

Radioactively Powered Emission from Neutron Star Mergers

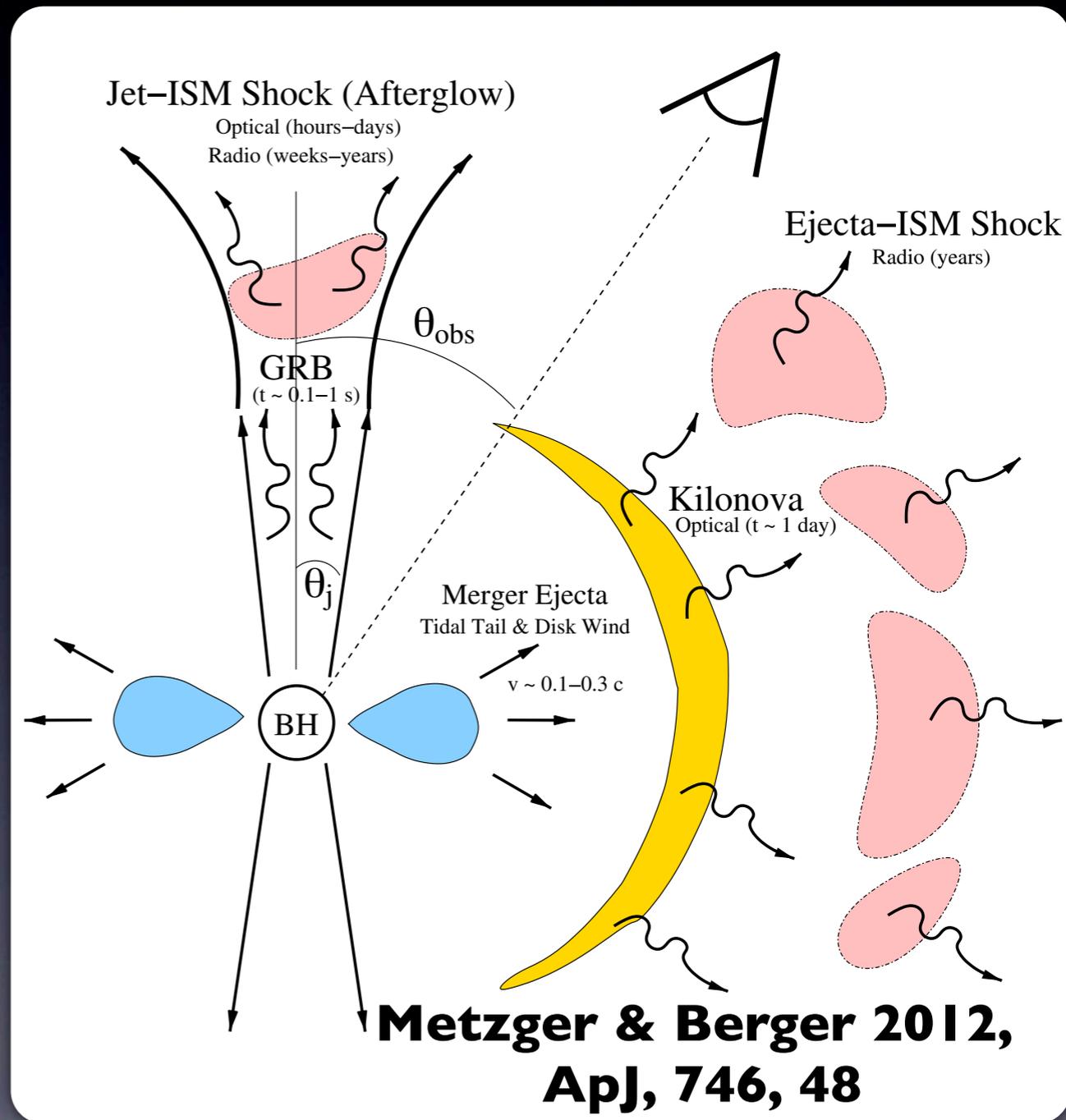
(kilonova, macronova, r-process nova, goldnova, ...)

Masaomi Tanaka (NAOJ)

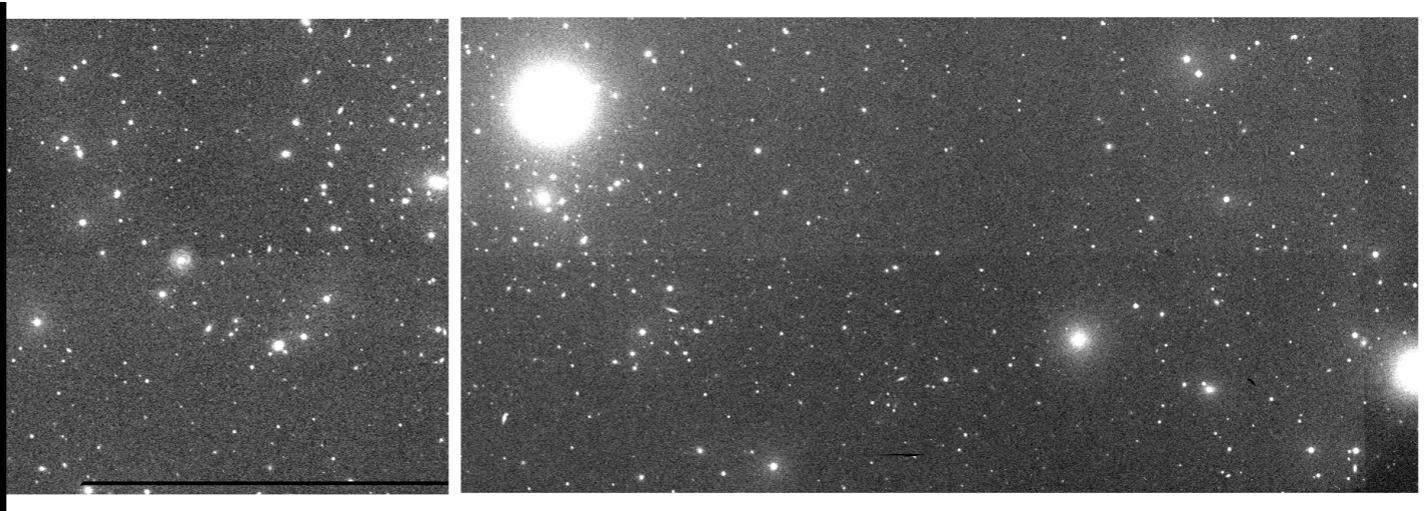
- MT & Hotokezaka 2013, ApJ, 775, L13
- MT, Hotokezaka, Kyutoku, Wanajo, Kiuchi, Sekiguchi, Shibata
2013, ApJ, in press (arXiv:1310.2774)
- Hotokezaka, Kyutoku, MT, Kiuchi, Sekiguchi, Shibata, Wanajo
2013, ApJ, 778, L16

EM emission from NS-NS merger = EM counterparts of GW sources

- On-axis short GRB
- Off-axis radio/optical afterglow
- Radioactive emission (r-process nuclei)



2 deg Kiso Supernova Survey (KISS)
Kiso 1 m Schmidt telescope



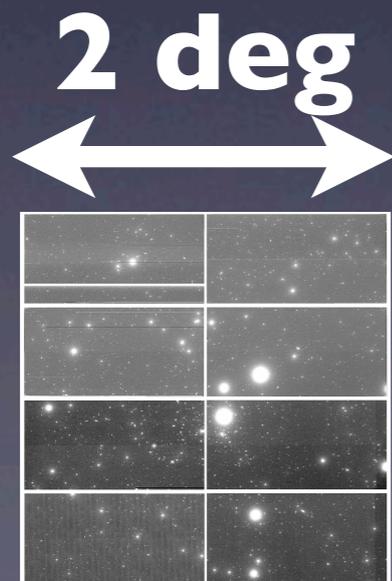
~ 200 galaxies / 4deg²
(< 200 Mpc)

GW alert error box

e.g. 10 deg x 10 deg

~ 5000 galaxies

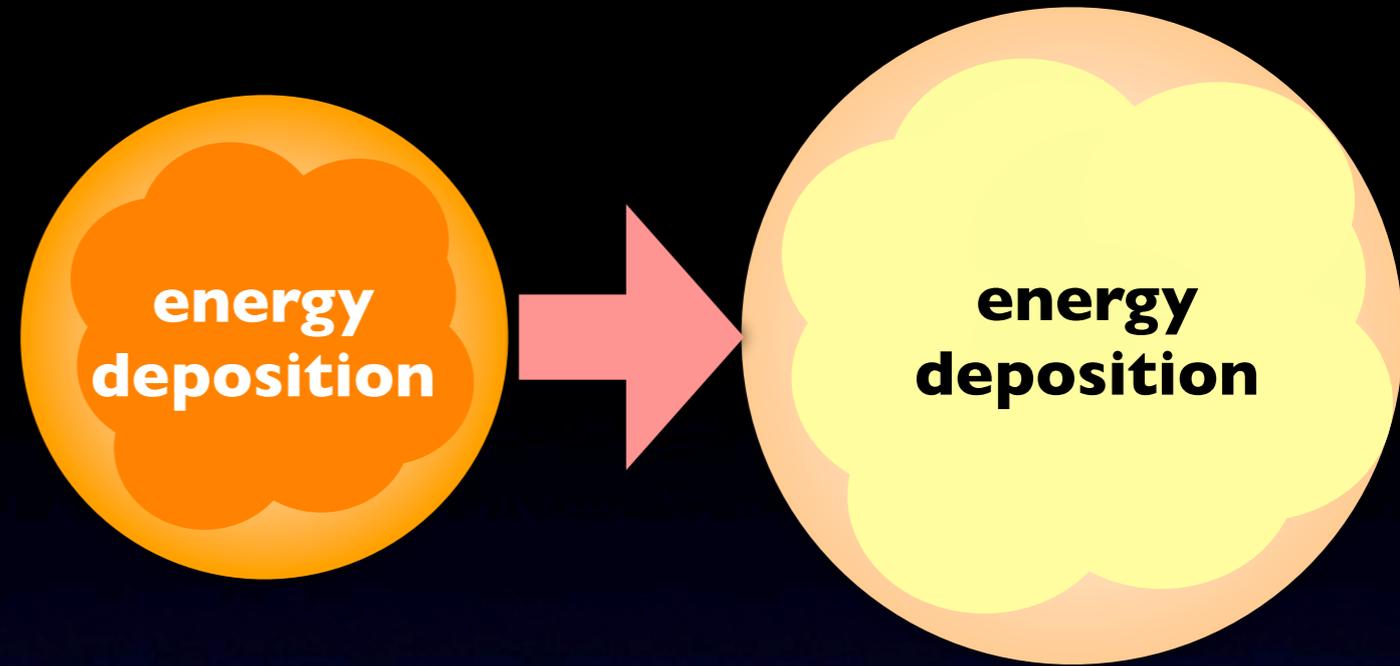
(< 200 Mpc)



