

# Probing neutron-proton dynamics by pions

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In order to investigate the nuclear symmetry energy at high density, we study the pion production in central collisions of neutron-rich nuclei  $^{132}\text{Sn}+^{124}\text{Sn}$  at 300 MeV/nucleon using a new approach by combining the antisymmetrized molecular dynamics (AMD) and a hadronic cascade model (JAM). The dynamics of neutrons and protons is solved by AMD, and then pions and  $\Delta$  resonances in the reaction process are handled by JAM. We see the mechanism how the  $\Delta$  resonance and pions are produced reflecting the dynamics of neutrons and protons. We also investigate the impacts of cluster correlations as well as of the high-density symmetry energy on the nucleon dynamics and consequently on the pion ratio. We find that the  $\Delta^-/\Delta^{++}$  production ratio agrees very well with the neutron-proton squared ratio  $(N/Z)^2$  in the high-density and high-momentum region. We show quantitatively that  $\Delta$  production ratio, and therefore  $(N/Z)^2$ , are directly reflected in the  $\pi^-/\pi^+$  ratio, with modification in the final stage of the reaction.

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## I. INTRODUCTION

It is one of the important subjects in nuclear physics and astrophysics to determine the nuclear symmetry energy at various densities. Constraints on the symmetry energy at densities lower than the normal density  $\rho_0$  have been obtained to some degree from nuclear physics experiments [1, 2]. On the other hand, intermediate-energy heavy-ion collisions are believed to be useful to investigate the symmetry energy at high densities. In fact, in central collisions at several hundred MeV/nucleon, transport calculations show that the maximum density around  $2\rho_0$  is reached when the system is compressed in an early stage of the reaction. The ratio of the neutron and proton densities in the compressed part is naturally sensitive to the symmetry energy at high densities [3, 4]. However, it is necessary to understand the link between the effects in the compression stage and the final observables, in order to extract the high-density symmetry energy from experimental data available at present and in future.

The  $\pi^-/\pi^+$  ratio has been proposed to be a good probe to constrain the high-density behavior of the symmetry energy [4]. In heavy-ion collisions, the pions are produced through the  $\Delta$  resonance formation in the nucleon-nucleon collisions that typically occur at early times in the compressed part of the system. If all the pions produced by the  $NN \rightarrow N\Delta \rightarrow NN\pi$  process are directly emitted, the expected ratio should be  $\pi^-/\pi^+ = (5N^2 + NZ)/(5Z^2 + NZ) \approx (N/Z)^2$ , where  $N$  and  $Z$  are the numbers of neutrons and protons that are relevant for the  $\Delta$  production. In another case of chemical equilib-

rium for  $NN \leftrightarrow N\Delta$  and  $\Delta \leftrightarrow N\pi$  reactions at a temperature  $T$ , the ratio  $\pi^-/\pi^+ \simeq e^{2(\mu_n - \mu_p)/T}$  [5, 6], where  $\mu_n$  and  $\mu_p$  are the neutron and proton chemical potentials, may also be related to the neutron and proton densities  $\rho_n/\rho_p$  or the symmetry energy [4, 7]. We also notice the relation  $e^{2(\mu_n - \mu_p)/T} \approx [f_n(\varepsilon)/f_p(\varepsilon)]^2$  when the neutron and proton phase-space densities are compared at the same single-particle energy  $\varepsilon$  satisfying  $\varepsilon - \mu_{n,p} \gg T$ . Thus, both in these extreme cases, the  $\pi^-/\pi^+$  ratio is related to some kind of neutron-to-proton squared ratio  $(N/Z)^2$  which is then supposed to be sensitive to the symmetry energy at high densities.

Some theoretical studies have been performed by different transport models to investigate the sensitivity of pion observables [4, 8–13]. At present, however, some of these results are contradicting to each other even qualitatively. The symmetry-energy dependence of the  $\pi^-/\pi^+$  ratio in Ref. [11] is opposite to that predicted by Refs. [4, 10], while Ref. [12] predicts that the  $\pi^-/\pi^+$  ratio for the total pion multiplicities does not depend on the density dependence of the symmetry energy. Moreover, the  $\pi^-/\pi^+$  ratio in central Au + Au collisions measured by the FOPI Collaboration [9, 14] is actually larger than the squared neutron-to-proton ratio of the total system,  $(N/Z)_{\text{sys}}^2$ , at low energies such as at 400 MeV/nucleon. Since the compressed part naturally becomes less neutron rich than the total system, we inevitably have  $\pi^-/\pi^+ > (N/Z)_{\text{sys}}^2 > (N/Z)^2$ , and therefore  $\pi^-/\pi^+$  cannot agree with  $(N/Z)^2$ , contradicting to the arguments in the previous paragraph, as long as  $N/Z$  is identified with the density ratio  $\rho_n/\rho_p$  in the com-

pressed part as is typically done in the literature. Most of the transport calculations underestimate the measured pion ratio, but some recent calculations [11, 12] predict the ratio as high as the experimental data. Nevertheless, investigations cannot be found in the literature on the mechanism to increase the pion ratio. Hence, our understanding of the pion ratio in relation to the symmetry energy and the nucleon dynamics is not complete particularly when we need to have a precise description to explain the absolute magnitude of  $\pi^-/\pi^+$ .

In this paper, we study the pion production in central collisions of neutron-rich nuclei  $^{132}\text{Sn} + ^{124}\text{Sn}$  at 300 MeV/nucleon, which is one of the systems to be measured at RIBF/RIKEN [15]. We develop and employ a new approach by combining the antisymmetrized molecular dynamics (AMD) [16] and a hadronic cascade model (JAM) [17]. The dynamics of neutrons and protons is solved by AMD, and then pions and  $\Delta$  resonances in the reaction process are handled by JAM. AMD calculations were performed for several cases with and without cluster correlations [18], and with two effective interactions corresponding to different density dependences of the symmetry energy. These different cases of AMD calculation yield different dynamics of neutrons and protons. The main aim here is to utilize these different cases to identify how the  $\Delta$  resonances and the emitted pions carry the information of the nucleon dynamics. We will also see the impacts of cluster correlations as well as of the high-density symmetry energy on the nucleon dynamics and consequently on the pion ratio.

## II. FORMULATION

In order to understand the  $\pi^-/\pi^+$  enhancement mechanism and to extract high-density symmetry energy, we need a reliable transport model of nucleons, clusters,  $\Delta$  resonances and pions. AMD has been demonstrated to be a reliable transport model of nucleon and clusters [16, 18]. It takes account of nucleon mean field effects, two-nucleon elastic scatterings, event-by-event fluctuations and antisymmetrization of nucleon wave functions. All of these ingredients could affect the  $N/Z$  ratio during the collision, and thus have influences on the  $\pi^-/\pi^+$  ratio. One can explain fragment mass distribution in heavy-ion collisions precisely, especially with cluster correlations. In the present AMD code, however,  $\Delta$  resonances and pions have not been incorporated. On the other hand, JAM is a reliable hadron transport model [17]. It has been successfully applied to  $pA$  and  $AA$  collisions in the energy range from 1 GeV/nucleon to 158 GeV/nucleon. We can describe hadron production spectra in JAM, and collective flows are also explained with the mean field effects switched on [19]. By comparison, antisymmetrization of nucleon wave functions is not included, and cluster correlation during the time evolution are not implemented. At present, there is no transport model which includes all of the above mentioned important ingredients; mean field

effects, two-body collisions including elastic and inelastic processes, event-by-event fluctuations, antisymmetrization of nucleon wave functions, and dynamical cluster correlations. Thus one of the best ways would be to combine AMD and JAM — nucleon transport in AMD and particle production in JAM.

### A. Perturbative treatment of pions and $\Delta$

We consider here a transport model which describes the reaction dynamics by the time evolution of one-body density matrices or corresponding phase-space distributions  $f_\alpha(\mathbf{r}, \mathbf{p}, t)$ , where the index  $\alpha$  stands for the particle species such as nucleons ( $N$ ),  $\Delta$  resonances ( $\Delta$ ) and pions ( $\pi$ ). It should be implicitly understood that the isospin (and spin) components of these particles are also distinguished by  $\alpha$ . The coupled equations for  $f_\alpha$  ( $\alpha = N, \Delta, \pi$ ) can be written in general as

$$\frac{\partial f_\alpha}{\partial t} = \dot{f}_\alpha^{\text{MF}}[f_N, f_\Delta, f_\pi] + I_\alpha[f_N, f_\Delta, f_\pi] \quad (1)$$

with the mean-field term  $\dot{f}_\alpha^{\text{MF}}$  and the collision term  $I_\alpha$ , both of which depend on the phase-space distributions of all the particles at the time  $t$  in general. The collision term includes at least the following processes

$$N + N \rightarrow N + N, \quad (2)$$

$$N + N \rightarrow N + \Delta, \quad (3)$$

$$N + \Delta \rightarrow N + N, \quad (4)$$

$$\Delta \rightarrow N + \pi, \quad (5)$$

$$N + \pi \rightarrow \Delta. \quad (6)$$

Now we limit our consideration to the sub-threshold or near-threshold cases where only a small number of  $\Delta$  production processes occur because the incident energy is not so high. Then we expect that the  $\Delta$  production can be treated perturbatively. If the perturbation parameter  $\lambda$  is multiplied to the probability of the  $NN \rightarrow N\Delta$  process,  $\Delta$  resonances and pions will appear in the first order of  $\lambda$ ,  $f_\Delta = \lambda f_\Delta^{(1)} + O(\lambda^2)$  and  $f_\pi = \lambda f_\pi^{(1)} + O(\lambda^2)$ . The zeroth order equation for the nucleon distribution  $f_N = f_N^{(0)} + O(\lambda)$  is

$$\frac{\partial f_N^{(0)}}{\partial t} = \dot{f}_N^{\text{MF}}[f_N^{(0)}, 0, 0] + I_N^{(\lambda=0)}[f_N^{(0)}, 0, 0], \quad (7)$$

where the collision term  $I_N^{(\lambda=0)}$  includes only the elastic  $NN \rightarrow NN$  scatterings. The distributions of  $\Delta$  resonances and pions can be solved by

$$\frac{\partial f_\Delta}{\partial t} = \dot{f}_\Delta^{\text{MF}}[f_N^{(0)}, f_\Delta, f_\pi] + I_\Delta[f_N^{(0)}, f_\Delta, f_\pi], \quad (8)$$

$$\frac{\partial f_\pi}{\partial t} = \dot{f}_\pi^{\text{MF}}[f_N^{(0)}, f_\Delta, f_\pi] + I_\pi[f_N^{(0)}, f_\Delta, f_\pi], \quad (9)$$

which are correct up to the first order of  $\lambda$ .

