

Numerical studies of JIMWLK (and glasma) evolution

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Outline

- ▶ CGC, Glasma, JIMWLK evolution
- ▶ JIMWLK equation in Langevin form
- ▶ Running coupling in JIMWLK T.L., H. Mäntysaari EPJC 2013
- ▶ JIMWLK as initial condition for CYM T.L., PLB 2011
- ▶ Wilson loop in glasma with JIMWLK (or MV) initial conditions
Dumitru, T.L., Nara, work in progress: more in Adrian's talk

JIMWLK [“gym-walk”] Jalilian-Marian, Iancu, McLerran, Weigert, Leonidov, Kovner

Gluon saturation, Glass and Glasma

Small x : the hadron/nucleus
wavefunction is characterized by
saturation scale $Q_s \gg \Lambda_{\text{QCD}}$.

Gluon saturation, Glass and Glasma

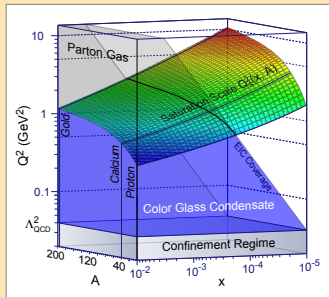
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$\mathbf{p} \sim Q_s$: strong fields $A_\mu \sim 1/g$

- ▶ occupation numbers $\sim 1/\alpha_s$
- ▶ classical field approximation.
- ▶ small α_s , but nonperturbative



Gluon saturation, Glass and Glasma

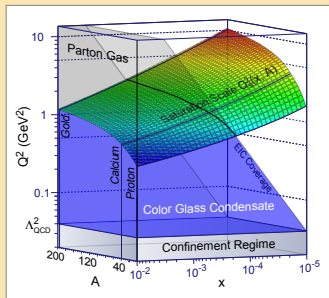
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CGC: Effective theory for wavefunction of nucleus

- ▶ Large x = color charge ρ , **probability** distribution $W_Y[\rho]$
- ▶ Small x = classical gluon field A_μ + quantum flucts.

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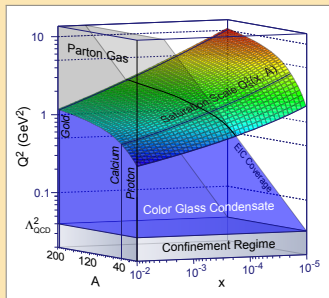
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Glasma field configuration of two colliding sheets of CGC.

Wilson line

Classical color field described as Wilson line

$$U(\mathbf{x}) = P \exp \left\{ ig \int dx^- A_{\text{cov}}^+(\mathbf{x}, x^-) \right\} \in \text{SU}(3)$$

Relation to color charge

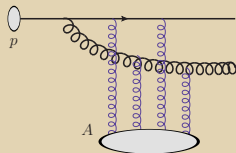
$$\nabla^2 A_{\text{cov}}^+(\mathbf{x}, x^-) = -g\rho(\mathbf{x}, x^-)$$

Example of usage: forward pA

- ▶ Quark from p (large x pdf) , radiate gluon(s)
- ▶ Eikonal propagation \implies Wilson lines $U(\mathbf{x})$

Need target expectation values of operators:

$$\text{Tr } U(\mathbf{x})U^\dagger(\mathbf{y}) \quad \text{Tr } U(\mathbf{x})U^\dagger(\mathbf{y})U(\mathbf{u})U^\dagger(\mathbf{v}) \quad \dots$$



See Bowen's talk

JIMWLK evolution

Classical color field described as Wilson line

$$U(\mathbf{x}) = P \exp \left\{ ig \int dx^- A^+(\mathbf{x}, x^-) \right\} \in \text{SU}(3)$$

- ▶ Energy dependent **probability** distribution $W_y[U]$ ($y \sim \ln \sqrt{s}$)
- ▶ Energy/rapidity dependence of $W_y[U]$ given by JIMWLK renormalization group equation

$$\partial_y W_y[U(\mathbf{x})] = \mathcal{H} W_y[U(\mathbf{x})]$$

JIMWLK Hamiltonian: (fixed coupling)

$$\mathcal{H} \equiv \frac{1}{2} \alpha_s \int_{\mathbf{xyz}} \frac{\delta}{\delta A_c^+(\mathbf{y})} \mathbf{e}^{ba}(\mathbf{x}, \mathbf{z}) \cdot \mathbf{e}^{ca}(\mathbf{y}, \mathbf{z}) \frac{\delta}{\delta A_b^+(\mathbf{x})},$$

$$\mathbf{e}^{ba}(\mathbf{x}, \mathbf{z}) = \frac{1}{\sqrt{4\pi^3}} \frac{\mathbf{x} - \mathbf{z}}{(\mathbf{x} - \mathbf{z})^2} (1 - U^\dagger(\mathbf{x})U(\mathbf{z}))^{ba}$$