Expected emission

Li & Paczynski 98
Kulkarni 05
Metzger+10

A. Bauswein's and B. Metzger's talks



Timescale

$$t_p \sim \underline{1 \text{ day}} \left(\frac{M}{0.01 M_\odot} \right)^{1/2} \left(\frac{v}{0.2c} \right)^{-1/2} \left(\frac{\kappa}{0.1 \text{ cm}^2 \text{ g}^{-1}} \right)^{1/2}$$

Luminosity

$$L \sim \underline{10^{42} \text{ erg s}^{-1}} \left(\frac{M}{0.01 M_\odot} \right)^{1/2} \left(\frac{v}{0.2c} \right)^{1/2} \left(\frac{\kappa}{0.1 \text{ cm}^2 \text{ g}^{-1}} \right)^{-1/2}$$

Opacity of Fe

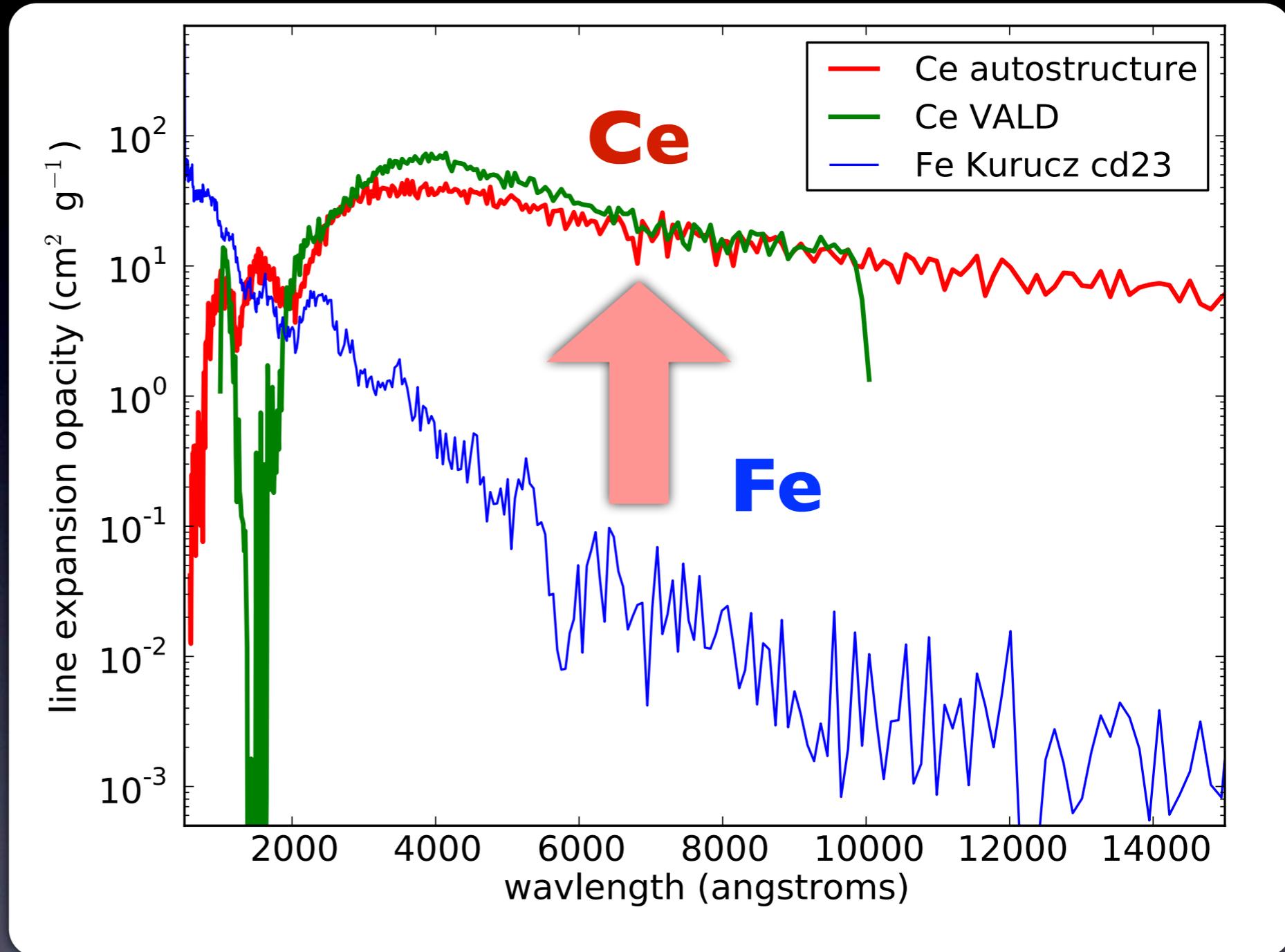
~ 20 mag at 200 Mpc
(1m-class telescopes)

Opacity of Lanthanoid Ce

$$\kappa = 10 \text{ cm}^2 \text{ g}^{-1}$$

$$\kappa = 1 \text{ cm}^2 \text{ g}^{-1}$$

$$\kappa = 0.1 \text{ cm}^2 \text{ g}^{-1}$$



High opacity (x 100) => Lower luminosity (x 1/10)

Kasen et al. 2013

Barnes & Kasen 2013

hydrogen 1 H 1.0079																	helium 2 He 4.0026	
lithium 3 Li 6.941	beryllium 4 Be 9.0122											boron 5 B 10.811	carbon 6 C 12.011	nitrogen 7 N 14.007	oxygen 8 O 15.999	fluorine 9 F 18.998	neon 10 Ne 20.180	
sodium 11 Na 22.990	magnesium 12 Mg 24.305											aluminium 13 Al 26.982	silicon 14 Si 28.086	phosphorus 15 P 30.974	sulfur 16 S 32.065	chlorine 17 Cl 35.453	argon 18 Ar 39.948	
potassium 19 K 39.098	calcium 20 Ca 40.078	scandium 21 Sc 44.956	titanium 22 Ti 47.867	vanadium 23 V 50.942	chromium 24 Cr 51.996	manganese 25 Mn 54.938	iron 26 Fe 55.845	cobalt 27 Co 58.933	nickel 28 Ni 58.693	copper 29 Cu 63.546	zinc 30 Zn 65.39	gallium 31 Ga 69.723	germanium 32 Ge 72.61	arsenic 33 As 74.922	selenium 34 Se 78.96	bromine 35 Br 79.904	krypton 36 Kr 83.798	
rubidium 37 Rb 85.468	strontium 38 Sr 87.62	yttrium 39 Y 88.906	zirconium 40 Zr 91.224	niobium 41 Nb 92.906	molybdenum 42 Mo 95.94	technetium 43 Tc [98]	ruthenium 44 Ru 101.07	rhodium 45 Rh 102.91	palladium 46 Pd 106.42	silver 47 Ag 107.87	cadmium 48 Cd 112.41	indium 49 In 114.82	tin 50 Sn 118.71	antimony 51 Sb 121.76	tellurium 52 Te 127.60	iodine 53 I 126.90	xenon 54 Xe 131.29	
caesium 55 Cs 132.91	barium 56 Ba 137.33	57-70 *	lutetium 71 Lu 174.97	hafnium 72 Hf 178.49	tantalum 73 Ta 180.95	tungsten 74 W 183.84	rhenium 75 Re 186.21	osmium 76 Os 190.23	iridium 77 Ir 192.22	platinum 78 Pt 195.08	gold 79 Au 196.97	mercury 80 Hg 200.59	thallium 81 Tl 204.38	lead 82 Pb 207.2	bismuth 83 Bi 208.98	polonium 84 Po [209]	astatine 85 At [210]	radon 86 Rn [222]
francium 87 Fr [223]	radium 88 Ra [226]	89-102 **	lawrencium 103 Lr [262]	rutherfordium 104 Rf [261]	dubnium 105 Db [262]	seaborgium 106 Sg [266]	bohrium 107 Bh [264]	hassium 108 Hs [269]	meitnerium 109 Mt [268]	ununillium 110 Uun [271]	unununium 111 Uuu [272]	ununbium 112 Uub [277]	ununquadium 114 Uuq [289]					

