

LAMBDA-NUCLEAR INTERACTION and **HYPERON PUZZLE in NEUTRON STARS**



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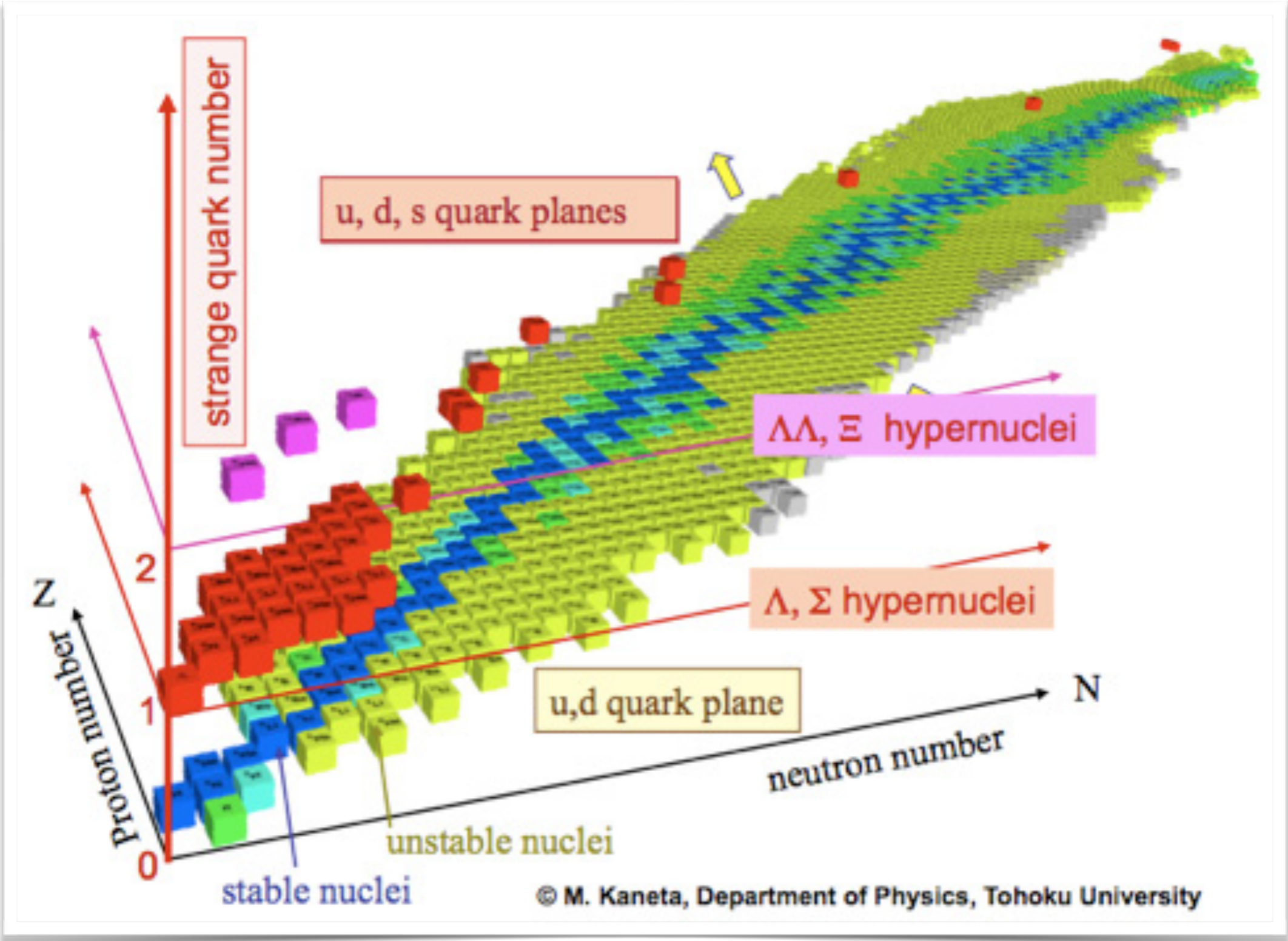


- **Equation of State of dense baryonic matter : constraints from massive neutron stars**
- **Hyperon-nucleon interactions from SU(3) chiral effective field theory**
- **Hyperon-NN three-body forces**
- **Emerging repulsions : suppression of hyperons in dense neutron matter**

Stefan Petschauer, Johann Haidenbauer, et al. :

Eur. Phys. J. **A** (2017) (arXiv: 1612.03758 [nucl-th]); Nucl. Phys. **A 957** (2017) 347;

Phys. Rev. **C93** (2016) 014001; Eur. Phys. J. **A52** (2016)15; Nucl. Phys. **A915** (2013) 24



Part I: Prologue

*Constraints on Equations of State
from
massive Neutron Stars*

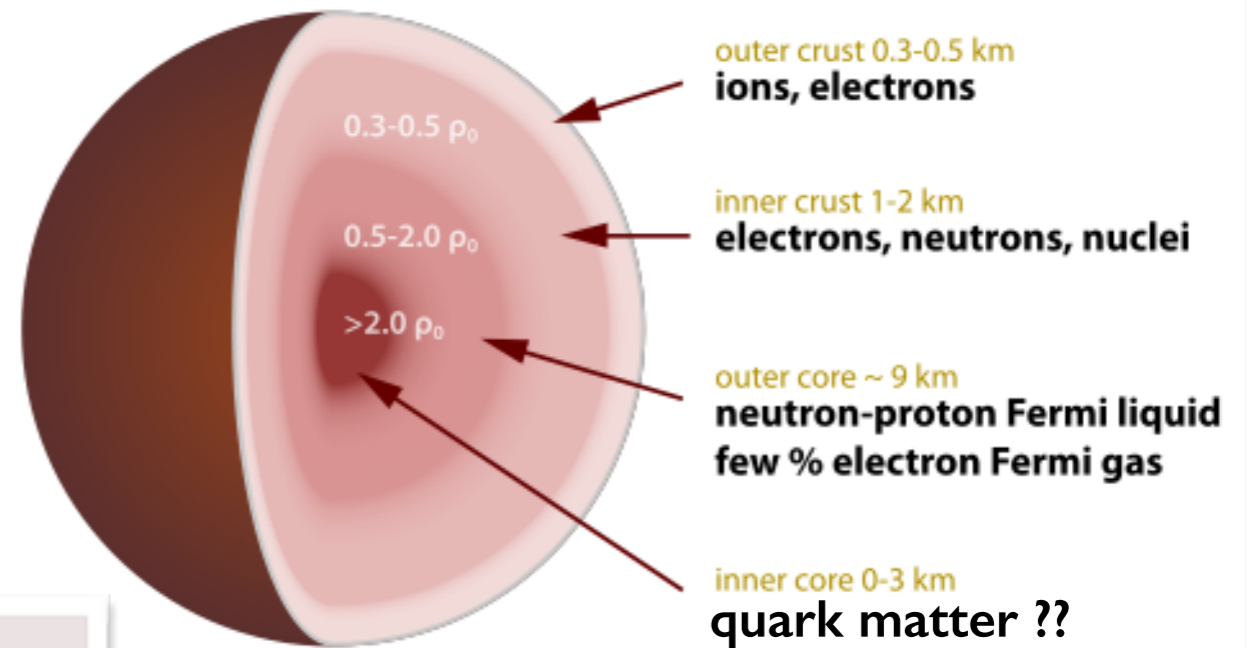
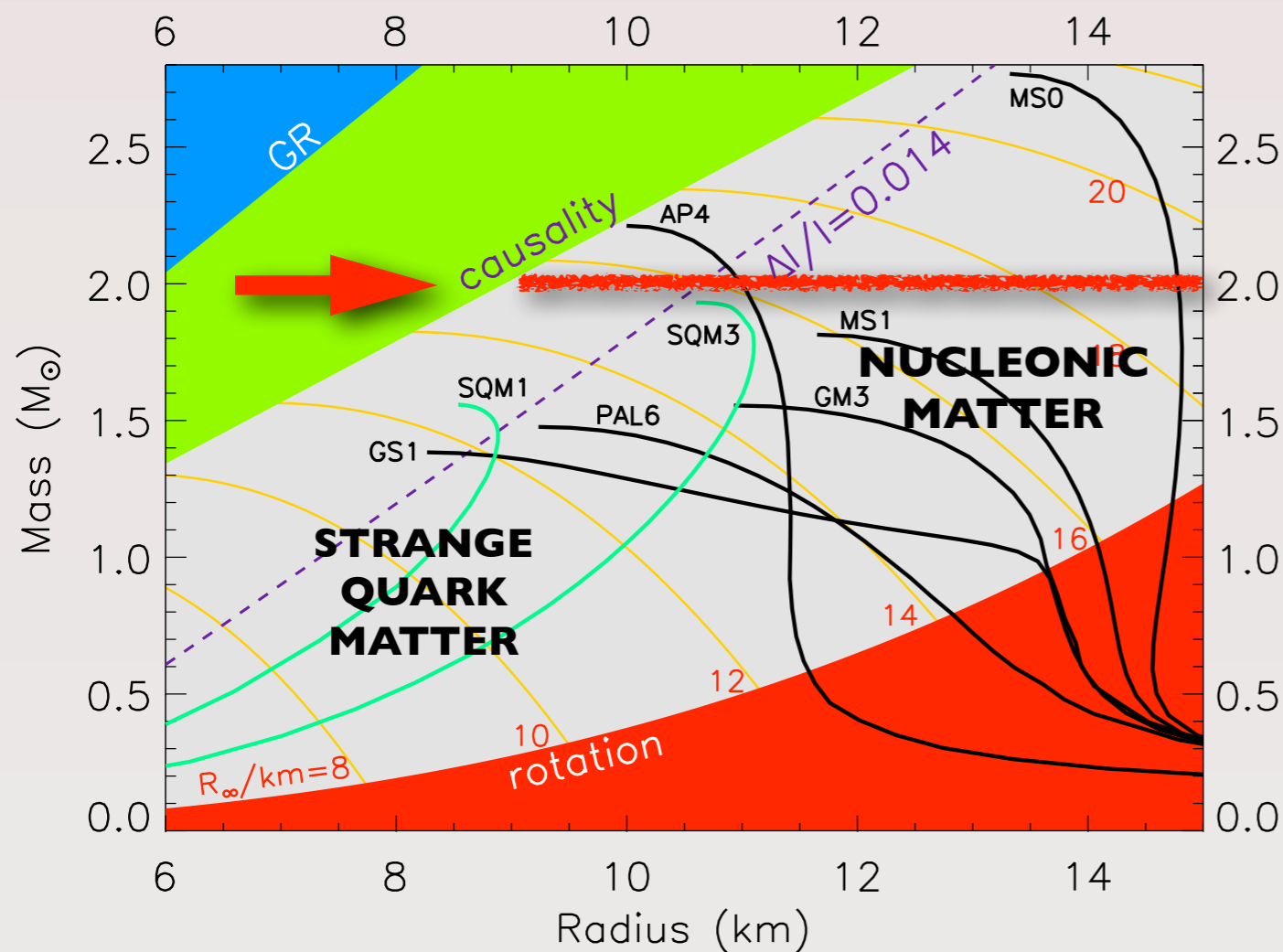


NEUTRON STARS and the EQUATION OF STATE of DENSE BARYONIC MATTER

J. Lattimer, M. Prakash

Phys. Reports 442 (2007) 109 Phys. Reports 621(2016) 127

Mass-Radius Relation



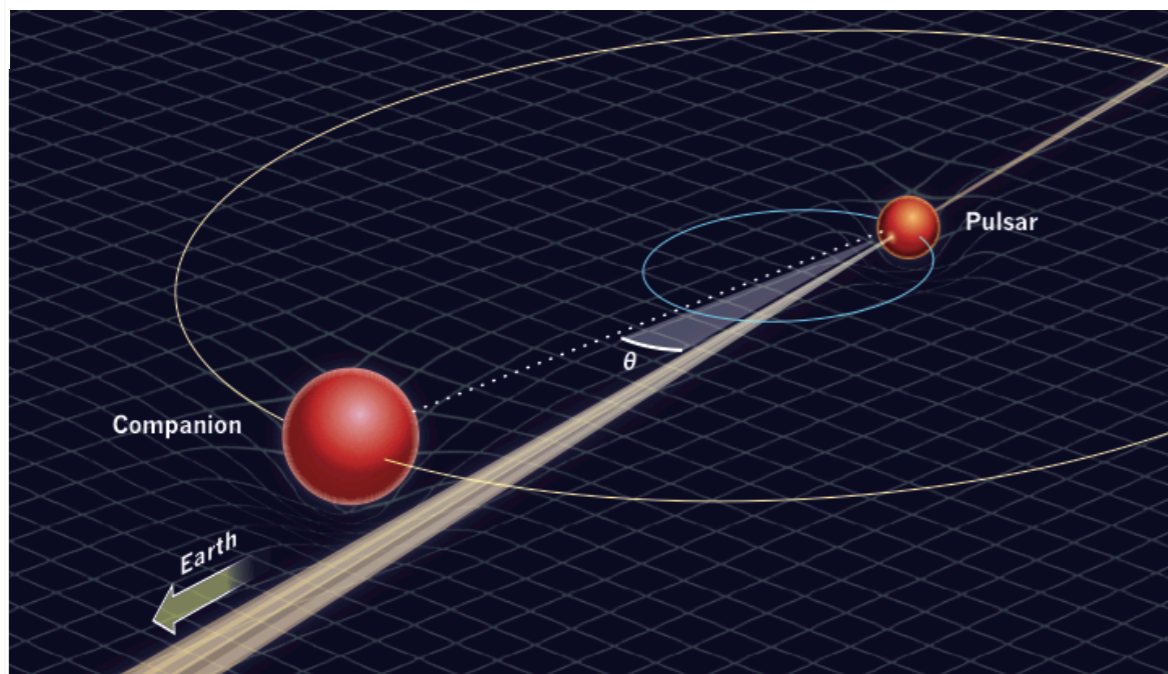
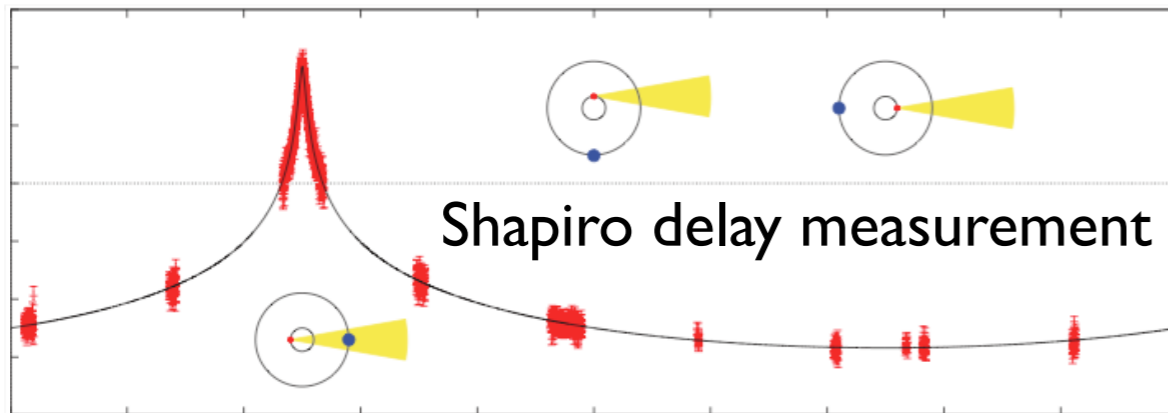
Tolman-Oppenheimer-Volkov Equations

$$\frac{dP}{dr} = -\frac{G}{c^2} \frac{(M + 4\pi Pr^3)(\mathcal{E} + P)}{r(r - GM/c^2)}$$

$$\frac{dM}{dr} = 4\pi r^2 \frac{\mathcal{E}}{c^2}$$

Constraints from massive NEUTRON STARS

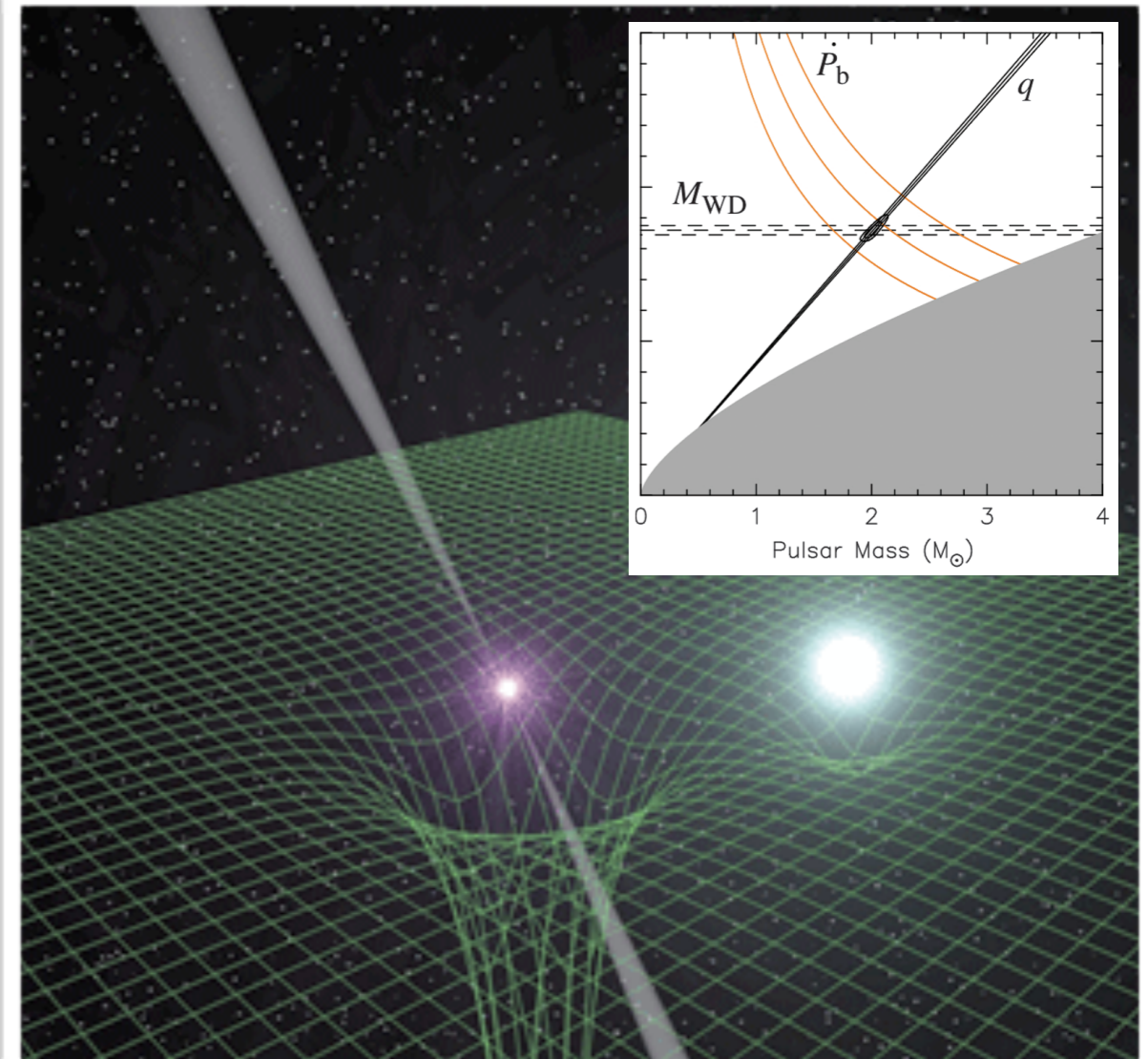
P.B. Demorest et al.
Nature 467 (2010) 1081



