#### Geometries of D1-D5-P states

#### Stefano Giusto

Università di Padova

#### Exotic structures of spacetime, YITP , March 2013

#### Based on

- S.G., L. Martucci, M. Petrini, R. Russo: arXiv:1306.1745
- S.G., R. Russo: arXiv:1311.5536

#### Introduction: the D1-D5-P system

• Paradigmatic example of BPS black hole in string theory (Strominger, Vafa)

Type II B

 $\mathbb{R}$   $\times$   $\mathbb{R}^4$   $\times$   $S^1$   $\times$   $T^4$ x<sup>i</sup> y za t D1: - • 0 D5: -0  $P: - \circ$ 0

### The "naive" D1-D5-P geometry

$$ds^{2} = -\frac{2}{\sqrt{Z_{1}Z_{2}}} dv \left( du + \frac{\mathcal{F}}{2} dv \right) + \sqrt{Z_{1}Z_{2}} ds_{4}^{2} + \sqrt{\frac{Z_{1}}{Z_{2}}} ds_{7^{4}}^{2}$$
$$v = \frac{t + y}{\sqrt{2}} \quad , \quad u = \frac{t - y}{\sqrt{2}} \quad , \quad ds_{4}^{2} = dx^{i} dx^{i}$$
$$Z_{i} = 1 + \frac{Q_{i}}{r^{2}} \quad , \quad \mathcal{F} = -\frac{Q_{p}}{r^{2}}$$

 $Q_1$ : D1 charge ,  $Q_2$ : D5 charge ,  $Q_p$ : P charge

- The geometry has an "extremal" horizon of finite area at r = 0.
- The geometry only depends on the charges: "no-hair theorem".

## The near-horizon decoupling limit

In the limit

$$r^2\,,\, Q_p\,\ll\, Q_1\,,\, Q_2 \quad \Rightarrow \quad Z_i o Z_i - 1$$

one obtains a geometry that is  $AdS_3 \times S^3 \times T^4$  asymptotically  $(r^2 \gg Q_p)$ 

$$ds^2 \approx -rac{r^2}{\sqrt{Q_1 Q_2}} 2 \, du \, dv + rac{\sqrt{Q_1 Q_2}}{r^2} \, dr^2 + \sqrt{Q_1 Q_2} \, d\Omega_3^2 + \sqrt{rac{Q_1}{Q_2}} \, ds_{T^4}^2$$

This limit is described by a CFT:

(4,4) sigma model with target space (a deformation of)  $(T^4)^{n_1 n_5}/S_{n_1 n_5}$ 

↓ D1-D5 CFT

4/25

## Microstate geometries

- There exist solutions of supergravity that reduce to the "naive" D1-D5-P geometry for large *r* up to  $O\left(\frac{1}{r^2}\right)$  F (and hence carry the same charges as the D1-D5-P black hole) but are completely smooth and have no horizon.
- In the near-horizon limit these geometries are dual to states of the D1-D5 CFT.



Conjecture: these geometries account for (a finite fraction of) the entropy of the D1-D5-P black hole.

(Mathur, Bena, Warner,...)

#### Our goal

To develop systematic methods to construct geometries that

- carry D1,D5,P charges and preserve 4 supersymmetries;
- are dual to states of the D1-D5 CFT.

Note: for geometries carrying only two charges (egs. D1-D5) this problem has been solved in full generality.

(Lunin, Mathur; Kanitscheider, Skenderis, Taylor)

- We will follow the steps:
  - find the general supergravity ansatz that preserves the same susy as D1-D5-P;
  - select solutions dual to known states of the D1-D5 CFT: apply symmetries of the CFT to known microstates.

