

# ***Vorticity and polarization in heavy-ion collisions***

---

**Takafumi Niida**



WAYNE STATE UNIVERSITY

***Hadron Interactions and Polarization from  
Lattice QCD, Quark model, and Heavy-ion collisions  
@YITP, Kyoto***



U.S. DEPARTMENT OF  
**ENERGY**

Office of  
Science

# Important features in non-central heavy-ion collisions

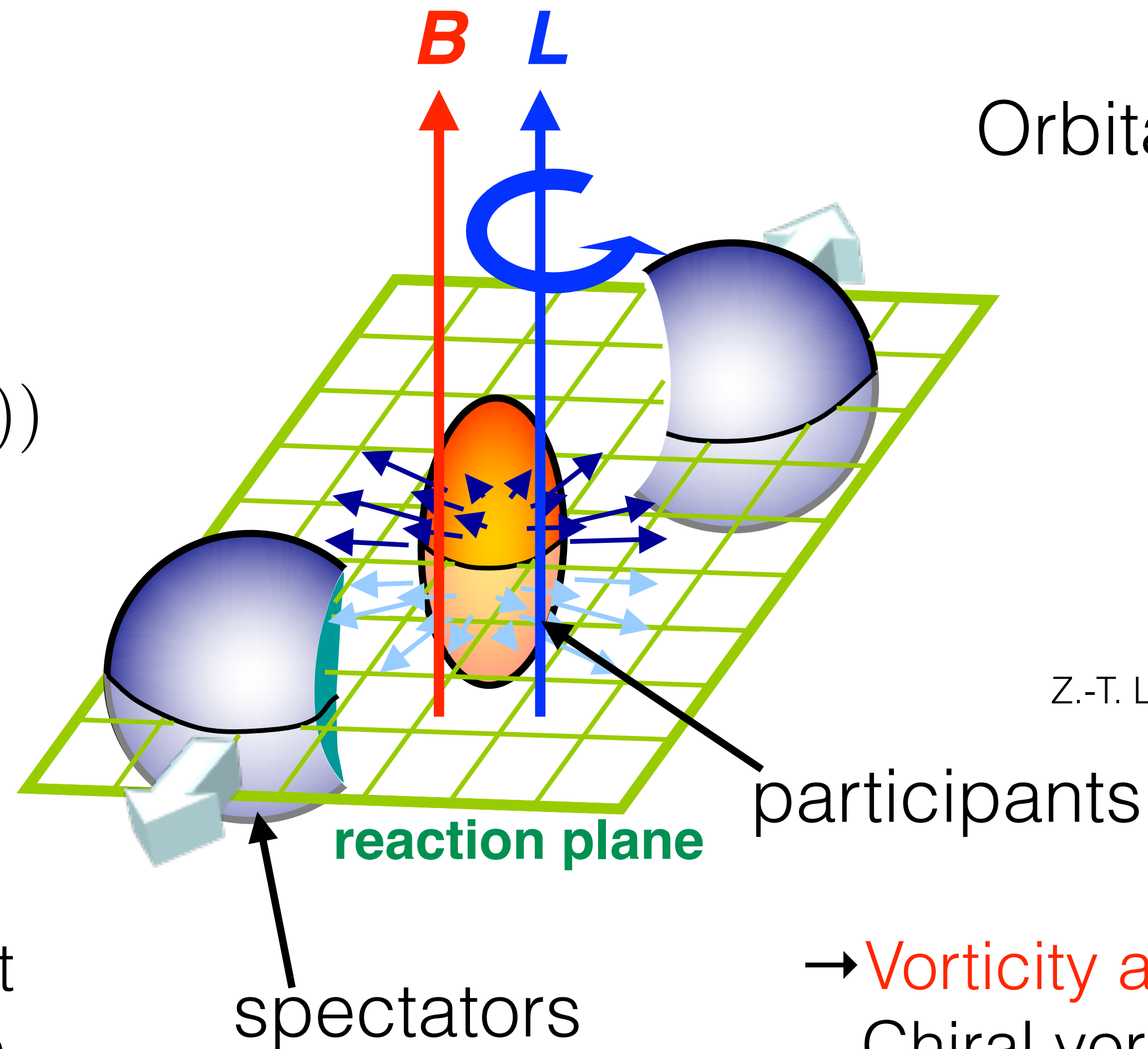
Strong magnetic field

$$B \sim 10^{13} \text{ T}$$

$$(eB \sim \text{MeV}^2 \ (\tau = 0.2 \text{ fm}))$$

Orbital angular momentum

$$L \sim 10^5 \hbar$$

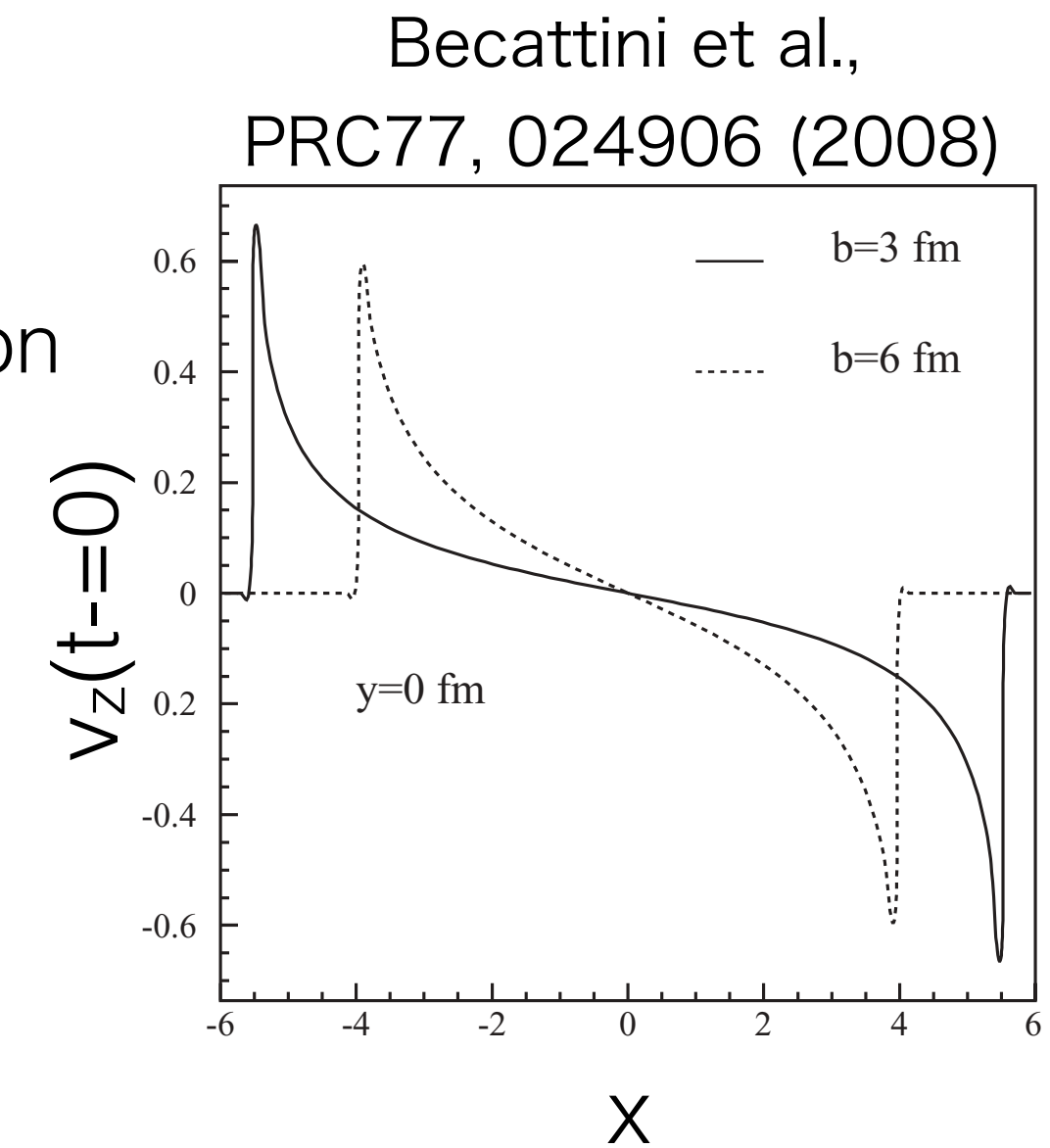
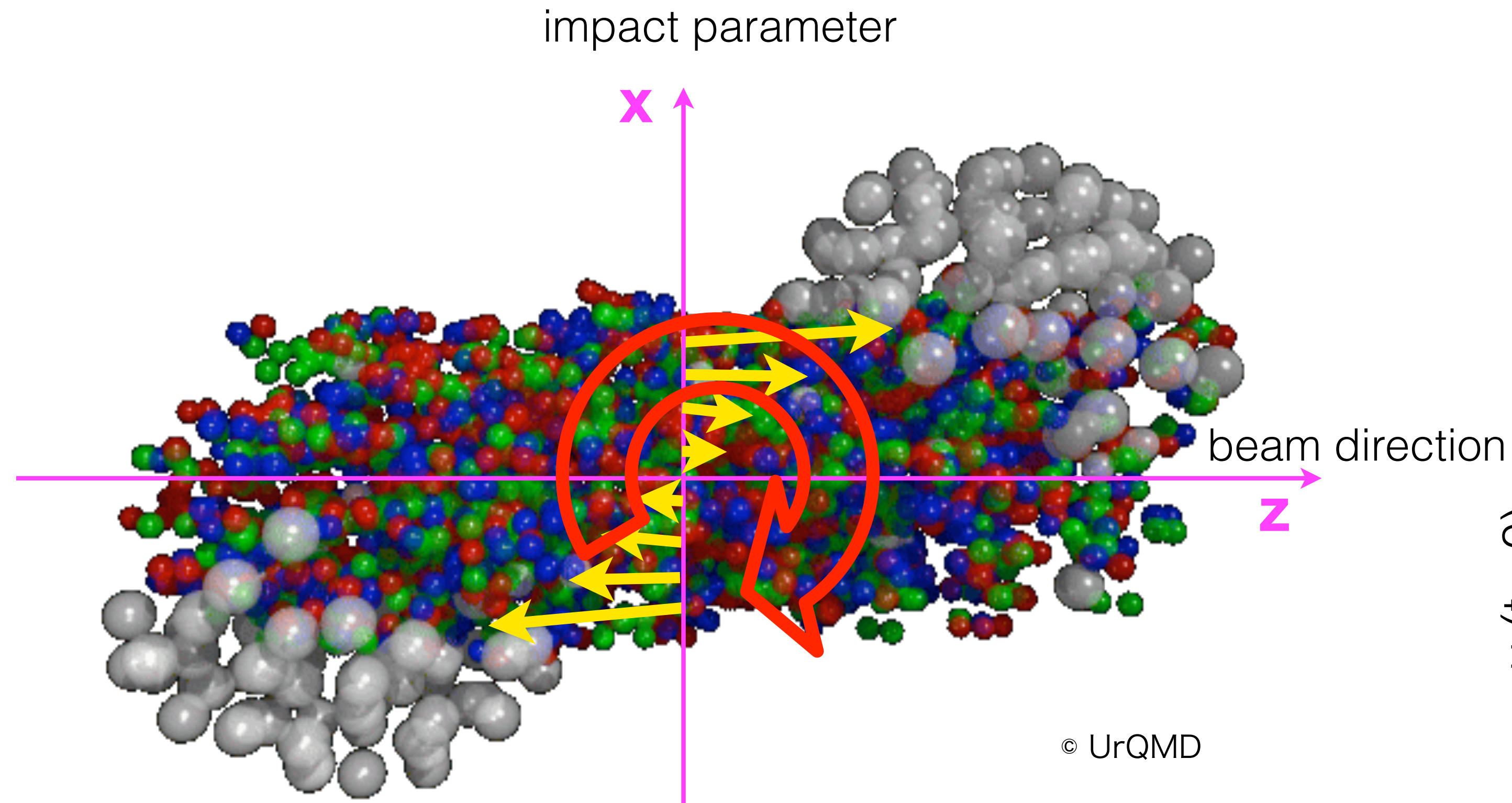


Z.-T. Liang and X.-N. Wang, PRL94, 102301 (2005)

→ Chiral magnetic effect  
Chiral magnetic wave

→ Vorticity and particle polarization  
Chiral vortical effect

# Vorticity in HIC

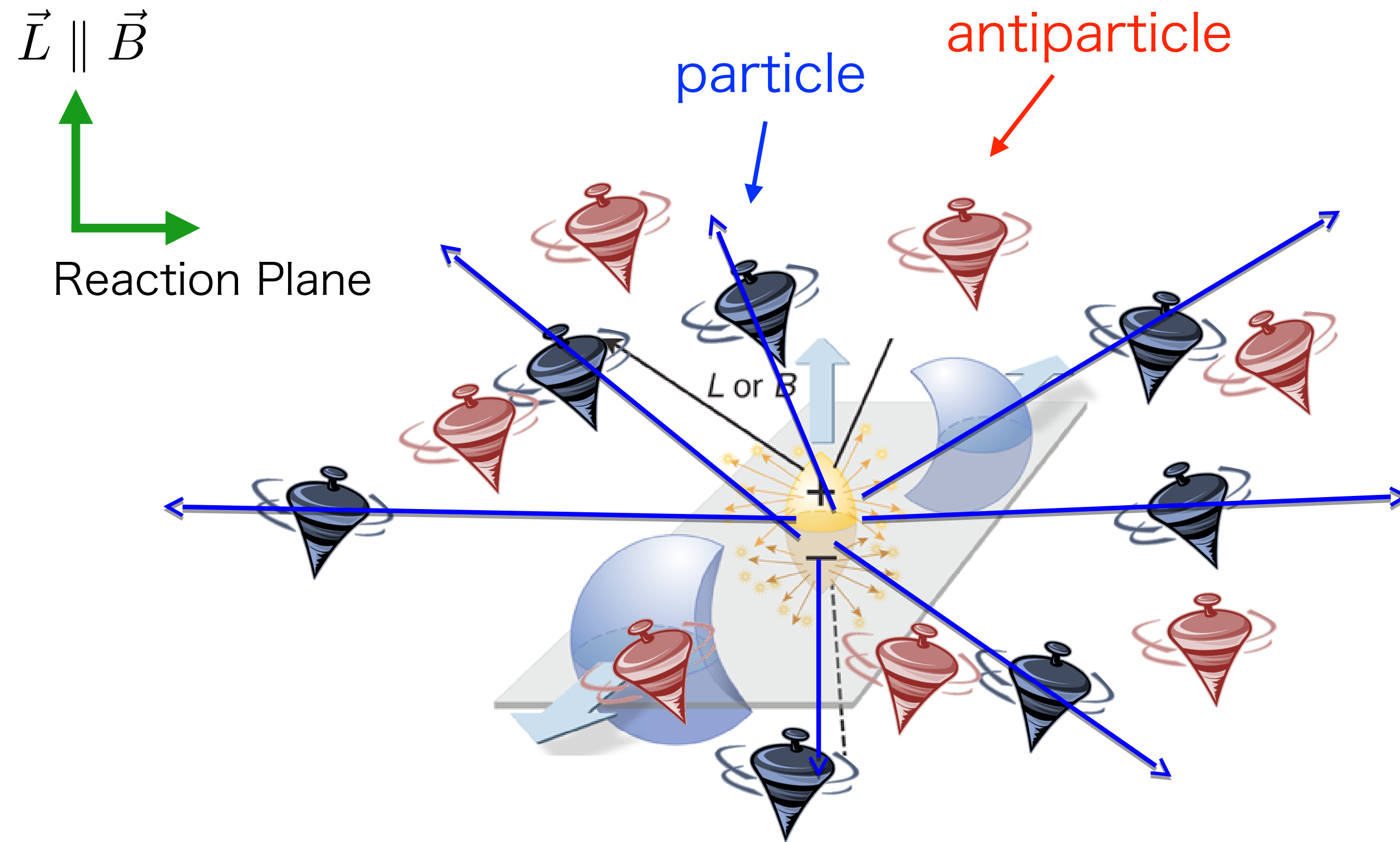


In non-central collisions,  
the initial collective longitudinal flow velocity depends on  $x$ .

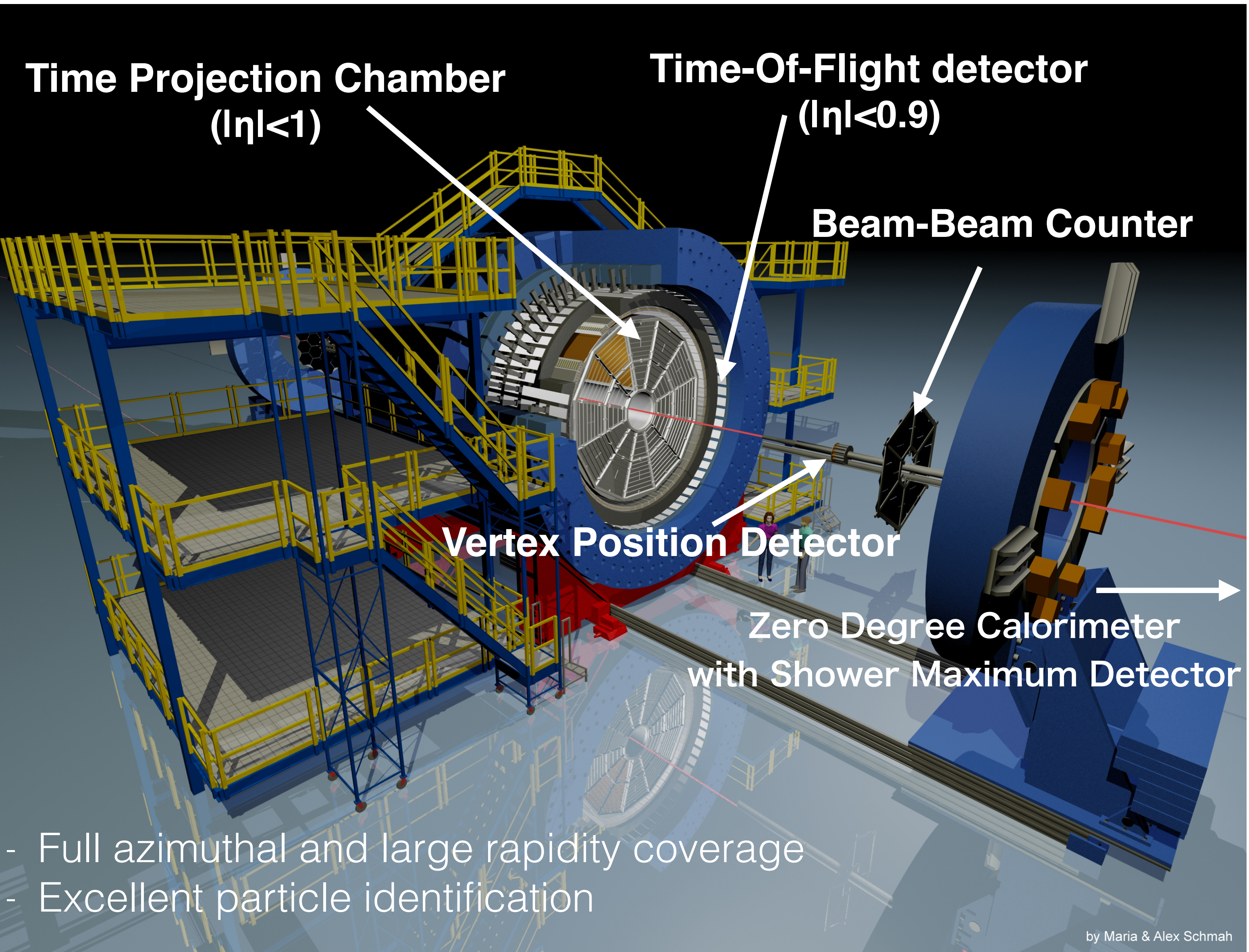
$$\omega_y = \frac{1}{2} (\nabla \times v)_y \approx -\frac{1}{2} \frac{\partial v_z}{\partial x}$$

# Global polarization

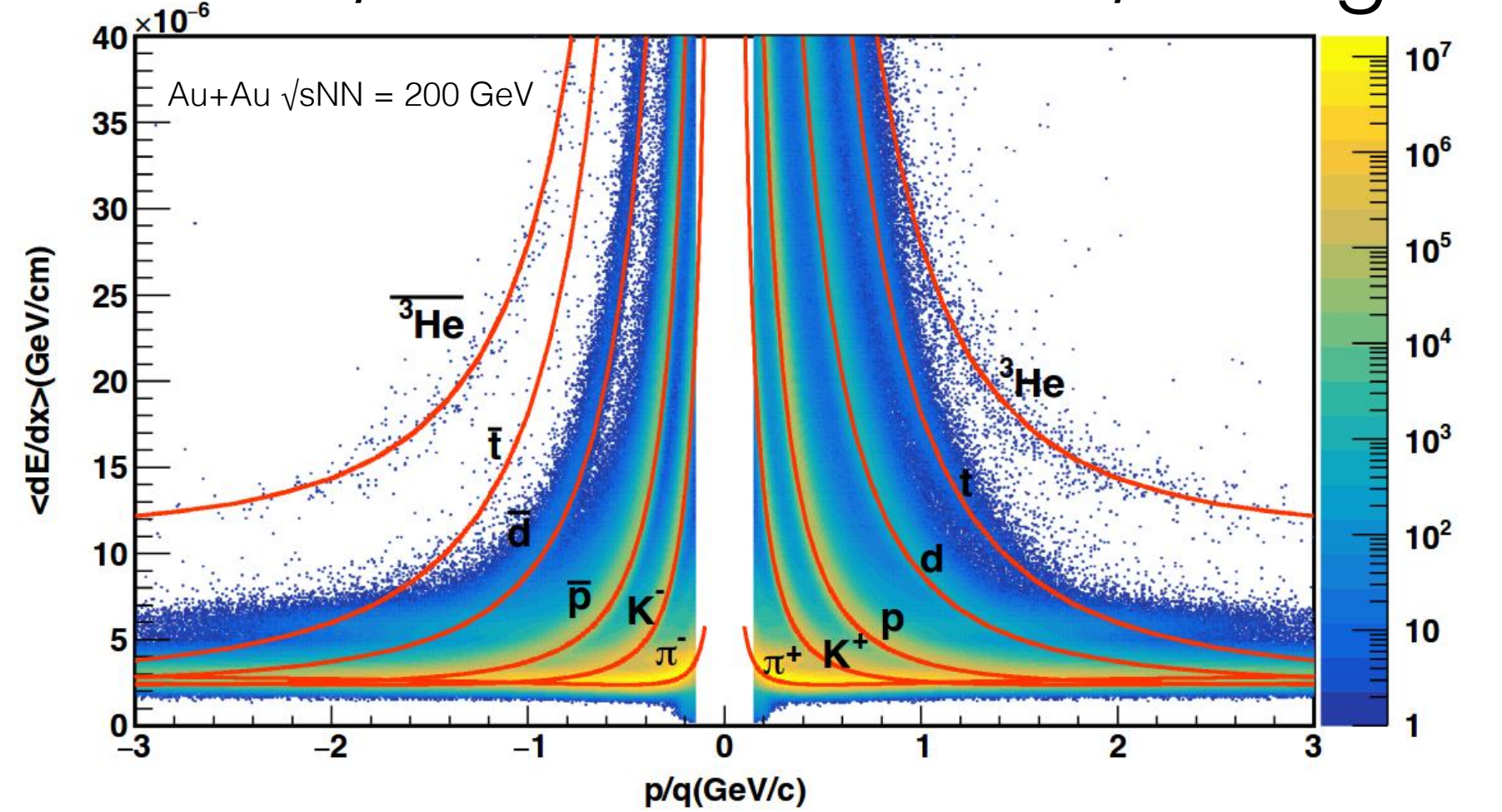
- Z.-T. Liang and X.-N. Wang, PRL94, 102301 (2005)
- S. Voloshin, nucl-th/0410089 (2004)



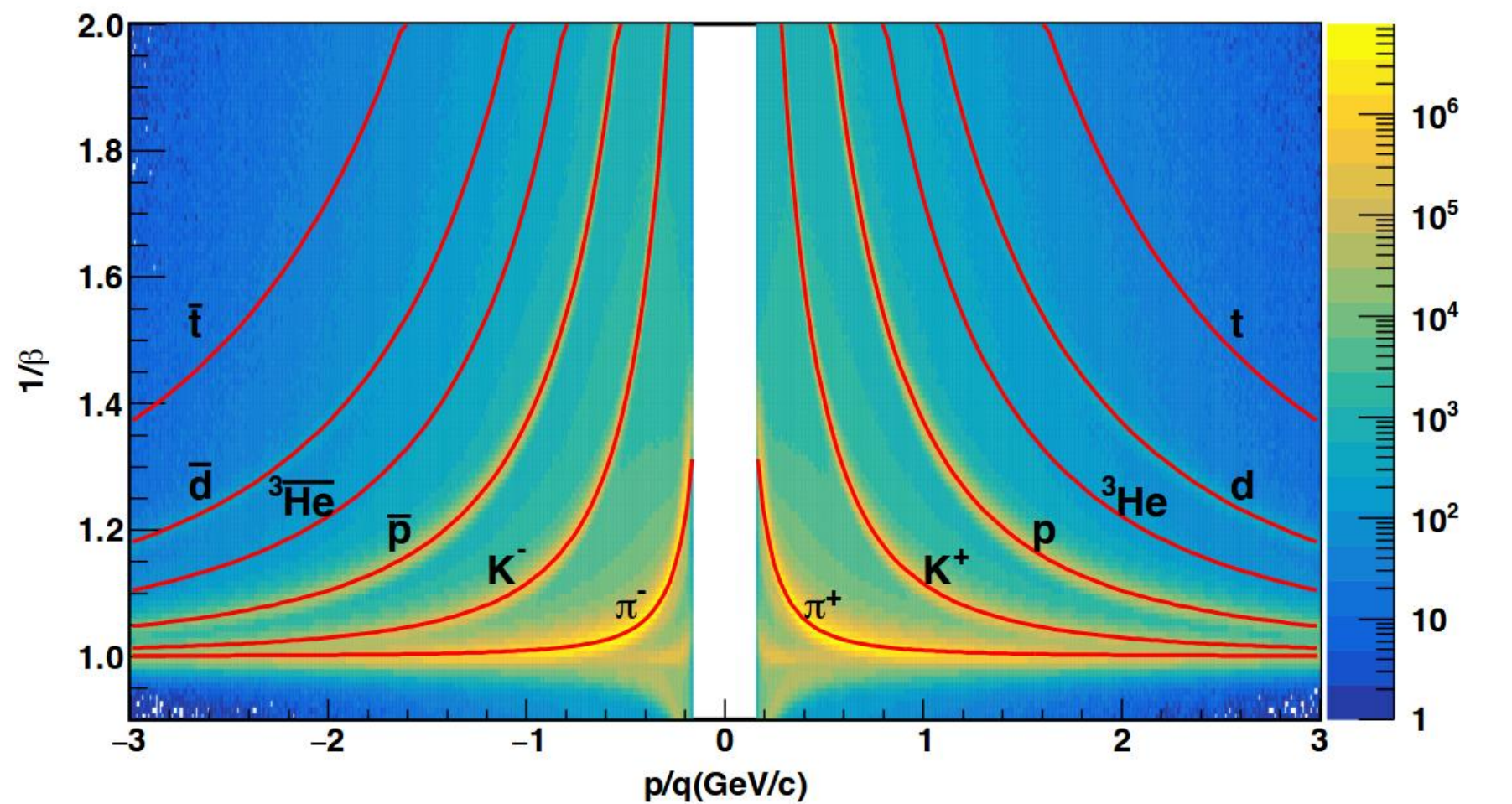
# STAR Detectors



TPC  $dE/dx$  vs momentum/charge



TOF  $1/\beta$  vs momentum/charge



# How to measure the polarization?

## Parity-violating decay of hyperons

Daughter baryon is preferentially emitted in the direction of hyperon's spin (opposite for anti-particle)

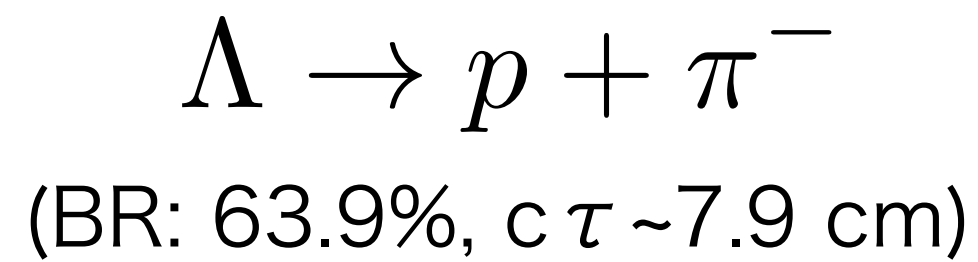
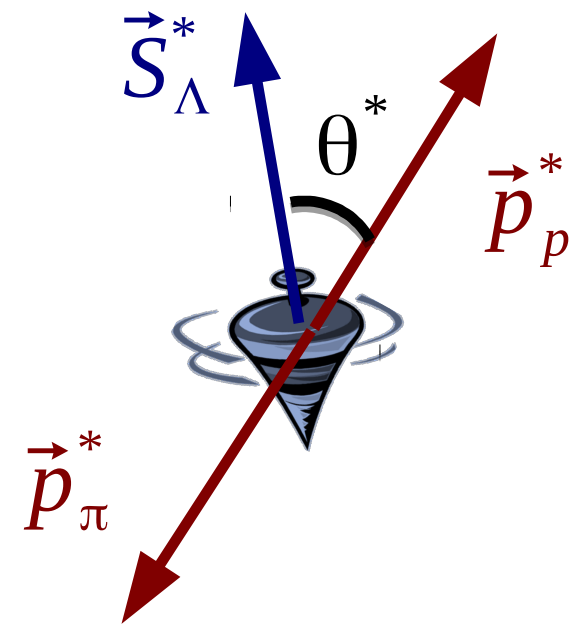
$$\frac{dN}{d\Omega^*} = \frac{1}{4\pi} (1 + \alpha_H \mathbf{P}_H \cdot \mathbf{p}_p^*)$$