Fokker-Planck and Langevin

Textbook example: two descriptions of Brownian motion

- ▶ 1-d diffusion eq. (\supset F-P eq.)

$$\partial_t P(x, t) = D \partial_x^2 P(x, t)$$

- ▶ $P(x, t)$ = probability for particle to be at location x at time t .
- ▶ For $x = 0$ at $t = 0$ solution:

$$P(x, t) = \frac{1}{\sqrt{4\pi Dt}} \exp \left\{ -\frac{x^2}{4Dt} \right\}$$

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- ▶ Langevin equation:

$$x(t) = \sqrt{2D} \eta(t)$$

$$\langle \eta(t) \eta(t') \rangle = \delta(t - t')$$

- ▶ $\langle x(t) \rangle = 0$

$$\langle x^2(t) \rangle = 2Dt$$

$$\langle x(t)x(t') \rangle = 2D \min(t, t')$$

\implies same as F.-P.

1d Brownian motion to JIMWLK

- ▶ Replace $x \implies U(\mathbf{x})$ and $t \implies y$.
- ▶ Constant $D \implies$ nonlinearity (U -dependence) in kernel
- ▶ $(N_c^2 - 1)N_{\perp}^2$ -dimensional nonlinear diffusion equation.
(N_{\perp}^2 = number of lattice points in transverse plane.)

Langevin formulation

Fokker-Planck \implies Langevin in JIMWLK Blaizot, Iancu, Weigert 2002

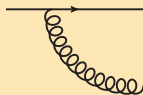
Original Langevin form: only right derivative ($\xi_z^{b,i}$ is noise)

$$U_{\mathbf{x}}(y + dy) = U_{\mathbf{x}}(y) \exp \left\{ it^a \int_{\mathbf{z}} \varepsilon_{\mathbf{x},\mathbf{z}}^{ab,i} \xi_z^{b,i} \sqrt{dy} + \sigma_{\mathbf{x}}^a dy \right\}.$$

Simpler, equivalent (for $dy \rightarrow 0$) form T.L., H.M.

$$U_{\mathbf{x}}(y + dy) = \exp \left\{ -i \frac{\sqrt{\alpha_s} dy}{\pi} \int_{\mathbf{z}} \mathbf{K}_{\mathbf{x}-\mathbf{z}} \cdot (U_{\mathbf{z}} \xi_{\mathbf{z}} U_{\mathbf{z}}^\dagger) \right\} \\ \times U_{\mathbf{x}}(y) \exp \left\{ i \frac{\sqrt{\alpha_s} dy}{\pi} \int_{\mathbf{z}} \mathbf{K}_{\mathbf{x}-\mathbf{z}} \cdot \xi_{\mathbf{z}} \right\},$$

$$K_{\mathbf{x}-\mathbf{z}}^i = \frac{(\mathbf{x} - \mathbf{z})^i}{(\mathbf{x} - \mathbf{z})^2}$$



$$i = x, y$$

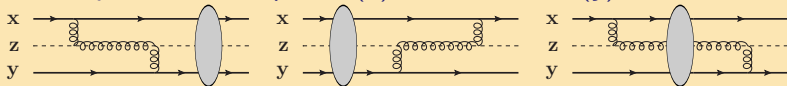
Fixed α_s noise: $\langle \xi_{\mathbf{x}}(y_m)_i^a \xi_{\mathbf{y}}(y_n)_j^b \rangle = \alpha_s \delta^{ab} \delta^{ij} \delta_{\mathbf{xy}}^{(2)} \delta_{mn}$, $\xi = \xi^a t^a$

