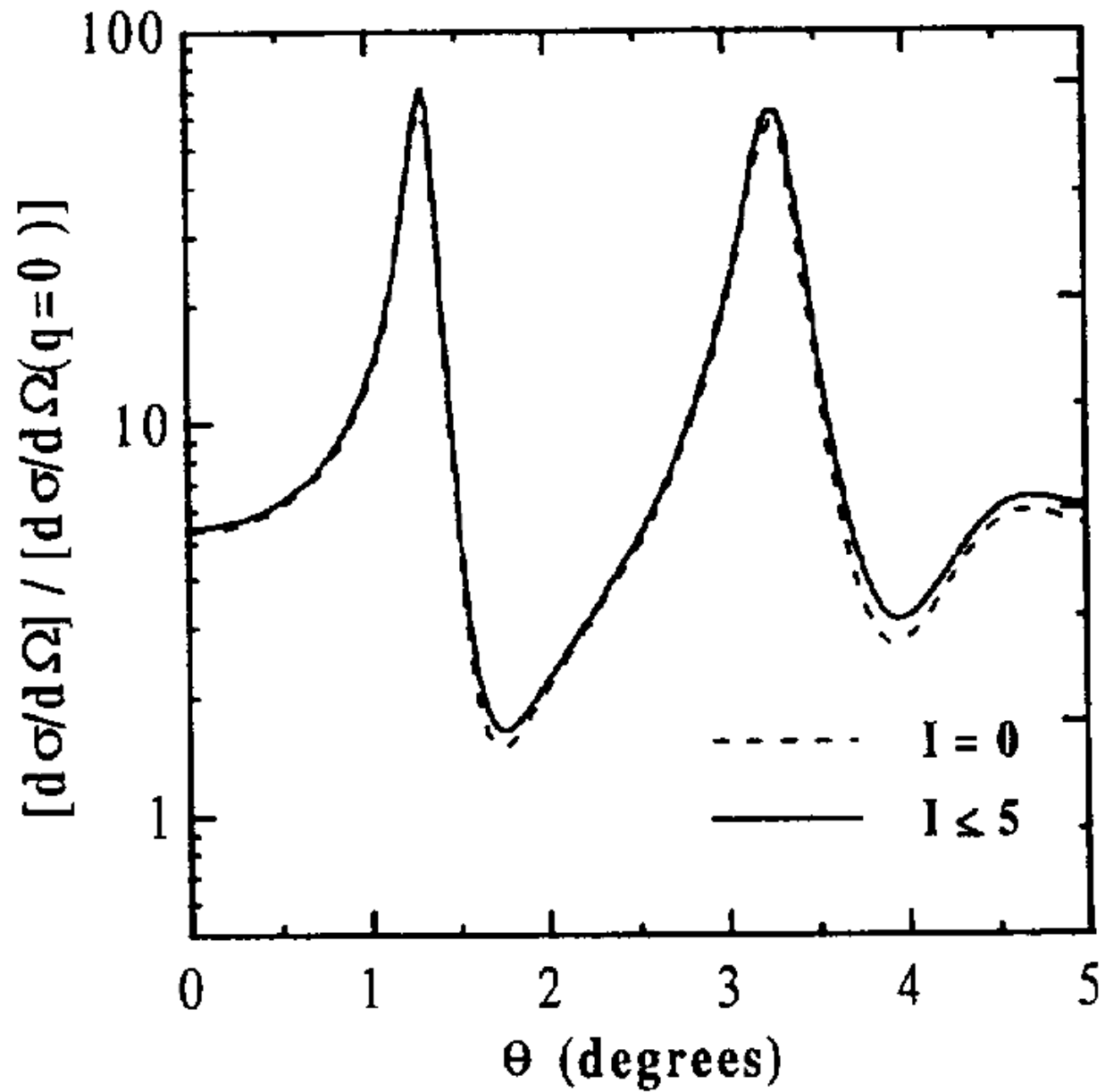
$$f_{aA \rightarrow bB}(\theta \sim 0) = \dots \tilde{\rho}_{\dots}^{aA}(0) \tilde{\rho}_{\dots}^{bB}(0) \times \int dp p V_{ST}^K(p) \times \int db b J_0(qb) e^{iX(b)}$$

$$\tilde{\rho}_{\dots}^{aA}(0) = \dots \langle A \| \sigma^S \tau \| a \rangle$$

$$\frac{d\sigma}{d\Omega}(\theta \sim 0^0) = \dots \left| \langle A \| \sigma^S \tau \| a \rangle \right|^2 \left| \langle B \| \sigma^S \tau \| b \rangle \right|^2$$

\implies • If $\left| \langle A \| \sigma^S \tau \| a \rangle \right|^2$ well known. E.g. $(a, A) = (n, p)$ then

Fermi and Gamow-Teller m.e. READ DIRECTLY from $\frac{d\sigma}{d\Omega}(\theta \sim 0^0)$



$^{13}\text{C}(^{13}\text{N}, ^{13}\text{C})^{13}\text{N}$ 70 MeV/nuc.

Bertulani, Lotti, PLB 402 (1997) 237



Relax

Rolling back. We forgot lots of things!



All the way back to 1911 (Rutherford?)

Rutherford was wrong*

Aguiar, Aleixo, Bertulani, PRC 42, 2180 (1990)

$$L = L^{LO} + L^{NLO} + L^{N^2LO} + \dots$$

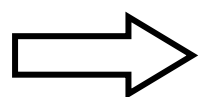
$$L^{LO} = \frac{1}{2} \mu c^2 \left(\frac{v}{c} \right)^2 - \frac{Z_1 Z_2 e^2}{r}$$

$$L^{NLO} = \frac{\mu^4 c^2}{8} \left[\frac{1}{m_1^3} - \frac{1}{m_2^3} \right] \left(\frac{v}{c} \right)^4 - \frac{\mu^2 Z_1 Z_2 e^2}{2m_1 m_2 r} \left[\left(\frac{v}{c} \right)^2 + \left(\frac{\mathbf{v} \cdot \mathbf{r}}{cr} \right)^2 \right]$$

$$L^{N^2LO} = \frac{\mu c^2}{512} \left(\frac{v}{c} \right)^6 + \frac{Z_1 Z_2 e^2}{16r}$$

$$\times \left[\frac{1}{8} \left\{ \left(\frac{v}{c} \right)^4 - 3 \left(\frac{v_r}{c} \right)^4 + 2 \left(\frac{v_r v}{c} \right)^2 \right\} + \frac{Z_1 Z_2 e^2}{\mu c^2 r} \left\{ 3 \left(\frac{v_r}{c} \right)^2 - \left(\frac{v}{c} \right)^2 \right\} + \frac{4 Z_1^2 Z_2^2 e^4}{\mu^2 c^4 r^2} \right]$$

Equations of motion

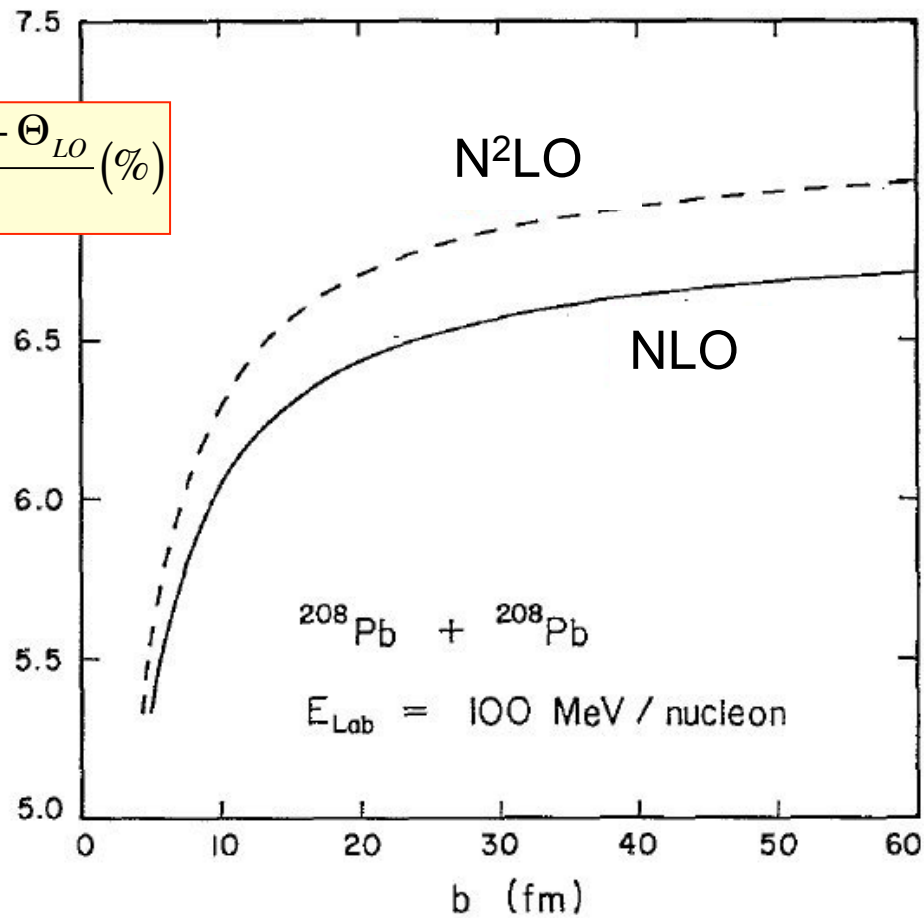


$\mathbf{r}(t)$, $\mathbf{p}(t)$, $\Theta(\mathbf{b})$,

$$\frac{d\sigma}{d\Omega}$$

* If one includes relativistic corrections

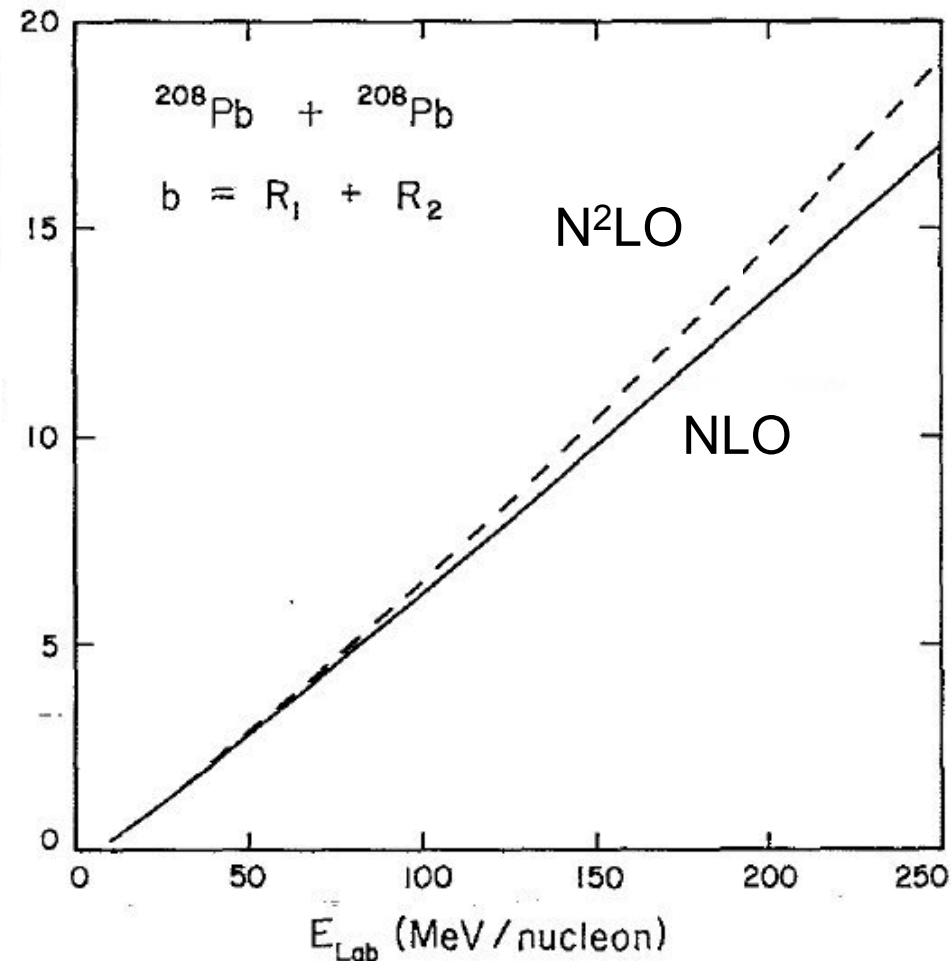
$$\frac{\Theta(E,b) - \Theta_{LO}}{\Theta_{LO}} (\%)$$

