

# Turbulent Thermalization

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arXiv:1303.5650, arXiv:1311.3005



Pb+Pb @  $\sqrt{s}$  = 2.76 ATeV

2010-11-08 11:30:46

Fill : 1482

Run : 137124

Event : 0x00000000D3BBE693

# Content



Initial state:  
Far from equilibrium



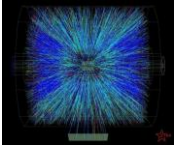
*Non-equilibrium  
dynamics*



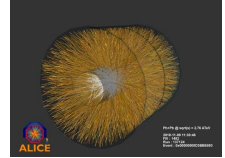
Final state:  
Thermal equilibrium

- I. Universal attractor in longitudinal expanding plasmas:  
Turbulence and `bottom up` thermalization
- II. Limitations of classical-statistical lattice simulations:  
A well understood quantum example

# Heavy Ion Collisions



- fluid-like behavior from very early time on
- very special transport properties, such as small  $\eta/s$



***How is local isotropization/thermalization achieved?***

## Strong correlations:

- a) strong coupling? Gauge-string duality: Heller, Janik, Witaszczyk; Chesler, Yaffe ...

**Sizeable anisotropy even at transition to hydrodynamic regime**

- b) weak coupling but highly occupied? CGC: McLerran, Venugopalan, ...

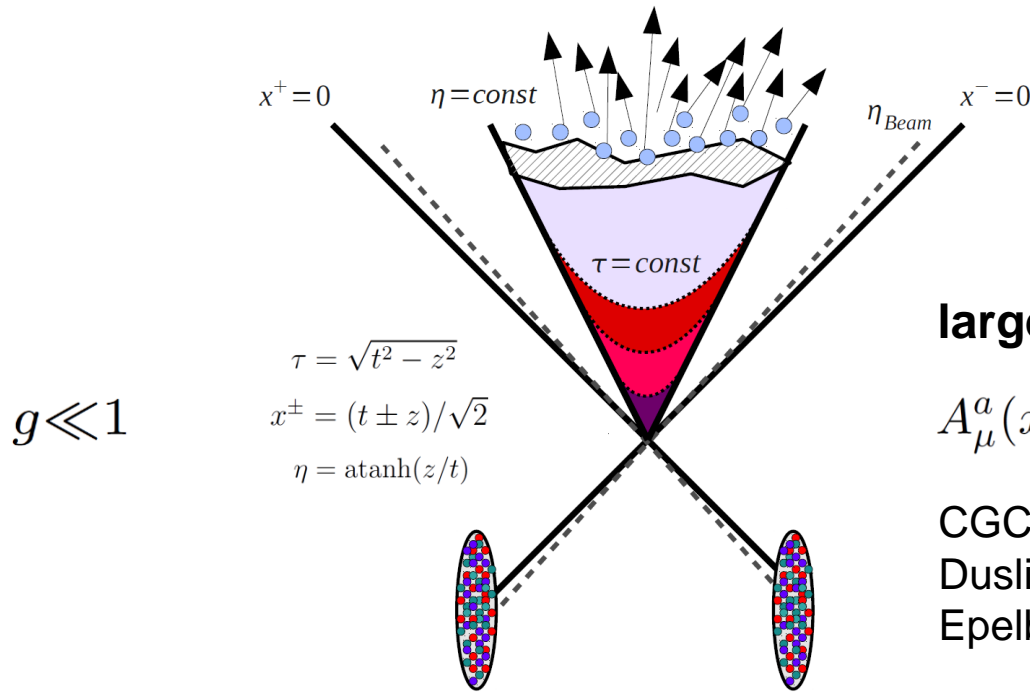
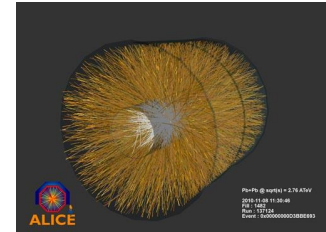
Energy density of gluons with typical momentum  $Q$  (at time  $\sim 1/Q$ )

$$\epsilon \sim \frac{Q^4}{\alpha_s} \quad \text{i.e. 'occupation numbers'} \quad f(p \lesssim Q) \sim \frac{1}{\alpha_s}$$

**Strongly correlated/nonperturbative even for weak coupling  $\alpha_s \ll 1$**

# High-energy/weak-coupling limit

Particle production in the presence of large fields



large initial fields:

$$A_\mu^a(x) = \langle \hat{A}_\mu^a(x) \rangle \sim \mathcal{O}(1/g)$$

CGC: Lappi, McLerran,  
 Dusling, Gelis, Venugopalan,  
 Epelbaum...

small initial (vacuum) fluctuations:

$$F_{\mu\nu}^{ab}(x, y) = \frac{1}{2} \left\langle \left\{ \hat{A}_\mu^a(x), \hat{A}_\nu^b(y) \right\} \right\rangle - A_\mu^a(x) A_\nu^b(y)$$

$$\sim \mathcal{O}(1)$$

→ **plasma instabilities!**

Mrowczynski; Rebhan,  
 Romatschke, Strickland;  
 Arnold, Moore, Yaffe ...

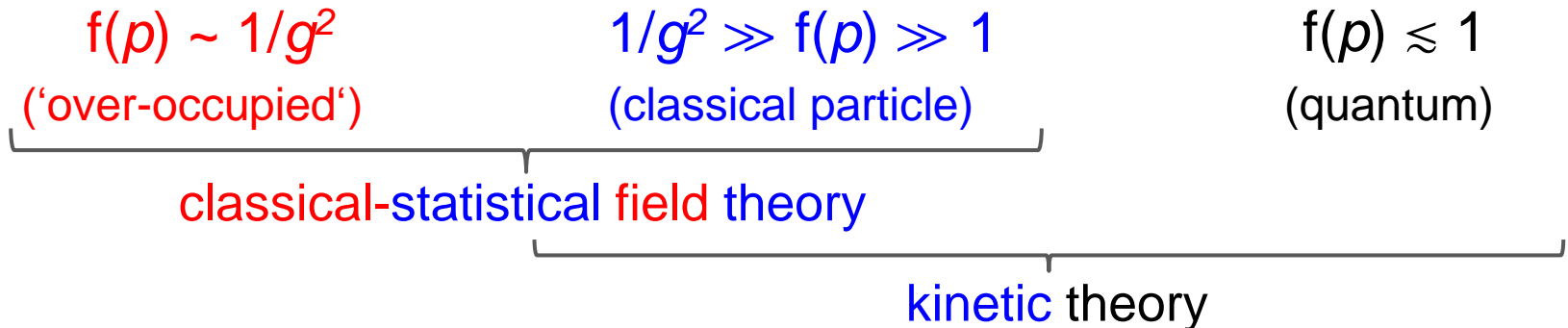
# Different weak-coupling thermalization scenarios

## Kinetic theory (parametric):

- Bottom-up isotropization  
BMSS: Baier, Mueller, Schiff, Son (2001)
- Rescattering due to plasma (Weibel) instabilities  
BD: Boedeker (2005); KM: Kurkela, Moore (2011)
- Transient Bose condensation+fixed anisotropy  
BGLMV: Blaizot, Gelis, Liao, McLerran, Venugopalan (2012)

## *How can we decide from first principles?*

→ for typical occupancies  $1 < f < 1/g^2$  a dual description is feasible:



# Non-linear evolution: Classical-statistical lattice gauge theory

**Wilson action:**

$$S[U] = -\beta_0 \sum_x \sum_i \left\{ \frac{1}{2\text{Tr}\mathbf{1}} (\text{Tr} U_{x,0i} + \text{Tr} U_{x,0i}^{-1}) - 1 \right\} \\ + \beta_s \sum_x \sum_{\substack{i,j \\ i < j}} \left\{ \frac{1}{2\text{Tr}\mathbf{1}} (\text{Tr} U_{x,ij} + \text{Tr} U_{x,ij}^{-1}) - 1 \right\}$$

Plaquette variables  $U_{x,\mu\nu} \equiv U_{x,\mu} U_{x+\hat{\mu},\nu} U_{x+\hat{\nu},\mu}^\dagger U_{x,\nu}^\dagger \approx \exp[-iga^2 F_{\mu\nu}(x)]$

Sampling introduces classical-statistical fluctuations ('loops to all orders')  
→ accurate for characteristic 'large fields/high occupation numbers':

anti-commutator  $\langle\langle A, A \rangle\rangle \gg \langle[A, A]\rangle$  commutator

***ab initio description of non-Abelian real-time dynamics!***

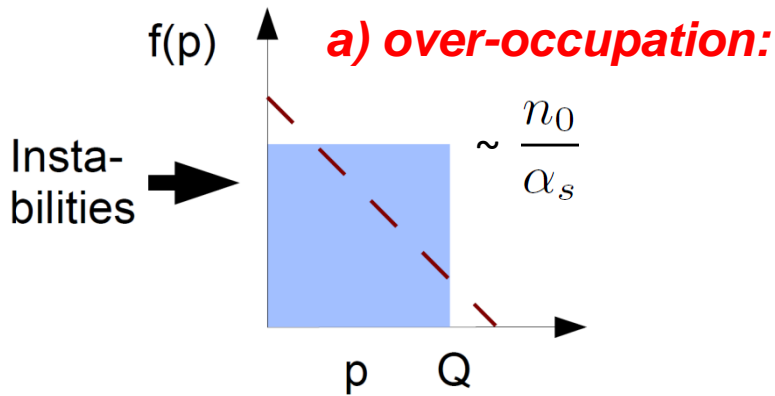
→ instability dynamics, turbulence:

Romatschke, Venugopalan; Berges, Boguslavski, Gelfand, Scheffler, Schlichting, Sexty;  
Kunihiro, Müller, Ohnishi, Schäfer, Takahashi, Yamamoto; Fukushima, Gelis, Eppelbaum;

...