* Lanthanide series

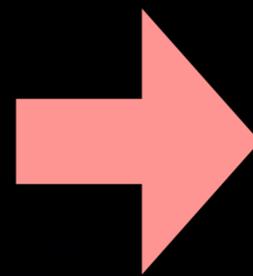
lanthanum 57 La 138.91	cerium 58 Ce 140.12	praseodymium 59 Pr 140.91	neodymium 60 Nd 144.24	promethium 61 Pm [145]	samarium 62 Sm 150.36	europium 63 Eu 151.96	gadolinium 64 Gd 157.25	terbium 65 Tb 158.93	dysprosium 66 Dy 162.50	holmium 67 Ho 164.93	erbium 68 Er 167.26	thulium 69 Tm 168.93	ytterbium 70 Yb 173.05
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** Actinide series

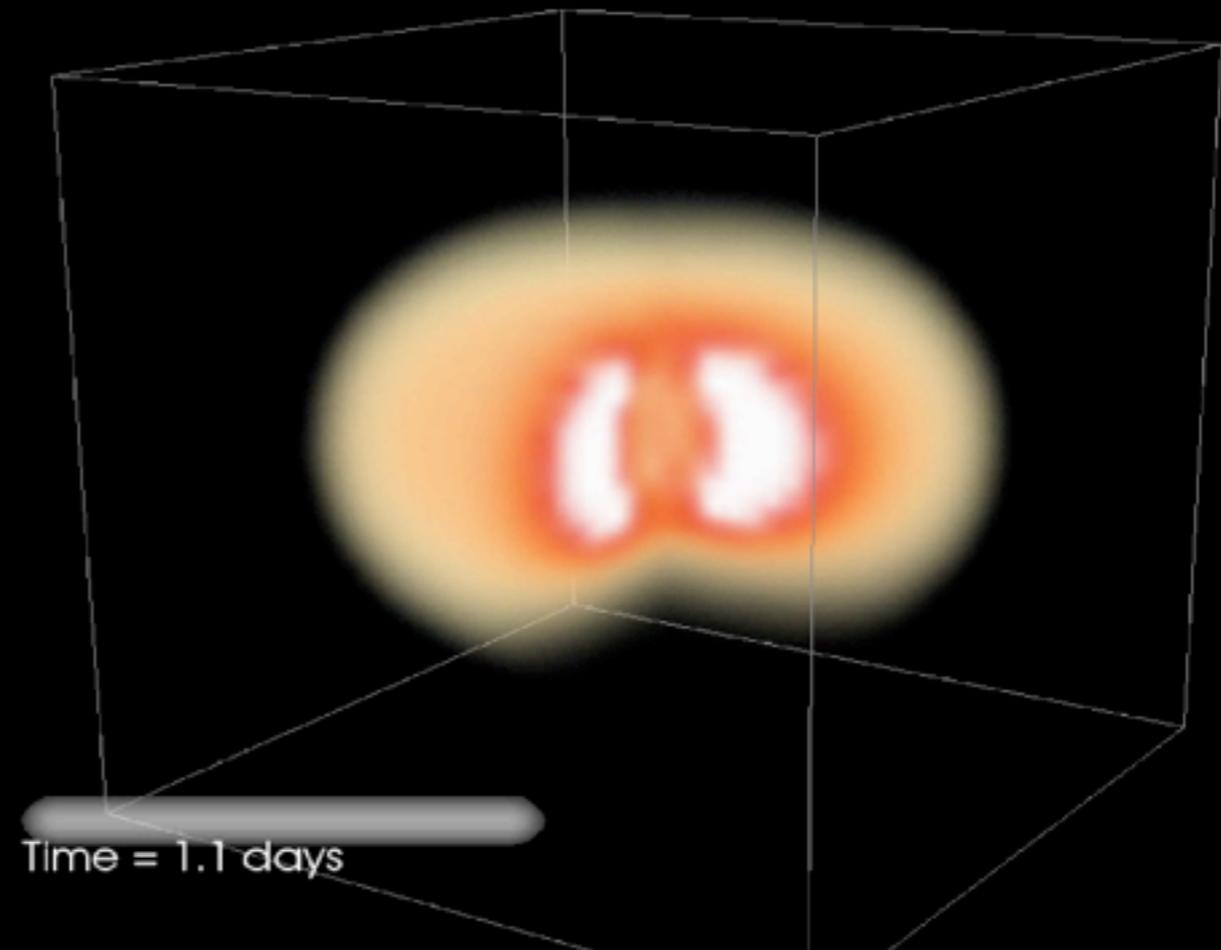
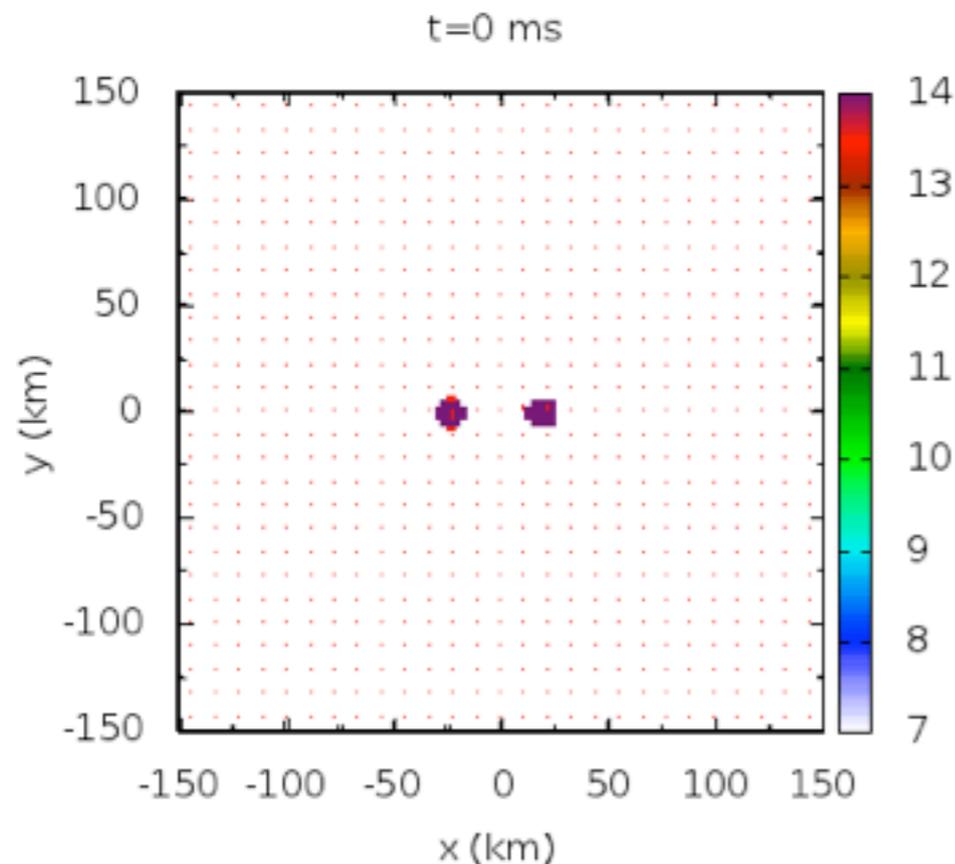
actinium 89 Ac [227]	thorium 90 Th 232.04	protactinium 91 Pa 231.04	uranium 92 U 238.03	neptunium 93 Np [237]	plutonium 94 Pu [244]	americium 95 Am [243]	curium 96 Cm [247]	berkelium 97 Bk [247]	californium 98 Cf [251]	einsteinium 99 Es [252]	fermium 100 Fm [257]	mendelevium 101 Md [258]	nobelium 102 No [259]
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**Radiative transfer in NS merger ejecta
with all the r-process elements**

Numerical relativity



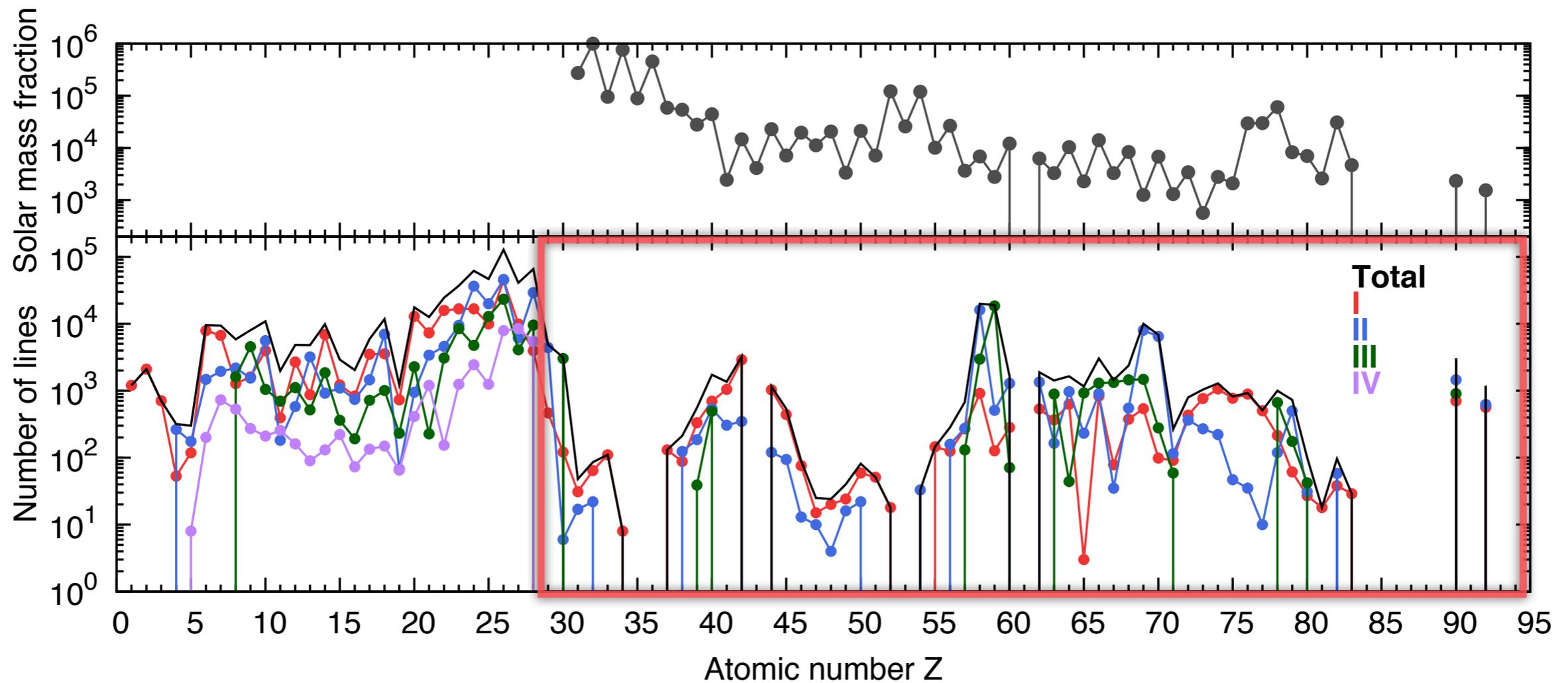
**3D, time-dependent,
multi-frequency
radiative transfer**