The equation (7) for the zeroth order of the nucleon distribution can be solved assuming a system composed of only nucleons without considering productions of other particles. In our present approach, we solve the nucleon dynamics by AMD [16, 18]. Then, for the calculated  $f_N^{(0)}$ , the equations (8) and (9) for  $\Delta$  resonances and pions are solved by another transport model JAM [17] which can handle particle productions. The information of nucleons in the JAM calculation is always replaced by  $f_N^{(0)}$  calculated by AMD. Namely the particle production is calculated by JAM based on the nucleon dynamics calculated by AMD.

The above treatment violates some conservation laws in the higher orders  $O(\lambda^2)$ . The result may be improved by introducing corrections for the conservation laws of baryon number, change and energy, modifying the nucleon information  $f_N^{(0)}$  in Eqs. (8) and (9). When a  $\Delta$  resonance exists in the JAM calculation, it should have been produced by a collision of two nucleons. Such a pair of two nucleons is chosen in the AMD calculation by taking into account the distance from the  $\Delta$  resonance, the phase space to produce a  $\Delta$  resonance, and the charge conservation condition. Then one of the two nucleons is annihilated (assuming that it is replaced by the  $\Delta$  resonance) and the charge and the momentum of the other nucleon is modified for the charge and energy conservations. In the pion case, we consider only the charge conservation by modifying the charge of a nucleon for each pion if necessary. We have checked the validity of this prescription by comparing two different calculations performed by the JAM code. The first calculation is done as in the AMD-JAM calculation described above but Eq. (7) is solved also by JAM by turning off all the inelastic  $NN$  collisions. The result is compared with the standard JAM calculation which solves all the particles as usual. With the present prescription for the conservation laws, the pion multiplicities and the pion ratios in the two calculations agree well within the errors of about 10% and 2%, respectively, at the incident energy 300 MeV/nucleon. These agreements get worse slightly at the incident energy 400 MeV/nucleon.

## B. AMD

AMD [16] describes the dynamics of a many-nucleon system by the time evolution of a Slater determinant of Gaussian wave packets

$$\langle \mathbf{r} | \varphi_j \rangle = e^{-\nu(\mathbf{r} - \mathbf{Z}_j / \sqrt{\nu})^2} \chi_{\alpha_j}, \quad (10)$$

where the wave packet centroid is denoted by  $\mathbf{Z}_j$  which is a complex vector, and the spin-isospin state  $\chi_{\alpha_j}$  takes  $p \uparrow$ ,  $p \downarrow$ ,  $n \uparrow$  and  $n \downarrow$ . The width parameter is chosen to be  $\nu = (2.5 \text{ fm})^{-2}$  as usual. The corresponding phase-

space distribution is

$$f_\alpha(\mathbf{r}, \mathbf{p}) = 8 \sum_{j \in \alpha} \sum_{k \in \alpha} e^{-2\nu(\mathbf{r} - \mathbf{R}_{jk})^2} \times e^{-(\mathbf{p} - \mathbf{P}_{jk})^2 / 2\hbar^2\nu} B_{jk} B_{kj}^{-1} \quad (11)$$

with  $\mathbf{R}_{jk} = (\mathbf{Z}_j^* + \mathbf{Z}_k) / \sqrt{\nu}$ ,  $\mathbf{P}_{jk} = 2i\hbar\sqrt{\nu}(\mathbf{Z}_j^* - \mathbf{Z}_k)$  and  $B_{jk} = \langle \varphi_j | \varphi_k \rangle$ .

### 1. Mean field term

The mean field term  $\hat{f}_\alpha^{\text{MF}}[f_N, 0, 0]$  is given by the equation of motion for the wave packet centroids  $\{\mathbf{Z}_j\}$  derived from the time-dependent variational principle.

In the present calculations, we employ the Skyrme SLy4 force [20] as the effective interaction with the spin-orbit term omitted. The corresponding nuclear-matter incompressibility is  $K = 230 \text{ MeV}$  at the saturation density  $\rho_0 = 0.160 \text{ fm}^{-3}$ . The nuclear-matter symmetry energy at  $\rho_0$  is  $S_0 = 32.0 \text{ MeV}$  with the slope parameter  $L = 46 \text{ MeV}$  (called ‘asy-soft’ or soft symmetry energy in Sec. III). In order to study the effect of the density dependence of the symmetry energy, we also perform calculations with a force obtained by changing the density dependent term in the SLy4 force

$$v_\rho^{(L=46)} = \frac{1}{6} t_3 (1 + x_3 P_\sigma) \rho(\mathbf{r}_1)^\alpha \delta(\mathbf{r}_1 - \mathbf{r}_2) \quad (12)$$

to

$$v_\rho^{(L=108)} = \frac{1}{6} t_3 (1 + x'_3 P_\sigma) \delta(\mathbf{r}_1 - \mathbf{r}_2) \rho(\mathbf{r}_1)^\alpha + \frac{1}{6} t_3 (x_3 - x'_3) \rho_0^\alpha P_\sigma \delta(\mathbf{r}_1 - \mathbf{r}_2). \quad (13)$$

By choosing  $x'_3 = -0.5$ , we have a force corresponding to  $L = 108 \text{ MeV}$  (called ‘asy-stiff’ or stiff symmetry energy) with the same equation of state of symmetric nuclear matter and with the same  $S_0$  as the original SLy4 force.

The Skyrme-type parametrization of the effective interaction is advantageous for the efficient AMD computation. However, the Skyrme forces have a quadratic momentum dependence of the mean field which is not valid at high energy collisions at several hundred MeV/nucleon. Therefore, the momentum dependence is corrected for the present calculations in a similar way to Ref. [21] in BUU-type calculations. The detailed formulation in the case of AMD is given in the Appendix A.

### 2. Collision term with and without clusters

The two-nucleon collision process corresponds to the collision term  $I_N^{(\lambda=0)}[f_N, 0, 0]$ . In AMD, a two-nucleon collision is treated as a stochastic transition from an AMD state  $|\Phi_i\rangle$  to another AMD state  $|\Phi_f\rangle$  specified by

the relative momentum between the scattered two nucleons ( $p_{\text{rel}}, \Omega$ ). The transition rate is expressed as

$$vd\sigma = \frac{2\pi}{\hbar} |\langle \Phi_f | V | \Phi_i \rangle|^2 \delta(E_f - E_i) \frac{p_{\text{rel}}^2 dp_{\text{rel}} d\Omega}{(2\pi\hbar)^3}. \quad (14)$$

In general, medium modification is introduced for the scattering matrix elements. However, in the present calculations, we employ the matrix elements in the free space. It should be noted that some medium effect still exists in the  $p_{\text{rel}}$ -dependence of the final state energy  $E_f$  in a similar way to a BUU calculation of Ref. [13]. The Pauli blocking for the scattered nucleons is taken into account.

In the usual treatment of two-nucleon collisions, only the states of the scattered two nucleons are changed in the final state  $|\Phi_f\rangle$  (see Ref. [16] for the precise description of the method which employs ‘physical coordinates’). On the other hand, an extension has been introduced [18] to allow direct formation of light clusters with  $A = 2, 3$  and 4 in the final state  $|\Phi_f\rangle$ . Namely, in the calculation with cluster correlations, when two nucleons  $N_1$  and  $N_2$  collide, we consider the process

$$N_1 + N_2 + B_1 + B_2 \rightarrow C_1 + C_2 \quad (15)$$

in which each of the scattered nucleons  $N_j$  ( $j = 1, 2$ ) may form a cluster  $C_j$  with a spectator particle  $B_j$ . This process includes the collisions without cluster formation as the special case of  $C_j = N_j$  with empty  $B_j$ . The transition rate of the cluster-forming process is given by Eq. (14) with the suitable choice of the final state  $|\Phi_f\rangle$ . When a cluster is formed, the corresponding wave packets are placed at the same phase-space point, i.e., the cluster internal state is represented by the harmonic-oscillator  $(0s)^n$  configuration. Denoting the initial and final states of the  $N_j + B_j$  system by  $|\varphi_j\rangle$  and  $|\varphi'_j\rangle$ , respectively, we have the transition rate

$$vd\sigma = \frac{2\pi}{\hbar} |\langle \varphi'_1 | \varphi_1^{\mathbf{q}} \rangle|^2 |\langle \varphi'_2 | \varphi_2^{-\mathbf{q}} \rangle|^2 \times |M|^2 \delta(E_f - E_i) \frac{p_{\text{rel}}^2 dp_{\text{rel}} d\Omega}{(2\pi\hbar)^3}, \quad (16)$$

where  $|\varphi_j^{\pm\mathbf{q}}\rangle = e^{\pm i\mathbf{q}\cdot\mathbf{r}_j} |\varphi_j\rangle$  are the states after the momentum transfer  $\pm\mathbf{q}$  to the nucleons  $N_j$  ( $j = 1, 2$ ), and  $(p_{\text{rel}}, \Omega)$  is the relative momentum between  $N_1$  and  $N_2$  in these states. The matrix element  $|M|^2$  is the same as for the usual two-nucleon collisions. We use an average value of  $|M|^2$  evaluated at  $p_{\text{rel}}$  and that evaluated at the initial relative momentum.