PSR J1614+2230

$$M = 1.97 \pm 0.04 M_{\odot}$$

J. Antoniadis et al.
Science 340 (2013) 6131

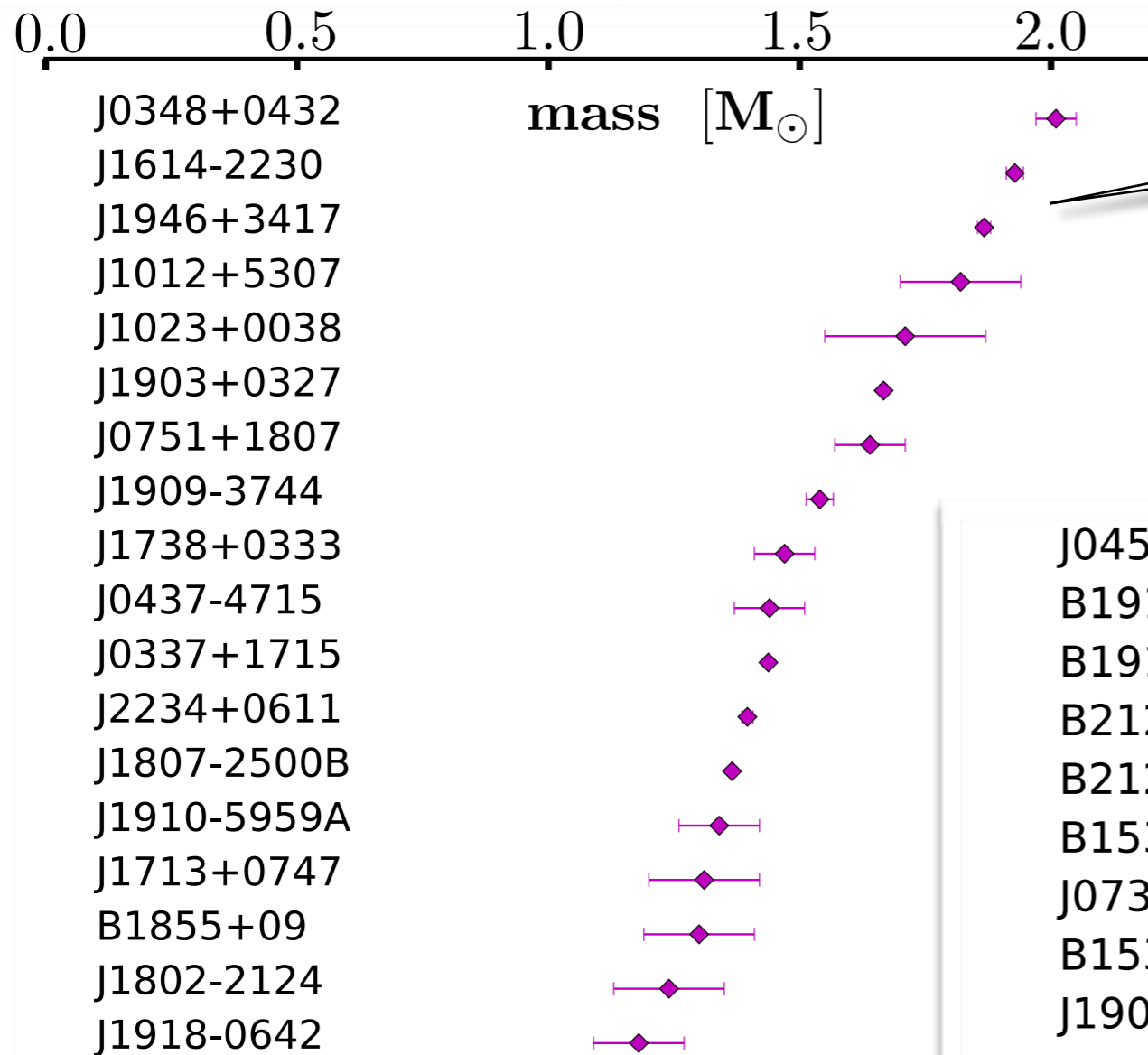


PSR J0348+0432

$$M = 2.01 \pm 0.04 M_{\odot}$$

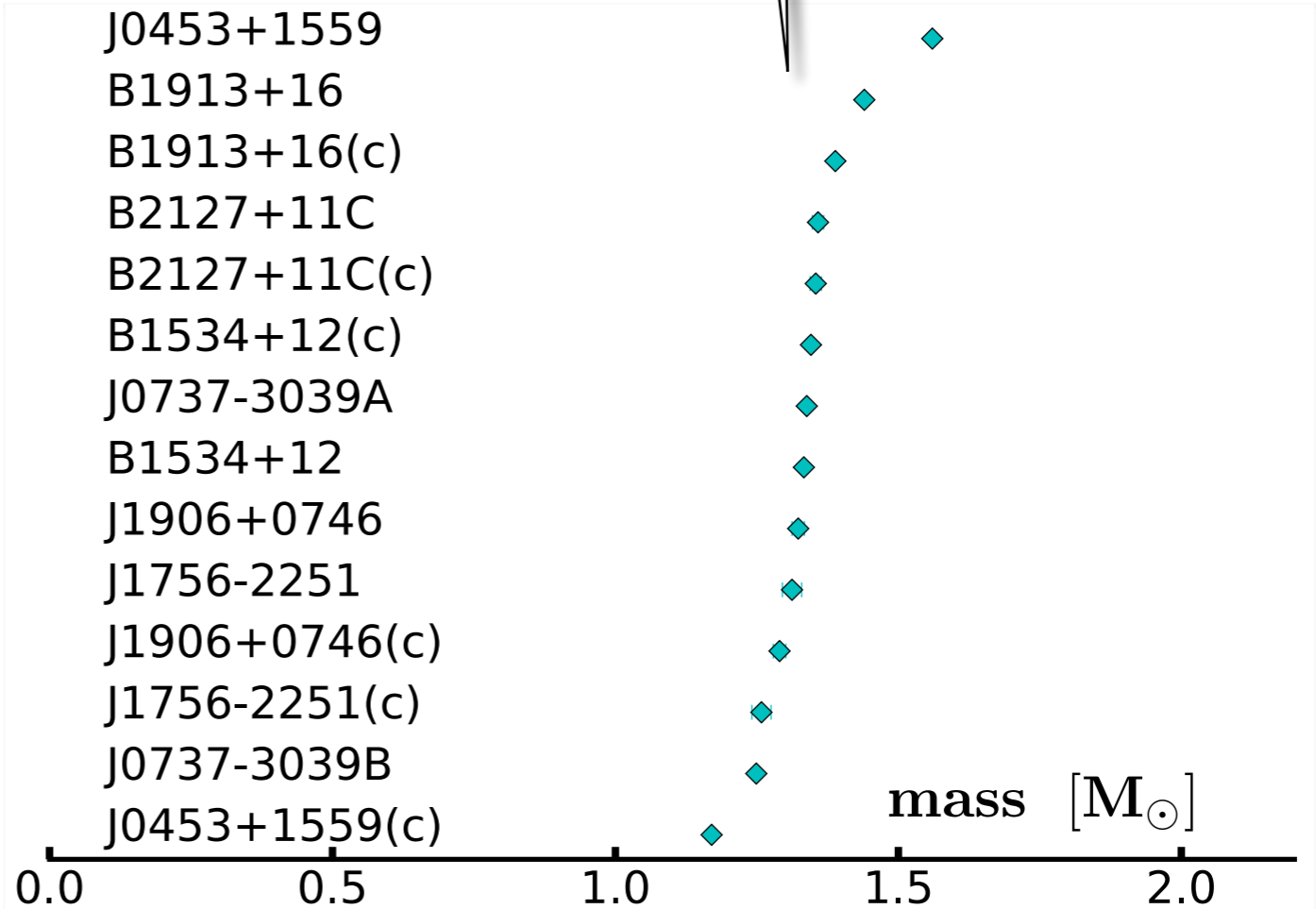
Population of MILLISECOND PULSARS

J. Antoniadis et al. arXiv:1605.01665



ms pulsars in binaries
e.g. with white dwarfs

double neutron stars

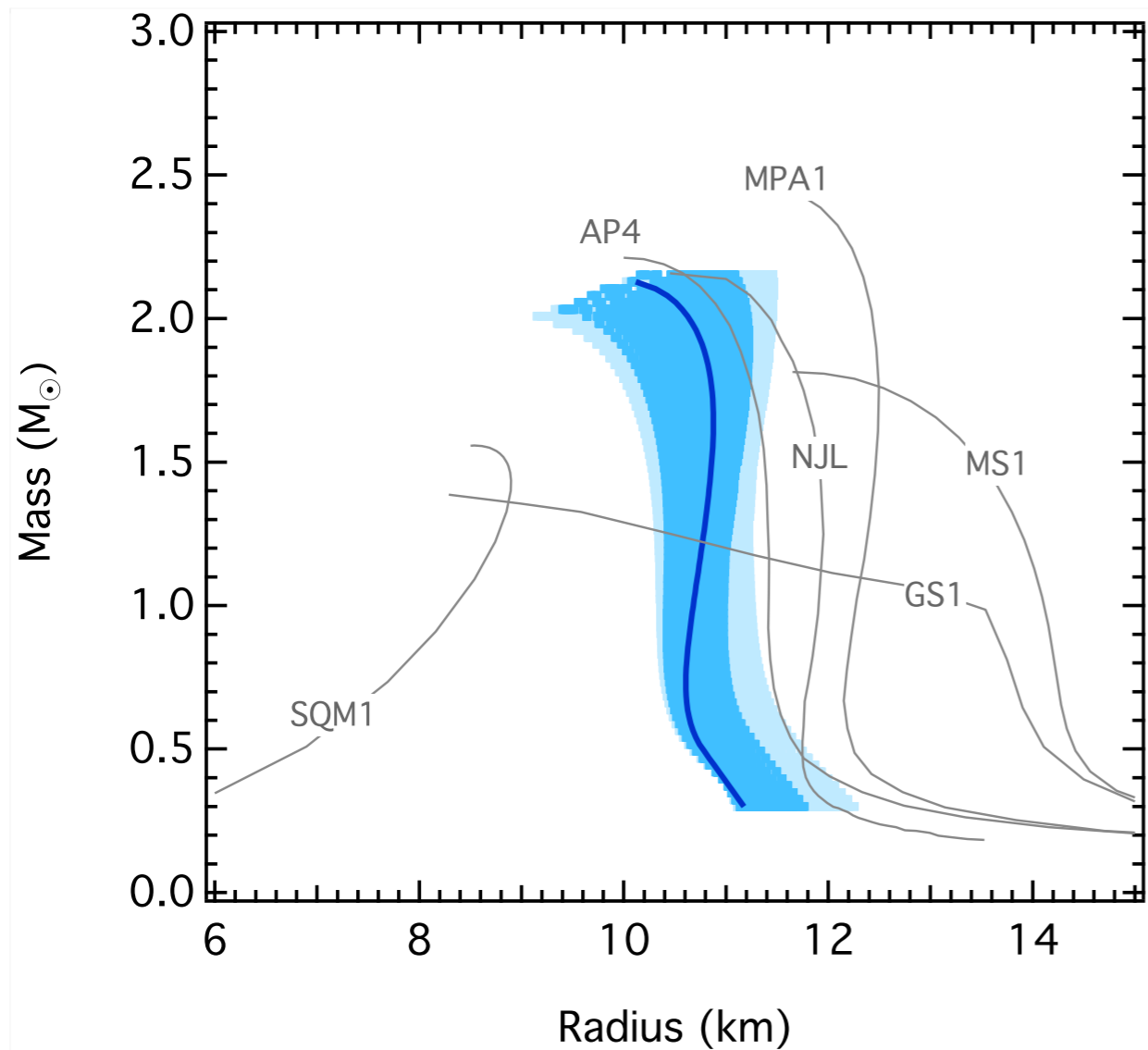


Note:
about **25%** of population are
massive n-stars ($M > 1.5 M_{\odot}$)

CONSTRAINTS from NEUTRON STARS

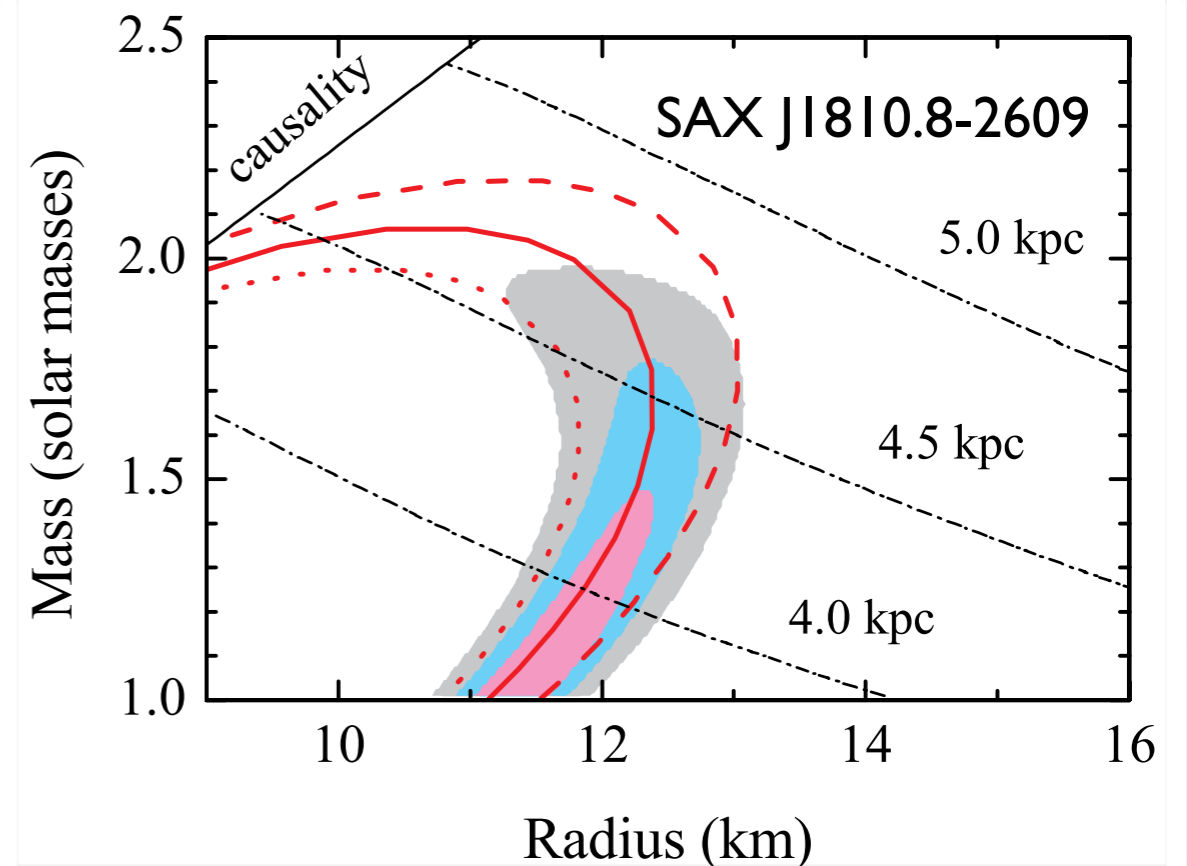
(contd.)

- Comprehensive analysis of 12 selected neutron stars



F. Özel, D. Psaltis, T. Güver, G. Baym, C. Heinke, S. Guillot
Astroph. J. 820 (2016) 28

- Atmosphere model fits of X-ray bursts



$$R = (11.5 - 13.0) \text{ km}$$

$$(M = 1.3 - 1.8 M_{solar})$$

J. Nättilä et al. : Astron. Astroph. 591 (2016) A25
V.F. Suleimanov et al. : arXiv:1611.09885

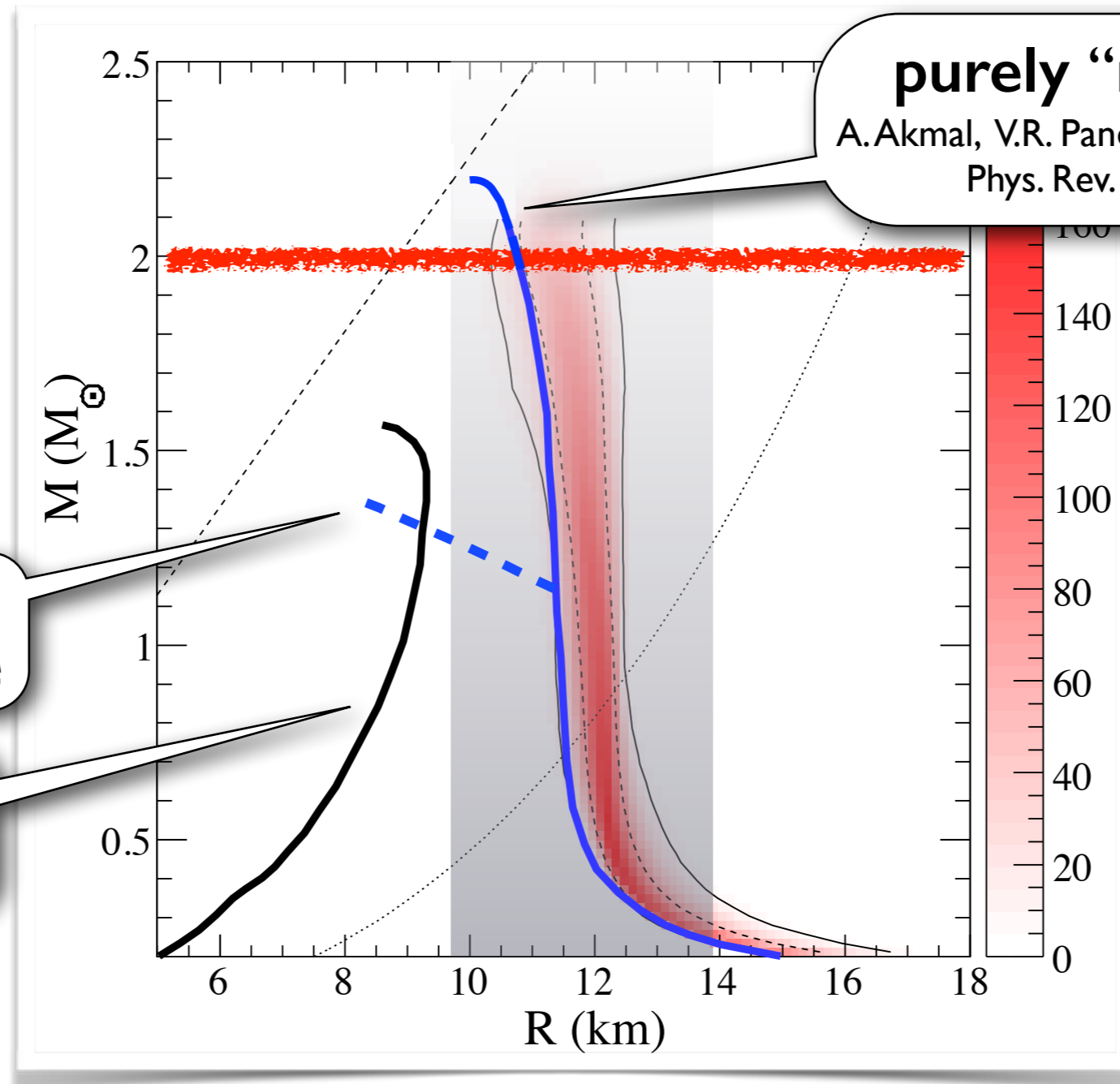
CONSTRAINTS from NEUTRON STARS

F. Özel, D. Psaltis: Phys. Rev. D80 (2009) 103003
 F. Özel, G. Baym, T. Güver: Phys. Rev. D82 (2010) 101301

Mass-Radius Relation

kaon condensate

quark matter



purely “nuclear” EoS
 A.Akmal, V.R. Pandharipande, D.G. Ravenhall
 Phys. Rev. C 58 (1998) 1804

K. Hebeler,
 J. Lattimer,
 Ch. Pethick,
 A. Schwenk:
 Phys. Rev. Lett.
 105 (2010) 161102

A.W. Steiner,
 J. Lattimer, E.F. Brown
 Astroph. J. 722 (2010) 33

- “Exotic” equations of state ruled out ?

NEUTRON STAR MATTER

from **Chiral EFT** and **FRG**

● Symmetry energy range: 30 - 35 MeV

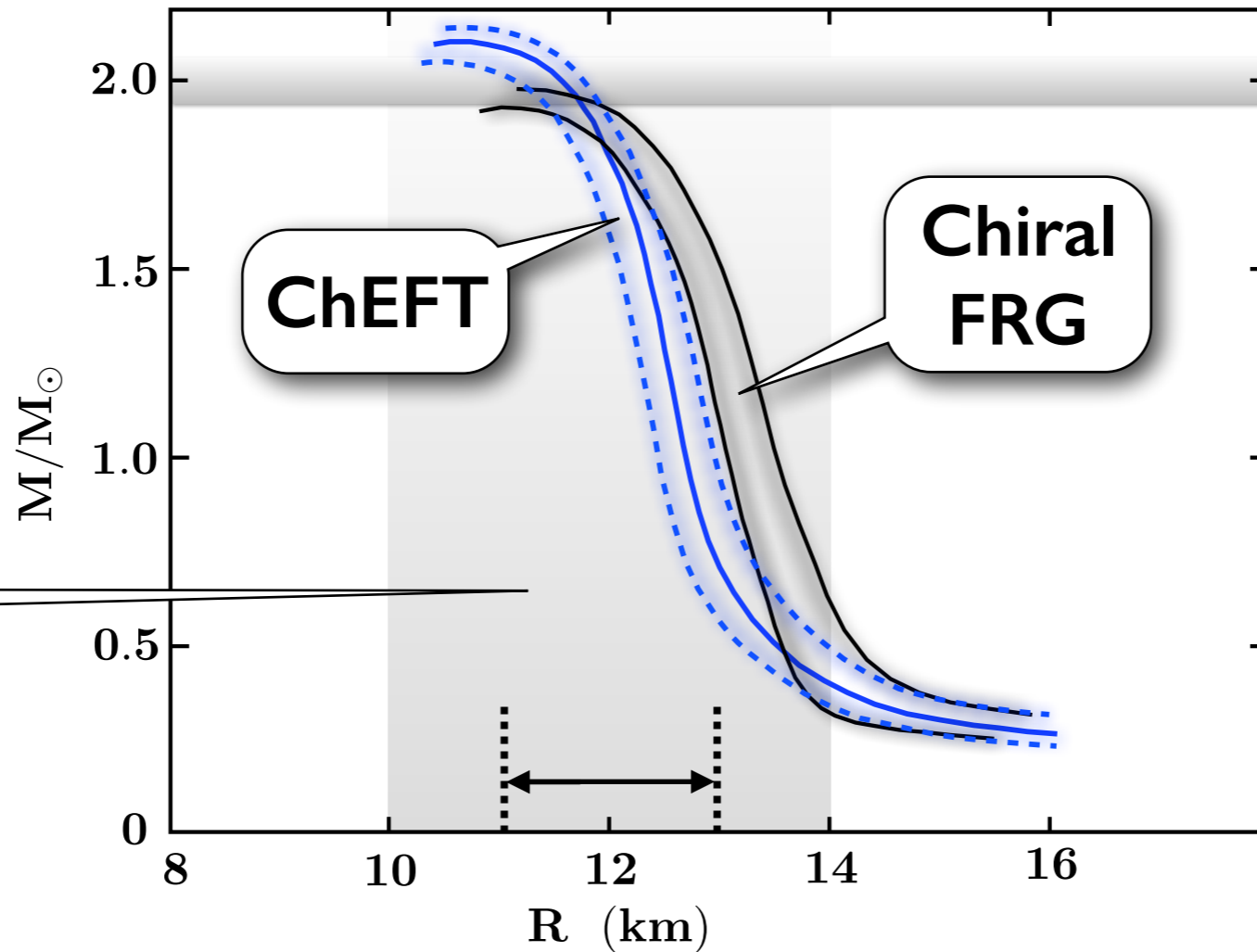
● Crust: SLy EoS

T. Hell, W.W.
Phys. Rev.
C90 (2014) 045801

M. Drews, W.W.

Phys. Rev.
C91 (2015) 035802

Prog. Part. Nucl. Phys.
93 (2017) 69



● Central core density

$$\rho_c \lesssim 5 \rho_0$$

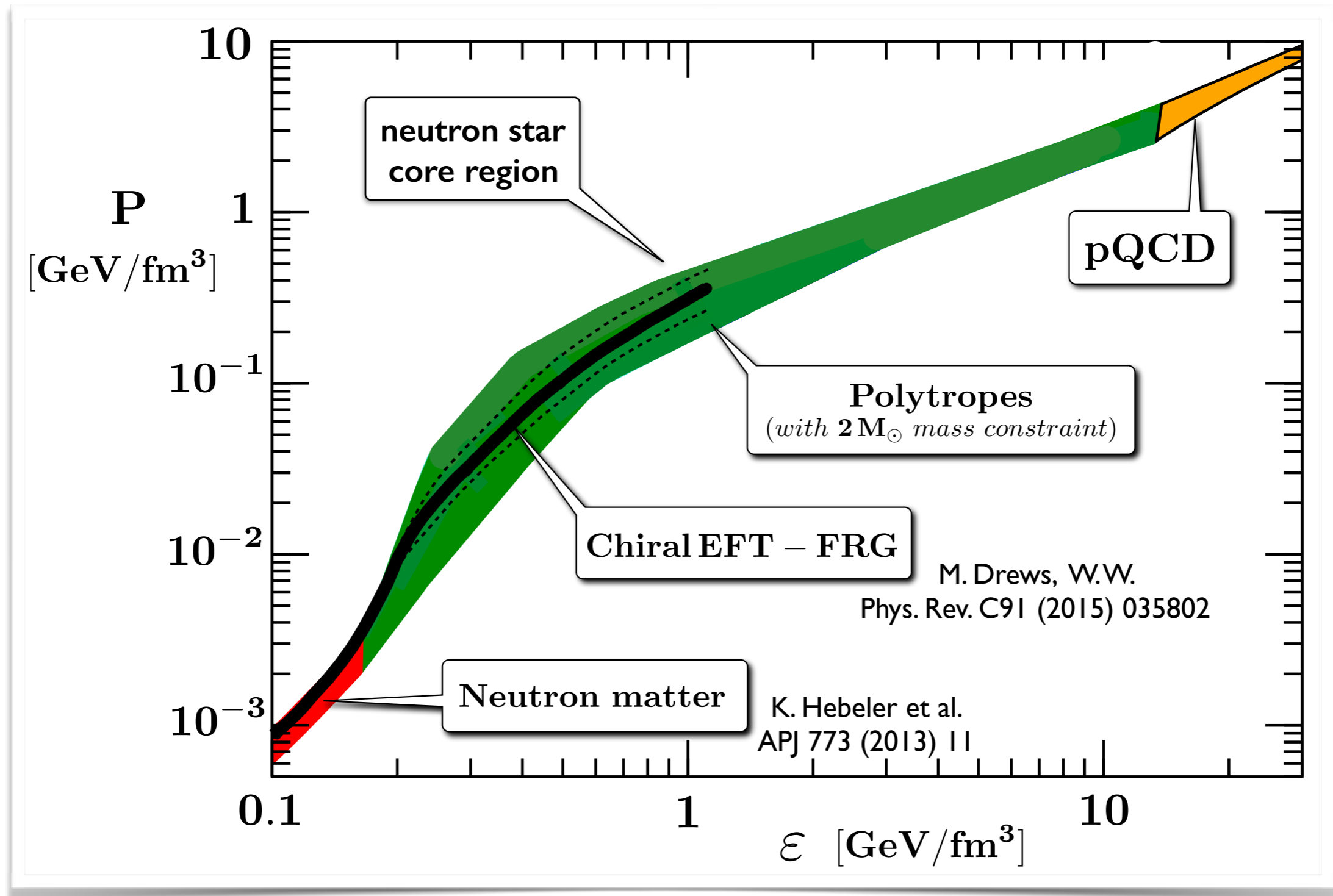
A.W. Steiner,
J.M. Lattimer, E.F. Brown
EPJ A52 (2016) 18