$P_H$ :  $\Lambda$  polarization

$p_p^*$ : proton momentum in the  $\Lambda$  rest frame

$\alpha_H$ :  $\Lambda$  decay parameter

$$(\alpha_\Lambda = -\alpha_{\bar{\Lambda}} = 0.642 \pm 0.013)$$

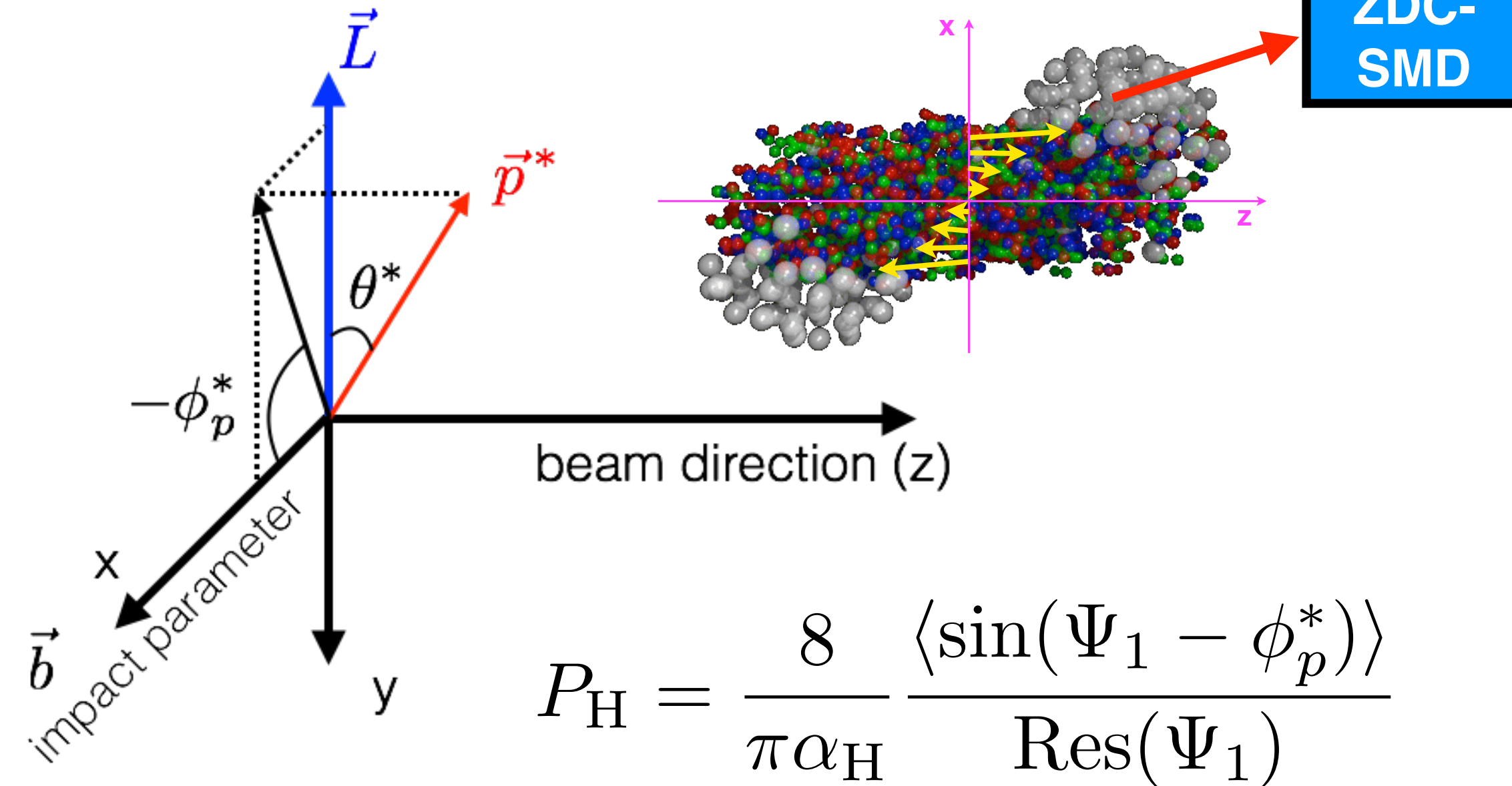


C. Patrignani et al. (PDG), Chin. Phys. C 40, 100001 (2016)

## Projection onto the transverse plane

Angular momentum direction can be determined by spectator deflection (spectators deflect outwards)

- S. Voloshin and TN, PRC94.021901(R)(2016)

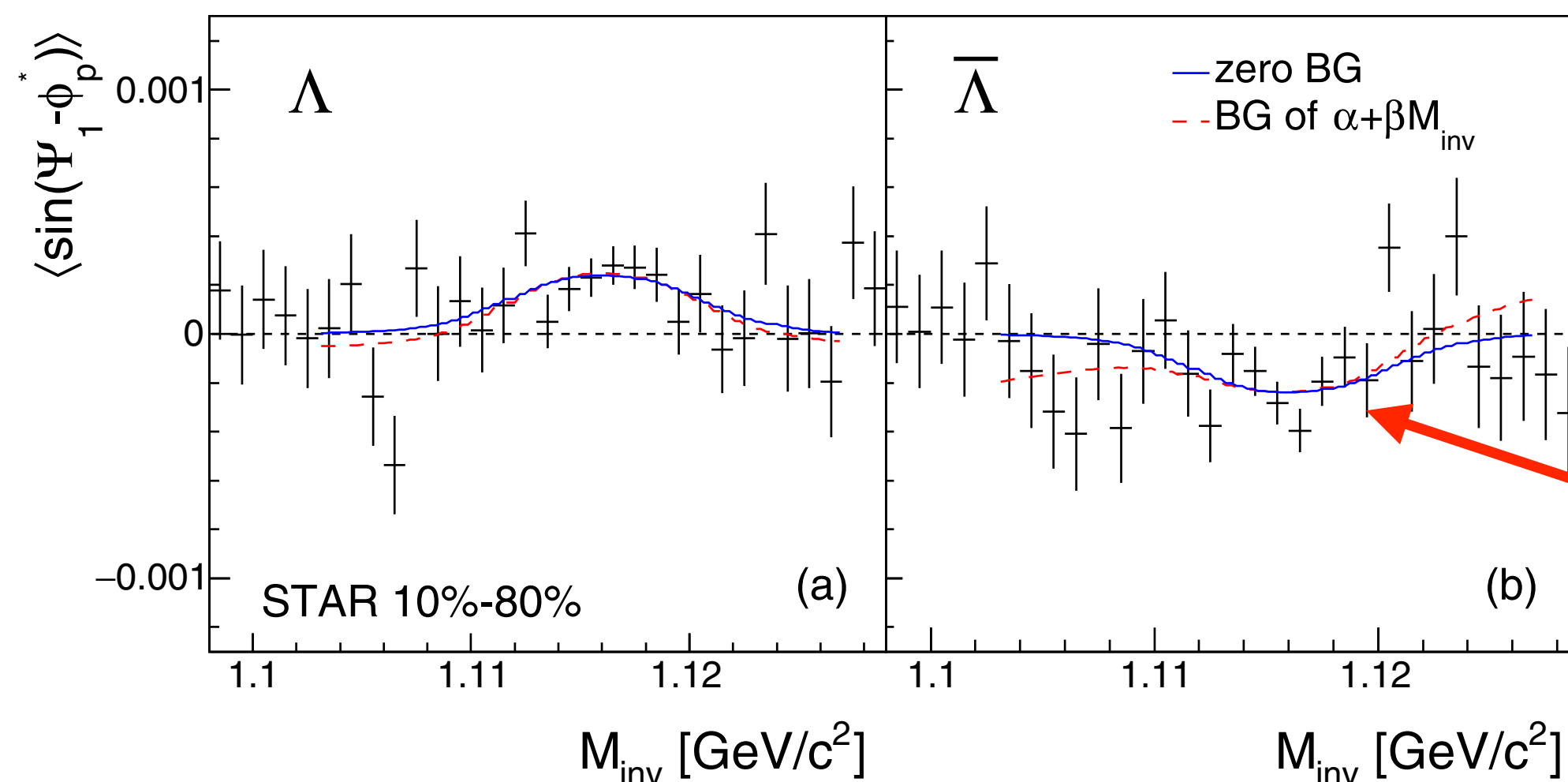
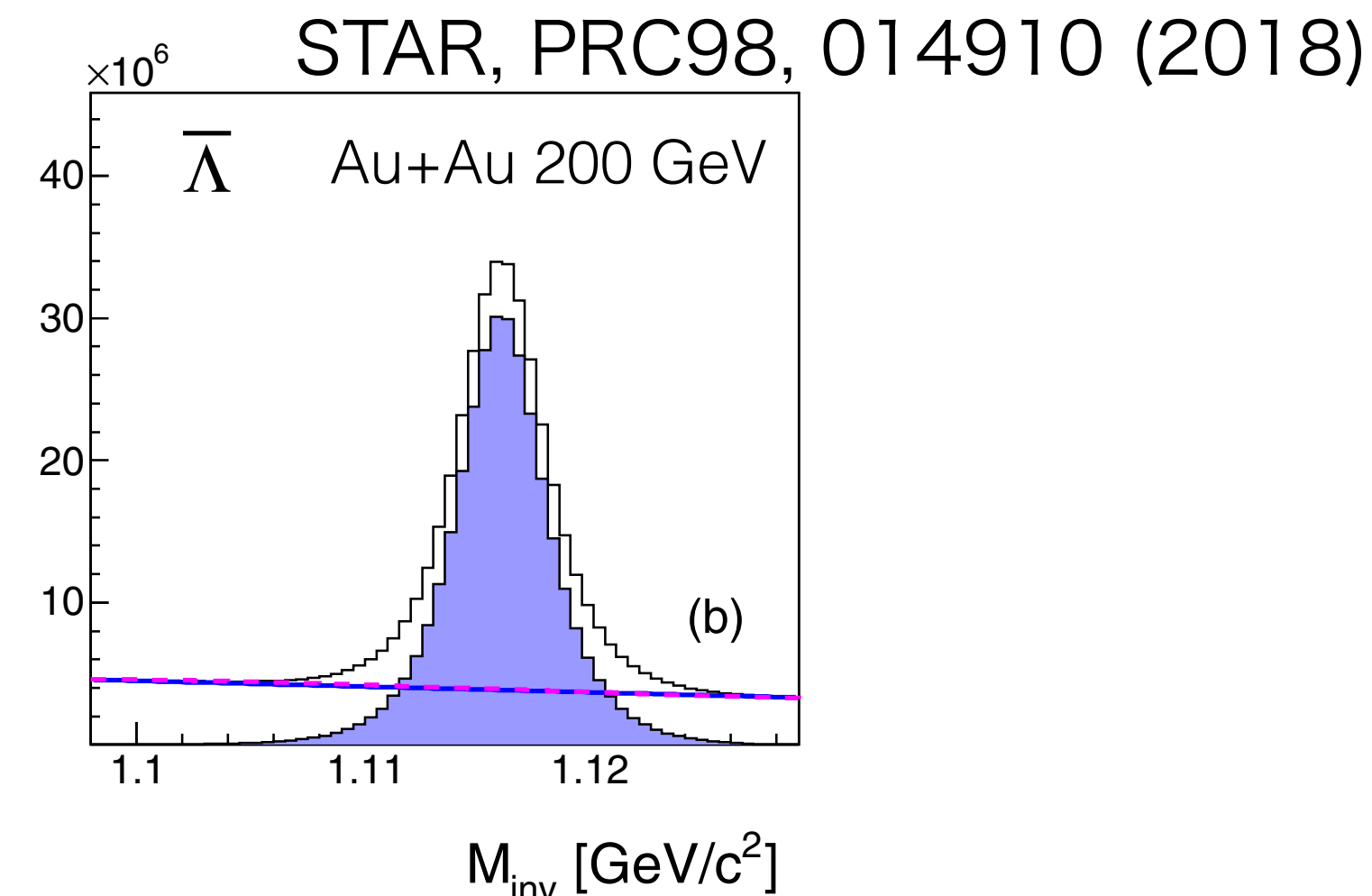
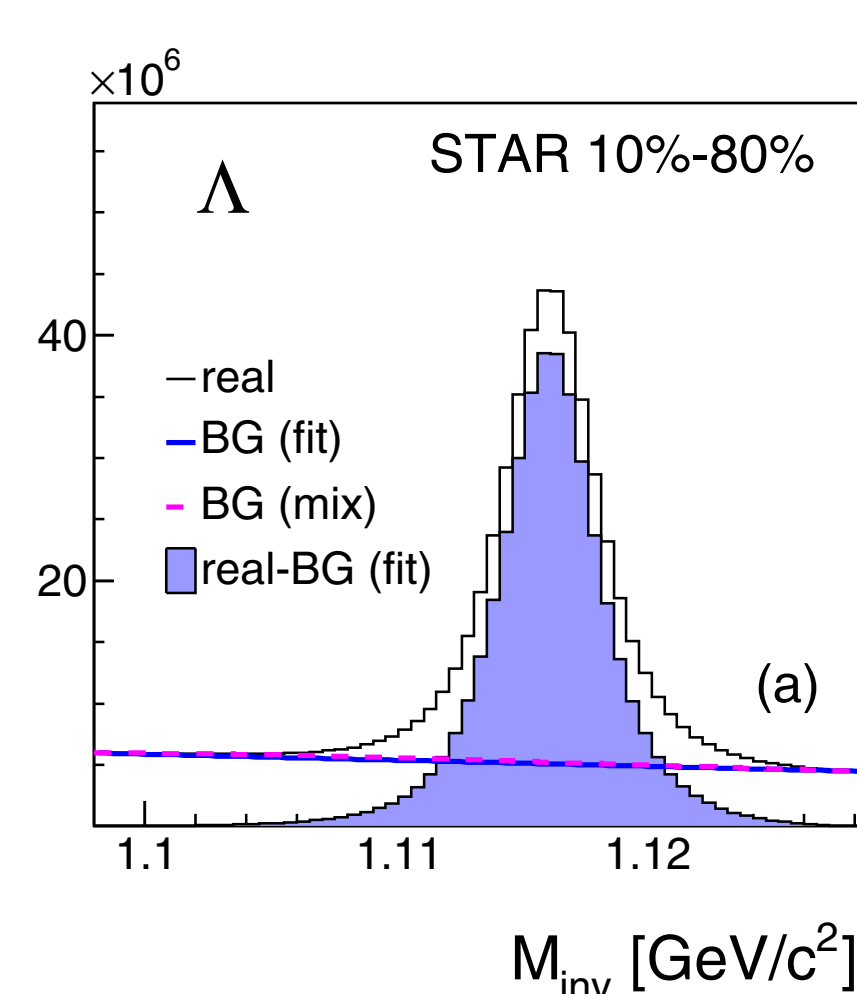
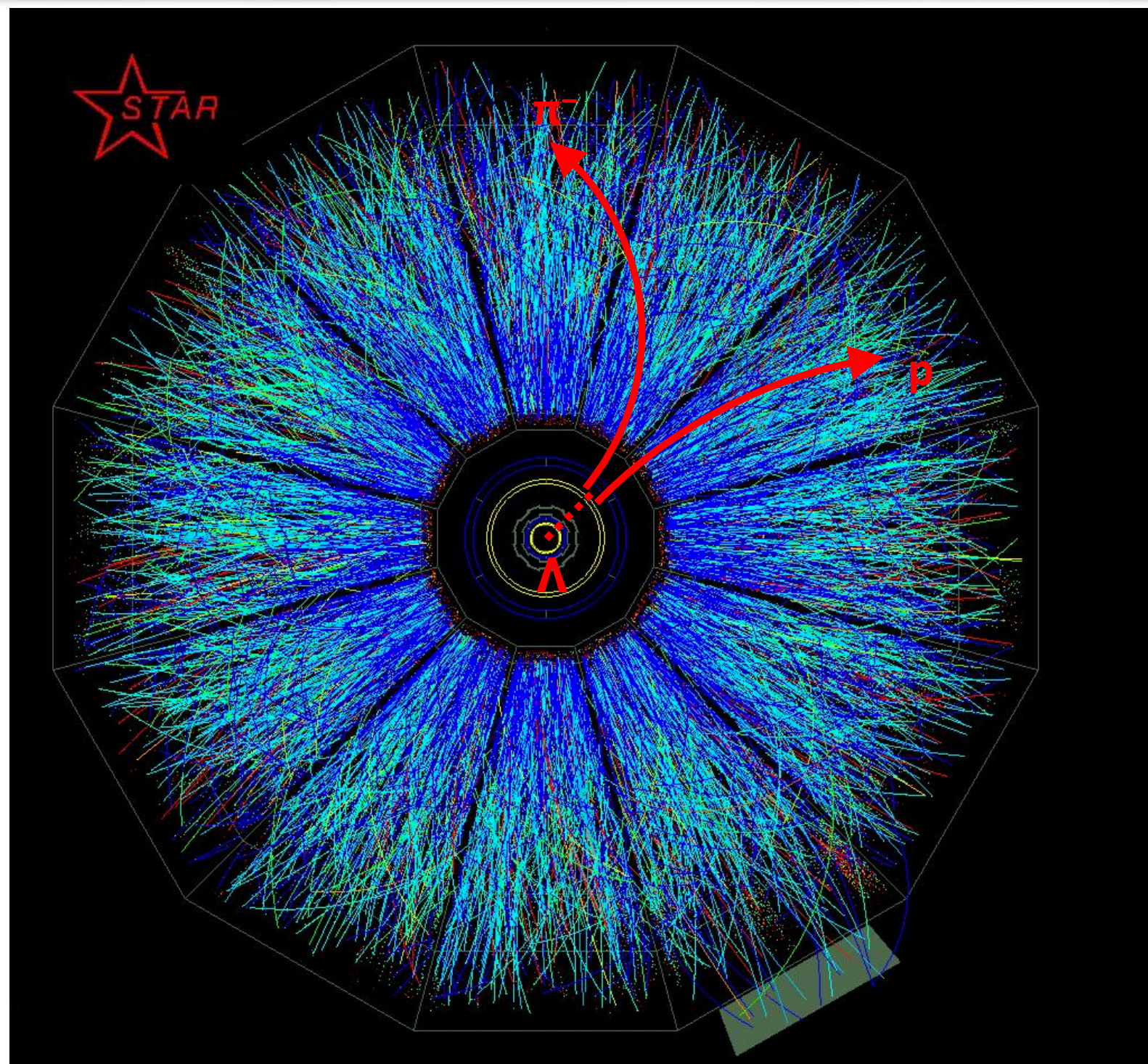


$\Psi_1$ : azimuthal angle of the impact parameter

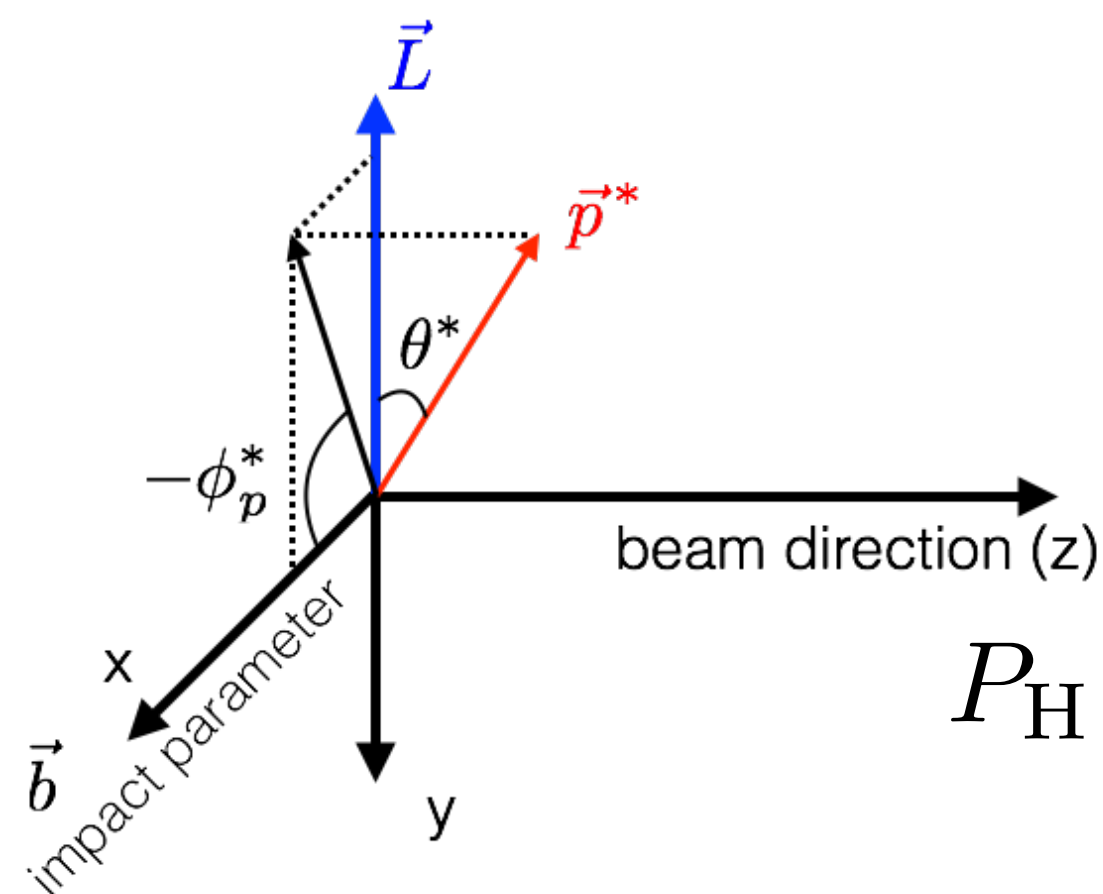
$\phi_p^*$ :  $\phi$  of daughter proton in  $\Lambda$  rest frame

STAR, PRC76, 024915 (2007)

# Signal extraction with $\Lambda$ hyperons



negative for anti- $\Lambda$   
 $\alpha_H = -\alpha_{\bar{H}}$



$$P_H = \frac{8}{\pi \alpha_H} \frac{\langle \sin(\Psi_1 - \phi_p^*) \rangle}{\text{Res}(\Psi_1)}$$

$$\begin{aligned} \langle \sin(\Psi_1 - \phi_p^*) \rangle^{\text{obs}} &= (1 - f^{\text{Bg}}(M_{\text{inv}})) \langle \sin(\Psi_1 - \phi_p^*) \rangle^{\text{Sg}} \\ &+ f^{\text{Bg}}(M_{\text{inv}}) \langle \sin(\Psi_1 - \phi_p^*) \rangle^{\text{Bg}}, \end{aligned}$$

# Feed-down effect

- Only ~25% of measured  $\Lambda$  and anti- $\Lambda$  are primary, while ~60% are feed-down from  $\Sigma^* \rightarrow \Lambda \pi$ ,  $\Sigma^0 \rightarrow \Lambda \gamma$ ,  $\Xi \rightarrow \Lambda \pi$
- Polarization of parent particle R is transferred to its daughter  $\Lambda$

$$\mathbf{S}_\Lambda^* = C \mathbf{S}_R^* \quad \langle S_y \rangle \propto \frac{S(S+1)}{3} \left( \omega + \frac{\mu}{S} B \right)$$

$C_{\Lambda R}$  : coefficient of spin transfer from parent R to  $\Lambda$

$S_R$  : parent particle's spin

$f_{\Lambda R}$  : fraction of  $\Lambda$  originating from parent R

$\mu_R$  : magnetic moment of particle R

Becattini, Karpenko, Lisa, Upsal, and Voloshin, PRC95.054902 (2017)

$$\begin{pmatrix} \varpi_c \\ B_c/T \end{pmatrix} = \begin{bmatrix} \frac{2}{3} \sum_R (f_{\Lambda R} C_{\Lambda R} - \frac{1}{3} f_{\Sigma^0 R} C_{\Sigma^0 R}) S_R(S_R + 1) & \frac{2}{3} \sum_R (f_{\Lambda R} C_{\Lambda R} - \frac{1}{3} f_{\Sigma^0 R} C_{\Sigma^0 R}) (S_R + 1) \mu_R \\ \frac{2}{3} \sum_{\bar{R}} (f_{\Lambda \bar{R}} C_{\Lambda \bar{R}} - \frac{1}{3} f_{\Sigma^0 \bar{R}} C_{\Sigma^0 \bar{R}}) S_{\bar{R}}(S_{\bar{R}} + 1) & \frac{2}{3} \sum_{\bar{R}} (f_{\Lambda \bar{R}} C_{\Lambda \bar{R}} - \frac{1}{3} f_{\Sigma^0 \bar{R}} C_{\Sigma^0 \bar{R}}) (S_{\bar{R}} + 1) \mu_{\bar{R}} \end{bmatrix}^{-1} \begin{pmatrix} P_\Lambda^{\text{meas}} \\ P_{\bar{\Lambda}}^{\text{meas}} \end{pmatrix}$$

Decay	C
Parity conserving: $1/2^+ \rightarrow 1/2^+ 0^-$	-1/3
Parity conserving: $1/2^- \rightarrow 1/2^+ 0^-$	1
Parity conserving: $3/2^+ \rightarrow 1/2^+ 0^-$	1/3
Parity-conserving: $3/2^- \rightarrow 1/2^+ 0^-$	-1/5
$\Xi^0 \rightarrow \Lambda + \pi^0$	+0.900
$\Xi^- \rightarrow \Lambda + \pi^-$	+0.927
$\Sigma^0 \rightarrow \Lambda + \gamma$	-1/3

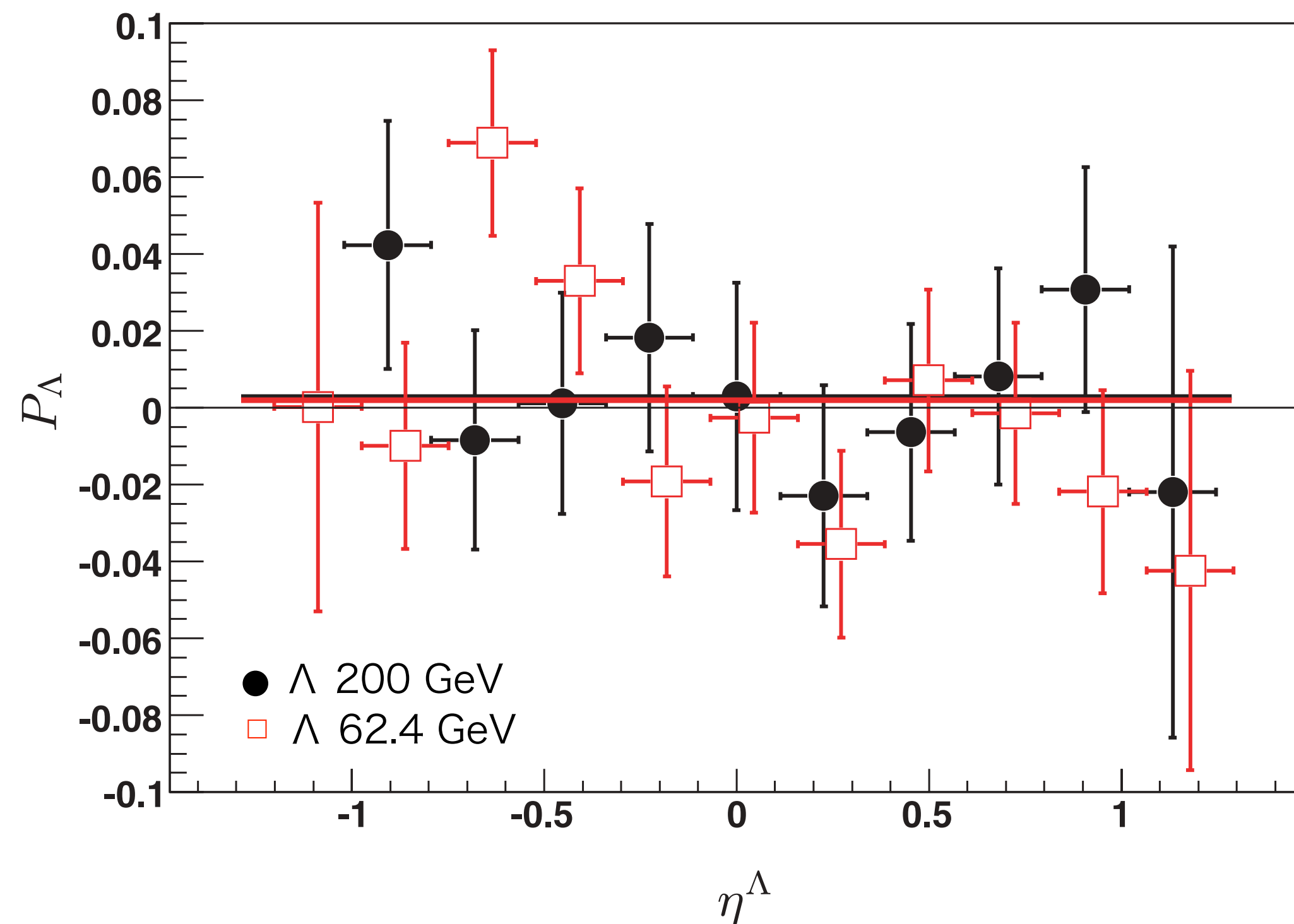
15%-20% dilution of primary  $\Lambda$  polarization  
(model-dependent)



# First paper on $\Lambda$ polarization from STAR

PHYSICAL REVIEW C 76, 024915 (2007)

## Global polarization measurement in Au+Au collisions



Au+Au collisions at  $\sqrt{s_{NN}} = 62.4$  and 200 GeV in 2004 with very limited statistics ( $\sim 9$ M events)

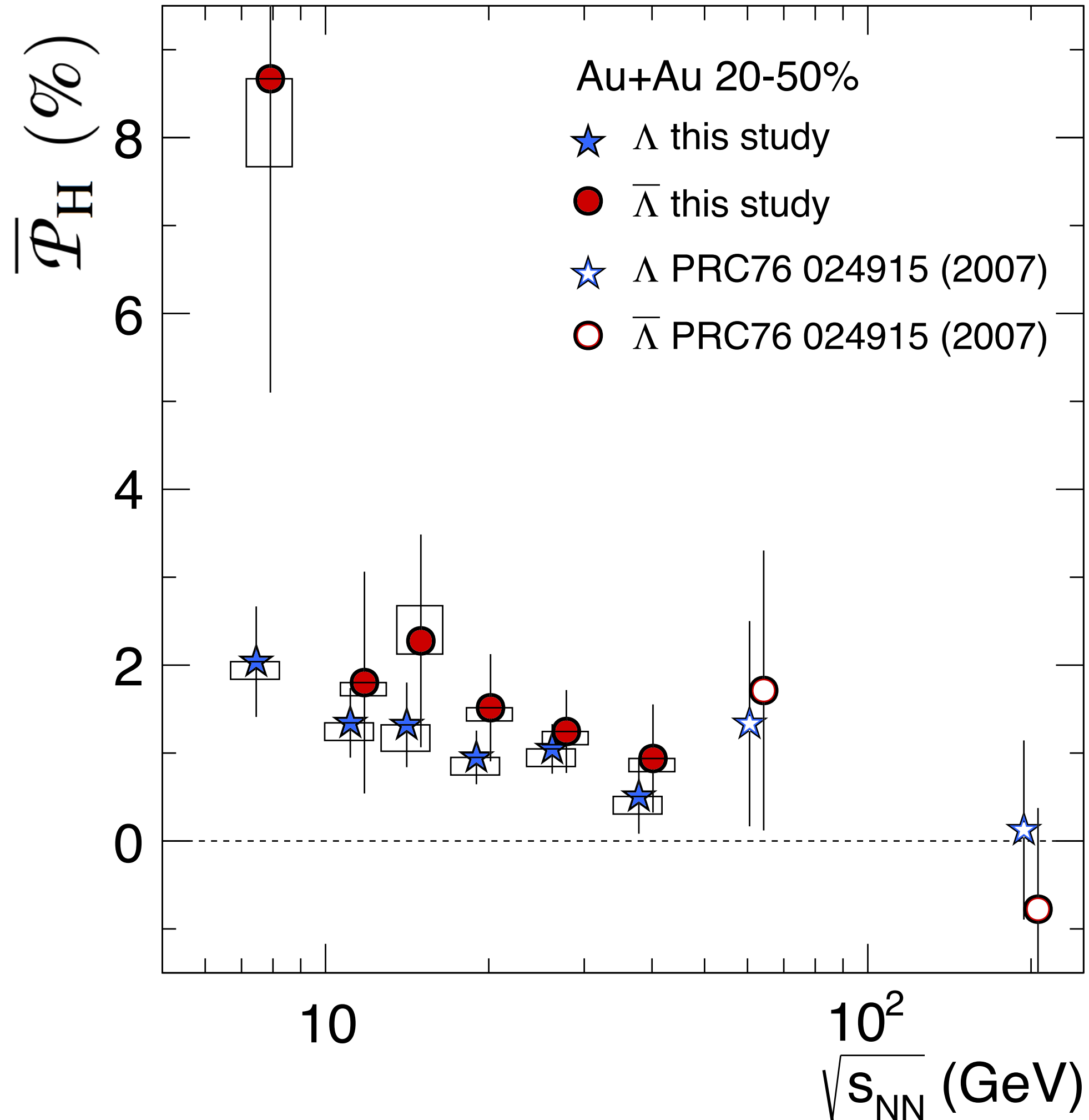
Results are consistent with zero...  
giving an upper limit of  $P_H < 2\%$

### III. CONCLUSION

The  $\Lambda$  and  $\bar{\Lambda}$  hyperon global polarization has been measured in Au+Au collisions at center-of-mass energies  $\sqrt{s_{NN}} = 62.4$  and 200 GeV with the STAR detector at RHIC. An upper limit of  $|P_{\Lambda, \bar{\Lambda}}| \leq 0.02$  for the global polarization of  $\Lambda$  and  $\bar{\Lambda}$  hyperons within the STAR detector acceptance is

# First observation of fluid vortices in HIC

STAR, Nature 548, 62 (2017)



#38



**The Fastest Fluid**

by Sylvia Morrow

Superhot material spins at an incredible rate.