### General D1-D5-P ansatz I

We apply methods of Generalized Geometry in 10D (Tomasiello)

- $\exists \epsilon_1, \epsilon_2$  Majorana-Weyl Killing spinors;
- the spinor bilinear K ≡ -<sup>1</sup>/<sub>2</sub>(ε<sub>1</sub>Γ<sup>M</sup>ε<sub>1</sub> + ε<sub>2</sub>Γ<sup>M</sup>ε<sub>2</sub>)∂<sub>M</sub> is a Killing vector;
- we restrict to the case *K* is null  $\Rightarrow K = \frac{\partial}{\partial u}$ ;
- other spinor bilinears

(*i*) 
$$\chi \equiv -\frac{1}{2} (\bar{\epsilon}_1 \Gamma^M \epsilon_1 - \bar{\epsilon}_2 \Gamma^M \epsilon_2) dx^M$$
 1 - form  
(*ii*)  $\Psi \equiv -32 \epsilon_1 \otimes \bar{\epsilon}_2 \Gamma_{(10)}$  polyform

satisfy

$$d\chi = i_{\mathcal{K}} \mathcal{H}$$
 ,  $(d - \mathcal{H}) \wedge (e^{-\phi} \Psi) = i_{\mathcal{K}} \mathcal{F} + \chi \wedge \mathcal{F}$ 

(*H*, *F* NSNS and RR field strengths);

#### General D1-D5-P ansatz II

- the missing susy equations come from the *v*-component of the gravitino variation ⇒ extra constraints;
- the previous equations plus the Bianchi identities for *H* and *F* imply all the equations of motion apart from the *vv*-component of Einstein equations ⇒ *R<sub>vv</sub>* + ... = 0;

• split the metric as  $10 \rightarrow \begin{array}{cc} (u,v) & x^i & z^a \\ 2 + 4 + 4 & and restrict (for simplicity) \\ to geometries that are homogeneous and isotropic with respect to <math>T^4 \Rightarrow$  non-trivial dependence on  $(v, x^i)$ ;

• impose that the susy preserved is of D1-D5 type:

$$\Gamma_{D_1} \epsilon_2 = \epsilon_1 \quad , \quad \Gamma_{D5} \epsilon_2 = \epsilon_1$$

#### General D1-D5-P ansatz III

the various conditions constrain the metric to be of the form

$$ds^{2} = -\frac{2\alpha}{\sqrt{Z_{1}Z_{2}}}(dv+\beta)\left[du+\omega+\frac{\mathcal{F}}{2}(dv+\beta)\right] + \sqrt{Z_{1}Z_{2}} ds_{4}^{2} + \sqrt{\frac{Z_{1}}{Z_{2}}} ds_{7^{4}}^{2}$$
  
with  $\alpha = \frac{Z_{1}Z_{2}}{Z_{1}Z_{2} - Z_{4}^{2}}$   
 $e^{2\phi} = \dots$ ,  $B = \dots$ ,  $F = \dots$ 

all fields of type IIB sugra can be excited;

### General D1-D5-P ansatz IV

$$ds^{2} = -\frac{2\alpha}{\sqrt{Z_{1}Z_{2}}}(dv+\beta) \Big[ du+\omega+\frac{\mathcal{F}}{2}(dv+\beta) \Big] + \sqrt{Z_{1}Z_{2}} \, ds_{4}^{2} + \sqrt{\frac{Z_{1}}{Z_{2}}} \, ds_{74}^{2}$$

 the sugra equations can be organized in such a way that the problem reduces to an almost linear one

(Bena, SG, Shigemori, Warner)

- the non-linear part involves  $ds_4^2$ ,  $\beta$ :
  - ds<sup>2</sup><sub>4</sub> admits an almost hyperkäler structure

 $J_A \wedge J_B = -2 \, \delta_{AB} \operatorname{vol}_4 \quad , \quad J_A = - *_4 J_A \quad , \quad dJ_A = \partial_{\nu} (\beta \wedge J_A)$ 

•  $D\beta = *_4 D\beta$  with  $D \equiv d - \beta \wedge \partial_v$ ;

the equations for Z<sub>1</sub>, Z<sub>2</sub>, Z<sub>4</sub>, ω, F... are linear (but inhomogeneous), if solved in the right order.