Hotokezaka et al. 2013

MT & Hotokezaka 2013

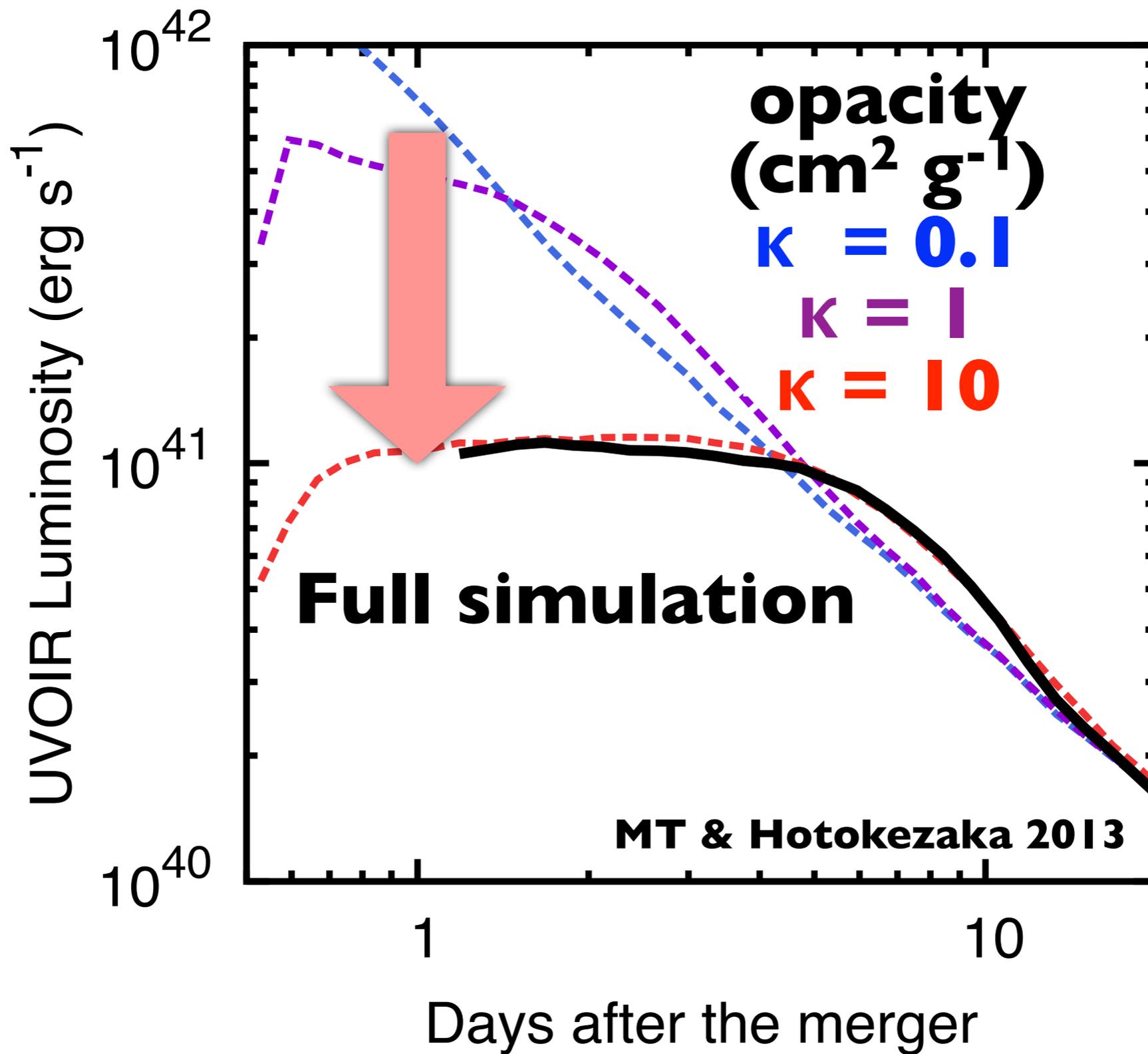
- 1. Opacity of r-process-dominated ejecta?**
- 2. Characteristic feature of NS merger?**
- 3. Dependence on EOSs and progenitors?**



**~500,000 transitions (up to Fe)
+ 100,000 transitions (r-process)**



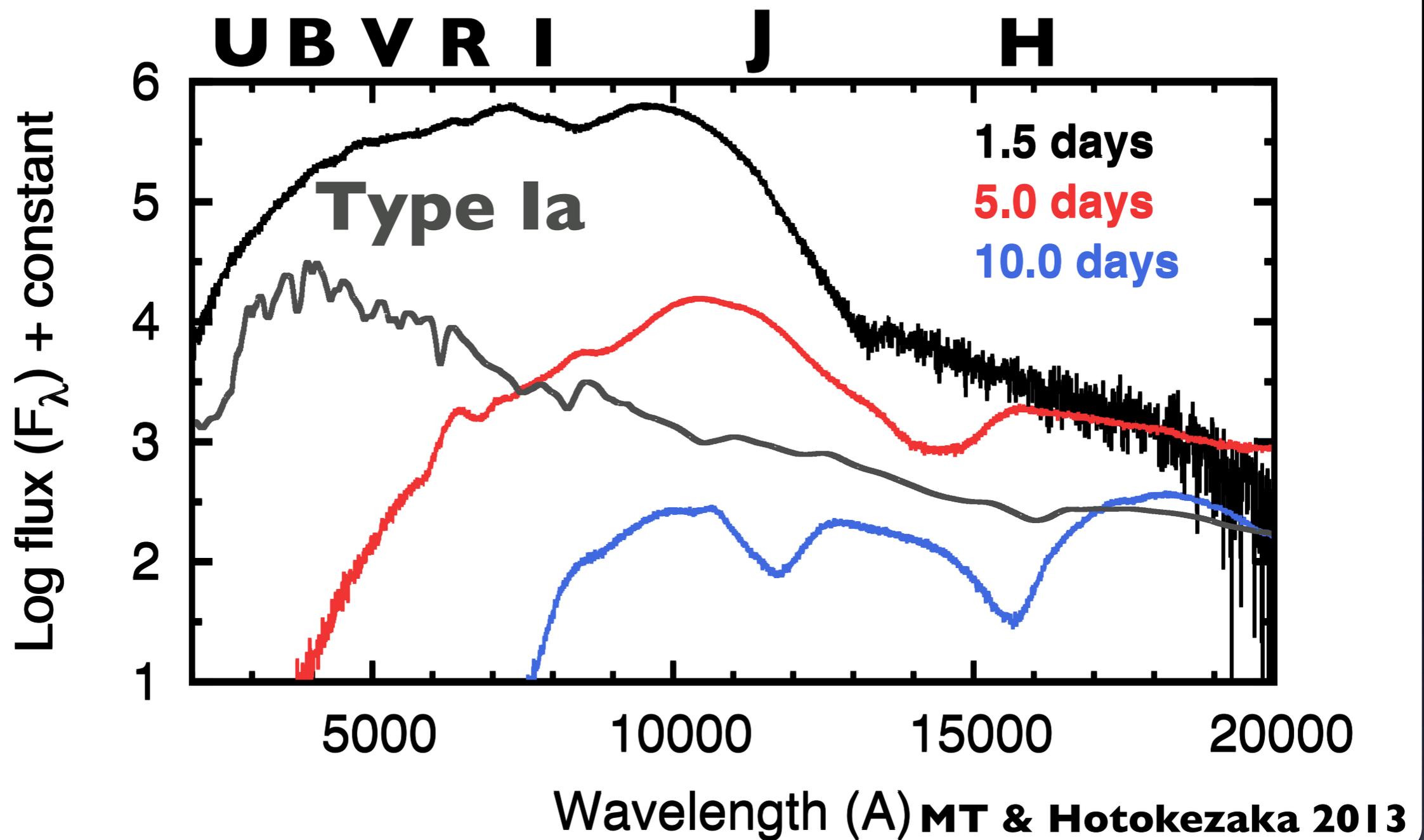
**Evaluate bound-bound opacity
(in each time step)**



