

# In-medium $\bar{K}$ & $\eta$ mesons

Mesic Nuclei, JU Krakow, Sept. 2013  
Hadrons in Nuclei, YITP Kyoto, Oct. 2013

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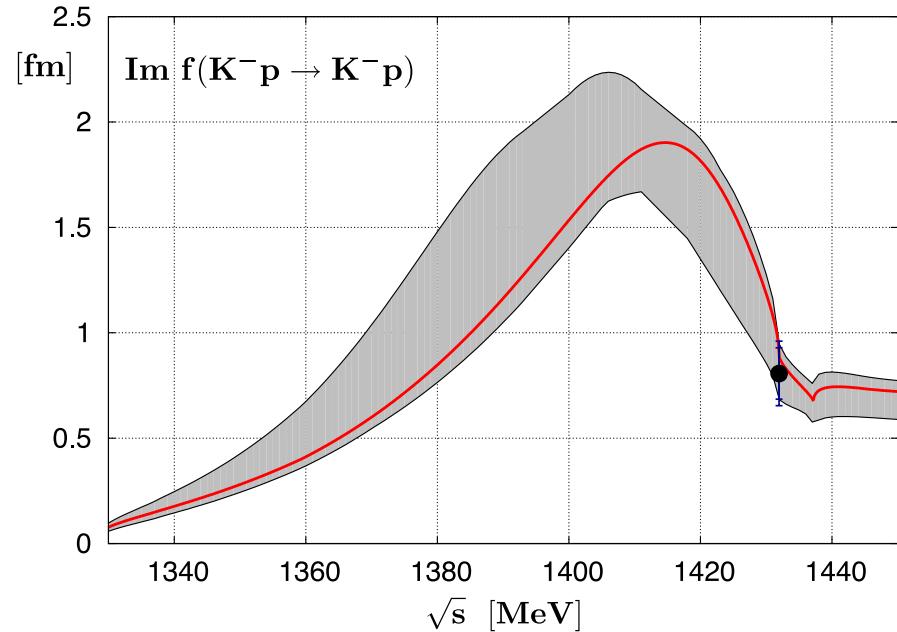
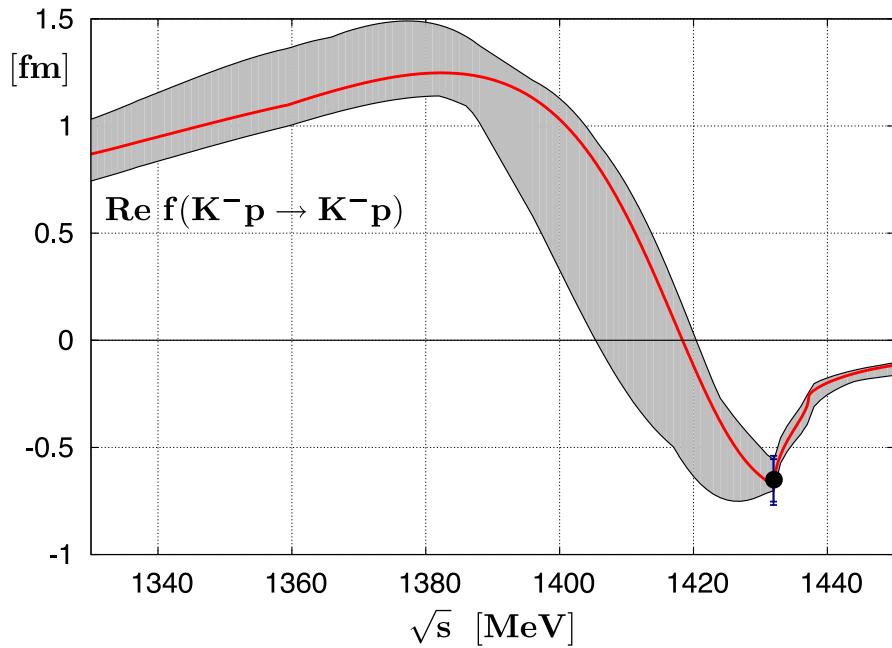
- $\bar{K}N - \pi Y$  chiral dynamics and its consequences
- $\bar{K}$  nuclear few-body systems
- $\bar{K}$ -nucleus potentials from  $K^-$  atoms

A.Gal in HYP2012 Proc., NPA 914 (2013) 270

- Quest for  $\eta$  nuclear quasibound states  
E.Friedman, A.Gal, J.Mareš, PLB 725 (2013) 334

# $\bar{K}N - \pi Y$ Chiral Dynamics

# $K^- p$ scattering amplitude from NLO chiral SU(3) dynamics



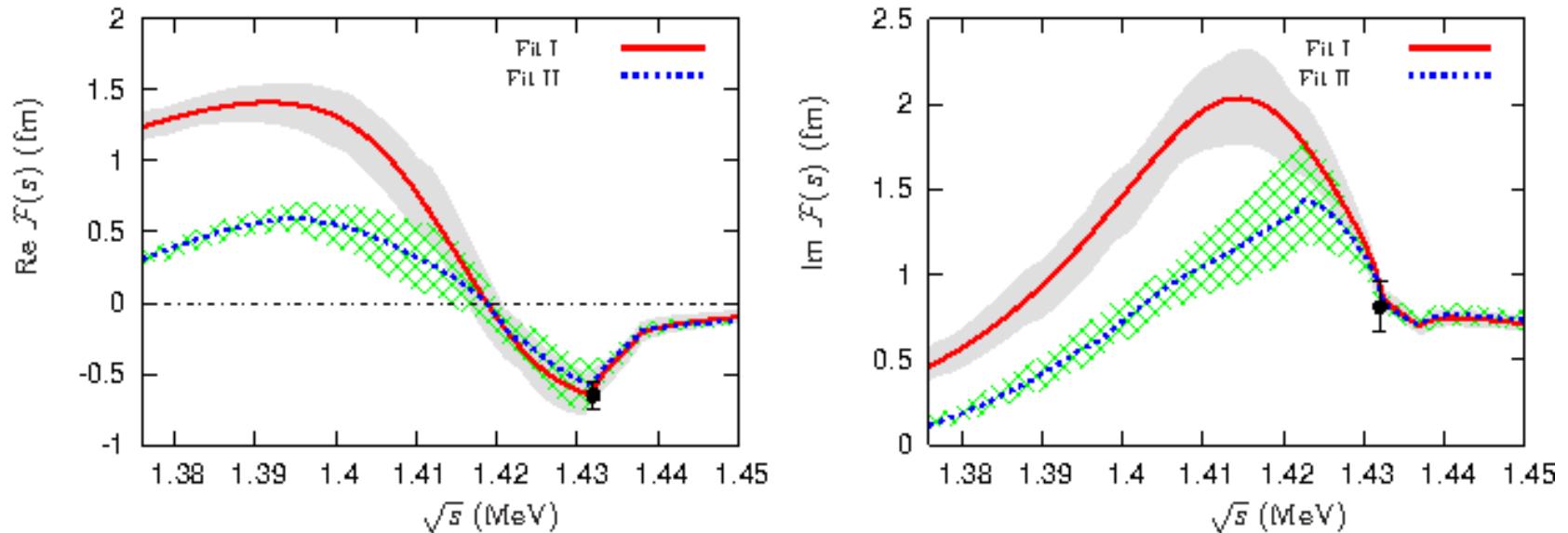
Y. Ikeda, T. Hyodo, W. Weise (IHW), PLB **706** (2011) 63; NPA **881** (2012) 98

Strong subthreshold  $K^- p$  attraction;  $\Lambda(1405)$  physics

Consequences for kaonic atoms and  $K^-$  nuclear quasibound states

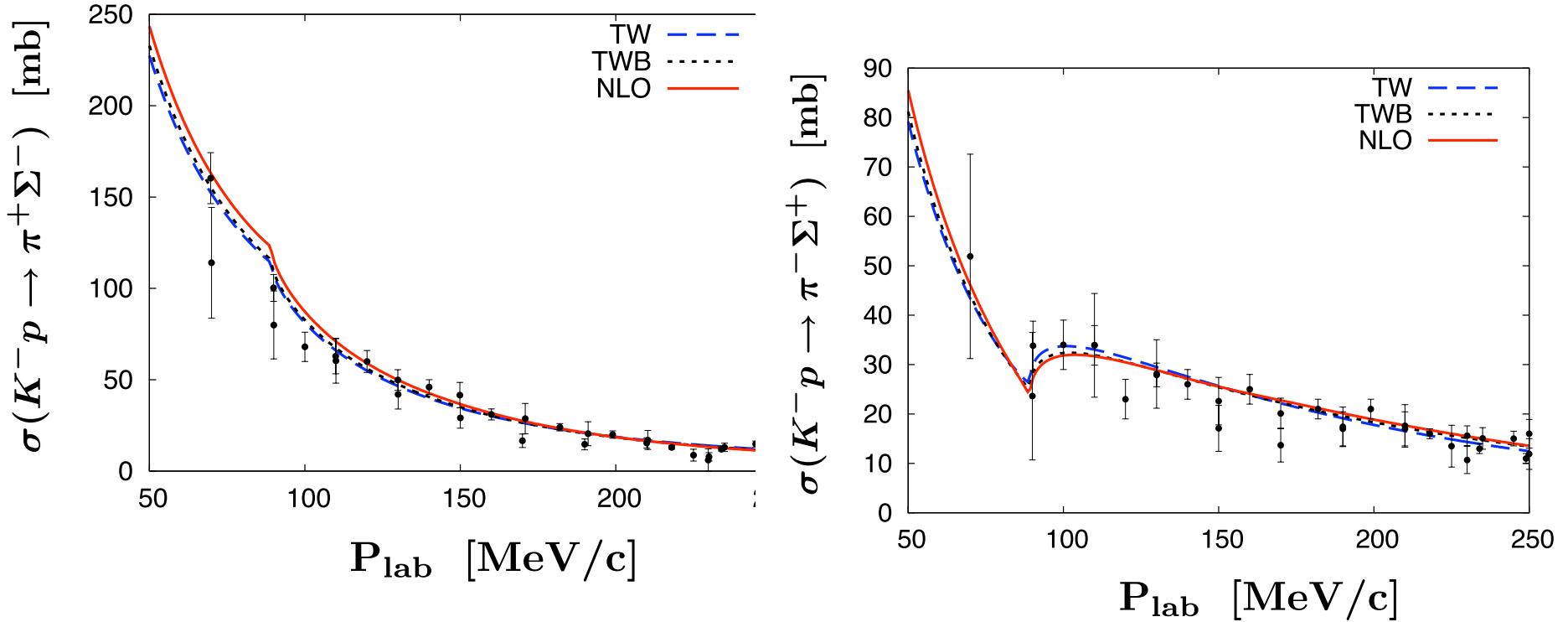
$K^-$  absorption might be governed by out-of-model  $K^- NN \rightarrow YN$