Multiply from left **and** right \implies remove deterministic term

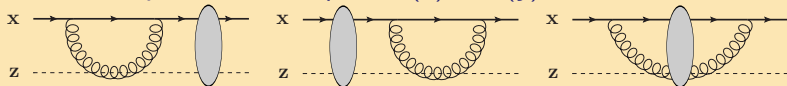
Interpreting JIMWLK: derive BK

$$U_{\mathbf{x}}(y + dy) = e^{-i\sqrt{\frac{\alpha_s dy}{\pi}} \int_z \mathbf{K}_{\mathbf{x}-z} \cdot (U_z \xi_z U_z^\dagger)} U_{\mathbf{x}} e^{i\sqrt{\frac{\alpha_s dy}{\pi}} \int_z \mathbf{K}_{\mathbf{x}-z} \cdot \xi_z},$$

- ▶ At $dy \rightarrow 0$ develop to $\mathcal{O}(\xi^2)$ and take expectation values.
- ▶ BK **Balitsky-Kovchegov** is equation for **dipole** $\hat{D}_{\mathbf{x},\mathbf{y}} = \text{Tr } U^\dagger(\mathbf{x})U(\mathbf{y})/N_c$
- ▶ Contract ξ 's from timestep of $U^\dagger(\mathbf{x})$ with one from $U(\mathbf{y})$: **real terms**



- ▶ Contract two ξ 's from timestep of $U^\dagger(\mathbf{x})$ or $U(\mathbf{y})$: **virtual terms**



- ▶ **Result**

$$\partial_y \hat{D}_{\mathbf{x},\mathbf{y}}(y) = \frac{\alpha_s N_c}{2\pi^2} \int_z \left(\mathbf{K}_{\mathbf{x}-z}^2 + \mathbf{K}_{\mathbf{y}-z}^2 - 2\mathbf{K}_{\mathbf{x}-z} \cdot \mathbf{K}_{\mathbf{y}-z} \right) \left[\hat{D}_{\mathbf{x},z} \hat{D}_{z,\mathbf{y}} - \hat{D}_{\mathbf{x},\mathbf{y}} \right].$$

Scale of running α_s in JIMWLK ?

BK for $\hat{D}_{\mathbf{x},\mathbf{y}}(y)$ describes dipole splitting $\mathbf{x} - \mathbf{y} \longrightarrow \mathbf{x} - \mathbf{z} ; \mathbf{z} - \mathbf{y}$

- ▶ α_s given by parent $\mathbf{x} - \mathbf{y}$: easy in BK, but funny in JIMWLK: Langevin is only for one Wilson line
- ▶ Daughter (scale in \mathbf{K}): easy to implement as $\sqrt{\alpha_s}$, but why?

$$\sqrt{\alpha_s} \mathbf{K}_{\mathbf{x}-\mathbf{z}} \rightarrow \sqrt{\alpha_s(\mathbf{x} - \mathbf{z})} \mathbf{K}_{\mathbf{x}-\mathbf{z}}$$

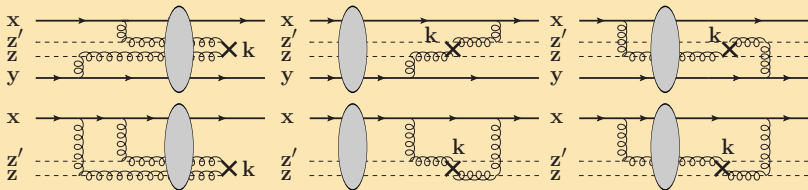
- ▶ Used in BK: combinations of these two Balitsky, Kovchegov, Weigert
- ▶ Suggestion T.L., H.Mäntysaari 2012 : α_s at k_T of radiated gluon.
- ▶ Easily implemented by new momentum space noise correlator

$$\langle \xi_{\mathbf{x}}(m)_i^a \xi_{\mathbf{y}}(n)_j^b \rangle \sim \alpha_s \delta_{\mathbf{xy}}^{(2)} = \alpha_s \int \frac{d^2 \mathbf{k}}{(2\pi)^2} e^{i\mathbf{k} \cdot (\mathbf{x}-\mathbf{y})}$$
$$\implies \int \frac{d^2 \mathbf{k}}{(2\pi)^2} e^{i\mathbf{k} \cdot (\mathbf{x}-\mathbf{y})} \alpha_s(\mathbf{k})$$