● Chiral many-body dynamics using “conventional” (pion & nucleon) degrees of freedom is consistent with neutron star constraints

NEUTRON STAR MATTER Equation of State

... and extrapolation
to PQCD limit

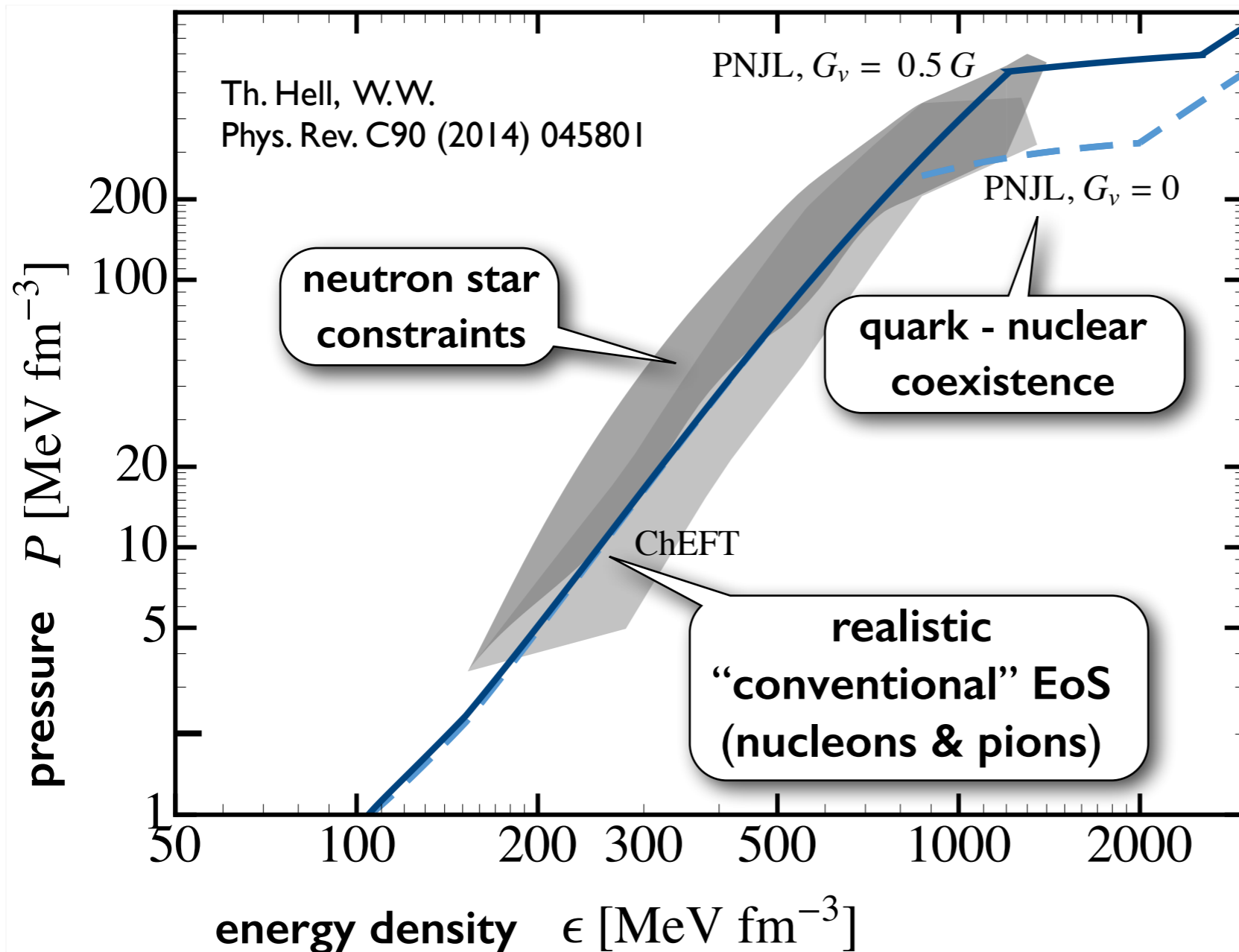
A. Kurkela et al.
Astroph. J. 789 (2014) 127



NEUTRON STAR MATTER

Equation of State

- In-medium **Chiral Effective Field Theory** up to 3 loops (reproducing thermodynamics of normal nuclear matter)
- 3-flavor PNJL (chiral quark) model at high densities (incl. strange quarks)



- conventional (hadronic) equation of state seems to work
- quark-nuclear coexistence can occur at baryon densities

$$\rho > 5 \rho_0$$

$$(\rho_0 = 0.16 \text{ fm}^{-3})$$

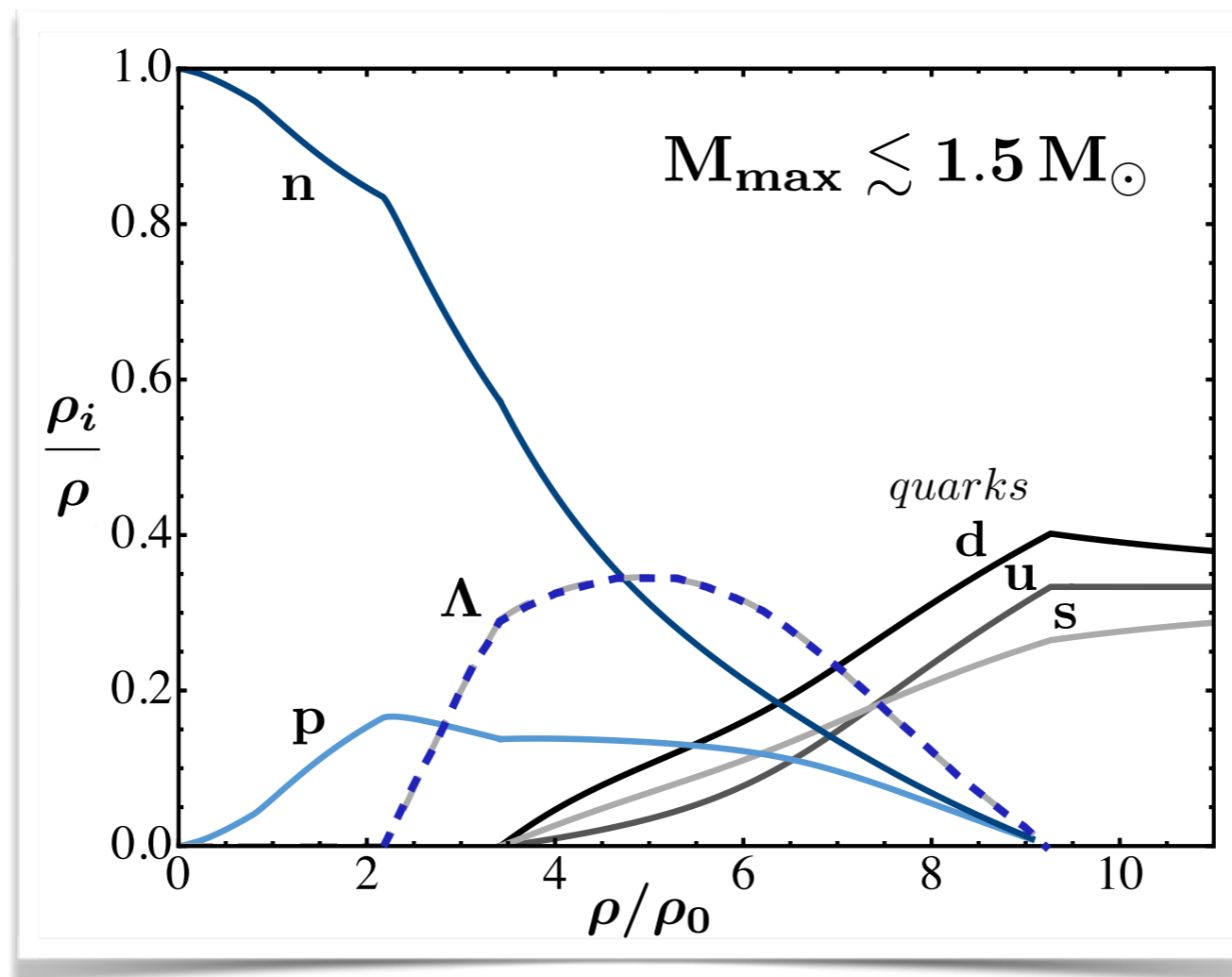
see also:

K. Masuda, T. Hatsuda, T. Takatsuka
PTEP (2013) 7, 073D01

NEUTRON STAR MATTER including **HYPERONS**

- In-medium **Chiral Effective Field Theory** (3-loops) plus Λ hyperons (incl. potential consistent with hypernuclei)
- 3-flavor **PNJL** model at high densities (incl. strange quarks)

Particle composition:
Fraction of particle species as function of baryon density

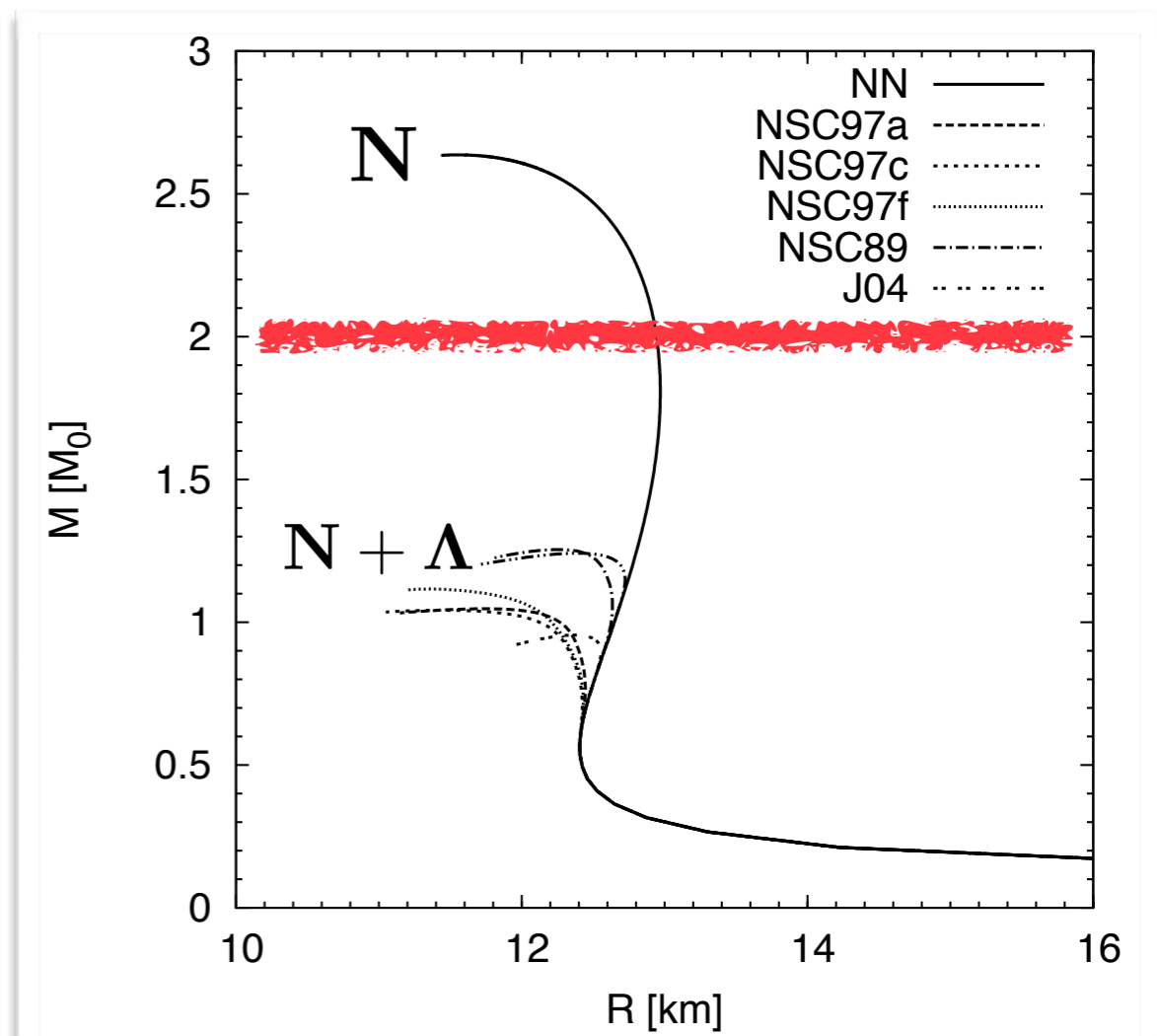
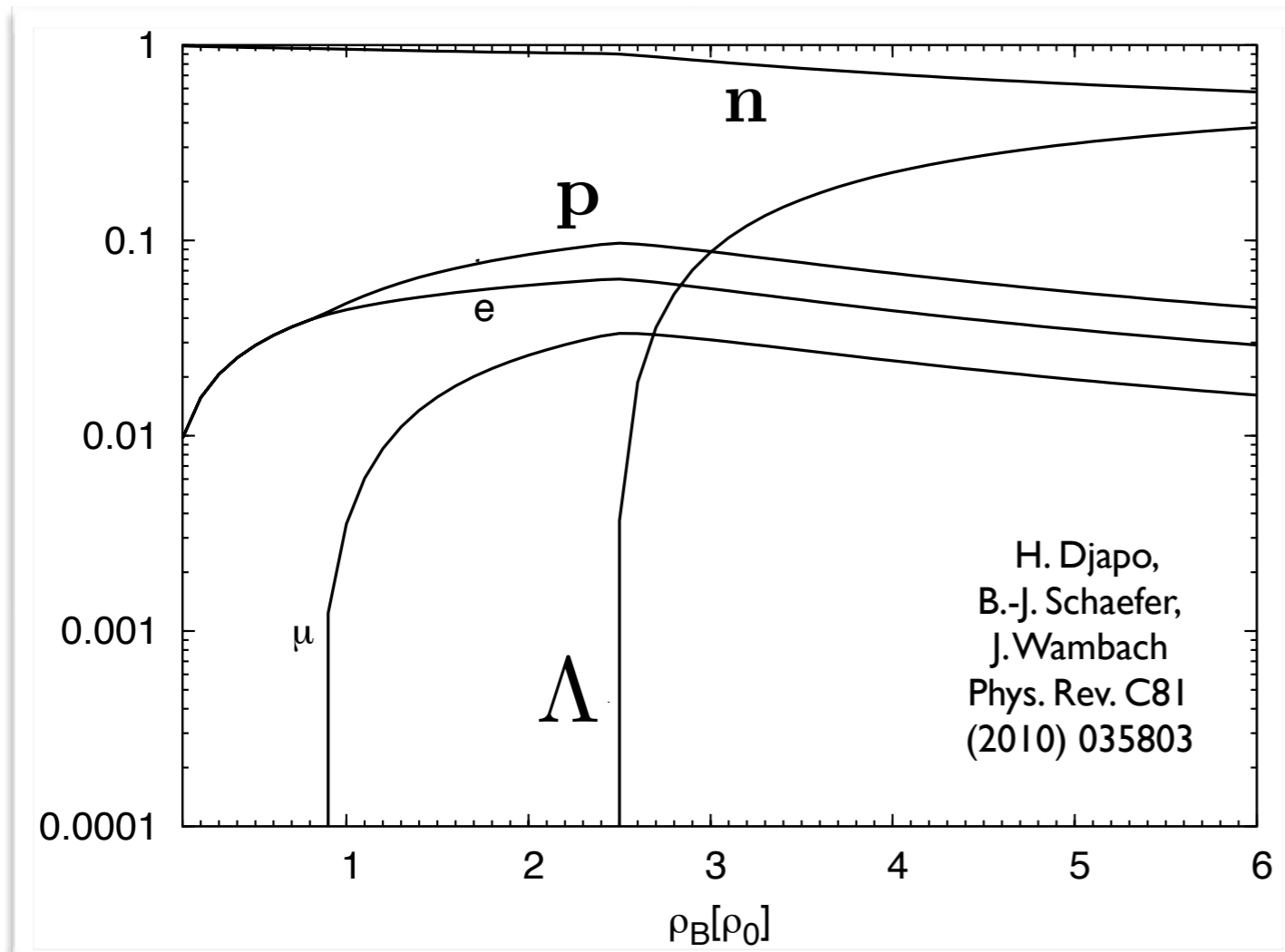


T. Hell, W.W.
 Phys. Rev. C90 (2014) 045801

occurrence of
 Λ hyperons
 $\mu_n = \mu_{\Lambda}$

- **Equation of state too soft** : maximum neutron star mass too low

NEUTRON STAR MATTER including **HYPERONS**



- Adding hyperons: equation of state far too soft
“Hyperon Puzzle”



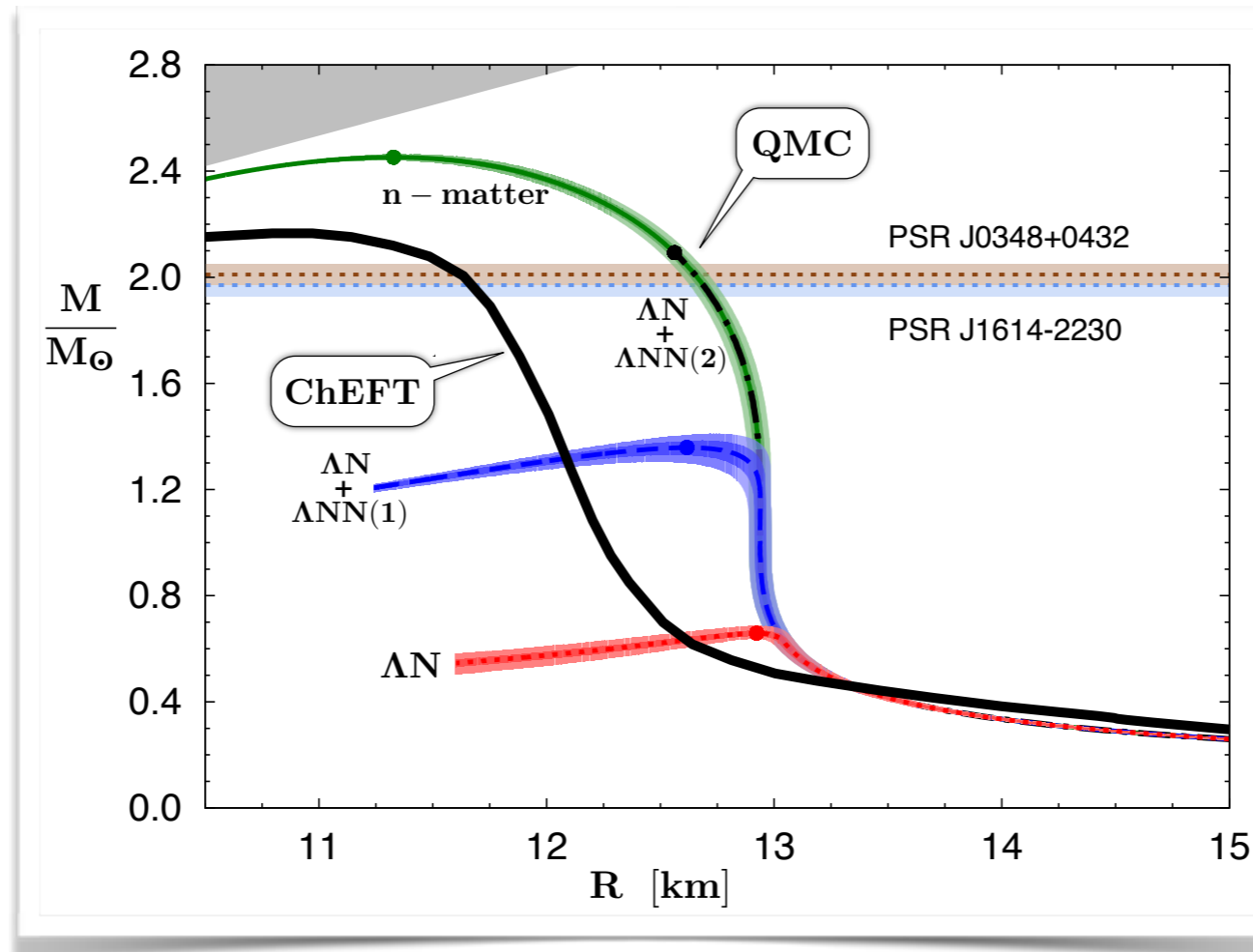
NEUTRON STAR MATTER including **HYPERONS**