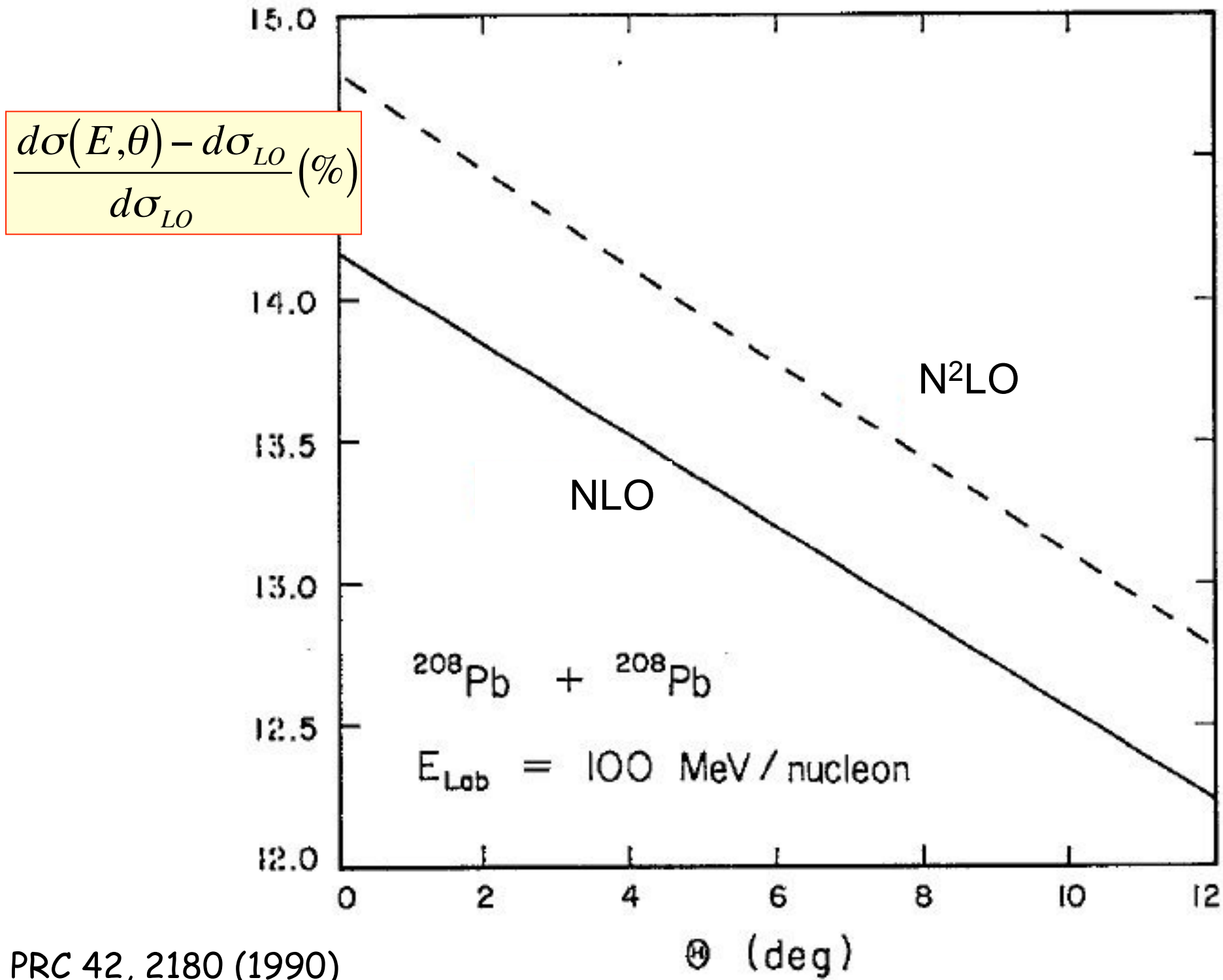


Deviations from Rutherford

important for elastic scattering:
experimental data often reported
as

$$\frac{d\sigma_{\text{elast}}}{d\sigma_{\text{Ruth}}}$$





So what?

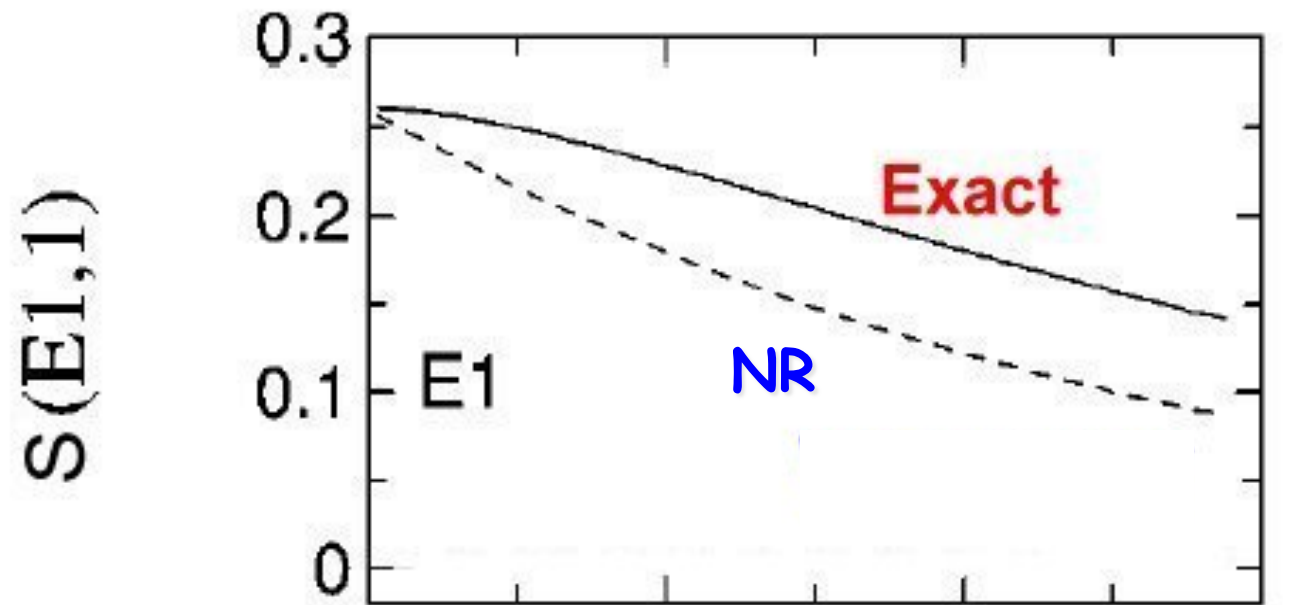


Aleixo, Bertulani, NPA 505, 448 (1989)

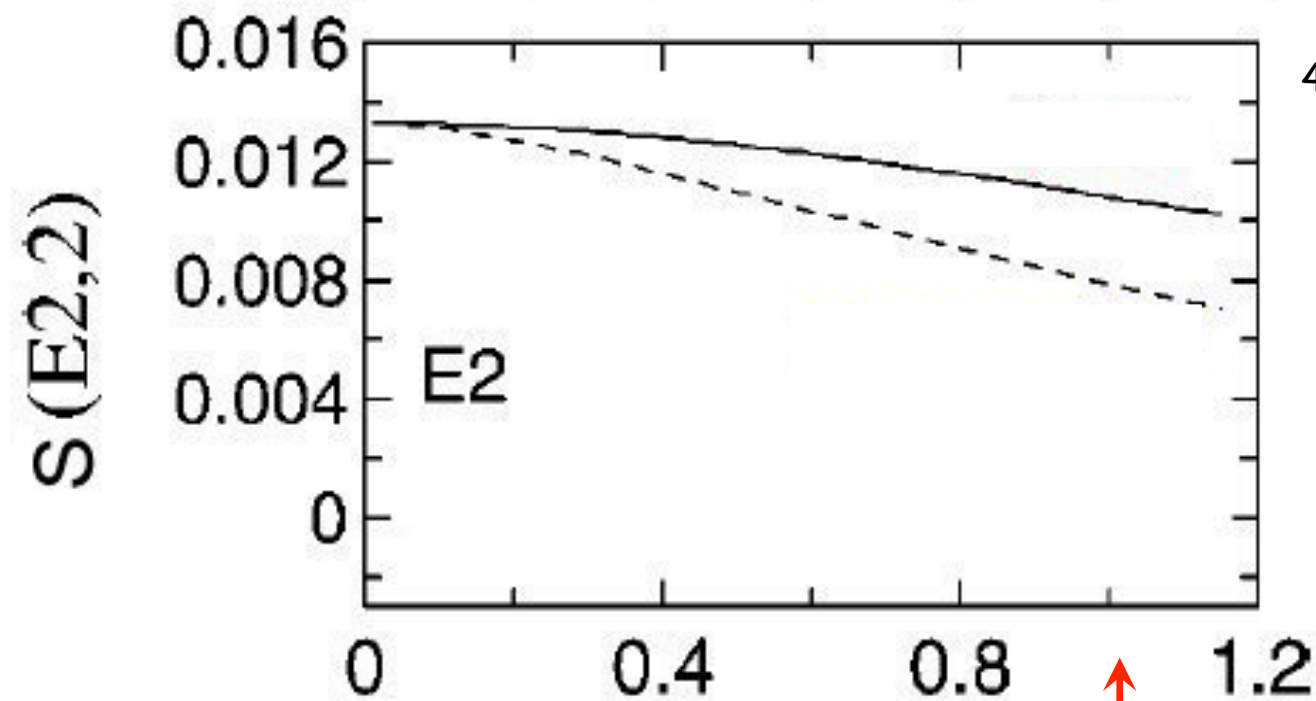
$$\begin{aligned}\frac{d\sigma}{d\Omega} &= \left(\frac{d\sigma}{d\Omega} \right)_{elast} \times P_{exc} \\ &= \sum_{\pi\lambda\mu} |S(\pi\lambda\mu)|^2 |M_{fi}(\pi\lambda, -\mu)|^2\end{aligned}$$

$$M_{fi}(\pi\lambda\mu) = \langle f | \text{EM Operator}(\lambda\mu) | i \rangle$$