Reinterpreting JIMWLK

$$U_{\mathbf{x}}(y + dy) = \exp \left\{ -i \frac{\sqrt{dy}}{\pi} \int_{\mathbf{z}} \mathbf{K}_{\mathbf{x}-\mathbf{z}} \cdot (U_{\mathbf{z}} \xi_{\mathbf{z}} U_{\mathbf{z}}^\dagger) \right\} \\ \times U_{\mathbf{x}}(y) \exp \left\{ i \frac{\sqrt{dy}}{\pi} \int_{\mathbf{z}'} \mathbf{K}_{\mathbf{x}-\mathbf{z}'} \cdot \xi_{\mathbf{z}'} \right\},$$

$$\langle \xi_{\mathbf{x}}(m)_i^a \xi_{\mathbf{y}}(n)_j^b \rangle \sim \int \frac{d^2 \mathbf{k}}{(2\pi)^2} e^{i\mathbf{k} \cdot (\mathbf{x}-\mathbf{y})} \alpha_S(\mathbf{k}) \equiv \tilde{\alpha}_{\mathbf{x}-\mathbf{y}}$$



- ▶ Breaks time-reversal-symmetry: choose scale as momentum of gluon either before or after the target
- ▶ Two gluon coordinates instead of one

Recovering BK

- ▶ Equation for dipole now involves higher point functions:

$$\partial_y \hat{D} = \frac{N_c}{2\pi^2} \int_{\mathbf{u}, \mathbf{v}} \tilde{\alpha}_{\mathbf{u}-\mathbf{v}} \left(\mathbf{K}_{\mathbf{x}-\mathbf{u}} \cdot \mathbf{K}_{\mathbf{x}-\mathbf{v}} + \mathbf{K}_{\mathbf{y}-\mathbf{u}} \cdot \mathbf{K}_{\mathbf{y}-\mathbf{v}} - 2\mathbf{K}_{\mathbf{x}-\mathbf{u}} \cdot \mathbf{K}_{\mathbf{y}-\mathbf{v}} \right) \\ \times \frac{1}{2} \left[\hat{D}_{\mathbf{x},\mathbf{u}} \hat{D}_{\mathbf{u},\mathbf{y}} + \hat{D}_{\mathbf{x},\mathbf{v}} \hat{D}_{\mathbf{v},\mathbf{y}} - \hat{D}_{\mathbf{x},\mathbf{y}} - \hat{D}_{\mathbf{v},\mathbf{u}} \hat{Q}_{\mathbf{x},\mathbf{v},\mathbf{u},\mathbf{y}} \right],$$

- ▶ But recall that α_s is a slowly varying function of the scale:

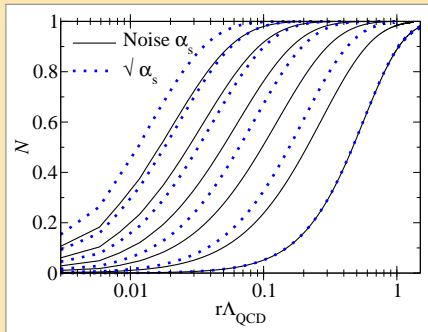
$$\tilde{\alpha}_{\mathbf{x}-\mathbf{y}} \equiv \int \frac{d^2 \mathbf{k}}{(2\pi)^2} e^{i\mathbf{k} \cdot (\mathbf{x}-\mathbf{y})} \alpha_s(\mathbf{k}) \sim \alpha_s \delta^2(\mathbf{x}-\mathbf{y})$$

$\Rightarrow \mathbf{u} \approx \mathbf{v}$ and structure simplifies to BK:

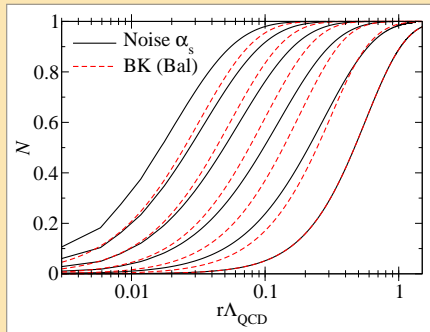
$$\frac{1}{2} \left[\hat{D}_{\mathbf{x},\mathbf{u}} \hat{D}_{\mathbf{u},\mathbf{y}} + \hat{D}_{\mathbf{x},\mathbf{v}} \hat{D}_{\mathbf{v},\mathbf{y}} - \hat{D}_{\mathbf{x},\mathbf{y}} - \hat{D}_{\mathbf{v},\mathbf{u}} \hat{Q}_{\mathbf{x},\mathbf{v},\mathbf{u},\mathbf{y}} \right] \approx \hat{D}_{\mathbf{x},\mathbf{u}} \hat{D}_{\mathbf{u},\mathbf{y}} - \hat{D}_{\mathbf{x},\mathbf{y}}$$

- ▶ Parametrically dominant length scale in coupling is “smallest dipole”, just like in Balitsky prescription for BK.