# $K^- p$ subthreshold ambiguity



Two NLO chiral-model fits by Guo-Oller, PRC 87 (2013) 035202

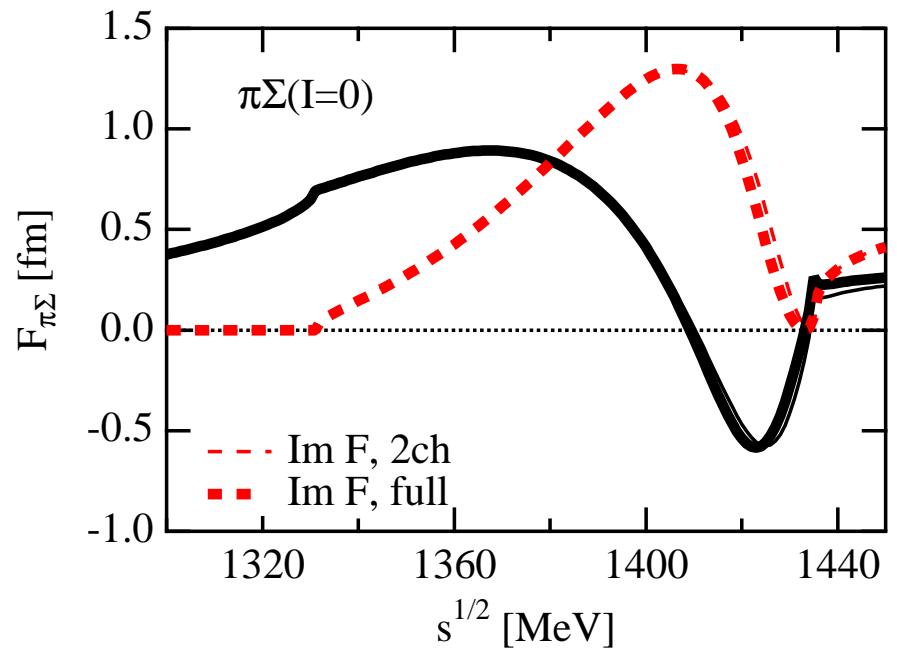
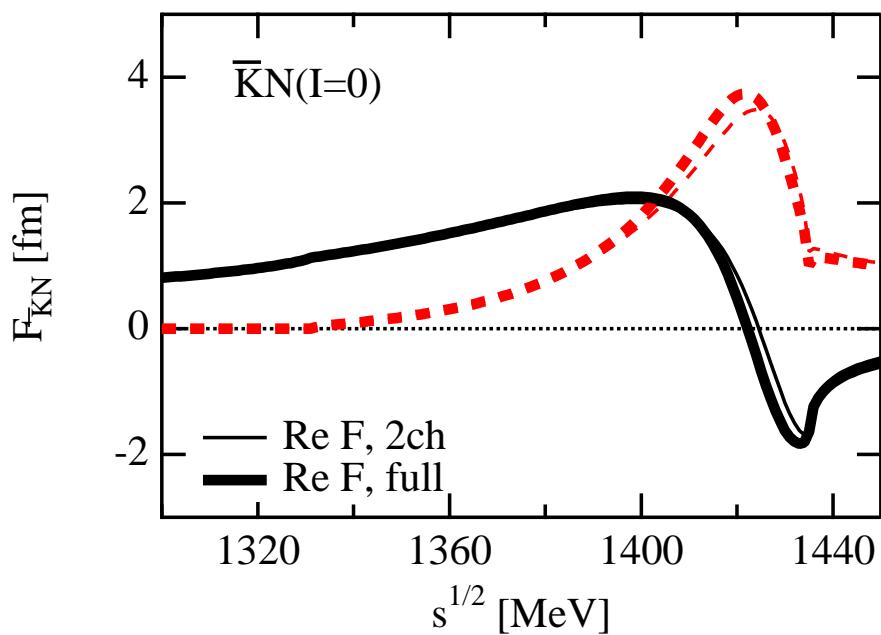
- Fit I: one value of meson weak-decay constant  $f = 125.7 \pm 1.1$  MeV.
- Fit II: separate fixed values for  $f_\pi$ ,  $f_K$ ,  $f_\eta$ .  
Fit II will create problems when confronted with kaonic-atom data.



$K^- p \rightarrow \pi^\pm \Sigma^\mp$  reaction data fitted by LEC of NLO scheme for  
 $\bar{K}N - \pi Y$  coupled channels ( $Y = \Lambda, \Sigma$ )

Y. Ikeda, T. Hyodo, W. Weise, NPA 881 (2012) 98

Large difference in cross sections  $\Rightarrow$  Strong isospin dependence



T. Hyodo, W. Weise, PRC 77 (2008) 035204

$I = 0$  coupled-channel amplitudes

Location of ‘resonances’:  $\bar{K}N \approx 1420$  MeV,  $\pi\Sigma \approx 1405$  MeV

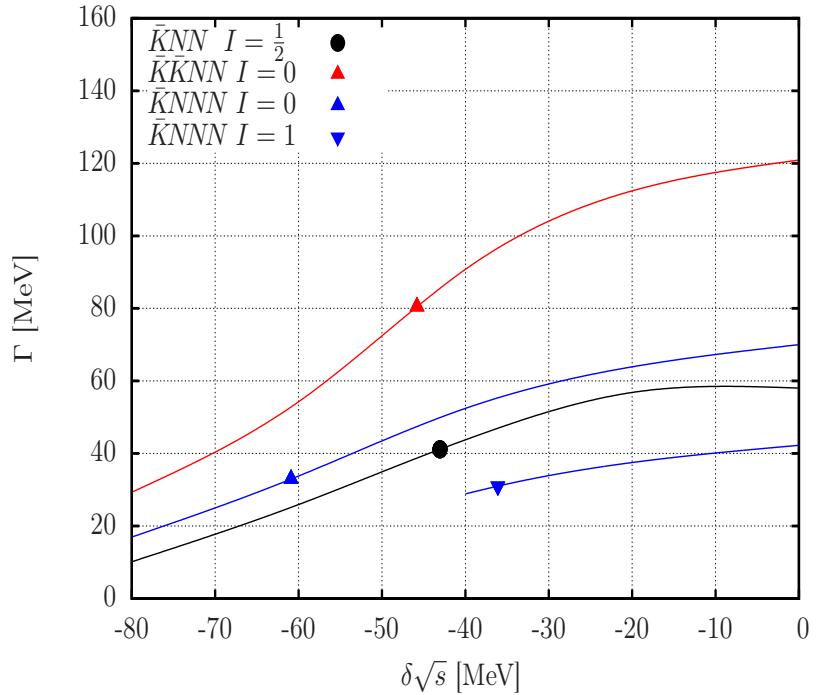
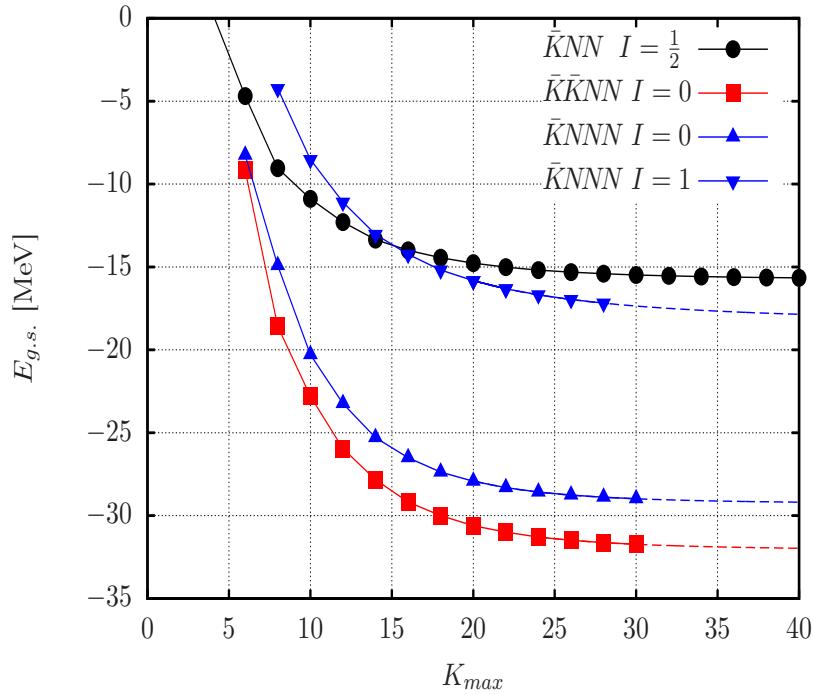
Are there two distinct ‘ $\Lambda(1405)$ ’ resonances?

# $\overline{K}$ nuclear few-body systems

# Energy dependence in $\bar{K}$ nuclear few-body systems

- $\Lambda(1405)$  induces strong energy dependence of the scattering amplitudes  $f_{\bar{K}N}(\sqrt{s})$  and the underlying effective single-channel input potentials  $v_{\bar{K}N}(\sqrt{s})$ .
- $s = (\sqrt{s_{\text{th}}} - B_K - B_N)^2 - (\vec{p}_K + \vec{p}_N)^2 \leq s_{\text{th}}$
- Expanding nonrelativistically near  $\sqrt{s_{\text{th}}} \equiv m_K + m_N$ :  
$$\delta\sqrt{s} = -\frac{B}{A} - \frac{A-1}{A}B_K - \xi_N \frac{A-1}{A}\langle T_{N:N} \rangle - \xi_K \left(\frac{A-1}{A}\right)^2 \langle T_K \rangle,$$
$$\delta\sqrt{s} \equiv \sqrt{s} - \sqrt{s_{\text{th}}}, \quad B_K = -E_K, \quad \xi_{N(K)} \equiv \frac{m_{N(K)}}{(m_N + m_K)}.$$
- Self-consistency: output  $\sqrt{s}$  from solving the Schroedinger equation identical with input  $\sqrt{s}$ .