The actual situation of a two-nucleon collision requires more considerations because there are many possible ways of forming a cluster for each  $N$  of the scattered nucleons  $N_1$  and  $N_2$ . For a scattered nucleon  $N$ , we first consider the possibility that  $N$  may form a cluster with one of the nucleons  $\{B_k; k = 1, 2, \dots\}$  which have the same spin-isospin state. This spin-isospin state that is studied first is randomly decided. The cluster-formed

state is denoted by  $|\Phi'_k\rangle$  which is obtained, by first changing the state to  $|\Phi^{\mathbf{q}}\rangle$  by the momentum transfer  $\mathbf{q}$  to  $N$ , and then moving the two wave packets of  $N$  and  $B_k$  to the same phase-space point without changing their center of mass. Since the different final states are not orthogonal  $\mathcal{N}_{kl} = \langle \Phi'_k | \Phi'_l \rangle \neq \delta_{kl}$ , the probability that  $N$  forms a cluster with one of  $\{B_k\}$  should be calculated as

$$P = \sum_{kl} \langle \Phi^{\mathbf{q}} | \Phi'_k \rangle \mathcal{N}_{kl}^{-1} \langle \Phi'_l | \Phi^{\mathbf{q}} \rangle = \sum_k |v_k|^2, \quad (17)$$

$$v_k = \sum_l \mathcal{N}_{kl}^{-1/2} \langle \Phi'_l | \Phi^{\mathbf{q}} \rangle, \quad (18)$$

This probability is calculated with an approximation that the many-body state is a direct product of wave packets centered at the ‘physical coordinates’ [16]. With the calculated probability  $P$ , a cluster will be formed with one of  $\{B_k\}$ . It is somewhat arbitrary which one of  $\{B_k\}$  should be chosen with what probability. In the present calculation, we choose  $B_k$  with the relative weight  $|v_k|^{2\gamma}$  with the parameter  $\gamma = 2.0$ . With the rest of the probability  $(1 - P)$ ,  $N$  does not form a cluster with a nucleon of this spin-isospin state. The procedure is repeated for other spin-isospin states for  $\{B_k\}$ . The particle  $N$  should be regarded as a cluster, instead of a scattered nucleon, if a (sub)cluster has been already formed in previous steps of the repetition. Thus the formation of light clusters is considered up to an  $\alpha$  particle. This procedure determines the probability  $P_{C_1 C_2}(p_{\text{rel}}, \Omega)$  for the combination of final clusters  $(C_1, C_2)$  as a function of the momentum transfer  $\mathbf{q}$  or  $(p_{\text{rel}}, \Omega)$ . It satisfies the normalization  $\sum_{C_1 C_2} P_{C_1 C_2}(p_{\text{rel}}, \Omega) = 1$ . The factor  $|\langle \varphi'_1 | \varphi_1^{\mathbf{q}} \rangle|^2 |\langle \varphi'_2 | \varphi_2^{-\mathbf{q}} \rangle|^2$  in Eq. (16) should be replaced by  $P_{C_1 C_2}(p_{\text{rel}}, \Omega)$ .

Even when the cluster formation is introduced, the many-body state is always represented by an AMD wave function which is a Slater determinant of nucleon wave packets. The time evolution of the many-body state is solved just as usual without depending on whether some of the wave packets form clusters due to collisions in the past (except for the cluster-cluster binding process in the next paragraph). This is in contrast to BUU by Danielewicz *et al.* [22] where clusters are treated as new particle species. In our approach, a nucleon in a formed cluster may collide with some other nucleon so that the cluster is broken. It may be the case that the scattered nucleon forms the same cluster as before, so that an elastic scattering of the cluster is possible. All of these kinds of processes are based on the nucleon-nucleon scattering matrix elements, without introducing parameters to control individual channels of cluster formation. In the present calculations with cluster correlations, however, the overall cluster production probability is suppressed, when the momentum transfer is extremely small, by a factor  $1 - \exp[-\mathbf{q}^2/(50 \text{ MeV}/c)^2]$ .

It has been found that we should take account of the correlations to form heavier fragments via coalescence of light clusters on top of the usual time evolution of AMD

[18]. This option of improvement has been turned on in the present calculations even though it will not strongly influence the following discussions in high energy collisions. The details are described in Appendix B for completeness.

It is experimentally clearly known that the clusters are important in many situations of heavy-ion collisions even though the incident energy is relatively high. For example, the FOPI data [14] show that only 21% of the total protons in the Au + Au system are emitted as free protons in central collisions at 250 MeV/nucleon, and all the other protons are bound in light clusters and heavier fragments. AMD without cluster correlations overestimates the proton multiplicity as many other transport models do. On the other hand, the AMD calculation with cluster correlations can reproduce this feature very well as shown in Ref. [18] for systems including the Au + Au central collisions at several hundred MeV/nucleon. Therefore, the calculation with clusters is believed to be much closer to the realistic case than without clusters. In the following section, however, we are going to show the results of both calculations with and without clusters, which are useful for the purpose to study the dependence of the pion production on the nucleon dynamics.

### 3. Test particles

The information of nucleon dynamics calculated by AMD is sent to the JAM calculation in the form of a set of test particles  $(\mathbf{r}_1, \mathbf{p}_1), (\mathbf{r}_2, \mathbf{p}_2), \dots, (\mathbf{r}_A, \mathbf{p}_A)$ . One test particle per nucleon is generated following the distribution function defined by Eq. (11). It should be noted that all kinds of quantum effects from the antisymmetrization of the many-body state are contained in this distribution function (or the Wigner function). In particular, it is not positive definite, and therefore in the phase-space region of  $f_\alpha(\mathbf{r}, \mathbf{p}) < 0$  the probability has to be replaced by zero, which can potentially introduce some inaccuracy of the test-particle representation. To check the accuracy, we compared the density profile for the ground state of the Au nucleus, to find no visible difference between the distribution of the generated test particles and the exact density profile. Therefore, this method of test particles should be sufficiently accurate in highly excited situations during heavy-ion collisions. The method to generate the test particles is described in Appendix C.

The set of test particles is sent to the JAM calculation at every 2 fm/c. We have checked that the result does not change when it is sent at every 1 fm/c.

The AMD calculation is much more time consuming than other transport models such as JAM. We typically generated 1000 AMD events for each case of the present calculations. However, we improve the statistics for the pion production by generating 500 JAM events from the same AMD event. As we will see in the next section, a sufficient statistical accuracy is obtained for the pion production with this limited number of AMD events.

## C. JAM

JAM is a transport model which is developed by Nara *et al.* [17]. This model has been successfully applied to high-energy collisions up to more than one hundred GeV/nucleon. In this model, in the energy domain relevant for the present work, the hadron-hadron reactions are treated by the cross sections based on experimental data and the detailed balance. In particular, in the present work, isospin symmetry is assumed. The cross section for the  $NN \rightarrow N\Delta$  process at the c.m. energy  $\sqrt{s}$  is written as [23, 24]

$$\sigma_{NN \rightarrow N\Delta} = \frac{C_I |\mathcal{M}|^2}{p_i s 16\pi} \times \int dm \frac{2}{\pi} \frac{m^2 \Gamma(m)}{(m^2 - m_\Delta^2)^2 + m^2 \Gamma(m)^2} p_f(m) \quad (19)$$

where  $p_i$  and  $p_f(m)$  are the initial and final momenta in the c.m. frame, and the Clebsh-Gordan factor is  $C_I = \frac{1}{4}$  or  $\frac{3}{4}$ . The matrix element  $|\mathcal{M}|$  is assumed as

$$|\mathcal{M}|^2 = A \frac{s \Gamma_\Delta^2}{(s - m_\Delta^2)^2 + s \Gamma_\Delta^2} \quad (20)$$

with  $\Gamma_\Delta = 0.118$  GeV and  $m_\Delta = 1.232$  GeV. This is a similar parametrization to UrQMD in Ref. [25]. The constant  $A/16\pi = 64400$  mb GeV<sup>2</sup> is determined to fit the data of the  $NN \rightarrow NN\pi$  cross sections. The distribution of the mass  $m$  of the  $\Delta$  resonance is determined by the integrand of Eq. (19).  $\Gamma(m)$  is the decay width for  $\Delta \rightarrow N\pi$  parametrized as in Refs. [17, 25]. In order to fit the data precisely near the pion threshold, non-resonant contributions and/or some components that violates isospin symmetry are necessary for the  $NN \rightarrow NN\pi$  processes, which are ignored, however, in the present work so that the pions are produced only through the formation of resonances. In our JAM calculation, the mean field for nucleons is not included.

To know how the JAM calculation reproduces the pion multiplicity of the experimental data, we have calculated Au + Au collisions at various incident energies for the impact parameters  $0 < b < 2$  fm. Figure 1 shows the energy dependence of the pion multiplicity. The solid lines indicate the pion multiplicities calculated by JAM. The points show the experimental data taken from the FOPI measurement of Au + Au collisions at 0.4, 0.8 and 1.5 GeV/nucleon [14]. We find that the JAM calculation almost reproduces the experimental data of pion multiplicities reasonably well. However, the calculation overestimates the pion multiplicities in particular at lower energies. This is probably because the JAM calculation here does not include the mean field potential, and therefore more energy is available for particle production without the cost of energy to compress the system [22, 26].