$$S(\pi\lambda\mu) = \text{orbital integrals}$$



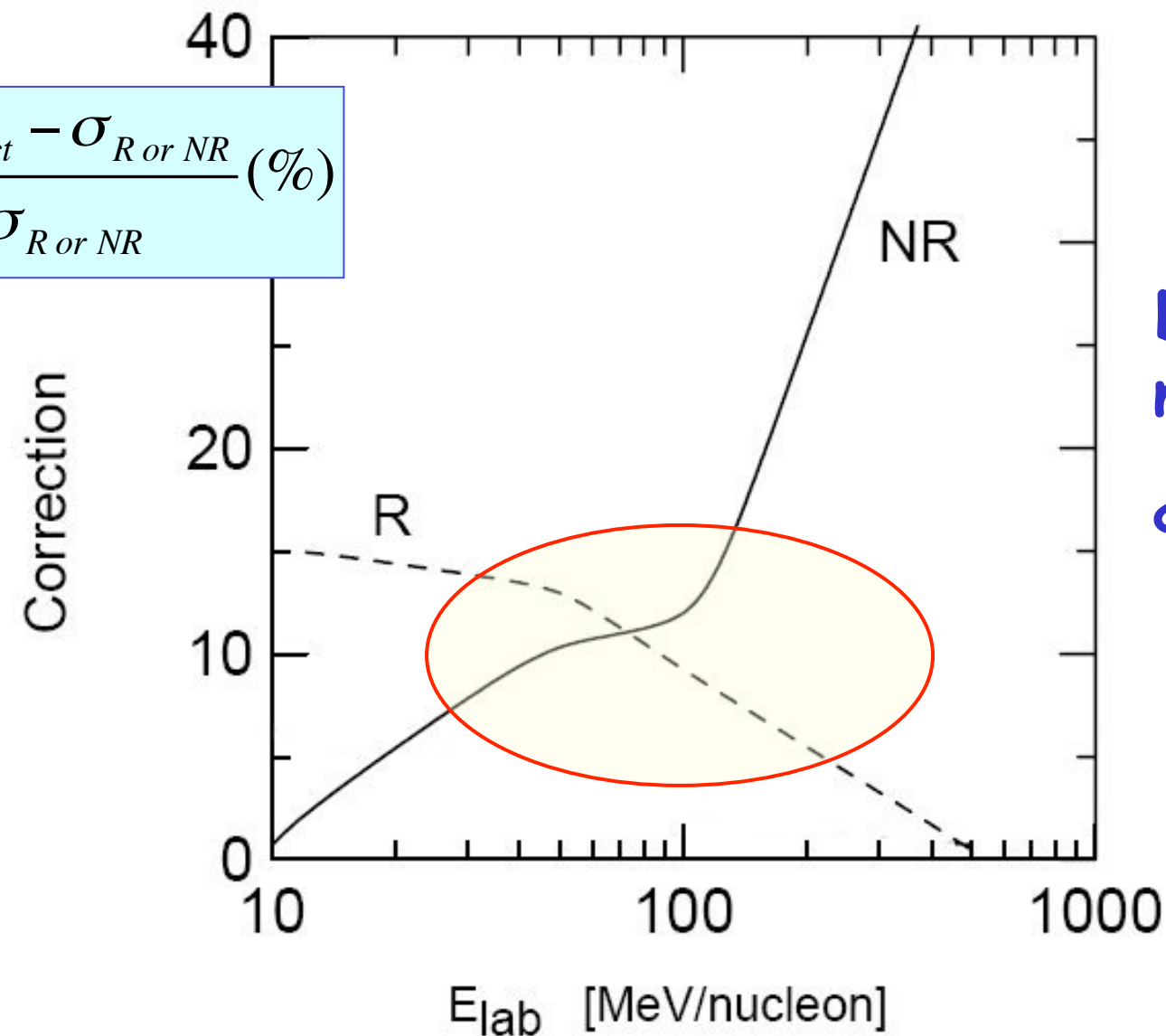
Deviations from
non-relativistic



^{40}S (100 MeV/nucleon) + Au

$$\xi = \frac{E_x b}{\gamma \hbar v}$$

Corrections important
large b 's, large E_x 's



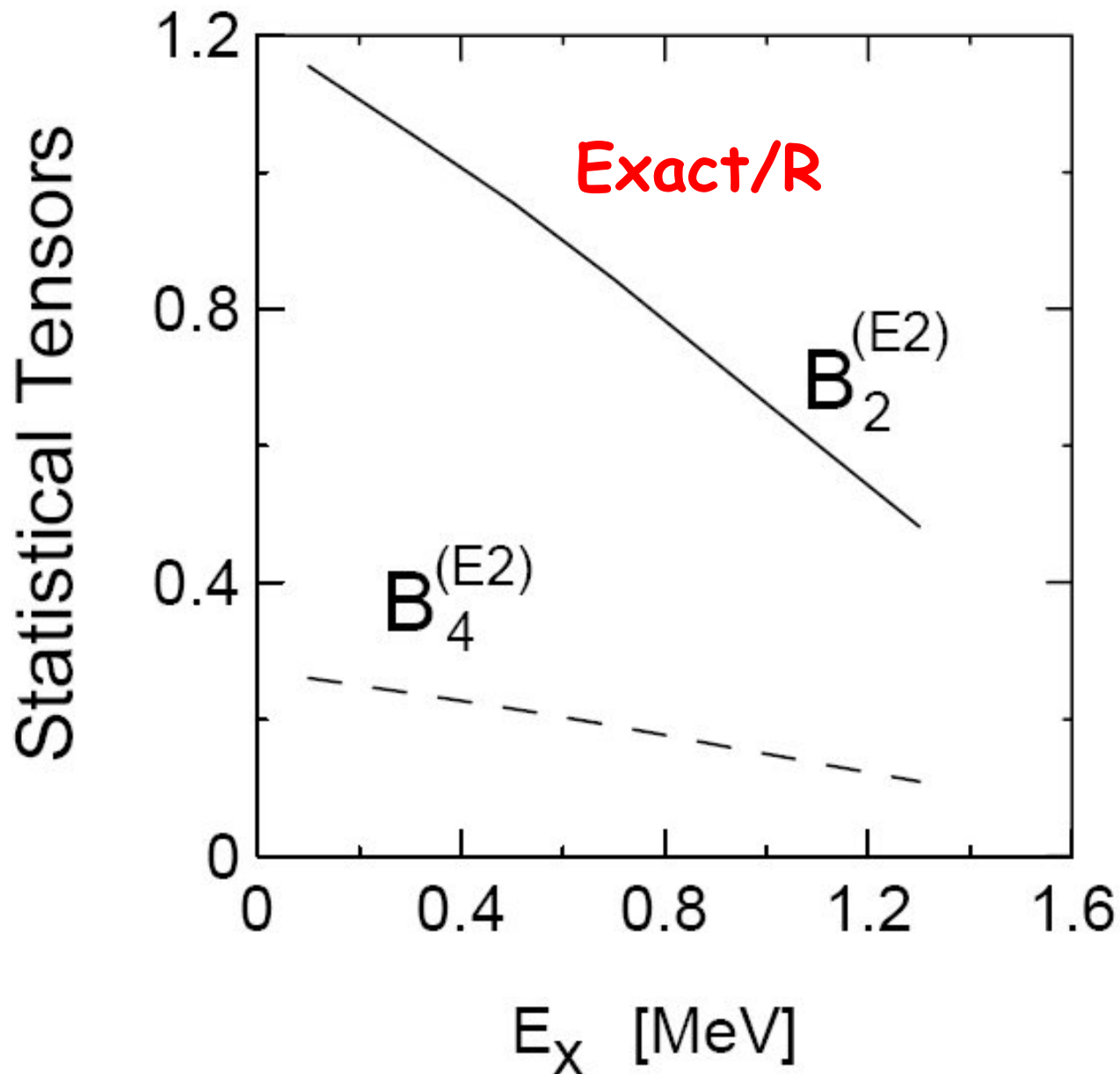
Deviations from
non-relativistic
& from relativistic

^{40}S (100 MeV/nucleon) + Au

$E_x = 0.89$ MeV

De-excitation by γ -ray emission

$$W_\gamma(\theta_\gamma) = 1 + \sum_{\kappa=2,4} B_\kappa Q_\kappa(E_\gamma) P_\kappa(\cos\theta_\gamma)$$



Deviations from
from relativistic
theory

^{38}S (100 MeV/nucleon) + Au

We all know that:

- Relativity obviously important at GANIL, GSI, MSU and RIKEN
(we are talking dynamics)
- ‘Rather’ easy to include for Coulomb interaction
(transformation properties of E/M fields well known)

How about nuclear interaction?

- Transformation properties of nucleus-nucleus potentials not exactly known
- Solution has to be based on QFT (QM + relativity)
- Can we save our DWBA, CC, or CDCC knowledge for something practical?

Way out?



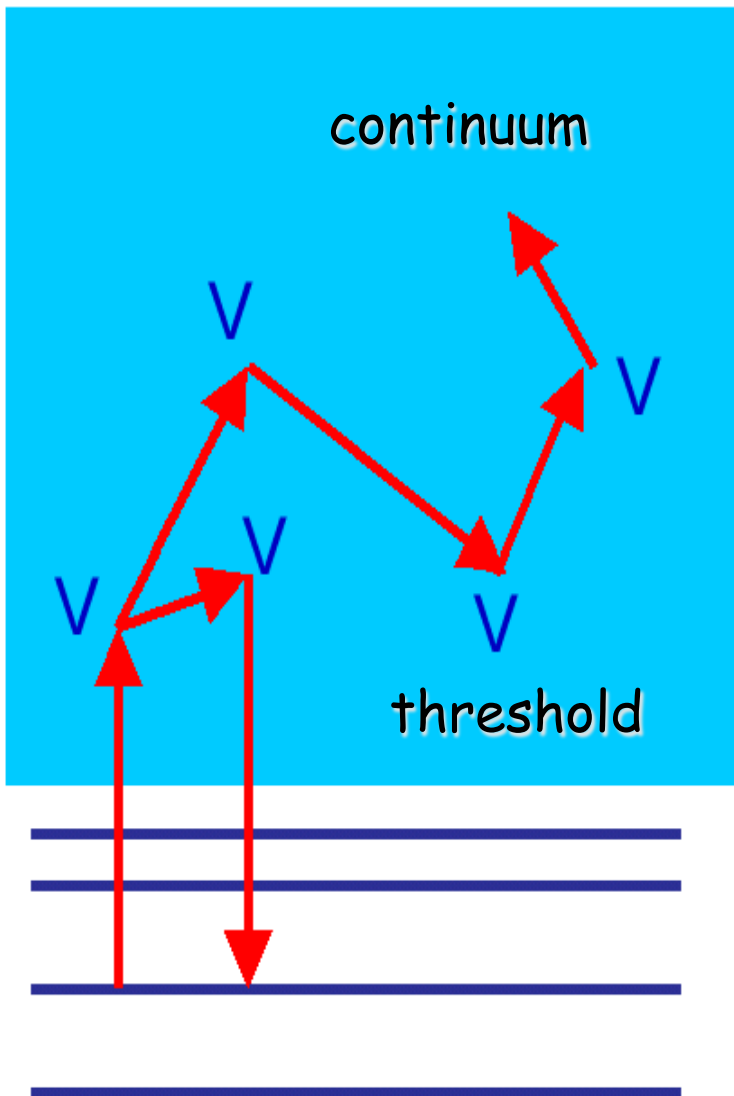
Continuum (CDCC)

$$|\varphi_b\rangle = |E_b, J_b M_b\rangle$$

$$|\varphi_{jJM}^{(c)}\rangle = \int \Gamma_j(E) |E, JM\rangle dE$$

$$\int \Gamma_i(E) \Gamma_j(E) dE = \delta_{ij}$$

continuum discretization



$$V_{\alpha\beta}(\mathbf{R}) = \langle \phi_\alpha(\mathbf{r}) | U(\mathbf{R}, \mathbf{r}) | \phi_\beta(\mathbf{r}) \rangle$$

$$\left[-\frac{\hbar^2}{2\mu} \nabla^2 - E \right] \chi_\alpha(\mathbf{R}) = - \sum_{\beta=0}^N V_{\alpha\beta}(\mathbf{R}) \chi_\beta(\mathbf{R})$$

Coupled-channels

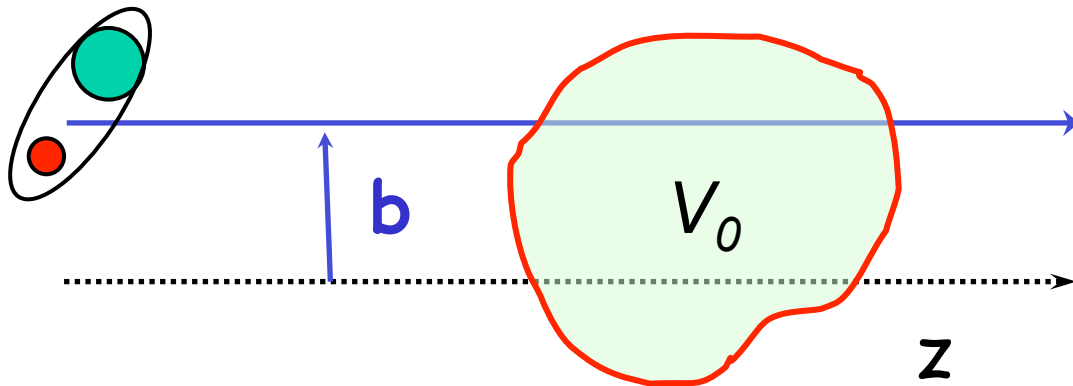
Bertulani, Canto, NPA 539, 163 (1992)
 $^{11}\text{Li} + ^{208}\text{Pb}$ (100 MeV/nucleon)

Relativistic-CDCC

Bertulani, PRL 94, 072701 (2005)

$$[\nabla^2 + k^2 - U] \Psi(\mathbf{R}, \mathbf{r}) = 0$$

$$U = V_0(2E - V_0)$$



$$\Psi(\mathbf{R}, \mathbf{r}) = \sum_{\alpha} S_{\alpha}(\mathbf{b}, z) e^{ik_{\alpha}z} \phi_{\alpha}(\mathbf{r}),$$

$$\mathbf{R} = (\mathbf{b}, z)$$

$$U \approx 2V_0E$$

$$\nabla^2 S \ll ik_z \partial_z S$$



$$iv \partial_z S_{\alpha}(\mathbf{b}, z) = \sum_{\beta} V_{\alpha\beta}(\mathbf{b}, z) S_{\beta}(\mathbf{b}, z) e^{i(k_{\beta} - k_{\alpha})z}$$

$$f_{\alpha}(\mathbf{Q}) = -\frac{ik}{2\pi} \int d\mathbf{b} e^{i\mathbf{Q} \cdot \mathbf{b}} [S_{\alpha}(\mathbf{b}, z = \infty) - \delta_{\alpha,0}]$$

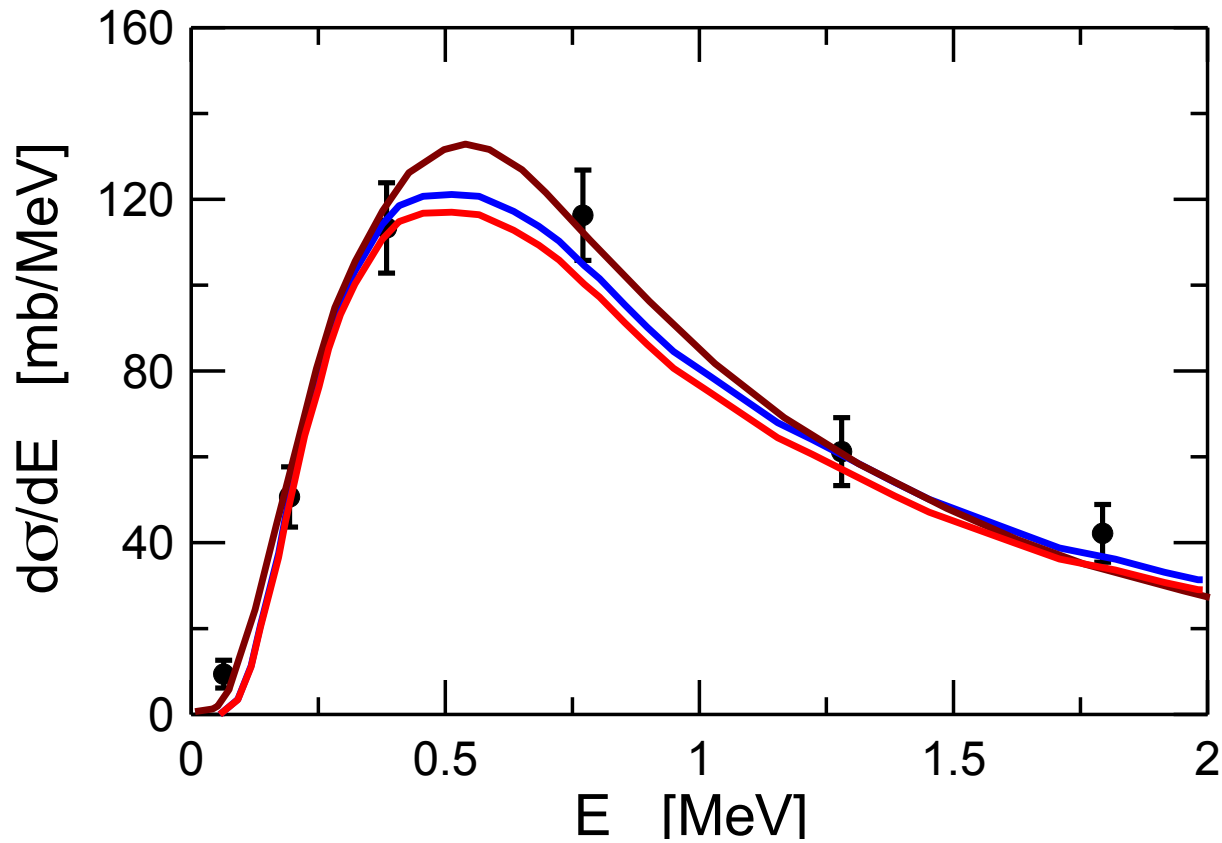
$$\mathbf{Q} = \mathbf{K}'_{\perp} - \mathbf{K}_{\perp} \quad \alpha = j l J M$$