# 3- & 4-body $B$ & $\Gamma$ calculated self-consistently



N. Barnea, A. Gal, E.Z. Liverts, PLB **712** (2012)

- Variational calculation in hyperspherical basis controlled by  $K_{max}$
- $\bar{K}N$  energy dependence [Hyodo–Weise, PRC 77 (2008) 035204] restrains  $B$  &  $\Gamma$  by treating  $\delta\sqrt{s_{\bar{K}N}}$  self-consistently
- $B$ (4-body) small w.r.t. non-chiral estimates of over 100 MeV

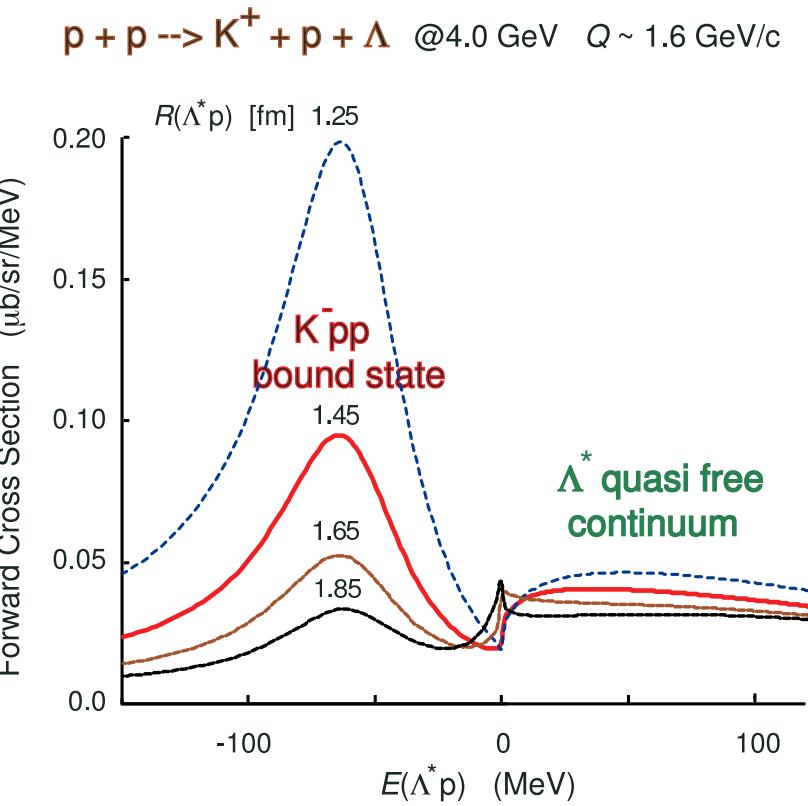
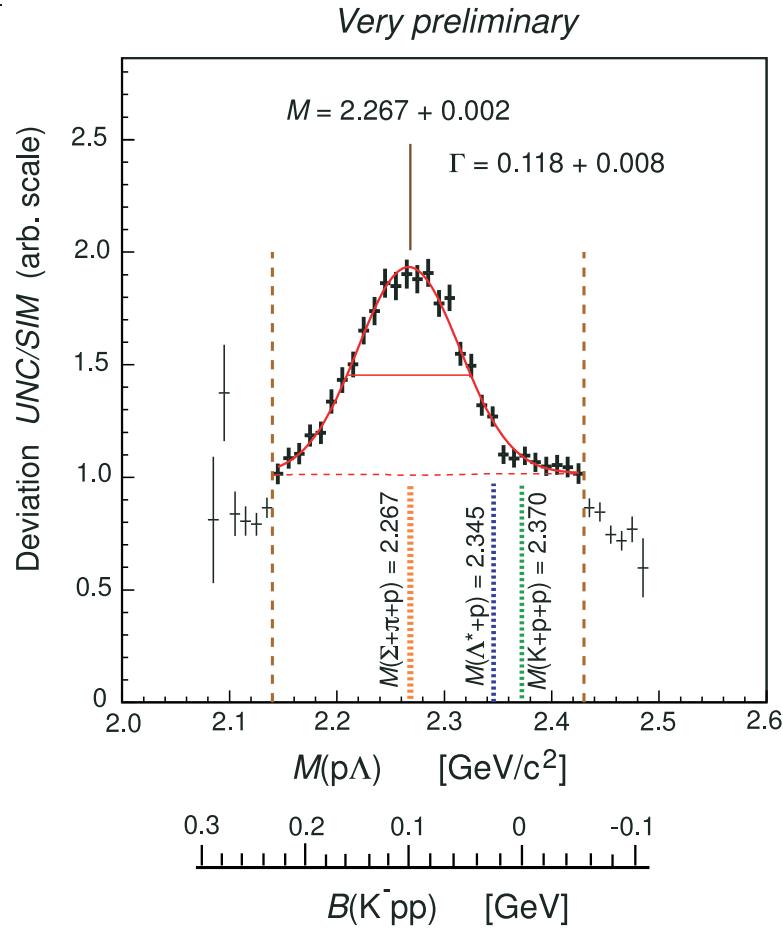
- $\bar{K}NN$ : is there an excited  $I = 1/2$  quasibound state ( $\bar{K}d$ , dominantly  $I_{NN} = 0$ ) on top of “ $K^-pp$ ” g.s. ?
- Bayar & Oset [NPA 881 (2012) 127]: **YES**, bound by about 9 MeV, from a peak in  $|T_{\bar{K}NN}|^2$  calculated in a fixed-scatterer approximation to Faddeev equations.
- Shevchenko [NPA 890-1 (2012) 50]: **UNLIKELY**, judging from the  $K^-d$  scattering length and effective range deduced from a  $\bar{K}NN$  Faddeev calculation.
- Barnea, Gal & Liverts do not find such a bound state below the  $\Lambda^*N$  threshold at  $B = 11.4$  MeV.

## $K^-pp$ calculated binding energies & widths (in MeV)

	chiral, energy dependent			non-chiral, static calculations			
	var. [1]	var. [2]	Fad. [3]	var. [4]	Fad. [5]	Fad. [6]	var. [7]
B	16	17–23	9–16	48	50–70	60–95	40–80
$\Gamma$	41	40–70	34–46	61	90–110	45–80	40–85

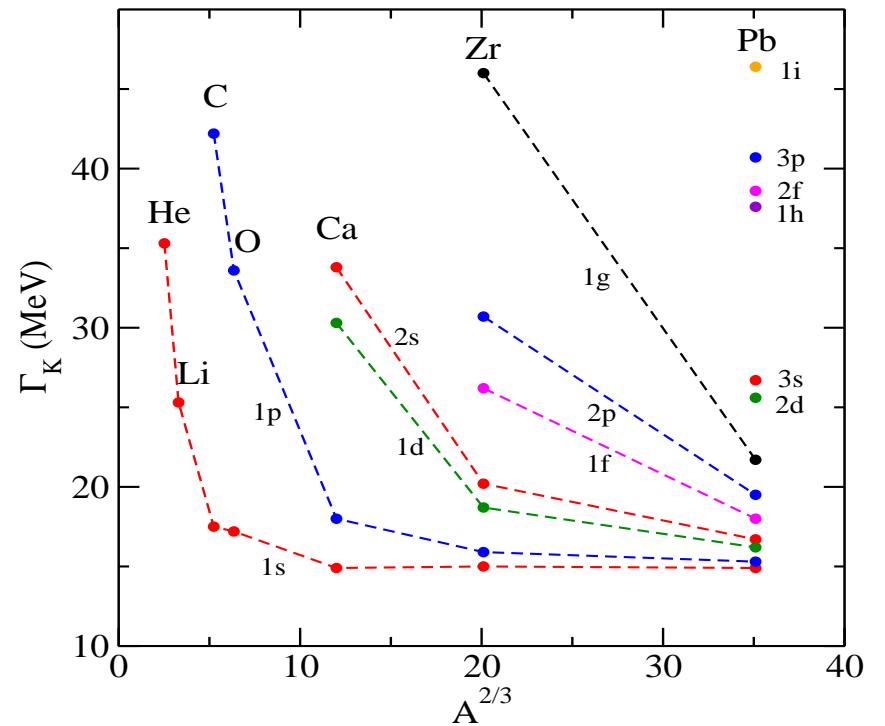
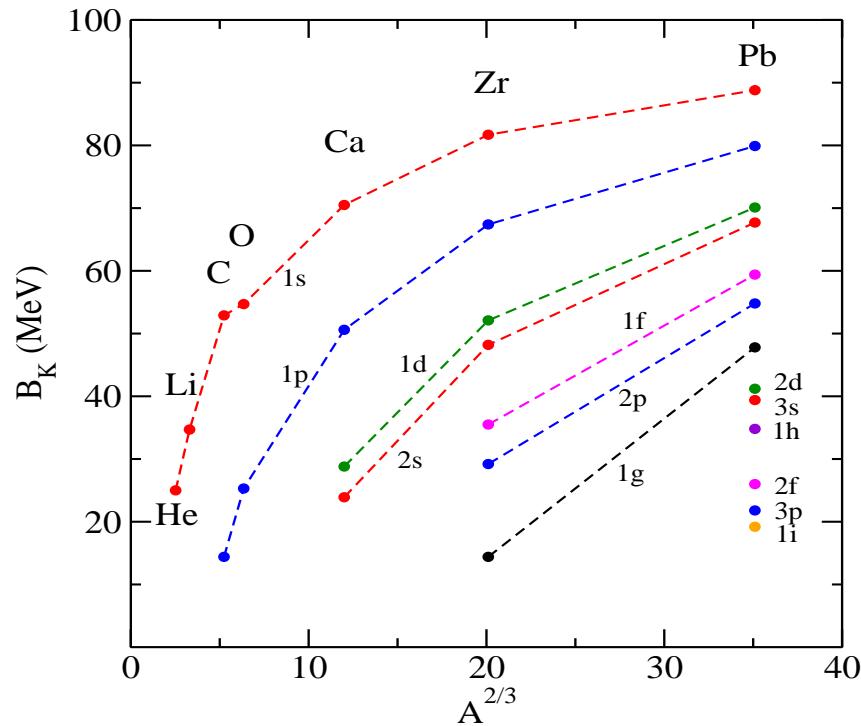
1. N. Barnea, A. Gal, E.Z. Liverts, PLB **712** (2012)
2. A. Doté, T. Hyodo, W. Weise, NPA **804** (2008) 197, PRC **79** (2009) 014003
3. Y. Ikeda, H. Kamano, T. Sato, PTP **124** (2010) 533
4. T. Yamazaki, Y. Akaishi, PLB **535** (2002) 70
5. N.V. Shevchenko, A. Gal, J. Mareš, PRL **98** (2007) 082301
6. Y. Ikeda, T. Sato, PRC **76** (2007) 035203, PRC **79** (2009) 035201
7. S. Wycech, A.M. Green, PRC **79** (2009) 014001 (including  $p$  waves)

Robust binding & large widths; chiral models give weak binding



Yamazaki et al. PRL 104 (2010) 132502, DISTO data reanalysis at 2.85 GeV  
**Broad  $K^-pp$  structure in  $pp \rightarrow \Lambda p K^+$  at  $\pi N \Sigma$  threshold**  
Forthcoming experiments:  $pp \rightarrow (K^-pp) + K^+$  at GSI  
 $K^-{}^3He \rightarrow (K^-pp) + n$  (E15) &  $\pi^+d \rightarrow (K^-pp) + K^+$  (E27) at J-PARC