Quantum Monte Carlo calculations using phenomenological hyperon-nucleon and hyperon-NN three-body interactions constrained by hypernuclei

ChEFT
calculations
“conventional”
n-star matter

—

T. Hell, W.W.
PRC90 (2014) 045801



QMC
computations
(hyper-neutron matter):

D. Lonardoni,
A. Lovato,
S. Gandolfi,
F. Pederiva
Phys. Rev. Lett.
114 (2015) 092301

Inclusion of hyperons: EoS too soft to support 2-solar-mass n-stars
unless: strong repulsion in **YN** and **YNN** ... interactions

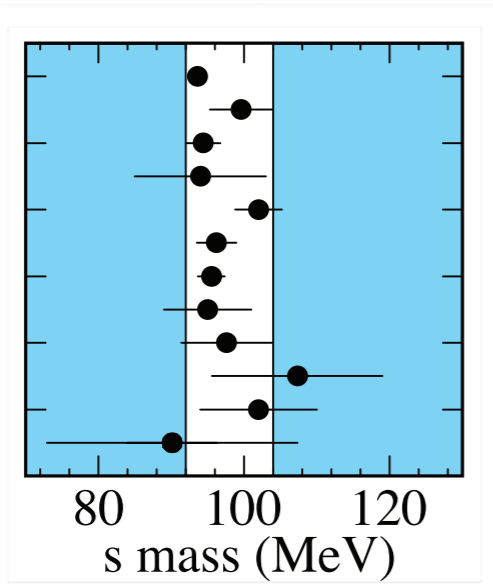
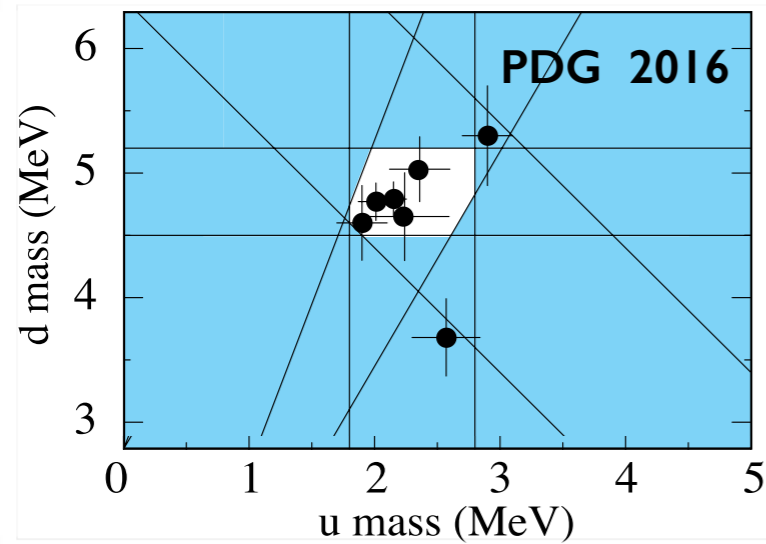
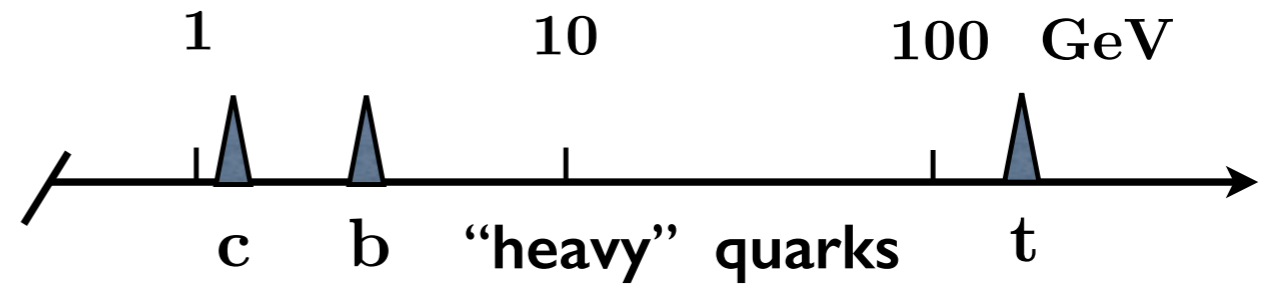
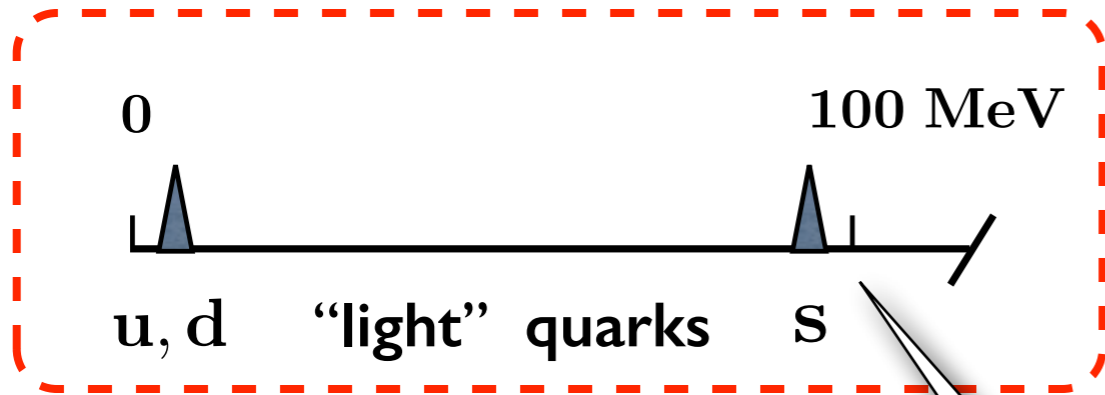
Part II

*Hyperon - Nucleon Interactions
from
Chiral $SU(3)$ Effective Field Theory*



Hierarchy of **QUARK MASSES** in **QCD**

- **Separation of Scales** -



Basic principles of LOW-ENERGY QCD :

Confinement of quarks & gluons in hadrons

Chiral Symmetry
 $SU(3)_L \times SU(3)_R$

Spontaneously broken (QCD dynamics)	Explicitly broken by non-zero quark masses
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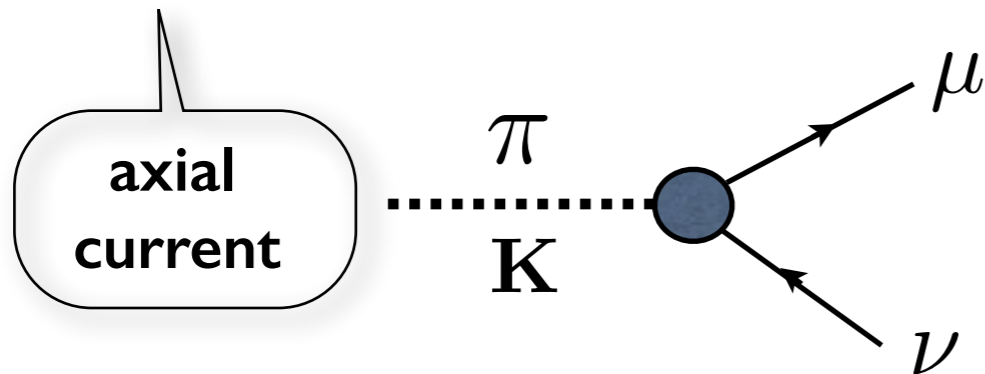
Spontaneously Broken CHIRAL $SU(3)_L \times SU(3)_R$ SYMMETRY

- NAMBU - GOLDSTONE BOSONS:**

Pseudoscalar $SU(3)$ meson octet $\{\phi_a\} = \{\pi, \mathbf{K}, \bar{\mathbf{K}}, \eta_8\}$

- DECAY CONSTANTS:**

$$\langle 0 | \mathbf{A}_a^\mu(0) | \phi_b(p) \rangle = i \delta_{ab} p^\mu f_b$$



Chiral limit: $f = 86.2 \text{ MeV}$

Order parameter :
 $4\pi f \sim 1 \text{ GeV}$

$$f_\pi = 92.21 \pm 0.16 \text{ MeV}$$

$$f_{\mathbf{K}} = 110.5 \pm 0.5 \text{ MeV}$$

- Gell-Mann,
Oakes,
Renner
relations**

$$m_\pi^2 f_\pi^2 = -\frac{m_u + m_d}{2} \langle \bar{u}u + \bar{d}d \rangle + \text{higher order corrections}$$

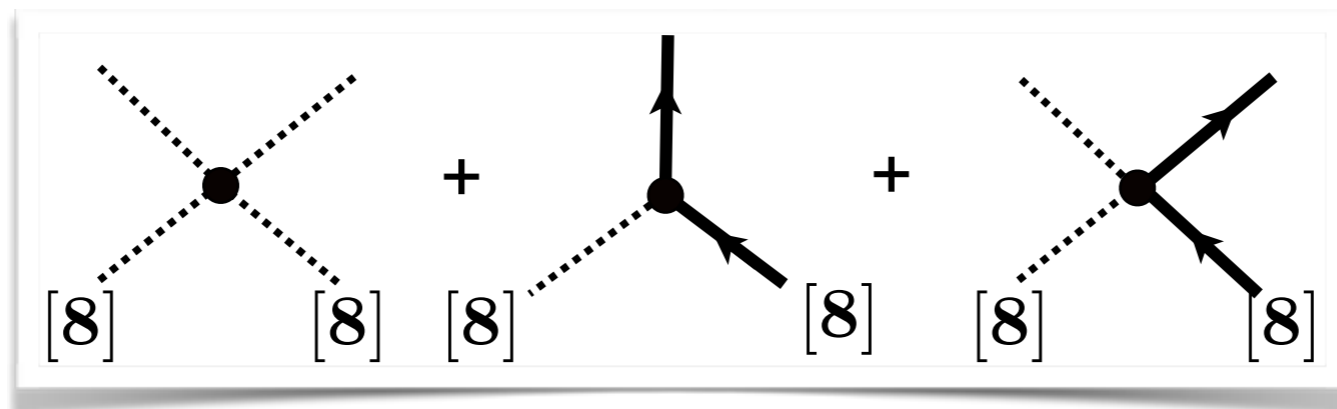
$$m_{\mathbf{K}}^2 f_{\mathbf{K}}^2 = -\frac{m_u + m_s}{2} \langle \bar{u}u + \bar{s}s \rangle$$

Chiral $SU(3)_L \times SU(3)_R$ Effective Field Theory

- Realization of **Low-Energy QCD** for energies / momenta
 $Q < 4\pi f \sim 1 \text{ GeV}$
- based on $SU(3)$ **Non-Linear Sigma Model** plus (heavy) **baryons**
- **Pseudoscalar meson octet** of $SU(3)_L \times SU(3)_R$
Nambu-Goldstone bosons coupled to **baryon octet**

$$P = \begin{pmatrix} \frac{\pi^0}{\sqrt{2}} + \frac{\eta}{\sqrt{6}} & \pi^+ & K^+ \\ \pi^- & -\frac{\pi^0}{\sqrt{2}} + \frac{\eta}{\sqrt{6}} & K^0 \\ K^- & \bar{K}^0 & -\frac{2\eta}{\sqrt{6}} \end{pmatrix} \quad B = \begin{pmatrix} \frac{\Sigma^0}{\sqrt{2}} + \frac{\Lambda}{\sqrt{6}} & \Sigma^+ & p \\ \Sigma^- & -\frac{\Sigma^0}{\sqrt{2}} + \frac{\Lambda}{\sqrt{6}} & n \\ -\Xi^- & \Xi^0 & -\frac{2\Lambda}{\sqrt{6}} \end{pmatrix}$$

meson-baryon
interaction
vertices



+ ... short distance
dynamics:
contact terms

Chiral $SU(3)_L \times SU(3)_R$ Effective Field Theory

- Starting point: **Meson-Baryon Lagrangian** (chiral limit)

$$\mathcal{L}_{\text{MB}} = \text{tr} \left(\bar{B} \left(i\gamma^\mu D_\mu - M_0 \right) B \right) - \frac{D}{2} \text{tr} \left(\bar{B} \gamma^\mu \gamma_5 \{u_\mu, B\} \right) - \frac{F}{2} \text{tr} \left(\bar{B} \gamma^\mu \gamma_5 [u_\mu, B] \right)$$

$$P = \begin{pmatrix} \frac{\pi^0}{\sqrt{2}} + \frac{\eta}{\sqrt{6}} & \pi^+ & K^+ \\ \pi^- & -\frac{\pi^0}{\sqrt{2}} + \frac{\eta}{\sqrt{6}} & K^0 \\ K^- & \bar{K}^0 & -\frac{2\eta}{\sqrt{6}} \end{pmatrix} \quad B = \begin{pmatrix} \frac{\Sigma^0}{\sqrt{2}} + \frac{\Lambda}{\sqrt{6}} & \Sigma^+ & p \\ \Sigma^- & -\frac{\Sigma^0}{\sqrt{2}} + \frac{\Lambda}{\sqrt{6}} & n \\ -\Xi^- & \Xi^0 & -\frac{2\Lambda}{\sqrt{6}} \end{pmatrix}$$

- Chiral covariant derivative: $D_\mu B = \partial_\mu B + [\Gamma_\mu, B]$

$$\Gamma_\mu = \frac{1}{2}(u^\dagger \partial_\mu u + u \partial_\mu u^\dagger) \quad u_\mu = i(u^\dagger \partial_\mu u - u \partial_\mu u^\dagger)$$

- Chiral (pseudoscalar Nambu-Goldstone boson) field :

$$U(x) = \mathbf{u}^2(x) = \exp \left(i \frac{\sqrt{2} P(x)}{f} \right) \quad \text{transforms as} \quad U \rightarrow R U L^\dagger$$

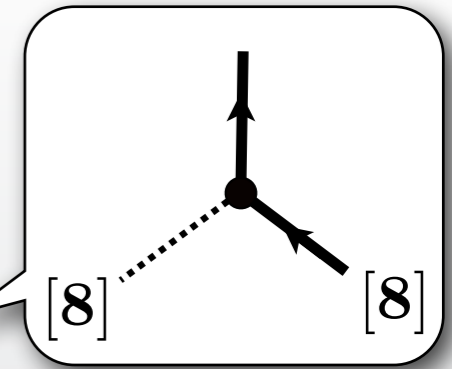
$$R \in SU(3)_R \quad L \in SU(3)_L$$

Chiral $SU(3)_L \times SU(3)_R$ Effective Field Theory

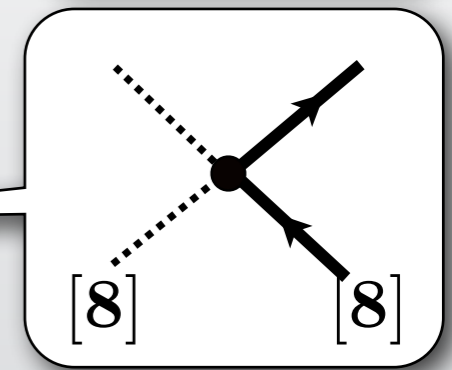
- **Interaction Lagrangian:** expand in powers of meson fields $P(x)$

$$\mathcal{L}_{int} = \mathcal{L}_1 + \mathcal{L}_2 + \dots + \text{mass terms}$$

$$\mathcal{L}_1 = -\frac{\sqrt{2}}{2f} \text{tr}(D\bar{B}\gamma^\mu\gamma_5\{\partial_\mu P, B\} + F\bar{B}\gamma^\mu\gamma_5[\partial_\mu P, B])$$



$$\mathcal{L}_2 = \frac{1}{4f^2} \text{tr}(i\bar{B}\gamma^\mu[[P, \partial_\mu P], B])$$

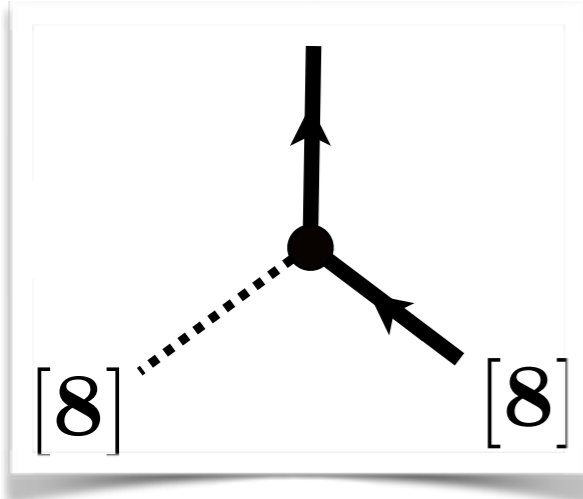


- **Input :** $F = 0.46$ $D = 0.81$ $f = 0.09 \text{ GeV}$
 $(g_A = F + D = 1.27)$
- **Physical meson and baryon masses** ($SU(3)$ breaking)



Chiral $SU(3)_L \times SU(3)_R$ Effective Field Theory : meson-baryon vertices

$$\begin{aligned}
 \mathcal{L}_1 = & -f_{NN\pi} \bar{N} \gamma^\mu \gamma_5 \tau N \cdot \partial_\mu \pi + i f_{\Sigma\Sigma\pi} \bar{\Sigma} \gamma^\mu \gamma_5 \times \Sigma \cdot \partial_\mu \pi \\
 & -f_{\Lambda\Sigma\pi} \left[\bar{\Lambda} \gamma^\mu \gamma_5 \Sigma + \bar{\Sigma} \gamma^\mu \gamma_5 \Lambda \right] \cdot \partial_\mu \pi - f_{\Xi\Xi\pi} \bar{\Xi} \gamma^\mu \gamma_5 \tau \Xi \cdot \partial_\mu \pi \\
 & -f_{\Lambda NK} \left[\bar{N} \gamma^\mu \gamma_5 \Lambda \partial_\mu K + \text{h.c.} \right] - f_{\Xi\Lambda K} \left[\bar{\Xi} \gamma^\mu \gamma_5 \Lambda \partial_\mu \bar{K} + \text{h.c.} \right] \\
 & -f_{\Sigma NK} \left[\bar{N} \gamma^\mu \gamma_5 \tau \partial_\mu K \cdot \Sigma + \text{h.c.} \right] - f_{\Sigma\Xi K} \left[\bar{\Xi} \gamma^\mu \gamma_5 \tau \partial_\mu \bar{K} \cdot \Sigma + \text{h.c.} \right] \\
 & -f_{NN\eta_8} \bar{N} \gamma^\mu \gamma_5 N \partial_\mu \eta - f_{\Lambda\Lambda\eta_8} \bar{\Lambda} \gamma^\mu \gamma_5 \Lambda \partial_\mu \eta \\
 & -f_{\Sigma\Sigma\eta_8} \bar{\Sigma} \cdot \gamma^\mu \gamma_5 \Sigma \partial_\mu \eta - f_{\Xi\Xi\eta_8} \bar{\Xi} \gamma^\mu \gamma_5 \Xi \partial_\mu \eta .
 \end{aligned}$$



$$\begin{array}{lll}
 f_{NN\pi} = G & f_{NN\eta_8} = \frac{1}{\sqrt{3}}(4\alpha - 1)G & f_{\Lambda NK} = -\frac{1}{\sqrt{3}}(1 + 2\alpha)G \\
 f_{\Xi\Xi\pi} = -(1 - 2\alpha)G & f_{\Xi\Xi\eta_8} = -\frac{1}{\sqrt{3}}(1 + 2\alpha)G & f_{\Xi\Lambda K} = \frac{1}{\sqrt{3}}(4\alpha - 1)G \\
 f_{\Lambda\Sigma\pi} = \frac{2}{\sqrt{3}}(1 - \alpha)G & f_{\Sigma\Sigma\eta_8} = \frac{2}{\sqrt{3}}(1 - \alpha)G & f_{\Sigma NK} = (1 - 2\alpha)G \\
 f_{\Sigma\Sigma\pi} = 2\alpha G & f_{\Lambda\Lambda\eta_8} = -\frac{2}{\sqrt{3}}(1 - \alpha)G & f_{\Sigma\Xi K} = -G
 \end{array}$$

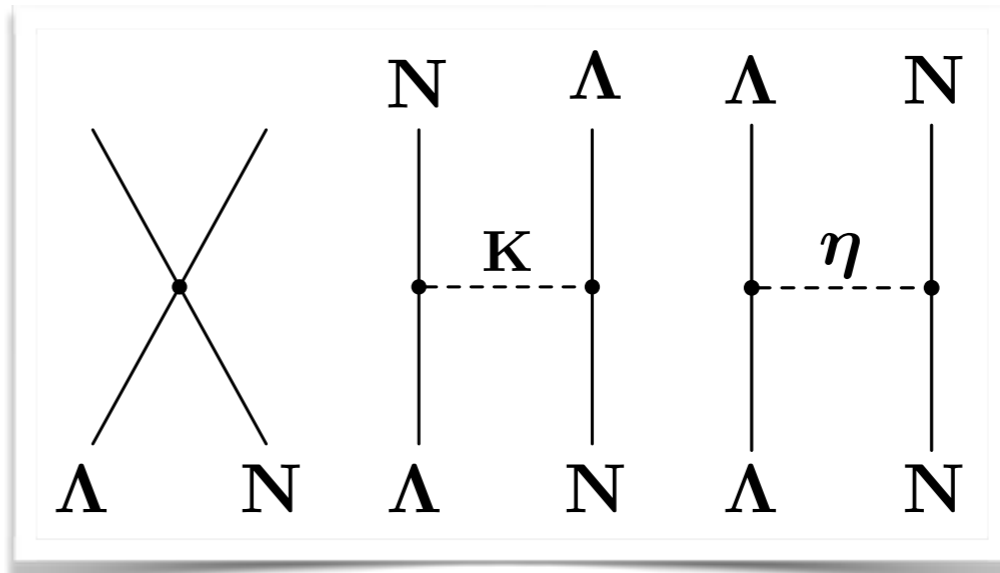
$$G = \frac{g_A}{2f} \simeq 7 \text{ GeV}^{-1} \simeq 1.4 \text{ fm} \qquad \alpha = \frac{F}{F + D} = 0.36$$