Higher opacity by factor of 100

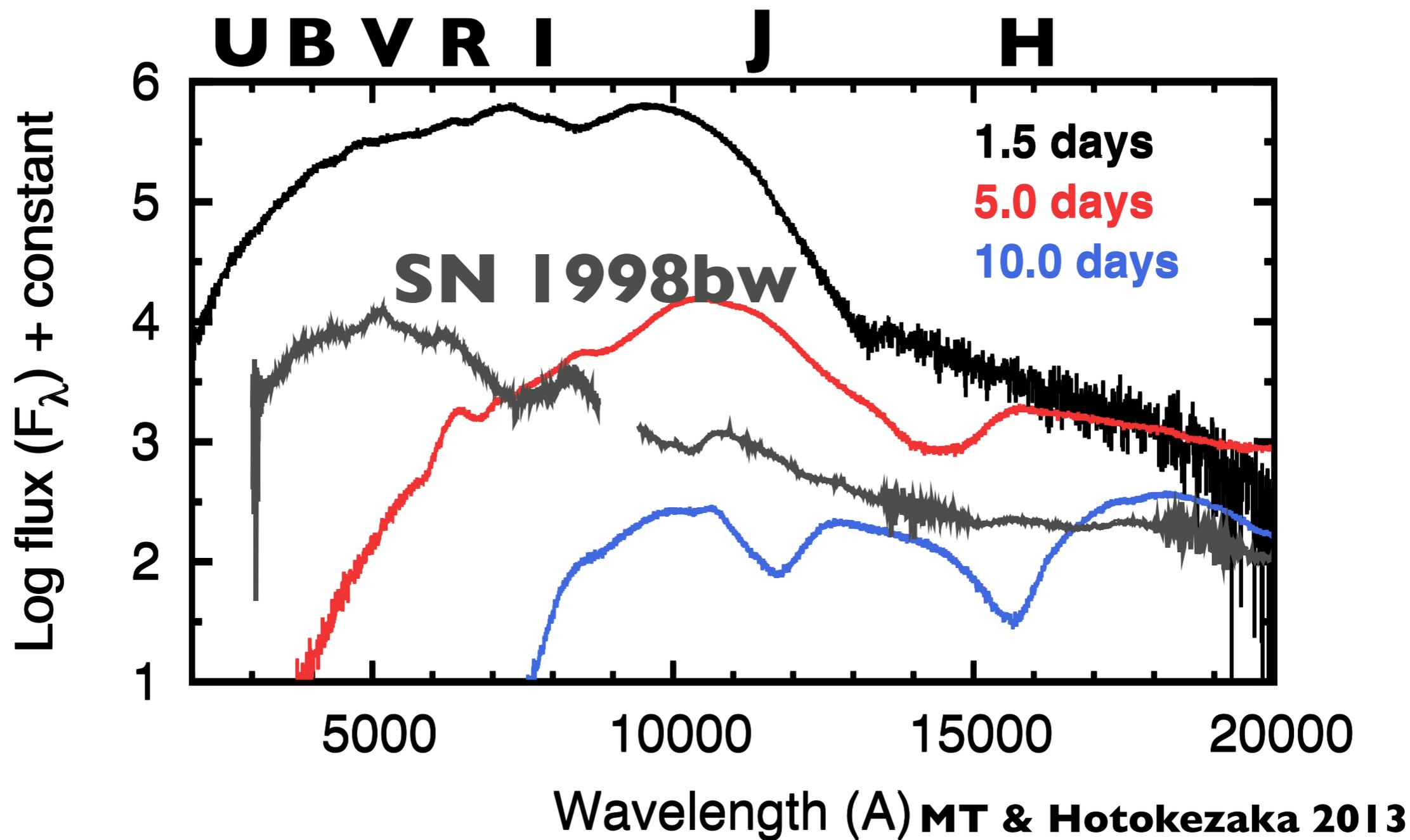
Fainter than previously expected by a factor of 10

(consistent with Kasen+13, Barnes & Kasen 13)



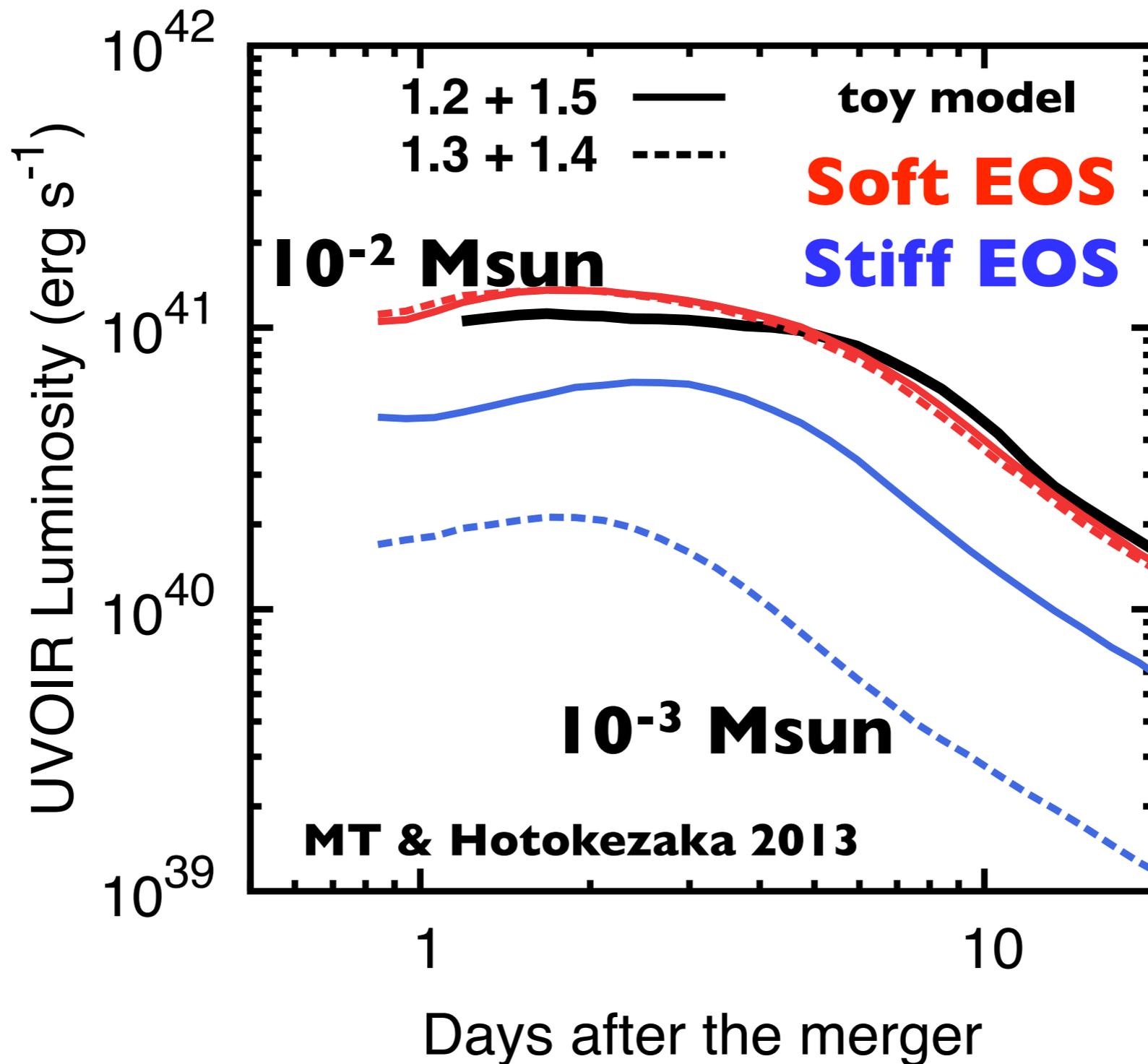
- **Very red SED (peak at NIR)**
- **Extremely broad-line (feature-less) spectra**

(Identification of r-process elements seems difficult)



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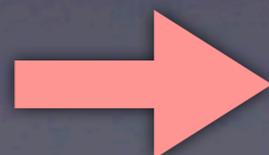