10/25

#### v-dependence

- The sugra equations simplify for *v*-independent geometries:
  - ds<sub>4</sub><sup>2</sup> is hyperkähler;
  - $d\beta = *_4 d\beta$ .
- On general grounds, geometries are expected to be *v*-independent if they are dual to eigenstates of momentum.
- Typically classical geometries are dual to coherent states (egs. e<sup>α a<sub>-n</sub></sup>|ψ<sub>0</sub>) with a<sub>-n</sub> carrying momentum n)
  - $\Rightarrow$  generic 3-charge microstates are *v*-dependent.

#### D1-D5 geometries (Lunin, Mathur; Kanitscheider, Skenderis, Taylor)

- D1-D5 microstates are v-independent.
- They have a flat 4D base:  $ds_4^2 = dx^i dx^i$ .
- The (*T*<sup>4</sup>-isometric) geometries are encoded in a profile in ℝ<sup>5</sup> (*F<sup>i</sup>*(*v'*), *F*(*v'*)) :

$$\beta = \frac{-A+B}{\sqrt{2}} , \ A = -\frac{Q_2}{L} \int_0^L \frac{dv' \dot{F}^i}{|x-F|^2} dx^i , \ dB = -*_4 dA$$
$$Z_2 = 1 + \frac{Q_2}{L} \int_0^L \frac{dv'}{|x-F|^2} , \ Z_1 = 1 + \frac{Q_2}{L} \int_0^L \frac{dv' |\dot{F}|^2}{|x-F|^2}$$
$$Z_4 = -\frac{Q_2}{L} \int_0^L \frac{dv' \dot{F}}{|x-F|^2} , \ \omega = -\frac{A+B}{\sqrt{2}}$$

•  $\beta \neq 0$  : KK-monopole dipole charge from binding D1 and D5.

## Adding momentum to D1-D5 I

• Start from a D1-D5 geometry and take its near-horizon limit:

 $Z_1 \rightarrow Z_1 - 1$  ,  $Z_2 \rightarrow Z_2 - 1 \Rightarrow$  asymptotically AdS geometry

 Act with an operator of the CFT chiral algebra that preserves susy and adds momentum:

•  $L_{-n}$ : Virasoro; •  $J_{-n}^{\alpha}$ :  $SU(2)_L$  R-symmetry; •  $J_{-n}^{z^a}$ :  $U(1)^4$   $T^4$  translations.

On the gravity side this is equivalent to performing a change of coordinates that does not vanish at the boundary of AdS.

At the perturbative level, this method was applied by Mathur, Saxena, Srivastava; Mathur, Turton; Shigemori.

## Adding momentum to D1-D5 II

 Rewrite metric and gauge fields in the form of the general ansatz and read-off ds<sup>2</sup><sub>4</sub>, β, Z<sub>I</sub>, ω, F.

One generates a *v*-dependent solution of the sugra equations that is asymptotically  $AdS_3 \times S^3 \times T^4$ .

 Extend the metric back to the asymptotically flat region, where the metric reduces to ℝ<sup>4,1</sup> × S<sup>1</sup> × T<sup>4</sup>.

Note: the replacement  $Z_1 \rightarrow Z_1 + 1$ ,  $Z_2 \rightarrow Z_2 + 1$ , leaving the remaining metric functions unchanged, does not produce a solution of the sugra constraints.

This is a non-trivial problem that can be solved exactly in some cases.

## Case 1: *v*-independent $ds_4^2$

- Assume that ds<sup>2</sup><sub>4</sub>, β do not depend on v after the change of coordinates.
- Assume also that  $ds_4^2$ ,  $\beta$  have a tri-holomorphic isometry  $\frac{\partial}{\partial \tau} \Rightarrow$

$$ds_4^2 = V^{-1}(d au + A)^2 + V ds_3^2 \quad , \quad *_3 dA = dV$$
  
 $eta = rac{K_3}{V}(d au + A) + \xi \quad , \quad *_3 d\xi = -dK_3$ 

- Egs:  $ds_4^2 = \text{flat } \mathbb{R}^4 \implies \tau = \phi + \psi$ ,  $V = \frac{1}{r}$
- The remaining metric coefficients may depend on v and  $\tau$ .