We have also checked the time evolution of the numbers of pions and  $\Delta$  resonances in the JAM calculations as shown in Fig. 2 for the central Au + Au collisions

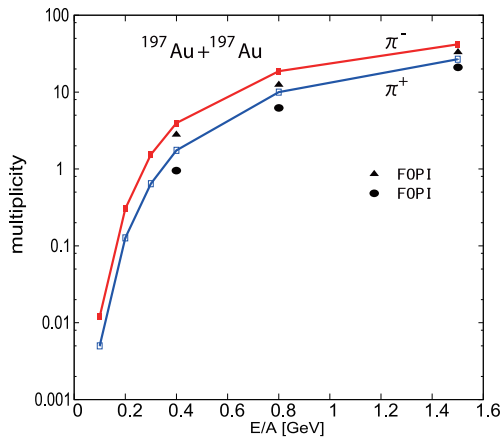


FIG. 1. Incident energy dependence of the charged pion multiplicities in central Au + Au collisions. The two lines indicate the multiplicities of  $\pi^-$  and  $\pi^+$ , respectively, calculated by JAM. Symbols represent the experimental data of the FOPI Collaboration [14].

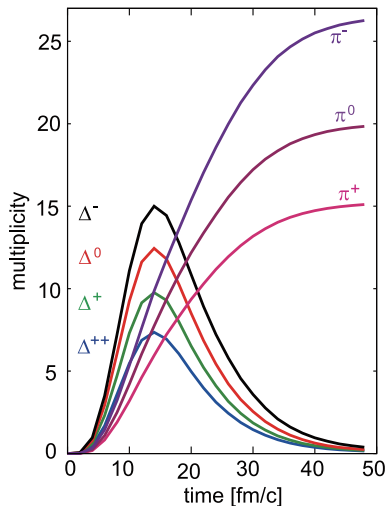


FIG. 2. Time evolution of the numbers of pions and  $\Delta$  resonances calculated by JAM for the central Au + Au collisions at 1 GeV/nucleon with the impact parameter  $b = 0$ .

at 1 GeV/nucleon. The result shows a similar behavior to a relativistic mean-field transport calculation reported in Fig. 1 of Ref. [27].

Thus JAM calculation provides a reasonable description for the pions and  $\Delta$  resonances produced in collisions at around 1 GeV/nucleon. It should be noted that in the AMD+JAM calculation the mean field for nucleons is taken into account in AMD and the ignored mean field in JAM does not actually influence the results.

The production and absorption reactions for  $\Delta$  and pions occur in the JAM calculation as in the free space without medium modification for the thresholds, while nucleons feel potential in the AMD calculation. This

corresponds to assuming that the potentials  $U_\tau^{(\Delta)}$  and  $U_\tau^{(\pi)}$  for  $\Delta$  and pions are related to the isospin-dependent nucleon potential  $U_\tau^{(N)} = U_0 + \tau U_{\text{sym}}$  as

$$U_\tau^{(\Delta)} = U_0 + \tau U_{\text{sym}}, \quad U_\tau^{(\pi)} = \tau U_{\text{sym}}, \quad (21)$$

where  $\tau$  is the isospin component. This is equivalent to the choice in the pBUU calculation of Ref. [12] if the momentum dependence is ignored. It should be noted that other transport calculations take a different choice [4, 8]. The details of the in-medium effects for  $\Delta$  and pions may influence the pion yields as investigated in equilibrium calculations [28, 29]. On the other hand, recently, the effects of the completely unknown symmetry (isovector) potential of the  $\Delta$  resonance on the pions in heavy-ion collisions have been studied in Ref [30]. It has been reported that these effects are negligible except for at deeply subthreshold energies and thus the pion ratio is still a good probe to investigate the high-density symmetry energy.

### III. RESULTS

We calculated collisions of  $^{132}\text{Sn} + ^{124}\text{Sn}$  at 300 MeV/nucleon for the impact parameters  $0 < b < 1$  fm. At the initial time  $t = 0$ , the centers of the two nuclei are separated by 15 fm. In order to investigate the relation between the high-density symmetry energy, the nucleon dynamics in a compressed neutron-rich system and the emitted pions, we are going to compare the results from five different cases as follows,

1. AMD + JAM with clusters (asy-soft)
2. AMD + JAM with clusters (asy-stiff)
3. AMD + JAM without clusters (asy-soft)
4. AMD + JAM without clusters (asy-stiff)
5. JAM (no mean field).

The first two cases are calculated with AMD with cluster correlations, and the next two cases are calculated without cluster correlations. For each of them, we have calculated with two different effective interactions ('asy-soft' and 'asy-stiff') for different density dependence of symmetry energy. We have also performed a standard JAM calculation without combining with AMD in order to clarify the effect of the mean field and the symmetry energy by the comparison with the AMD+JAM calculations.

#### A. Neutron-proton dynamics

To see the dynamics of neutrons and protons in the five different calculations, Fig. 3 shows some information on

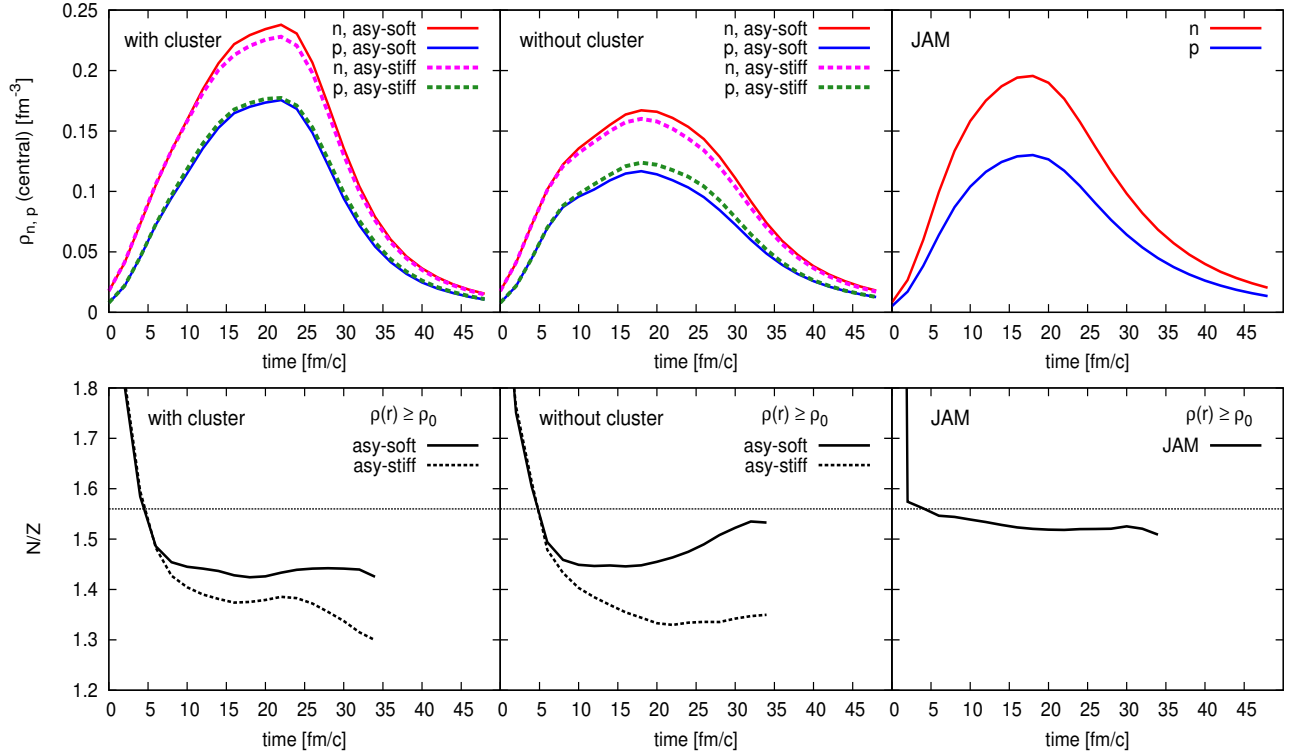


FIG. 3. Time evolution of the neutron and proton densities in the central part of the system within the radius of 2 fm (upper) and the ratio of neutrons and protons (lower) in central collisions of  $^{132}\text{Sn} + ^{124}\text{Sn}$  at 300 MeV/nucleon. The left and middle panels show the results of AMD+JAM with clusters and without clusters, respectively. The right panels are for the simple JAM calculation. Two different lines (solid and dotted) correspond to the two different density dependence of symmetry energy. The horizontal lines in the bottom panels represent the ratio of the total system  $(N/Z)_{\text{sys}} = 1.56$ .

the time evolution of the densities of protons and neutrons. The upper panels show the neutron and proton densities in the radius of 2 fm from the center-of-mass of the system. The maximum density  $\rho = \rho_n + \rho_p \sim 2\rho_0$  is reached at  $t \simeq 20$  fm/c. However, we find that the maximum density is higher with cluster correlation than without it. The time of the maximum density also depends on the cluster correlation. One of the possible reasons for this is that the cluster formation has an effect to gather nucleons spatially so that the compression of the central part of the system continues longer. In the case of JAM calculation, the higher maximum density is reached at an earlier time than in the AMD without clusters, which is reasonable because the mean-field potential for nucleons is not included in JAM.

The lower panels of Fig. 3 show the time evolution of the neutron-to-proton ratio  $N/Z$  calculated for the central region within a radius  $r_0$  from the center of mass. The radius  $r_0$  is determined at each time in each event by the condition  $\rho(r_0) \approx \rho_0$ , where the spherically averaged density  $\rho(r)$  is evaluated by using the set of test particles. The results with the soft and stiff symmetry energies are shown by the solid and dashed lines, respectively, for each case of with and without cluster correlation. In all the cases, the  $N/Z$  ratio of the compressed part (at  $t \gtrsim 10$

fm/c) becomes smaller than the that of the total system  $(N/Z)_{\text{sys}} = 1.56$ , which is consistent with the symmetry energy effect that does not favor high-density neutron-rich matter. We can see clearly that this effect to reduce  $N/Z$  of the compressed part is stronger with the stiff symmetry energy. This symmetry energy effect is consistent with the results of other transport models [4] at least qualitatively. However, there may be model dependence in the quantitative values of  $N/Z$ . In fact, in our calculations here, the symmetry energy effect is stronger without cluster correlation than with cluster correlation.