- Positive polarization signal at lower energies!
- polarization looks to increase in lower energies
- anti- $\Lambda$  is systematically larger than  $\Lambda$

Becattini, Karpenko, Lisa, Uppal, and Voloshin, PRC95.054902 (2017)

$\mu_\Lambda$ :  $\Lambda$  magnetic moment  
T: temperature at thermal equilibrium

$$P_\Lambda \simeq \frac{1}{2} \frac{\omega}{T} + \frac{\mu_\Lambda B}{T}$$

$$P_{\bar{\Lambda}} \simeq \frac{1}{2} \frac{\omega}{T} - \frac{\mu_\Lambda B}{T}$$

$$\omega = (P_\Lambda + P_{\bar{\Lambda}}) k_B T / \hbar$$

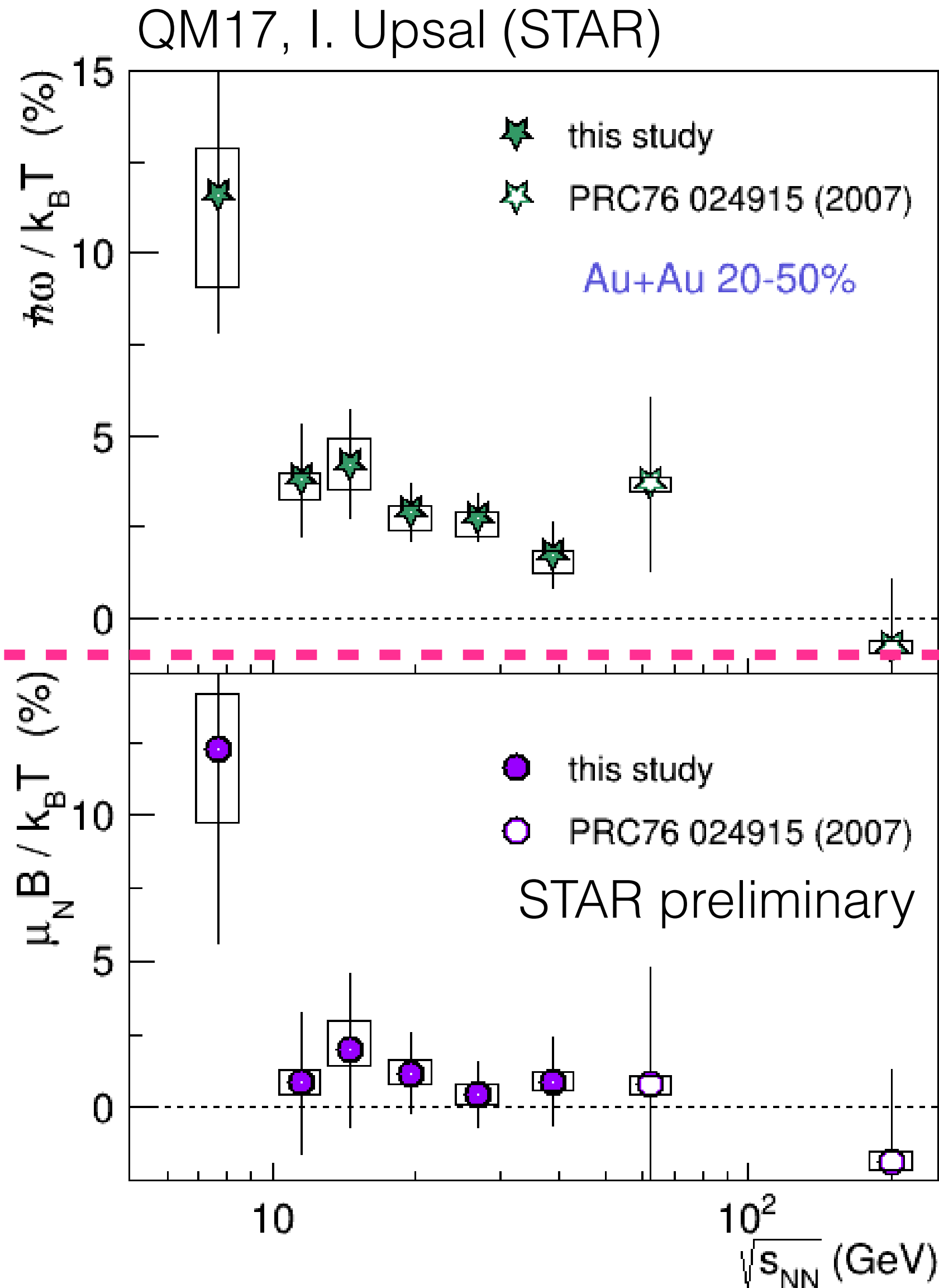
$$\sim 0.02-0.09 \text{ fm}^{-1}$$

$$\sim 0.6-2.7 \times 10^{22} \text{ s}^{-1}$$

(T=160 MeV)

The most vortical fluid ever observed!

# Possible probe of magnetic field



Becattini, Karpenko, Lisa, Upsal, and Voloshin, PRC95.054902 (2017)

$$P_\Lambda \simeq \frac{1}{2} \frac{\omega}{T} + \frac{\mu_\Lambda B}{T}$$

$$P_{\bar{\Lambda}} \simeq \frac{1}{2} \frac{\omega}{T} - \frac{\mu_\Lambda B}{T}$$

$\mu_\Lambda$ :  $\Lambda$  magnetic moment

$$B = (P_\Lambda - P_{\bar{\Lambda}}) k_B T / \mu_N$$

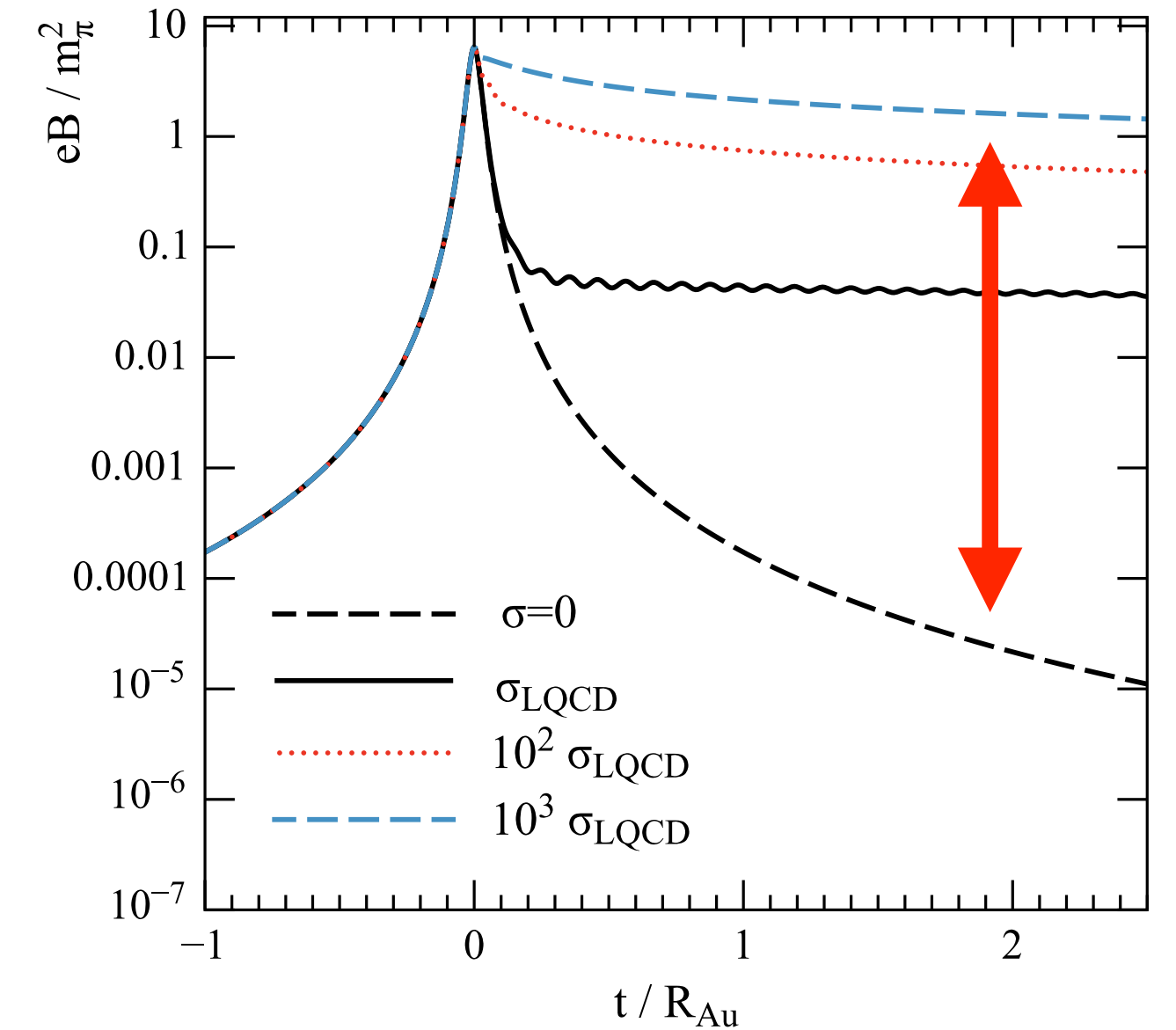
$$\sim 5.0 \times 10^{13} \text{ [Tesla]}$$

nuclear magneton  $\mu_N = -0.613\mu_\Lambda$

Extracted B-field at freeze-out assuming local thermal equilibrium, although it's consistent with zero.

Need more data!  $\rightarrow$  BES-II and Isobaric collisions

McLerran and Skokov, Nucl. Phys. A929, 184 (2014)

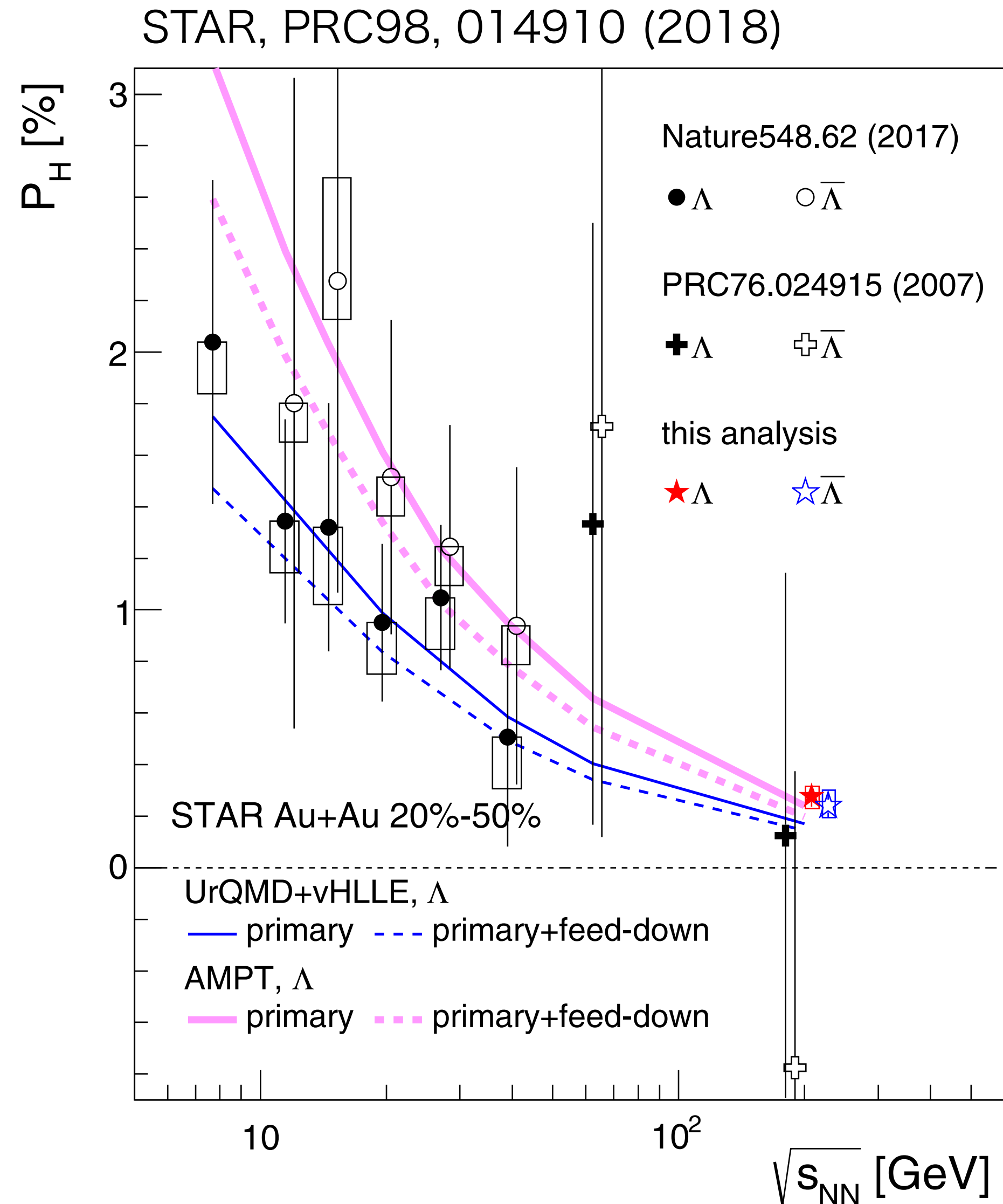


conductivity increases B-lifetime

$$B \sim 10^{13} \text{ T}$$

$$(eB \sim \text{MeV}^2 \text{ } (\tau = 0.2 \text{ fm}))$$

# Positive signal at $\sqrt{s_{NN}} = 200 \text{ GeV}$



$$P_H(\Lambda) [\%] = 0.277 \pm 0.040(\text{stat}) \pm_{0.049}^{0.039} (\text{sys})$$

$$P_H(\bar{\Lambda}) [\%] = 0.240 \pm 0.045(\text{stat}) \pm_{0.045}^{0.061} (\text{sys})$$

- 5-7 $\sigma$  significance, comparable to the combined result of 7.7-39 GeV

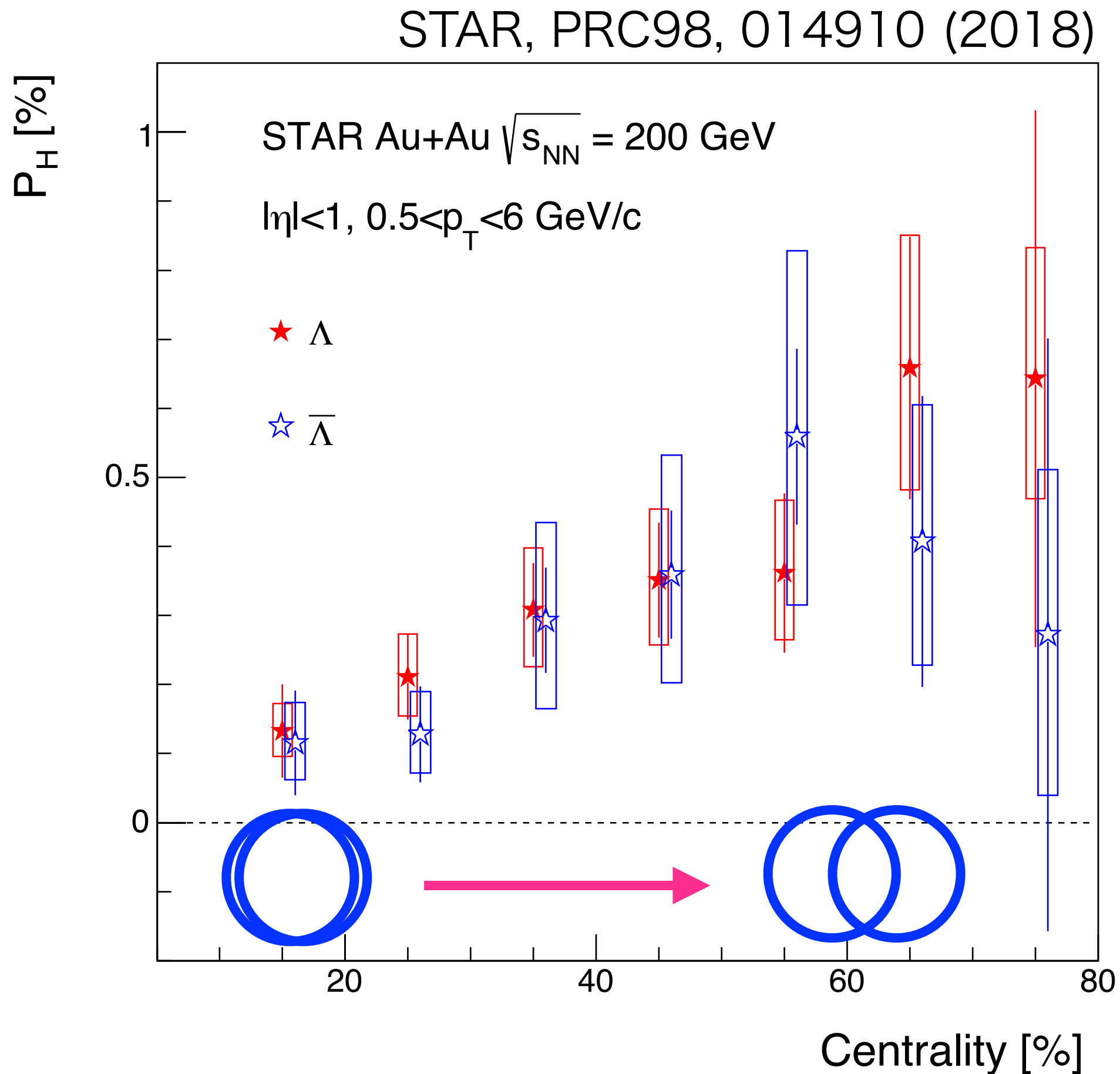
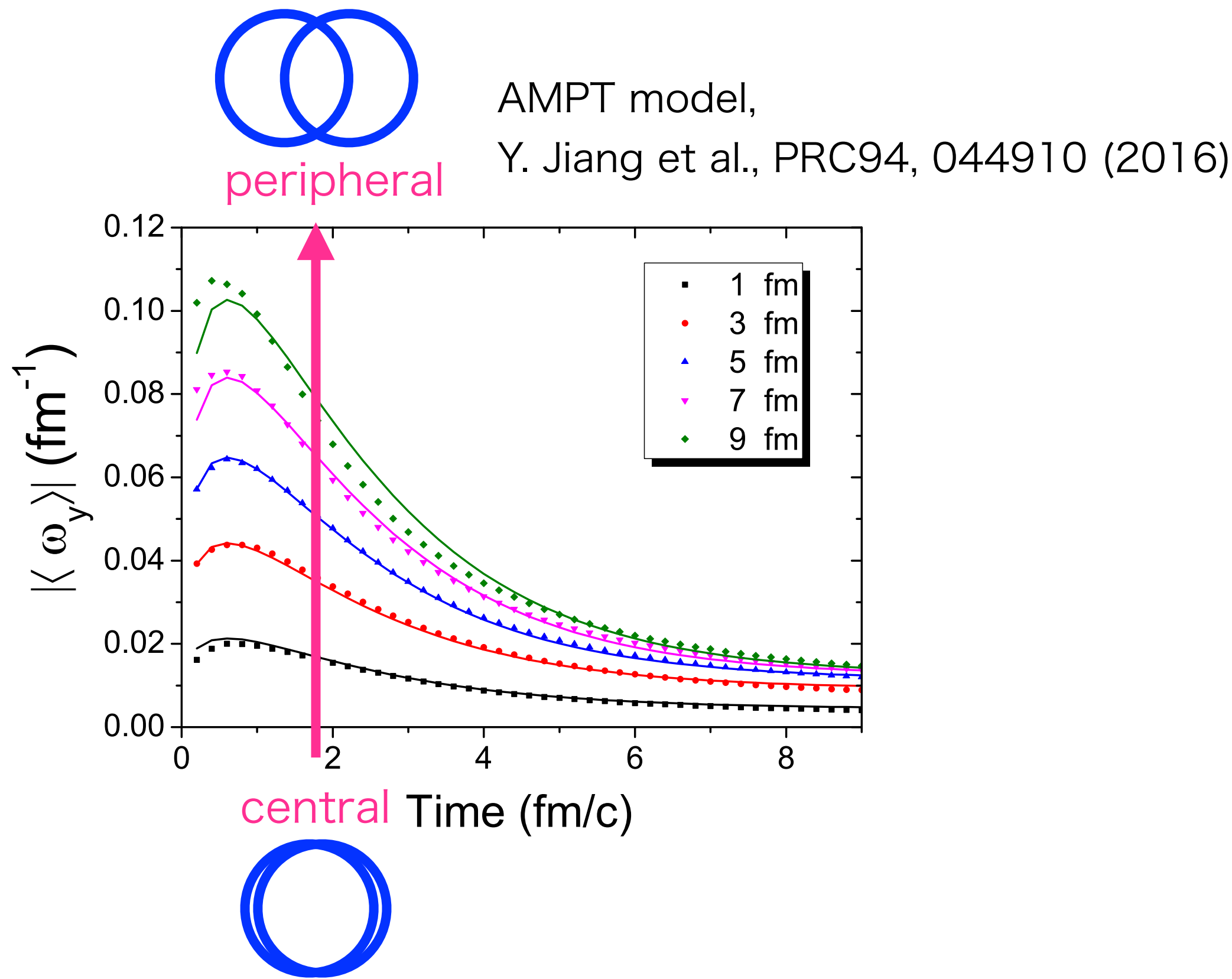
- Feed-down ~15%-20% reduction of  $P_H$  (model-dependent)

Becattini, Karpenko, Lisa, Upsal, and Voloshin, PRC95.054902 (2017)

UrQMD+vHLLE: I. Karpenko and F. Becattini, EPJC(2017)77:213

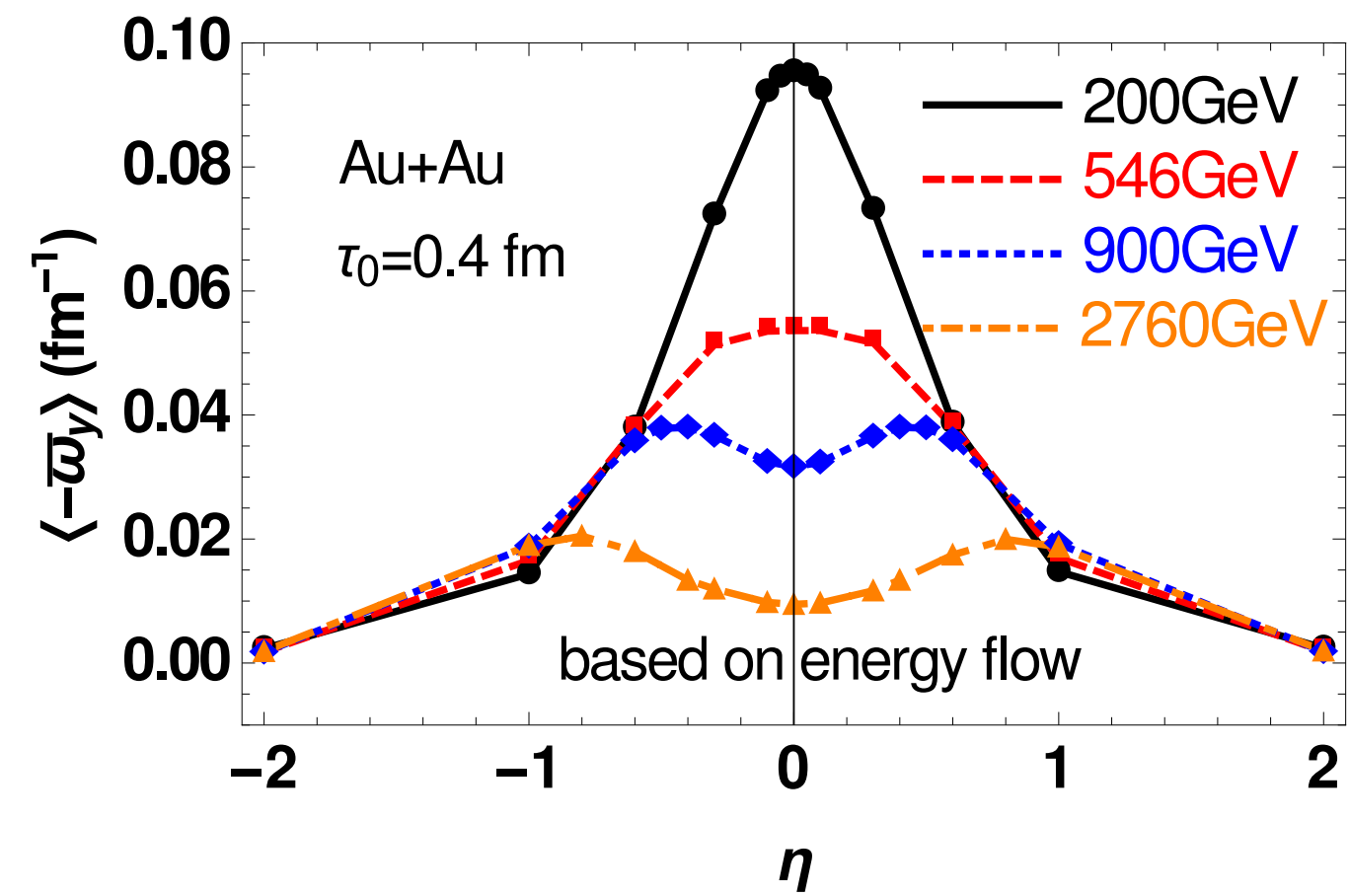
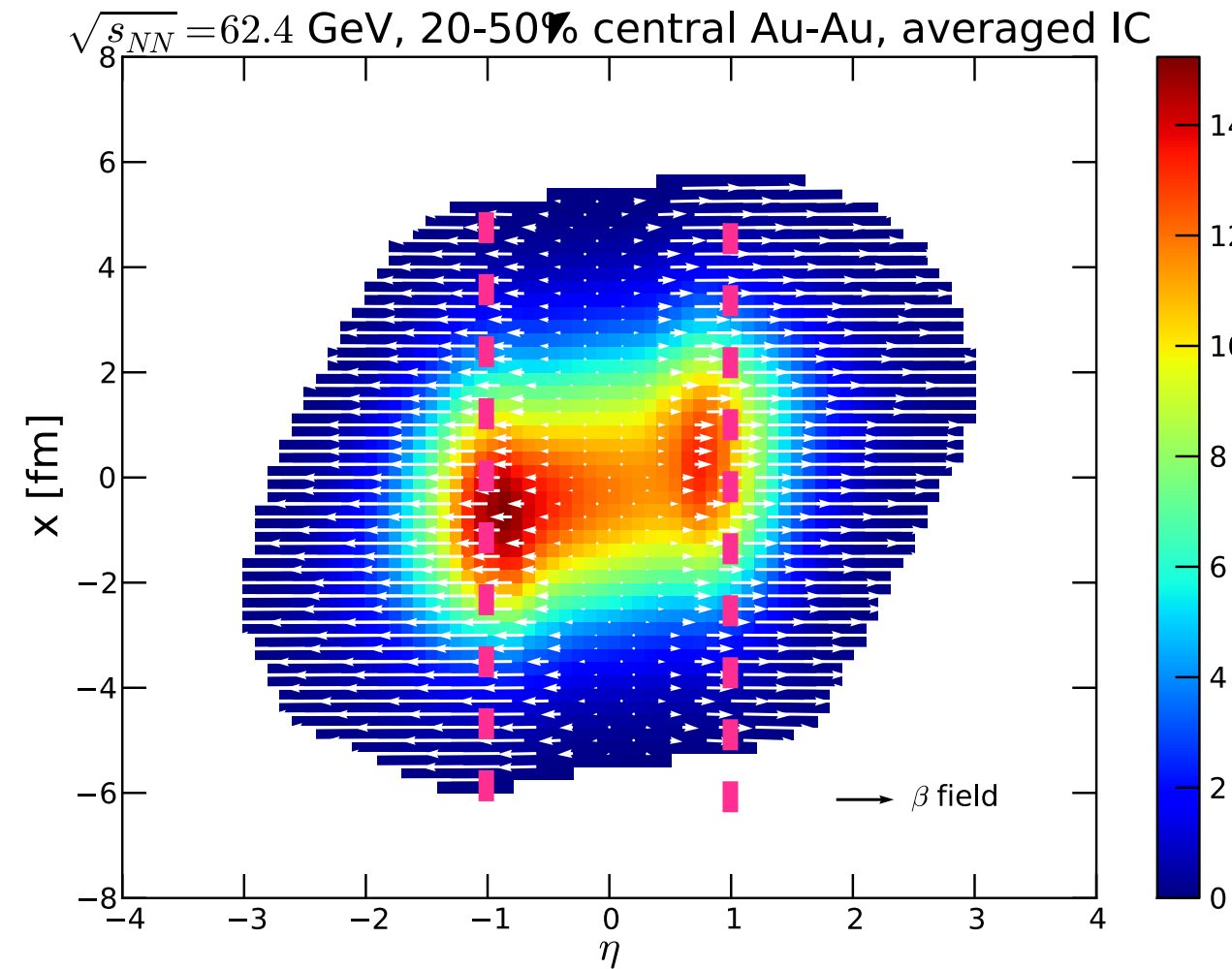
AMPT: H. Li et al., Phys. Rev. C 96, 054908 (2017)

# Centrality dependence of $P_H$



In most central collision  $\rightarrow$  no initial angular momentum  
 As expected, the polarization decreases in more central collisions

# $\eta$ dependence of $P_H$

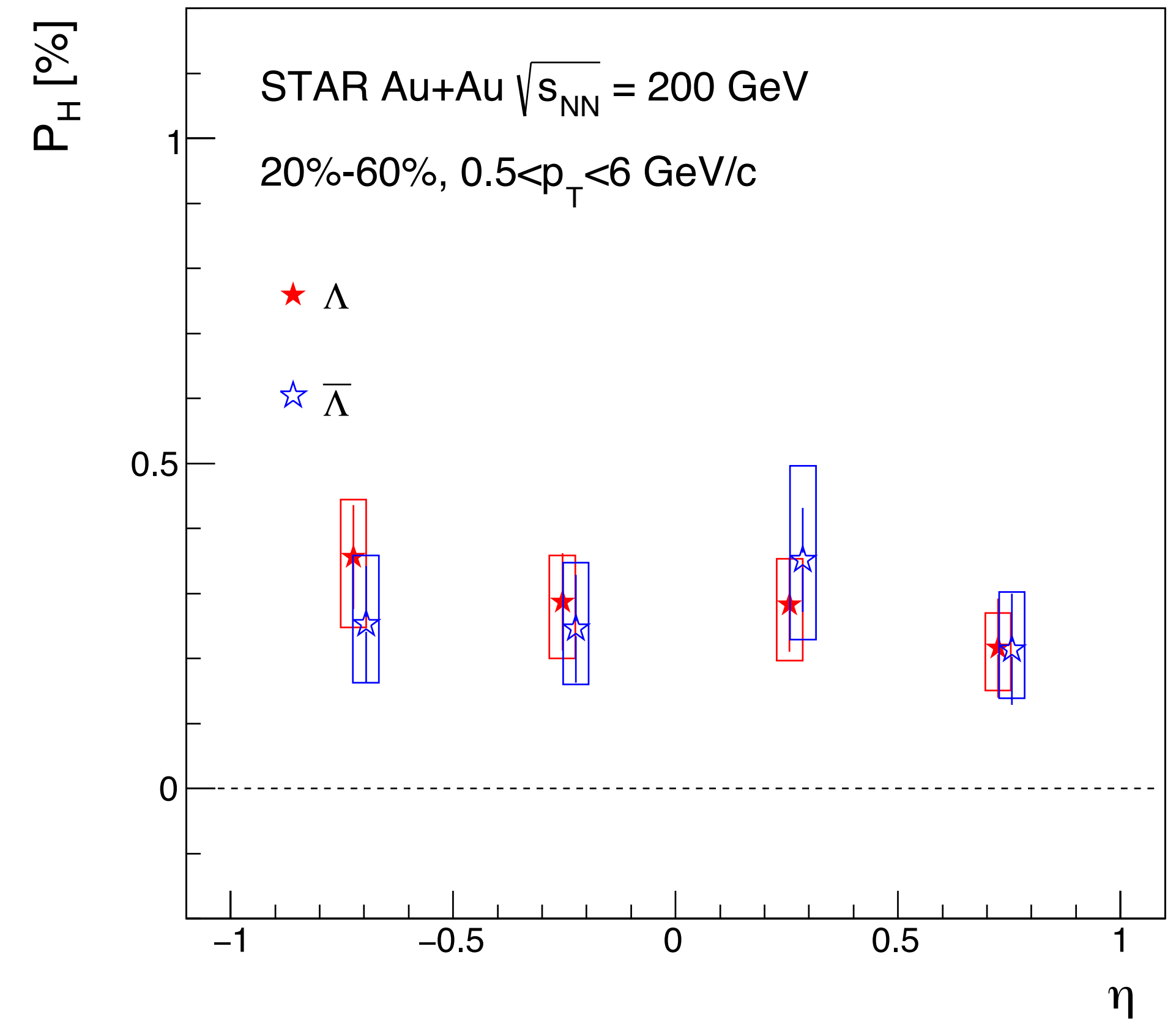


- Shear flow structure/initial flow velocity would be stronger in forward/backward region
- Expect rapidity dependence of the polarization