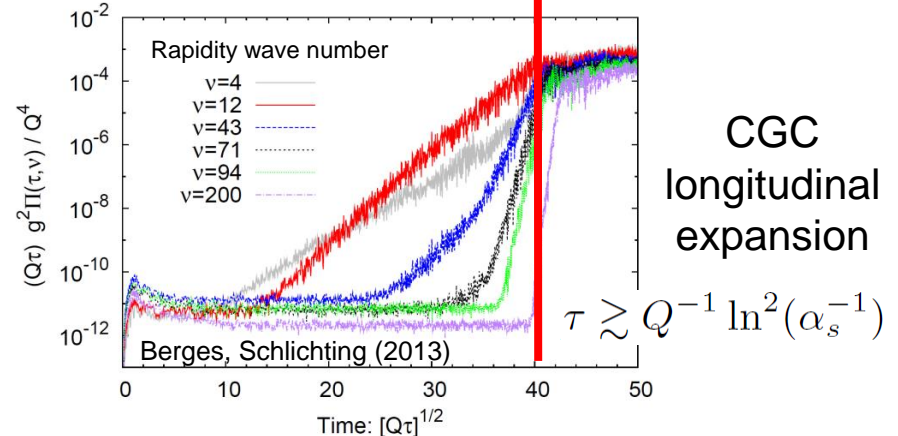
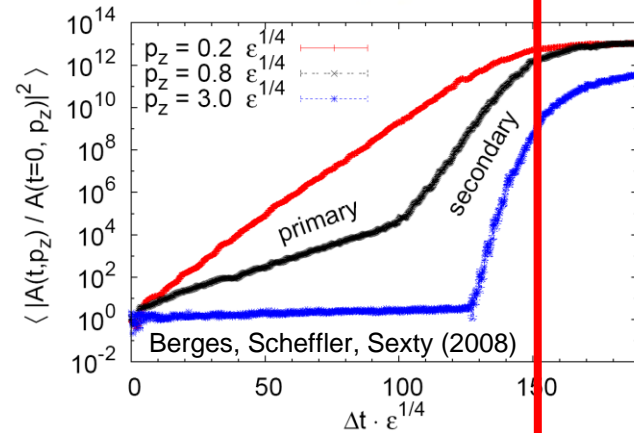
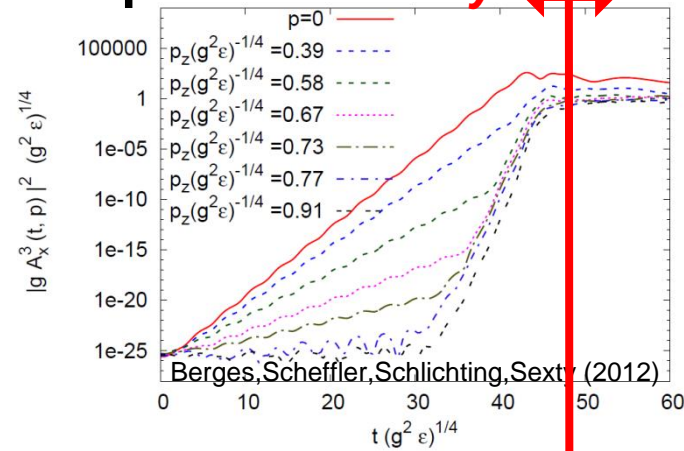
# Plasma instabilities at early times

- quasi-exponential growth of initial fluctuations
- clear separation between instability and scaling (turbulent) regime
- characteristic distributions after instabilities saturate; schematic:

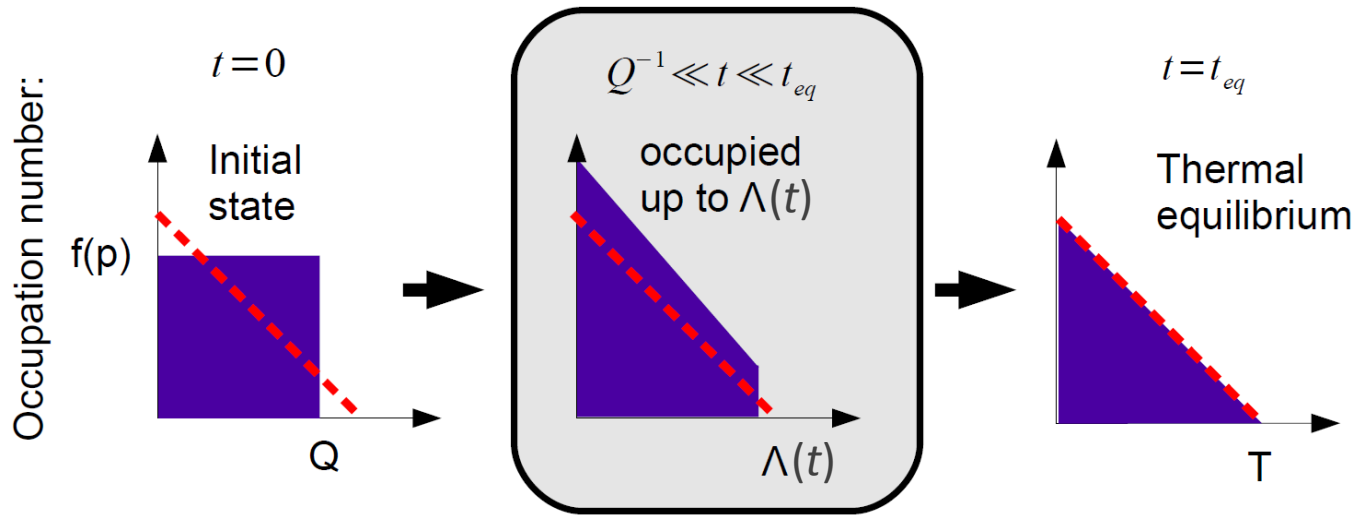


**b) anisotropy:**  $p \approx \sqrt{p_T^2 + (\xi_0 p_z)^2}$

Examples: **instability** ←→ **turbulence**



# Thermalization process – schematic



$$Q\tau_0 \sim \ln^2(\alpha_s^{-1}) \rightarrow$$

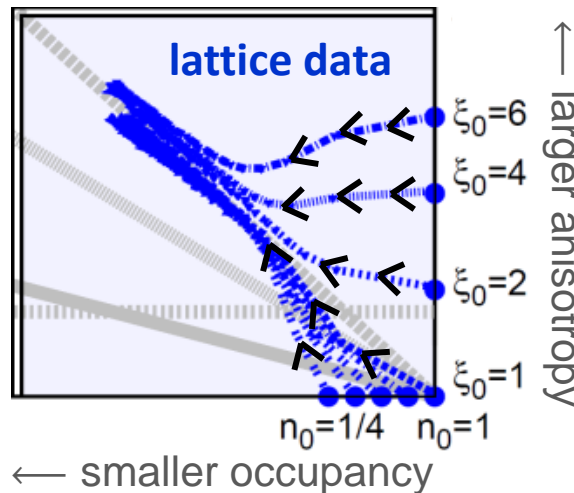
**Universal  
attractor**

**several stages (BMSS):**

$$Q\tau \sim \alpha^{-3/2}, \alpha^{-5/2}, \alpha^{-13/5}$$

attractor (classical)  
isotropization (quantum)  
thermalization (quantum)

different initial occupancies  
& anisotropies lead to  
same universal properties!