## B. $\Delta$ and pions

In Fig. 4, we show the reaction rates of  $\Delta$  production ( $NN \rightarrow N\Delta$ ) and absorption ( $N\Delta \rightarrow NN$ ) in the upper and middle panels, respectively. The time evolution of these rates for  $\Delta^-$ ,  $\Delta^0$ ,  $\Delta^+$  and  $\Delta^{++}$  are shown by the four lines. We can see that about 70% of the produced  $\Delta$  resonances are absorbed and turned back to nucleons. We also show the numbers of existing  $\Delta$  resonances and pions in the lower panel of Fig. 4.

In Fig. 5, we show the time evolution of the  $\pi_{\text{like}}^-/\pi_{\text{like}}^+$  ratio for the five different cases. The  $\pi_{\text{like}}$  particles are

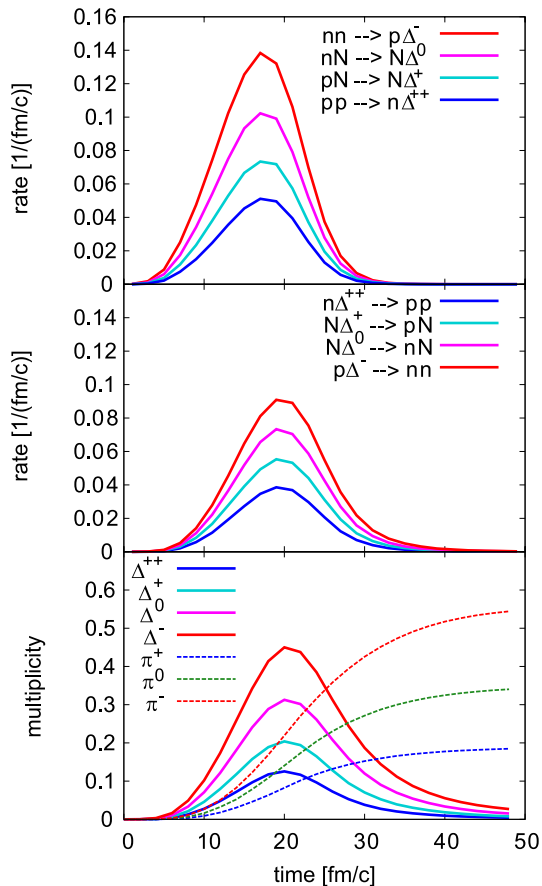


FIG. 4. Reaction rates of  $\Delta$  production (upper) and  $\Delta$  absorption (middle), and the numbers of existing  $\Delta$  resonances and pions (lower) as functions of time, in the AMD+JAM calculation with clusters (asy-soft) for central collisions of  $^{132}\text{Sn} + ^{124}\text{Sn}$  at 300 MeV/nucleon. For the production and absorption of  $\Delta^+$  and  $\Delta^0$ , the line shows the sum of the reaction rates with a neutron ( $N = n$ ) and with a proton ( $N = p$ ).

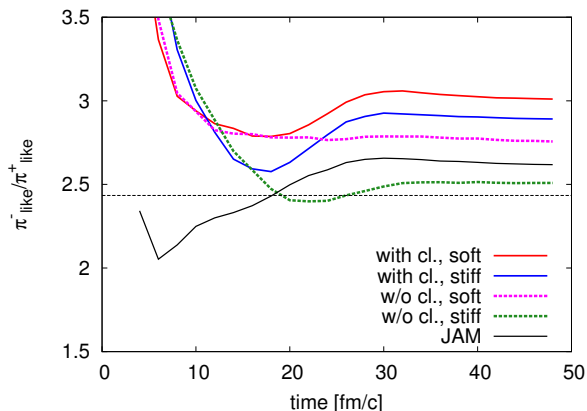


FIG. 5. The time evolution of the  $\pi_{\text{like}}^-/\pi_{\text{like}}^+$  ratios in central collisions of  $^{132}\text{Sn} + ^{124}\text{Sn}$  at 300 MeV/nucleon in the five cases of calculation. The horizontal line represents the  $(N/Z)_{\text{sys}}^2$  ratio of the total system.

defined, including the  $\Delta$  resonances depending on the branching ratio to decay into the pion, as

$$\pi_{\text{like}}^- = \pi^- + \Delta^- + \frac{1}{3}\Delta^0, \quad (22)$$

$$\pi_{\text{like}}^0 = \pi^0 + \frac{2}{3}\Delta^0 + \frac{2}{3}\Delta^+, \quad (23)$$

$$\pi_{\text{like}}^+ = \pi^+ + \Delta^{++} + \frac{1}{3}\Delta^+. \quad (24)$$

Our calculation predicts that the evolution of the pion-like ratio has a dependence on the symmetry energy. The pion-like ratio calculated with the soft symmetry energy is larger than that with the stiff symmetry energy in both of calculations with and without cluster correlations. This result seems to be similar to the predictions reported in Refs. [4, 10] qualitatively. We also find that the pion-like ratio depends on the cluster correlations and that the symmetry-energy effect appears in the pion-like ratio stronger in the case without cluster correlations. This is consistent with the result of the neutron-proton dynamics as shown in the lower panels of Fig. 3, suggesting a possibility that the pion production is really related to the neutron-proton dynamics.

The pion-like ratio reaches the final  $\pi^-/\pi^+$  value at around  $t \simeq 30$  fm/c. In all the five cases, the predicted final  $\pi^-/\pi^+$  ratios become larger than  $(N/Z)_{\text{sys}}^2 = 2.4336$  of the total system, which is consistent with the experimental observation for Au + Au system at 400 MeV/nucleon [9] and is suggesting that the relation  $\pi^-/\pi^+ \approx (N/Z)^2$  does not hold if  $(N/Z)$  is taken from Fig. 3. On the other hand, the behaviors of the pion-like ratio before  $t \simeq 30$  fm/c are complicated. The origin of these behaviors will be better understood through the analysis in the next subsections.

### C. Relation of nucleon dynamics and $\Delta$ production

As mentioned above, the calculated pion ratio becomes larger than the  $(N/Z)^2$  ratio of the compressed part of the system. To find the origin of this effect, we investigate what kind of information of neutrons and protons is carried by  $\Delta$  resonances. Since  $\Delta^-$  and  $\Delta^{++}$  are produced only by the  $nn \rightarrow p\Delta^-$  and  $pp \rightarrow n\Delta^{++}$  reactions, respectively, we expect that the  $\Delta^-/\Delta^{++}$  ratio of the production rates of these resonances should be most directly linked to some kind of  $(N/Z)^2$  ratio of nucleons. Figure 6 shows the ratio of the production reaction rates of  $\Delta^-$  and  $\Delta^{++}$  as functions of time for the five different calculations. The  $\Delta$  production is peaked around  $t = 15$ - $20$  fm/c as shown in Fig. 4, and the ratio does not have a significant meaning at very early and late times. This  $\Delta$  production ratio is compared with Fig. 7(a) which shows the same information as in the lower panels of Fig. 3 but shows the squared ratio  $(N/Z)^2$ . We find clearly that the  $\Delta$  production ratio is much larger than this  $(N/Z)^2$  ratio in the high-density region of the system, except for the JAM case in which both ratios are close to the  $(N/Z)_{\text{sys}}^2$  ratio of the total system. The relative ordering of the ratios for the five cases also disagrees between Fig. 6 and



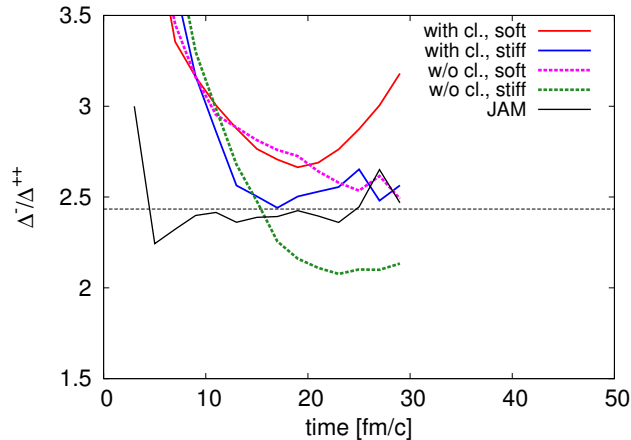


FIG. 6. The time evolution of the  $\Delta^-/\Delta^{++}$  ratio of the  $\Delta$  production rates. The five different lines show the calculations. The horizontal line represents the  $(N/Z)_{\text{sys}}^2$  ratio of the total system.

Fig. 7(a). This result shows that the  $\Delta$  production, and therefore the pion production, are not simply linked to the  $N/Z$  ratio of the high-density part of the system.

Each panel of Fig. 7 shows the time evolution of the squared neutron-to-proton ratio  $(N/Z)^2$  calculated for the nucleons that satisfy the following condition.

- (a) Nucleons in the sphere  $\rho(r) \geq \rho_0$  centered at the center-of-mass of the system.
- (b) Nucleons with high momenta  $|\mathbf{p} - \mathbf{p}_{\text{rad}}| \geq p_{\text{cut}}$  in the sphere  $\rho(r) \geq \rho_0$  centered at the center-of-mass of the system. We take  $p_{\text{cut}} = 480$  MeV/c. The collective radial momentum  $\mathbf{p}_{\text{rad}}$  is subtracted from the nucleon momentum  $\mathbf{p}$ .