I. Karpenko and F. Becattini, EPJC(2017)77:213

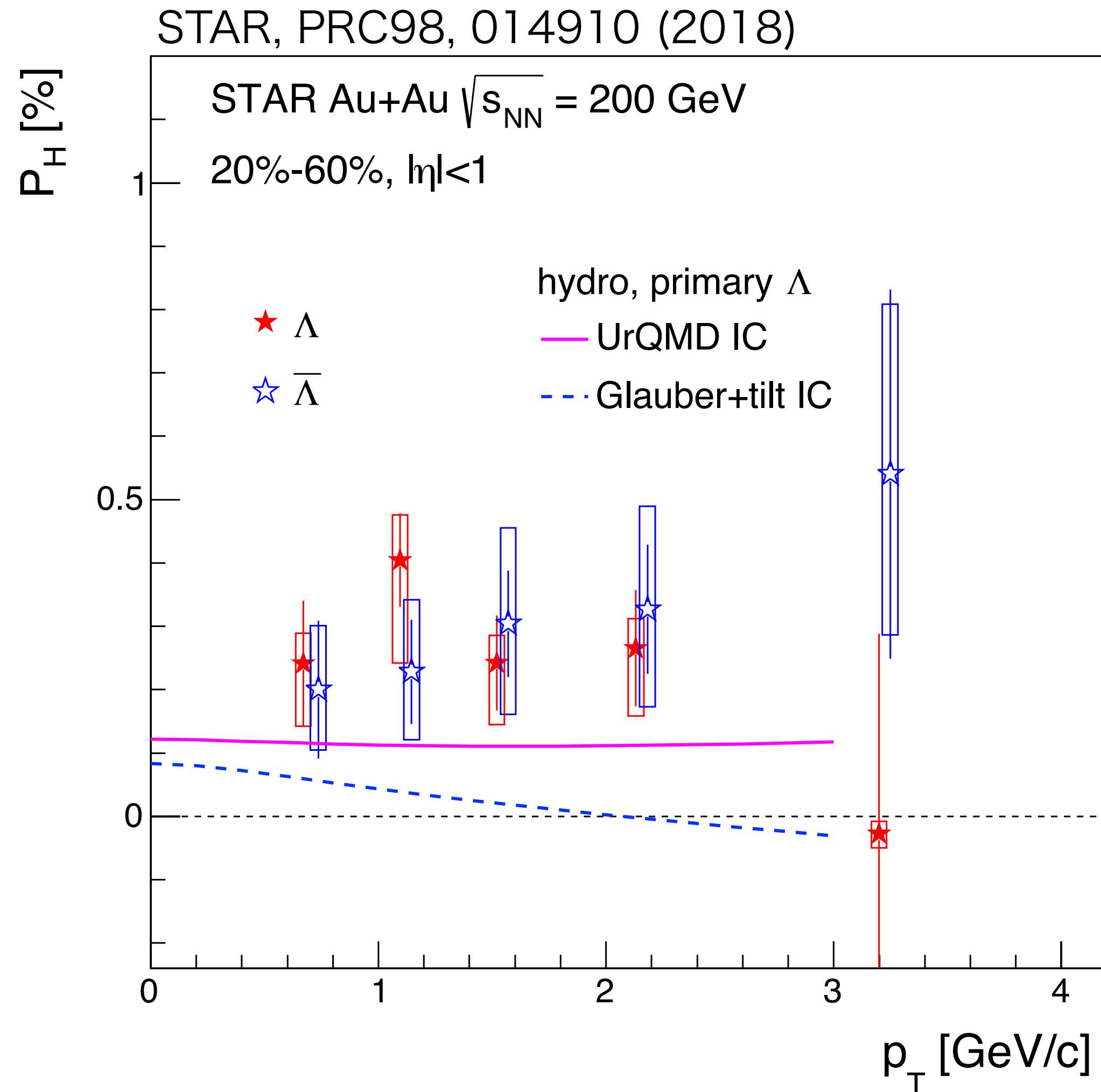
W.-T. Deng and X.-G. Huang, arXiv:1609.01801

STAR, PRC98, 014910 (2018)



- The data do not show significant  $\eta$  dependence
  - Maybe due to baryon transparency at higher energy
  - Also due to event-by-event C.M. fluctuations

# $p_T$ dependence of $P_H$



- No significant  $p_T$  dependence, as expected from the initial angular momentum of the system
- Hydrodynamic model underestimates the data. Initial conditions affect the magnitude and dependence on  $p_T$

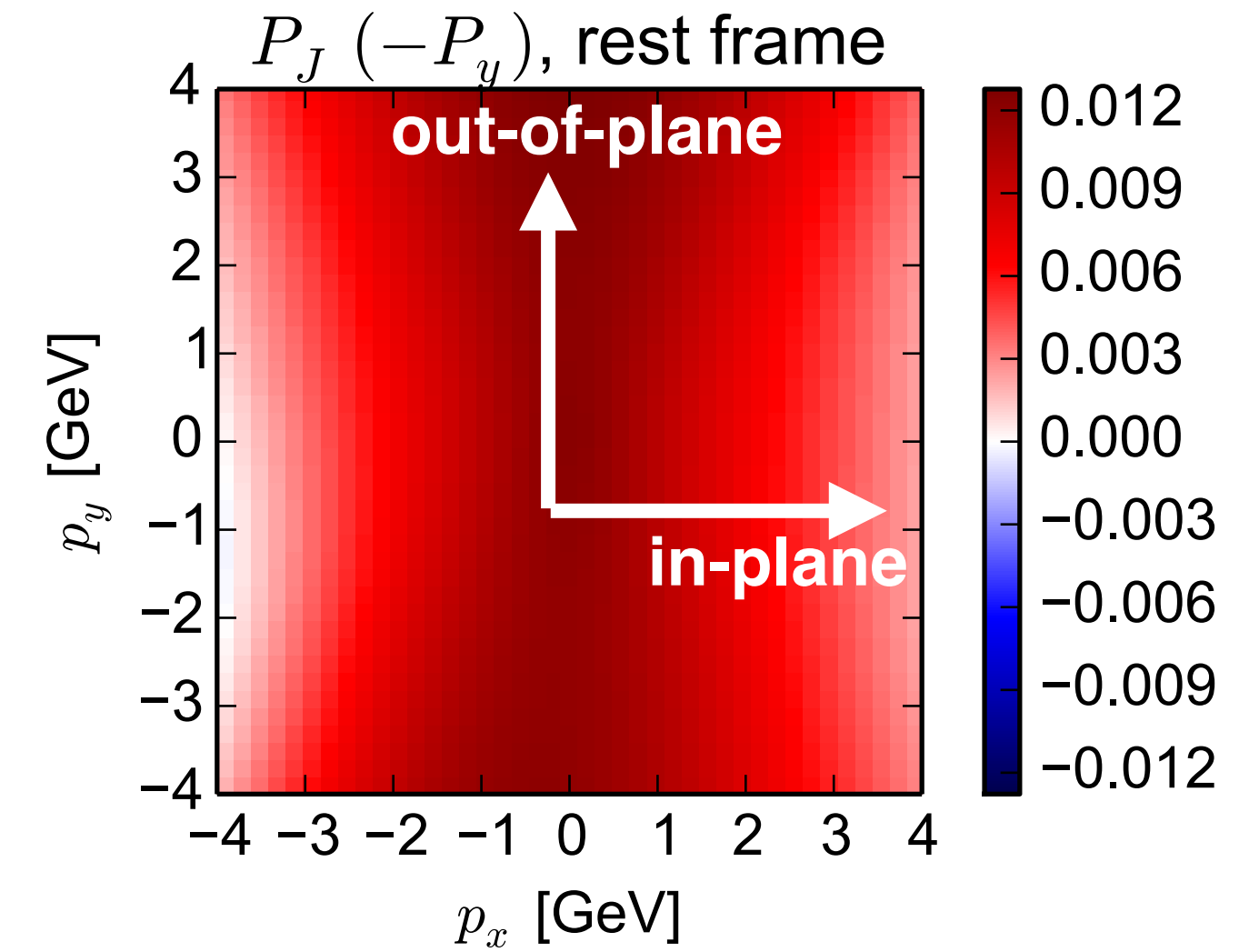
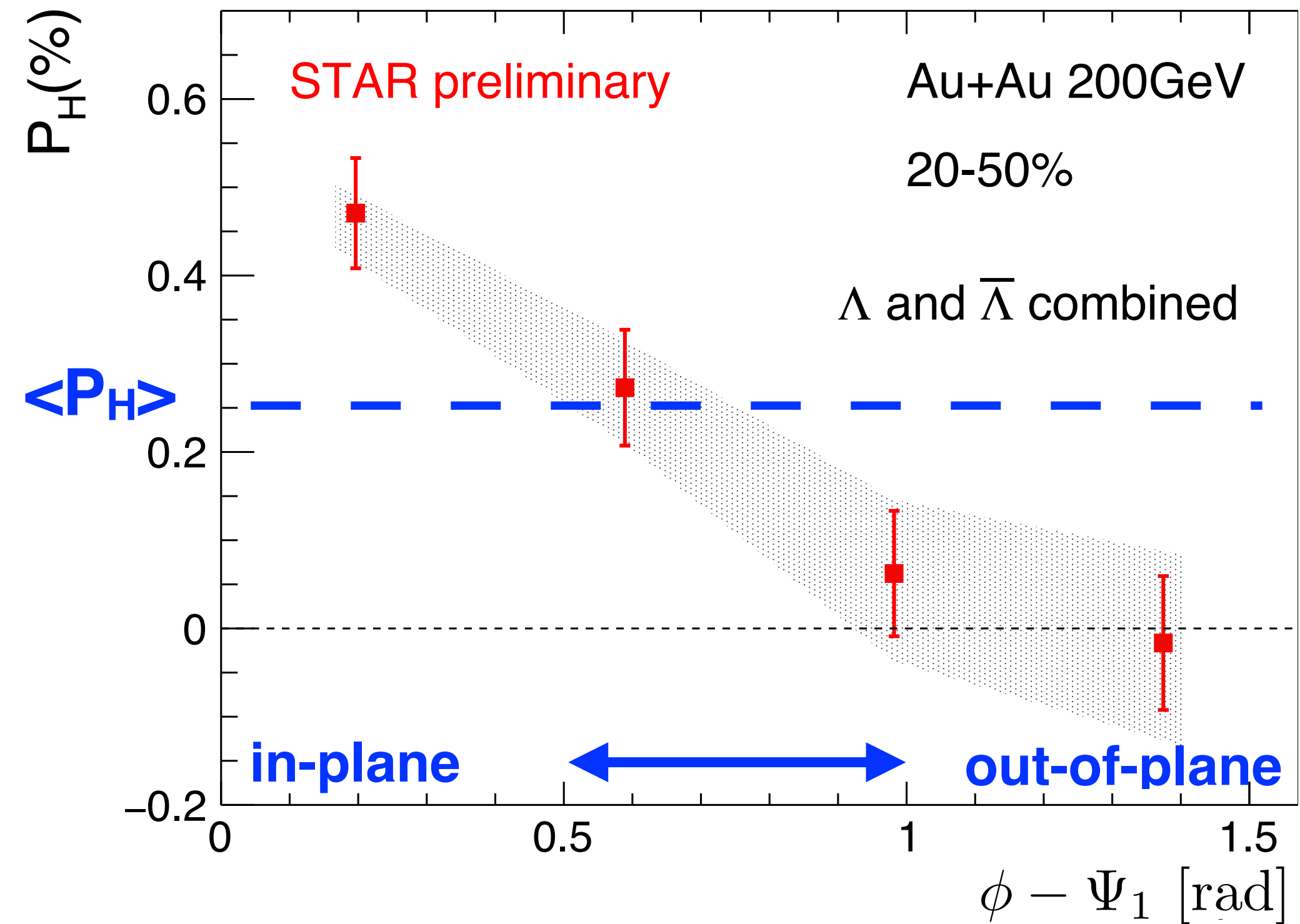
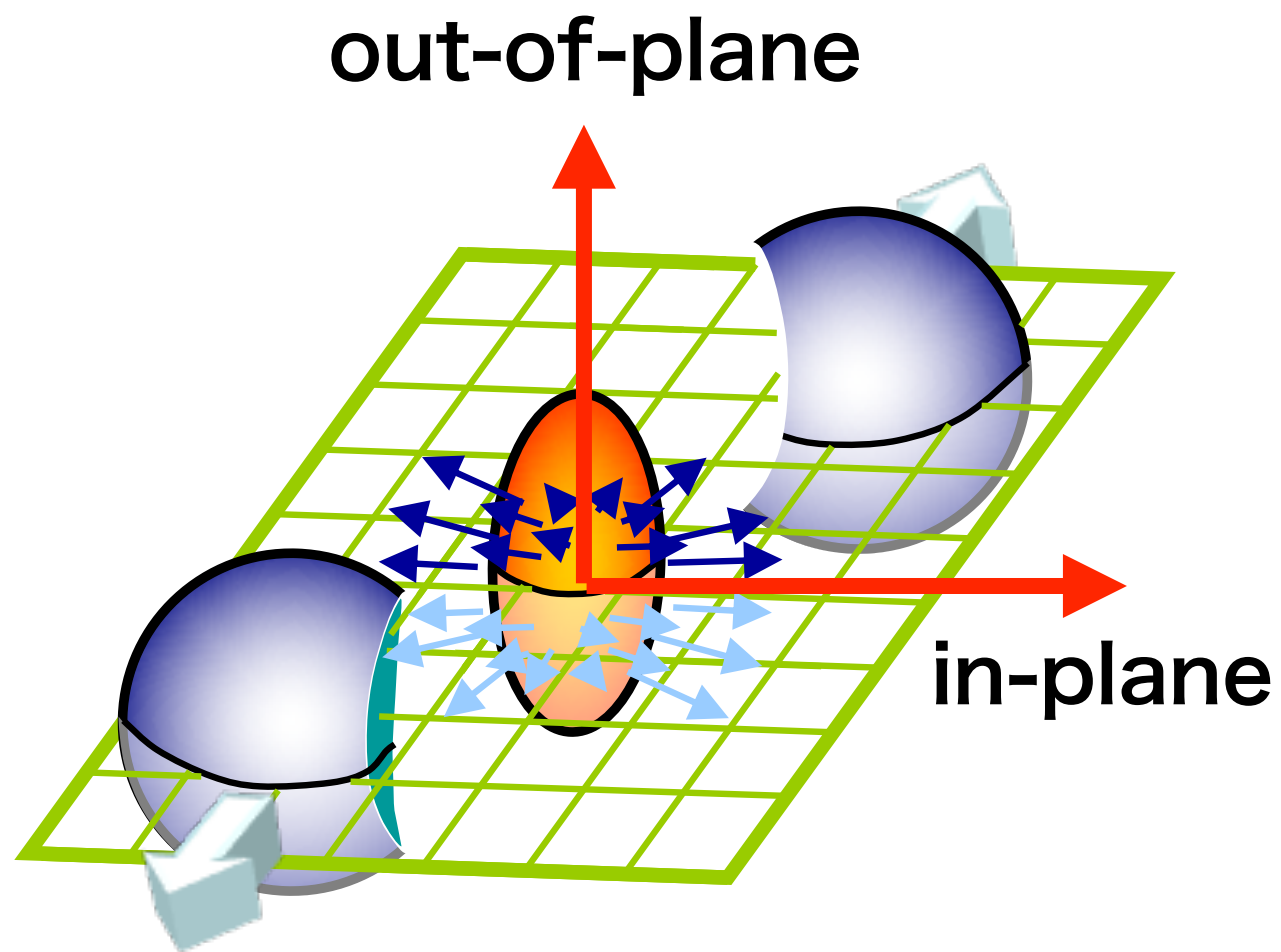
3D viscous hydrodynamic model with 2 initial conditions (ICs)

- UrQMD IC

- Glauber with source tilt IC

F. Becattini and I. Karpenko, PRL120.012302, 2018

# Azimuthal angle dependence of $P_H$



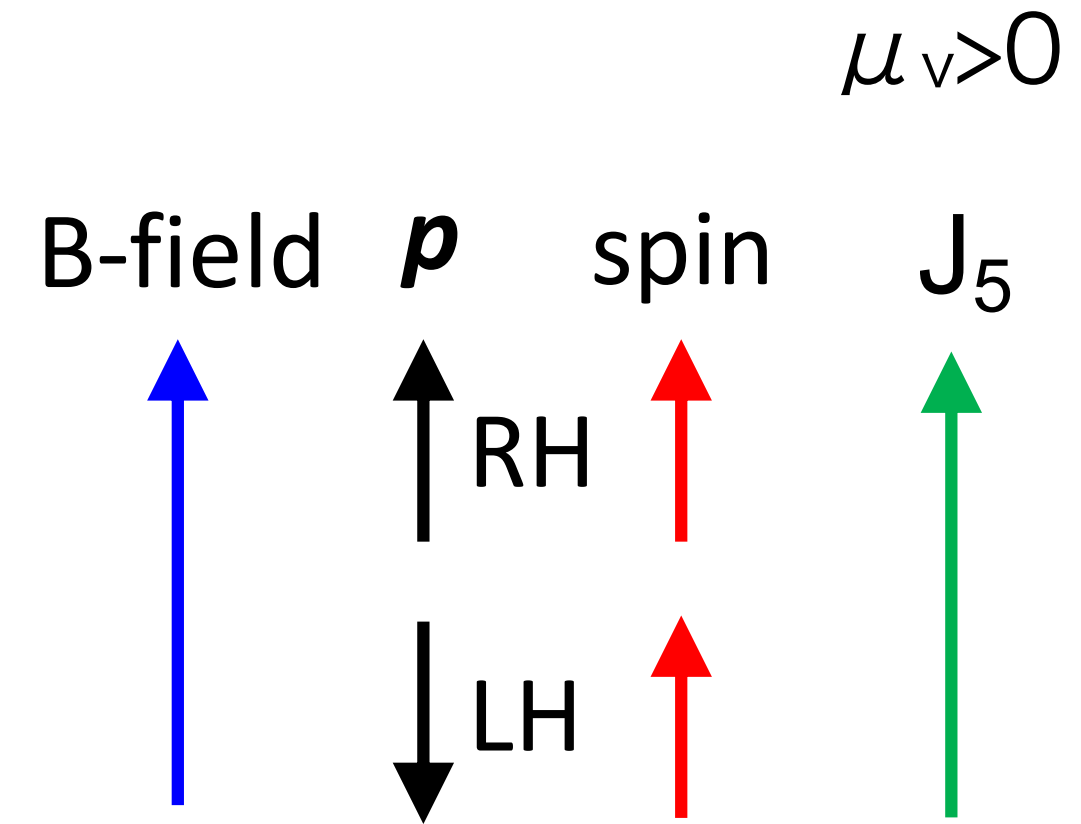
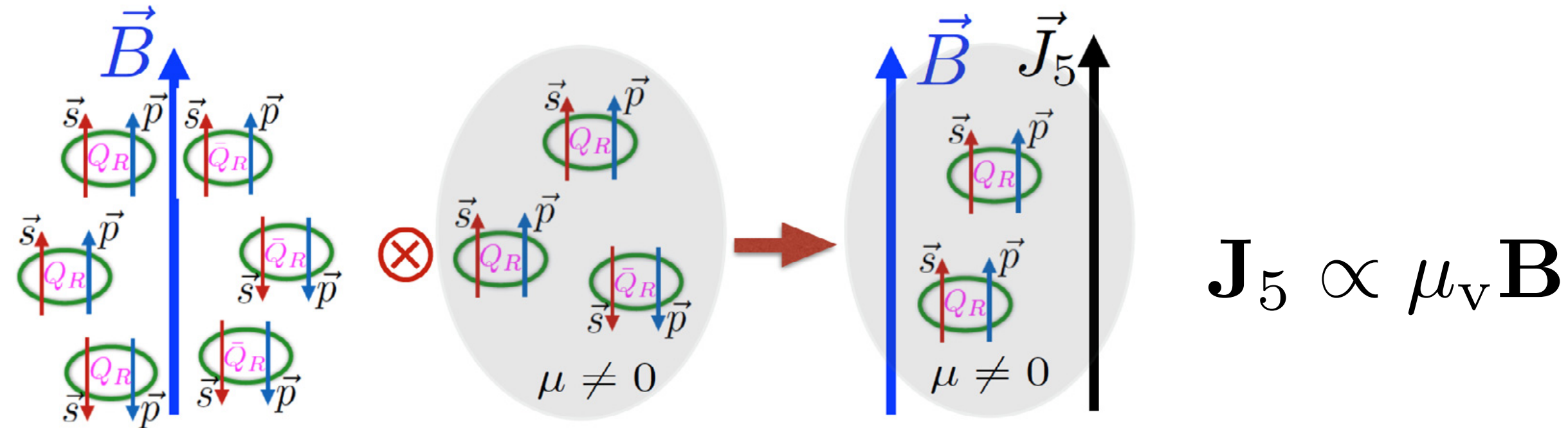
I. Karpenko and F. Becattini, EPJC(2017)77:213

- ◆ Larger polarization in in-plane than in out-of-plane
- ◆ Opposite to hydrodynamic model! (larger in out-of-plane)



# $\Lambda$ polarization vs. charge asymmetry?

## Chiral Separation Effect

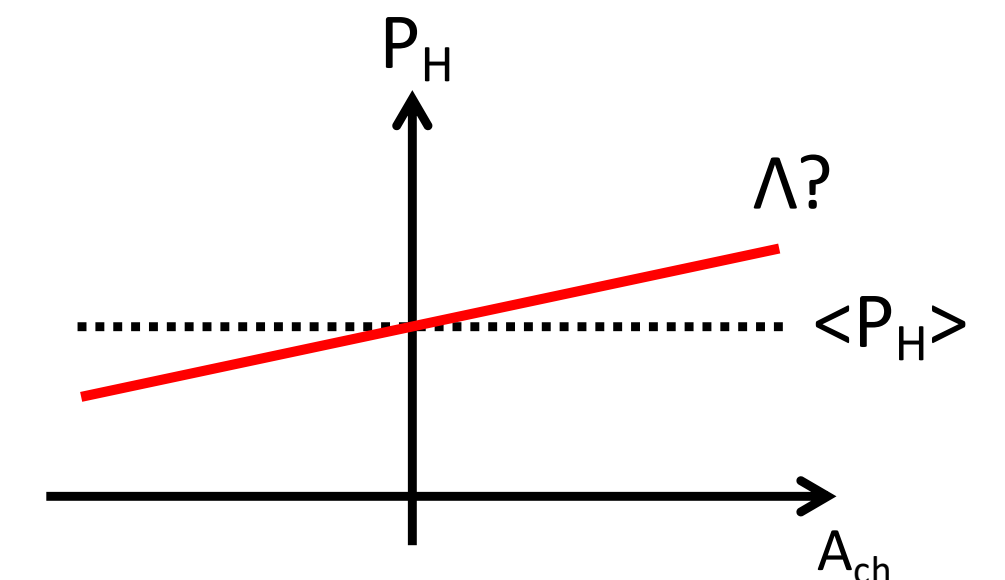


B-field + massless quarks + non-zero  $\mu_v \rightarrow$  axial current  $J_5$   
 (spin alignment + spin and momentum in (anti)parallel for RH(LH) quarks)

- $\Lambda$  polarization may have a contribution from the axial current  $J_5$  induced by B-field (Chiral Separation Effect), S. Shlichting and S. Voloshin
- Use charge asymmetry  $A_{ch}$  instead of  $\mu_v$

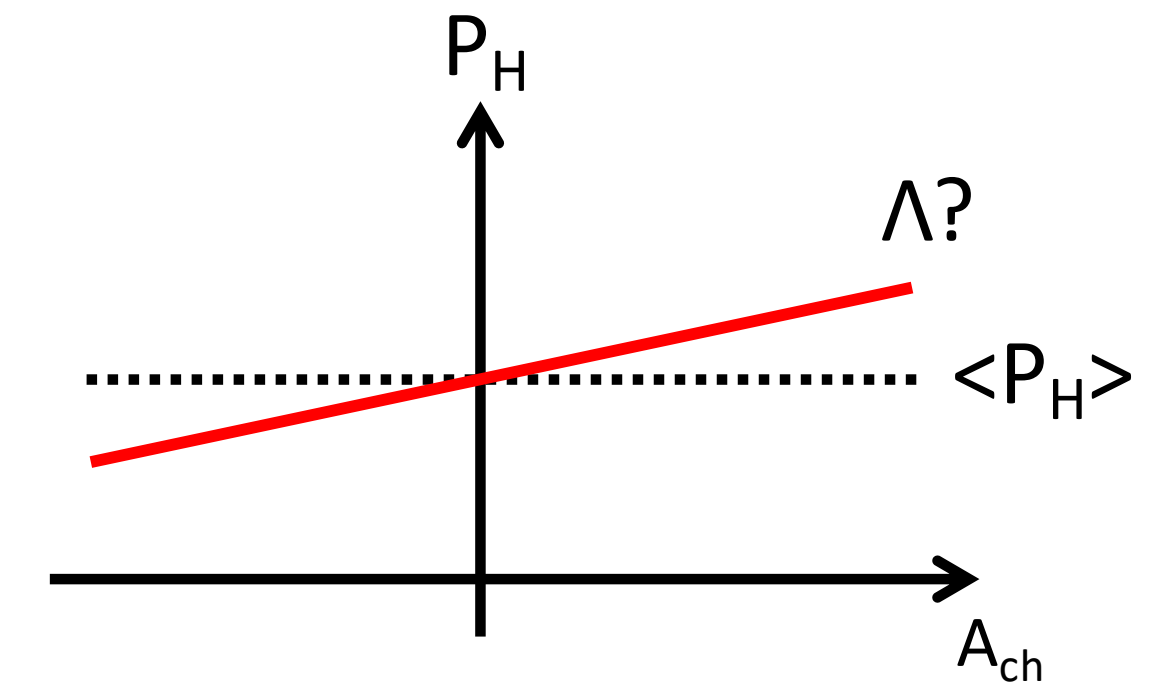
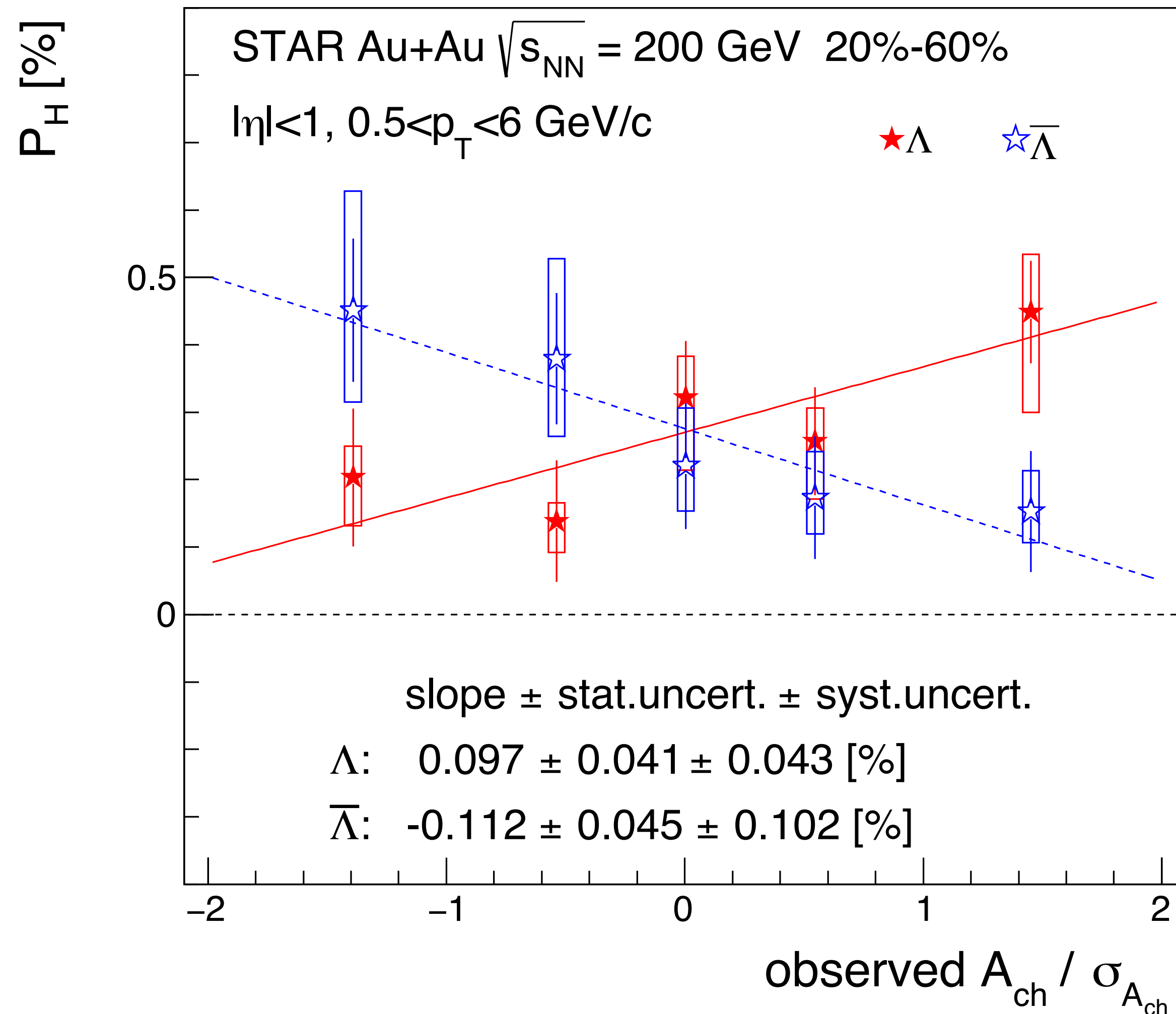
what's the expectation?  
 true for u-quark but also for  $\Lambda$ ?

$$\mu_v/T \propto \frac{\langle N_+ - N_- \rangle}{\langle N_+ + N_- \rangle} = A_{ch}$$



# $\Lambda$ polarization vs charge asymmetry?

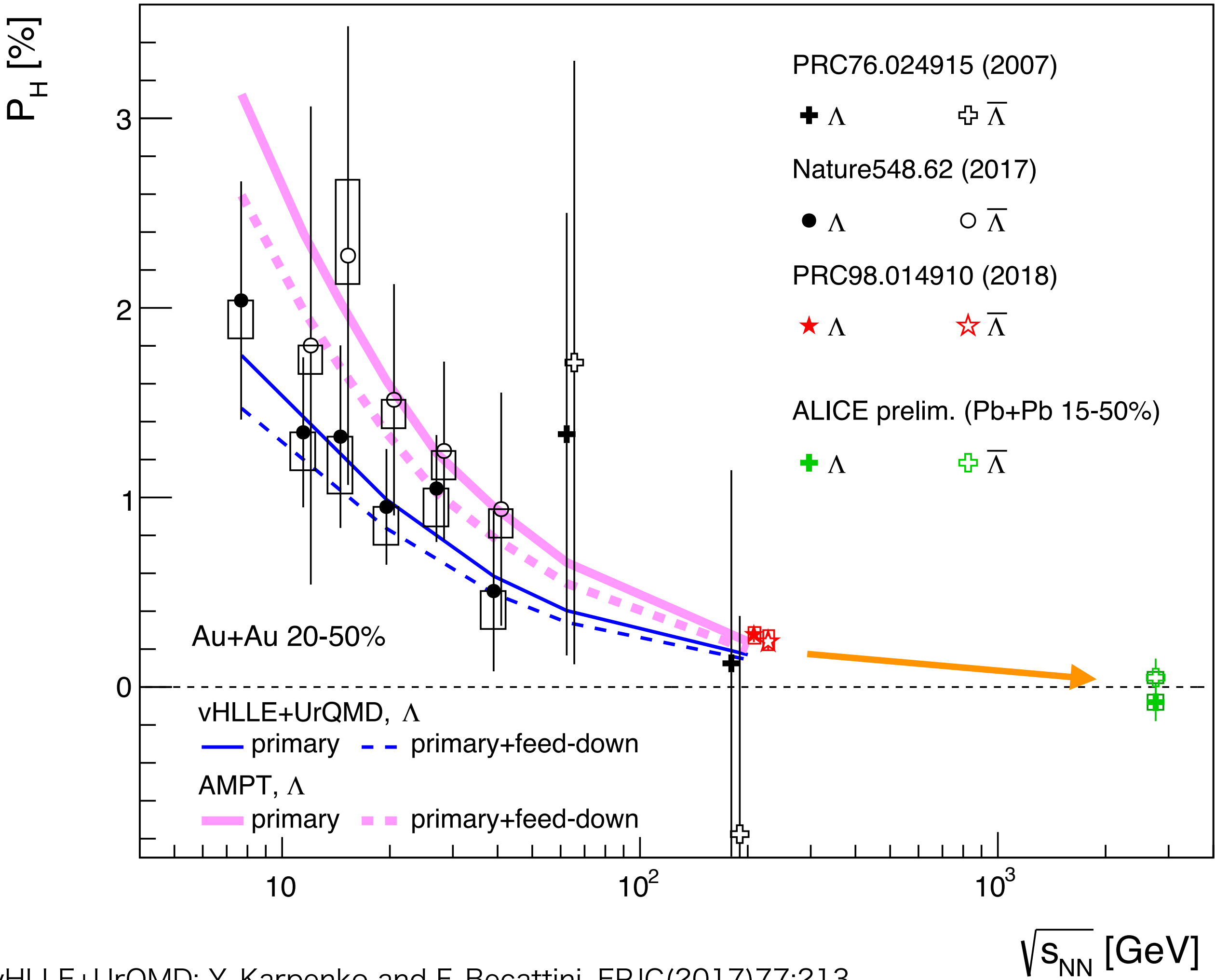
STAR, PRC98, 014910 (2018)



Slopes of  $\Lambda$  and anti- $\Lambda$  seem to be different.  
 (statistical significance is  $\sim 2\sigma$  level)

Possibly a contribution from the axial current?