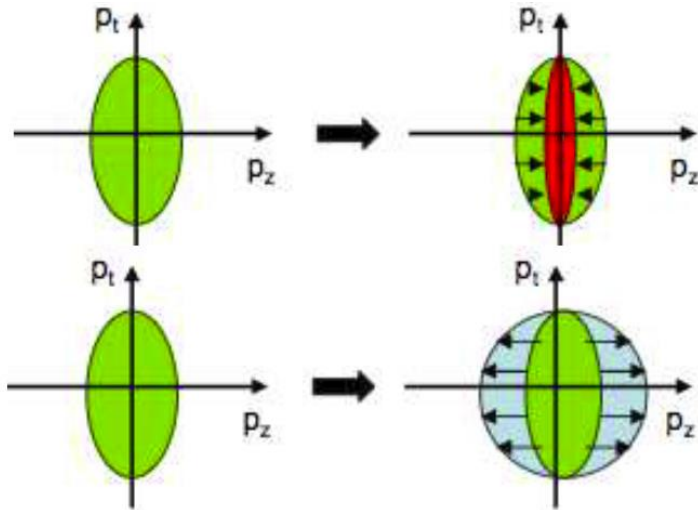


↑ larger anisotropy

← smaller occupancy



# Thermalization of expanding systems



## ***Longitudinal Expansion:***

- Red-shift of longitudinal momenta  $p_z$   
→ increase of anisotropy
- Dilution of the system

## ***Interactions:***

- Isotropize the system

- Redshift and dilution require extremely large lattices
- Will report here on real-time classical-statistical simulations on

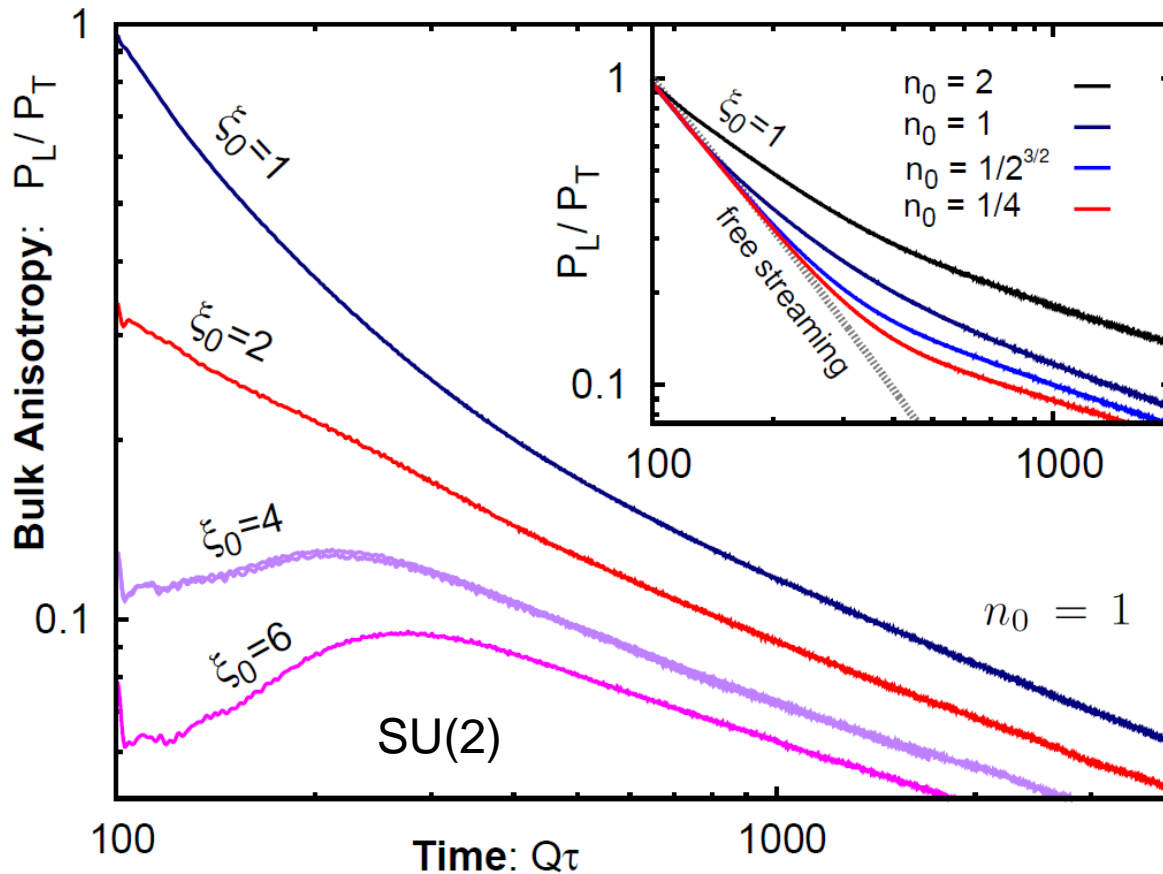
***256<sup>2</sup> × 4096 (!) spatial lattices***

-- fully capture important infrared dynamics

# Longitudinal expanding non-Abelian plasma: Anisotropy

Initial gluon distributions:  $f(p_T, p_z, \tau_0) = \frac{n_0}{2g^2} \Theta \left( Q - \sqrt{p_T^2 + (\xi_0 p_z)^2} \right)$

occupancy parameter  $\rightarrow$   $n_0$   
anisotropy parameter  $\rightarrow$   $\xi_0$

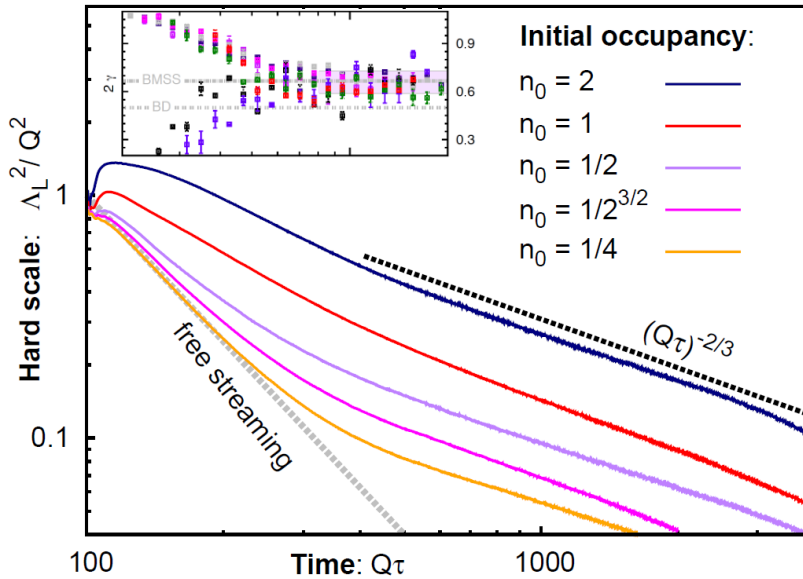


Large initial anisotropy leads to transient isotropy increase ( $\rightarrow$  instabilities).

Smaller initial occupancy leads to transient period of free streaming.

Continues to be **strongly correlated** throughout the entire turbulent stage!

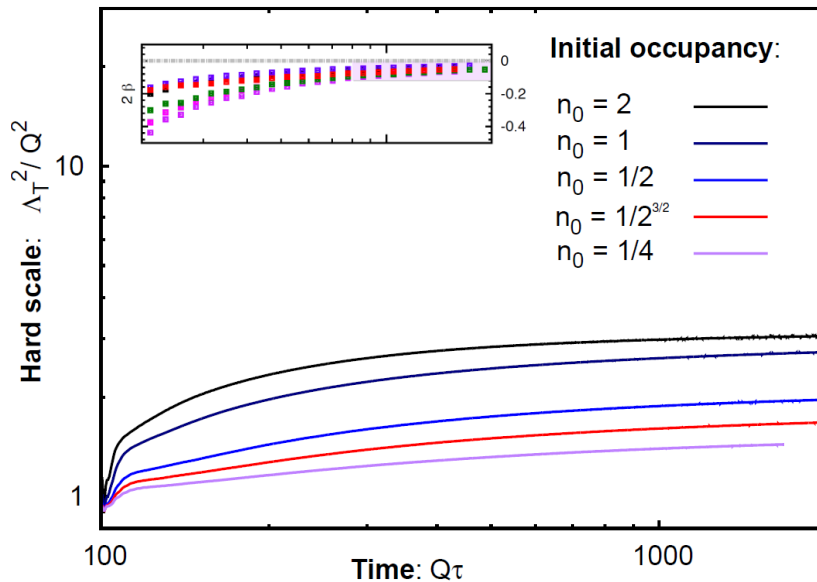
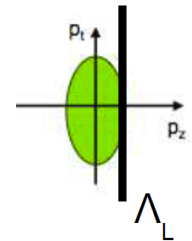
# Gauge invariant hard scales: Universal power laws