V_0 = time-like
component of 4-vector



Relativistic CDCC
= Lorentz invariant

Pb(^8B , $p^7\text{Be}$) at 83 MeV/nucleon



DATA: Davids et al, 2002

— LO
— all orders
— all orders
— relativistic

*$V_0 = \text{Coulomb} + \text{nuclear}$
with relativistic
corrections*

4-10% effect

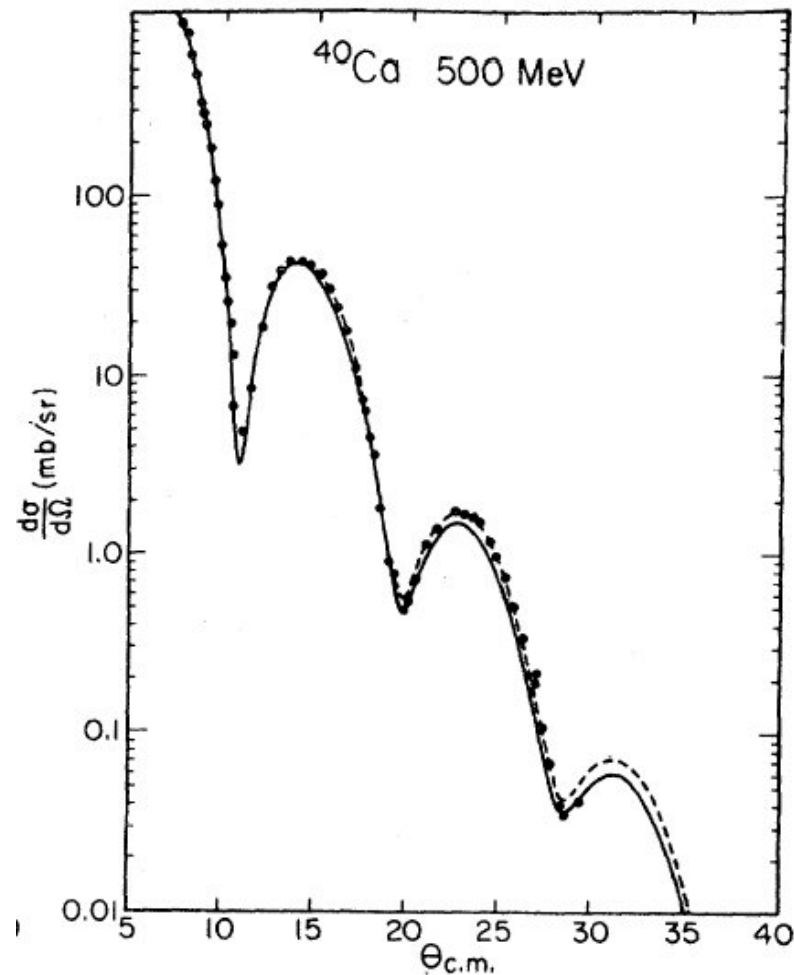
Nuclear



Clue: Proton-nucleus scattering at intermediate energies⁴⁸

- meson exchange, two-nucleon interaction
- mean field approximation, U_0 (ω exchange), U_S (2π exchange)

$$\left[E - V_C - U_0 - \beta (mc^2 + U_S) \right] \Psi = -i\hbar c \alpha \cdot \nabla \Psi$$



non-relativistic reduction

$$\left[-\frac{\hbar^2}{2m} \nabla^2 + U_{cent} + \left(\frac{\hbar}{2mc} \right)^2 \frac{1}{r} \frac{d}{dr} U_{SO} \sigma \cdot \mathbf{L} \right] \phi = E \phi$$

$$U_{cent} = m^* (U_0 + U_S) + \dots$$

$$m^* = 1 - \frac{U_0 - U_S}{2mc^2} + \dots$$

$$U_{SO} = U_0 - U_S + \dots$$

Arnold, Clark, PLB 84, 46 (1979)

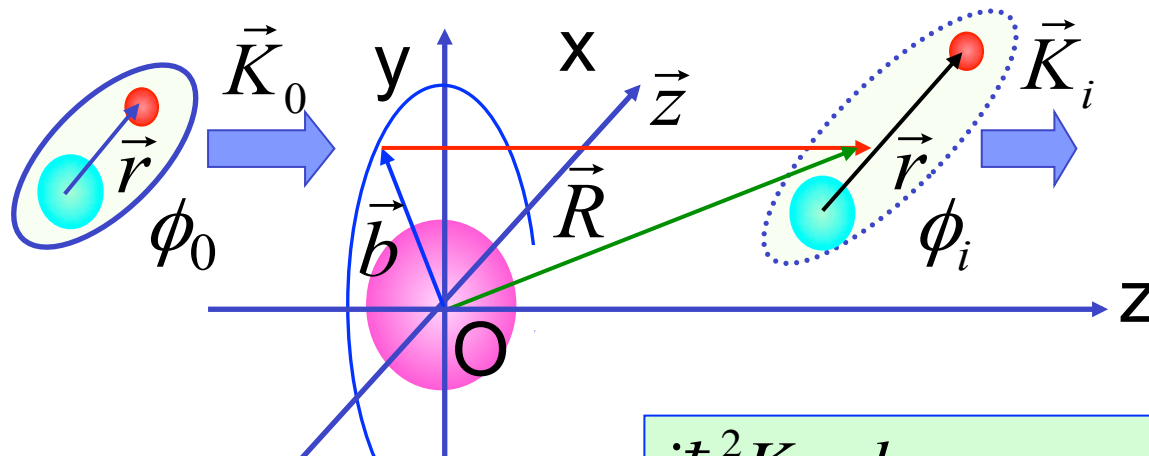
Transition: low to high energies

Eikonal scattering waves $\hat{S}_i(K_i, \vec{R})$

$$\psi^{E-CDCC} = \sum_i \hat{\phi}_i(\vec{r}) \hat{S}_i(b, z) \exp(i\vec{K}_i \cdot \vec{R})$$

$$K_i = \sqrt{2\mu_R(E - \varepsilon_i)} / \hbar,$$

Energy conservation



● Boundary condition

$$\hat{S}_i(b, z) \xrightarrow{z \rightarrow -\infty} \delta_{i,0}$$

$$\Delta \hat{S}_i(b, z) \cong 0 \quad \longrightarrow$$

$$\frac{i\hbar^2 K_i}{\mu_R} \frac{d}{dz} \hat{S}_i^{(b)}(z) = \sum_{i'} F_{ii'}^{(b)}(z) \hat{S}_{i'}^{(b)}(z) e^{i(K_{i'} - K_i)z}$$

Eikonal scattering amplitude transformed into QM form

$$f_{i,0}^E = \sum_L f_L^E \equiv \sum_L \frac{2\pi}{iK_i} \sqrt{\frac{2L+1}{4\pi}} i^m Y_{Lm}(\Omega) [S_{i,0}^{b(L;i)} - \delta_{i,0}]$$

Hybrid scattering amplitude is given by

$$f_{i,0}^H \equiv \sum_{L=0}^{L_C} f_L^Q + \sum_{L=L_C+1}^{L_{\max}} f_L^E$$

Ogata et al,
PRC68, 064609 (2003)

Relativistic CDCC

Form factor of non-rel. E-CDCC

$$F_{c'c}^{(b)}(Z) = \langle \Phi_{c'} | U_{1A} + U_{2A} | \Phi_c \rangle_r e^{-i(m-m')\phi} = \sum_{\lambda} F_{c'c}^{(b);\lambda}(Z)$$

Lorentz transform of form factor and coordinates

$$F_{c'c}^{(b);\lambda}(Z) \rightarrow f_{\lambda, m'-m} \gamma F_{c'c}^{(b)\lambda}(\gamma Z)$$

$$f_{\lambda, m'-m}^{\text{Coul}} = \begin{cases} 1/\gamma & (\lambda=1, m'-m=0) \\ \gamma & (\lambda=2, m'-m=\pm 1) \\ 1 & (\text{otherwise}) \end{cases}$$

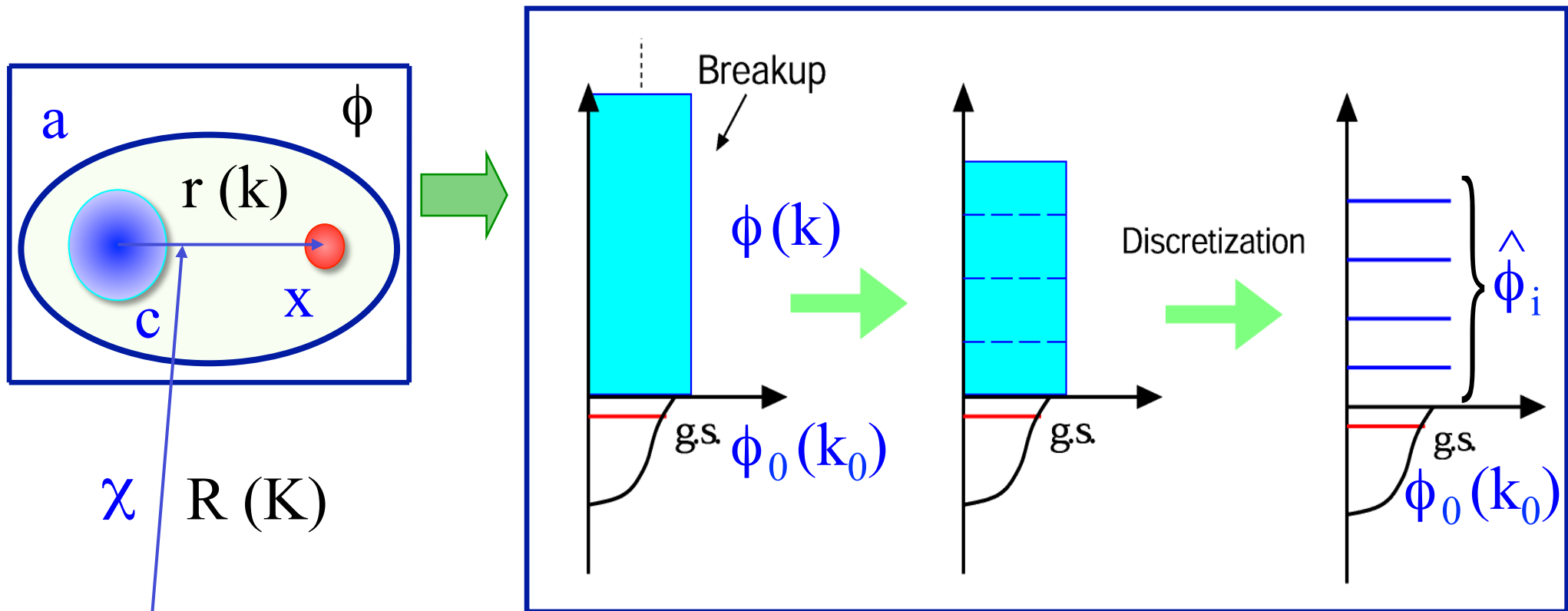
$$f_{\lambda, m'-m}^{\text{nucl}} = 1$$

Assumptions

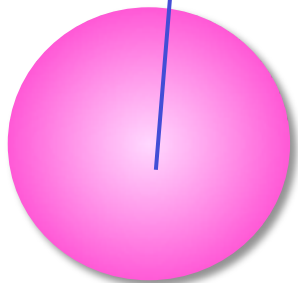
- ✓ Point charges for 1, 2 and A
- ✓ Neglecting far-field ($r_i > R$) contribution
- ✓ Correction to nuclear form factor