For the condition (a), we choose the nucleons in the high-density central region within a radius  $r_0$  from the center of mass. The radius  $r_0$  is determined by the condition  $\rho(r_0) \approx \rho_0$  as described in the subsection A. For the condition (b), we choose only the nucleons with momenta  $|\mathbf{p} - \mathbf{p}_{\text{rad}}|$  larger than  $p_{\text{cut}} = 480$  MeV/c in the center-of-mass system, in addition to the condition (a). The collective radial momentum  $\mathbf{p}_{\text{rad}} = p_{\text{rad}}(r)\mathbf{r}/r$  is subtracted for this condition, where  $p_{\text{rad}}(r)$  is the radial momentum component averaged for the nucleons on the sphere of the radius  $r$ . It is natural to consider this kind of momentum condition since a sufficient energy is required to excite a  $\Delta$  resonance in a two-nucleon collision. Our choice of  $p_{\text{cut}}$  corresponds to  $p_{\text{cut}}^2/m_N = 245$  MeV while  $m_\Delta - m_N = 293$  MeV. We have also checked that the changes of  $p_{\text{cut}}$  by  $\pm 20$  MeV/c result in  $\pm 3\%$  differences in  $(N/Z)^2$  (for the average values shown in Fig. 8).

The panels of Fig. 7 correspond to the results of the time evolution of the  $(N/Z)^2$  ratio calculated for the nucleons satisfying the conditions (a) and (b), respectively. We find that when the high momentum condition is imposed, the  $(N/Z)^2$  ratio changes drastically compared to

that without the momentum condition. The  $(N/Z)^2$  ratio in condition (b) becomes larger than that in condition (a).

By comparing the  $\Delta^-/\Delta^{++}$  production ratio in Fig. 6 and the  $(N/Z)^2$  ratio in Fig. 7(a) for each condition, we have already seen that  $\Delta^-/\Delta^{++}$  and  $(N/Z)^2$  do not agree if the nucleons are selected only by the high-density condition (a). On the other hand, in the result of the condition (b),  $\Delta^-/\Delta^{++}$  is quite similar to  $(N/Z)^2$ , i.e., the relation  $\Delta^-/\Delta^{++} \simeq (N/Z)^2$  holds as a function of time. Thus, we can conclude that  $\Delta$  resonances, and hopefully pions, carry direct information on nucleons in high density and high momentum region of the one-body phase space. We also mention that the agreement is not as perfect as in the case (b) at  $t = 20$ -30 fm/c if the collective radial momentum is not subtracted to define the condition.

#### D. From nucleons to pion ratios

To discuss the relation between the dynamics of nucleons and the final pion ratio, we show in Fig. 8 a summary of different ratios in five calculations.

For the time-dependent nucleon ratio  $[N(t)/Z(t)]^2$  shown in Fig. 7, we here define a representative  $(N/Z)^2$  ratio as

$$\left(\frac{N}{Z}\right)^2 = \frac{\int_0^\infty N(t)^2 dt}{\int_0^\infty Z(t)^2 dt}, \quad (25)$$

where  $N(t)$  and  $Z(t)$  indicate the numbers of neutrons and protons as functions of time which satisfy the conditions described in the previous subsection. This ratio carries the information in the compression stage because  $N(t)$  and  $Z(t)$  take large values only if the system has a large high-density region  $\rho > \rho_0$ . In the first two columns of Fig. 8, the  $(N/Z)_\rho^2$  and  $(N/Z)_{\rho,p}^2$  ratios show the calculated representative values when the nucleons are selected by the conditions (a) and (b), respectively. As we have already seen in the previous subsection, the  $(N/Z)^2$  increases by choosing the high momentum part of the phase space. The effect is stronger in the calculations with cluster correlations. This is understandable because, with cluster correlations, many  $\alpha$  clusters are formed which contain the same number of neutrons and protons and have relatively low momentum per nucleon, and therefore the remaining part of high-momentum nucleons becomes neutron rich. Another important point found here is that the symmetry energy effect, namely the difference between stiff and soft symmetry energies, is smaller when the cluster correlation is turned on. This is also reasonable because the cluster correlation forces some neutrons and protons to move together, and therefore the different forces acting on neutrons and protons are averaged out to some degree.

As a representative value of the  $\Delta$  production ratio, we show, in the third column of Fig. 8, the  $\Delta^-/\Delta^{++}$  ratio

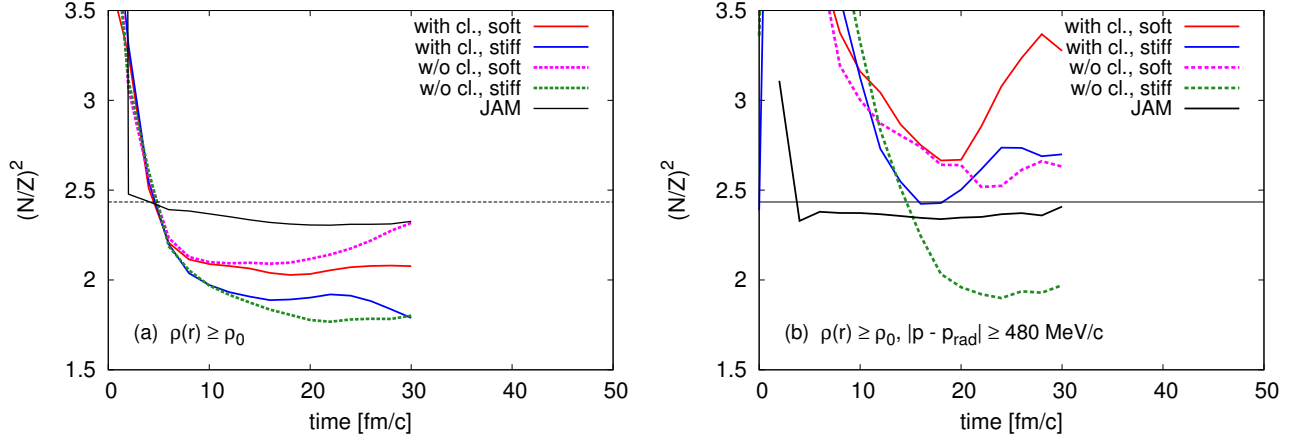


FIG. 7. The time evolution of the squared ratio of neutron and proton  $(N/Z)^2$ . The left and right figures show the  $(N/Z)^2$  ratio calculated for the nucleons in the condition (a) and (b), respectively. See details in the text.

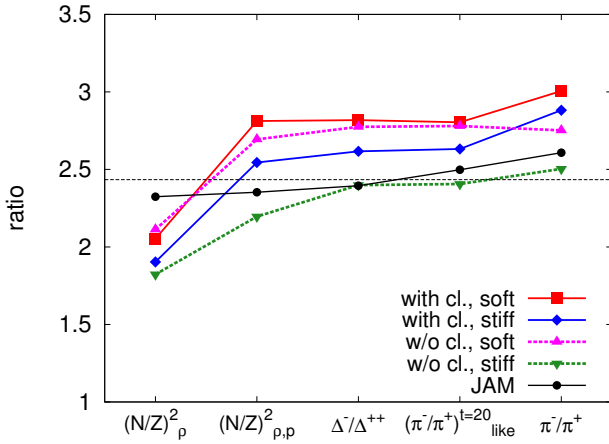


FIG. 8. The nucleon ratios  $(N/Z)_\rho^2$  and  $(N/Z)_{\rho,p}^2$  [Eq. (25)] in high-density region with and without high-momentum condition, respectively, the  $\Delta^-/\Delta^{++}$  production ratio [Eq. (26)], the pion-like ratio at  $t = 20$  fm/c, and the final  $\pi^-/\pi^+$  ratio. Each line connects the ratios for each of the five cases of calculation for central collisions of  $^{132}\text{Sn} + ^{124}\text{Sn}$  at 300 MeV/nucleon. The horizontal line represents the  $(N/Z)_{\text{sys}}^2$  ratio of the total system. The statistical uncertainties in the final  $\pi^-/\pi^+$  ratio are smaller than 0.02.

of the total production numbers

$$\frac{\Delta^-}{\Delta^{++}} = \frac{\int_0^\infty (nn \rightarrow p\Delta^-) dt}{\int_0^\infty (pp \rightarrow n\Delta^{++}) dt}, \quad (26)$$

where  $(nn \rightarrow p\Delta^-)$  and  $(pp \rightarrow n\Delta^{++})$  indicate the reaction rates of the  $\Delta$  production at each time shown in the upper panel of Fig. 4. We can see that the  $\Delta^-/\Delta^{++}$  ratio is different from the  $(N/Z)_\rho^2$  ratio, while the  $(N/Z)_{\rho,p}^2$  ratio is almost equal to the  $\Delta^-/\Delta^{++}$  ratio. These results are consistent with the comparison of Fig. 6 and Fig. 7.

It might not be straightforward, in principle, how the  $\Delta$  production ratio is related to the pion ratio because many of the produced  $\Delta$  resonances are absorbed by  $N\Delta \rightarrow NN$  reactions as we have seen in Fig. 4. However, we find that the behavior of the  $\Delta$  production ratio before  $t \approx 20$  fm/c in Fig. 6 is similar to the pion-like ratio in Fig. 5 calculated from the  $\Delta$  resonances and pions that exist at each time. The ratio of numbers of  $\Delta$  resonances at  $t = 20$  fm/c is  $\Delta^{++} : \Delta^+ : \Delta^0 : \Delta^- \approx 1 : 1.62 : 2.49 : 3.58$  in the lower panel of Fig. 4 for the case of soft symmetry energy with cluster correlations. The  $\Delta^-$  to  $\Delta^{++}$  ratio is significantly larger than the  $\Delta$  production ratio because the neutron-richness of the system influences the isospin dependence of the  $\Delta$  absorption rates<sup>1</sup>. The pion-like ratio calculated from only these  $\Delta$  resonances is  $(\pi^-/\pi^+)_{\text{like}}' = (\Delta^- + \frac{1}{3}\Delta^0)/(\Delta^{++} + \frac{1}{3}\Delta^+) = 2.86$  which happens to be similar to the  $\Delta$  production ratio. The fourth column of Fig. 8 shows the pion-like ratio at  $t = 20$  fm/c (see Fig. 5). An interesting observation is that the dependence of the pion-like ratio on the symmetry energy and the cluster correlations is quite similar to that of the  $\Delta$  production ratio and therefore to that of  $(N/Z)_{\rho,p}^2$ . This suggests that the information on the high-density nucleon dynamics remains in the pion-like ratio at  $t = 20$  fm/c, without being much influenced by the  $\Delta$  absorption.