# Go to the LHC energy



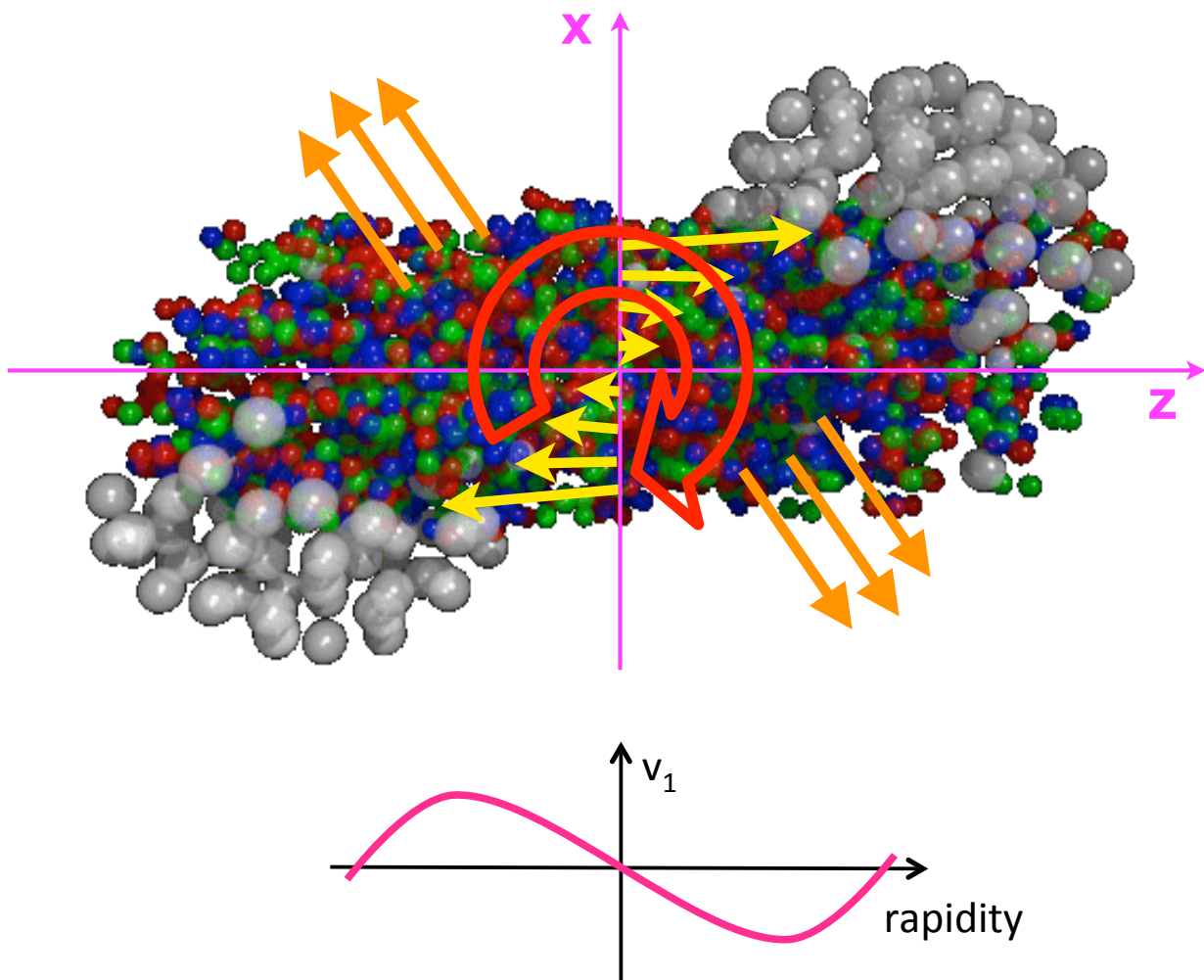
$$P_H(\Lambda) [\%] = -0.08 \pm 0.10 \text{ (stat)} \pm 0.04 \text{ (syst)}$$

$$P_H(\bar{\Lambda}) [\%] = 0.05 \pm 0.10 \text{ (stat)} \pm 0.03 \text{ (syst)}$$

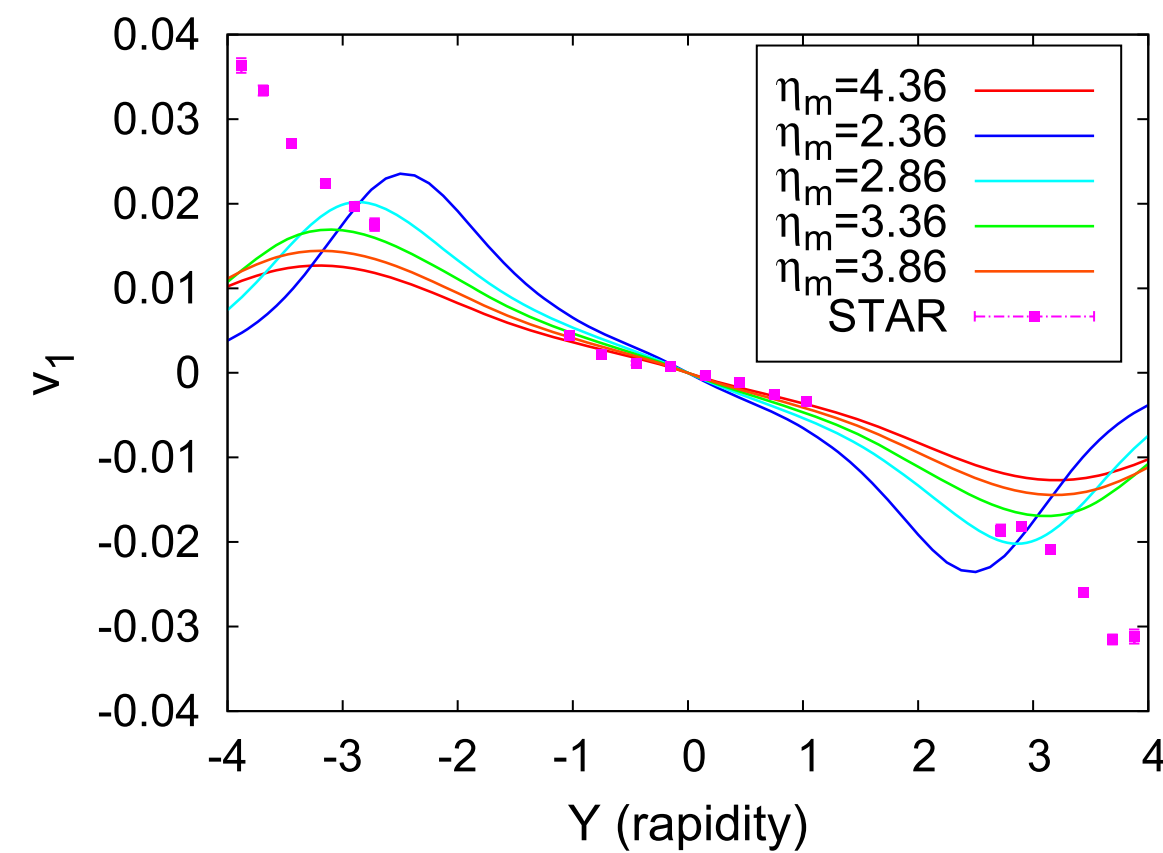
M. Konyushikhin (ALICE), QCD Chirality Workshop 2017

vHLLE+UrQMD: Y. Karpenko and F. Becattini, EPJC(2017)77:213  
 AMPT: H. Li et al., Phys. Rev. C 96, 054908 (2017)

# dv<sub>1</sub>/dy vs polarization in data



F. Becattini et al., Eur. Phys. J. C (2015)75:406

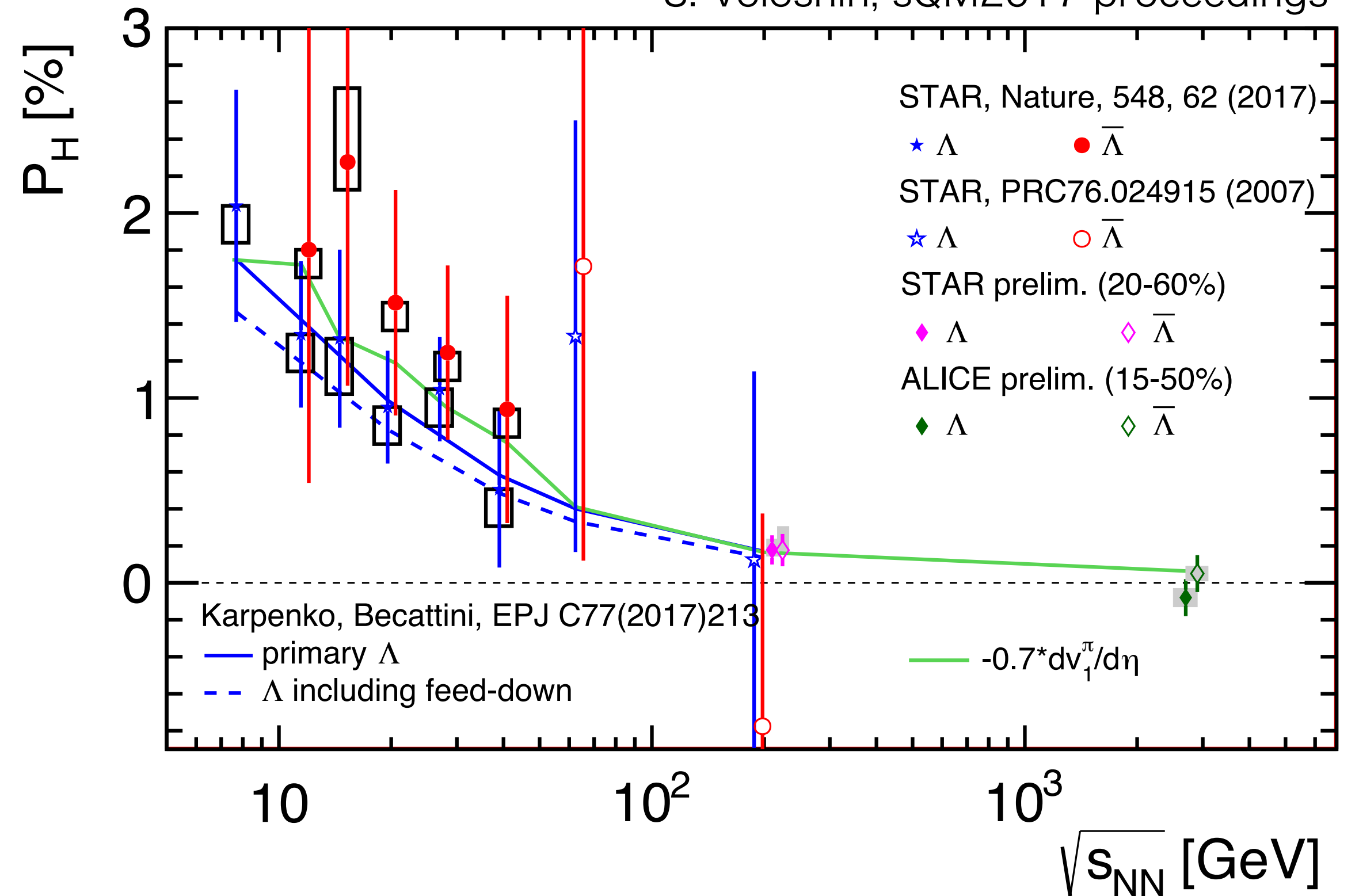


**Fig. 6** Directed flow of pions for different values of  $\eta_m$  parameter with  $\eta/s = 0.1$  compared with STAR data [22]

$$\frac{dN}{d\phi} \propto (1 + 2v_1 \cos(\phi - \Psi) + 2v_2 \cos(2\phi - 2\Psi) + \dots)$$

- Vorticity is likely related to the directed flow.
- The tilted source accounting for vorticity provides a better description of v<sub>1</sub>!

S. Voloshin, sQM2017 proceedings

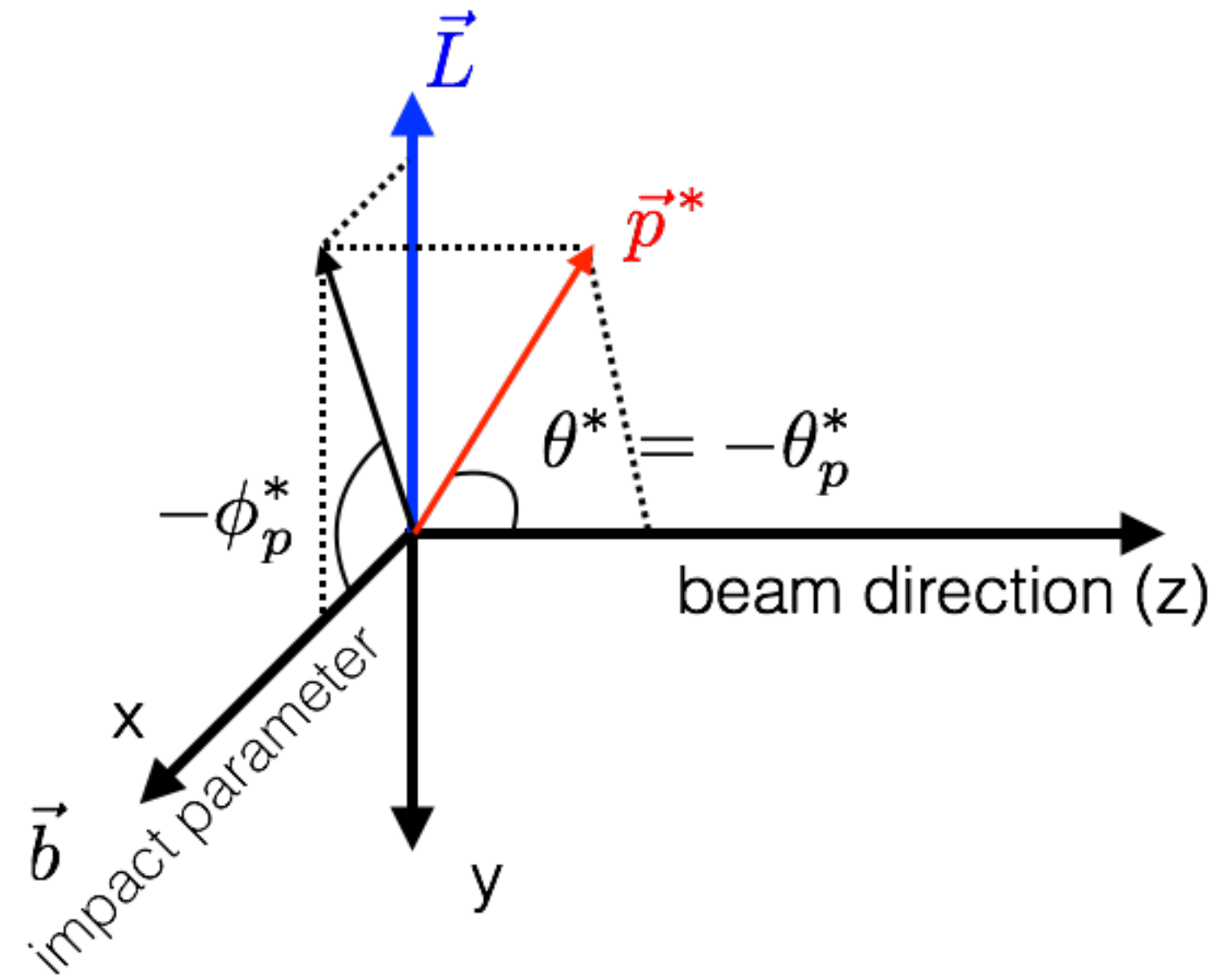
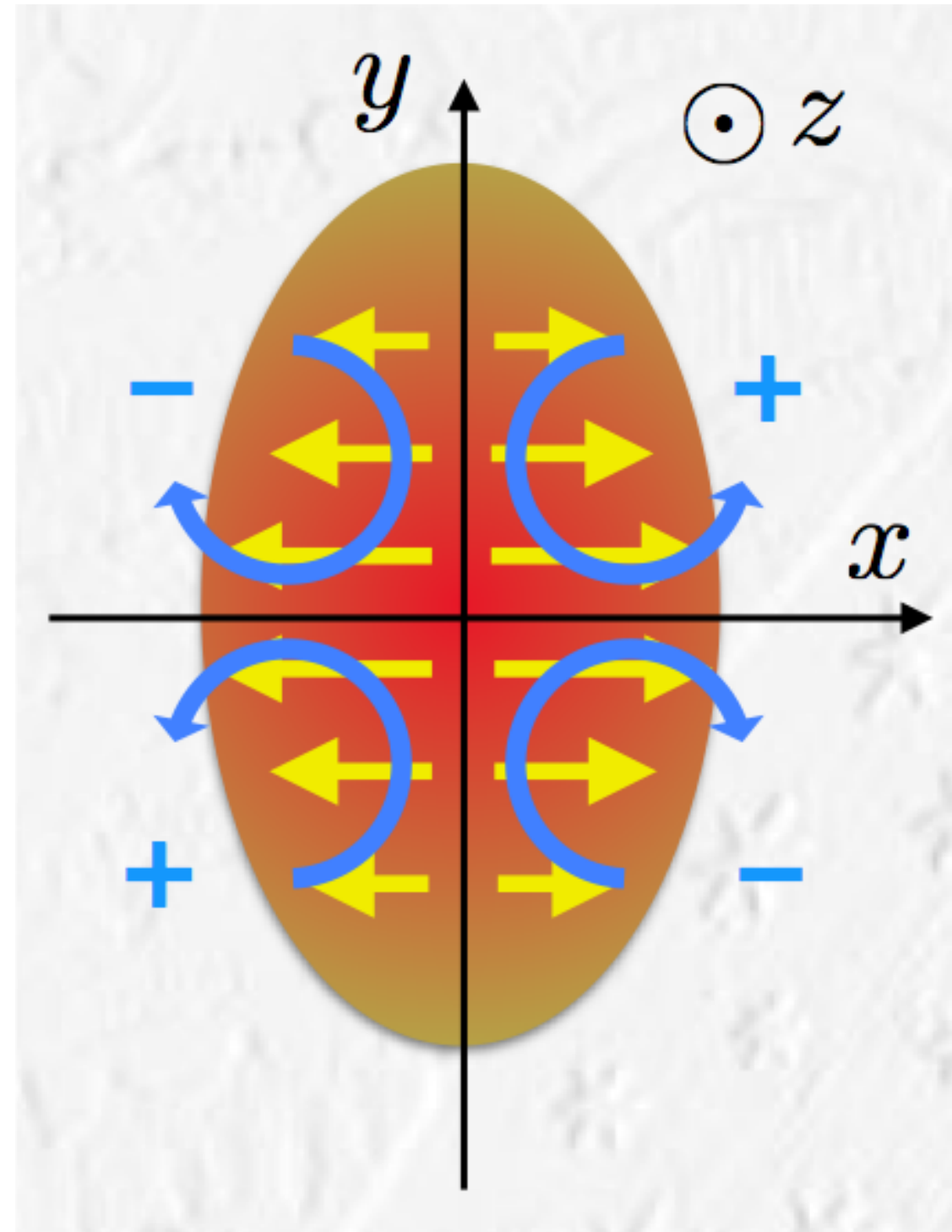


A similar energy dependence of dv<sub>1</sub>/dy to the polarization!

# Polarization along the beam direction

S. Voloshin, SQM2017

F. Becattini and I. Karpenko, PRL120.012302 (2018)



$$\begin{aligned} \frac{dN}{d\Omega^*} &= \frac{1}{4\pi} (1 + \alpha_H \mathbf{P}_H \cdot \mathbf{p}_p^*) \\ \langle \cos \theta_p^* \rangle &= \int \frac{dN}{d\Omega^*} \cos \theta_p^* d\Omega^* \\ &= \alpha_H P_z \langle (\cos \theta_p^*)^2 \rangle \\ \therefore P_z &= \frac{\langle \cos \theta_p^* \rangle}{\alpha_H \langle (\cos \theta_p^*)^2 \rangle} \\ &= \frac{3 \langle \cos \theta_p^* \rangle}{\alpha_H} \quad (\text{if perfect detector}) \end{aligned}$$

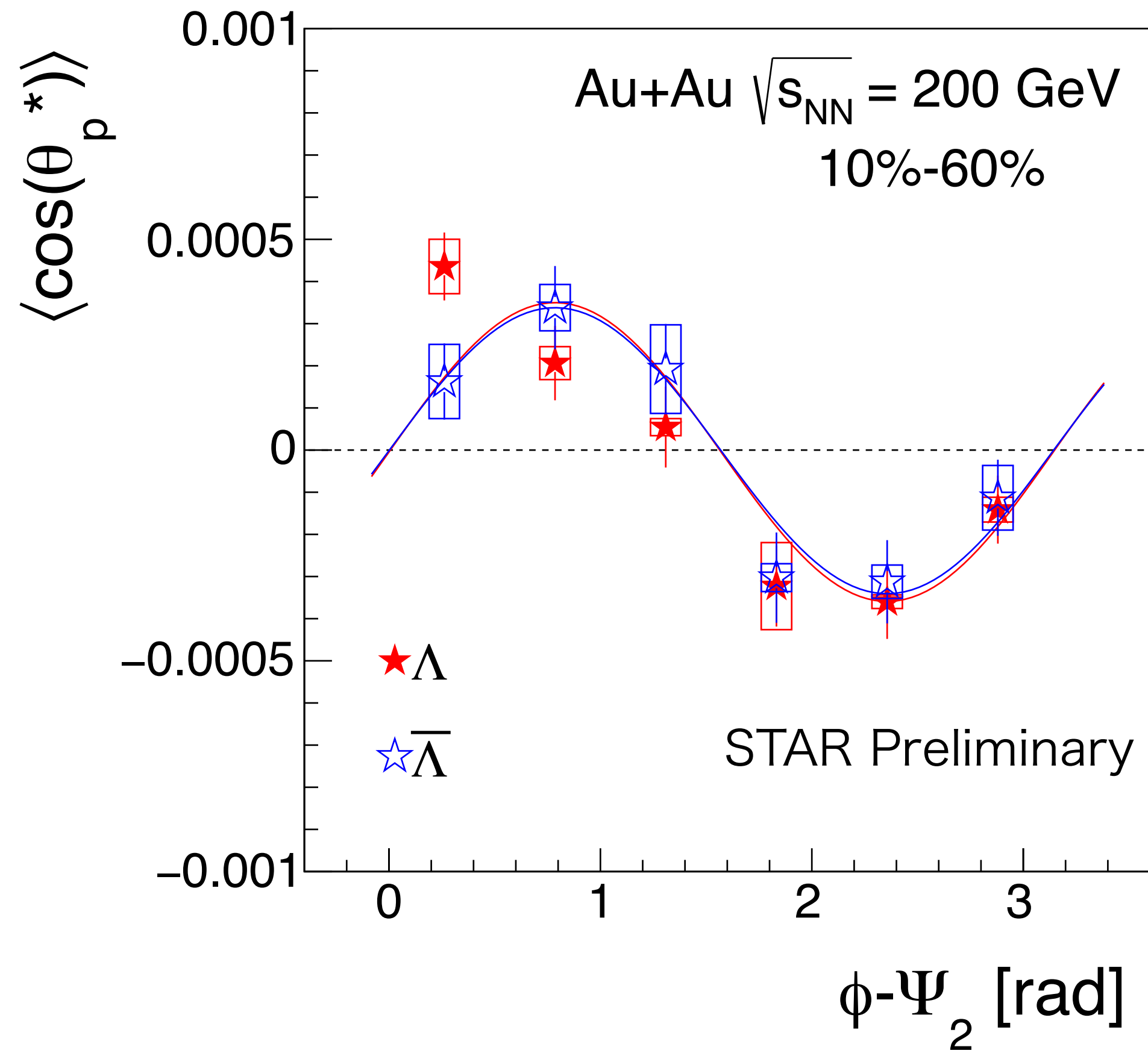
$\alpha_H$ : hyperon decay parameter

$\theta_p^*$ :  $\theta$  of daughter proton in  $\Lambda$  rest frame

Stronger flow in in-plane than in out-of-plane could make local polarization along beam axis!