Chiral SU(3) Effective Field Theory and Hyperon-Nucleon Interactions

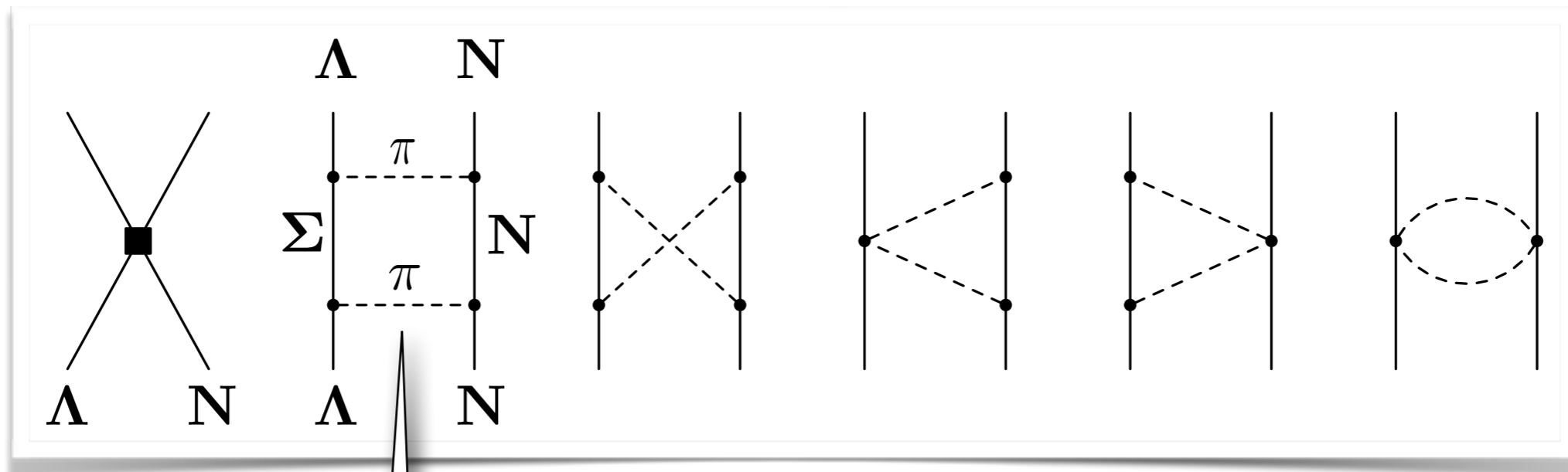
J. Haidenbauer, S. Petschauer, N. Kaiser, U.-G. Meißner, A. Nogga, W.W.: Nucl. Phys. A 915 (2013) 24

Example:
 ΛN
interaction

● Leading order (LO)



● Next-to-leading order (NLO)



2nd order tensor force



Hyperon - Nucleon Interaction Contact Terms



$$V_{BB \rightarrow BB}^{(0)} = C_S + C_T \boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2$$

$$V_{BB \rightarrow BB}^{(2)} = C_1 \mathbf{q}^2 + C_2 \mathbf{k}^2 + (C_3 \mathbf{q}^2 + C_4 \mathbf{k}^2) \boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2 + \frac{i}{2} C_5 (\boldsymbol{\sigma}_1 + \boldsymbol{\sigma}_2) \cdot (\mathbf{q} \times \mathbf{k})$$

$$+ C_6 (\mathbf{q} \cdot \boldsymbol{\sigma}_1) (\mathbf{q} \cdot \boldsymbol{\sigma}_2) + C_7 (\mathbf{k} \cdot \boldsymbol{\sigma}_1) (\mathbf{k} \cdot \boldsymbol{\sigma}_2) + \frac{i}{2} C_8 (\boldsymbol{\sigma}_1 - \boldsymbol{\sigma}_2) \cdot (\mathbf{q} \times \mathbf{k})$$

- **SU(3) symmetry** reduces number of independent constants

$$\mathbf{8} \otimes \mathbf{8} = \mathbf{27} \oplus \mathbf{8}_s \oplus \mathbf{1} \oplus \mathbf{10} \oplus \mathbf{10}^* \oplus \mathbf{8}_a$$

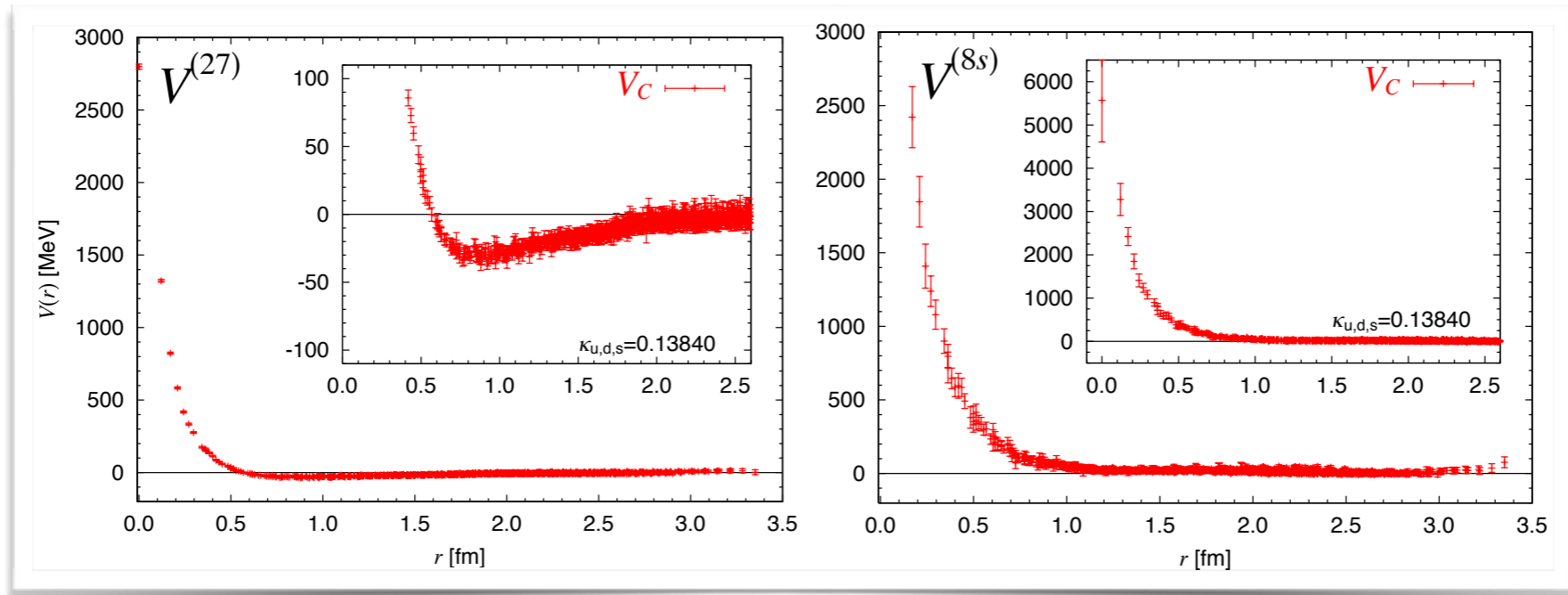
S	Channel	I	$V_{1S_0, 3P_0, 3P_1, 3P_2}$	$V_{3S_1, 3S_1-3D_1, 1P_1}$
0	$NN \rightarrow NN$	0	—	C^{10^*}
	$NN \rightarrow NN$	1	C^{27}	—
-1	$\Lambda N \rightarrow \Lambda N$	$\frac{1}{2}$	$\frac{1}{10} (9C^{27} + C^{8_s})$	$\frac{1}{2} (C^{8_a} + C^{10^*})$
	$\Lambda N \rightarrow \Sigma N$	$\frac{1}{2}$	$\frac{3}{10} (-C^{27} + C^{8_s})$	$\frac{1}{2} (-C^{8_a} + C^{10^*})$
	$\Sigma N \rightarrow \Sigma N$	$\frac{1}{2}$	$\frac{1}{10} (C^{27} + 9C^{8_s})$	$\frac{1}{2} (C^{8_a} + C^{10^*})$
	$\Sigma N \rightarrow \Sigma N$	$\frac{3}{2}$	C^{27}	C^{10}

S. Petschauer,
N. Kaiser
Nucl. Phys.
A 916 (2013) 1-29



Hyperon - Nucleon Interactions from Lattice QCD

$$\Lambda N(^1S_0) = \frac{9}{10}[27] + \frac{1}{10}[8_s]$$



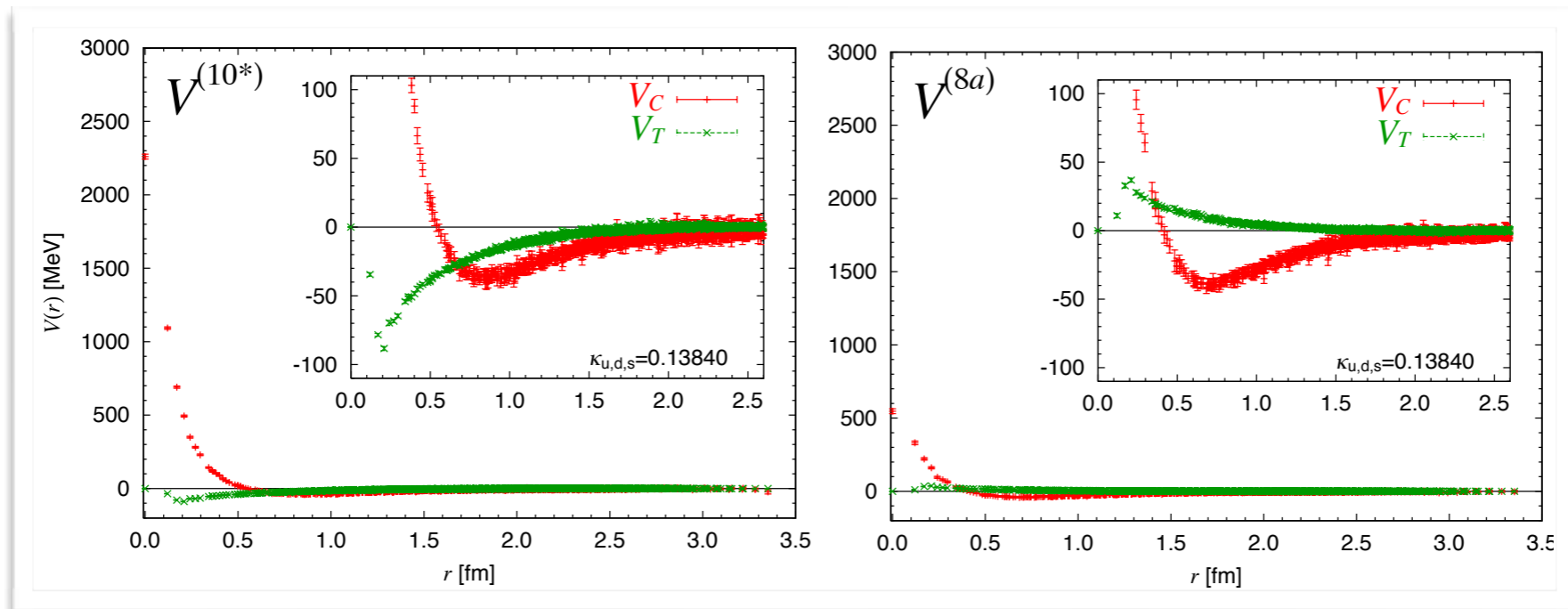
$m_{ps} = 0.47 \text{ GeV}$

T. Inoue et al.

(HAL QCD)
PTP 124 (2010) 591

Nucl. Phys.
A881 (2012) 28


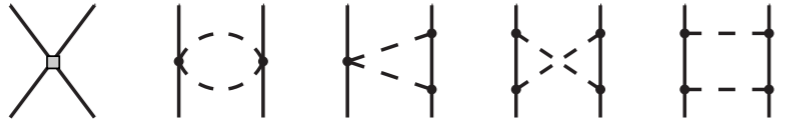
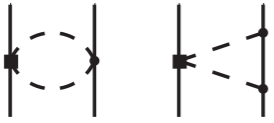
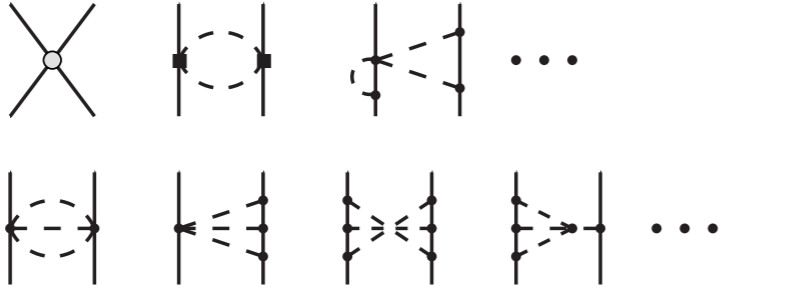
$$\Lambda N(^3S_1) = \frac{1}{2}[10^*] + \frac{1}{2}[8_a]$$



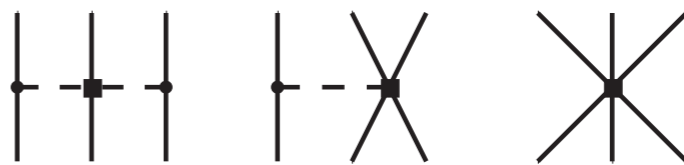
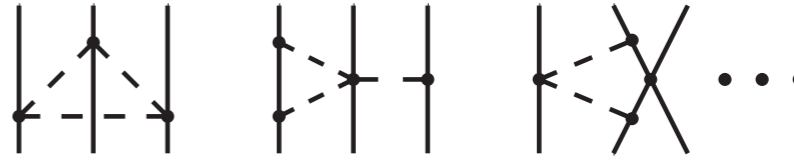
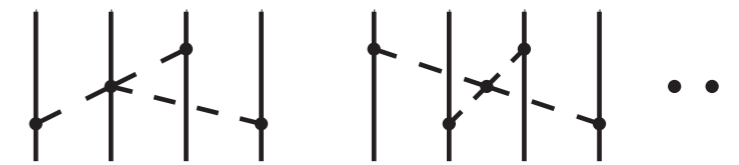
towards
physical
quark
masses

● note: strong short-distance repulsive interaction

BARYON-BARYON INTERACTIONS from CHIRAL EFFECTIVE FIELD THEORY

	BB interactions
LO	
NLO	
N ² LO	
N ³ LO	

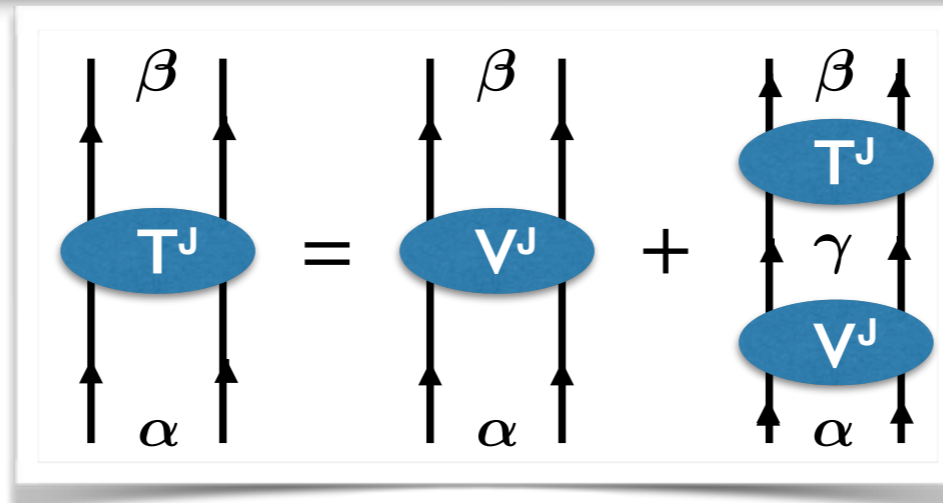
- Systematically organized hierarchy in powers of $\frac{Q}{\Lambda}$ (Q: momentum, energy, pion mass)

3 – body forces	
N ² LO	
N ³ LO	
4 – body forces	
N ³ LO	

- NN interaction state-of-the-art: N⁴LO plus convergence tests at N⁵LO
- YN interaction (limited data base): NLO plus three-body forces



Coupled-Channels Lippmann-Schwinger Equation



- Partial waves (LS) $_J$, baryon-baryon channels α, β

$$\mathbf{T}_{\beta\alpha}^J(p_f, p_i; \sqrt{s}) = \mathbf{V}_{\beta\alpha}^J(p_f, p_i) + \sum_{\gamma} \int_0^{\infty} \frac{dp p^2}{(2\pi)^3} \mathbf{V}_{\beta\gamma}^J(p_f, p) \frac{2\mu_{\gamma}}{p_{\gamma}^2 - p^2 + i\epsilon} \mathbf{T}_{\gamma\alpha}^J(p, p_i; \sqrt{s})$$

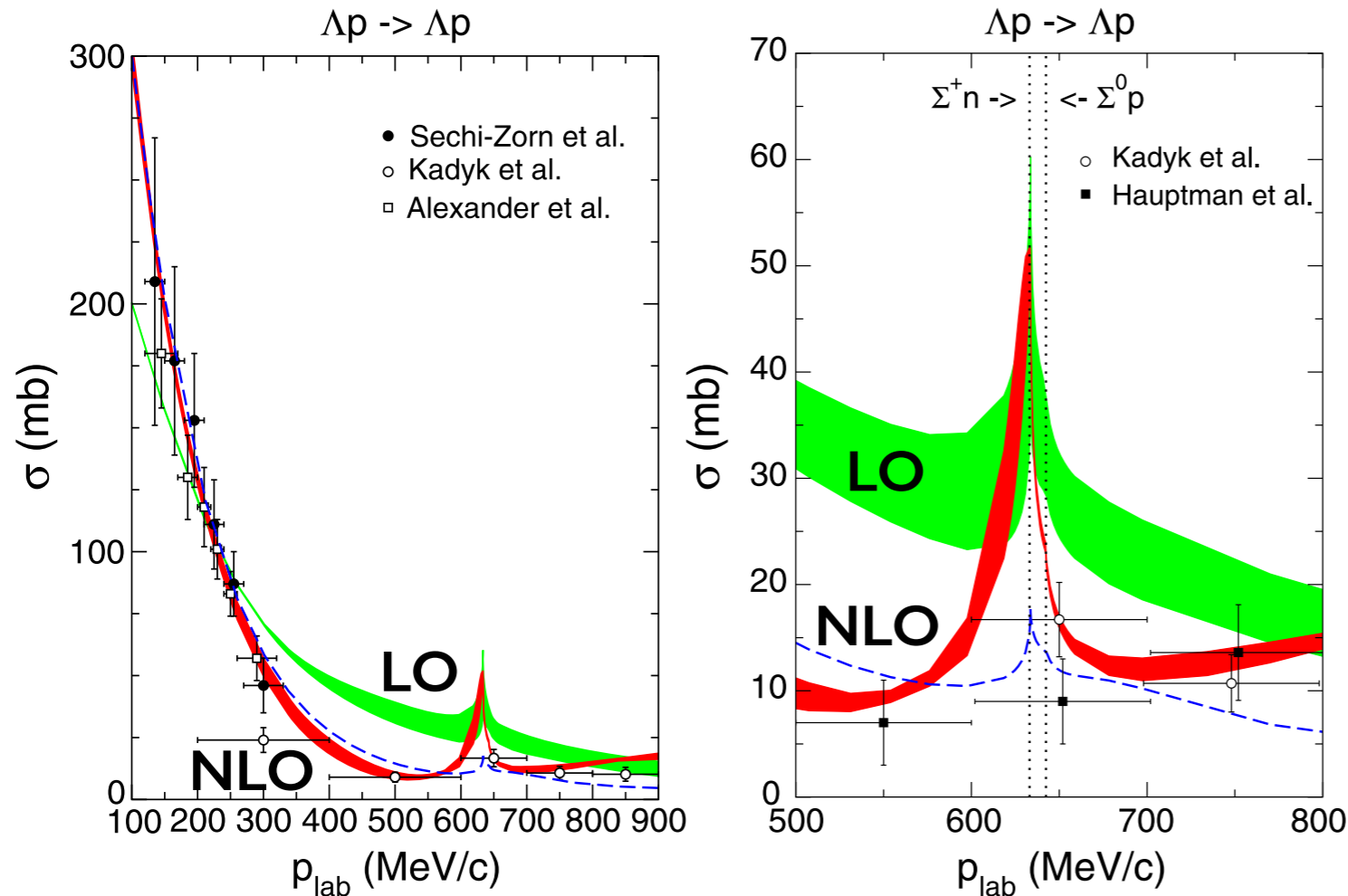
- On-shell momentum of intermediate channel γ determined by :

$$\sqrt{s} = \sqrt{M_{\gamma,1}^2 + p_{\gamma}^2} + \sqrt{M_{\gamma,2}^2 + p_{\gamma}^2}$$