# Case 1: v-independent ds<sub>4</sub><sup>2</sup> - An example

Start from the D1-D5 geometry associated with the profile

$$F^{1} + iF^{2} = a e^{\frac{2\pi i v'}{L}}$$
,  $F^{3} = F^{4} = 0$ ,  $\mathcal{F} = -b \sin\left(\frac{2\pi v'}{L}\right)$ 

The geometry depends on the angle of the (1, 2) plane but  $\beta$  does not.

- Apply the operator  $e^{\chi(J_{-1}^+ J_1^-)} = e^{\Sigma} e^{\chi(J_0^+ J_0^-)} e^{-\Sigma}$  $\phi \to \phi + \frac{t}{R}, \psi \to \psi + \frac{y}{R} \leftarrow \text{spectral flow rotation in } \mathbb{R}^4$ Note: when b = 0 the action of this operator is trivial.
- One can see that  $ds_4^2$  and  $\beta$  are left invariant, while the other metric coefficients transform non-trivially under this operator.

# Case 1: v-independent ds<sub>4</sub><sup>2</sup> - The solution I

 Niehoff, Warner found a general way to solve the sugra constraints in Case 1:

(almost) all metric coefficients can be expressed algebraically in terms of harmonic functions for the covariant Laplacian  $*_4D *_4D$ .

Egs: 
$$Z_1 = L_1 + \frac{K_2 K_3}{V}$$
,  $Z_2 = L_2 + \frac{K_1 K_3}{V}$  with  
 $\partial_{\tau} K_{2,1} + \partial_{\nu} L_{1,2} = 0$ ,  $*_4 D *_4 D K_{1,2} = *_4 D *_4 D L_{1,2} = 0$ 

To find  $\omega$  one has to solve a system of partial differential equations.

# Case 1: v-independent ds<sub>4</sub><sup>2</sup> - The solution II

 To construct an asymptotically flat 3-charge geometry one starts from the geometric data generated by the chiral algebra transformation, extracts the corresponding harmonic functions, and replaces

#### $L_1 \rightarrow 1 + L_1 \quad , \quad L_2 \rightarrow 1 + L_2$

The coefficients of other harmonic functions have to be modified to preserve regularity and absence of Dirac-Misner strings.

- There is a unique asymptotically flat solution that solves all the regularity constraints and matches the near-horizon solution.
- The asymptotic charges of the solution match the values expected from the CFT (more on this in Rodolfo's talk).

## Case 2: v-dependent $ds_4^2$

- After the change of coordinates  $ds_4^2$  and  $\beta$  are *v*-dependent.
- An example:
  - Start from the 2-charge metric (in the near-horizon limit) with

 $F^{1} + iF^{2} = a e^{\frac{2\pi i v'}{L}}, F^{3} = F^{4} = \mathcal{F} = 0;$ 

• Apply the transformation  $e^{\epsilon (J_{-n}^3 - J_n^3)}$  which is equivalent to

(\*) 
$$\psi \to \psi - \epsilon \sin\left(\frac{n\sqrt{2}v}{R}\right), \phi \to \phi - \epsilon \sin\left(\frac{n\sqrt{2}v}{R}\right) \Leftrightarrow x^i \to x^i - f^i(v)$$

- The coordinate transformation generates the 3-charge near-horizon geometry; ds<sup>2</sup><sub>4</sub>, β depend on v.
- The Niehoff, Warner method does not apply directly: how to construct the asymptotically flat solution?
- Note: the transformation (\*) cannot be applied to the asymptotically flat solution because it screws up the asymptotic structure.

# Case 2: v-dependent $ds_4^2$ - The solution

- Strategy: apply the Niehoff, Warner method before the change of coordinates (when ds<sup>2</sup><sub>4</sub> and β are ν-independent) to generate a new non-explicitly asymptotically flat solution that becomes explicitly asymptotically flat after the transformation (\*).
- The new solution needs to have the following asymptotic limit

 $Z_{1,2} 
ightarrow 1$  ,  $\omega 
ightarrow -\dot{f}_i(v) dx^i$  ,  $\mathcal{F} 
ightarrow -|\dot{f}(v)|^2$ 

Note: these boundary conditions force all metric coefficients (apart from  $ds_4^2$ ,  $\beta$ ) to be *v*-dependent.