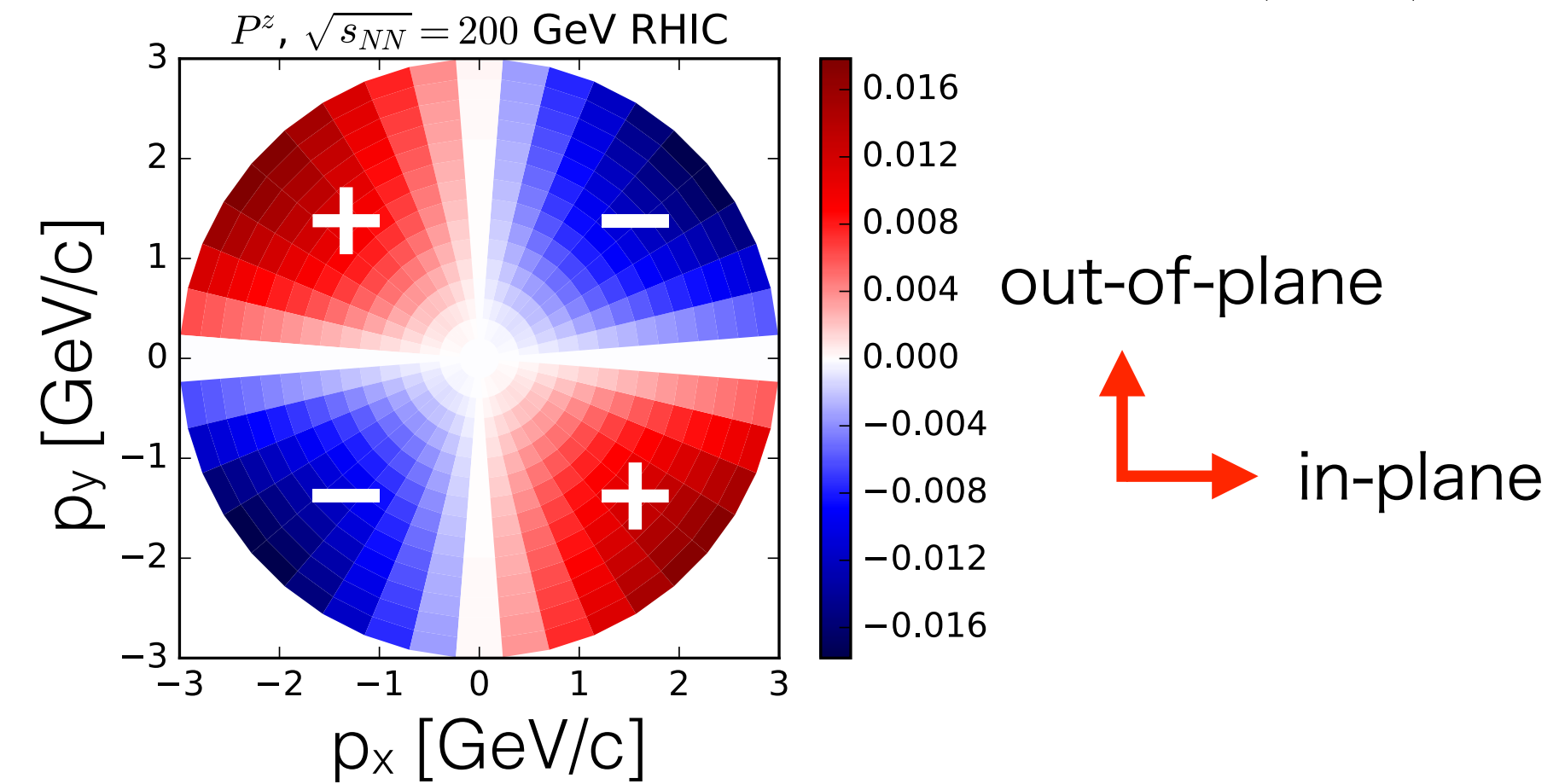
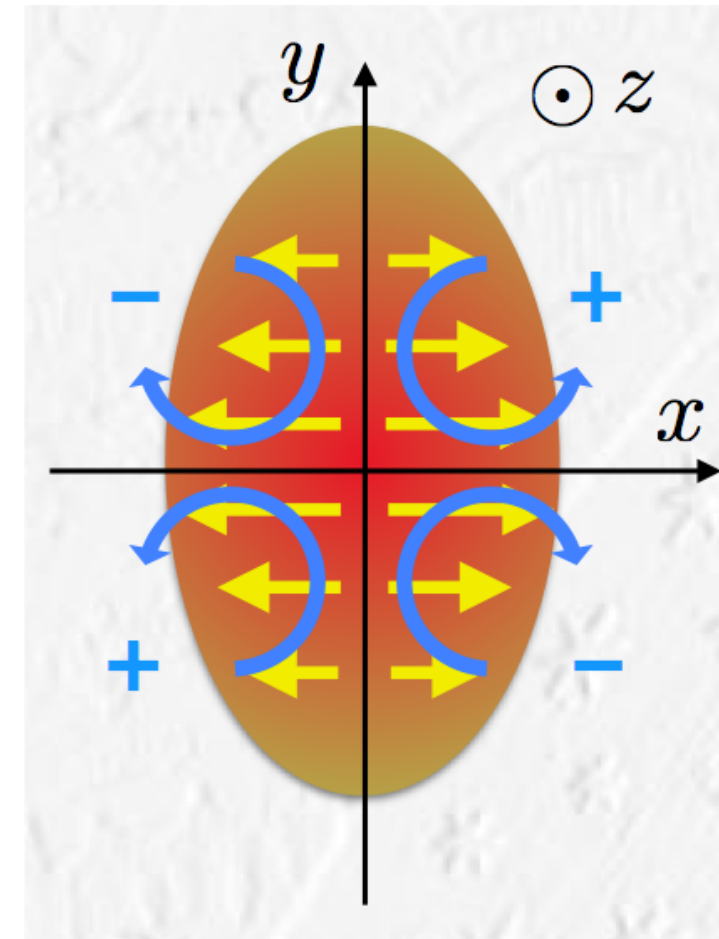
Longitudinal component,  $P_z$ , can be expressed with  $\langle \cos \theta_p^* \rangle$ .  $\langle (\cos \theta_p^*)^2 \rangle$  accounts for an acceptance effect

# Polarization along the beam direction



- Effect of  $\Psi_2$  resolution is not corrected here

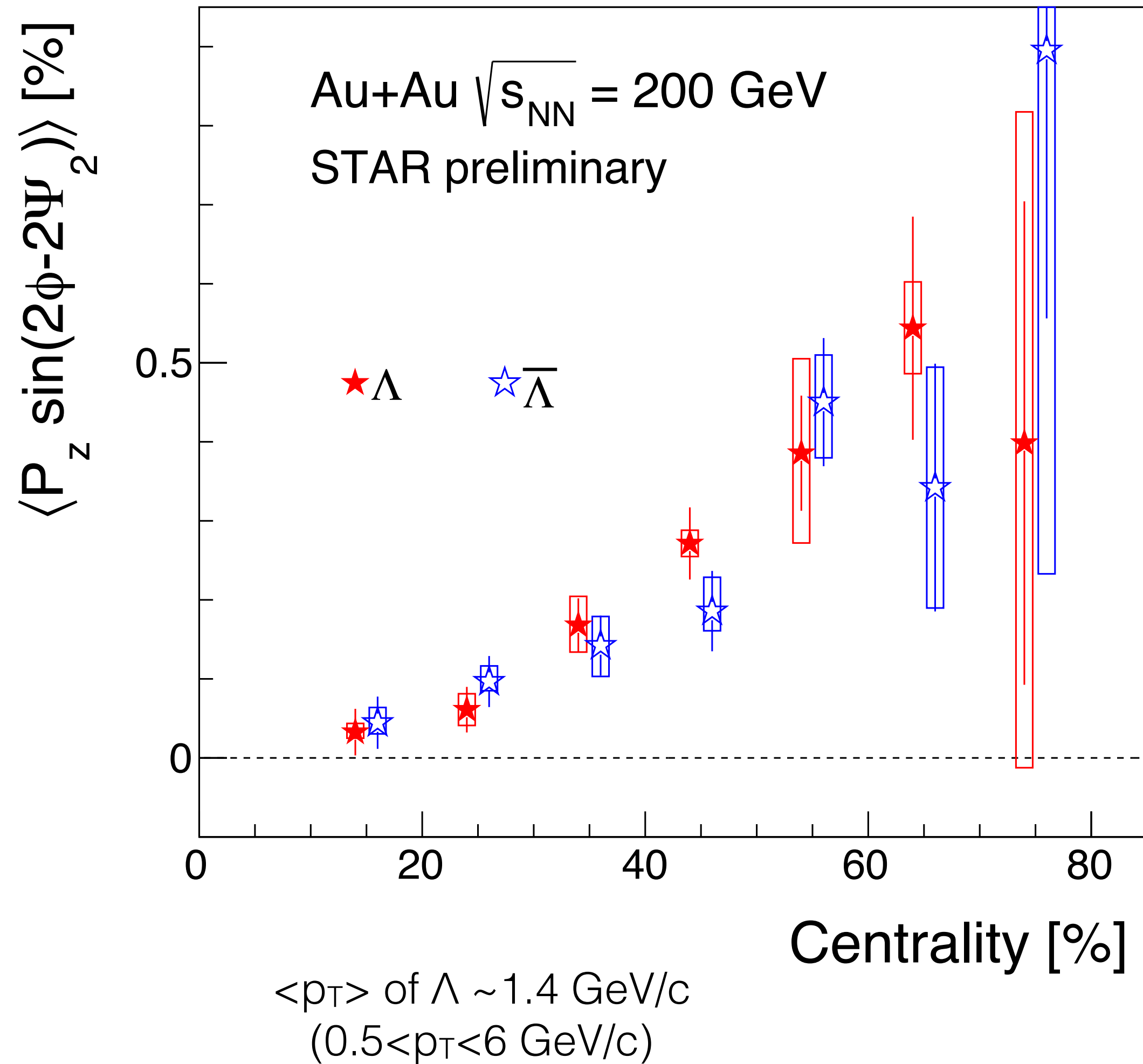
S. Voloshin, SQM2017



Hydro calculation of  $P_z$   
F. Becattini and I. Karpenko,  
PRL.120.012302 (2018)

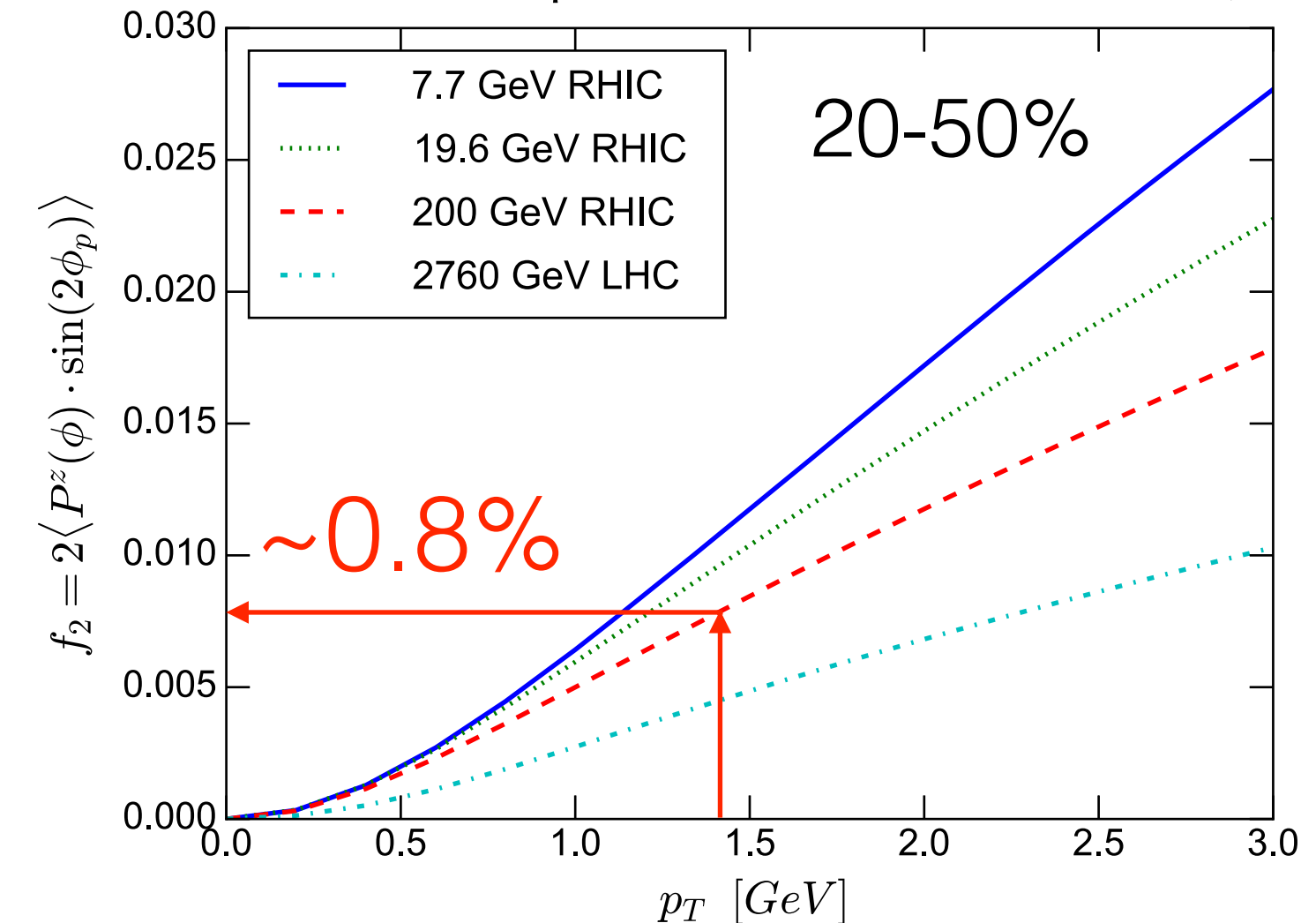
- Sine structure as expected from the elliptic flow!
- Opposite sign to hydrodynamic model and a transport model (AMPT)
  - Hydro model: F. Becattini and I. Karpenko, PRL.120.012302 (2018)
  - AMPT model: X. Xia, H. Li, Z. Tang, Q. Wang, arXiv:1803.0086

# Centrality dependence of $P_z$ modulation



- Strong centrality dependence as in  $v_2$
- Similar magnitude to the global polarization
- $\sim 5$  times smaller magnitude than the hydro and AMPT with the opposite sign!

F. Becattini and I. Karpenko, PRL.120.012302 (2018)



# Sign problem in $P_z$

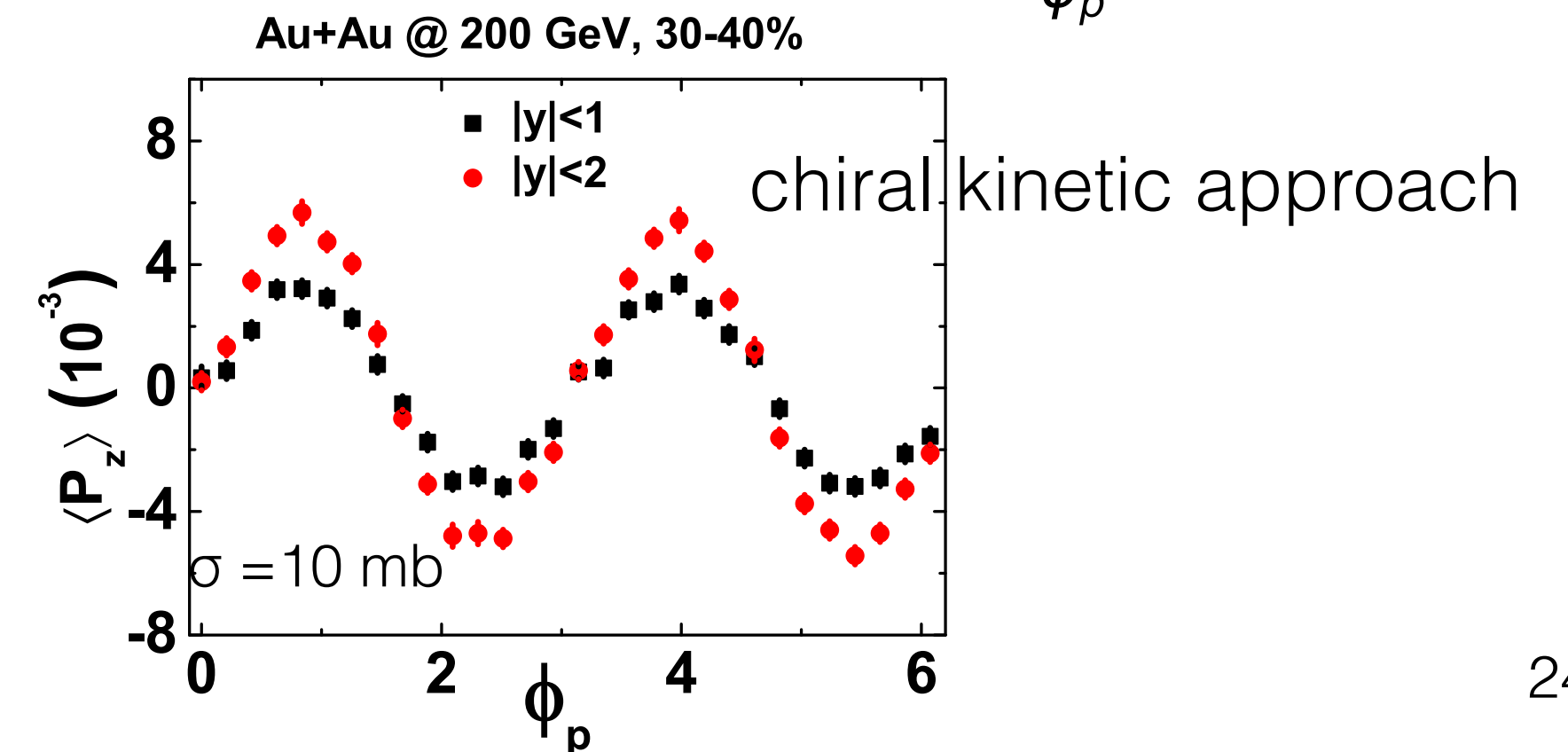
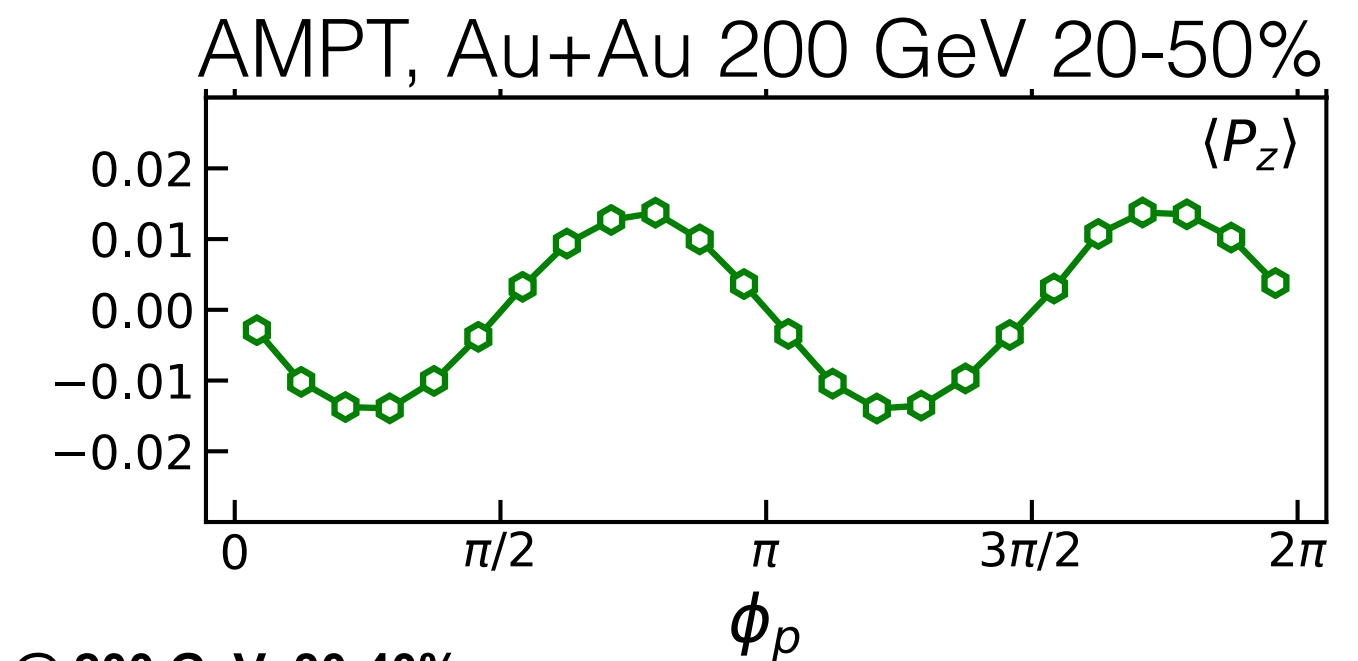
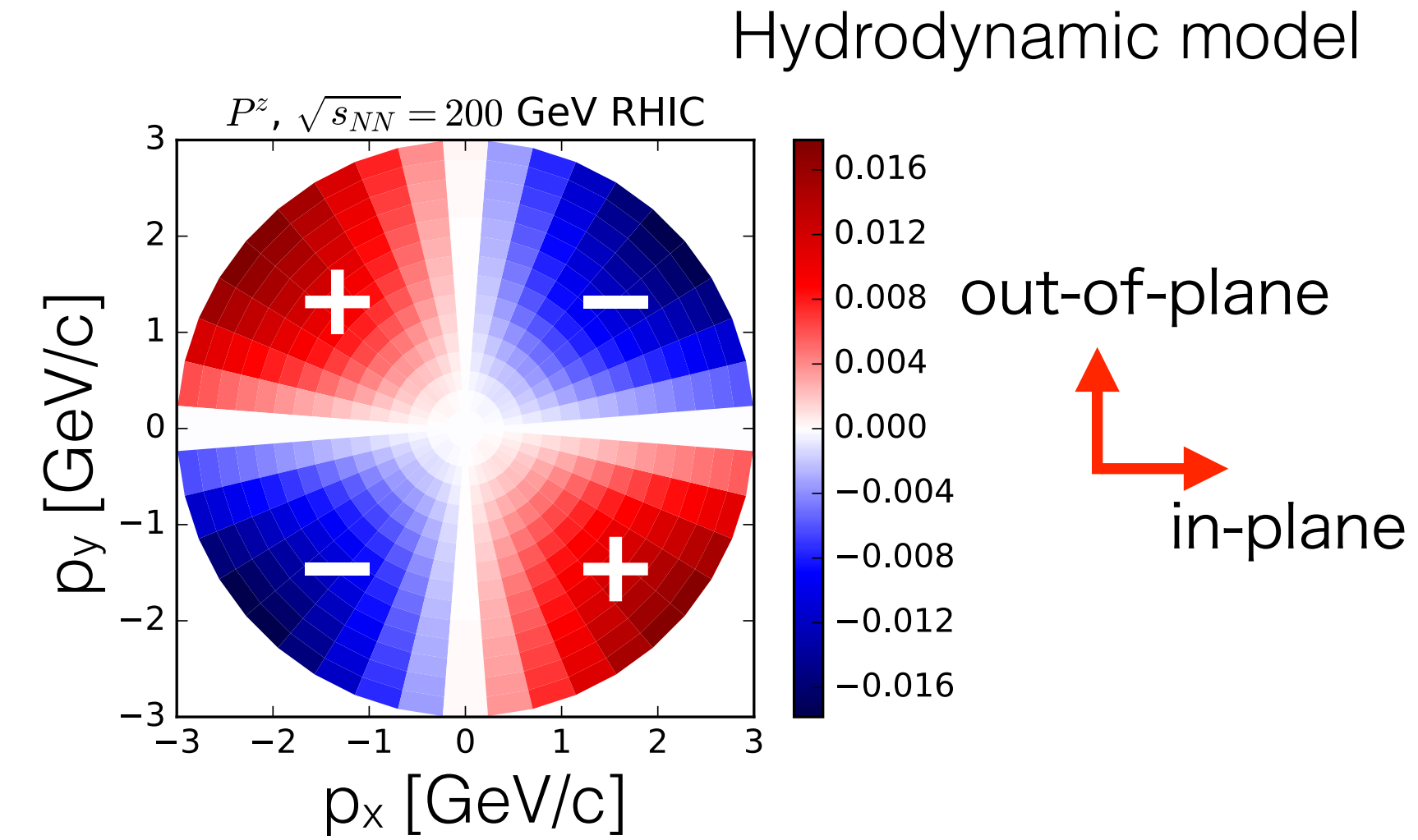
Opposite sign to hydrodynamic model and AMPT model

- F. Becattini and I. Karpenko, PRL.120.012302 (2018)  
3D viscous hydrodynamic model with UrQMD initial condition assuming a local thermal equilibrium
- AMPT: X. Xia, H. Li, Z. Tang, Q. Wang, PRC98.024905 (2018)

Same sign as chiral kinetic approach

- Y. Sun and C.-M. Ko, arXiv:1810.10359
- Assuming non-equilibrium of spin degree of freedom
- Smaller quark scattering cross section changes the sign

Suggest incomplete thermal equilibrium of spin degree of freedom as it develops later in time unlike the global polarization?





# Contributions to $P_z$ in hydro

I. Karpenko, QM2018

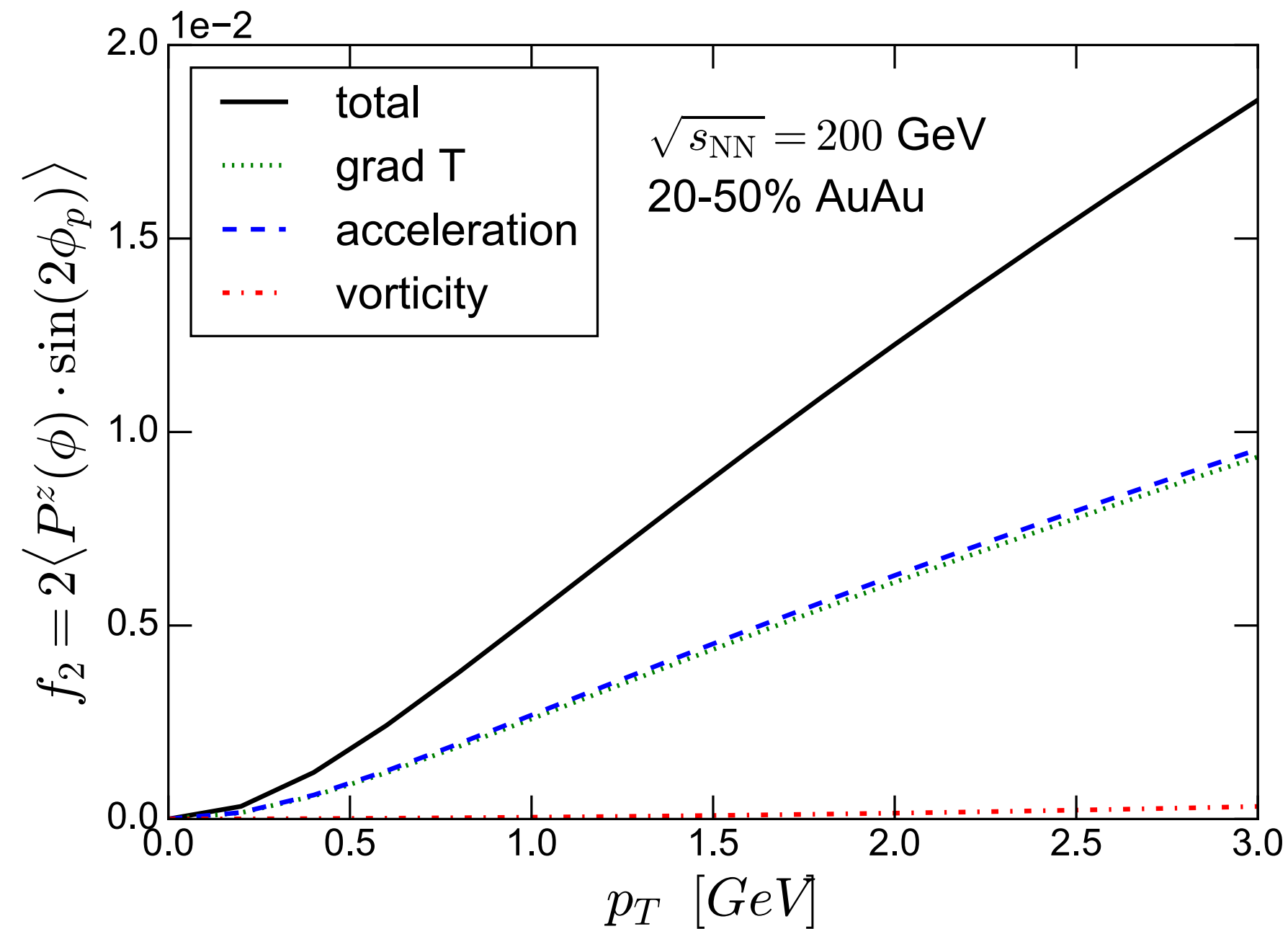
$$S^\mu \propto \varepsilon^{\mu\rho\sigma\tau} \omega_{\rho\sigma} p_\tau = \varepsilon^{\mu\rho\sigma\tau} (\partial_\rho \beta_\sigma) p_\tau = \underbrace{\varepsilon^{\mu\rho\sigma\tau} p_\tau \partial_\rho \left( \frac{1}{T} \right) u_\sigma}_{\text{grad}T} + \underbrace{\frac{1}{T} 2 [\omega^\mu (u \cdot p) - u^\mu (\omega \cdot p)]}_{\text{"NR vorticity"}} + \underbrace{\varepsilon^{\mu\rho\sigma\tau} p_\tau A_\sigma u_\rho}_{\text{acceleration}}$$

temperature gradient

kinematic vorticity

relativistic term

Longitudinal quadrupole  $f_2$ :



$P_z$  dominated by temperature gradient and relativistic term, but not by kinematic vorticity based on the hydro model.

How small is the kinematic vorticity?

Can we estimate it with the blast-wave model?

# Blast-wave model

- Hydro inspired model parameterized with freeze-out condition assuming the longitudinal boost invariance
  - Freeze-out temperature  $T_f$
  - Radial flow rapidity  $\rho_0$  and its modulation  $\rho_2$
  - Source size  $R_x$  and  $R_y$

$$\rho(r, \phi_s) = \tilde{r}[\rho_0 + \rho_2 \cos(2\phi_b)]$$

$$\tilde{r}(r, \phi_s) = \sqrt{(r \cos \phi_s)^2 / R_x^2 + (r \sin \phi_s)^2 / R_y^2}$$

- Calculate vorticity at the freeze-out using the parameters extracted from spectra,  $v_2$ , and HBT fit

$$\langle \omega_z \sin(2\phi) \rangle = \frac{\int d\phi_s \int r dr I_2(\alpha_t) K_1(\beta_t) \omega_z \sin(2\phi_b)}{\int d\phi_s \int r dr I_0(\alpha_t) K_1(\beta_t)}$$

$$\omega_z = \frac{1}{2} \left( \frac{\partial u_y}{\partial x} - \frac{\partial u_x}{\partial y} \right),$$