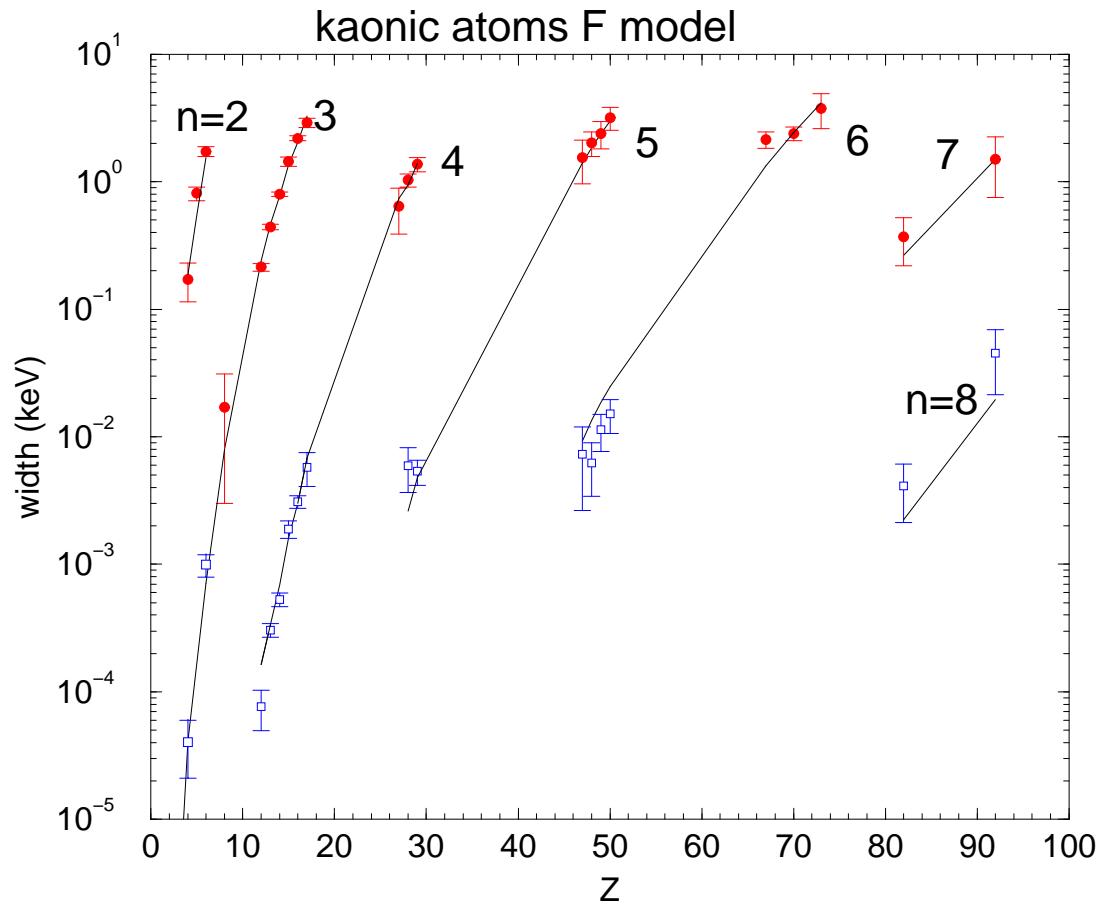
# RMF quasibound spectra calculated self-consistently (NLO30 '+' SE')



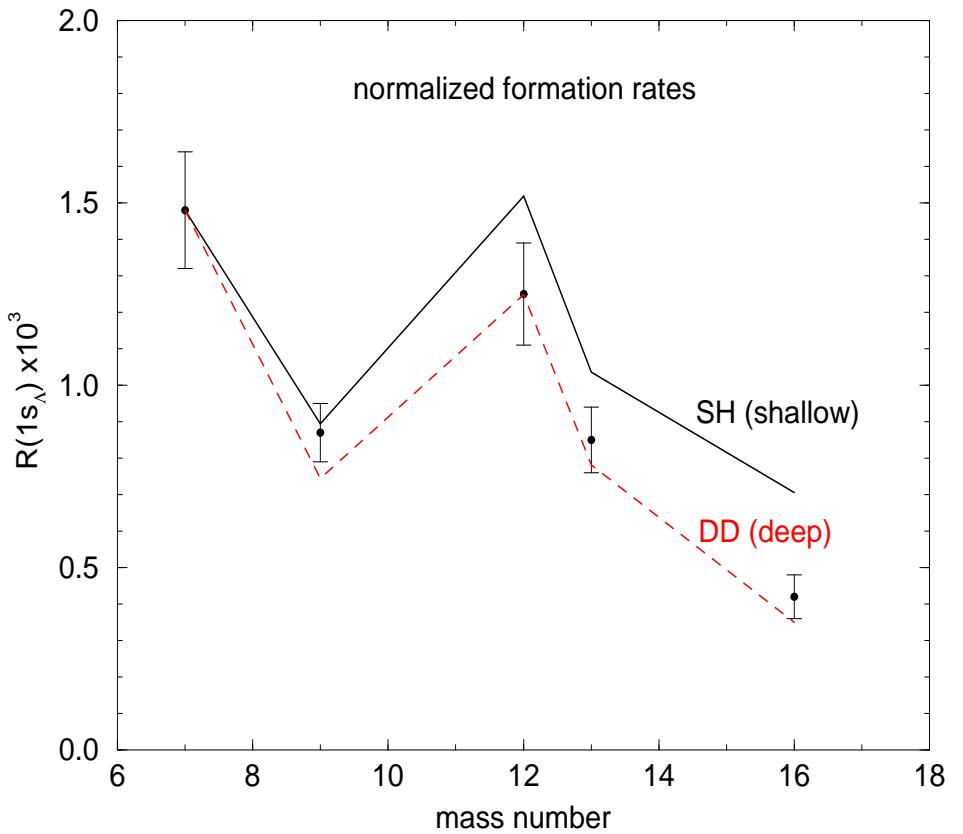
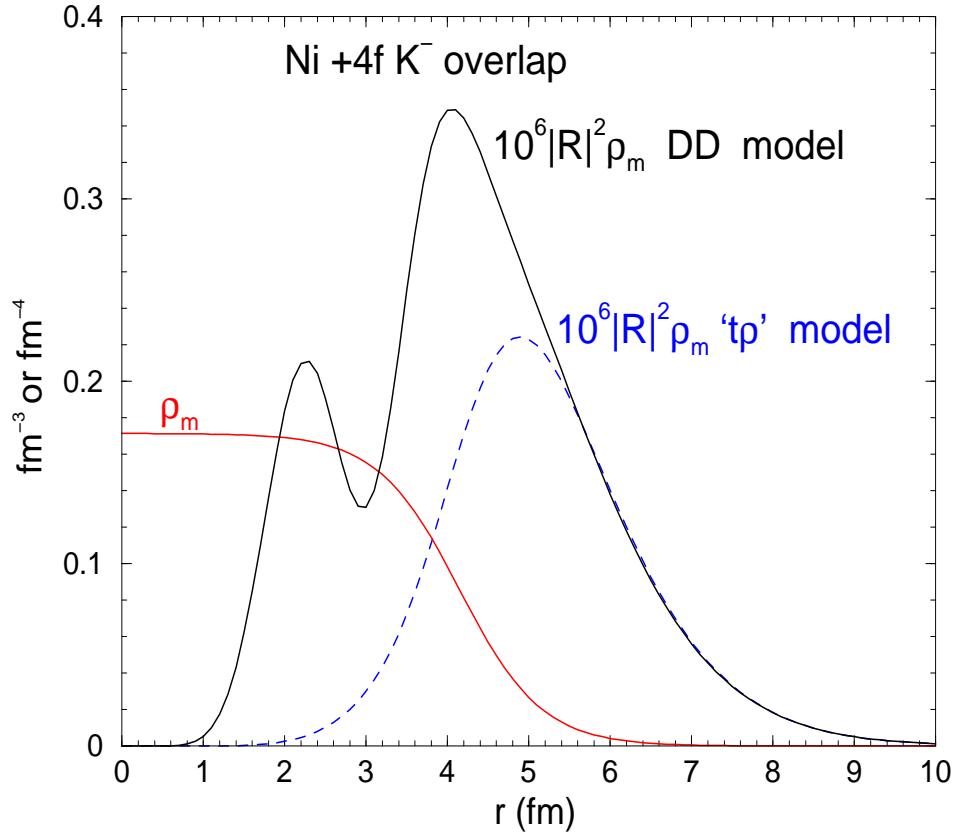
D. Gazda, J. Mareš, NPA 881 (2012) 159

- NLO30 is a chirally motivated coupled channel separable model with in-medium versions [A. Cieplý, J. Smejkal, NPA 881 (2012) 115]
- $\Gamma_K$  due only to  $K^- N \rightarrow \pi Y$  (no  $K^- NN \rightarrow YN$ ) decay modes
- Self consistency: deep  $K^-$  levels are narrower than shallow ones

What do  $K^-$  atoms tell us?



$K_{\text{atom}}^-$  widths across the periodic table in model F (deep pot.)  
 Lowest  $\chi^2$  phenom. model,  $\chi^2 = 84$  per 65 data points,  
 J. Mareš, E. Friedman, A. Gal, NPA 770 (2006) 84.



Left:  $K^-$ -Ni 4f atomic wavefunction overlap with nuclear density for deep potential, revealing a nuclear  $\ell = 3$  quasibound state.

Right: FINUDA  $1s_\Lambda$  formation rates in  $K^-_{stop}$  capture in nuclei [Cieplý-Friedman-Gal-Krejčířík, PLB 698 (2011) 226].

Deep  $K^-$  nuclear potential is favored.

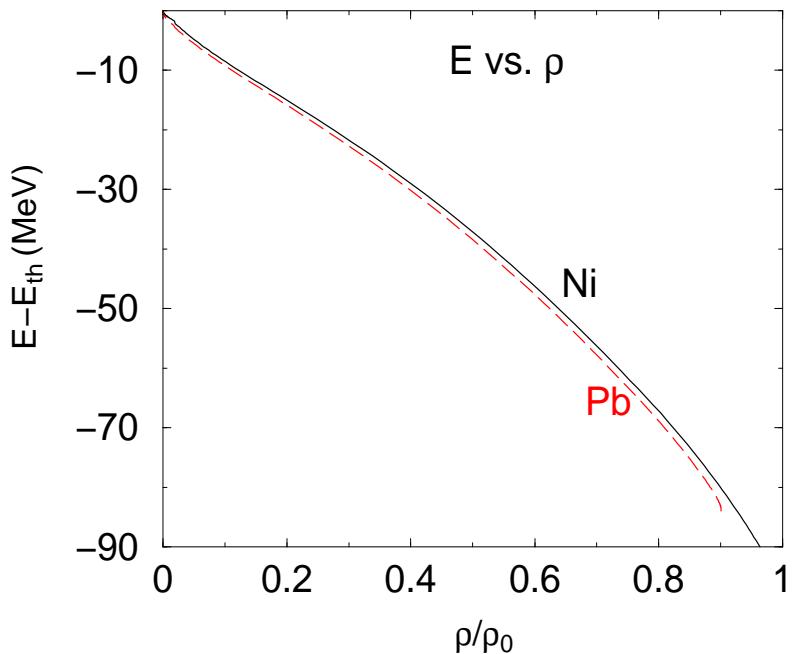
## Self-consistency requirement imposed in recent $K^-$ atom calculations