- Relativistic kinematics relating lab. and c.m. momenta

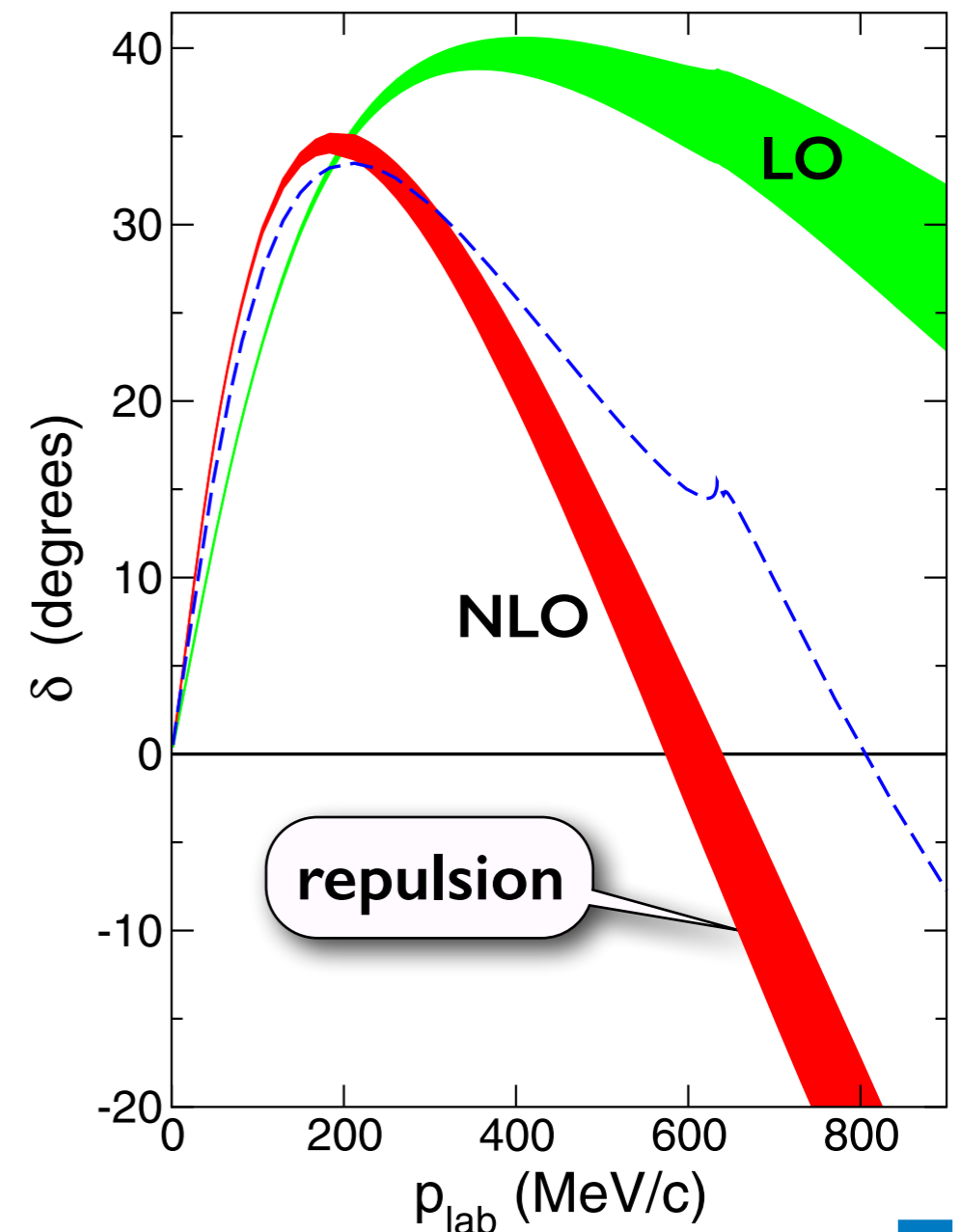


Hyperon - Nucleon Interaction from Chiral SU(3) EFT



J. Haidenbauer, S. Petschauer, N. Kaiser,
U.-G. Meißner, A. Nogga, W.W.
Nucl. Phys. A 915 (2013) 24

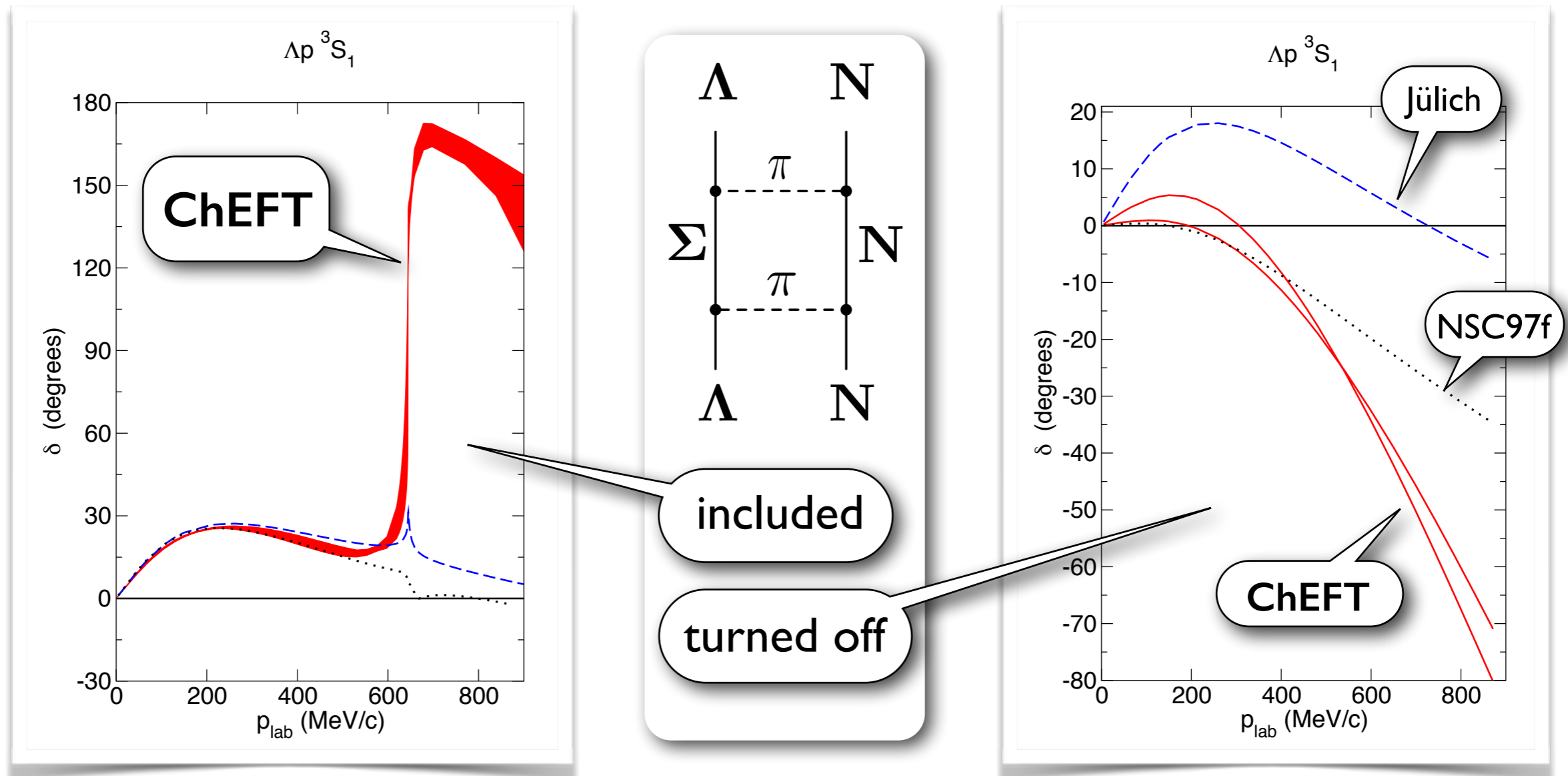
Λp 1S_0 phase shift



- moderate attraction at low momenta
→ relevant for hypernuclei
- strong repulsion at higher momenta
→ relevant for dense baryonic matter

Hyperon - Nucleon Interaction (contd.)

- Triplet-S channel and $\Lambda N \leftrightarrow \Sigma N$ coupling (2nd order tensor force)

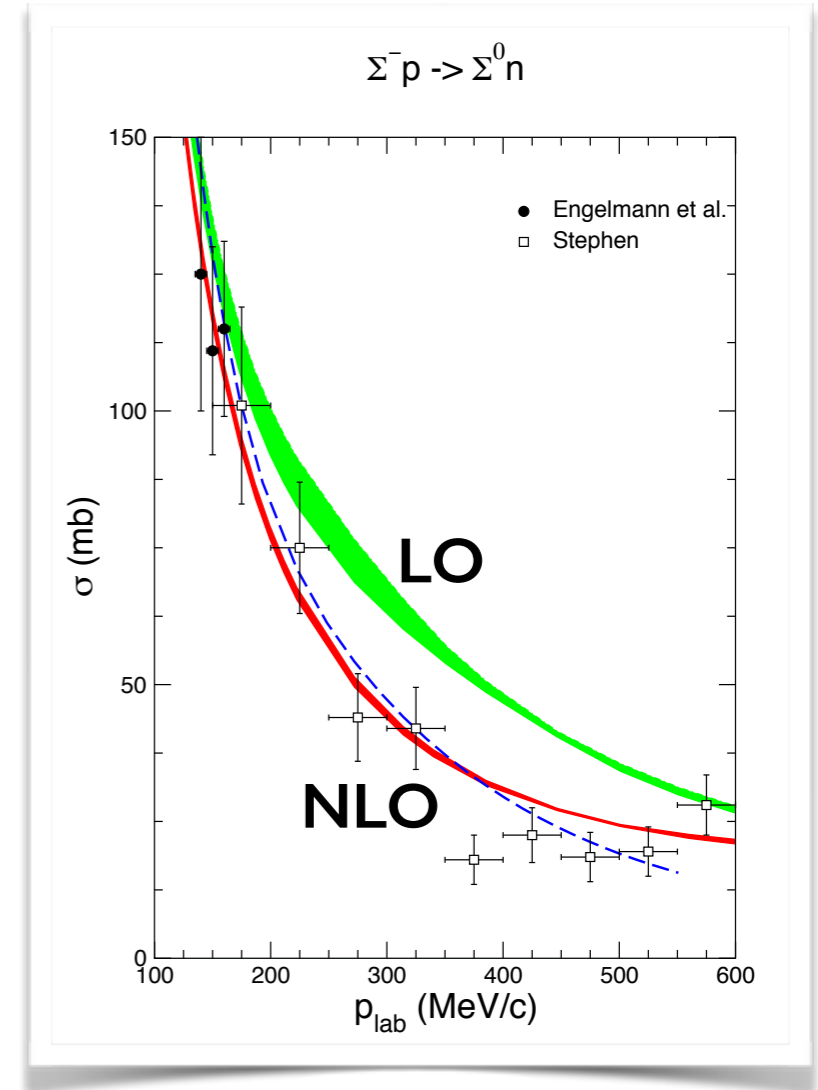
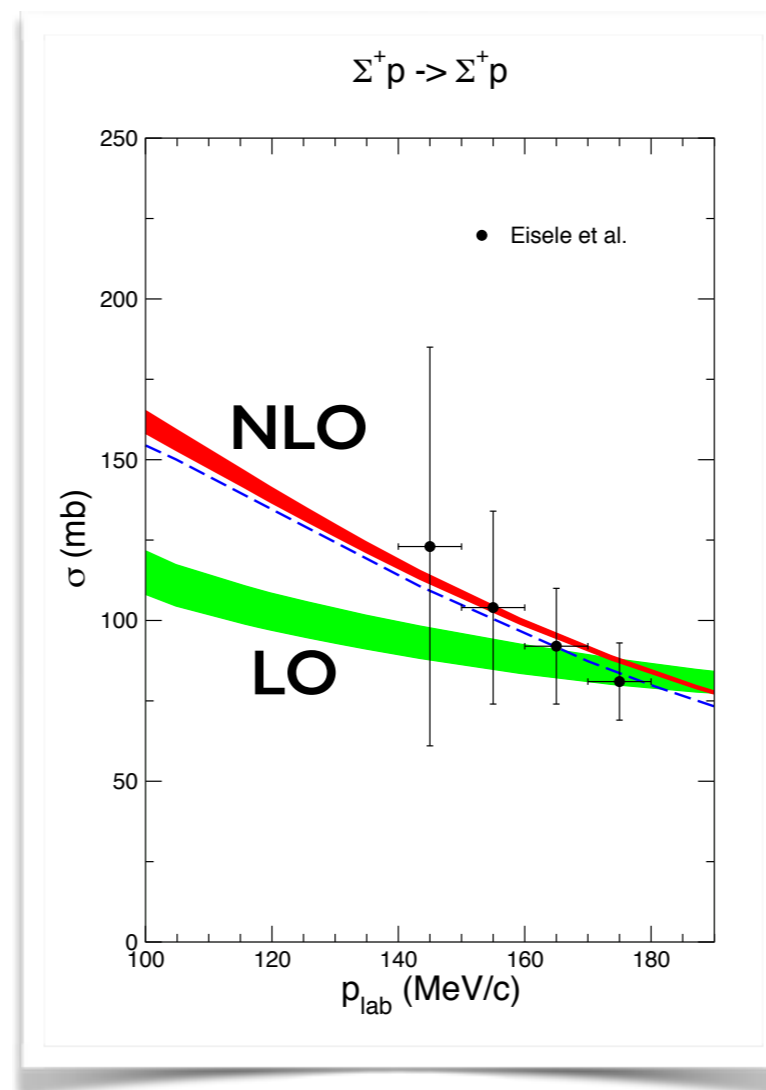
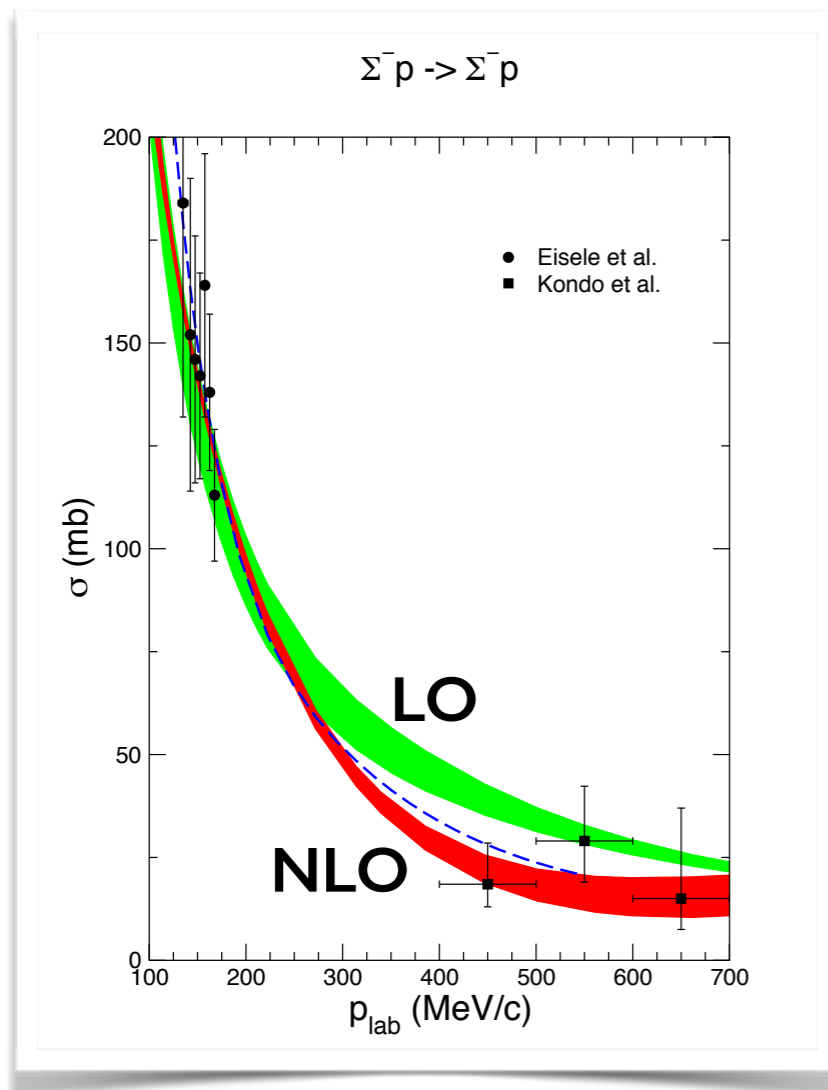


- In-medium (Pauli) suppression of $\Lambda N \leftrightarrow \Sigma N$ coupling :
increasing repulsion with rising density

Hyperon - Nucleon Interaction (contd.)

J. Haidenbauer, S. Petschauer, N. Kaiser,
U.-G. Meißner, A. Nogga, W.W.
Nucl. Phys. A 915 (2013) 24

- ΣN elastic and charge exchange scattering

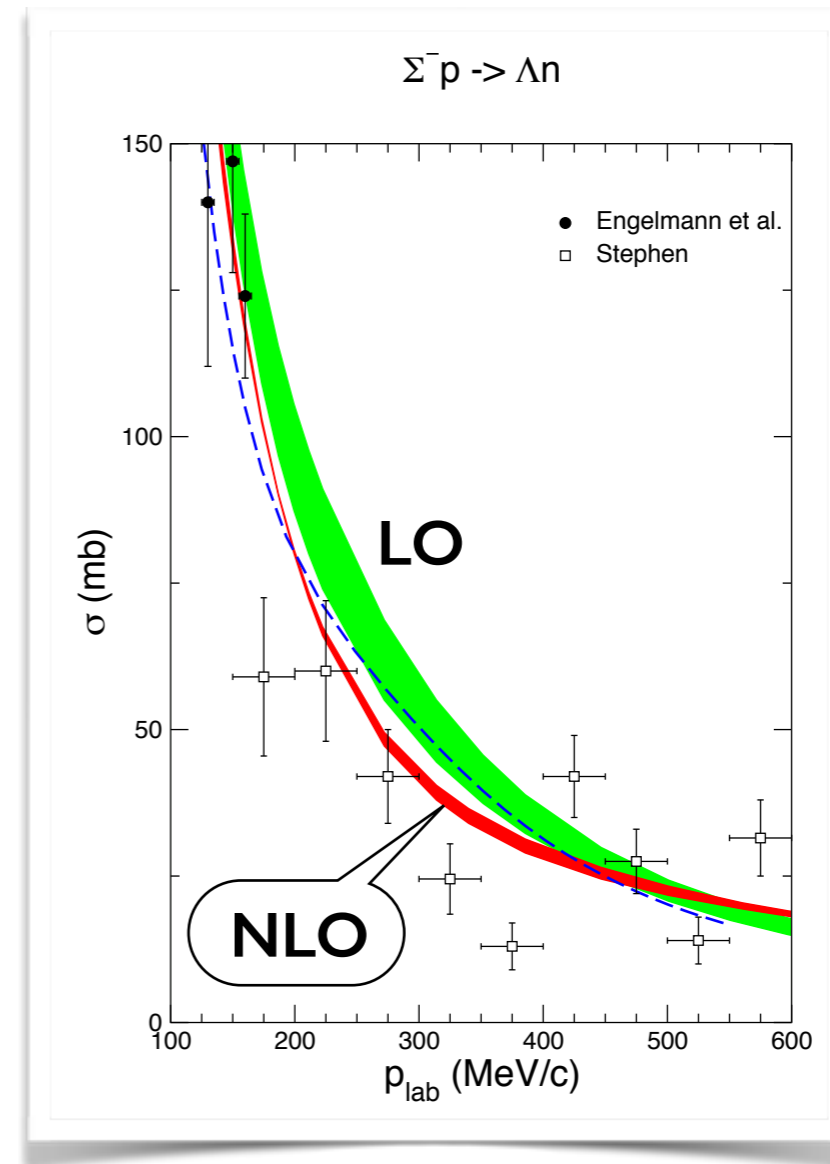
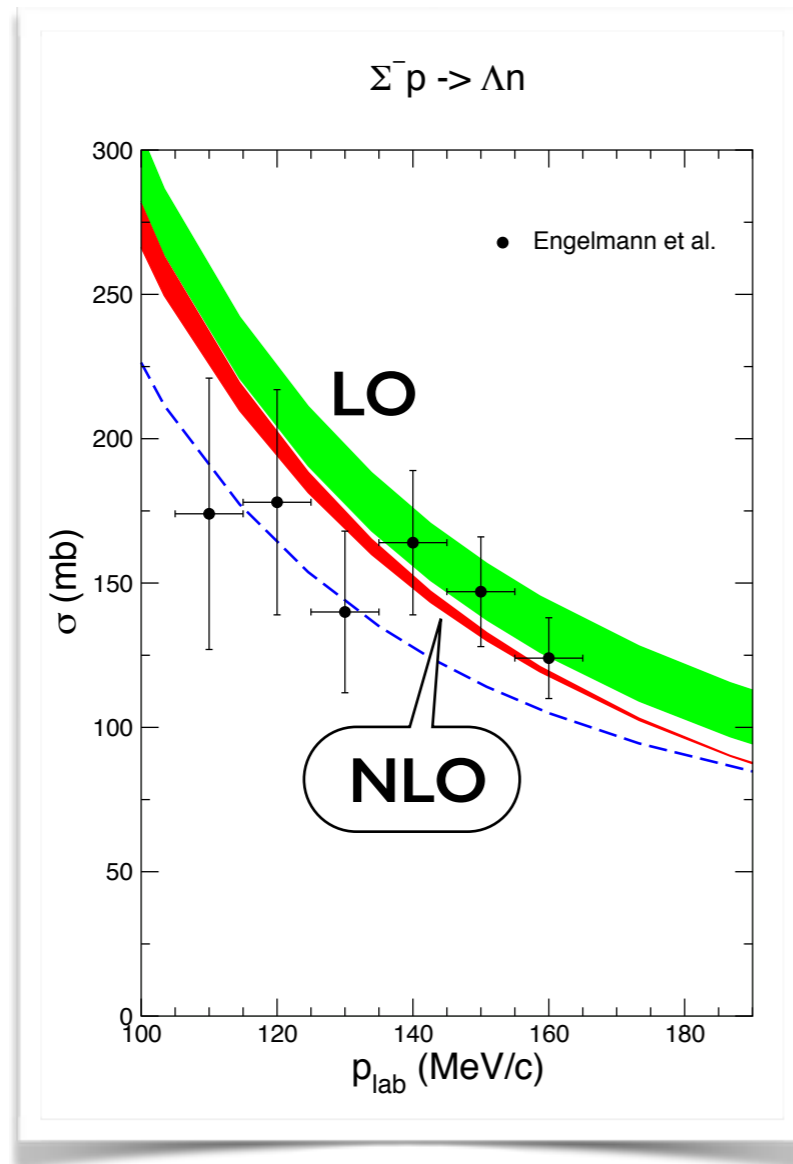


- Quest for much improved hyperon-nucleon scattering data base !

Hyperon - Nucleon Interaction (contd.)

J. Haidenbauer, S. Petschauer, N. Kaiser,
U.-G. Meißner, A. Nogga, W.W.
Nucl. Phys. A 915 (2013) 24

● $\Sigma N \rightarrow \Lambda N$ reaction



● Quest for much improved hyperon-nucleon scattering data base !