Comparison BK/JIMWLK



Evolution with our prescription is slower than with $\sqrt{\alpha_s}$. This is good, data favors slower evolution

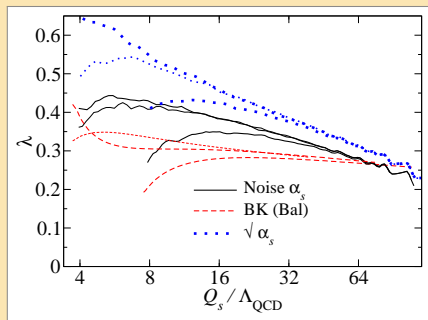


But this is still faster than with Balitsky prescription in BK (Although parametrically dominant scales are the same.)

Note: rcBK fits to HERA data need to take $\Lambda_{\text{QCD}} \approx 50\text{MeV}$ to make evolution slow enough.

Evolution speed

$$\lambda \equiv \frac{d \ln Q_s^2}{dy}$$

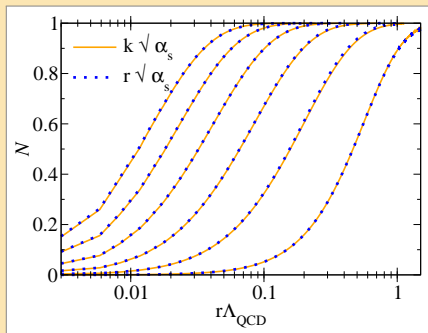


At very IR scales also dependence on how the Landau pole is regulated (different line shapes)

Side note: scale in coordinate vs momentum space

If running coupling depends only on scale in \mathbf{K} ($\sqrt{\alpha_s}$ -prescription), can use either coordinate or momentum space:

$$\sqrt{\alpha_s(\mathbf{x})} \frac{\mathbf{x}}{\mathbf{x}^2} \quad \text{vs.} \quad \sqrt{\alpha_s(\mathbf{k})} \frac{\mathbf{k}}{\mathbf{k}^2}$$



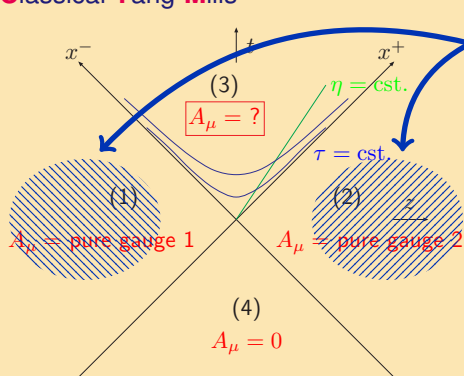
Numerically verified identification [Kovchegov, Weigert](#) (for this kernel!)

$$\ln \frac{\mathbf{k}^2}{\Lambda_{\text{QCD}}^2} \sim \ln \frac{4e^{-2\gamma_E}}{r^2\Lambda_{\text{QCD}}^2}$$

Gluon fields in AA collision

Classical Yang-Mills

2 pure gauges

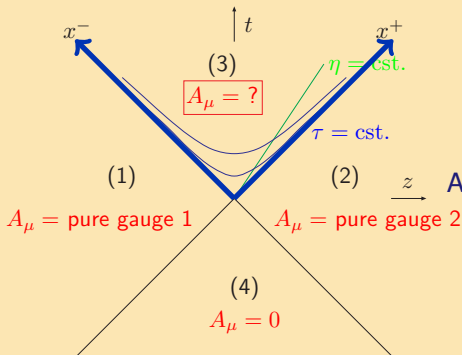


$$A_{(1,2)}^i = \frac{i}{g} U_{(1,2)}(\mathbf{x}) \partial_i U_{(1,2)}^\dagger(\mathbf{x})$$

$$U_{(1,2)}(\mathbf{x}) = P e^{ig \int dx^- \frac{\rho(\mathbf{x}, x^-)}{\nabla^2}}$$