[Cieplý-Friedman-Gal-Gazda-Mareš, PLB 702 (2011) 402]:

$$\sqrt{s_{K^-N}} \rightarrow E_{\text{th}} - B_N - B_K - \xi_N \frac{p_N^2}{2m_N} - \xi_K \frac{p_K^2}{2m_K}$$

$$\xi_{N(K)} = \frac{m_{N(K)}}{(m_N + m_K)}$$

$$\frac{p_K^2}{2m_K} \sim -V_{K^-} \approx 100 \text{ MeV}$$

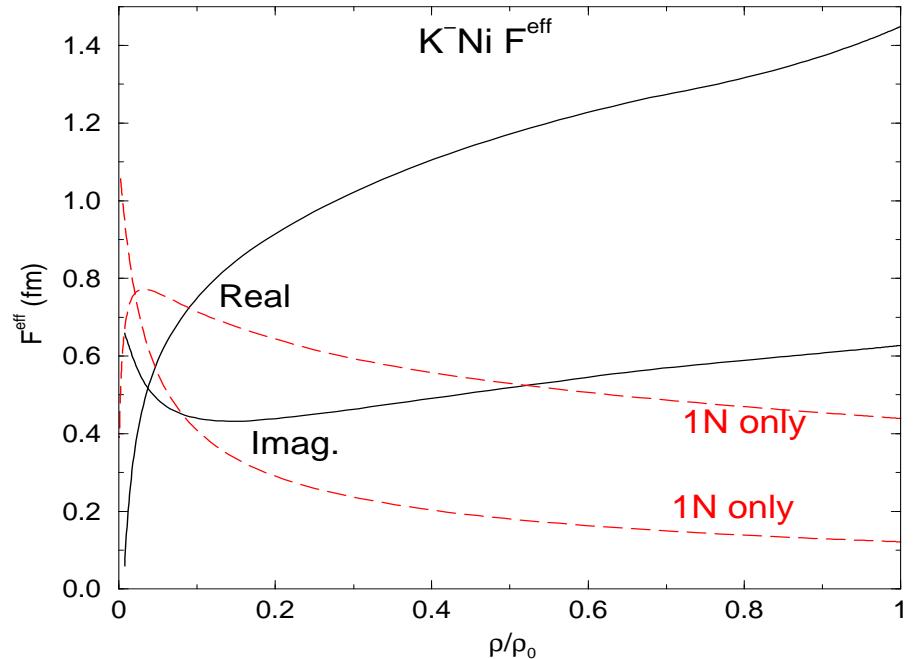
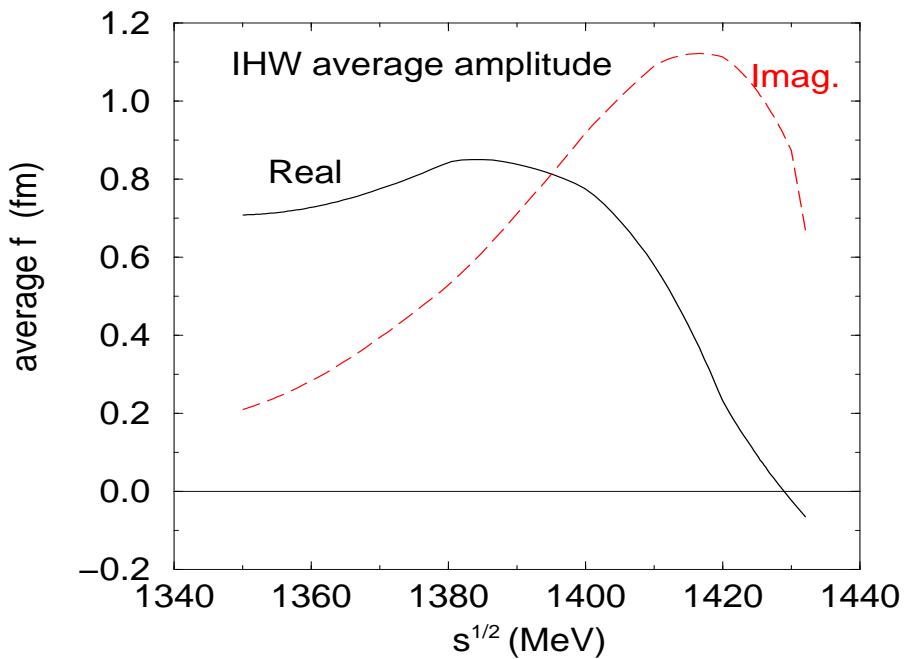


$K^-$  is not at rest!

Friedman-Gal, NPA 899 (2013) 60

$K^-N$  subthreshold energy *vs*  
nuclear density in  $K^-$  atoms.

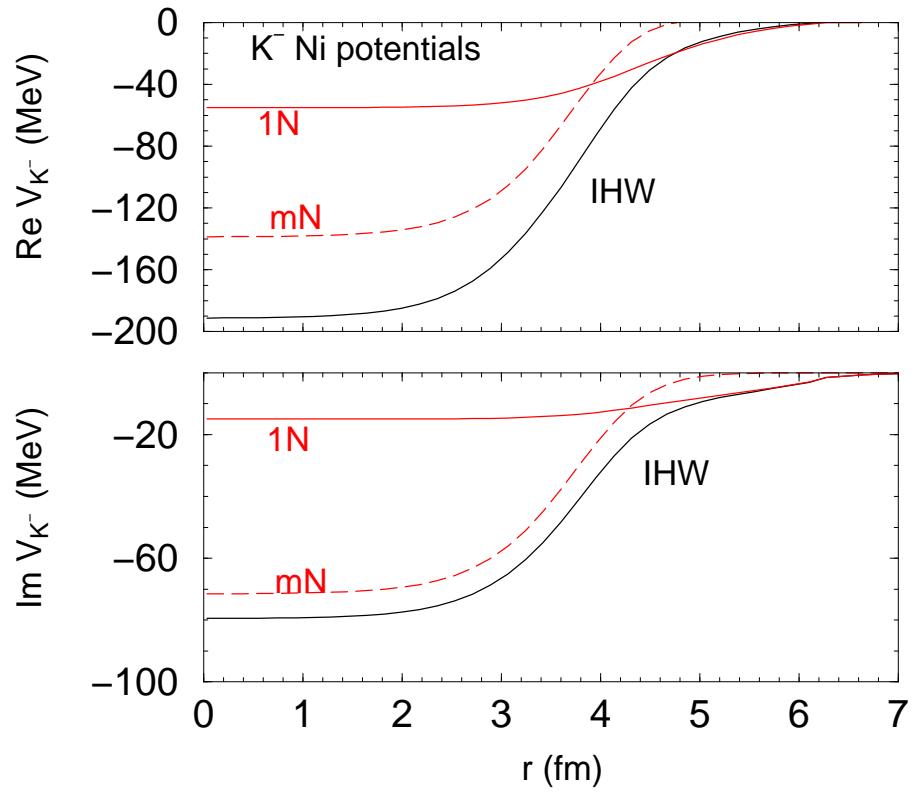
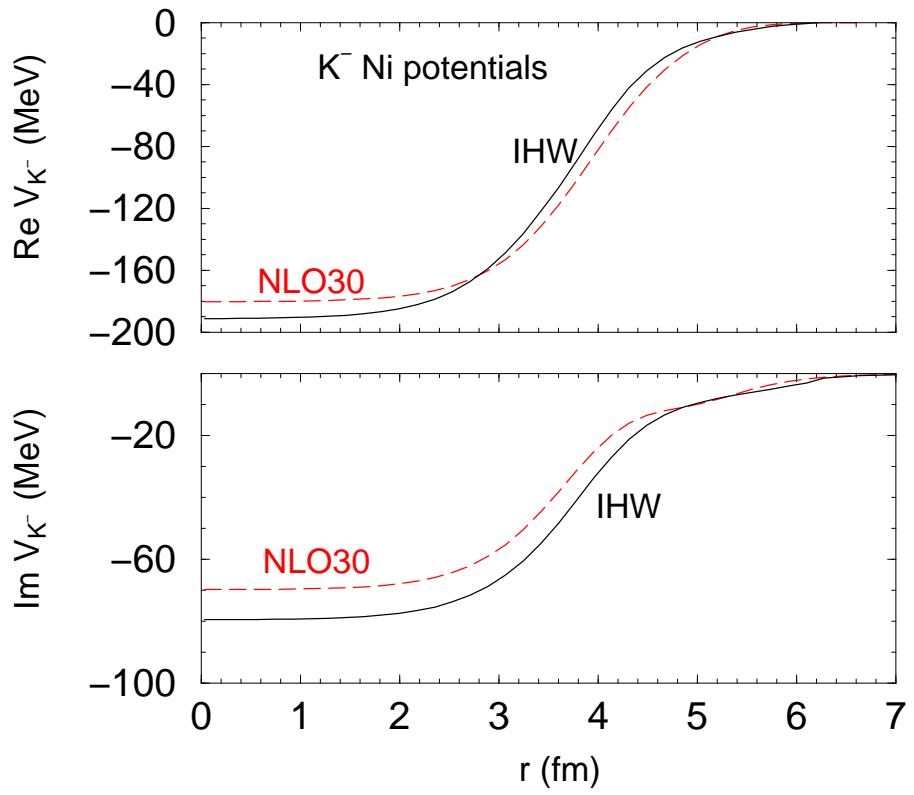
**A dominant in-medium effect**



Left: IHW free-space input  $f_{K^-N}$

Right: atomic-fit output  $\mathcal{F}_{\text{tot}}^{\text{eff}}$

- Subthreshold energy shift is applied self consistently to in-medium 1N amplitude plus  $(2+\dots)N$  phenomenological amplitude.
- Multiple-scattering inclusion of in-medium correlations.
- $K^-$ -atom best-fit:  $\chi^2/N_{\text{data}} = 118/65$  [Friedman-Gal, NPA 899 (2013) 60].