Radius of
1.35 Msun NS

R = 11.1 km

R = 13.6 km

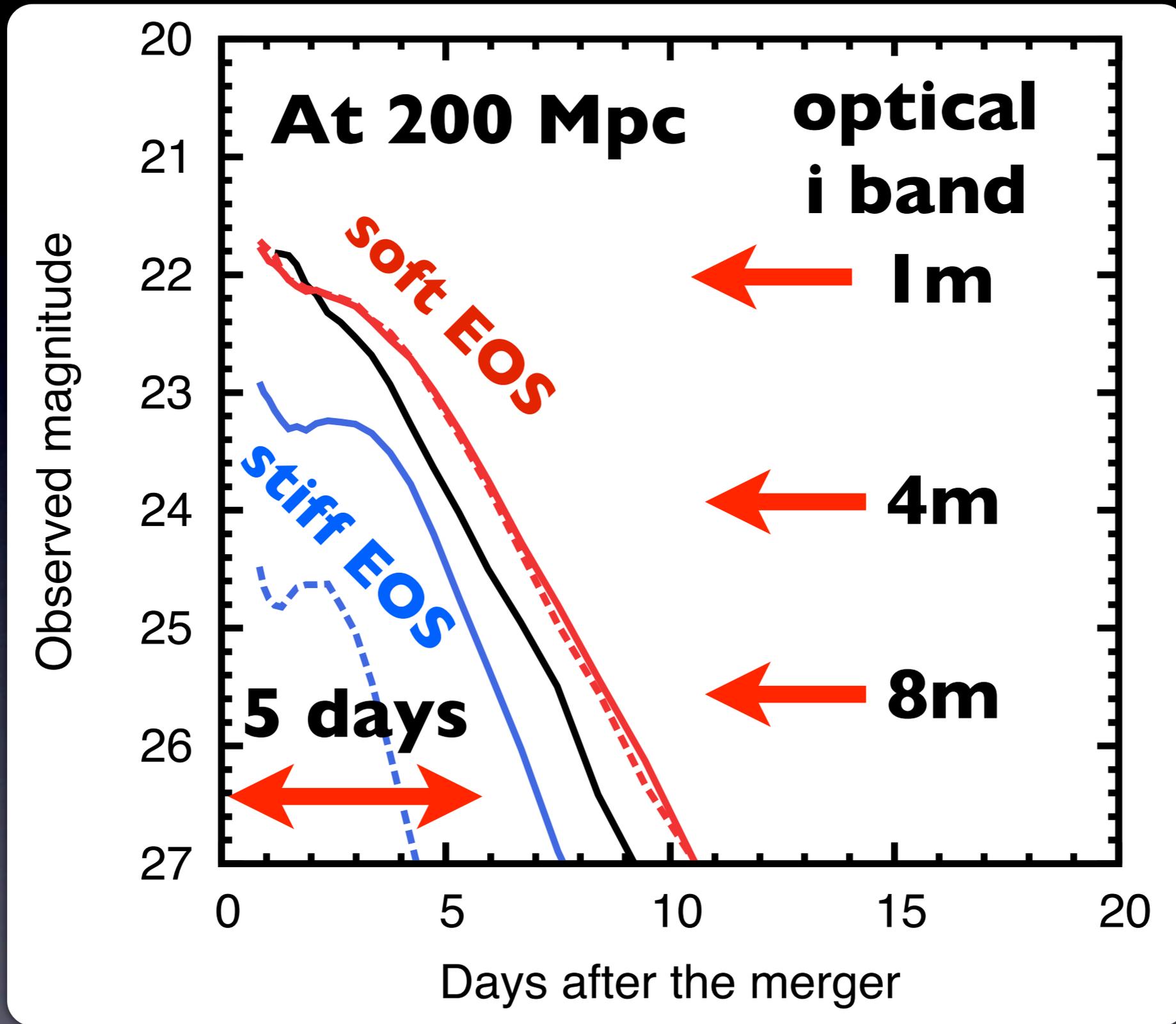
Softer EOS/Higher NS mass ratio



Brighter emission

(Hotokezaka+13; Bauswein+13, M. Shibata's, A. Bauswein's talks)