Part III

Hyperon Interactions in Nuclear and Neutron Matter

- **YNN three-body forces from Chiral SU(3) EFT**
- **Density dependence of
 Λ -nuclear single particle potential**
- **Towards a solution of the
“hyperon puzzle” in neutron stars ?**

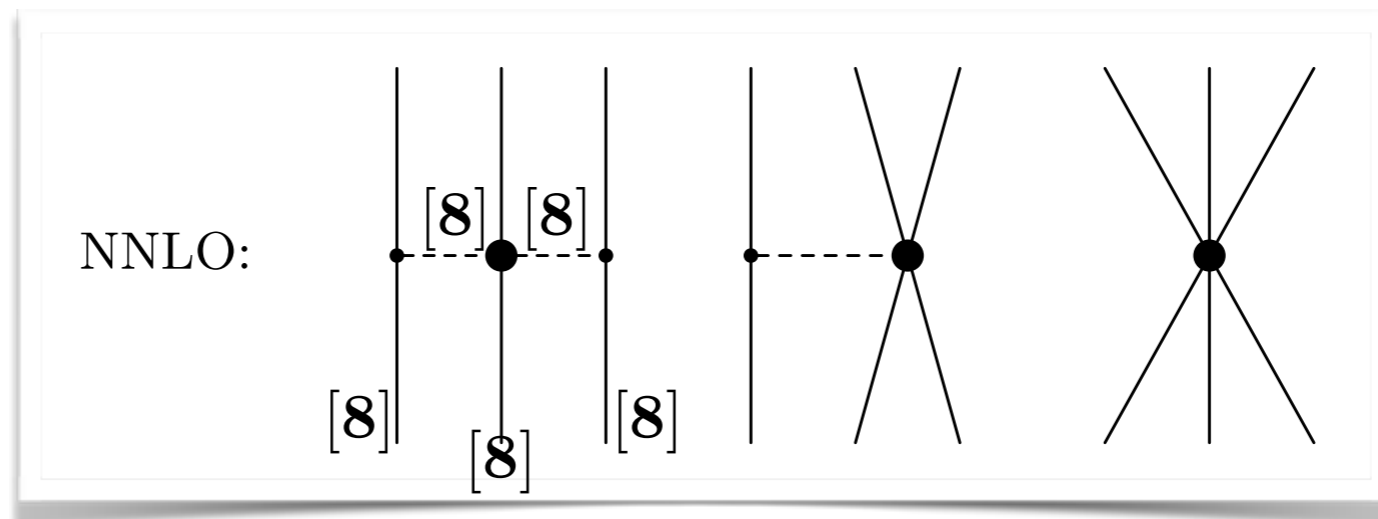


HYPERON - NUCLEON - NUCLEON THREE-BODY FORCES from CHIRAL SU(3) EFT

S. Petschauer et al. Phys. Rev. C93 (2016) 014001

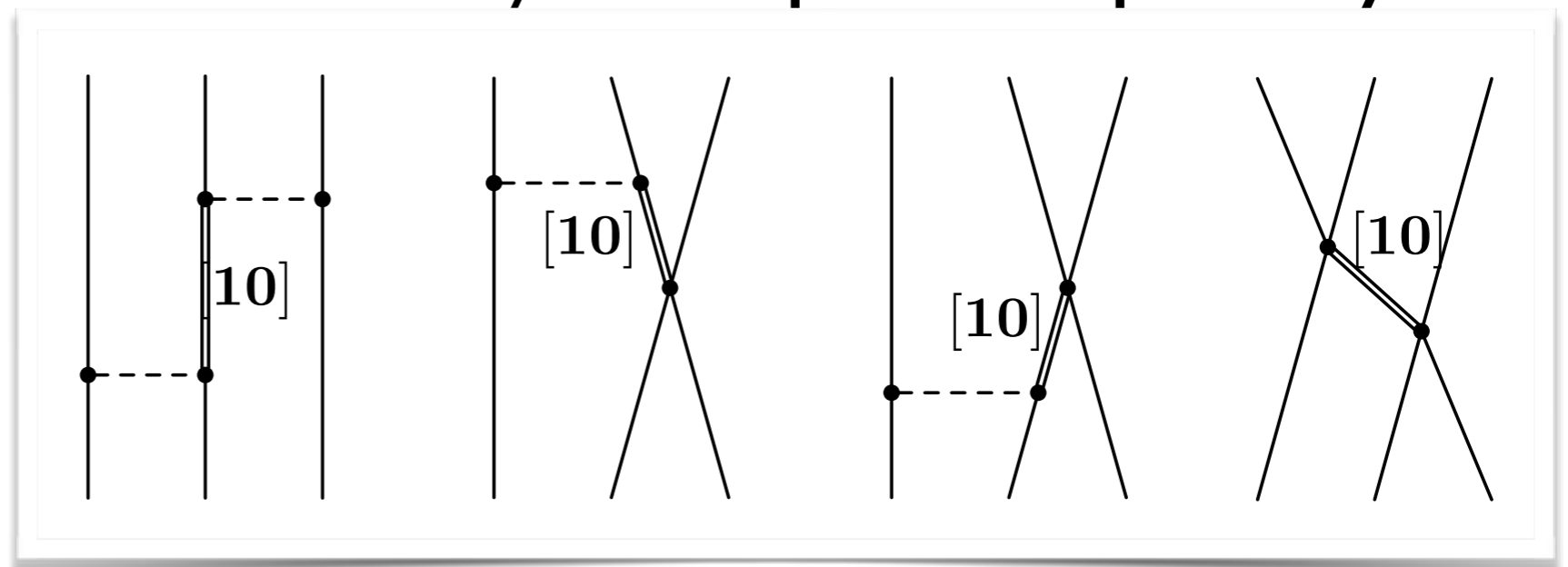
- **Chiral SU(3) Effective Field Theory:**
interacting pseudoscalar meson & baryon octets + contact terms

3-baryon
sector:



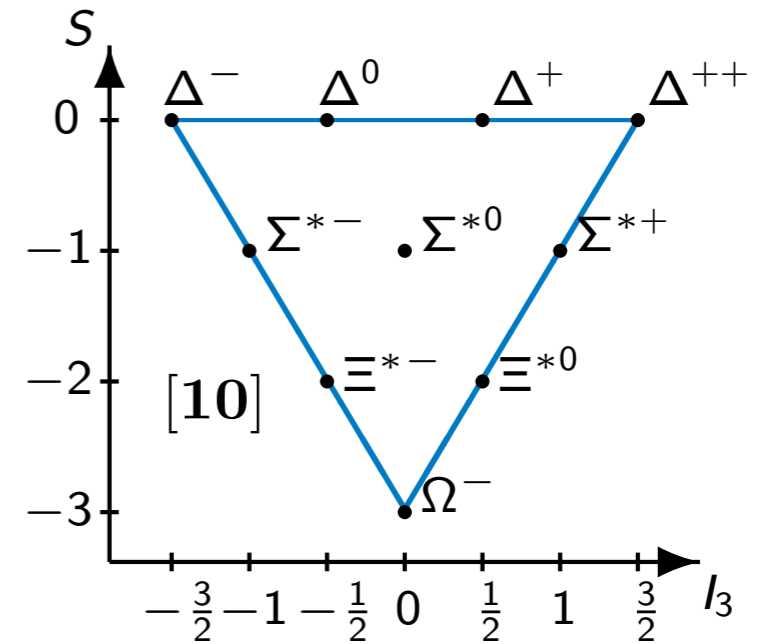
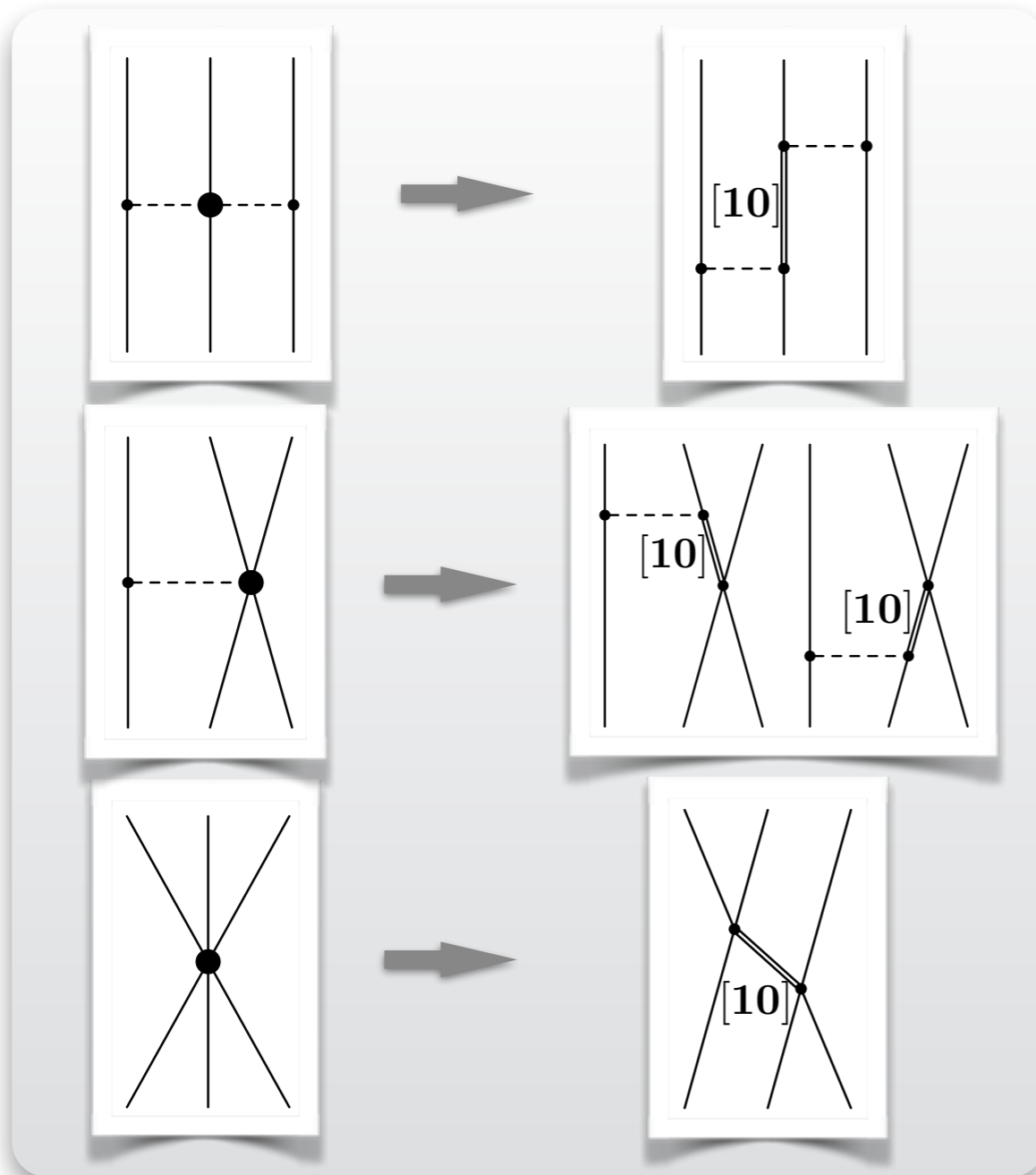
- **Chiral SU(3) Effective Field Theory with explicit decuplet baryons:**

explicit
baryon decuplet :
promotion to **NLO**

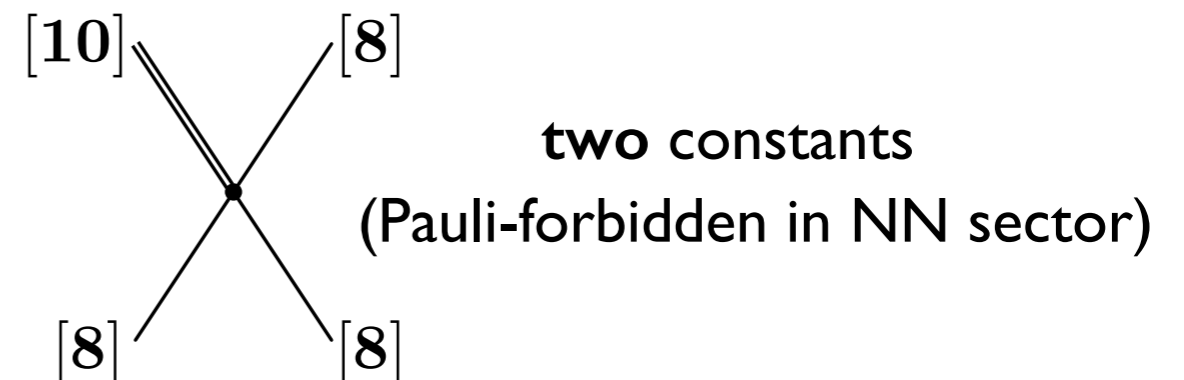
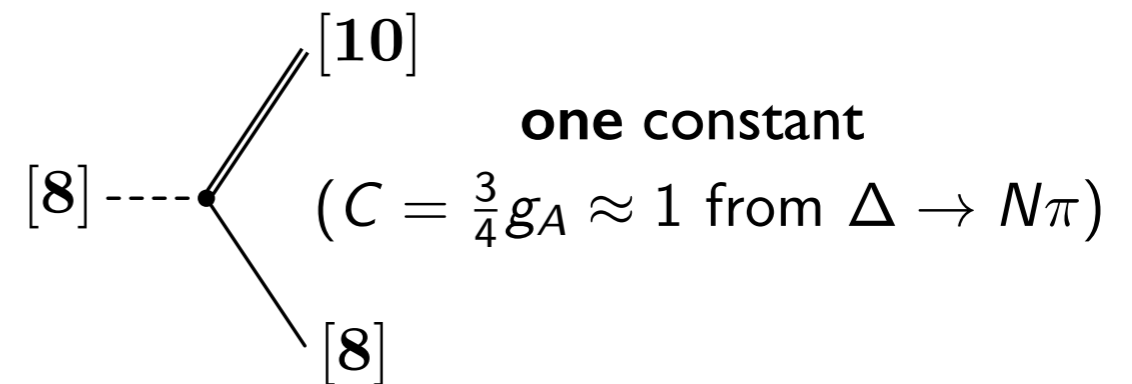


Decuplet Dominance in YNN three-body forces

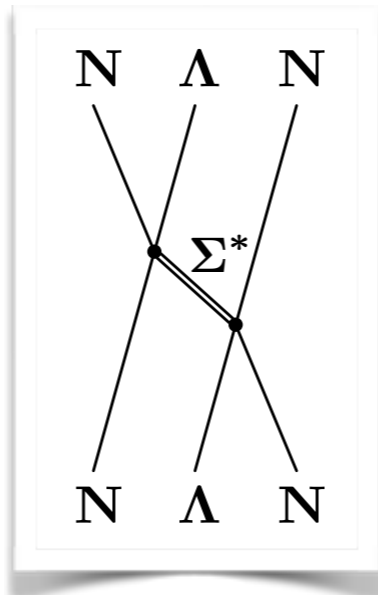
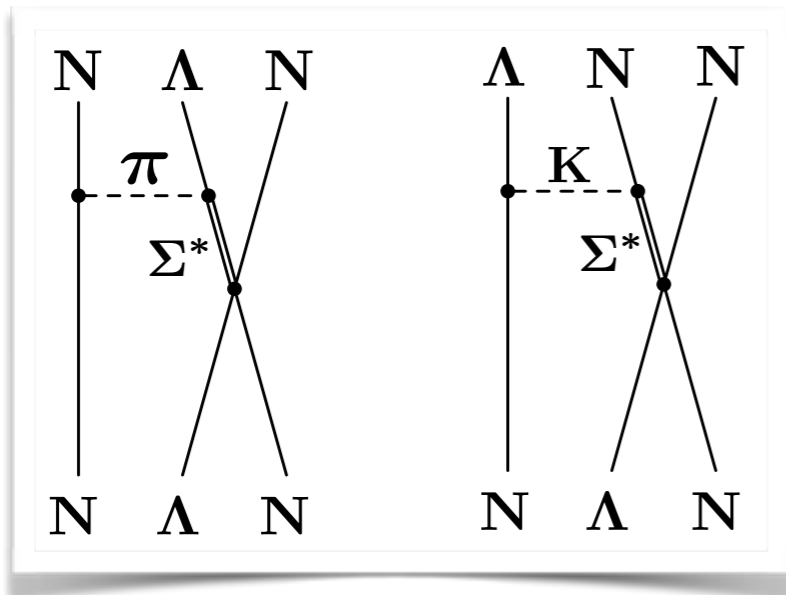
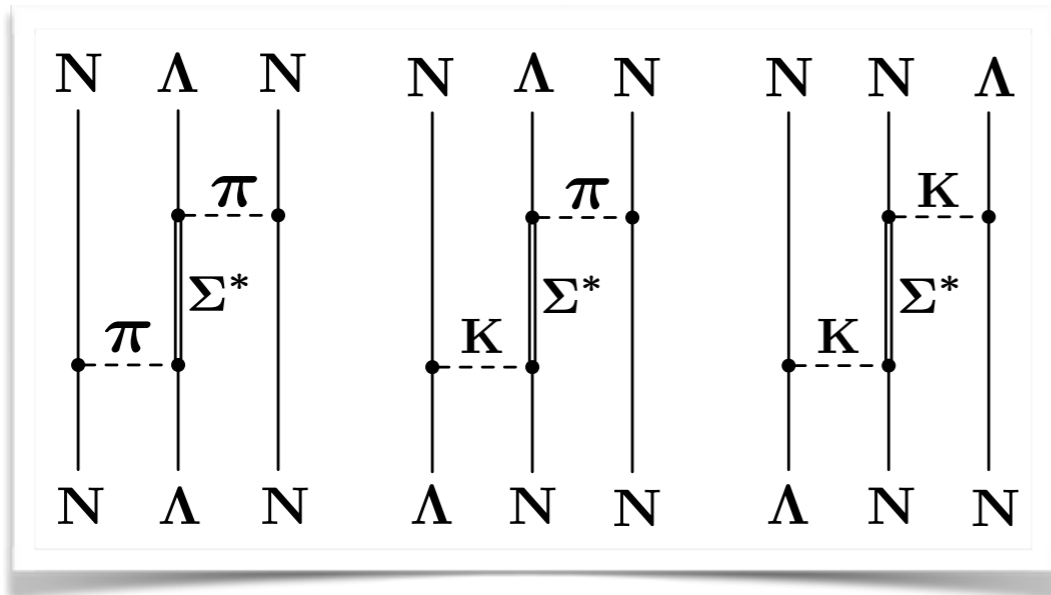
- Estimates of YNN 3-body interactions assuming dominant decuplet (Σ^* , Δ) intermediate states



- ... much reduced set of parameters
basic vertices :



Example : ΛNN three-body interactions



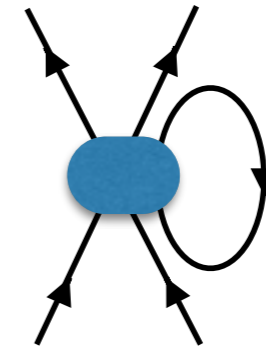
transition	type	B^*
$NNN \rightarrow NNN$	$\pi\pi$	Δ
$\Lambda NN \rightarrow \Lambda NN$	$\pi\pi$	Σ^*
	πK	Σ^*
	KK	Σ^*
	π	Σ^*
	K	Σ^*
	ct	Σ^*
$\Lambda NN \leftrightarrow \Sigma NN$	$\pi\pi$	Δ, Σ^*
	πK	Δ, Σ^*
	$\pi\eta$	Σ^*
	KK	Σ^*
	$K\eta$	Σ^*
	π	Δ, Σ^*
	K	Σ^*
	η	Σ^*
	ct	Σ^*

Density-dependent EFFECTIVE HYPERON - NUCLEON INTERACTION from CHIRAL THREE-BARYON FORCES

S. Petschauer, J. Haidenbauer, N. Kaiser, U.-G. Meißner, W.W.

Nucl. Phys. A957 (2017) 347

$$V_{12} = \sum_B \text{tr}_{\sigma_3} \int_{|\vec{k}| \leq k_f^B} \frac{d^3k}{(2\pi)^3} V_{123}$$



- Λn density-dependent effective interaction

$$V_{\Lambda n}^{\text{eff}, \pi\pi} = \frac{C^2 g_A^2}{2f^4 \Delta} [\rho_n + 2\rho_p] + \mathcal{F}(k_F^p, k_F^n; p, q)$$

repulsive

$$V_{\Lambda n}^{\text{eff}, \pi} = \frac{CH g_A}{9f^2 \Delta} [\rho_n + 2\rho_p] + \mathcal{G}(k_F^p, k_F^n; p, q)$$

+/-

$$V_{\Lambda n}^{\text{eff}, ct} = \frac{H^2}{18\Delta} [\rho_n + 2\rho_p]$$

repulsive

- Decuplet-octet mass difference $\Delta = M_{[10]} - M_{[8]} = 270 \text{ MeV}$

- Coupling parameters : $C = \frac{3}{4} g_A \simeq 1$ $-\frac{1}{f^2} \lesssim H \lesssim +\frac{1}{f^2}$ (dim. arguments natural size)

Density-dependent EFFECTIVE HYPERON - NUCLEON INTERACTION from CHIRAL THREE-BARYON FORCES

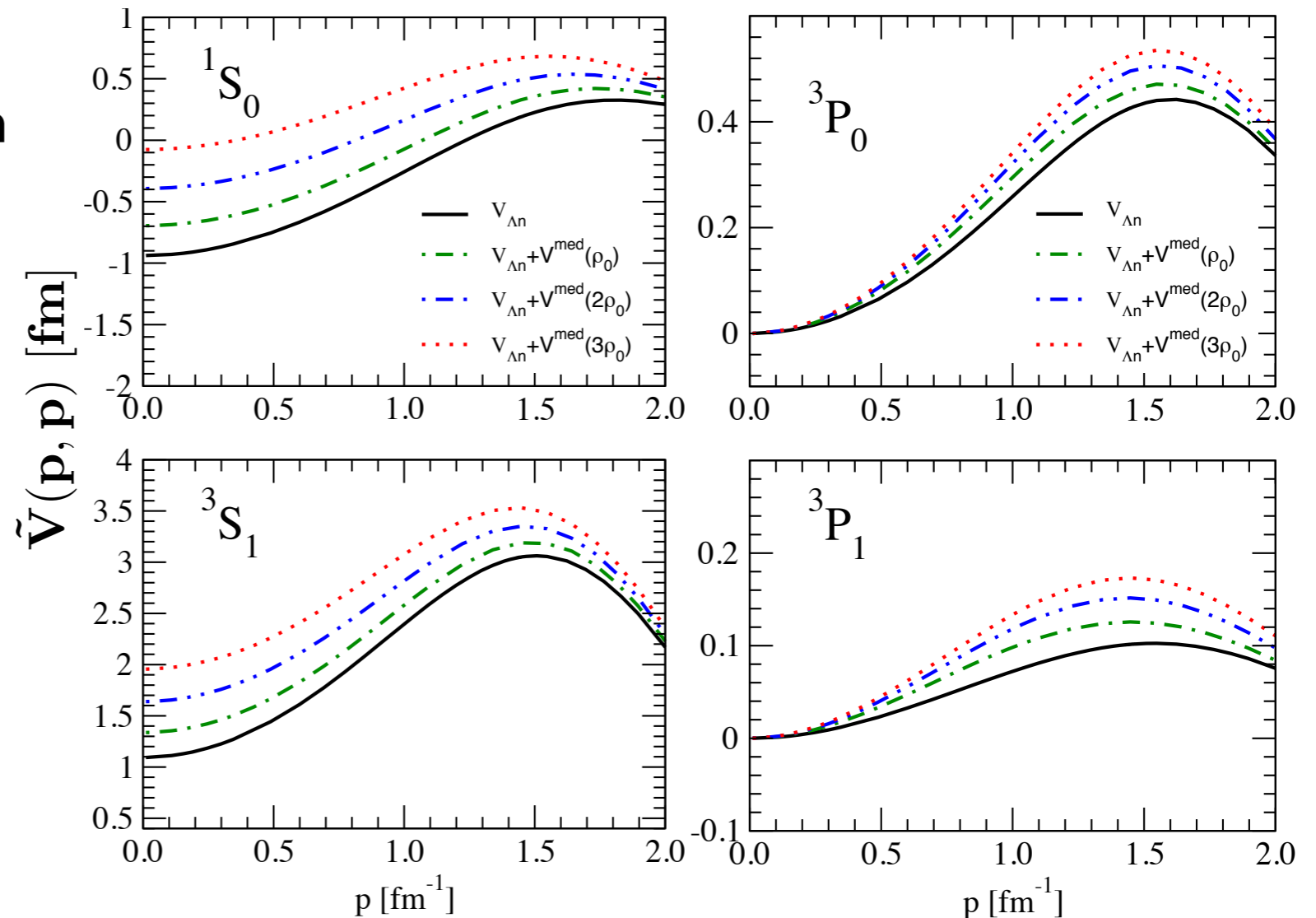
S. Petschauer, J. Haidenbauer, N. Kaiser, U.-G. Meißner, W.W.
 NPA957 (2017) 347