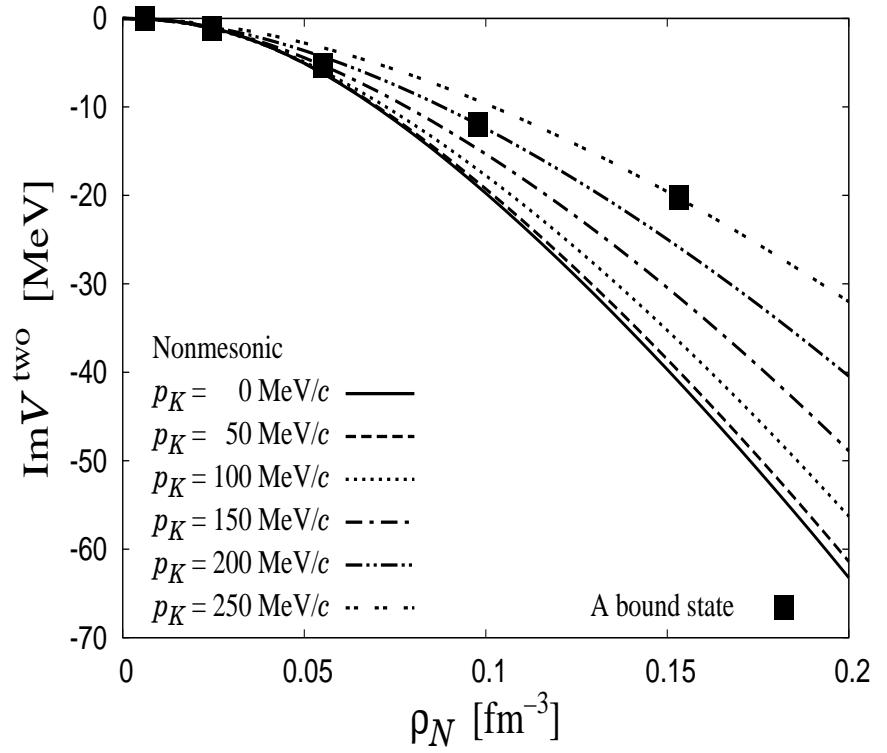
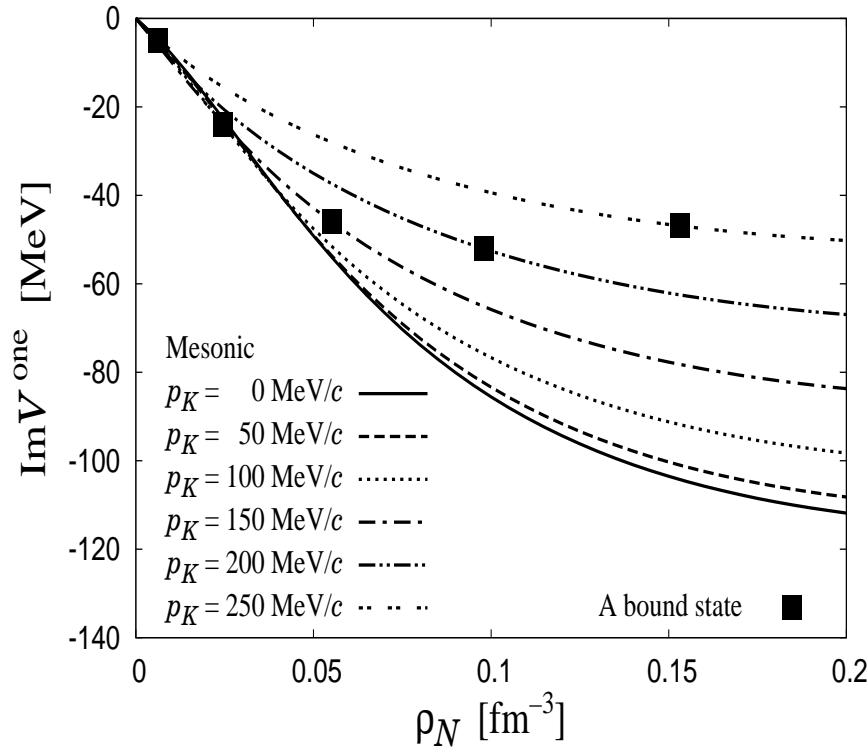


Kaonic-atom best-fit  $V_{K^-}$  for Ni & its non-additive breakdown into in-medium **1N** and phenomenological **m(any)N** contributions.

NLO30: A. Cieply, J. Smejkal, NPA **881** (2012) 115 (in-medium).

IHW: Y. Ikeda, T. Hyodo, W. Weise, NPA **881** (2012) 98.

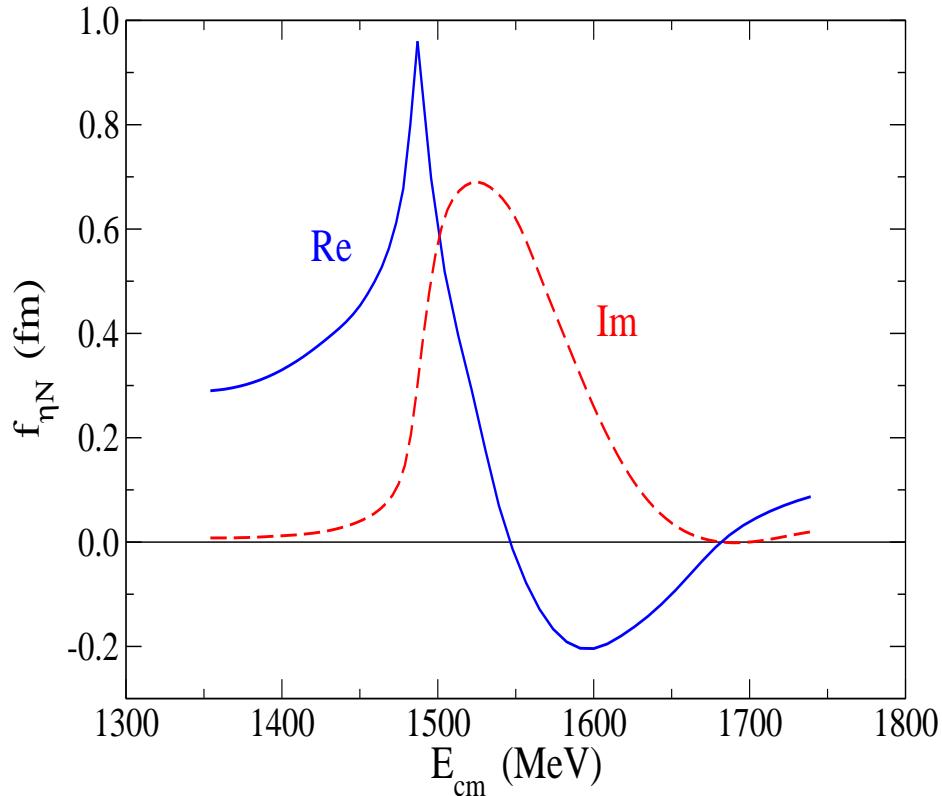
Figures taken from Friedman-Gal, NPA **899** (2013) 60.



$K^-$  nuclear 1N (left) and 2N (right) absorptive potentials,  
 both calculated in a chiral unitary approach [PRC 86 (2012) 065205]  
 by Sekihara, Yamagata-Sekihara, Jido, Kanada-En'yo.  
 Note: empirical 25% 2N:1N BR is reached at too high density.

$\eta$  nuclear quasibound states

## $f_{\eta N}(\sqrt{s})$ from $K$ -matrix & $N^*(1535)$ chiral models



$a_{\eta N}$ model dependence					
$a$ (fm)	M1	M2	GW	GR	CS
Re	0.22	0.38	0.96	0.26	0.67
Im	0.24	0.20	0.26	0.24	0.20

Mai et al. PRD 86 (2012) 094033  
 Green-Wycech PRC 71 (2005) 014001  
 Garcia-Recio et al. PLB 550 (2002) 47  
 Cieply-Smejkal arXiv:1308.4300, NPA

- Re  $a$  varies between 0.2 to 1.0 fm; Im  $a$  stable 0.2–0.3 fm.
- M1, M2, GW free-space models; GR, CS in-medium.
- In-medium: energy dependence, Pauli blocking, self-energies.

# In-medium $\eta N$ amplitudes

Friedman-Gal-Mareš, PLB 725 (2013) 334

Cieplý-Friedman-Gal-Mareš, in preparation

- KG equation and self-energies:

$$[\nabla^2 + \tilde{\omega}_\eta^2 - m_\eta^2 - \Pi_\eta(\omega_\eta, \rho)] \psi = 0$$

$$\tilde{\omega}_\eta = \omega_\eta - i\Gamma_\eta/2, \quad \omega_\eta = m_\eta - B_\eta$$

$$\Pi_\eta(\omega_\eta, \rho) \equiv 2\omega_\eta V_\eta = -4\pi \frac{\sqrt{s}}{m_N} f_{\eta N}(\sqrt{s}, \rho) \rho$$

- Pauli blocking (Waas-Rho-Weise NPA 617 (1997) 449):

$$f_{\eta N}^{\text{WRW}}(\sqrt{s}, \rho) = \frac{f_{\eta N}(\sqrt{s})}{1 + \xi(\rho)(\sqrt{s}/m_N)f_{\eta N}(\sqrt{s})\rho}, \quad \xi(\rho) = \frac{9\pi}{4p_F^2}$$

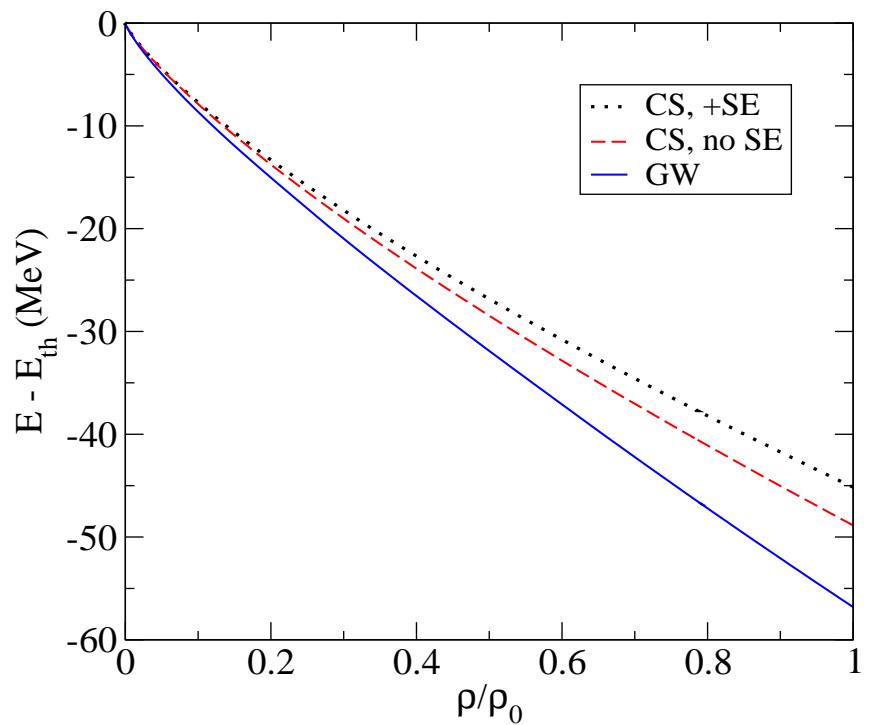
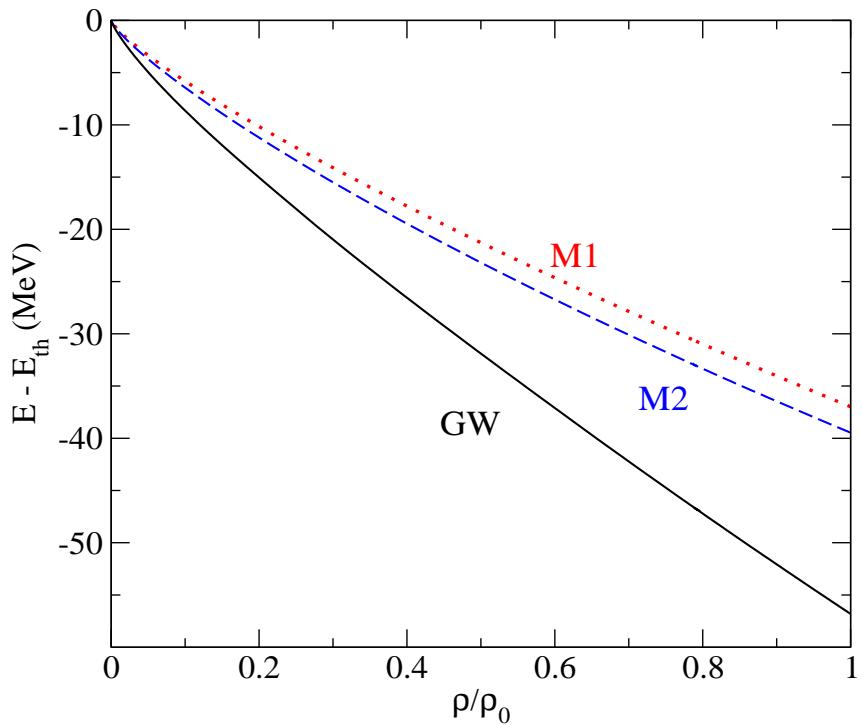
- $N^*(1535) \Rightarrow$  energy dependent  $f_{\eta N}(\sqrt{s})$ .