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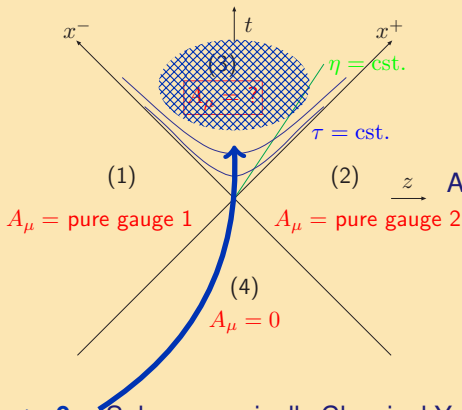
At $\tau = 0$:

$$A^i \Big|_{\tau=0} = A_{(1)}^i + A_{(2)}^i$$

$$A^\eta \Big|_{\tau=0} = \frac{ig}{2} [A_{(1)}^i, A_{(2)}^i]$$

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$\tau > 0$ Solve numerically Classical Yang-Mills **CYM** equations.
This is the **glasma** field \implies Then average over ρ .

Gluons with $p \sim Q_s$ — strings of size $R \sim 1/Q_s$

Gluon spectrum in the glasma

T.L., *Phys.Lett.* **B703** (2011) 325 ; 1st direct use of JIMWLK in CYM calculation

Q_s is only dominant scale

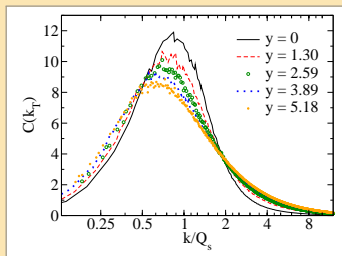
Parametrically gluon spectrum $\frac{dN_g}{dy d^2\mathbf{x} d^2\mathbf{p}} = \frac{1}{\alpha_s} f\left(\frac{p}{Q_s}\right)$

Gluon spectrum in the glasma

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ugd for 1 nucleus

$$C(\mathbf{k}) = \frac{k^2}{N_c} \text{Tr} \langle U(\mathbf{k}) U^\dagger(\mathbf{k}) \rangle$$

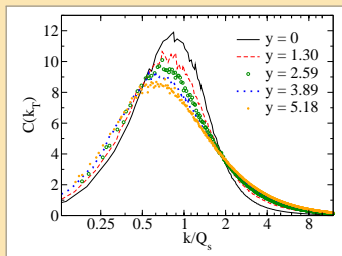
becomes **harder** with evolution.

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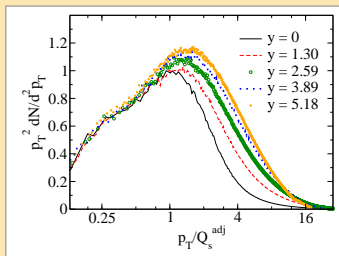
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becomes **harder** with evolution.



Produced gluon spectrum:
harder at higher \sqrt{s}
(Here: midrapidity, $y \equiv \ln \sqrt{s/s_0}$)

Gluon multiplicity and mean p_T

Q_s is only dominant scale

$$\text{Parametrically } \frac{dN_g}{dy d^2\mathbf{x}} = c_N \frac{C_F}{2\pi^2\alpha_s} Q_s^2 \quad \langle p \rangle \sim Q_s$$

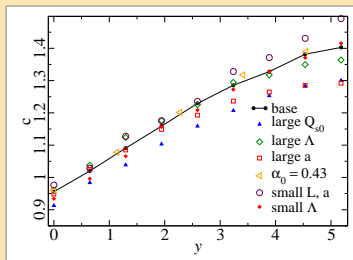
Note: in full CYM total gluon multiplicity is IR finite, no cutoff.