• Result: there is a unique solution with the correct asymptotic behavior that is regular.

#### Case 2: relation with "Vachaspati transform"

- The solution generating technique we employ to add momentum to D1-D5 geometries is analogous to the one used by Dabholkar et al. and Callan at al. to construct D1-P geometries:

  - start from the "naive" D1 solution; turn on  $\omega = -\dot{f}_i(v)dx^i$ ,  $\mathcal{F} = -|\dot{f}(v)|^2$ ;
  - perform the coordinate shift  $x^i \rightarrow x^i f^i(v)$ .
- One important technical difference: the D1 solution has

 $\beta = \mathbf{0} \Rightarrow \frac{\partial}{\partial u}$  is hypersurface orthogonal

 $\Rightarrow$  constant  $\omega$  and  $\mathcal{F}$  solve the equations of motion.

 Moreover, for D1-P solutions all equations are linear and homogeneous  $\Rightarrow$  one can superpose solutions

### Multiply-wound profiles

• One can also considers profiles that are multiply-wound along S<sup>1</sup> :

$$f^i\Big(\mathbf{v}+2\pi\,\mathbf{w}\,\frac{R}{\sqrt{2}}\Big)=f^i(\mathbf{v})\,,\quad\mathbf{w}>1$$

- This is crucial to obtain the full D1-P entropy.
- The shift  $x^i \rightarrow x^i f^i(v)$  is not defined globally on spacetime.
- An equivalent description:  $f^i(v) \rightarrow f^i_{\alpha}(v)$ ,  $\alpha = 1, ..., w$  with

$$f^{i}_{\alpha}\Big(\mathbf{v}+2\pi\,\frac{\mathbf{R}}{\sqrt{2}}\Big)=f^{i}_{\alpha+1}(\mathbf{v})$$

- For D1-P: use linearity to superpose solutions for different α.
- For D1-D5-P: the problem is non-linear.
   How to construct geometries for multiply-wound profiles?

### Summary

- We have determined the general class of susy solutions carrying D1,D5,P charges.
- Within this class, we have constructed geometries dual to BPS states of the D1-D5 CFT with momentum.
- The class of microstates we can construct is very particular: chiral algebra descendants of RR ground states.
- It is non-trivial that a unique, regular and asymptotically flat solution exists with the quantum numbers predicted by the CFT.

### Outlook

- Main open problems:
  - can one extend the construction of microstate geometries to more generic states?
     Note: constructing geometries corresponding to multi-wound profiles would be a big step in this direction;
  - can microstate geometries account for a finite fraction of the 3-charge entropy?
- Extract more informations from the geometries we have:
  - compute holographically 1-point functions of chiral operators in the 3-charge microstates and compare with the CFT;

(Kanitscheider, Skenderis, Taylor)

• compute holographically the entanglement entropy in the microstate geometries.

(Ryu, Takayanagi)

### Entanglement entropy in black hole microstates

- Most computations of entanglement entropy are done in the vacuum or the thermal state  $\Rightarrow AdS_3 \times S^3 \times T^4$ .
- $T^4$ -isotropic microstates can be easily reduced on  $T^4$  $\Rightarrow ds_6^2$  6D Einstein metric.
- In the near-horizon limit ds<sup>2</sup><sub>6</sub> is asymptotically AdS<sub>3</sub> × S<sup>3</sup> but depends non-trivially on S<sup>3</sup>.
- The Ryu, Takayanagi holographic recipe for computing entanglement entropy applies to asymptotically AdS<sub>3</sub> spaces.
- A small generalization: consider a 1D domain *A* and a geodesic (with respect to  $ds_6^2$ ) 4-manifold  $\Gamma_A = \gamma_A \times S^3$ , where  $\gamma_A$  is a curve at t = const. such that  $\partial \gamma_A = \partial A$ ; the entanglement entropy is

$$S_{A} = \frac{\operatorname{Area}(\Gamma_{A})}{4G_{6}}$$

 Can one compare with the D1-D5 CFT prediction? (work in progress with R. Russo)