- The typical **longitudinal momentum** of hard excitations exhibits a **universal scaling** behavior

$$\Lambda_L^2(\tau) \propto Q^2 (Q\tau)^{-2\gamma}$$

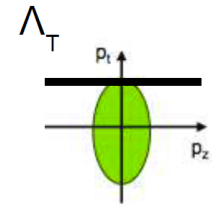
$$2\gamma = 0.67 \pm 0.07$$



- The typical **transverse momentum** of hard excitations remains approximately **constant**

$$\Lambda_T^2(\tau) \propto Q^2 (Q\tau)^{-2\beta}$$

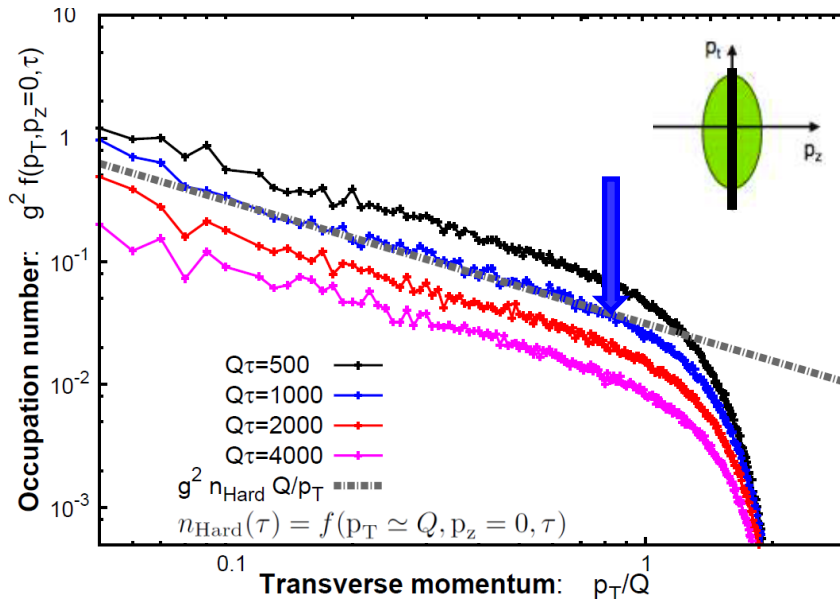
$$2\beta \simeq 0$$



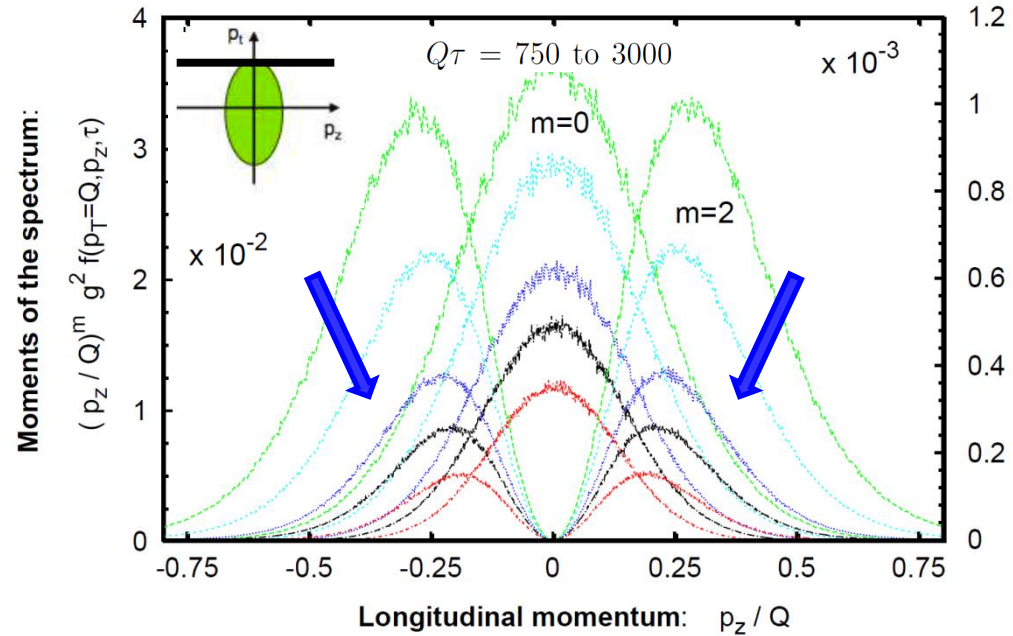
# Transverse and longitudinal spectra

(Coulomb type gauge)

## Transverse spectrum

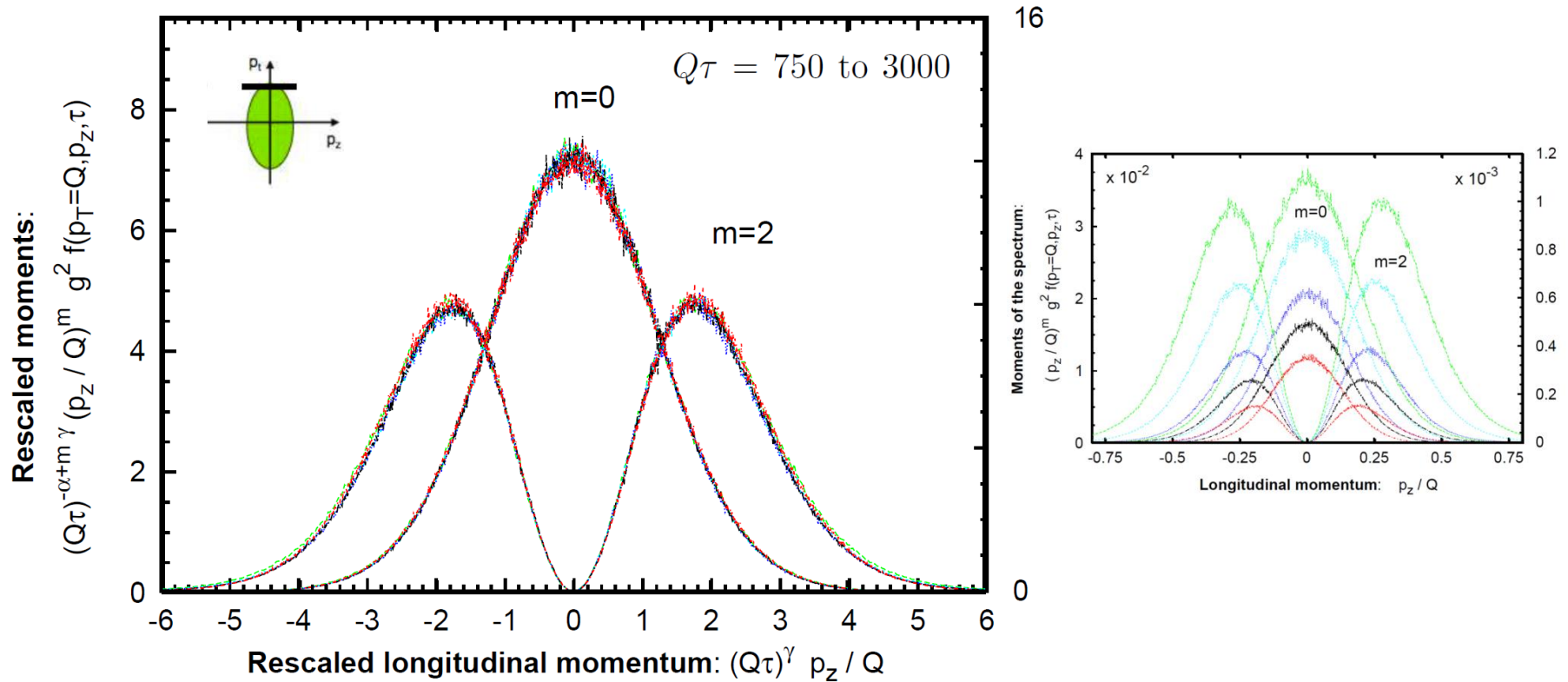


## Longitudinal spectrum



***Thermal-like transverse shape  $\sim 1/p_T$  even as longitudinal distribution is being 'squeezed'***

# Nonthermal fixed point: Self-similar evolution!



**The spectrum shows a self-similar evolution with universal scaling exponents  $\alpha, \beta, \gamma$  and scaling function  $f_S$ :**

$$f(p_T, p_z, \tau) = (Q\tau)^\alpha f_S \left( \underbrace{(Q\tau)^\beta p_T, (Q\tau)^\gamma p_z}_{\text{stationary fixed-point distribution}} \right)$$

stationary fixed-point distribution

# Nature of nonthermal fixed point: wave turbulence

**Boltzmann equation** with generic collision term for longitudinal expansion:

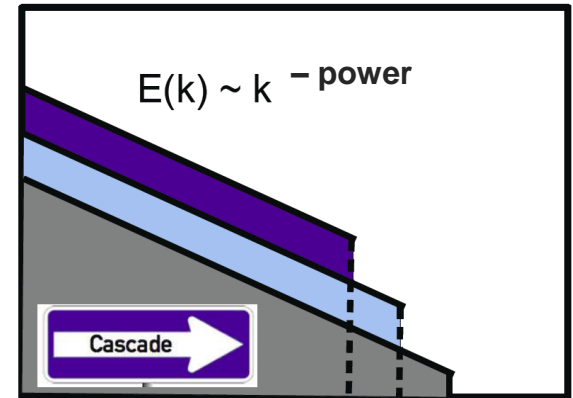
$$\left[ \partial_t - \frac{p_z}{t} \partial_{p_z} \right] f(p_T, p_z, t) = C[p_T, p_z, t; f]$$