u: local flow velocity,  $I_n$ ,  $K_n$ : modified Bessel functions

F. Retiere and M. Lisa, PRC70.044907 (2004)

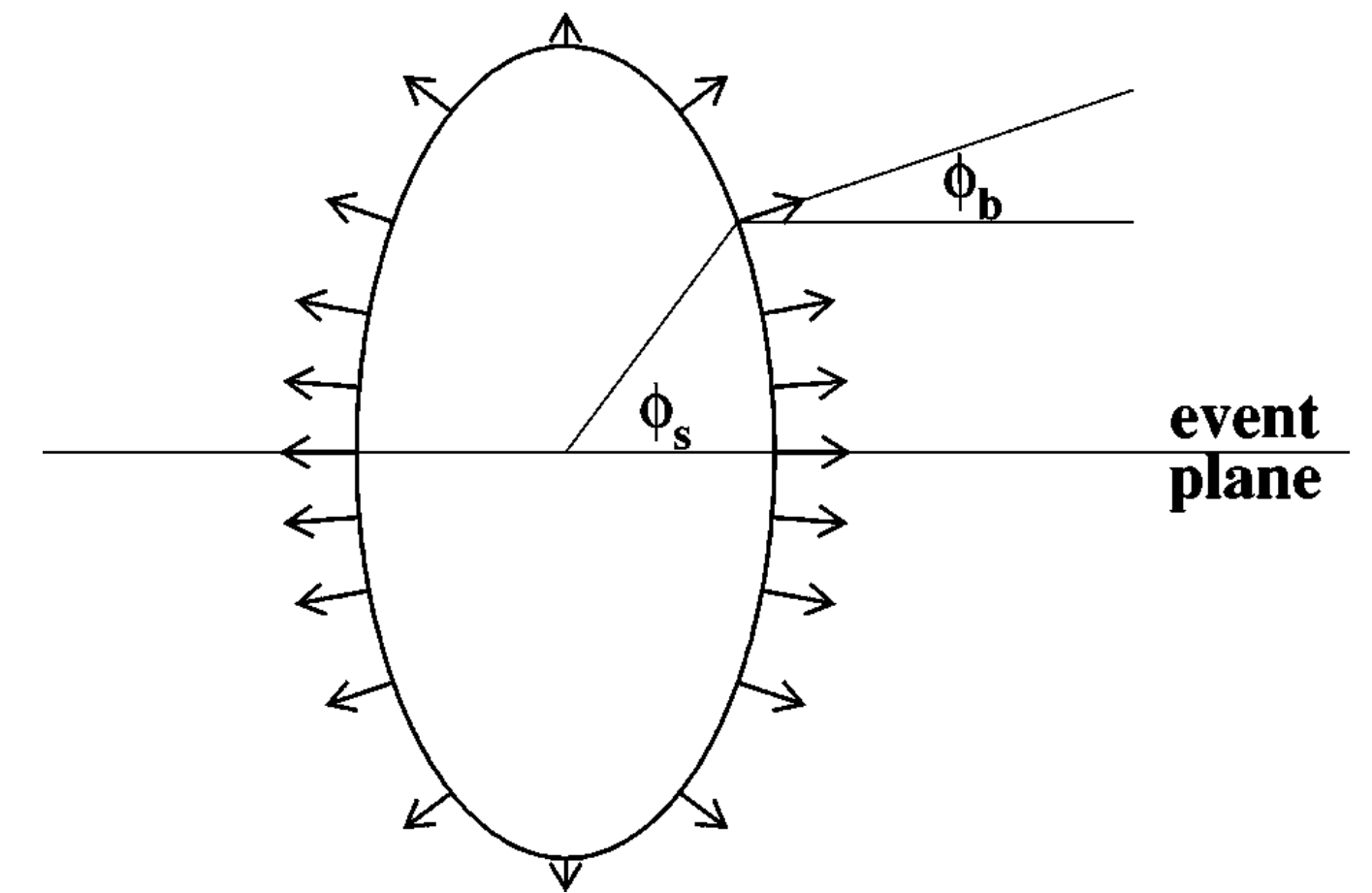
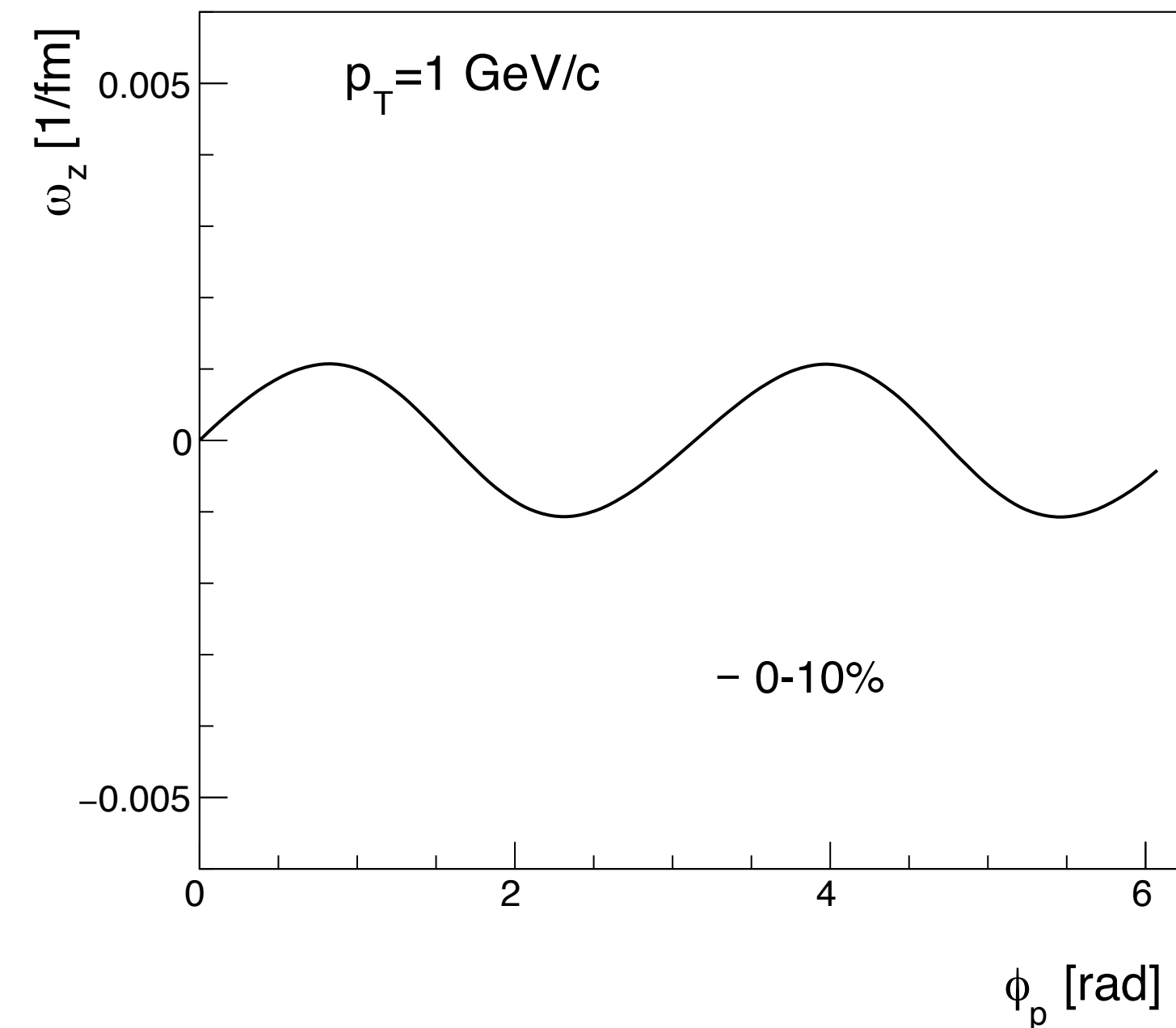
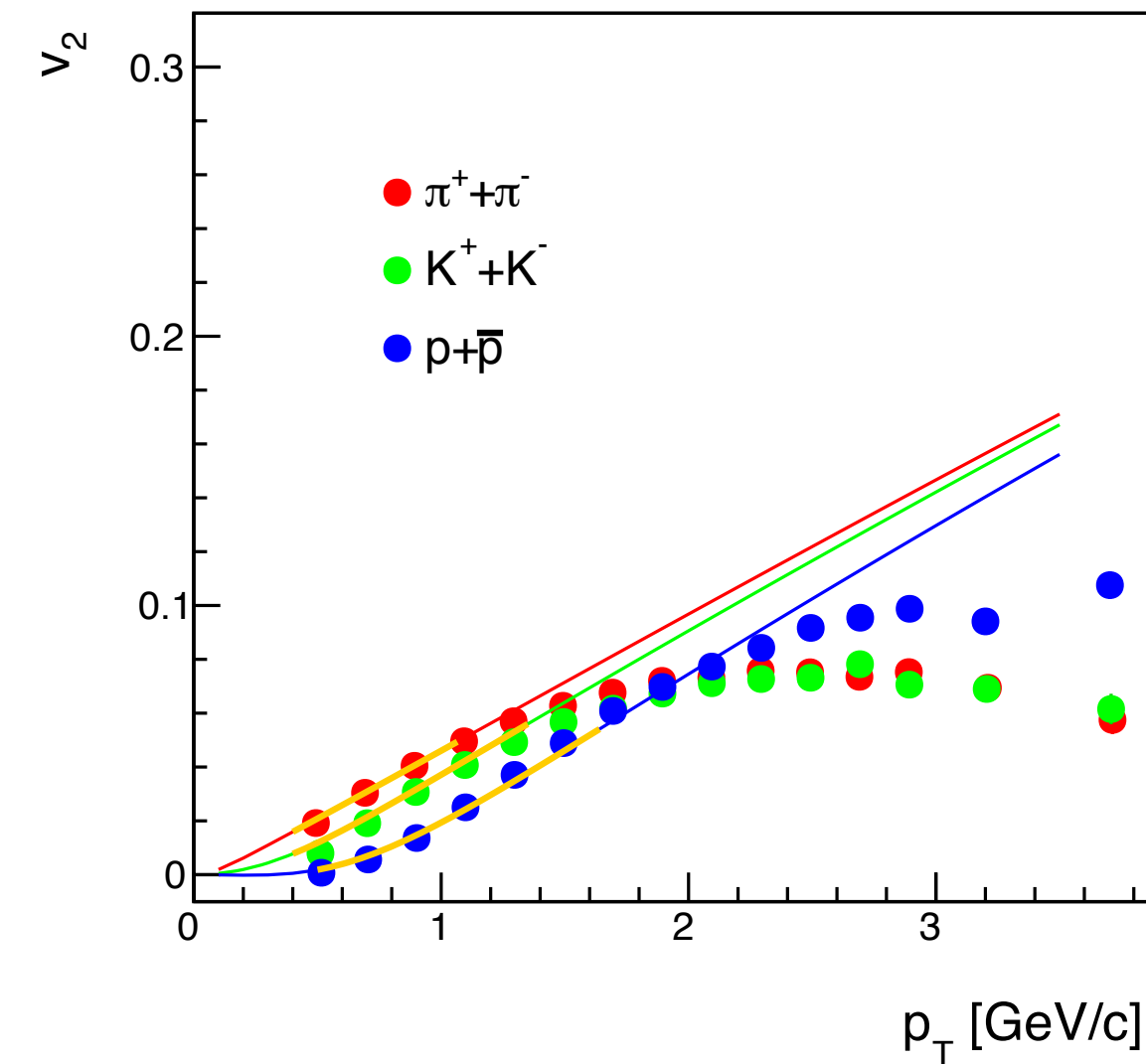
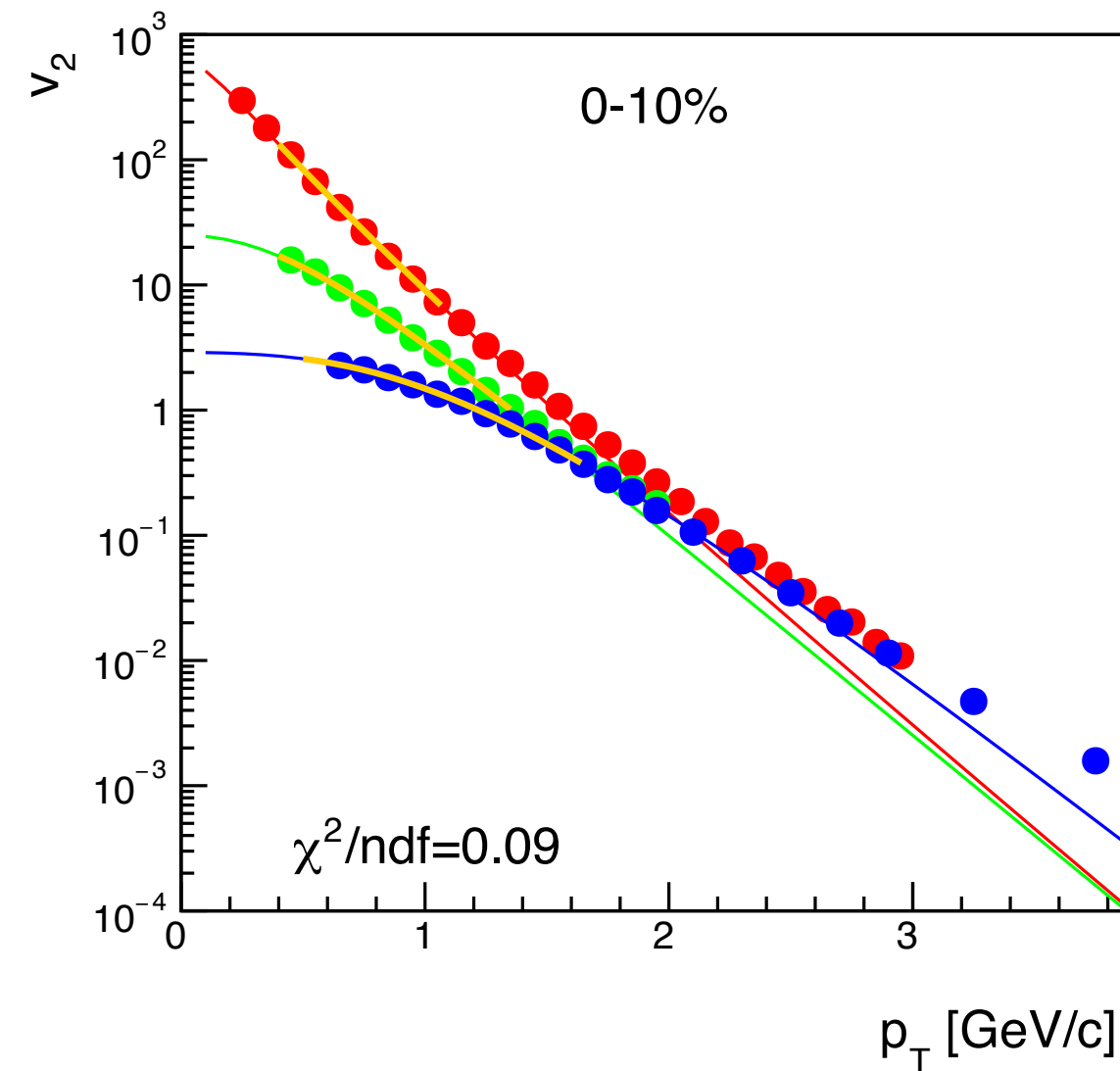


FIG. 2. Schematic illustration of an elliptical subshell of the source. Here, the source is extended out of the reaction plane ( $R_y > R_x$ ). Arrows represent the direction and magnitude of the flow boost. In this example,  $\rho_2 > 0$  [see Eq. (4)].

$\phi_s$ : azimuthal angle of the source element  
 $\phi_b$ : boost angle perpendicular to the elliptical subshell

# $\omega_z$ and $P_z$ from the BW model

e.g. Blast-wave fit to spectra and  $v_2$



Data:

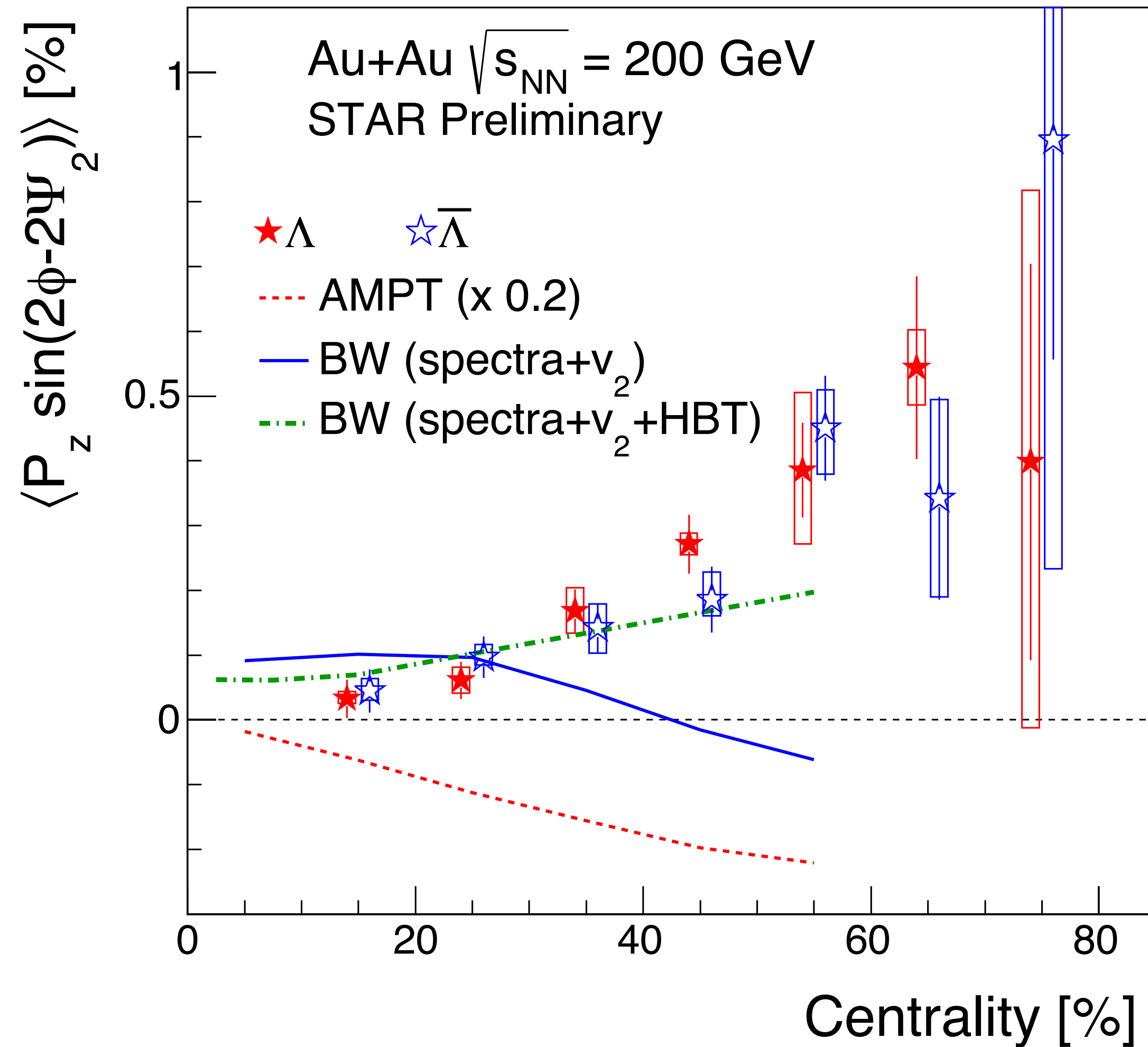
PHENIX, PRC69.034909 (2004)

PHENIX, PRC93.051902(R) (2016)

Calculated vorticity  $\omega_z$  shows the sine modulation. Assuming a local thermal equilibrium, z-component of polarization is estimated as follows:

$$P_z \approx \omega_z / (2T)$$

# $P_z$ modulation from the BW model



X. Xia, H. Li, Z. Tang, Q. Wang, PRC98.024905 (2018)

T. Niida, S. Voloshin, A. Dobrin, and R. Bertens,  
in preparation

BW parameters obtained with HBT: STAR, PRC71.044906 (2005)

# Summary

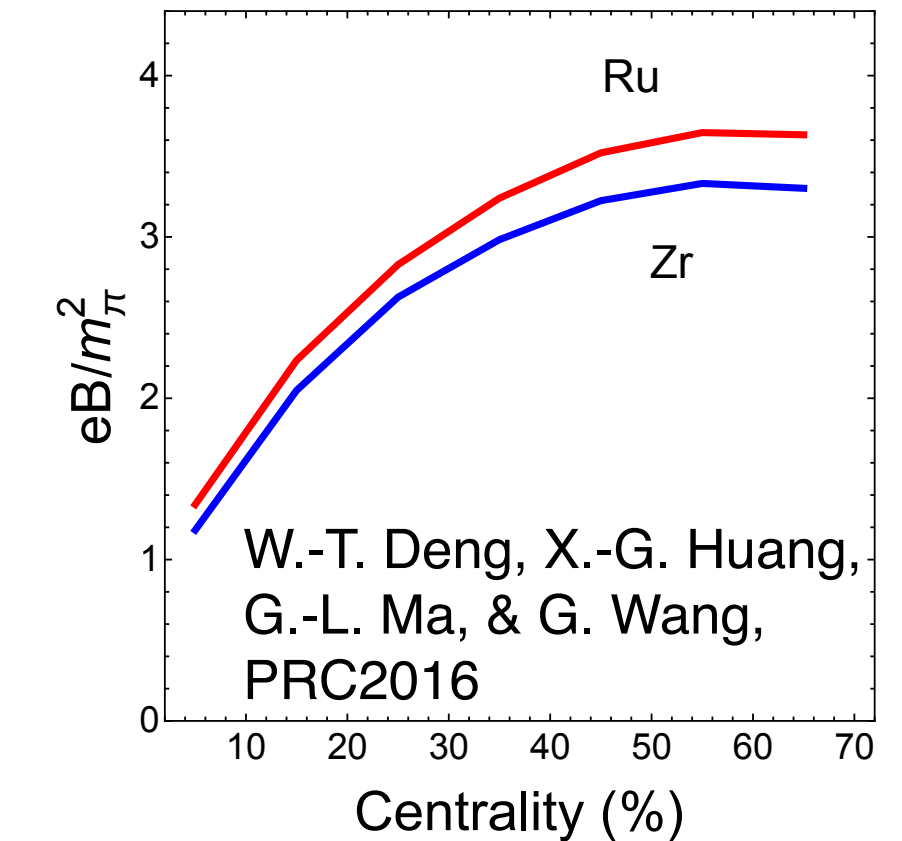
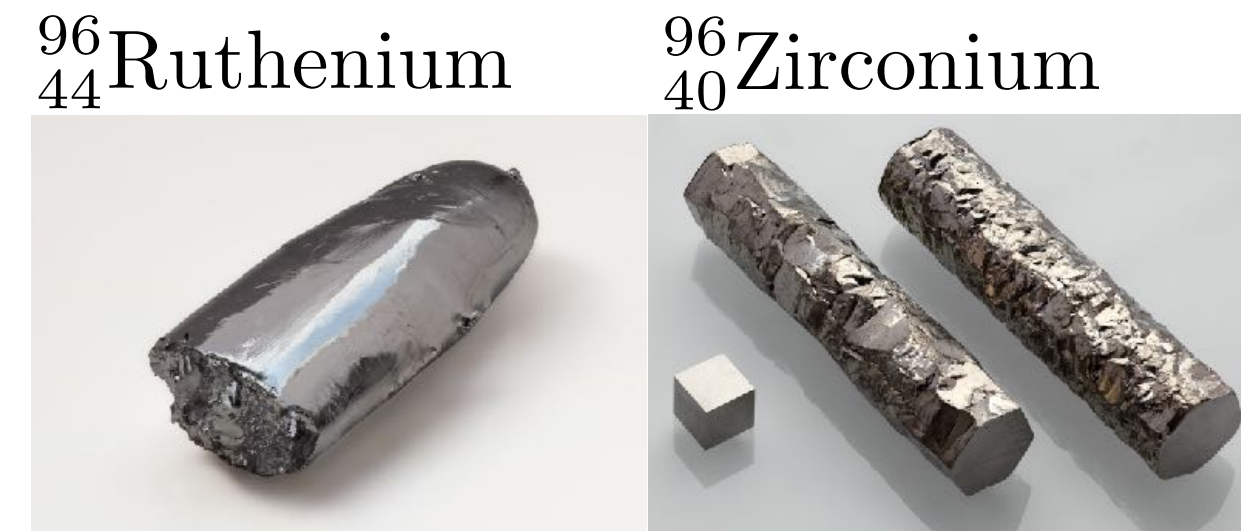
---

- Observation of positive  $\Lambda$  global polarization at  $\sqrt{s_{NN}} = 7.7-62.4$  GeV, and later at 200 GeV
  - Indicating the thermal vorticity of the system in HIC  $\omega \sim 10^{22} \text{ s}^{-1}$  ( $T=160$  MeV)
  - Polarization decreases at higher energies, and
  - Larger signal in more peripheral collisions but no significant dependence on  $p_T$   
→ **Quantitatively consistent with hydrodynamic and AMPT models**
  - Larger signal in in-plane than in out-of-plane  
→ **Disagree with hydrodynamic model**
  - Charge-asymmetry dependence ( $\sim 2\sigma$  level) in the polarization  
→ **A possible relation to the axial current induced by B-field**
  
- $\Lambda$  polarization along the beam direction at  $\sqrt{s_{NN}} = 200$  GeV
  - Quadrupole structure relative to the 2<sup>nd</sup>-order event plane, as expected from the elliptic flow  
→ **Qualitatively consistent with a picture of the elliptic flow but agree/disagree among the data and theoretical calculations in the sign**
  - Strong centrality dependence as in the elliptic flow
  - The blast-wave model predicts the same sign and similar magnitude to the data

# Outlook

□ Isobar collision data (Ru+Ru, Zr+Zr) already taken in 2018!

- Same mass number but different number of protons
  - 10% difference in the magnetic field
  - More  $P_H$  splitting btw  $\Lambda$  and anti- $\Lambda$  in Ru?



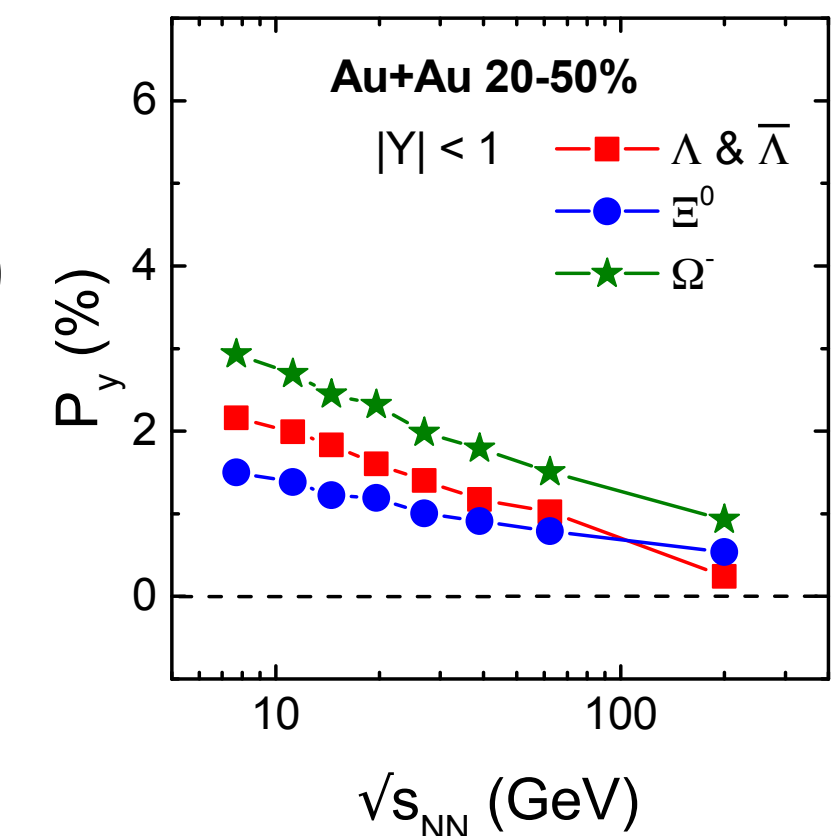
□ New 27 GeV data taken in 2018! (x10 events with  $\sim 1.5$  better EP resolution)

- Possible probe of the magnetic field from  $\Lambda$  vs anti- $\Lambda$  global polarization

□ Beam Energy Scan II (2019+) with STAR detector upgrade

- x10 events for  $\sqrt{s_{NN}} = 7.7-19.6$  GeV (collider mode) +  $\sqrt{s_{NN}} = 3-7.7$  GeV (Fixed target)
- How about at forward/backward rapidity? How about for multi-strangeness?

D.-X. Wei *et al.*, arXiv:1810.00151

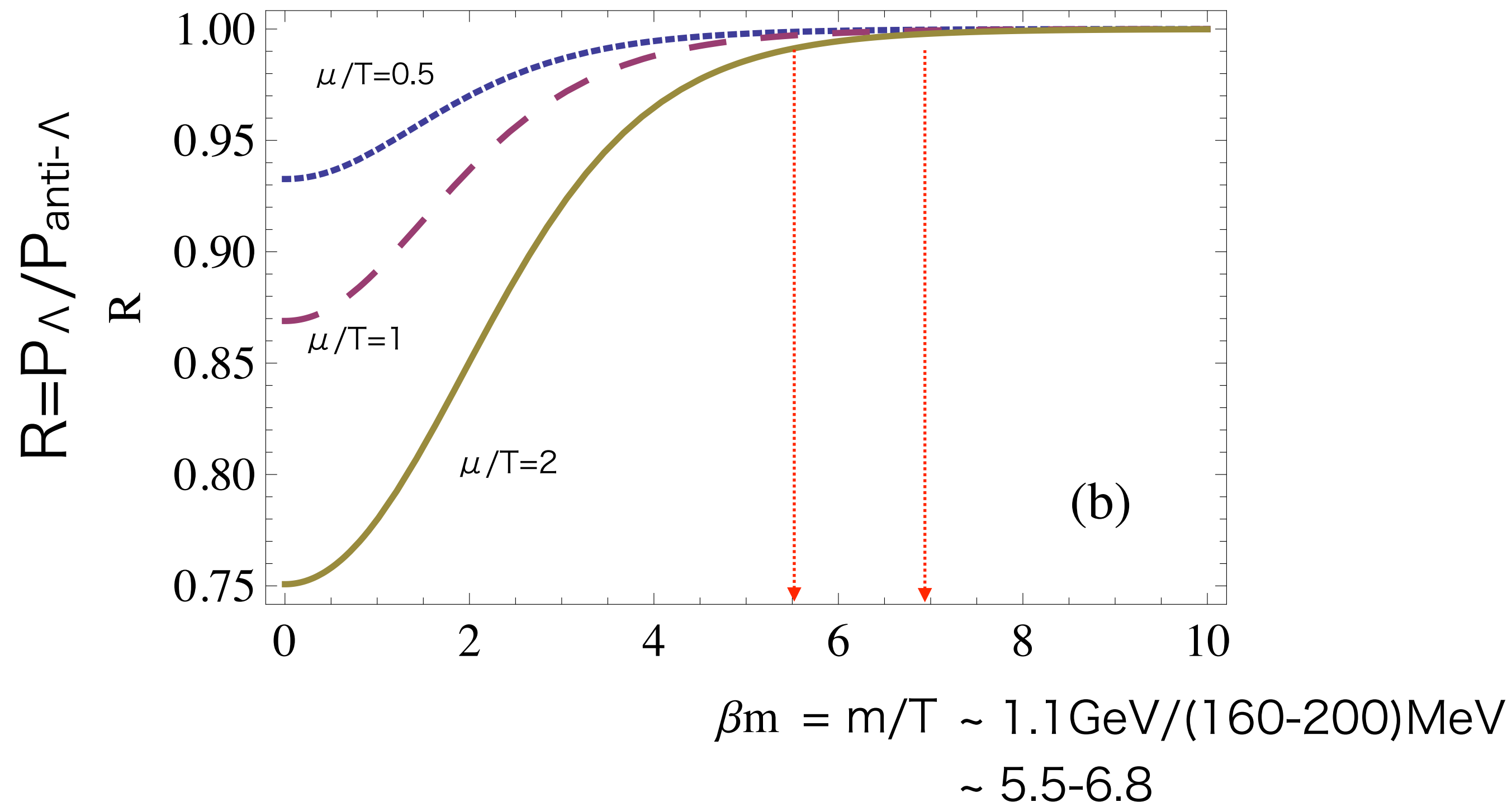


# ***Back up***

---

# Effect of non-zero chemical potential

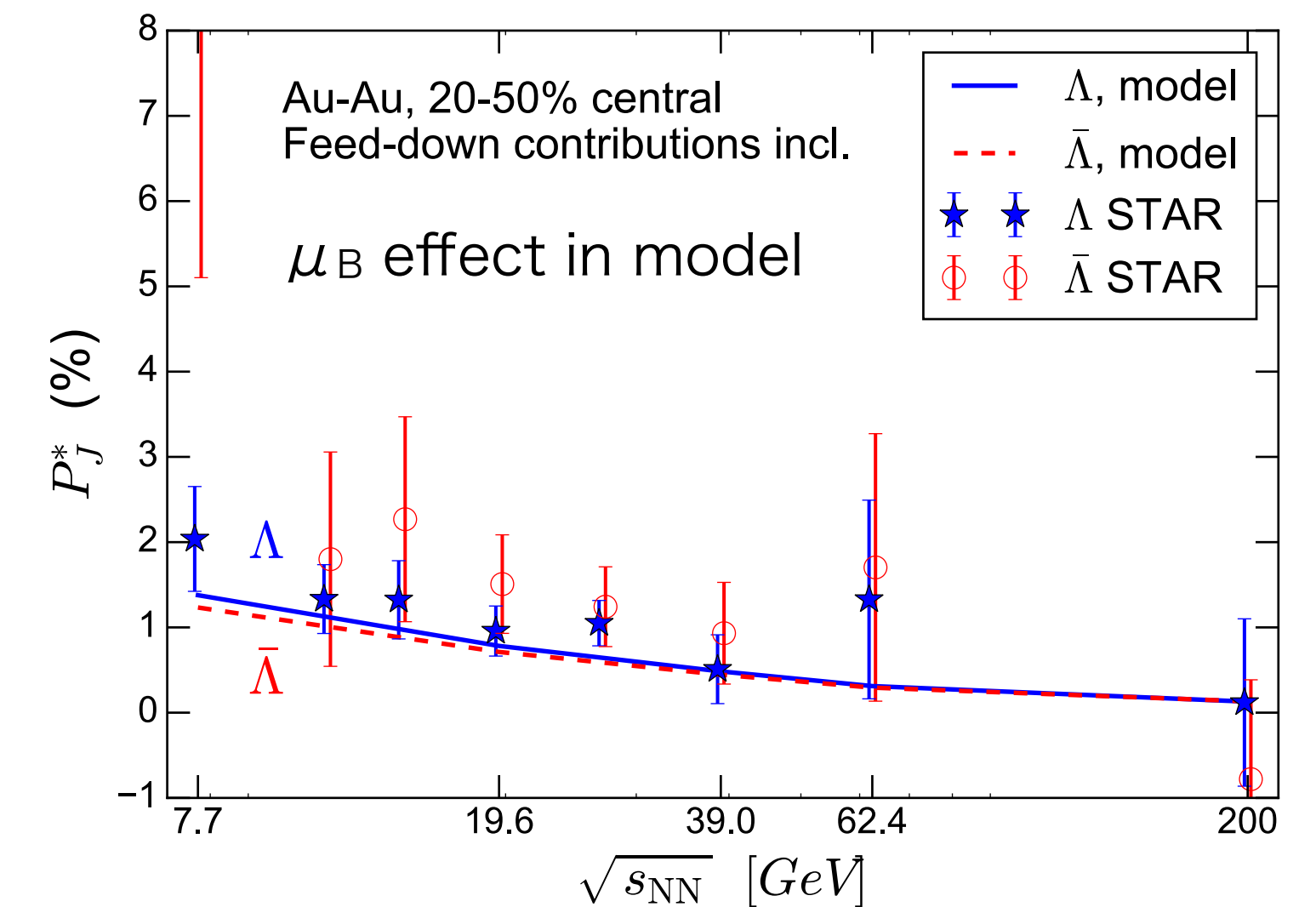
R. Fang, L. Pang, Q. Wang, and X. Wang, PRC94, 024904 (2016)



Non-zero chemical potential makes difference in polarization between  $\Lambda$  and anti- $\Lambda$ , but the effect seems to be small.