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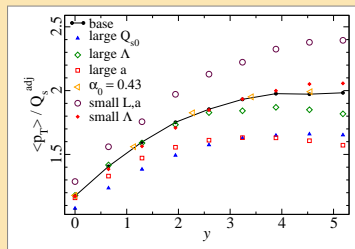
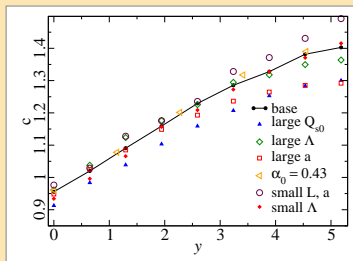
Scaled multiplicity grows with \sqrt{s}
(Midrapidity, $y \equiv \ln \sqrt{s/s_0}$)

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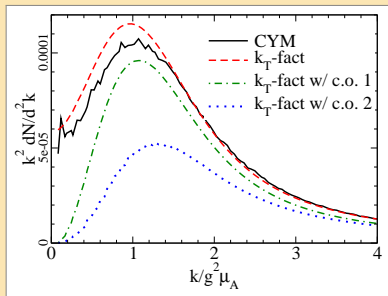
Harder gluon spectrum
 \Rightarrow higher $\langle p \rangle / Q_s$ as scaling
regime sets in.

(Large lattice cutoff effects.)

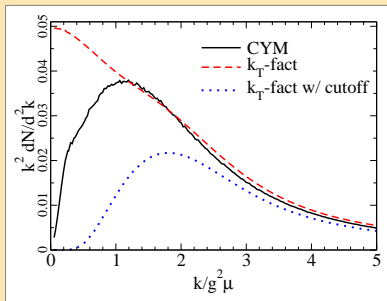
Side note: CYM vs. k_T -factorization

Blaizot, T.L., Mehtar-Tani 2010

$$\frac{dN}{d^2\mathbf{p}dy} = \frac{\#}{\alpha_s} \frac{1}{\mathbf{p}^2} \int_{\mathbf{k}} \left[\theta(\mathbf{p} - \mathbf{k}) \right] \phi_y(\mathbf{k}) \phi_y(\mathbf{p} - \mathbf{k})$$



pA



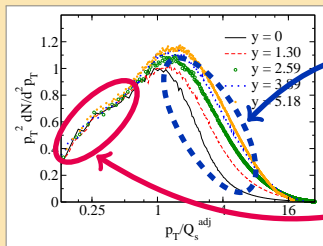
AA

k_T -factorization works only for $p_T \gtrsim Q_s$

- ▶ OK for high- p_T spectra
- ▶ Not for total gluon multiplicity

(Suggested interpretation Levin, 2010 Sudakov suppression factor.)

Universality in the IR spectrum?



- ▶ Scaled gluon spectrum in the UV depends on anomalous dimension \Rightarrow different for MV, JIMWLK
- ▶ IR seems to scale

Gauge inv. probe for $p \lesssim Q_s$

Spatial Wilson loop

$$W(A) = \frac{1}{N_c} \text{Tr} \mathbb{P} \exp \left\{ ig \oint_A dx \cdot \mathbf{A} \right\}$$

2d lattice: transverse links:

$$\uparrow = U_i(\mathbf{x}) = \exp \{ igaA_i \}$$

$$W(A) = \frac{1}{N_c} \text{Tr} \left[\begin{array}{c} \leftarrow \leftarrow \leftarrow \\ \uparrow \uparrow \uparrow \\ \rightarrow \rightarrow \rightarrow \\ \downarrow \downarrow \downarrow \end{array} \right]$$

Measure Wilson loops

Calculation is simple:

- ▶ Construct initial glasma fields at $\tau = 0$ using e.g.
 - ▶ MV model
 - ▶ rcJIMWLK
 - ▶ fcJIMWLK

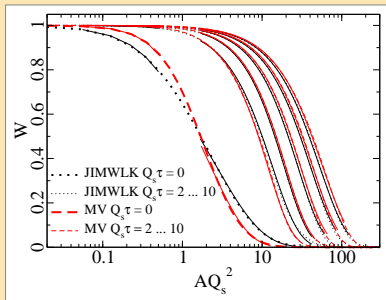
(Try to have same $Q_s a$ to minimize lattice effects)

- ▶ Evolve forward in τ
- ▶ Measure $W(A)$

Behavior in both UV ($AQ_s^2 \lesssim 1$) and IR ($AQ_s^2 \gtrsim 1$) parametrized as

$$W = \exp \{ -C(AQ_s)^{\gamma} \}$$

(Fits restricted to $e^{-2.5} < AQ_s^2 < e^{-0.5}$ (UV) and $e^{0.5} < AQ_s^2 < e^5$ (IR))