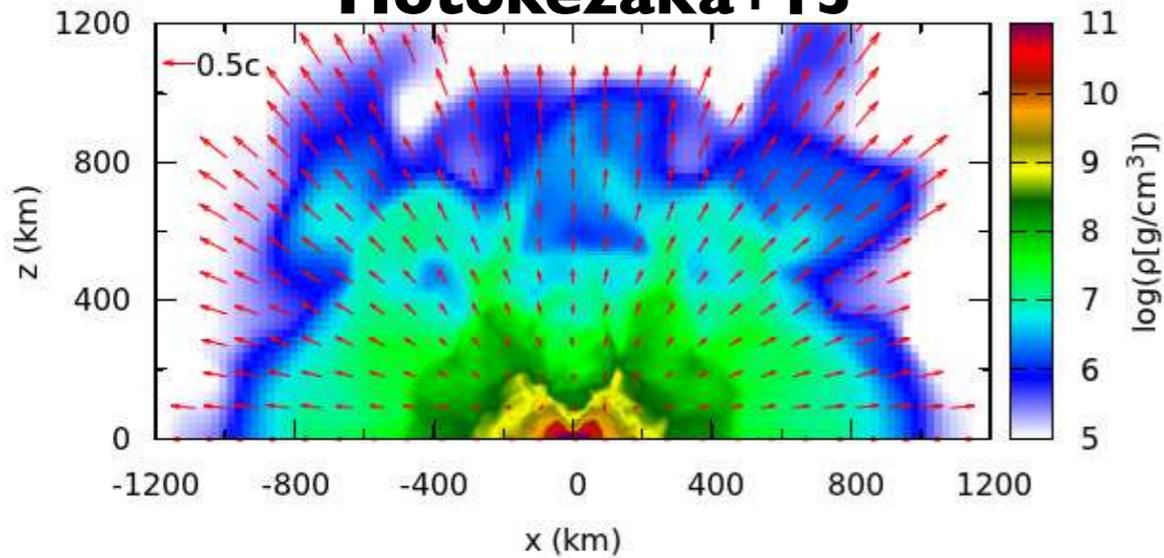
Observing strategy after GW detection



**Best with i band (0.8 μm)
NIR (~ 21 AB mag in NIR)**

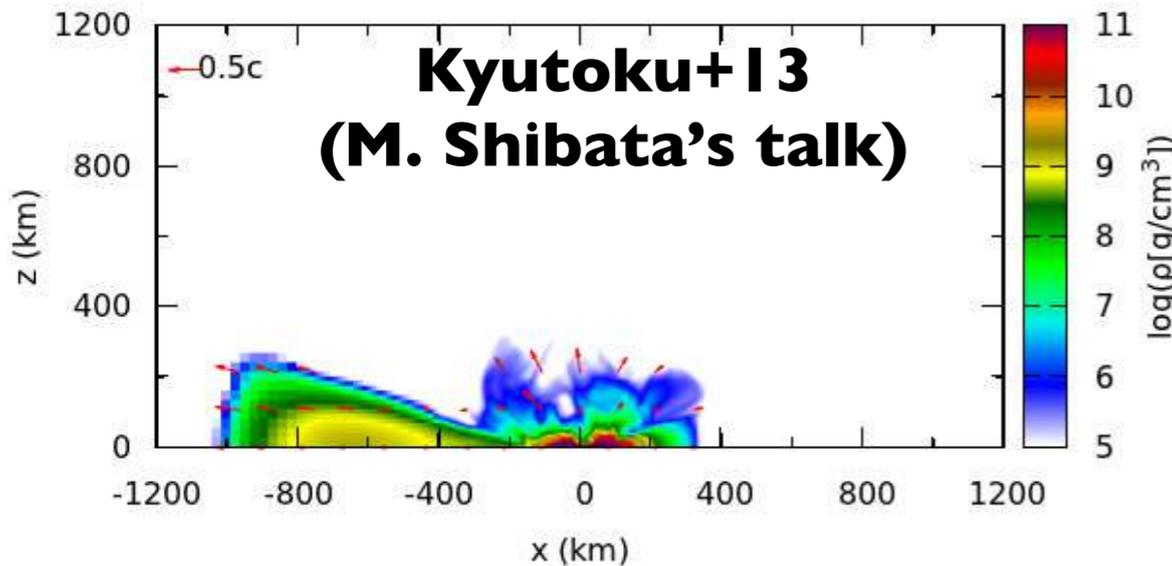
NS-NS

Hotokezaka+I3



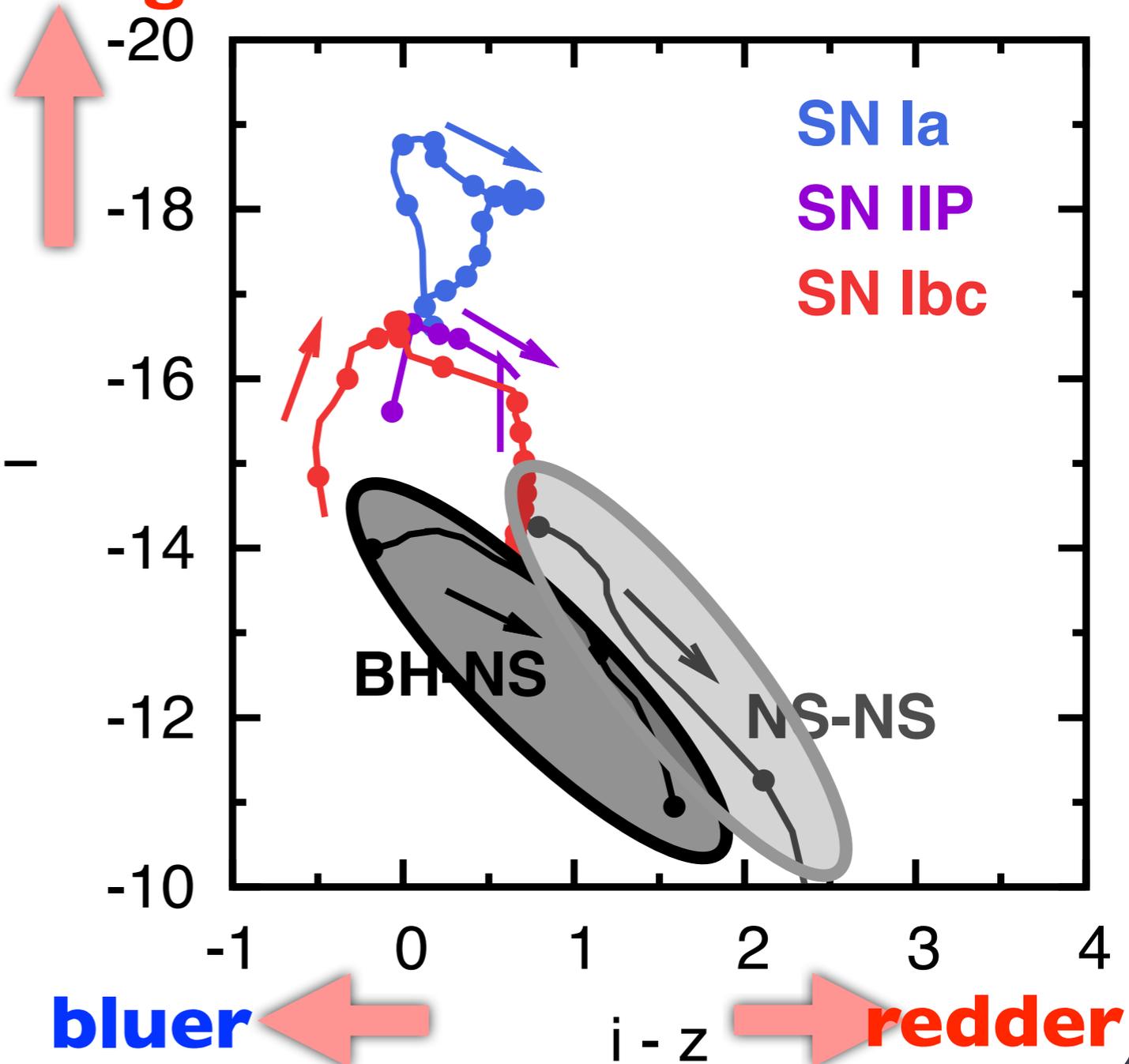
BH-NS

Kyutoku+I3 (M. Shibata's talk)