**Self-similar evolution:**

$$f(p_T, p_z, t) = t^\alpha f_S(t^\beta p_T, t^\gamma p_z)$$

$$C[p_T, p_z, t; f] = t^\mu C[t^\beta p_T, t^\gamma p_z; f_S]$$

**Turbulent Thermalization**



→ **a) fixed point equation for stationary distribution:**

$$\alpha f_S(p_T, p_z) + \beta p_T \partial_{p_T} f_S(p_T, p_z) + (\gamma - 1) p_z \partial_{p_z} f_S(p_T, p_z) = C[p_T, p_z; f_S]$$

→ **b) scaling condition:**

$$\alpha - 1 = \mu(\alpha, \beta, \gamma)$$

# Nonthermal fixed point

Interpret scaling condition with **energy/number conserving\*** Fokker-Planck-type dynamics for the collisional broadening of longitudinal momentum distribution:

$$C^{(\text{elast})}[p_T, p_z; f] = \hat{q} \partial_{p_z}^2 f(p_T, p_z, t)$$

with **momentum diffusion parameter**:  $\hat{q} \sim \alpha_S^2 \int \frac{d^2 p_T}{(2\pi)^2} \int \frac{dp_z}{2\pi} f^2(p_T, p_z, t)$

→ 1)  $\mu = 3\alpha - 2\beta + \gamma$        $\xrightarrow{\alpha - 1 = \mu(\alpha, \beta, \gamma)}$        $2\alpha - 2\beta + \gamma + 1 = 0$

2) number conservation       $\longrightarrow$        $\alpha - 2\beta - \gamma + 1 = 0$

3) energy conservation       $\longrightarrow$        $\alpha - 3\beta - \gamma + 1 = 0$

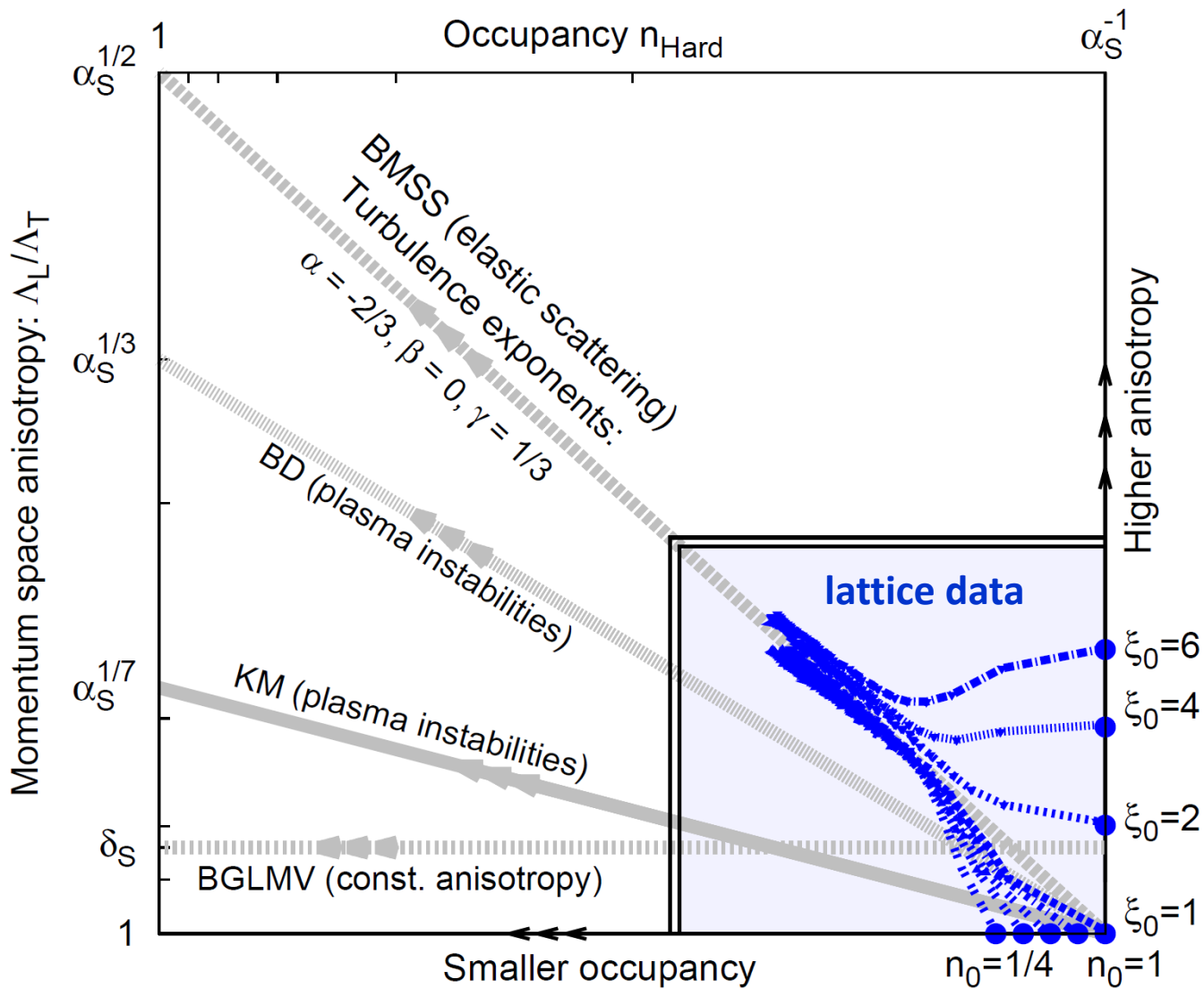
$\alpha = -2/3, \quad \beta = 0, \quad \gamma = 1/3$

*remarkable agreement with lattice data!*

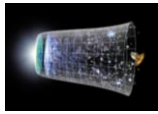
\*cf. early stages of Baier, Mueller, Schiff, Son, PLB 502 (2001) 51 ('BMSS')

# Universal attractor

Evolution in the 'anisotropy-occupancy plane'

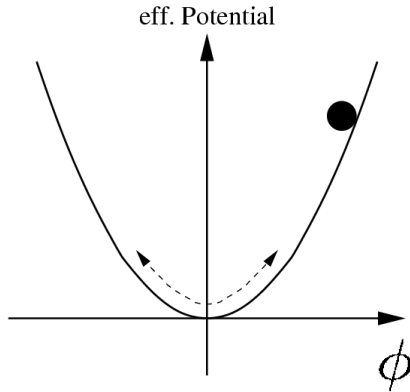






# A well understood *quantum* example

## Early universe preheating:



Kofman, Linde, Starobinsky, PRL 73 (1994) 3195

Scalar  $\lambda\Phi^4$  inflaton:  $\lambda \ll 1$

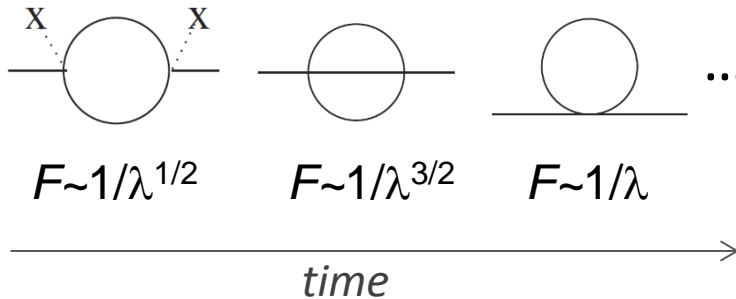
- **large** initial field  $\phi = \langle \Phi \rangle \sim 1/\lambda^{1/2}$
- **small** fluctuation  $F \sim \langle \{\Phi, \Phi\} \rangle - \phi\phi \sim 1$

*Instability:*  $F(t) \sim e^{\gamma t} \quad (\gamma > 0)$

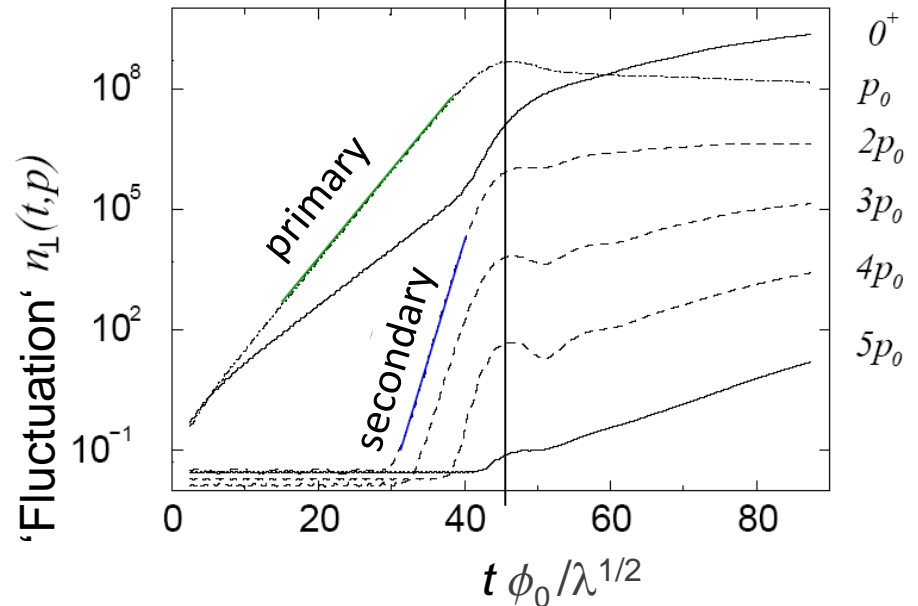
## Quantum field theory:

Berges, Serreau, PRL 91 (2003) 111601

### *Dynamical power counting (2PI):*



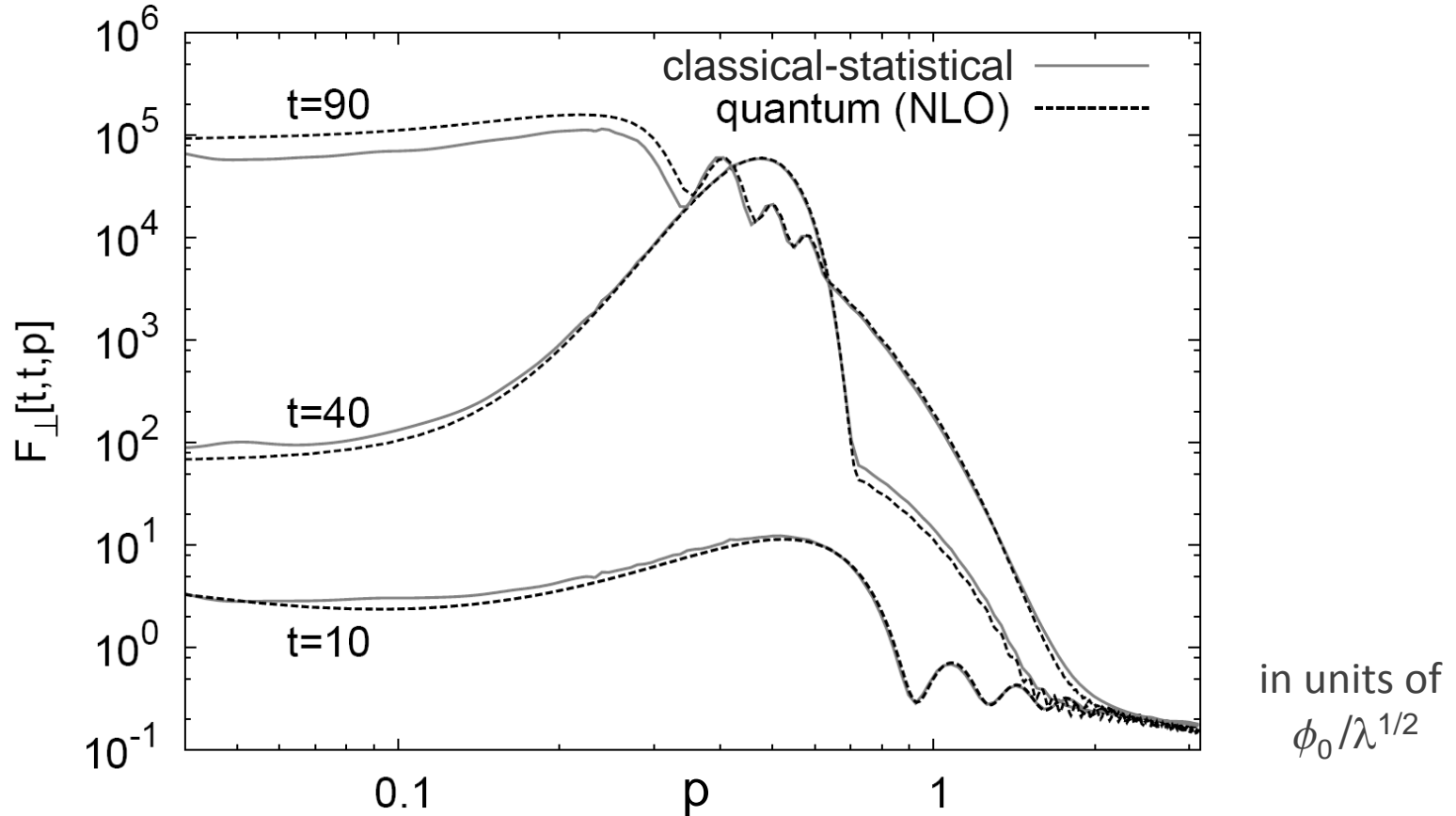
**instability**  $\longleftrightarrow$  **turbulence**



# Build up of fluctuations

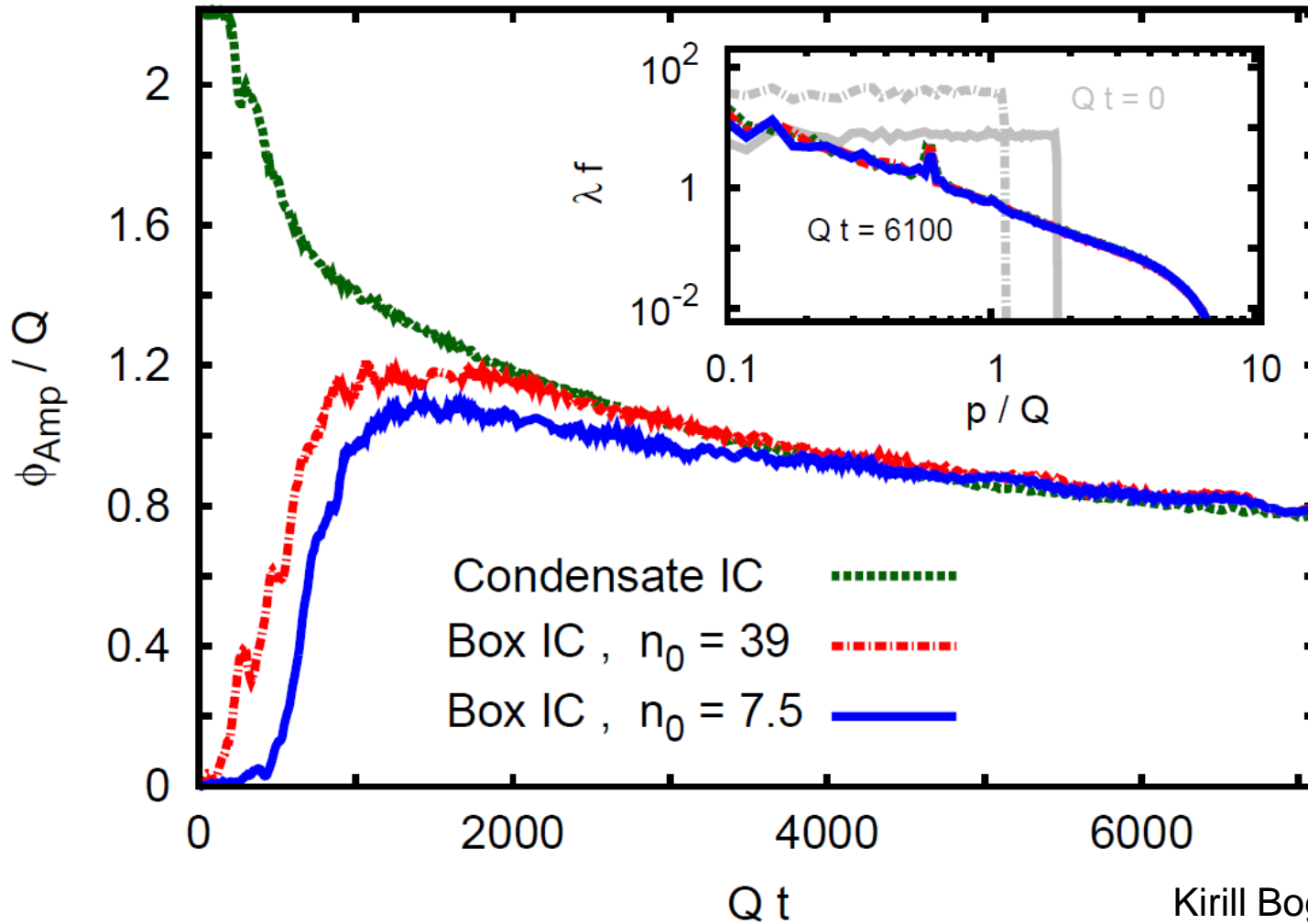
Build up of fluctuations during instability regime:

large initial field  $\phi \sim 1/\lambda^{1/2}$   $t \sim \log 1/\lambda$   $\rightarrow$  large fluctuations  $F \sim 1/\lambda$