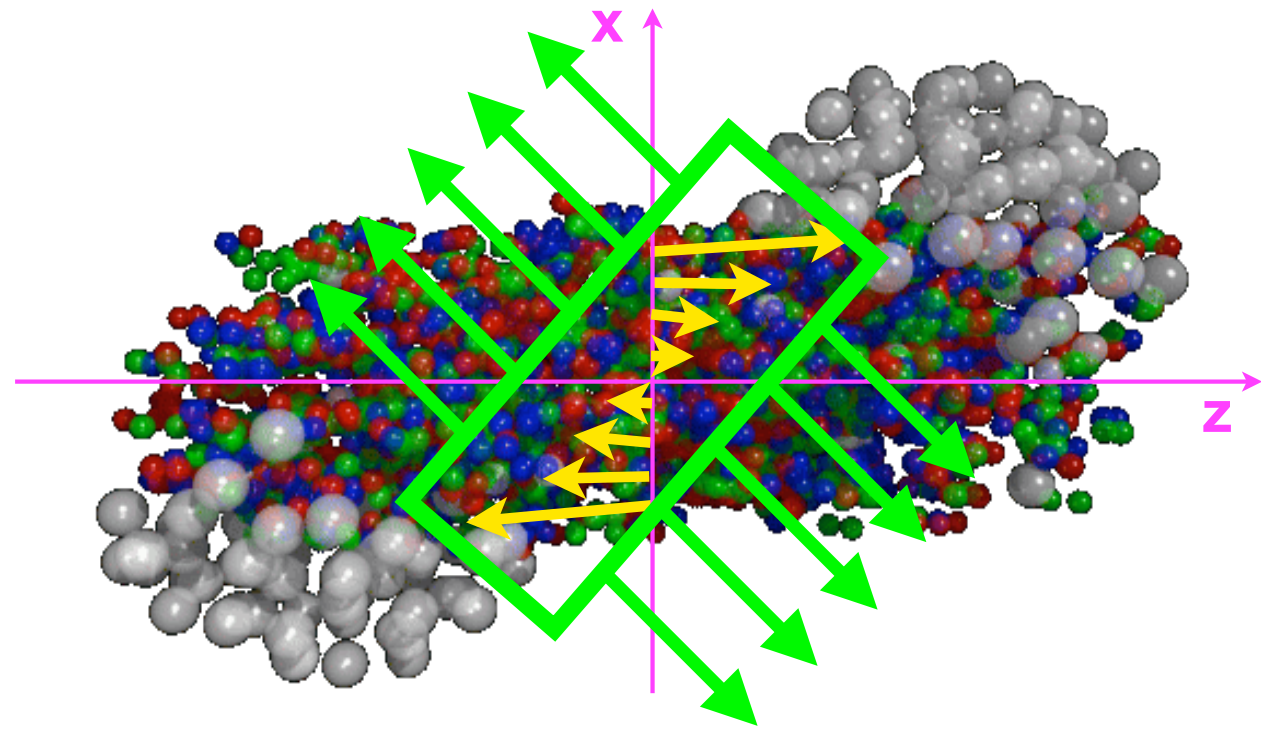
Y. Karpenko, sQM2017

$\Lambda$  and  $\bar{\Lambda}$ : UrQMD+vHLLÉ vs experiment





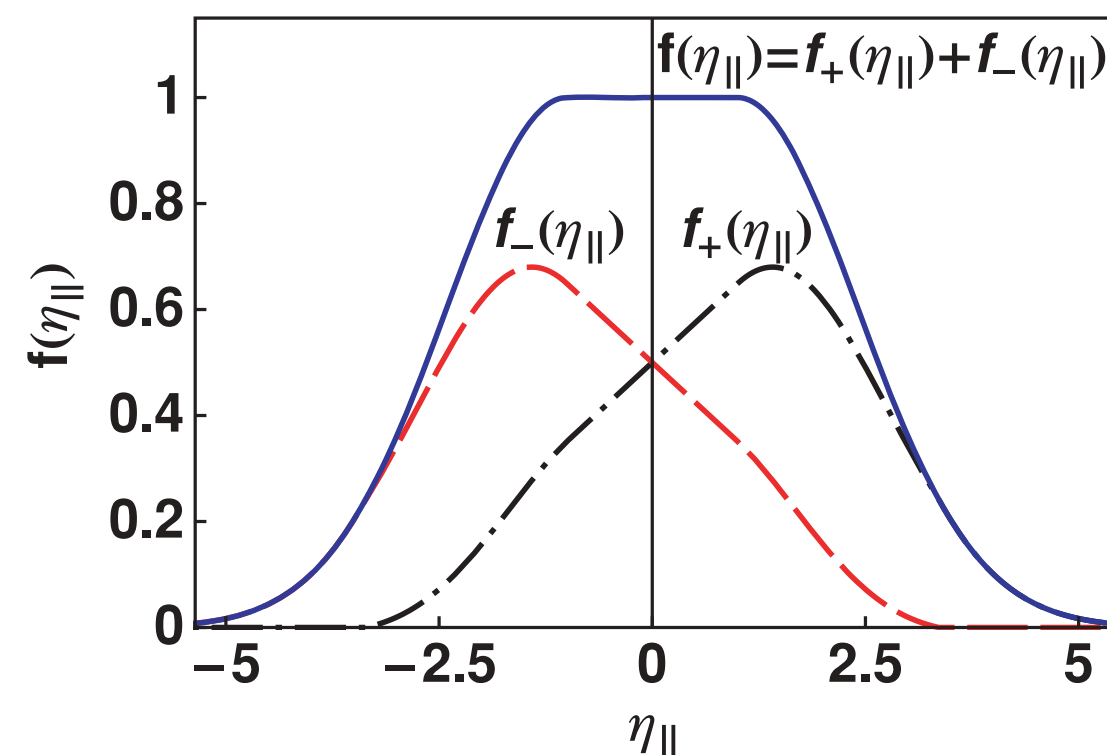
# What's the origin of $v_1$ ?



Vorticity is likely related to directed flow

Preferential emission of forward(backward) going participants results in the initial source tilt. The initial tilt with an expansion leads to a vorticity, and creates  $\eta$  dependence of directed flow.

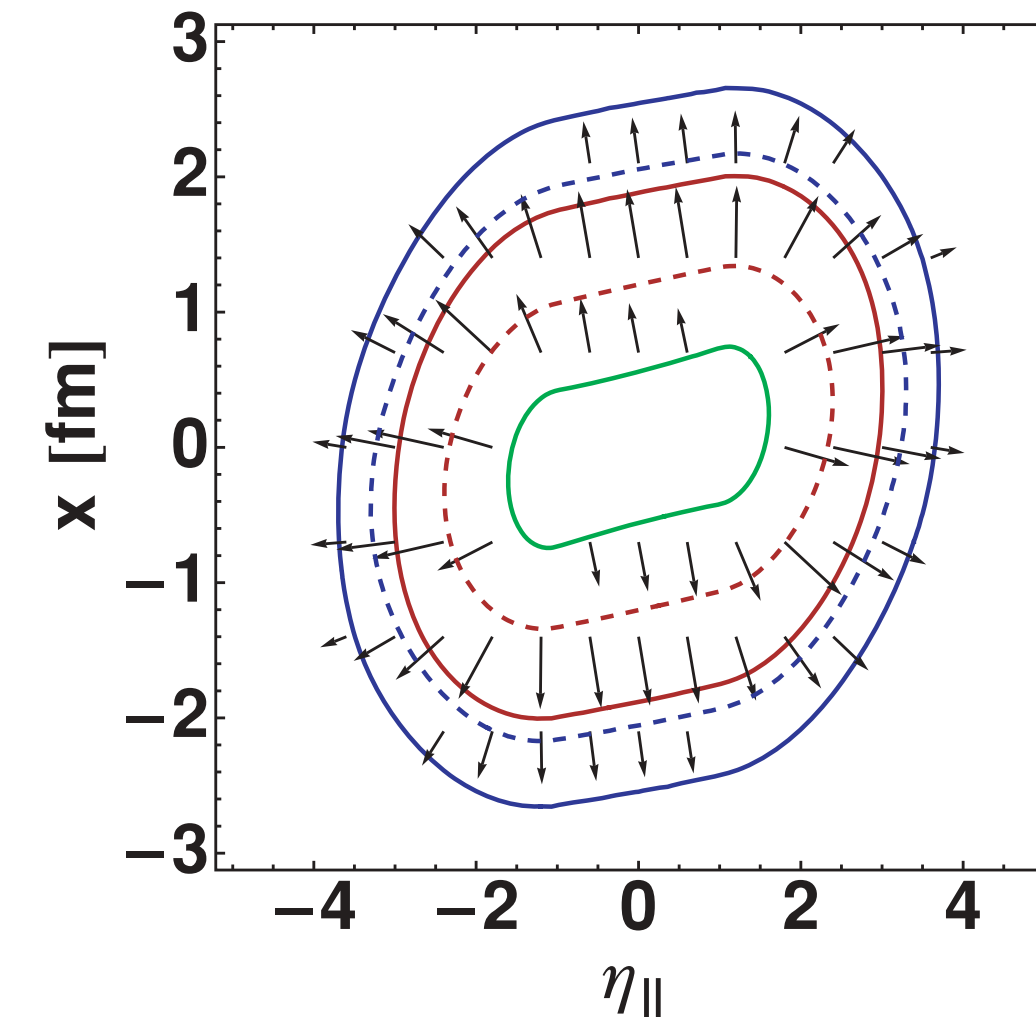
Bozek and Wyslciel, PRC81.054902 (2010)



$$f_+(\eta_{\parallel}) = f(\eta_{\parallel})f_F(\eta_{\parallel})$$

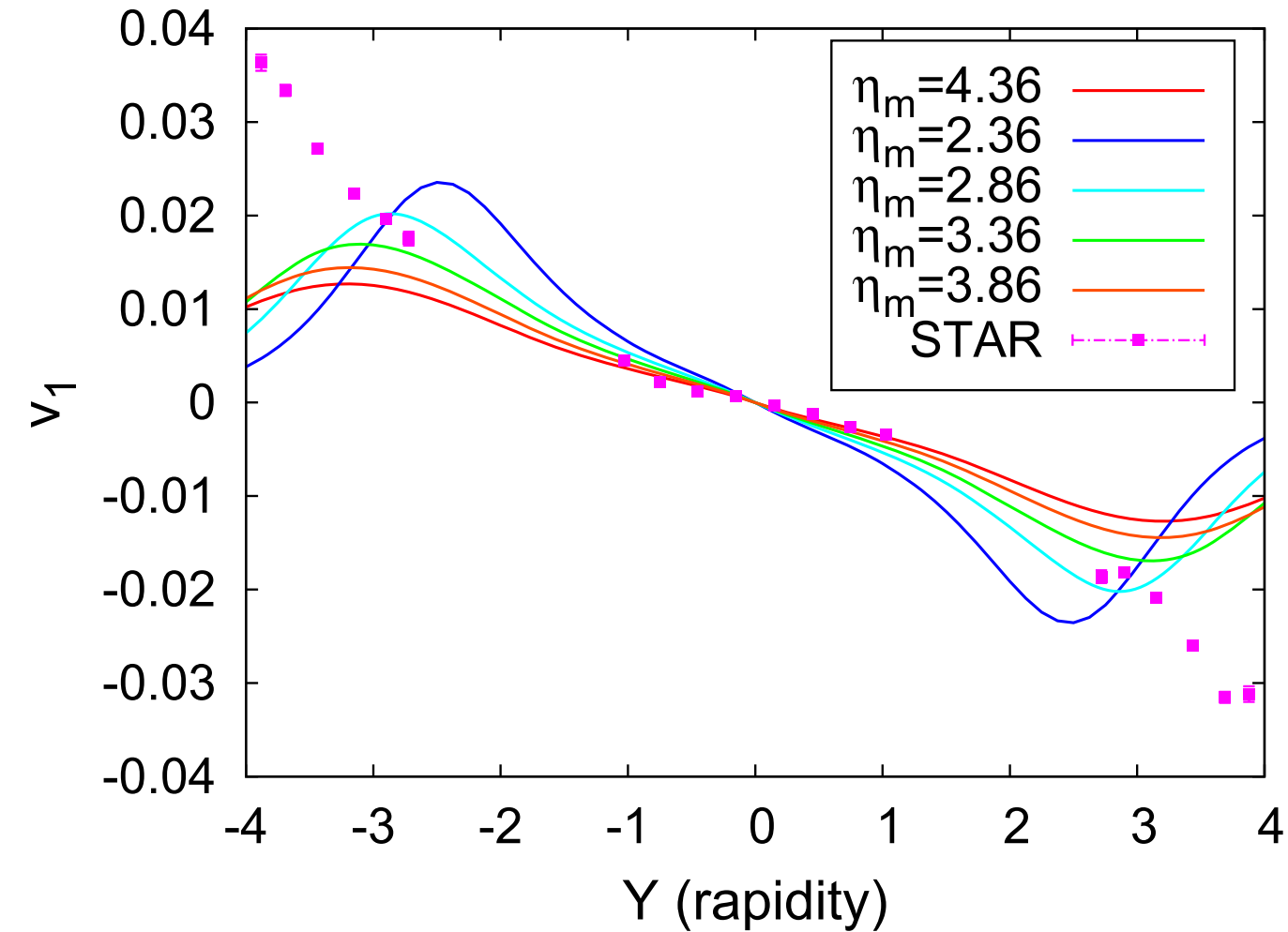
$$f_F(\eta_{\parallel}) = \begin{cases} 0 & \eta_{\parallel} < -\eta_m \\ \frac{\eta_{\parallel} + \eta_m}{2\eta_m} & -\eta_m \leq \eta_{\parallel} \leq \eta_m \\ 1 & \eta_m < \eta_{\parallel} \end{cases}$$

wounded nucleon model

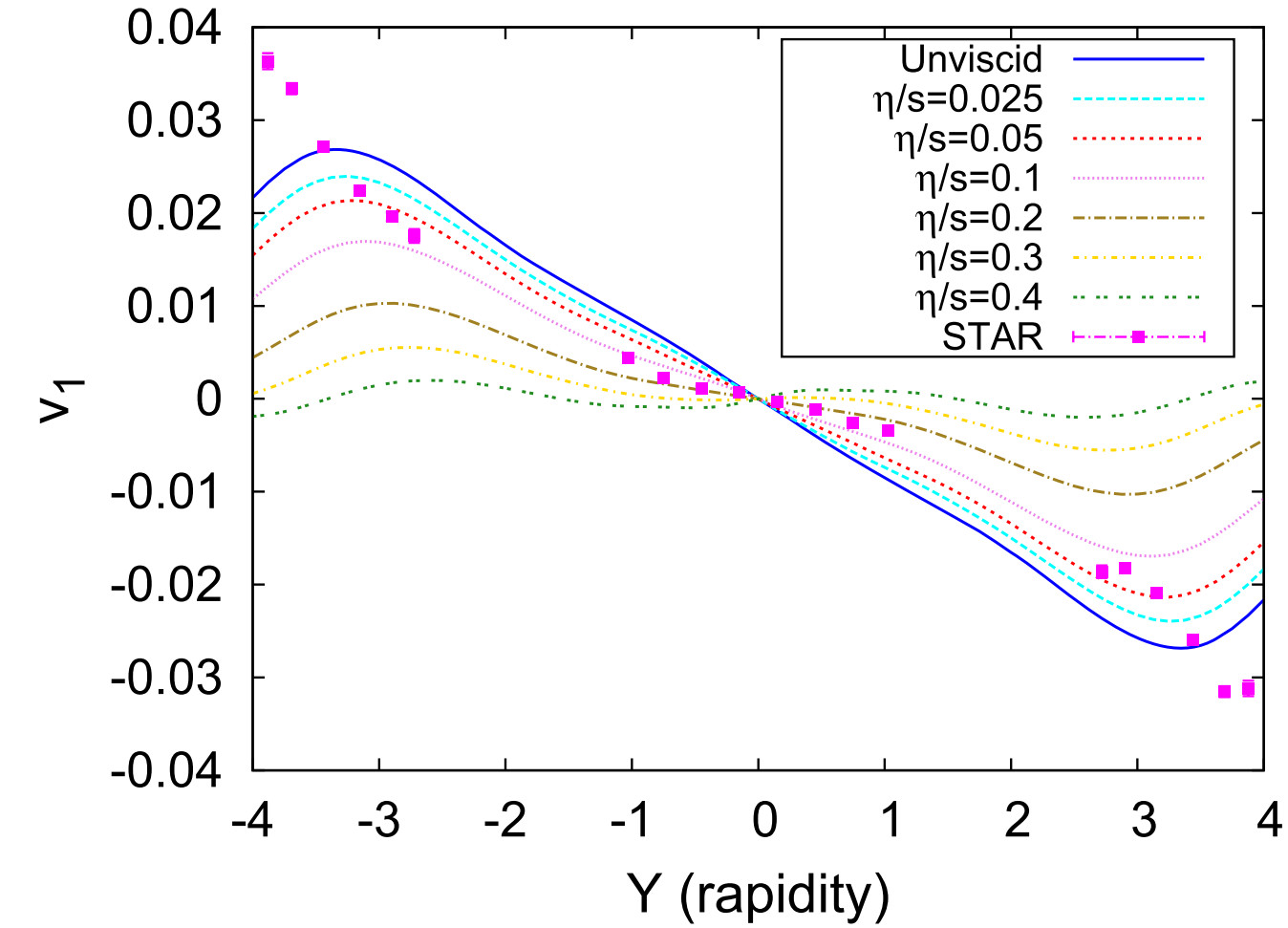


# Hydro with the “tilted” source

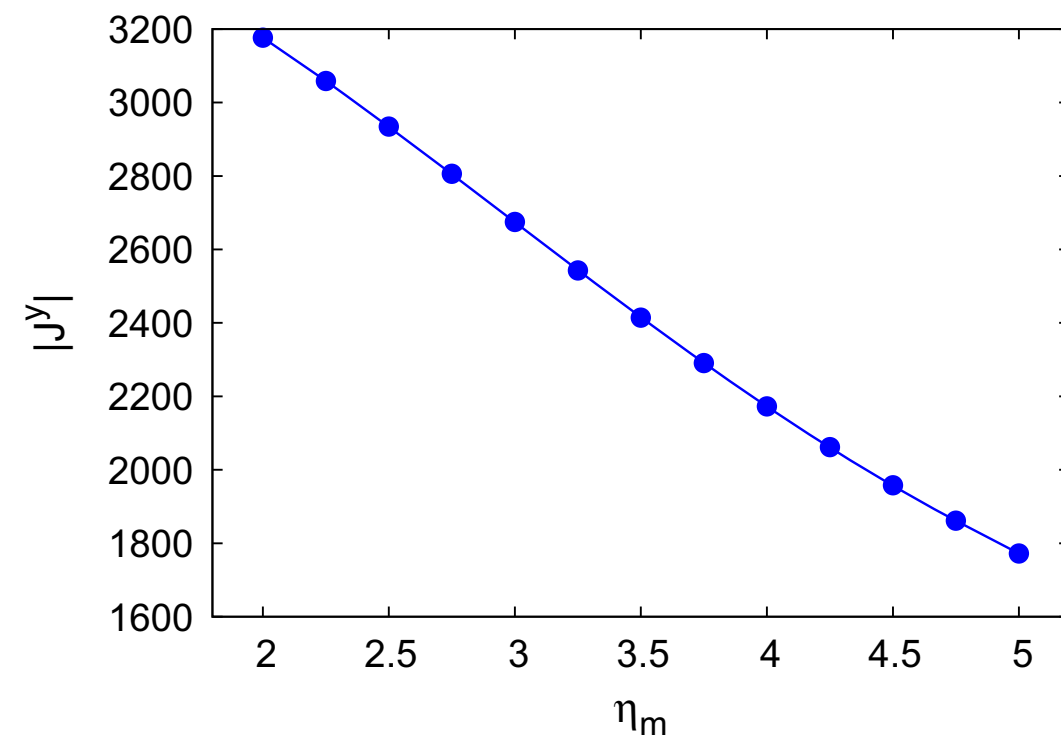
F. Becattini et al., Eur. Phys. J. C (2015)75:406



**Fig. 6** Directed flow of pions for different values of  $\eta_m$  parameter with  $\eta/s = 0.1$  compared with STAR data [22]



**Fig. 7** Directed flow of pions for different values of  $\eta/s$  with  $\eta_m = 2.0$  compared with STAR data [22]



**Fig. 9** Angular momentum (in  $\hbar$  units) of the plasma with Bjorken initial conditions as a function of the parameter  $\eta_m$

The tilted source which accounts for vorticity provides a better description of  $v_1$ !

# Chiral Magnetic Effect (CME)

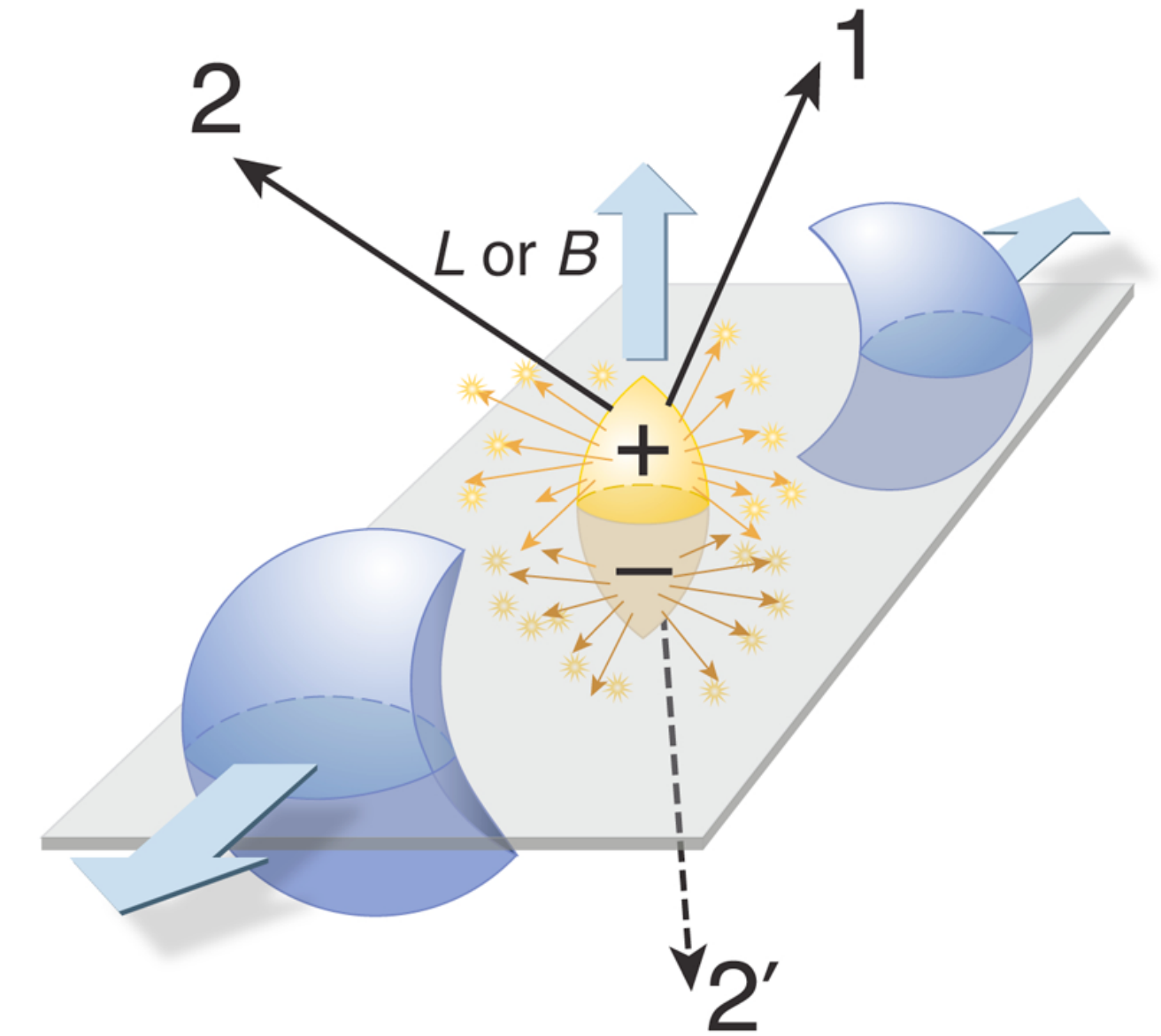
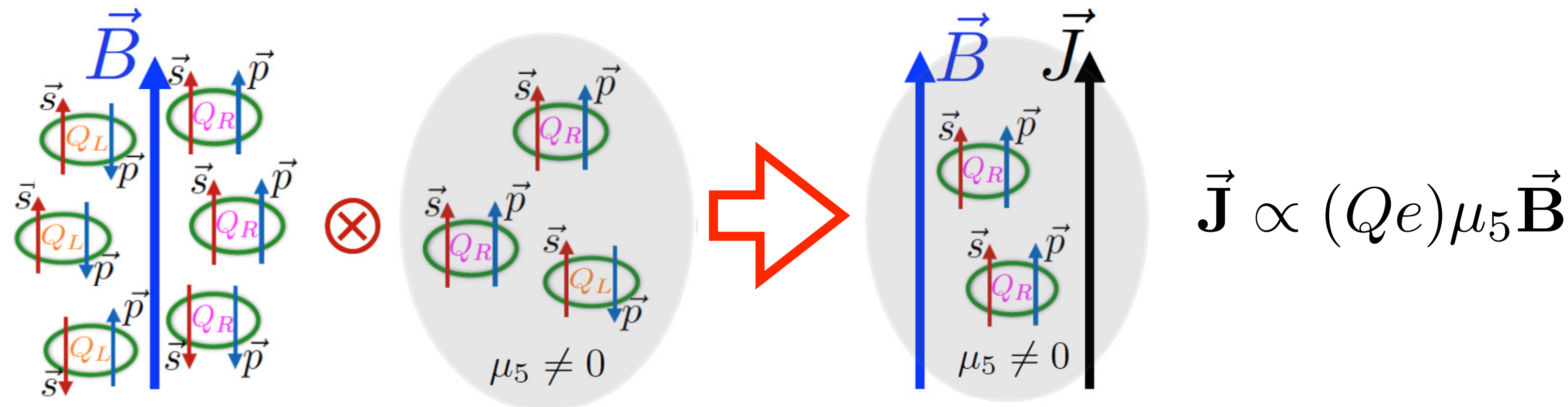
D. Kharzeev, R. Pisarski, M. Tytgat, PRL81, 512 (1998)  
D. Kharzeev, PPNP75(2014)133-151

Magnetic field + massless quarks + chirality imbalance

*spin alignment  
(opposite direction  
for opposite sign)*

*spin and momentum in (anti-)parallel  
for right(left)-handed quarks*

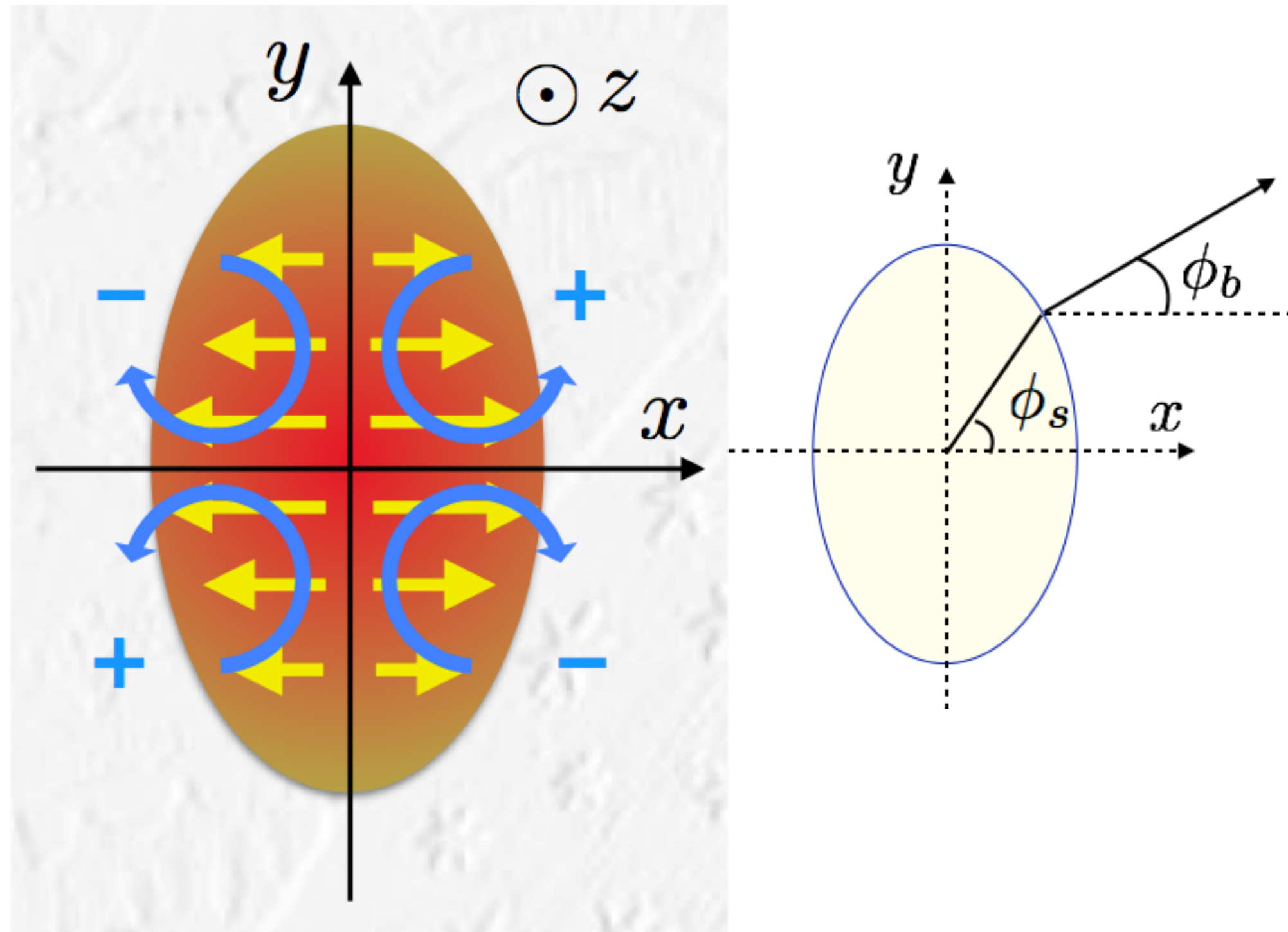
*right-handed quarks  $\neq$  left-handed quarks*



Induction of electric current along the magnetic field,  
called *Chiral Magnetic Effect (CME)*

# Blast-wave parameterization

S. Voloshin, arXiv:1710.08934



$$r_{max} = R[1 - a \cos(2\phi_s)],$$

$$\rho_t = \rho_{t,max}[r/r_{max}(\phi_s)][1 + b \cos(2\phi_s)] \approx \rho_{t,max}(r/R)[1 + (a + b) \cos(2\phi_s)].$$

$$\omega_z = 1/2(\nabla \times \mathbf{v})_z \approx (\rho_{t,nmax}/R) \sin(n\phi_s)[b_n - a_n].$$

$a_n$ : spatial anisotropy

$b_n$ : flow anisotropy

$R$ : reference source radius

$\rho_t$ : transverse flow velocity

Quadrupole or sine structure of  $\omega_z$  is expected.

# ***Systematic uncertainties***

---

## Case of 200 GeV as an example

- Event plane determination: ~22%
- Methods to extract the polarization signal: ~21%
- Possible contribution from the background: ~13%
- Topological cuts: <3%
- Uncertainties of the decay parameter: ~2% for  $\Lambda$ , ~9.6% for anti- $\Lambda$
- Extraction of  $\Lambda$  yield (BG estimate): <1%

Also, the following studies were done to check if there is no experimental effect:

- Two different polarities of the magnetic field for TPC
- Acceptance effect
- Different time period during the data taking
- Efficiency effect