HR diagram

brighter



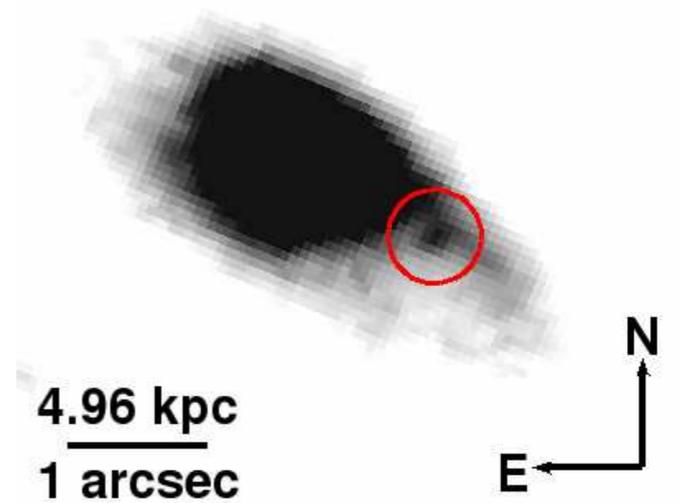
Application to GRB 130603B

Berger+13

Tanvir+13

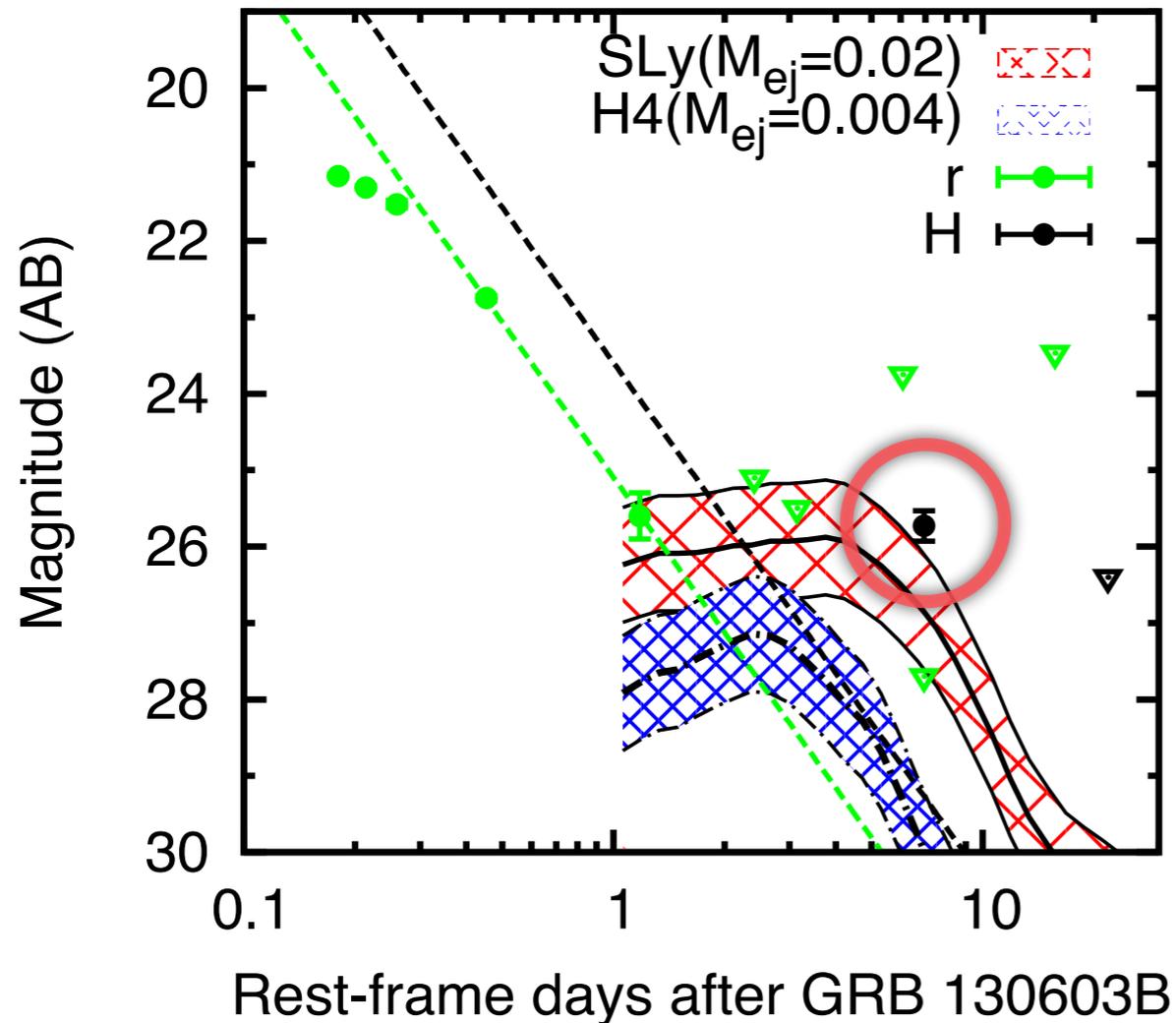
A. Fruchter's talk

WFC3/F160W

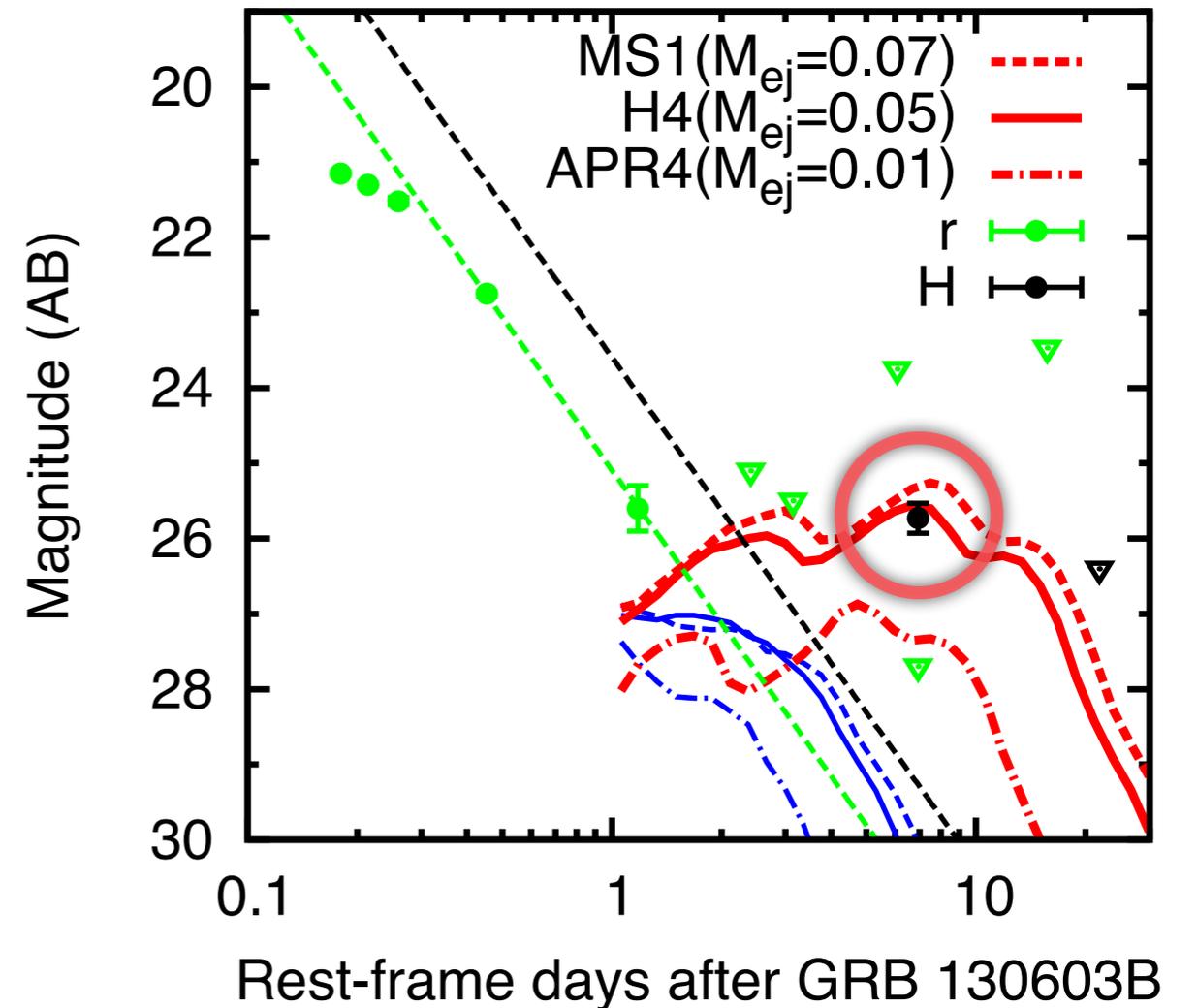


NS-NS

Hotokezaka+13



BH-NS



Constraints on Progenitors/EOSs

Efficient mass ejection

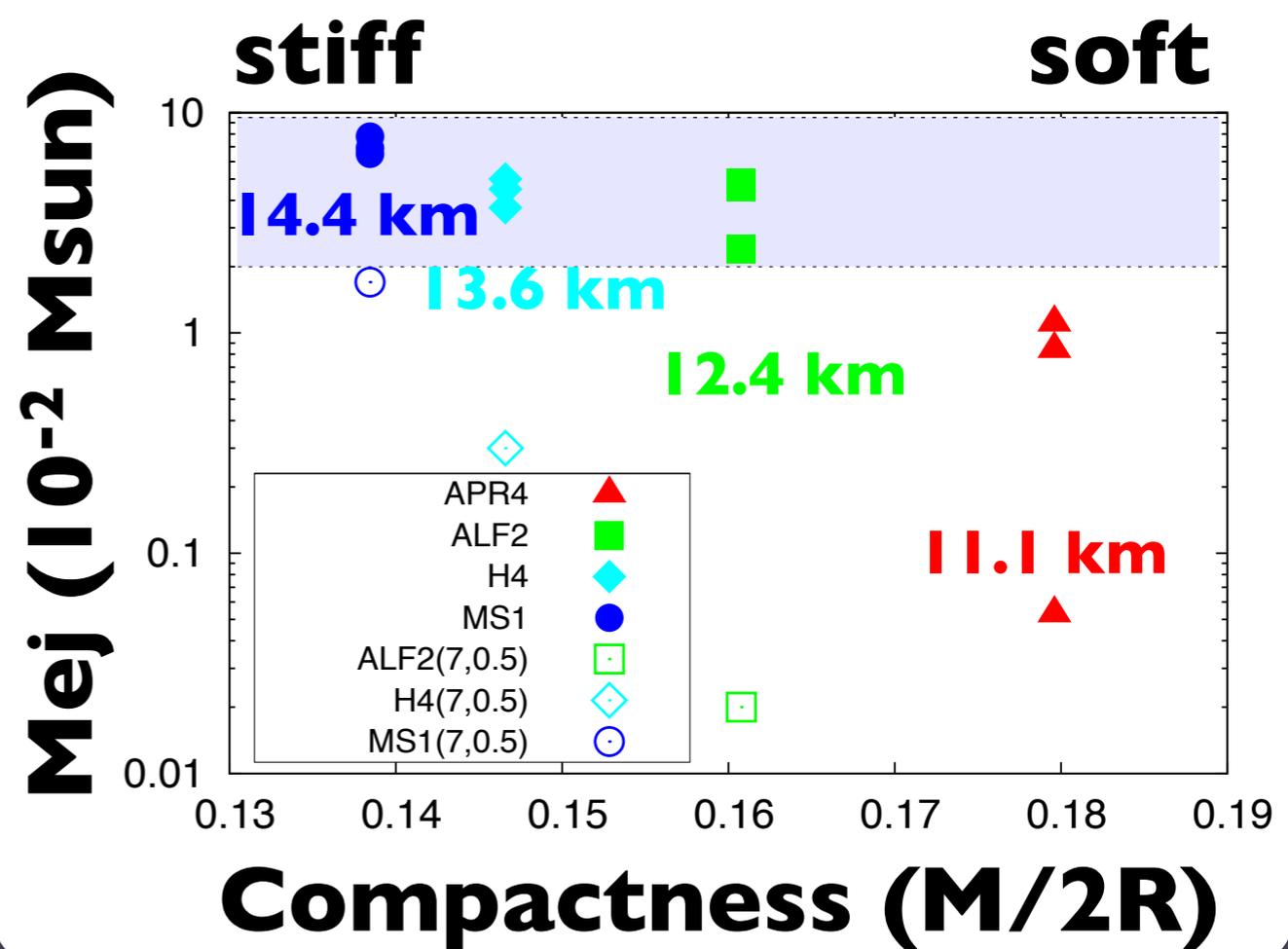
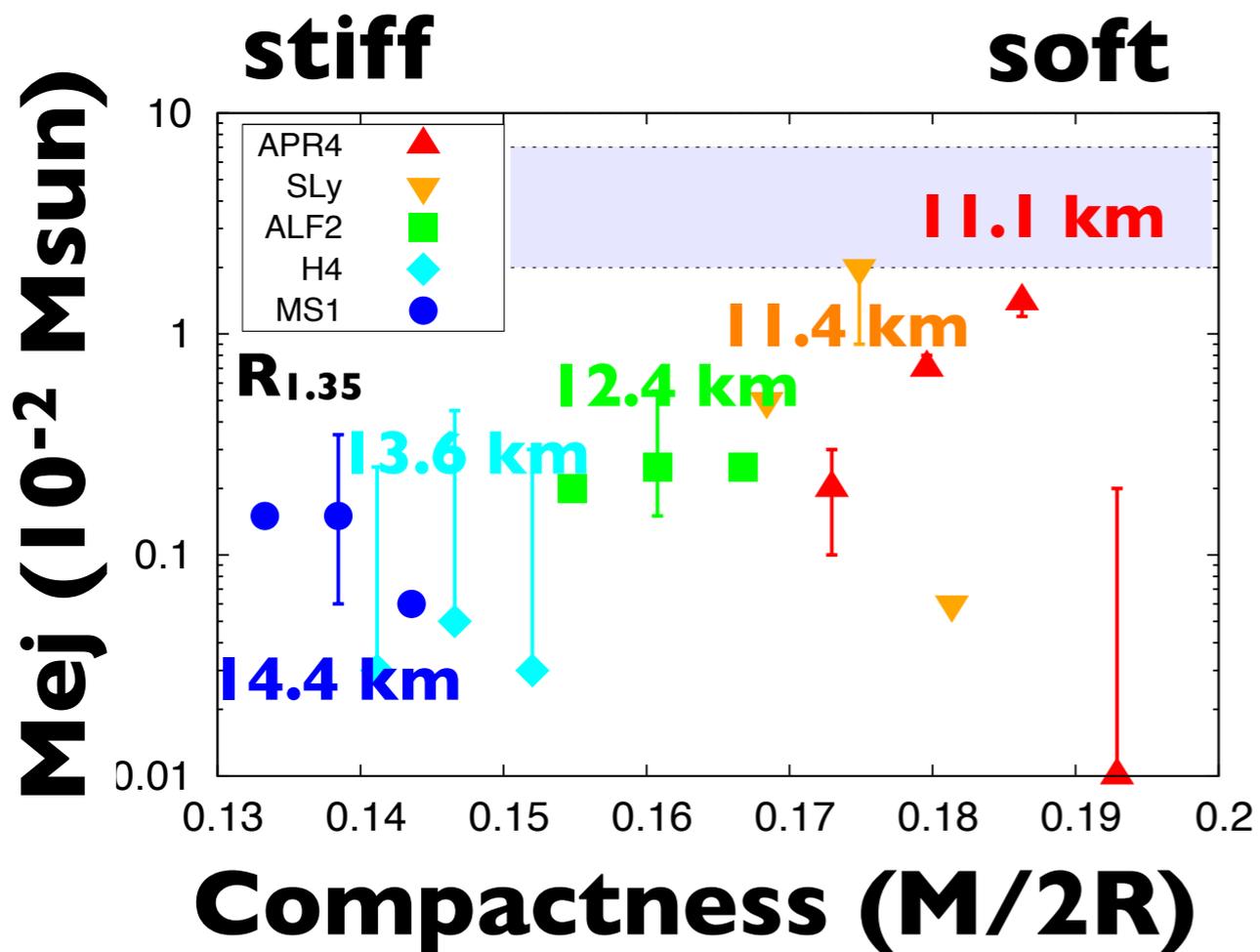
$M_{ej} \sim 0.02 M_{sun}$

NS-NS

=> **soft EOS**

BH-NS

=> **stiff EOS**



Summary

- **EM emission from NS-NS/BH-NS mergers**
 - **Accurate localization of GW sources**
- **“Full” radiative transfer simulations**
 - **Higher opacity than Fe by a factor of 100**
 - **SED peak at near-IR**
 - **Bluer emission from BH-NS mergers**
- **Observing strategy**
 - **22-25 mag (i band) => 4-8m class telescopes**
 - **Extremely broad-line spectra**
- **GRB 130603B**
 - **Constraints on progenitors and EOSs**