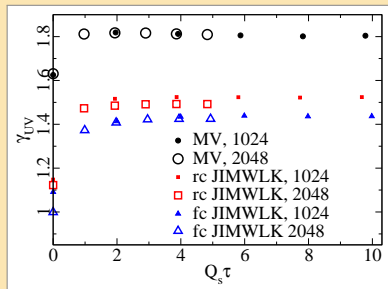
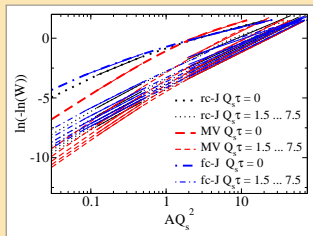


Wilson loop scaling exponents

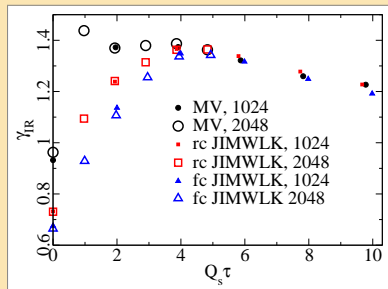
In practice fit logarithm

$$W = \exp \{-C(AQ_s)^\gamma\}$$

$$\iff \ln(-\ln W) = \gamma \ln(AQ_s) + \ln C$$



UV: remembers initial condition



IR: **nontrivial universal behavior**

Magnetic field correlator

Wilson loop measures magnetic flux:

$$W(A) = \frac{1}{N_c} \text{Tr} \mathbb{P} \exp \left\{ ig \oint_A \mathbf{dx} \cdot \mathbf{A} \right\} = \frac{1}{N_c} \text{Tr} \exp \left\{ ig \int d^2 \mathbf{x} B_z(\mathbf{x}) \right\}$$

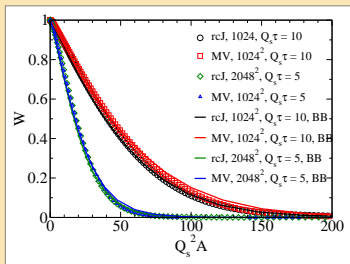
If magnetic field consists of uncorrelated Gaussian domains:

$$\langle W(A) \rangle = \exp \left\{ -\frac{1}{2} \frac{1}{N_c} \text{Tr} \left\langle \left[\int d^2 \mathbf{x} g B_z(\mathbf{x}) \right]^2 \right\rangle \right\}$$

(Connect B :s at different locations with gauge link for gauge invariance)

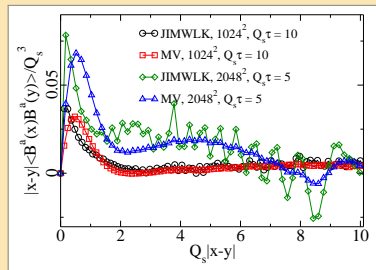
$\Rightarrow W(A)$ related to $\langle B(\mathbf{x})B(\mathbf{y}) \rangle$

Check: compare $W(A)$ with reconstruction from BB -correlator



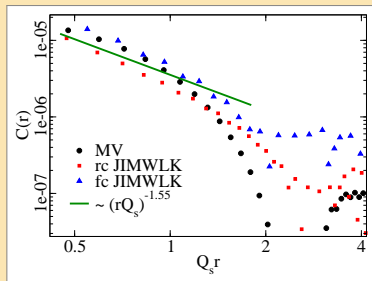
Magnetic field correlator

However: no obvious scaling
seen in BB -correlator



Same on log plot

$$C(|\mathbf{x}-\mathbf{y}|) \equiv \text{Tr} \left\langle [B(\mathbf{x})B(\mathbf{y})]_{\text{gauge link}} \right\rangle$$



Straight line: $\sim (rQ_s)^{-1.55}$.

(For $C(r) \sim (rQ_s)^{-\alpha}$ one would get $\gamma = 2 - \alpha/2 \iff \alpha = 4 - 2\gamma$;

from $W(A)$ measured $\gamma = 1.22$)

Conclusions

- ▶ JIMWLK can actually be applied; for some observables BK is not enough
- ▶ Running coupling, going towards NLO . . . necessary for realistic phenomenology
- ▶ Provides initial state for AA collision in CYM: glasma
- ▶ Universal behavior in the for $p_T \ll Q_s$ seen both in gluon spectrum and spatial Wilson loop.