In medium  $\Rightarrow$  go subthreshold:  $\delta\sqrt{s} = \sqrt{s} - \sqrt{s_{\text{th}}}$

$$\delta\sqrt{s} \approx -B_N \frac{\rho}{\bar{\rho}} - \xi_N B_\eta \frac{\rho}{\rho_0} - \xi_N T_N \left(\frac{\rho}{\rho_0}\right)^{2/3} + \xi_\eta \text{Re } V_\eta(\sqrt{s}, \rho)$$

**Self-consistency relationship between  $\delta\sqrt{s}$  &  $\rho$**

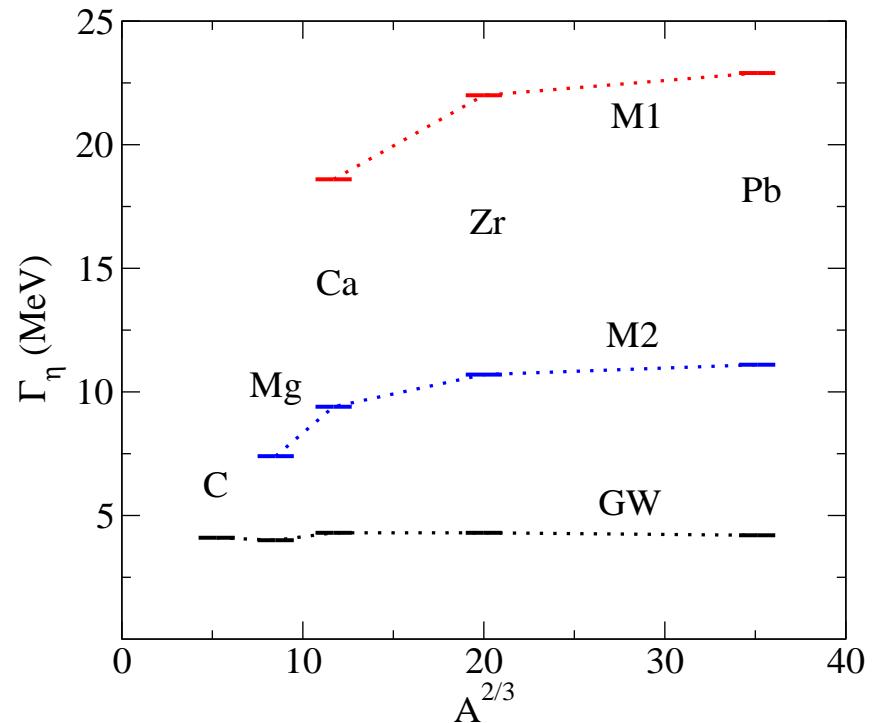
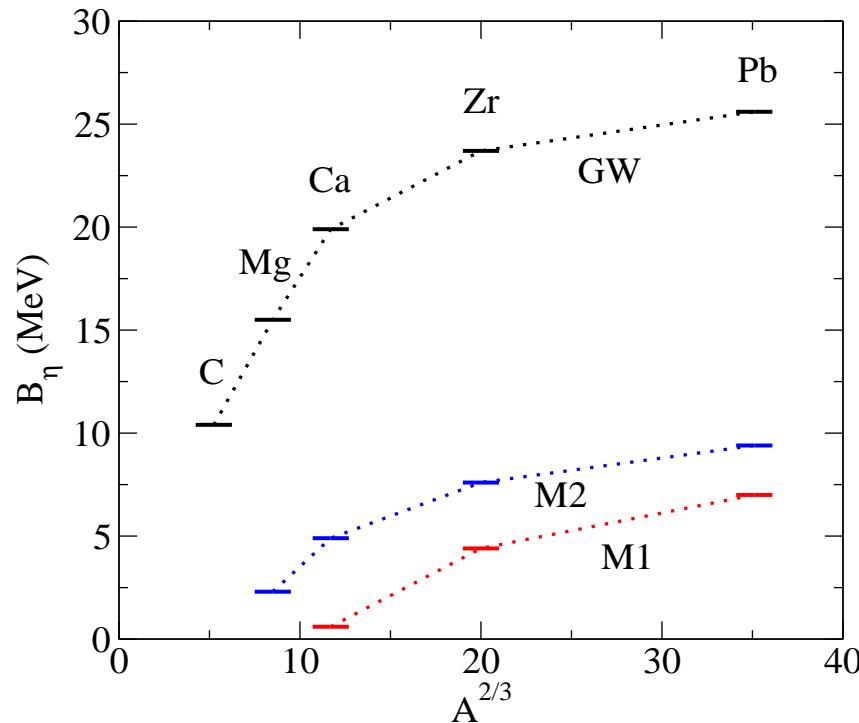
# Self-consistency relationship



$\delta\sqrt{s}$  vs.  $\rho$  for  $1s_\eta$  bound state in Ca using in-medium  $f_{\eta N}$

- 40–60 MeV subthreshold energy shifts at nuclear matter density  $\rho_0$ , larger than shifting down by  $B_\eta$  (GR) or by 30 MeV (Haider-Liu)
- Larger  $\text{Re } a_{\eta N} \Rightarrow$  larger  $\delta\sqrt{s} = E - E_{\text{th}}$

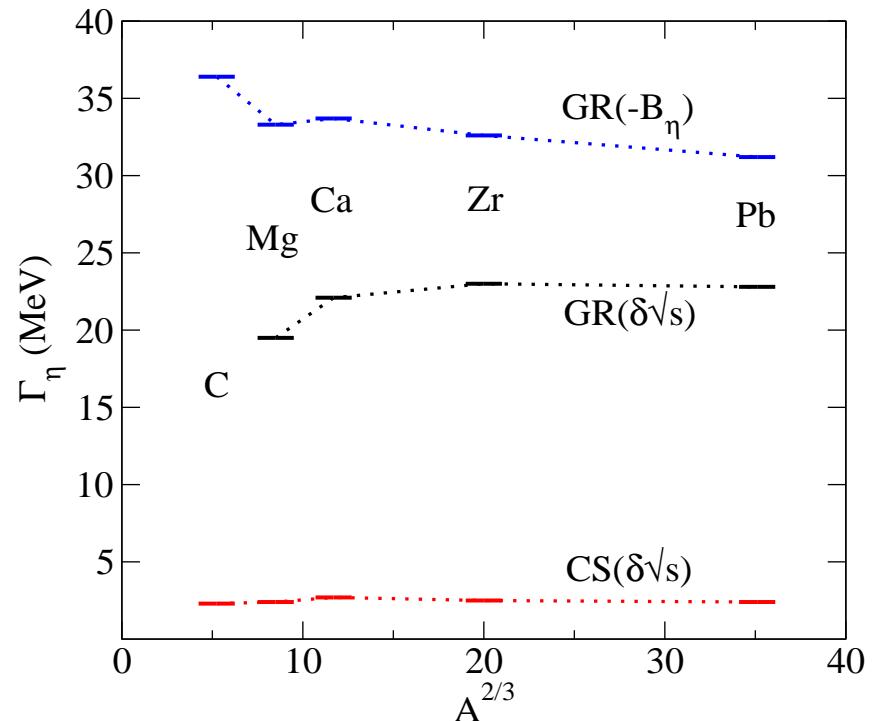
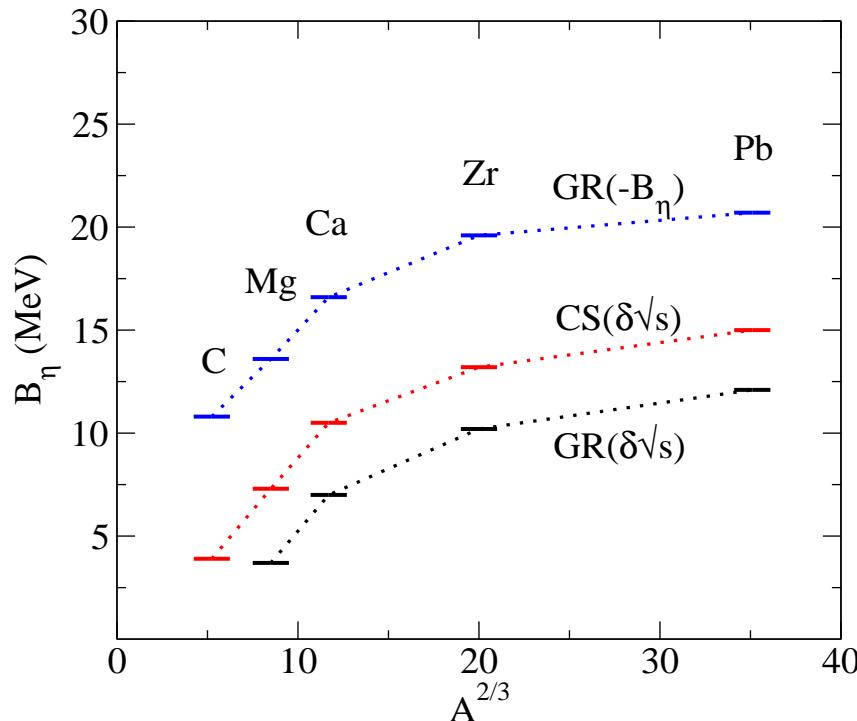
# Model dependence I



Binding energy and width of  $1s_\eta$  bound states across the periodic table using WRW Pauli-blocked  $f_{\eta N}$