- ΛNN three-body force transformed into density-dependent effective two-body interaction

- Λn effective interaction in neutron matter

- Momentum-space potentials $(H = +\frac{1}{f^2})$

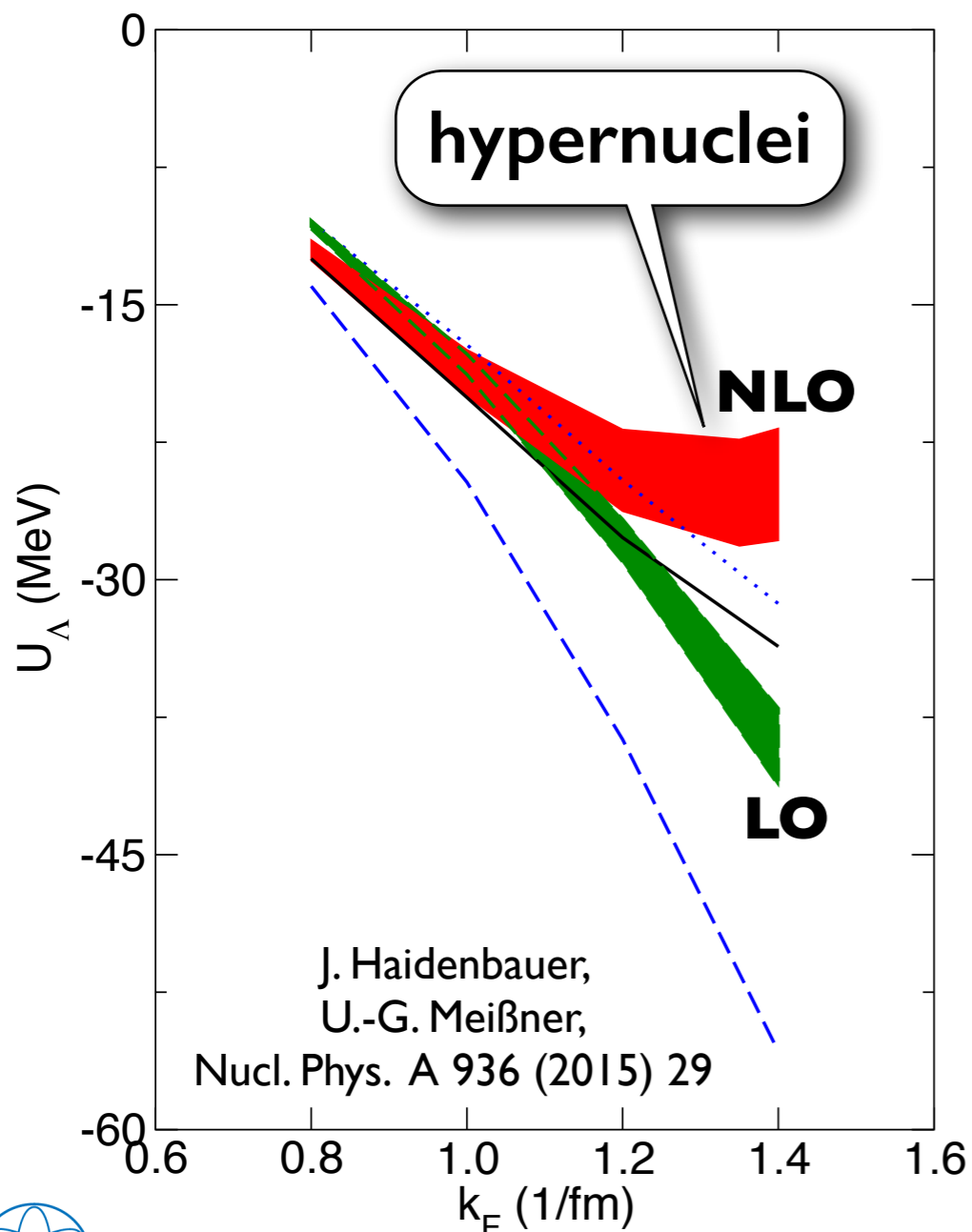
increasing repulsion
 \sim proportional to density



Density dependence of Λ single particle potential

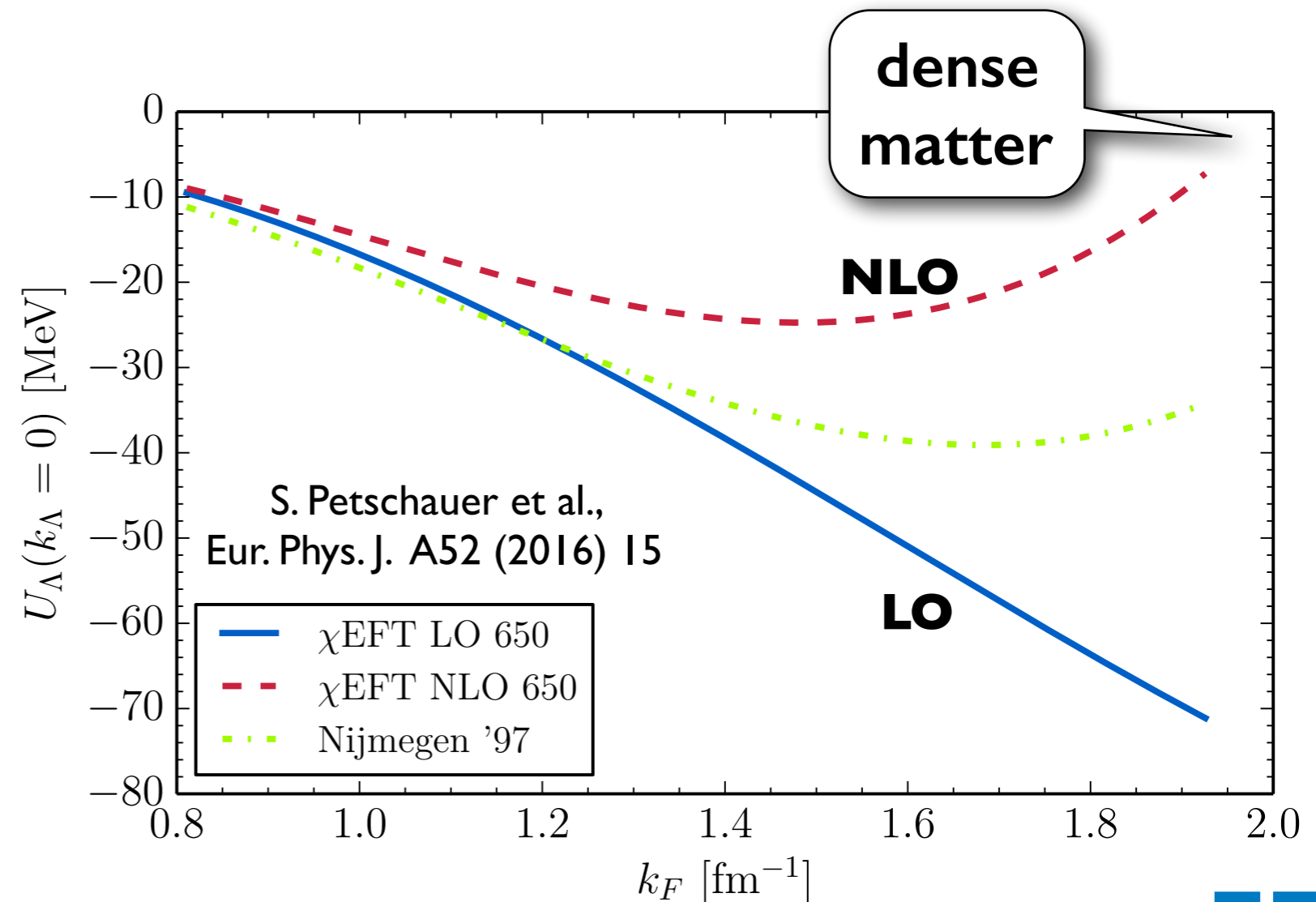
Λ in symmetric nuclear matter - YN two-body interactions only

Brueckner calculations
using chiral SU(3) interaction



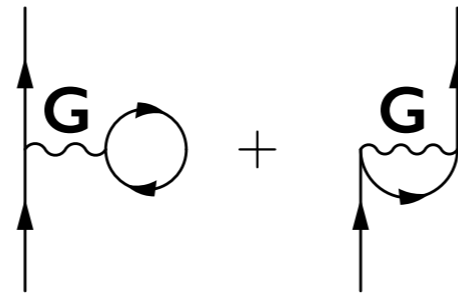
$$G(\omega) = V + V \frac{Q}{e(\omega) + i\epsilon} G(\omega)$$

(“continuous choice”)

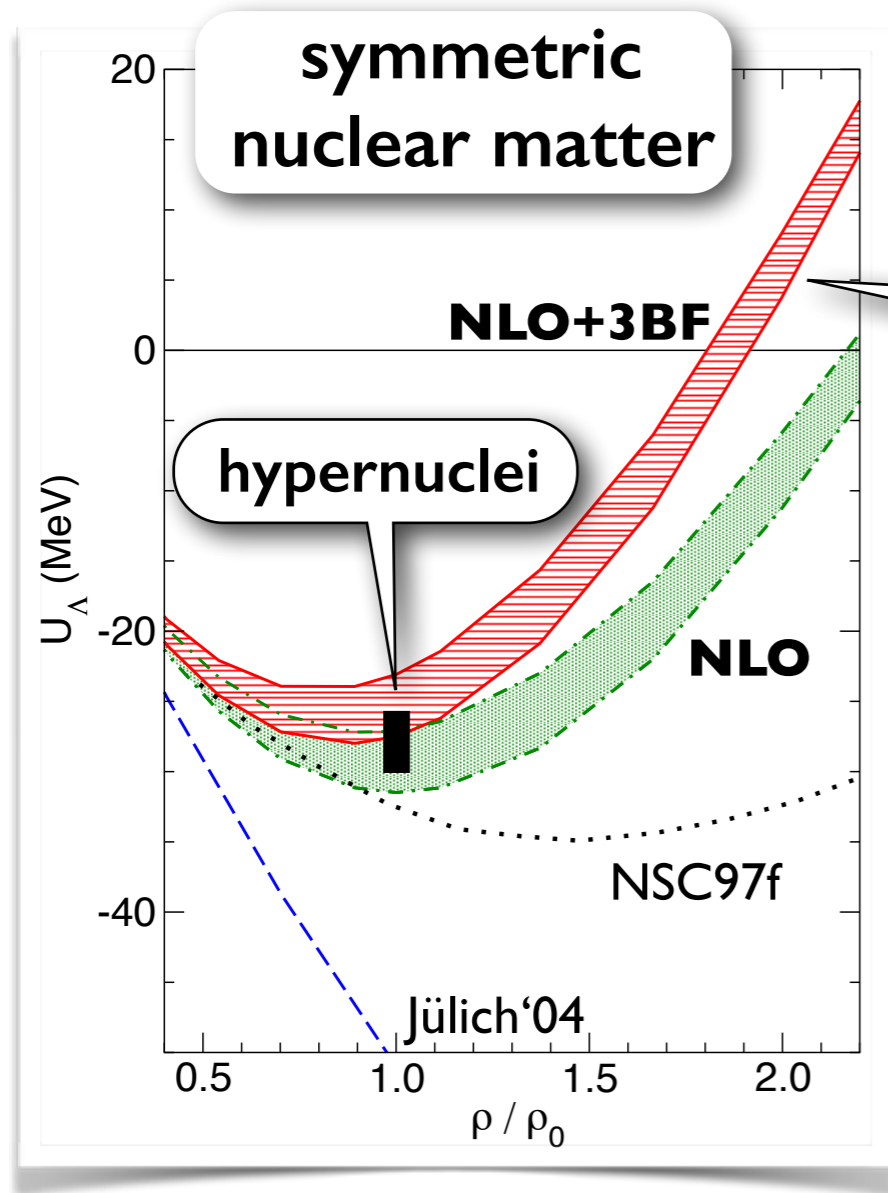


Density dependence of Λ single particle potential

- Brueckner calculations using chiral SU(3) interactions

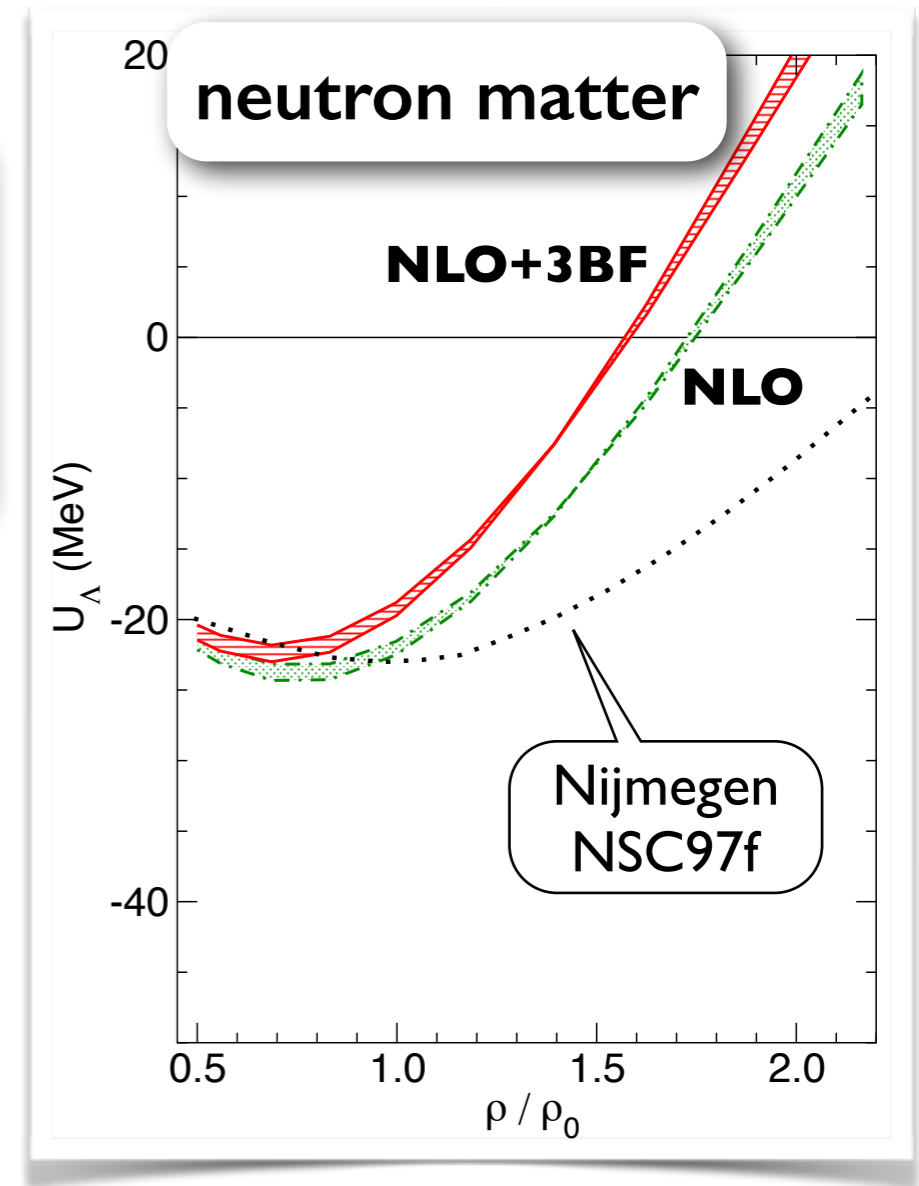


$$G(\omega) = V + V \frac{Q}{e(\omega) + i\epsilon} G(\omega)$$



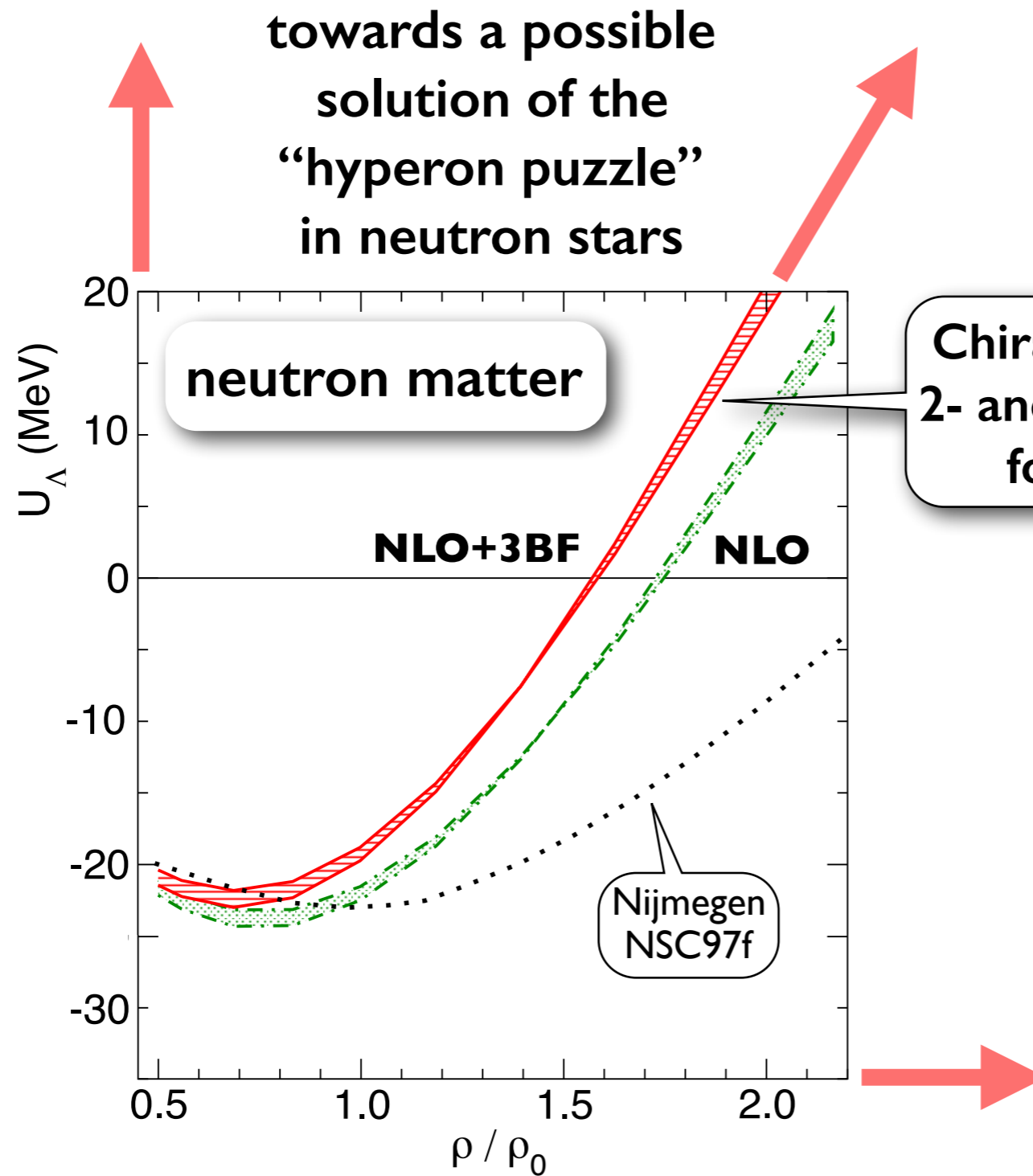
**Chiral SU(3)
2- and 3-body
forces**
($H = -\frac{1}{f^2}$)

J. Haidenbauer,
U.-G. Meißner,
N. Kaiser,
W.W.
arXiv:1612.03758
EPJA (2017)
to appear



- ... towards a possible solution of the “hyperon puzzle”

Hyperons in the core of neutron stars ?



- Quick estimate

$$\mathcal{E} = (E_n + E_\Lambda)\rho$$

$$E_\Lambda \simeq \frac{3(k_F^\Lambda)^2}{10 M_\Lambda^*} + U_\Lambda(\rho)$$

If E_Λ grows as fast as

$$\mu_n = \frac{\partial \mathcal{E}}{\partial \rho_n}$$

with increasing density:
no hyperons
in n-star matter



SUMMARY

- ★ **Constraints on dense baryon matter equation-of-state from neutron stars :**
 - ▶ very stiff EoS required !
 - ▶ “non-exotic” EoS (nuclear chiral dynamics) seems to work
 - ▶ hyperon puzzle:
naively adding hyperons implies far too soft EoS
- ★ **Progress in constructing hyperon-nuclear interactions from Chiral SU(3) Effective Field Theory**
 - ▶ YN two-body interactions at NLO
 - ▶ importance of $\Lambda N \leftrightarrow \Sigma N$ (2nd order pion exchange tensor force)
 - ▶ YNN three-body forces
- ★ **Single particle potential of a Λ in nuclear and neutron matter**
 - ▶ moderately attractive at low density (hypernuclei)
 - ▶ strongly repulsive at high density
... towards solution of “hyperon problem” in neutron stars