The pion-like particles have to go through the exterior region of the expanding system. The symmetry-energy

<sup>1</sup> If the system were in chemical equilibrium with a temperature  $T$  and neutron and proton chemical potentials  $\mu_n$  and  $\mu_p$ , one would expect the  $\Delta$  multiplicity ratio might follow  $\Delta^-/\Delta^{++} \simeq e^{3(\mu_n - \mu_p)/T} \simeq (N/Z)_{\rho,p}^3$  for  $(m_\Delta - \mu_{n,p})/T \gg 1$ . In our simulation, however,  $\Delta^-/\Delta^{++} = 3.58$  is not as large as  $(N/Z)_{\rho,p}^3 = 4.7$ . This implies that dynamical effects in the neutron and proton distributions are important and/or the full chemical equilibrium for  $\Delta$  resonances is not achieved.

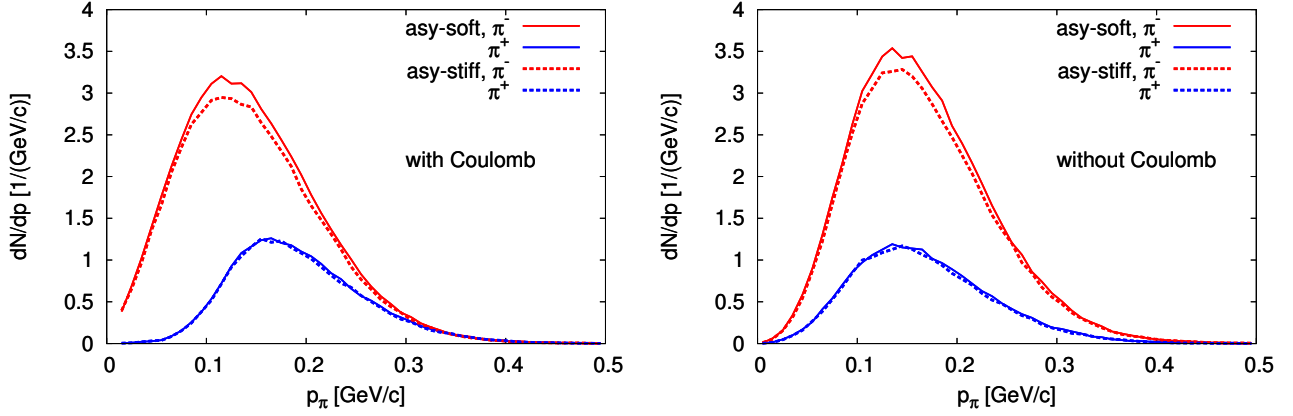


FIG. 9. Momentum distribution of  $\pi^-$  and  $\pi^+$  in the center-of-mass frame calculated with AMD+JAM with clusters for central collisions of  $^{132}\text{Sn} + ^{124}\text{Sn}$  at 300 MeV/nucleon. The left and right panels show the results with and without Coulomb force for pions, respectively.

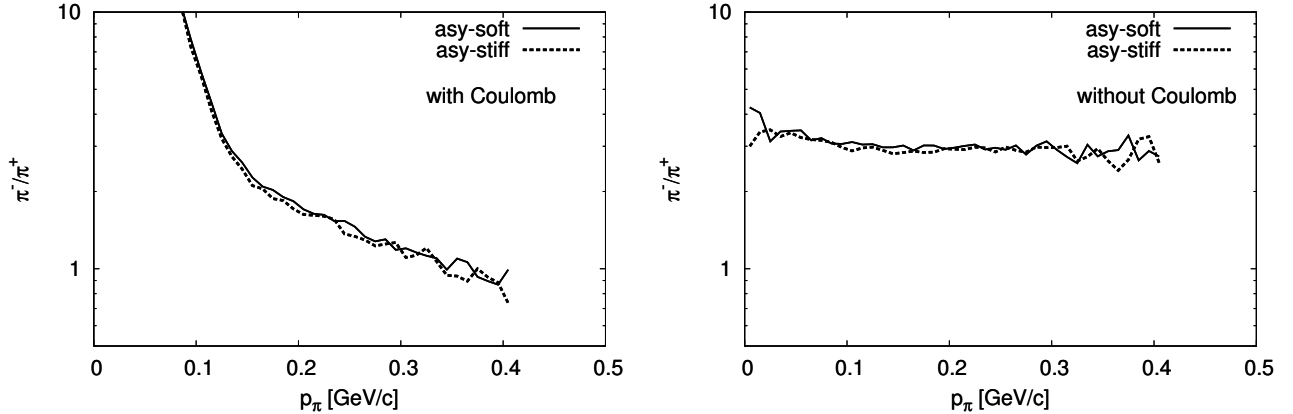


FIG. 10. The same as Fig. 9 but the  $\pi^-/\pi^+$  spectral ratio is shown.

effect on the nucleon  $N/Z$  ratio in the exterior region should be opposite to that in the inner region because the total numbers of neutrons and protons are (almost) conserved. As shown in the rightmost column of Fig. 8, in the final stage of the reaction, the pion ratio is modified to some degree from the pion-like ratio at  $t = 20$  fm/ $c$  to the final  $\pi^-/\pi^+$  ratio. In each case of with and without clusters, the symmetry energy effect at  $t = 20$  fm/ $c$  is reduced in the final ratio to about 70% of the value at  $t = 20$  fm/ $c$ . We also find the effect of clusters tends to raise the pion ratio in the final stage, which is probably because the interior/exterior part becomes less/more neutron rich when more  $\alpha$  clusters, with the same number of neutrons and protons contained, are formed with relatively low velocities.

### E. Pion spectra

Finally, we investigate the pion spectra. Fig. 9 shows the pion spectra with or without Coulomb force for charged pions. Coulomb force evidently changes the pion spectrum because  $\pi^+$  is accelerated and  $\pi^-$  is decelerated. The  $\pi^-/\pi^+$  spectral ratio shown in Fig. 10 can become very large at low pion momenta due to the Coulomb effect.

The symmetry-energy effect, however, appears in our present results as a simple normalization factor for the pion spectra and the spectral ratio, and thus we cannot obtain information on the symmetry energy effect more than found in the ratio of the total pion multiplicities. This point does not agree with what has been found in the pBUU calculation [12, 31] in which the symmetry-energy effect is strong in the high momentum part of the spectra.

#### IV. SUMMARY

The mechanism of pion production was studied with a new approach by combining two transport models AMD and JAM. For central  $^{132}\text{Sn} + ^{124}\text{Sn}$  collisions at 300 MeV/nucleon, the production of  $\Delta$  resonances and pions are treated as perturbation. Two different AMD calculations with and without cluster correlations were performed, not only to investigate the effect of clusters but also to study the correlations between the nucleon dynamics and  $\Delta$  and pion production. We found that the  $\Delta^-/\Delta^{++}$  production ratio agrees very well with the neutron-proton squared ratio  $(N/Z)^2$  in the high-density and high-momentum region of the one-body phase space. We also found that the  $\Delta$  production ratio, and therefore  $(N/Z)^2$ , are directly reflected in the  $\pi^-/\pi^+$  ratio. The effect of the high-density symmetry energy in the  $\pi^-/\pi^+$  ratio is modified in the final stage of the reaction, with a large part of the effect still remaining, which is qualitatively similar to the case of BUU in the literature [4].

If the AMD calculations with and without clusters are regarded as two different models, the present results show the value of the pion ratio is model-dependent as well as the nucleon dynamics is. Evidently models should be constrained by other observables such as the multiplicities and the spectra of clusters in order to extract the symmetry energy from the pion ratio. It may also possible to find a combination of observables that has less model dependence, such as by taking the ratio of an observable from different reaction systems [3]. It is, nevertheless, preferable to clarify the origin of the different predictions of different models by investigating the dynamics in detail as has been done here.

It is, of course, interesting and necessary to extend the present study to other reaction systems with different neutron-proton asymmetries and for different energies. Such calculations are in progress, and the results will be reported elsewhere.

#### ACKNOWLEDGMENTS

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#### Appendix A: Momentum dependent term

The momentum-dependent term of the interaction energy density for a Skyrme effective interaction can be written by using the phase-space distribution function  $f_\alpha(\mathbf{r}, \mathbf{p})$  ( $\alpha = p \uparrow, p \downarrow, n \uparrow, n \downarrow$ ) as

$$\mathcal{E}_\tau(\mathbf{r}) = \frac{1}{2} \sum_{\alpha\beta} U_{\alpha\beta}^\tau \int \frac{d\mathbf{p}_1}{(2\pi\hbar)^3} \int \frac{d\mathbf{p}_2}{(2\pi\hbar)^3} (\mathbf{p}_1 - \mathbf{p}_2)^2 f_\alpha(\mathbf{r}, \mathbf{p}_1) f_\beta(\mathbf{r}, \mathbf{p}_2). \quad (\text{A1})$$

The coefficients  $U_{\alpha\beta}^\tau$  are related to the Skyrme parameters by

$$U_{\alpha\beta}^\tau = \frac{1}{4} t_1 \langle \alpha\beta | (1 + x_1 P_\sigma) | \alpha\beta - \beta\alpha \rangle + \frac{1}{4} t_2 \langle \alpha\beta | (1 + x_2 P_\sigma) | \alpha\beta + \beta\alpha \rangle. \quad (\text{A2})$$

Employing the fact that the momentum dependence is quadratic, it is possible to write

$$\mathcal{E}_\tau(\mathbf{r}) = \sum_{\alpha\beta} U_{\alpha\beta}^\tau (\tau_\alpha(\mathbf{r}) \rho_\beta(\mathbf{r}) - \mathbf{J}_\alpha(\mathbf{r}) \cdot \mathbf{J}_\beta(\mathbf{r})), \quad (\text{A3})$$

where several kinds of densities are defined by

$$\rho_\alpha(\mathbf{r}) = \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} f_\alpha(\mathbf{r}, \mathbf{p}), \quad (\text{A4})$$

$$\mathbf{J}_\alpha(\mathbf{r}) = \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} (\mathbf{p} - \bar{\mathbf{p}}(\mathbf{r})) f_\alpha(\mathbf{r}, \mathbf{p}), \quad (\text{A5})$$

$$\tau_\alpha(\mathbf{r}) = \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} (\mathbf{p} - \bar{\mathbf{p}}(\mathbf{r}))^2 f_\alpha(\mathbf{r}, \mathbf{p}). \quad (\text{A6})$$

Although  $\bar{\mathbf{p}}(\mathbf{r})$  can be arbitrarily chosen for Eq. (A3) to hold, we will choose the local average momentum

$$\bar{\mathbf{p}}(\mathbf{r}) = \frac{1}{\sum_\alpha \rho_\alpha(\mathbf{r})} \sum_\alpha \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} \mathbf{p} f_\alpha(\mathbf{r}, \mathbf{p}), \quad (\text{A7})$$

so that the term  $\mathbf{J}_\alpha \cdot \mathbf{J}_\beta$  in Eq. (A3) can be neglected keeping the Galilei invariance.