- Larger  $\text{Re } a_{\eta N} \Rightarrow$  larger  $B_\eta$
- Widths are unrelated to  $\text{Im } a_{\eta N}$

# Model dependence II

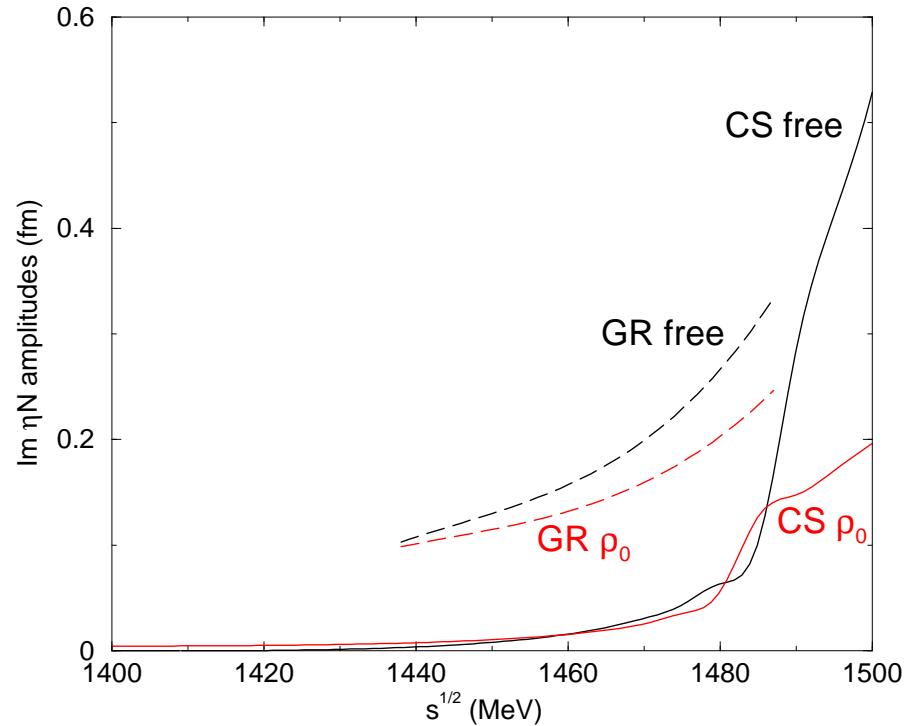
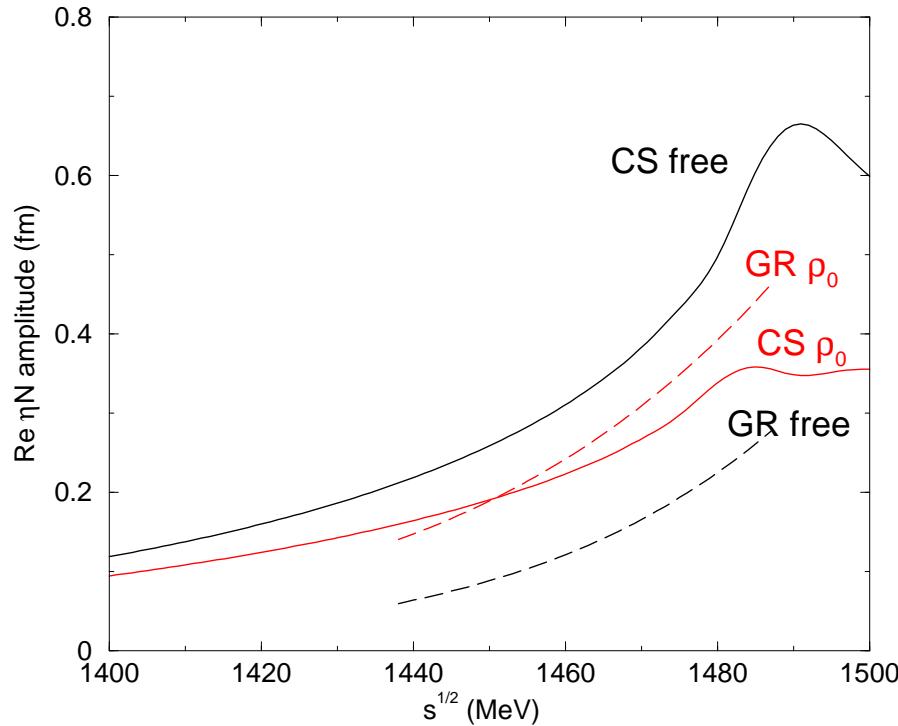


Sensitivity of calculated  $B_{1s_\eta}$  and  $\Gamma_{1s_\eta}$  to version of self-consistency

- $\delta\sqrt{s}$  recipe reduces both  $B_{1s_\eta}$  and  $\Gamma_{1s_\eta}$  w.r.t.  $-B_{1s_\eta}$  recipe
- GR's widths are too large to resolve  $\eta$  bound states

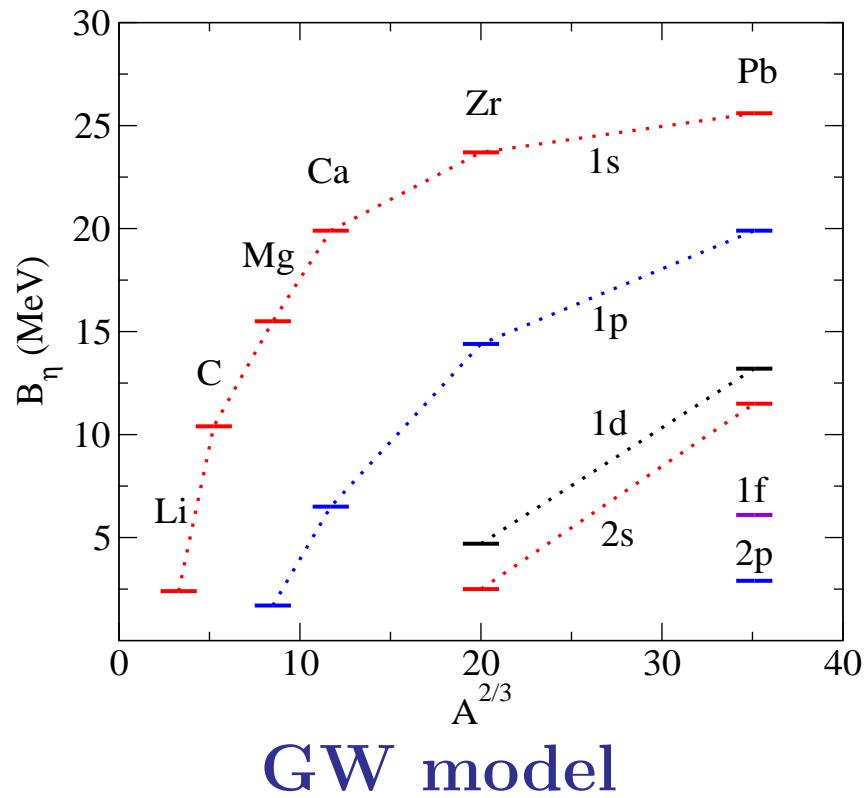
Why  $\Gamma_\eta(\text{GR}) \gg \Gamma_\eta(\text{CS})$  for similar  $\text{Im } a_{\eta N}$ ?

# Energy dependence of free-space & in-medium amplitudes

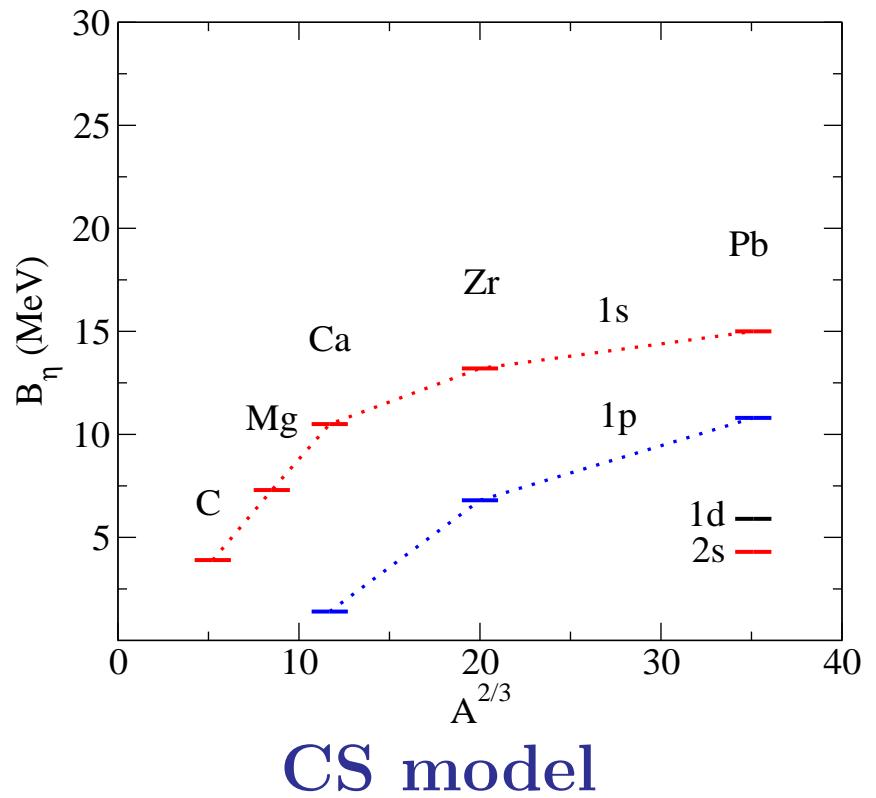


- Subthreshold  $\text{Re } f_{\eta N}$  similar in both in-medium models in spite of large free-space difference at threshold
- Subthreshold  $\text{Im } f_{\eta N}$  differ widely, which explains the huge difference between  $\Gamma_\eta(\text{GR})$  and  $\Gamma_\eta(\text{CS})$

# Model predictions for small widths



GW model



CS model

- more theoretical work is needed to figure out what makes subthreshold values of  $\text{Im } f_{\eta N}$  sufficiently small to generate small widths.

# Summary

- Large widths,  $\Gamma_{\bar{K}} > 50$  MeV, expected for single- $\bar{K}$  quasibound nuclear states. Focus on light systems.  
Searches for  $K^- pp$  are underway in GSI and J-PARC.
- Major issues: (i) how deep is  $\bar{K}$  nuclear spectrum?  
(ii) how big is  $\Gamma(\bar{K}NN \rightarrow YN)$  w.r.t.  $\Gamma(\bar{K}N \rightarrow \pi Y)$ ?
- Subthreshold behavior of  $f_{\eta N}$  is crucial in studies of  $\eta$ -nuclear bound states to decide whether (i) such states exist, (ii) can they be resolved (widths?), and (iii) which nuclear targets and reactions to try?

Thanks to my collaborators N. Barnea, A. Cieplý,  
E. Friedman, D. Gazda, J. Mareš