Ogata, Bertulani, PTP 121 (2009), 1399
PTP, 123 (2010) 701

Theory movie in next 5 transparencies (enjoy!)©

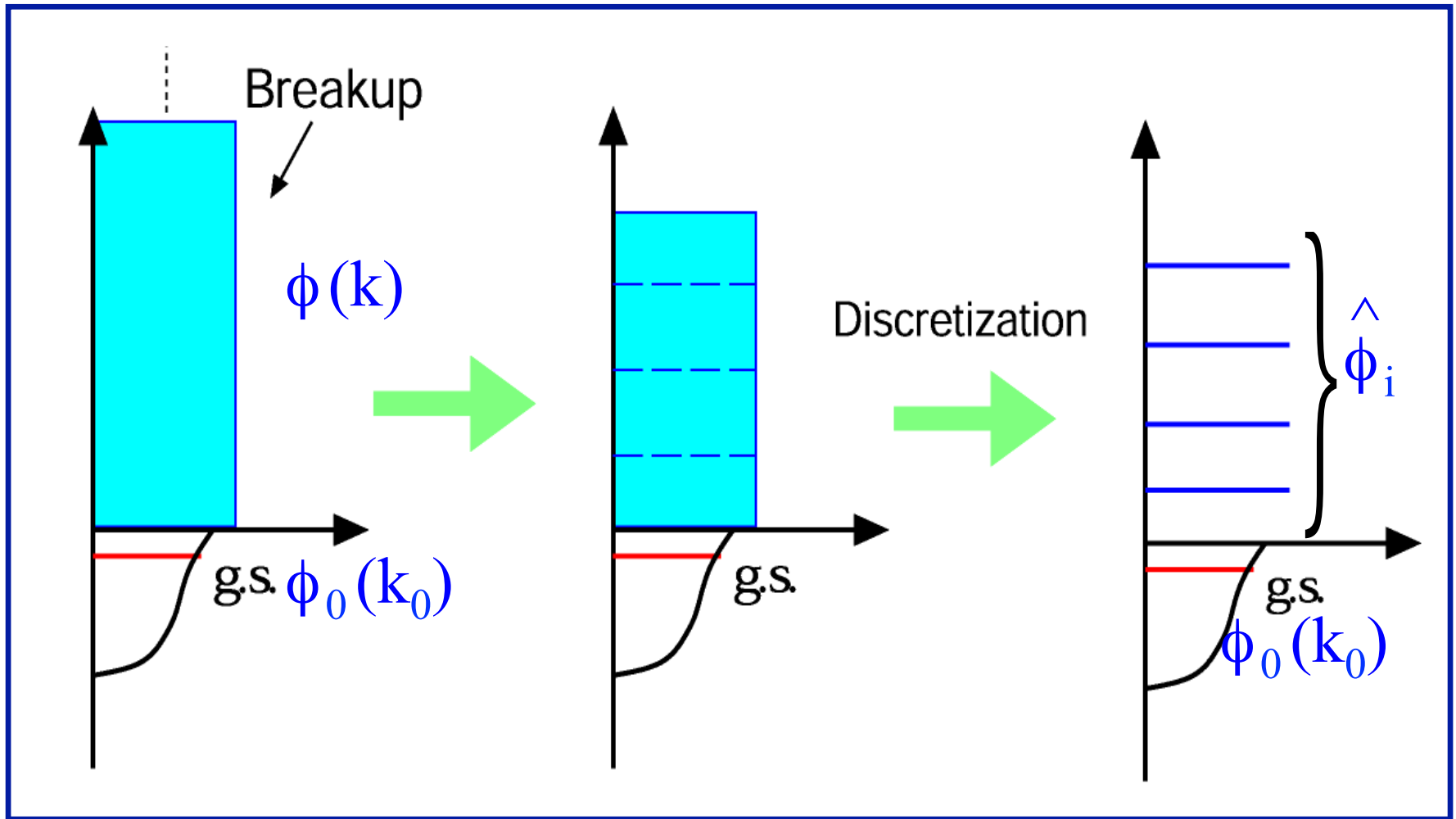


χ $R(K)$

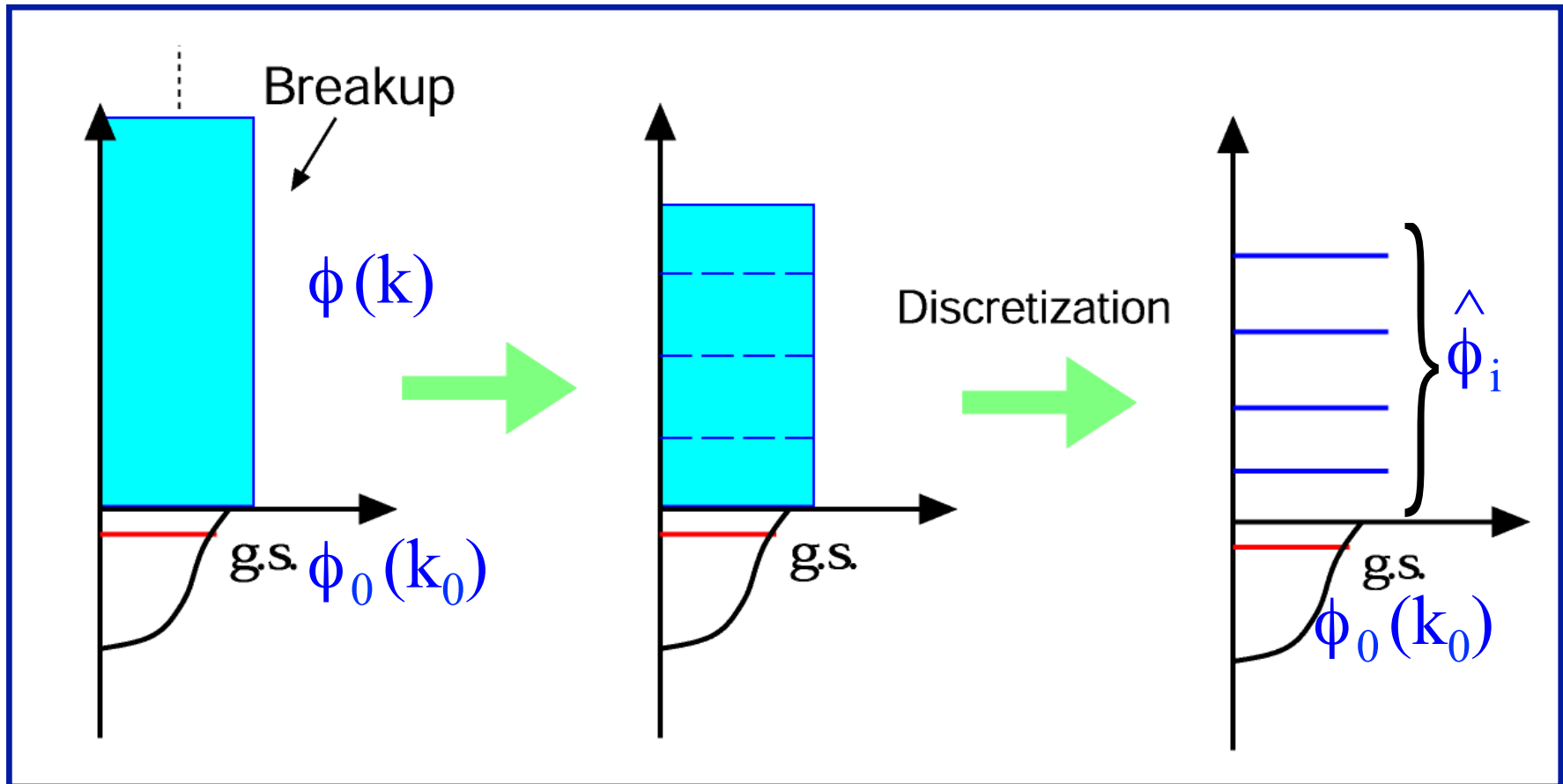


A

$$\psi(\vec{r}, \vec{R}) = \phi_0(k_0, \vec{r})\chi_0(K_0, \vec{R}) + \int_0^\infty \phi(k, \vec{r})\chi(K, \vec{R})dk$$

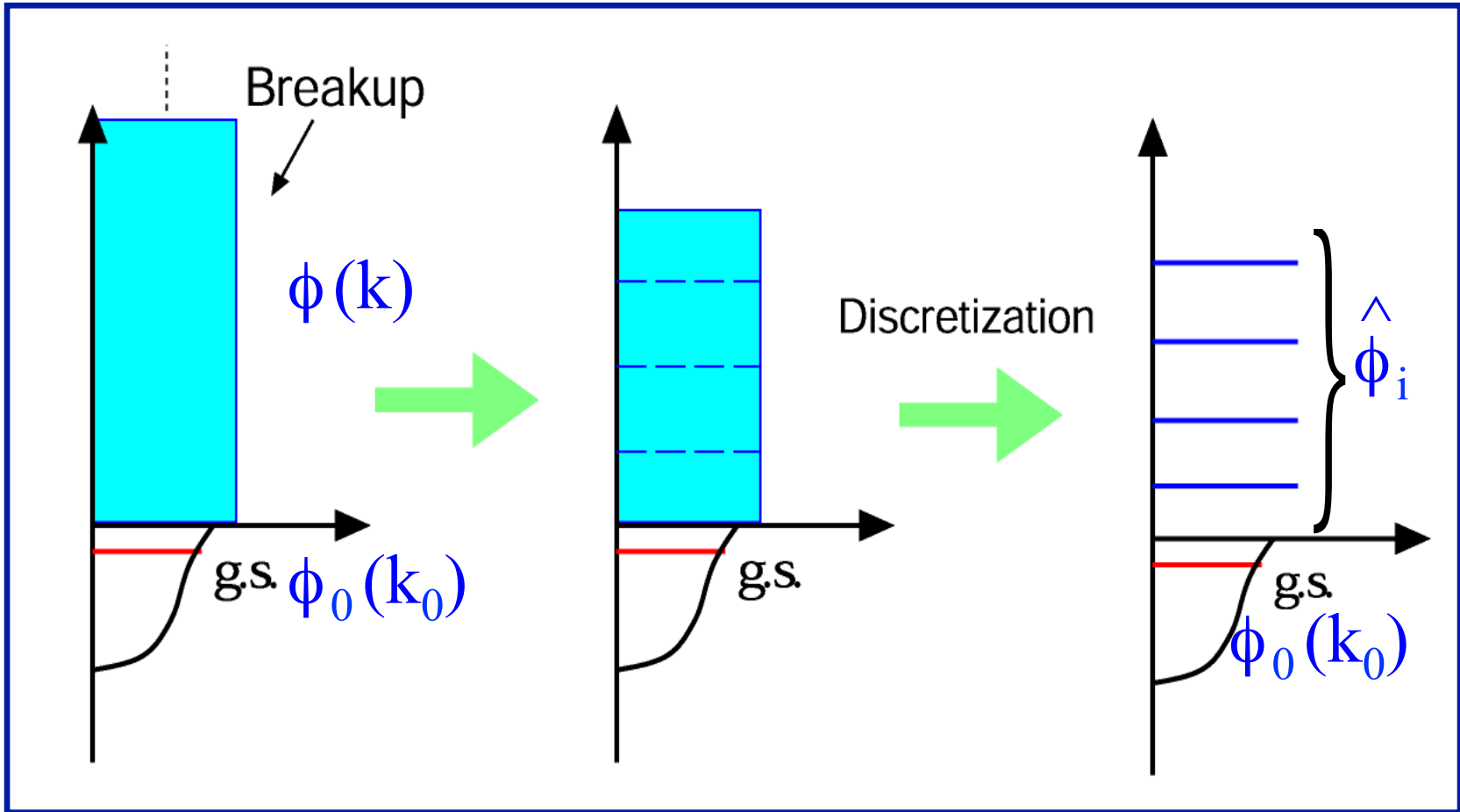


$$\psi(\vec{r}, \vec{R}) = \phi_0(k_0, \vec{r}) \chi_0(K_0, \vec{R}) + \int_0^\infty \phi(k, \vec{r}) \chi(K, \vec{R}) dk$$



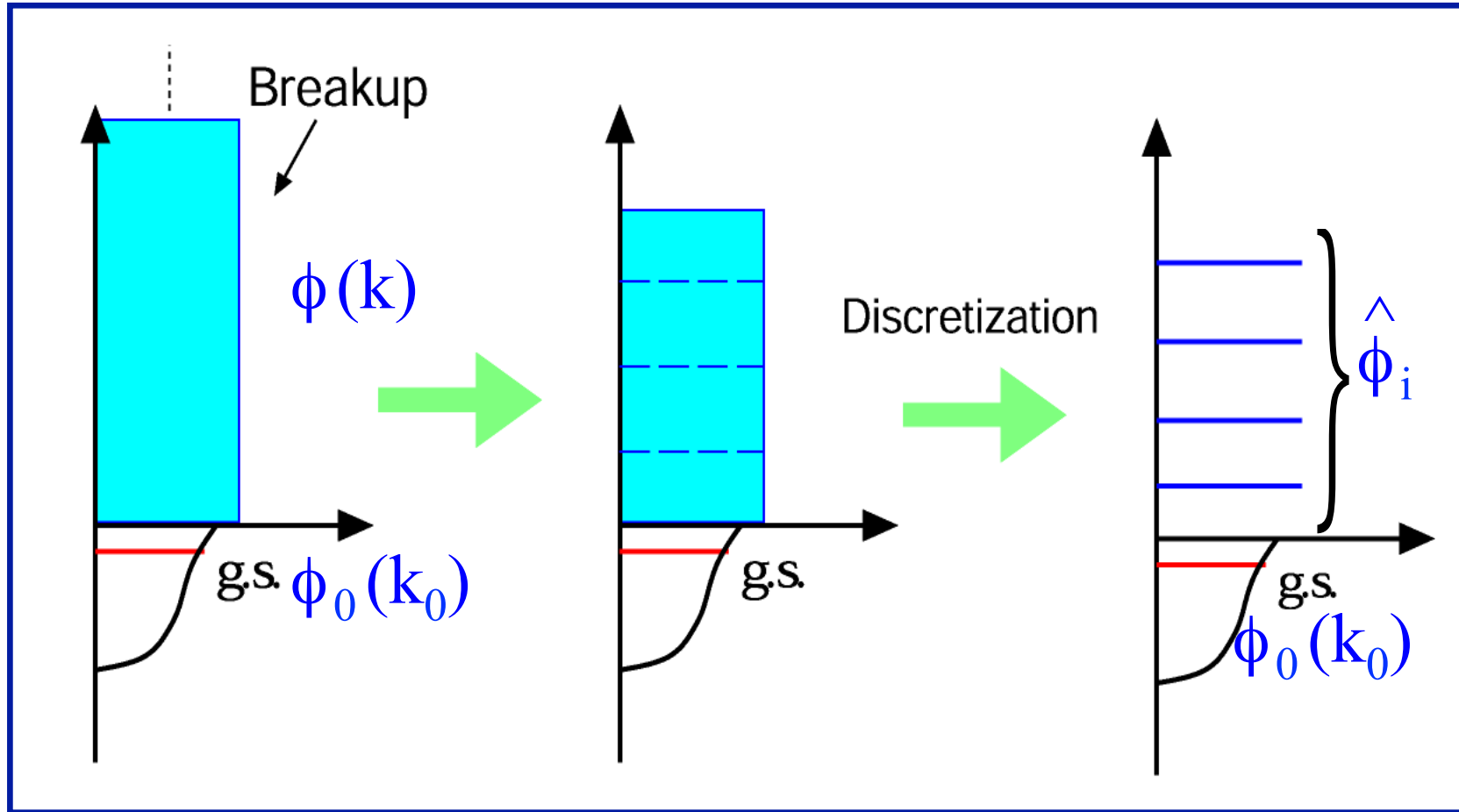
$$\psi(\vec{r}, \vec{R}) = \phi_0(k_0, \vec{r})\chi_0(K_0, \vec{R}) + \int_0^\infty \phi(k, \vec{r})\chi(K, \vec{R})dk$$