For the distribution function of Eq. (11) in the case of AMD, we can analytically perform the momentum integration to have

$$\tau_\alpha(\mathbf{r}) = \left(\frac{2\nu}{\pi}\right)^{\frac{3}{2}} \sum_{j,k \in \alpha} [(\mathbf{P}_{jk} - \bar{\mathbf{p}}(\mathbf{r}))^2 + 3\hbar^2\nu] \times e^{-2\nu(\mathbf{r} - \mathbf{R}_{jk})^2} B_{jk} B_{kj}^{-1}. \quad (\text{A8})$$

For the application to high energy collisions, we now modify the quadratic momentum dependence to the same momentum dependence as Ref. [21] by modifying  $\tau_\alpha(\mathbf{r})$  to

$$\tilde{\tau}_\alpha(\mathbf{r}) = \left(\frac{2\nu}{\pi}\right)^{\frac{3}{2}} \sum_{j,k \in \alpha} \left[ \frac{(\mathbf{P}_{jk} - \bar{\mathbf{p}}(\mathbf{r}))^2}{1 + (\mathbf{P}_{jk} - \bar{\mathbf{p}}(\mathbf{r}))^2/\Lambda^2} + 3\hbar^2\nu \right] \times e^{-2\nu(\mathbf{r} - \mathbf{R}_{jk})^2} B_{jk} B_{kj}^{-1}, \quad (\text{A9})$$

and then use the modified momentum-dependent part of the interaction energy density

$$\tilde{\mathcal{E}}_\tau(\mathbf{r}) = \sum_{\alpha\beta} U_{\alpha\beta}^\tau \tilde{\tau}_\alpha(\mathbf{r}) \rho_\beta(\mathbf{r}). \quad (\text{A10})$$

In the present work, we choose the parameter  $\Lambda = 395$  MeV/ $c$  for the momentum scale.

## Appendix B: Correlations to bind clusters

Many of light nuclei (Li, Be *etc.*) have only one or a few bound states which may be regarded as bound states of internal clusters. The quantum-mechanical probability of forming such a nucleus is not consistent with the semiclassical phase space with which it can be formed in the standard treatment of AMD. Therefore, for a better description, inter-cluster correlation is introduced as a stochastic process of binding clusters.

The basic idea is to replace the radial component of the relative momentum between clusters by zero if moderately separated clusters ( $2.5 < R_{\text{rel}} < 7$  fm) are moving away from each other with a small relative kinetic energy [ $\mathbf{R}_{\text{rel}} \cdot \mathbf{V}_{\text{rel}} > 0$  and  $\frac{1}{2}\mu(V_{\text{rel}\parallel}^2 + 0.25V_{\text{rel}\perp}^2) < 7$  MeV]. In addition to these conditions, linking is allowed only if each of the two clusters is one of the three closest clusters of the other when the distance is measured by  $[(\mathbf{R}_{\text{rel}}/3 \text{ fm})^2 + (\mathbf{V}_{\text{rel}}/0.25c)^2]^{1/2}$ , so that linking usually occurs in dilute environment. Non-clustered nucleons are treated here in the same way as clusters but two nucleons are not allowed to be linked. Two clusters also should not be linked if they can form an  $\alpha$  or lighter cluster due to the combination of their spins and isospins. It is possible that more than two clusters are linked by this condition. However, only in the case that the mass number of the linked system is  $\leq 10$ , the binding is performed for the linked system by eliminating the radial velocities of clusters in the center-of-mass frame of the linked system.

The energy conservation should be achieved by scaling the relative radial momentum between the center-of-mass of the linked system and a third cluster. A reasonable way to choose a third cluster may be to find a cluster which has participated in a collision that formed one of the clusters in the linked system. However, since we do not keep the full history of collisions in our computation, we choose a cluster that has the minimal value of

$$(r + 7.5 \text{ fm})(1.2 - \cos^2 \theta) / \min(\varepsilon_{\parallel}, 5 \text{ MeV}) \quad (\text{B1})$$

as the third cluster for energy conservation, where  $r$  and  $\varepsilon_{\parallel}$  are the distance and the radial component of the kinetic energy for the relative motion between the linked system and the third cluster. The factor with the angle  $\theta$  between the relative coordinate ( $\mathbf{r}$ ) and velocity ( $\mathbf{v}$ ) is introduced so as to favor the case of  $\mathbf{r} \parallel \mathbf{v}$ .

## Appendix C: Test particles for the AMD phase-space distribution

Here we describe a method to generate test particles following the one-body phase-space distribution function  $f(\mathbf{r}, \mathbf{p})$  given by Eq. (11). We generate  $A$  test particles, where  $A$  is the number of nucleons (for each spin-isospin state) in the system.

Let us first consider a distribution given by a sum of Gaussian distributions

$$g(\mathbf{r}, \mathbf{p}) = \sum_{i=1}^A g_i(\mathbf{r}, \mathbf{p}), \quad (\text{C1})$$

with

$$g_i(\mathbf{r}, \mathbf{p}) = (2\alpha)^3 e^{-2\nu\alpha(\mathbf{r}-\mathbf{R}_i)^2 - \alpha(\mathbf{p}-\mathbf{P}_i)^2/2\hbar^2\nu}, \quad (\text{C2})$$

where the ‘physical coordinates’ [16] are used as the centroids ( $\mathbf{R}_i, \mathbf{P}_i$ ). The case of  $\alpha = 1$  corresponds to the usual wave-packet molecular dynamics (MD) without antisymmetrization. A natural idea in MD is to sample a test particle from each Gaussian distribution  $g_i(\mathbf{r}, \mathbf{p})$ . It should be noted that this MD sampling introduces many-body correlations among test particles, which is different from sampling  $A$  test particles independently following the total distribution  $g(\mathbf{r}, \mathbf{p})$ .

Our aim here is to extend the MD sampling, with reasonable correlations, for the total one-body distribution  $f(\mathbf{r}, \mathbf{p})$ . We may decompose  $f(\mathbf{r}, \mathbf{p})$  into  $A$  terms as

$$f(\mathbf{r}, \mathbf{p}) = \sum_{i=1}^A \hat{f}(\mathbf{r}, \mathbf{p}) g_i(\mathbf{r}, \mathbf{p}) \quad (\text{C3})$$

with  $\hat{f}(\mathbf{r}, \mathbf{p}) = f(\mathbf{r}, \mathbf{p})/g(\mathbf{r}, \mathbf{p})$ . The average number of test particles to be generated for each term should be

$$\bar{N}_i = \int \hat{f}(\mathbf{r}, \mathbf{p}) g_i(\mathbf{r}, \mathbf{p}) \frac{d\mathbf{r}d\mathbf{p}}{(2\pi\hbar)^3} \approx \frac{1}{n} \sum_{j=1}^n \hat{f}(\mathbf{r}_j, \mathbf{p}_j) \quad (\text{C4})$$

which is evaluated by sampling many points ( $\mathbf{r}_j, \mathbf{p}_j$ ) ( $j = 1, 2, \dots, n$ ) from the Gaussian distribution  $g_i(\mathbf{r}, \mathbf{p})$ . With these average numbers  $\bar{N}_i$ , the actual numbers of test particles  $N_i$ , which should be integers, are randomly determined in such a way that  $\sum_{i=1}^A N_i = A$  and  $N_i = \text{floor}(\bar{N}_i)$  or  $\text{floor}(\bar{N}_i) + 1$ .

For each term of Eq. (C3),  $N_i$  test particles should be sampled with the relative weight function  $\hat{f}(\mathbf{r}, \mathbf{p}) g_i(\mathbf{r}, \mathbf{p})$ , which is a straightforward numerical procedure. However, we may introduce additional correlations among test particles by modifying  $\hat{f}$  as

$$\hat{f}(\mathbf{r}, \mathbf{p}) := (1 - e^{-2\nu\beta(\mathbf{r}-\mathbf{r}_k)^2 - \beta(\mathbf{p}-\mathbf{p}_k)^2/2\hbar^2\nu}) \hat{f}(\mathbf{r}, \mathbf{p}) \quad (\text{C5})$$

when a test particle ( $\mathbf{r}_k, \mathbf{p}_k$ ) is generated. Test particles are generated sequentially in a random order, and the modification of  $\hat{f}$  by the  $k$ -th test particle influences only the test particles generated after  $k$ .

In the present work, we have chosen the parameters  $\alpha = \frac{2}{3}$ ,  $\beta = 2$  and  $n = 20$ .

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