**Good agreement of classical-statistical and quantum for large  $\phi$ , large  $F$**

# Inensitivity to initial conditions

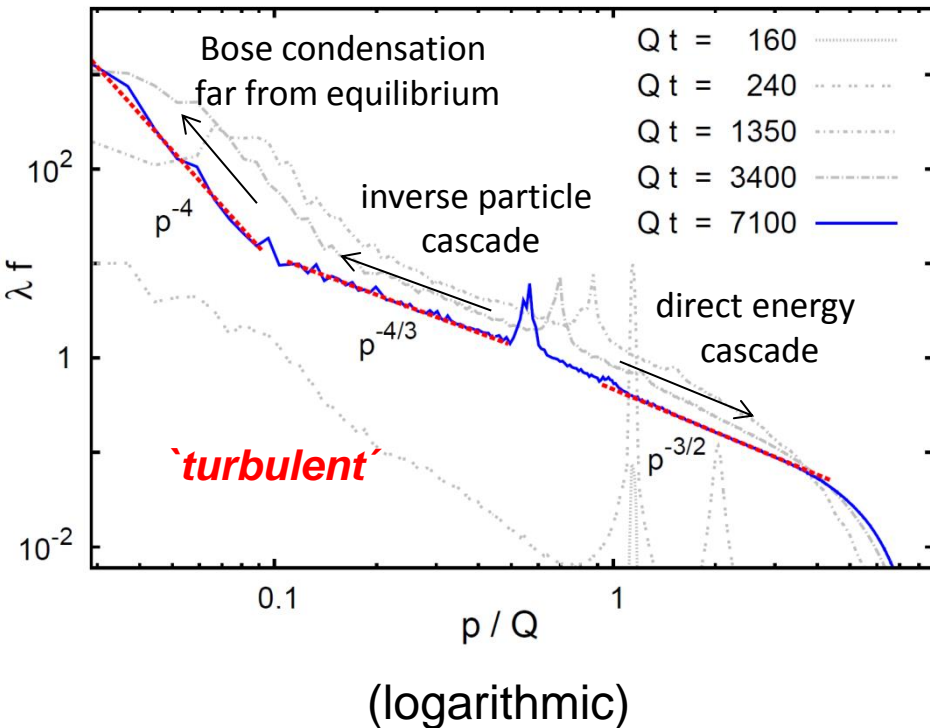


***Details about initial conditions are lost after instability period***

# Limitations of classical-statistical simulations

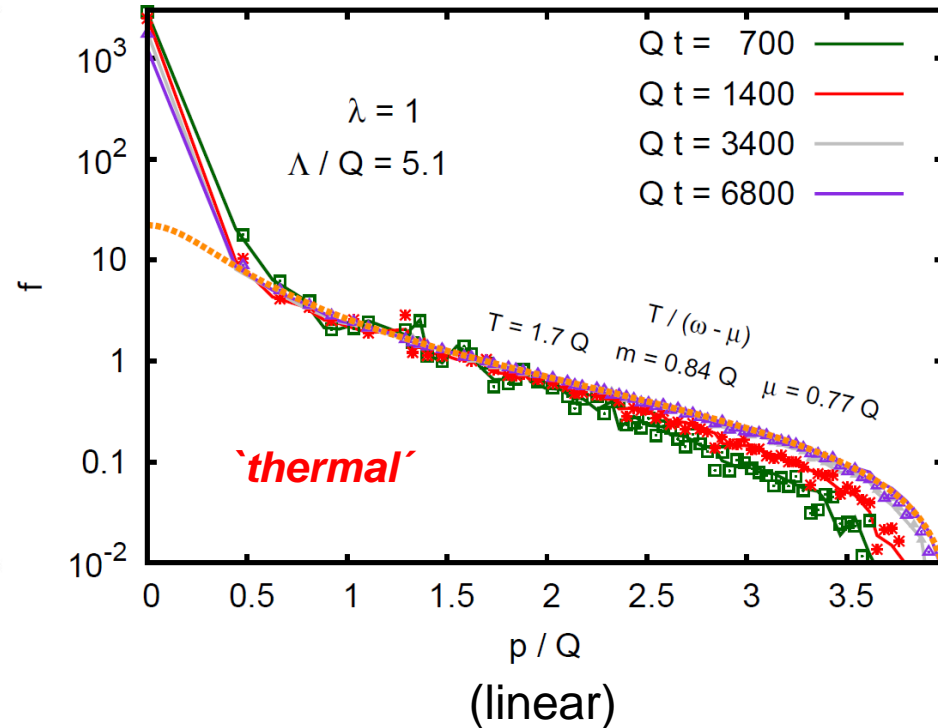
Comparing classical-statistical results at different couplings:

Turbulent thermalization  
at weak coupling  $\lambda \ll 1$



Kirill Boguslavski

$\lambda = 1$  thermalization scenario in  
Epelbaum, Gelis, NPA 872 (2011) 210

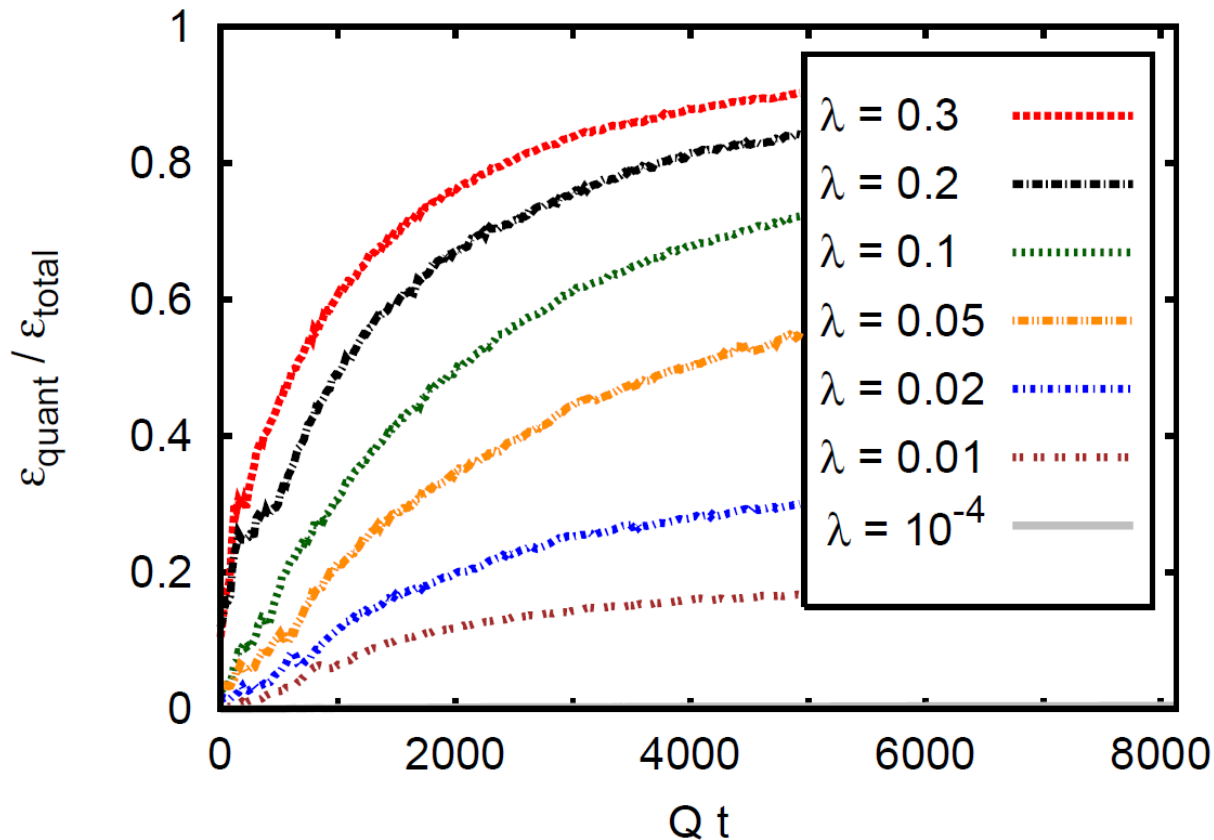


data provided by Thomas Epelbaum

See also Micha, Tkachev (2003): UV; Berges, Rothkopf, Schmidt (2008): IR

## Some diagnostics

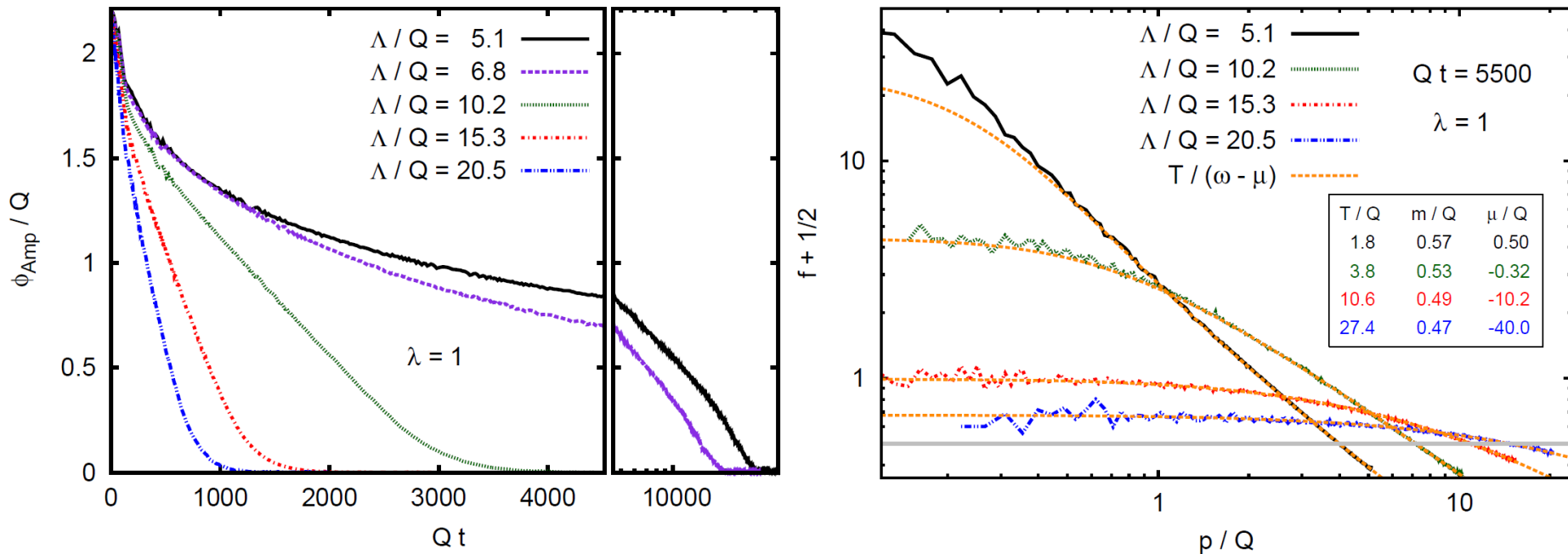
$$\epsilon_{\text{quant}} = \int \frac{d^3 p}{(2\pi)^3} p f(p) \Theta(1 - f)$$



***Already for rather small couplings a sizeable fraction of modes does not fulfill classicality condition  $f(p) > 1$***

# Cutoff dependence

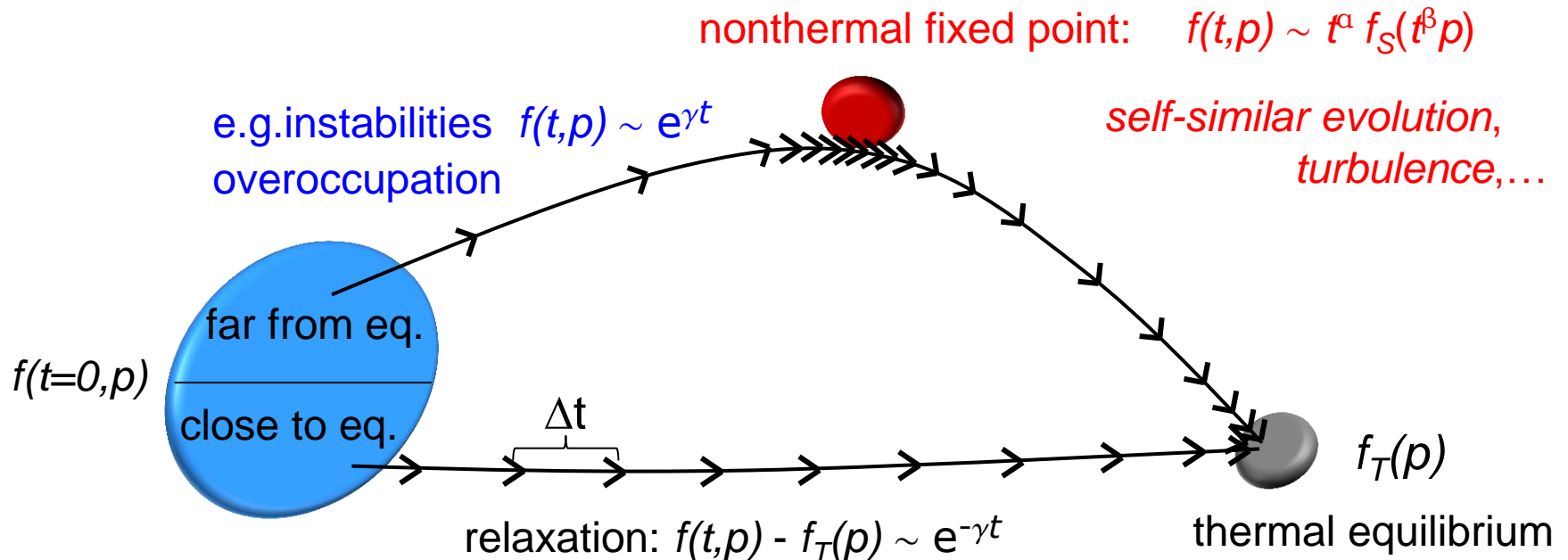
- $\lambda \ll 1$  results are insensitive to variation of UV cutoff
- Beyond their range of validity classical-statistical simulations show strong cutoff dependencies:



See also Aarts, Berges (2001); Smit, Arrizabalaga, Tranberg (2004)...

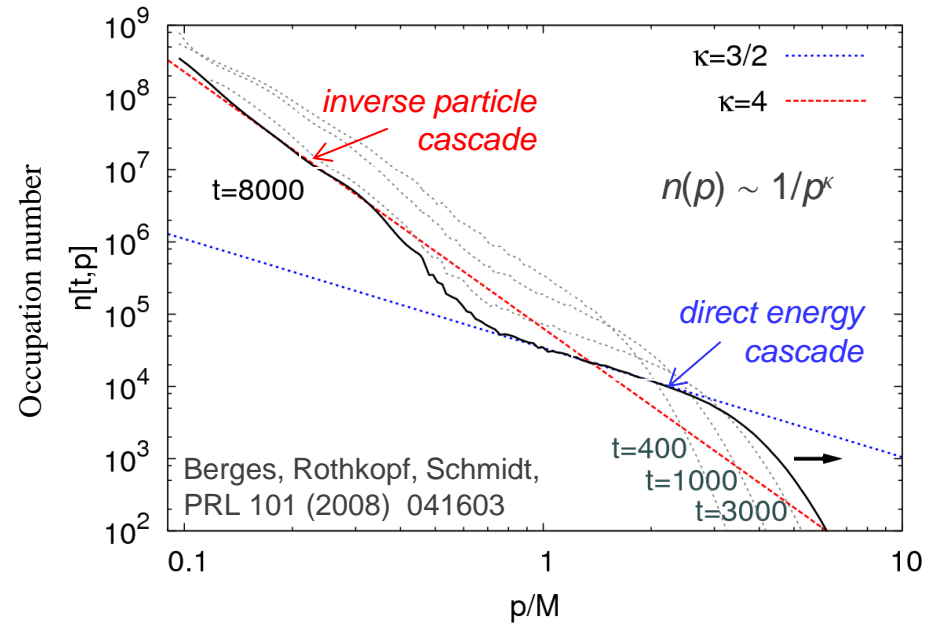
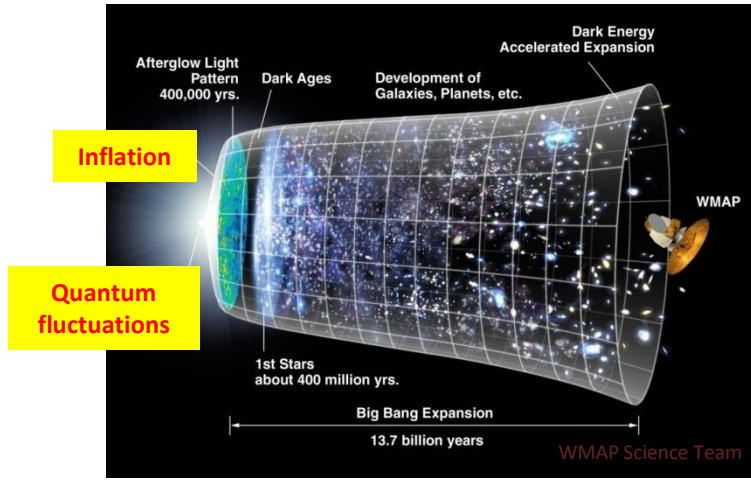
# Conclusions

- Universal attractor** → ‘bottom up’ thermalization  
 BMSS:  $\tau_{\text{iso.}} \sim Q^{-1} \alpha_S^{-5/2}$  ;  $T_{\text{therm}} \sim \alpha_S^{2/5} Q$
- Strongly correlated dynamics** throughout the entire turbulent regime, despite very weak coupling (universal scaling, ‘self-tuned criticality’)!



# Universality far from equilibrium

- Reheating dynamics after chaotic inflation



- Superfluid turbulence in a cold Bose gas