Truncation and Discretization



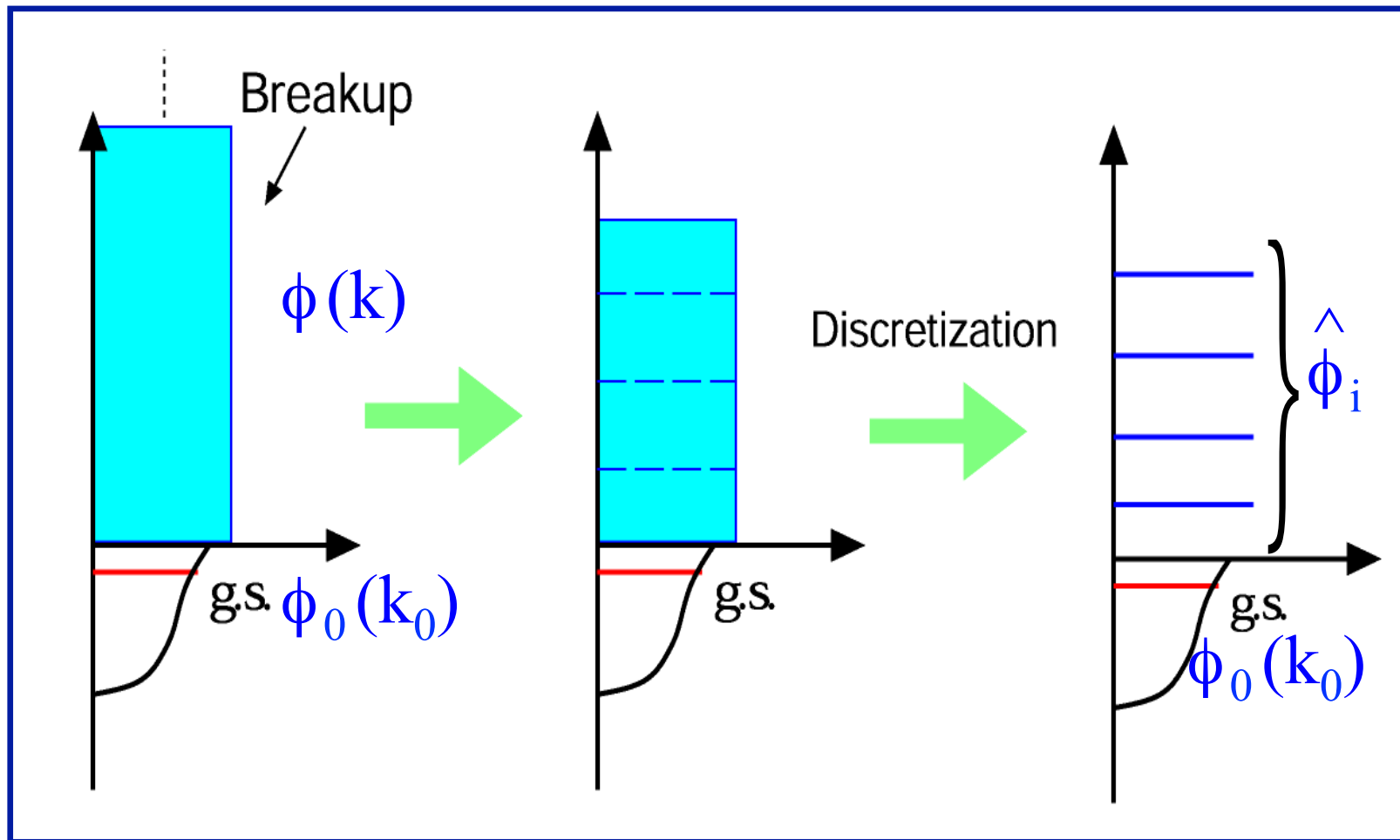
$$\psi(\vec{r}, \vec{R}) \cong \phi_0(k_0, \vec{r}) \chi_0(K_0, \vec{R}) + \underbrace{\sum_{i=1}^{i_{\max}} \int_{k_{i-1}}^{k_i} \phi(k, \vec{r}) \chi(K, \vec{R}) dk}_{\text{Truncation and Discretization}}$$

Truncation and Discretization



$$\psi(\vec{r}, \vec{R}) \cong \phi_0(k_0, \vec{r}) \chi_0(K_0, \vec{R}) + \sum_{i=1}^{i_{\max}} \chi(K_i, \vec{R}) \int_{k_{i-1}}^{k_i} \phi(k, \vec{r}) dk$$


 Truncation and Discretization



$$\psi(\vec{r}, \vec{R}) \cong \phi_0(k_0, \vec{r}) \chi_0(K_0, \vec{R}) + \sum_{i=1}^{i_{\max}} \chi(K_i, \vec{R}) \int_{k_{i-1}}^{k_i} \phi(k, \vec{r}) dk$$

THE END.

$$\psi^{\text{CDCC}}(\vec{r}, \vec{R}) = \sum_{i=0}^{i_{\max}} \hat{\phi}_i(\vec{r}) \hat{\chi}_i(K_i, \vec{R})$$

Truncation and Discretization

Reaction

$^{208}\text{Pb}(^8\text{B}, ^7\text{Be}+p)$ at 250 A MeV and 100 A MeV

$^{208}\text{Pb}(^{11}\text{Be}, ^{10}\text{Be}+n)$ at 250 A MeV and 100 A MeV

Projectile wave function and distorting potential

Standard Woods-Saxon

Modelspace

^8B

$$l_{\max} = 3$$

$$N_s = 20, N_{p-d} = 10,$$

$$N_f = 5$$

$$\varepsilon_{\max} = 10 \text{ MeV}$$

$$r_{\max} = 200 \text{ fm}$$

$$R_{\max} = 500 \text{ fm}$$

$$N_{\text{ch}} = 138$$

^{11}Be

$$l_{\max} = 3$$

$$N_{s,p} = 20, N_d = 10,$$

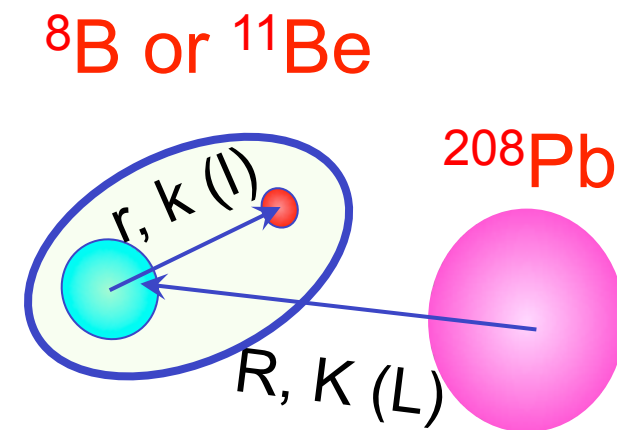
$$N_f = 5$$

$$\varepsilon_{\max} = 10 \text{ MeV}$$

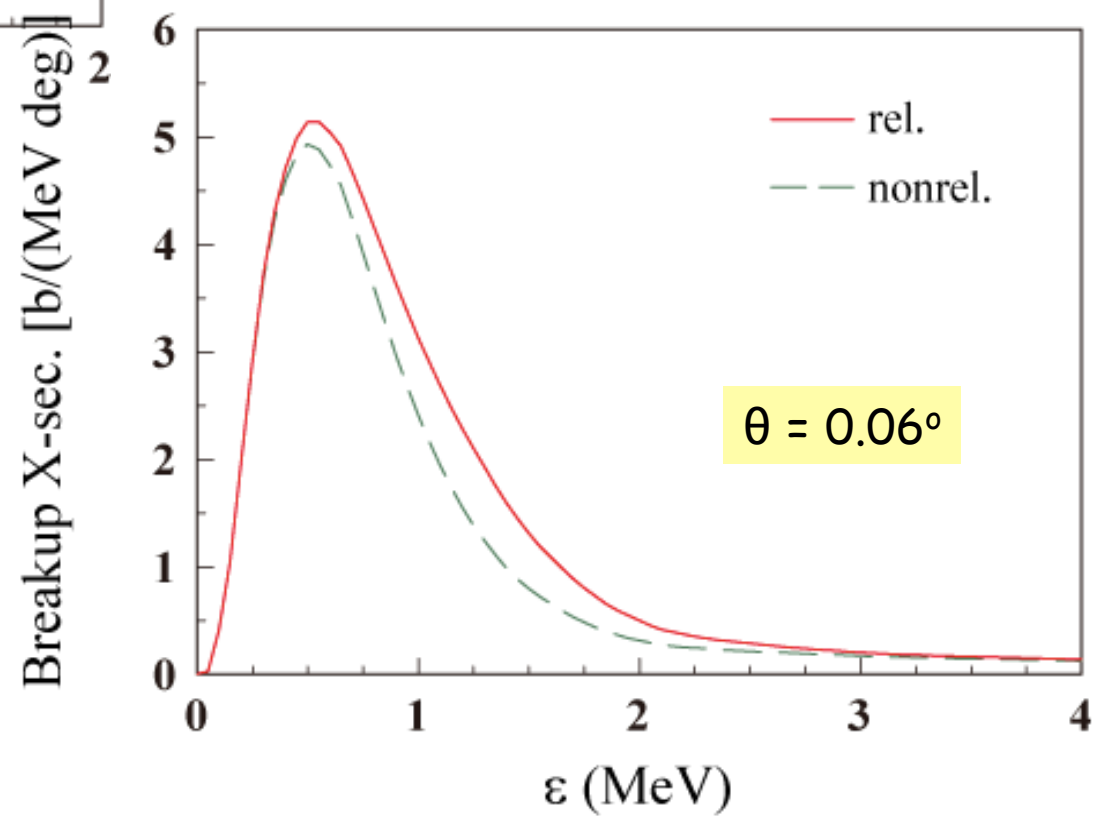
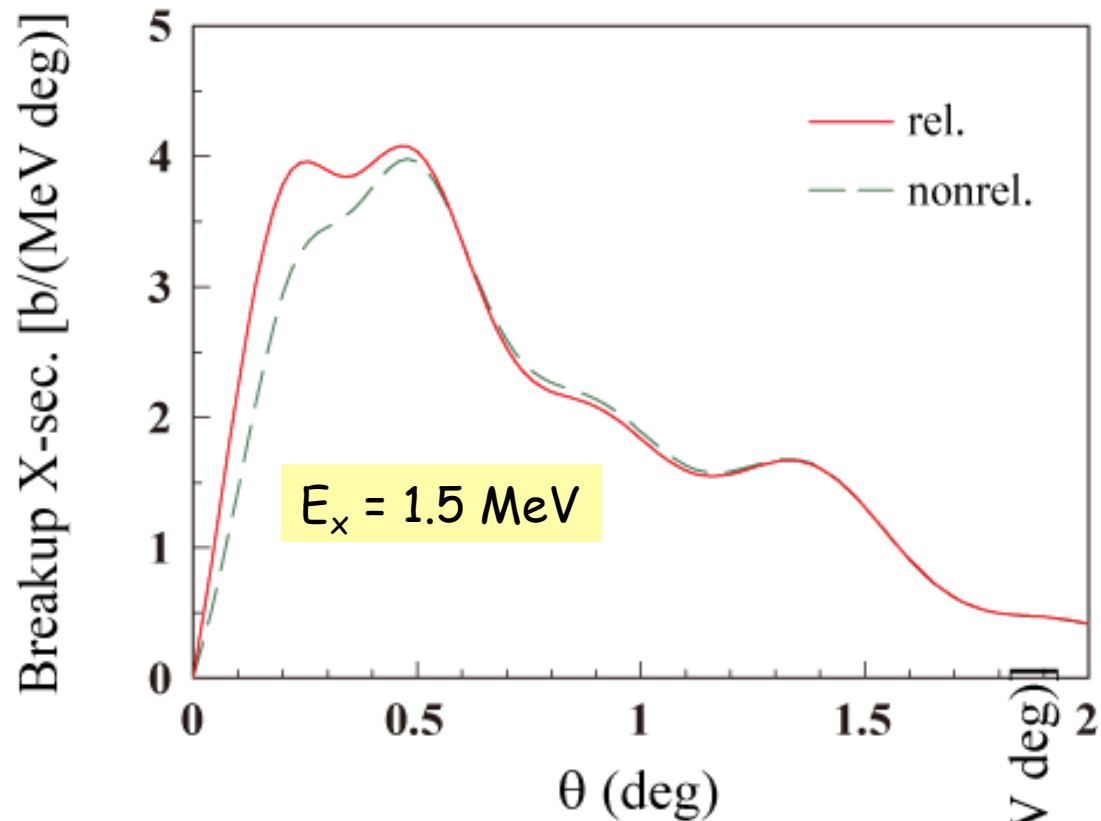
$$r_{\max} = 200 \text{ fm}$$

$$R_{\max} = 450 \text{ fm}$$

$$N_{\text{ch}} = 166$$



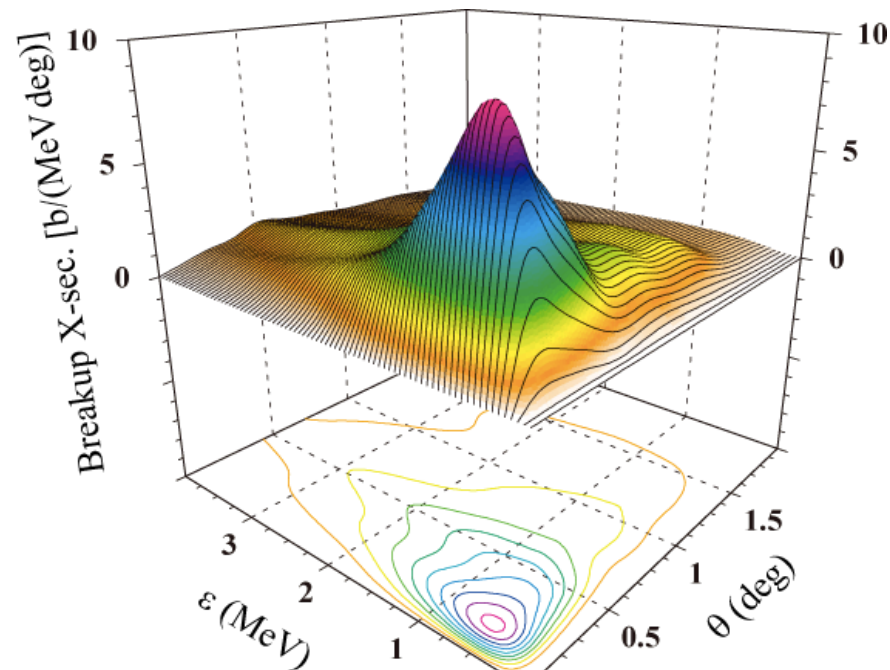
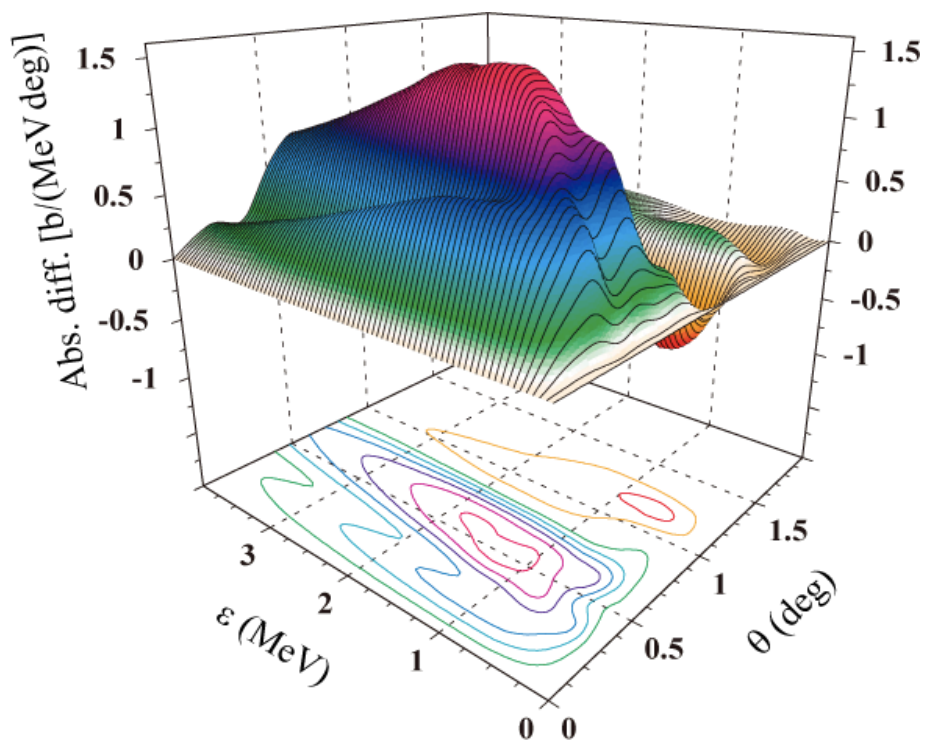
Pb(^8B , $p^7\text{Be}$) at 250 MeV/nucleon



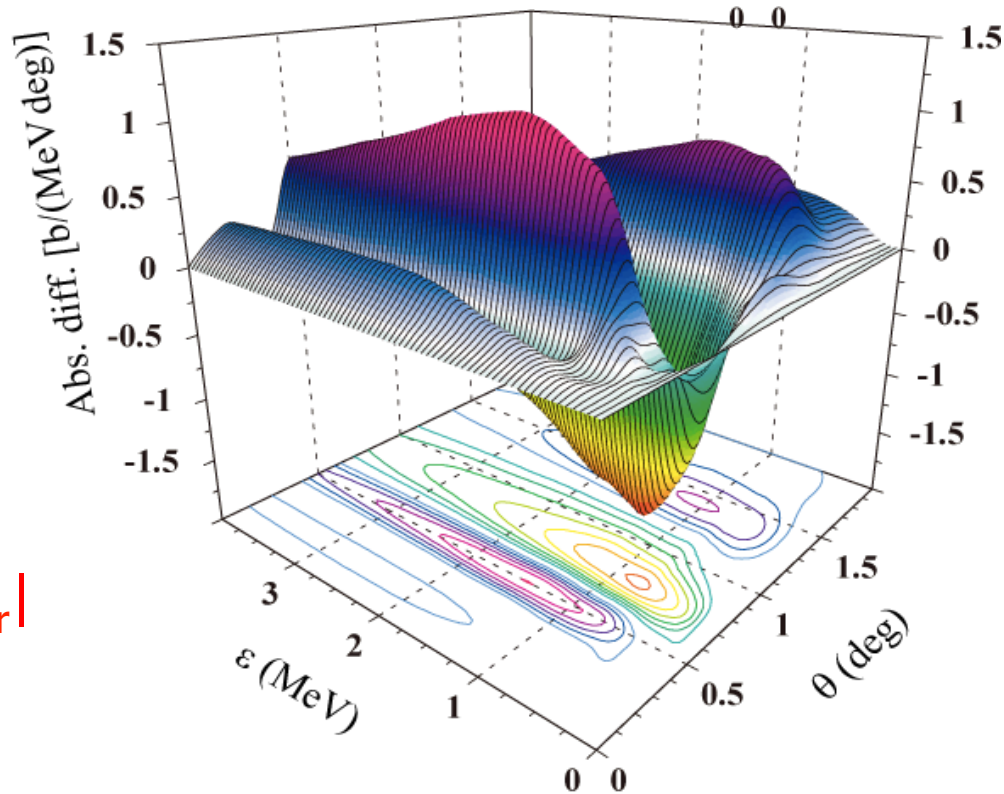
Pb(^8B , p ^7Be) at 250 MeV/nucleon

all orders

$|\sigma_{\text{all}} - \sigma_{\text{NR}}|$



$|\sigma_{\text{all}} - \sigma_{\text{no-nuclear}}|$



Things can get worse



Relativistic MF nucleus-nucleus potential

Long, Bertulani, PRC 83, 024907 (2011).

σ, ω, ρ and γ exchange

$$E = \int d^3r \sum_a \bar{\psi}_a (-i\boldsymbol{\gamma} \cdot \nabla + M) \psi_a$$

$$+ \frac{1}{2} \sum_{\phi=\sigma,\omega,\rho,\gamma} \int d^3r d^3r' \sum_{ab} \bar{\psi}_a(\mathbf{r}) \bar{\psi}_b(\mathbf{r}') \Gamma_\phi(\mathbf{r}, \mathbf{r}') D_\phi(\mathbf{r} - \mathbf{r}') \psi_a(\mathbf{r}) \psi_b(\mathbf{r}')$$

$$\Gamma_\phi(\mathbf{r}, \mathbf{r}') = -g_\sigma(\mathbf{r}) g_\sigma(\mathbf{r}')$$

$$\Gamma_\omega(\mathbf{r}, \mathbf{r}') = -\left(g_\omega \gamma^\mu\right)_\mathbf{r} \cdot \left(g_\omega \gamma_\mu\right)_{\mathbf{r}'}$$

$$\Gamma_\rho(\mathbf{r}, \mathbf{r}') = -\left(g_\rho \gamma^\mu \vec{\tau}\right)_\mathbf{r} \cdot \left(g_\rho \gamma_\mu \vec{\tau}\right)_{\mathbf{r}'}$$

$$\Gamma_\gamma(\mathbf{r}, \mathbf{r}') = \frac{e^2}{4} \left[\gamma^\mu (1 - \tau_z)\right]_\mathbf{r} \cdot \left[\gamma_\mu (1 - \tau_z)\right]_{\mathbf{r}'}$$

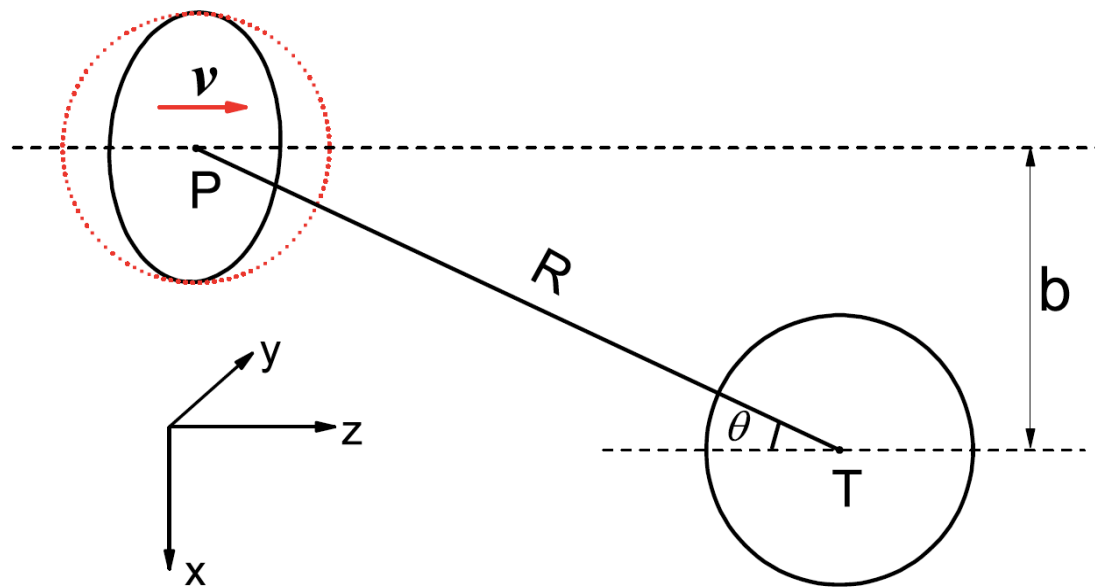
$$D_\phi = \frac{1}{4\pi} \frac{e^{m_\phi |\mathbf{r} - \mathbf{r}'|}}{|\mathbf{r} - \mathbf{r}'|}$$

$$D_\gamma = \frac{1}{|\mathbf{r} - \mathbf{r}'|}$$

Lorentz transform

$$x_p = x_t + b, \quad y_p = y_t$$

$$z_p = \gamma(z_t + R \cos \theta)$$



$$E(A_t, A_p, v) = E(A_t) + E(A_p, v) + \mathbf{E}(A_t, A_p, v)$$

$$\mathbf{E}(A_t, A_p, v) = \sum_{\phi=\sigma, \omega, \rho, \gamma} \int d^3 r \int d^3 r' \sum_{ab} \bar{\psi}_{t,a}(\mathbf{r}) \bar{\psi}_{p,b}(\mathbf{r}') \Gamma_{\phi}(\mathbf{r}, \mathbf{r}') D_{\phi}(\mathbf{r} - \mathbf{r}') \psi_{t,a}(\mathbf{r}) \psi_{p,b}(\mathbf{r}')$$

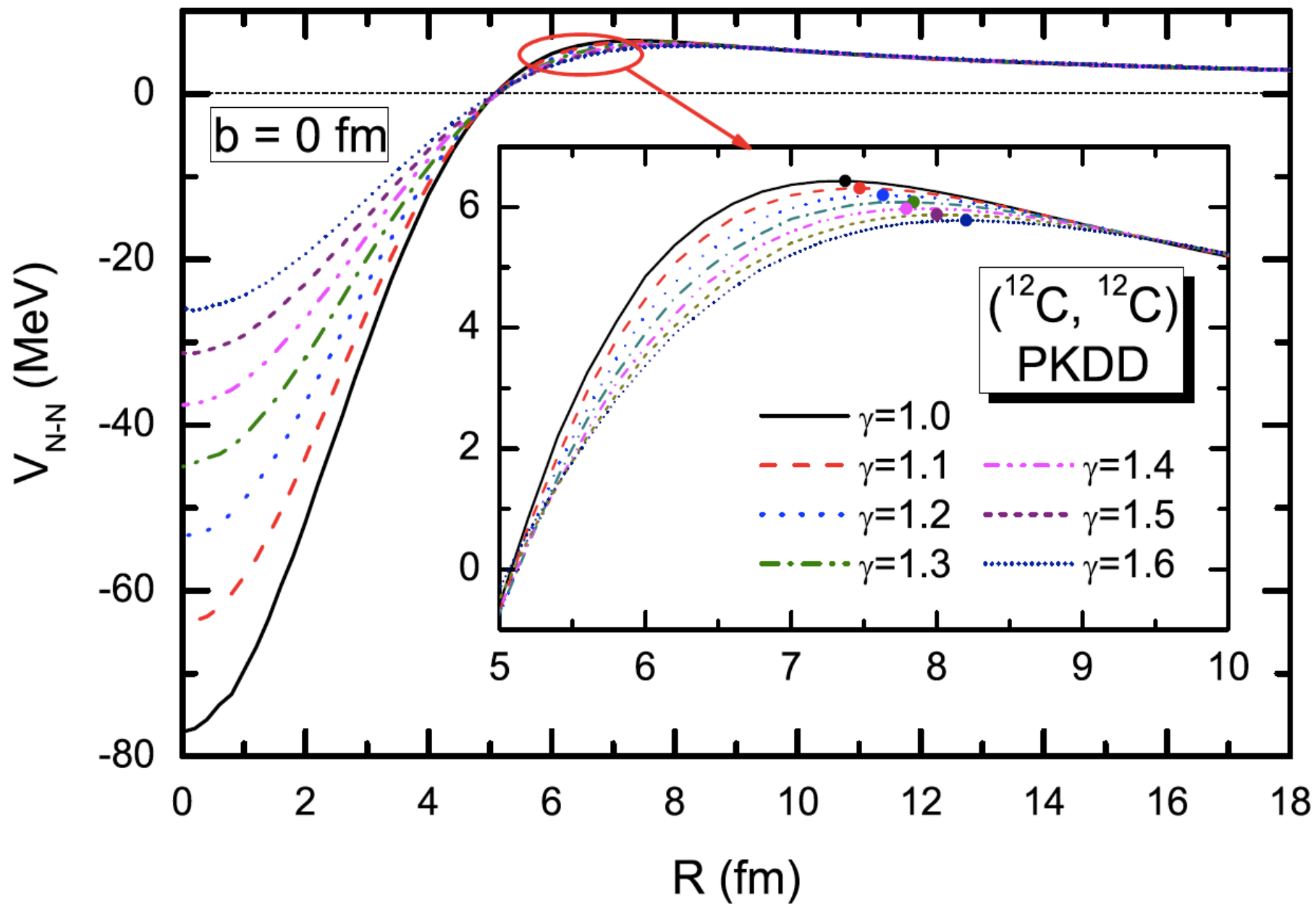
Ex: σ and ω contributions

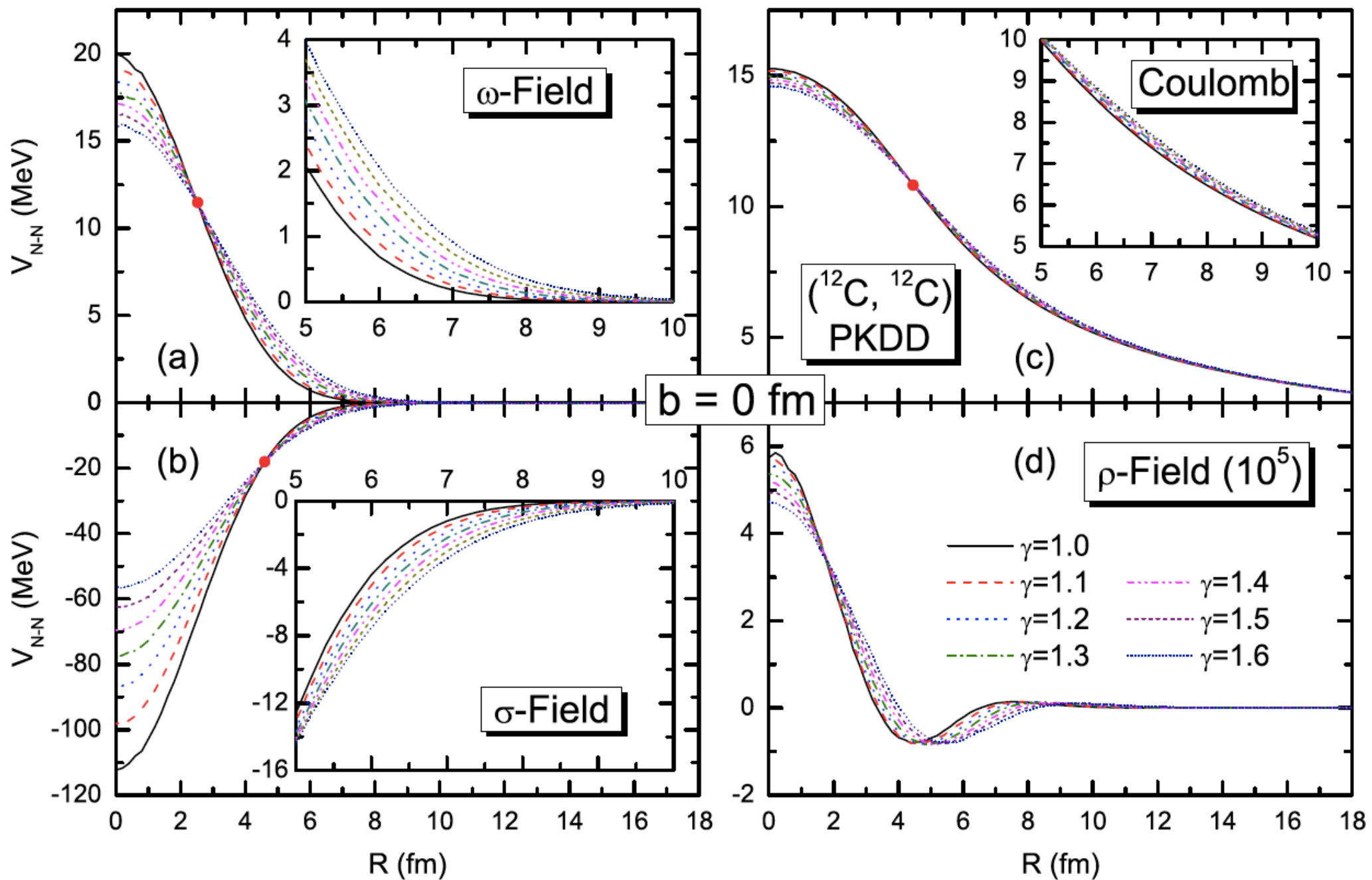
$$\mathbf{E}_{\sigma} = -\frac{1}{\gamma} \int d^3 r_t \int d^3 r'_p g_{\sigma}(\mathbf{r}_t) \rho_{s,t}(\mathbf{r}_t) D_{\sigma}(\mathbf{r} - \mathbf{r}') \rho_{s,p}(\mathbf{r}'_p) g_{\sigma}(\mathbf{r}'_p)$$

$$\mathbf{E}_{\omega} = \int d^3 r_t \int d^3 r'_p g_{\omega}(\mathbf{r}_t) \rho_{b,t}(\mathbf{r}_t) D_{\omega}(\mathbf{r} - \mathbf{r}') \rho_{b,p}(\mathbf{r}'_p) g_{\omega}(\mathbf{r}'_p)$$

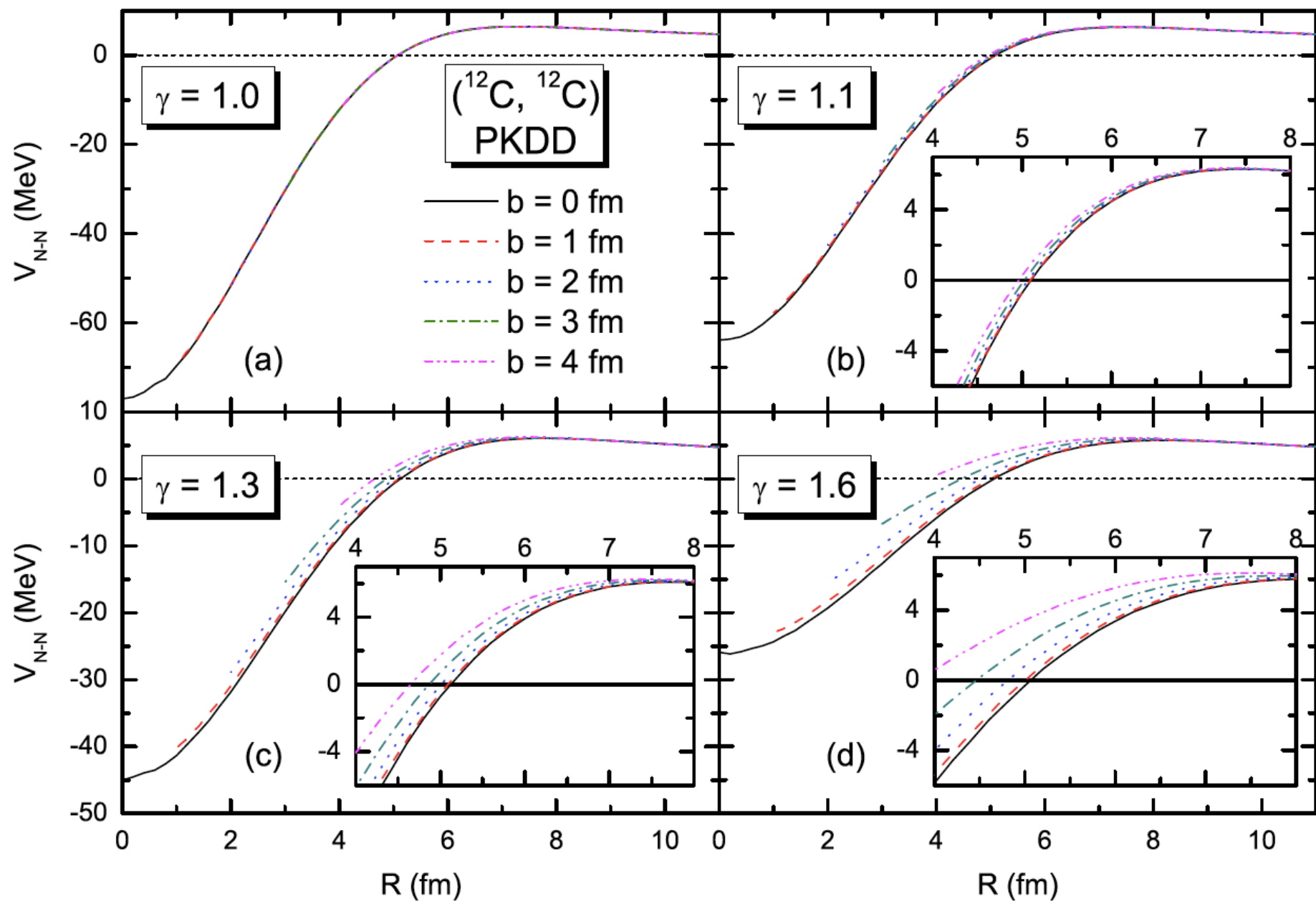
$$\rho_s(\mathbf{r}) = \sum_a \bar{\psi}_a(\mathbf{r}) \psi_a(\mathbf{r}), \quad \rho_b(\mathbf{r}) = \sum_a \bar{\psi}_a(\mathbf{r}) \gamma^0 \psi_a(\mathbf{r})$$

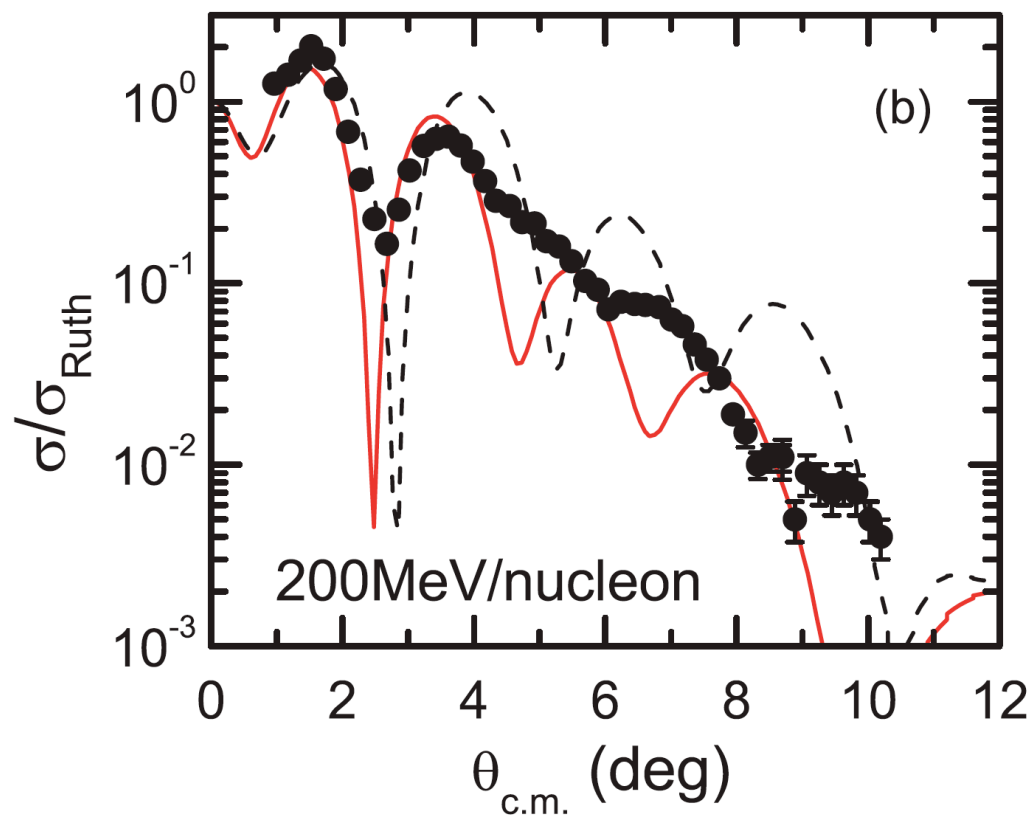
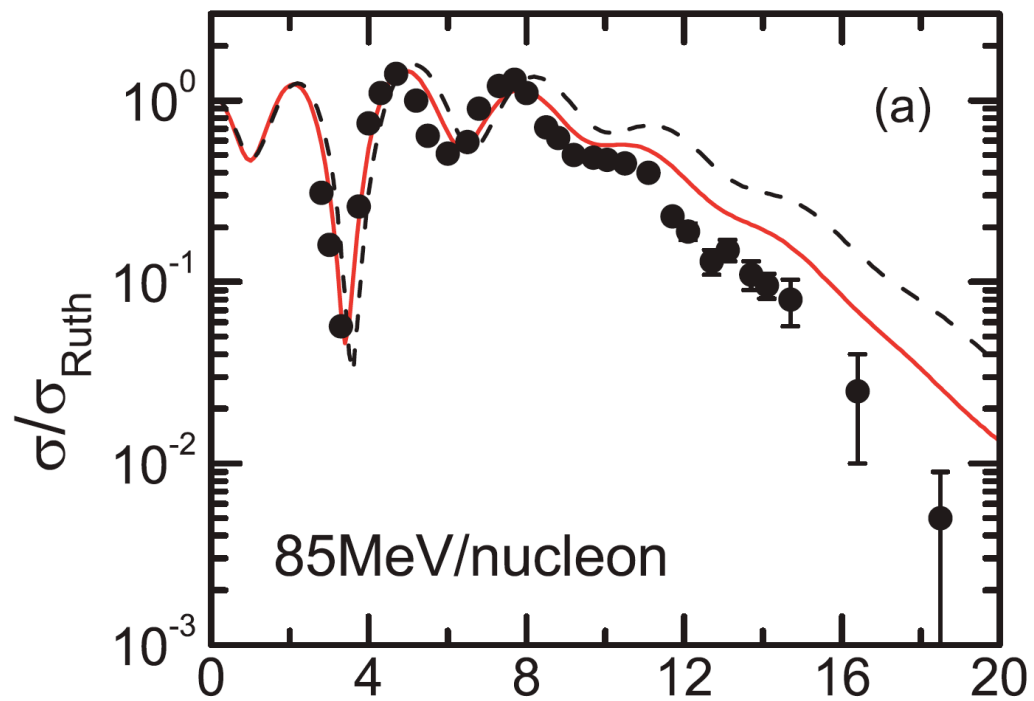
Projectile densities boosted to the target frame





Dependence on energy and impact parameter





OUT OF TIME