

2016年11月15-17日

東京大学大学院理学系研究科物理学専攻

物理学特別講義 B XVI

宇宙大規模構造と観測的宇宙論

Observational cosmology with  
large-scale structure

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# Overview

# Large-scale structure

Spatial matter inhomogeneities over Mpc  $\sim 10^3$  Mpc

Mpc =  $10^6$  parsec  $\sim 3 \cdot 10^6$  light years  
(c.f., 40kpc for size of Milky Way)

- Hierarchical clustering of matter distribution:

galaxy  $\subset$  group / cluster  $\subset$  supercluster

- Contain rich cosmological information

primordial fluctuations,

structure formation

dynamics of cosmic expansion

- Traditionally traced by galaxy redshift surveys

(other LSS probes are gravitational lensing, Lyman-alpha forest)

# Observing large-scale structure

Intensive use of telescope is necessary



8.2m

Very Large Telescope (Chile)



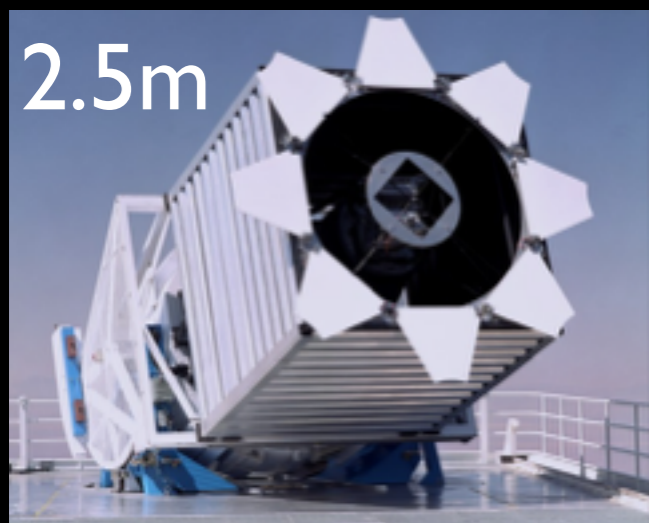
8.2m

Subaru Telescope (Hawaii)



3.6m

Canada-France-Hawaii  
Telescope (Hawaii)



2.5m

Sloan Digital Sky Survey  
@ APO (New Mexico)

Blanco telescope  
@ CTIO (Chile)



4m

[https://en.wikipedia.org/wiki/Very\\_Large\\_Telescope](https://en.wikipedia.org/wiki/Very_Large_Telescope)  
<http://www.sdss.org/instruments/>  
<http://subarutelescope.org/Information/Download/DImage/index.html>  
<http://www.cfht.hawaii.edu/en/news/CFHT30/#wallpaper>  
<http://www.darkenergysurvey.org/DECam/index.shtml>

# Redshift

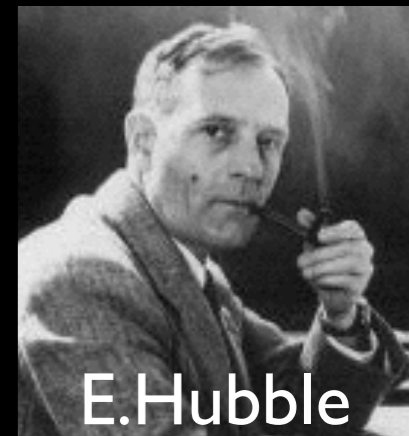
A key measurement to probe 3D view of large-scale structure

Distant galaxies looks **redder** than nearby galaxies  
due to **cosmic expansion**

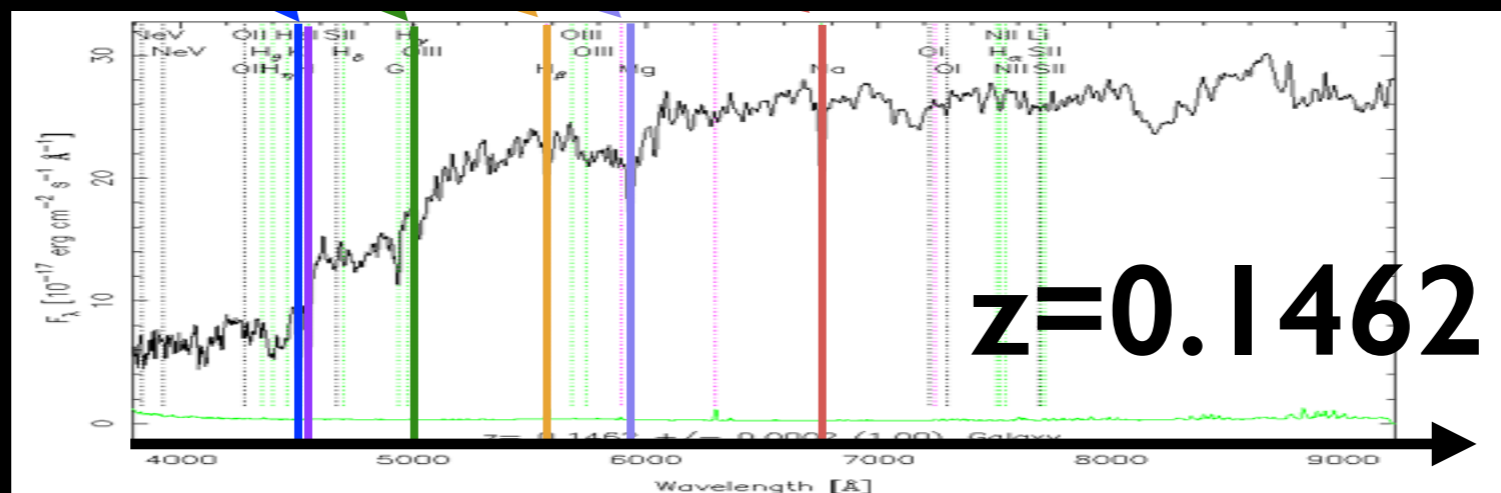
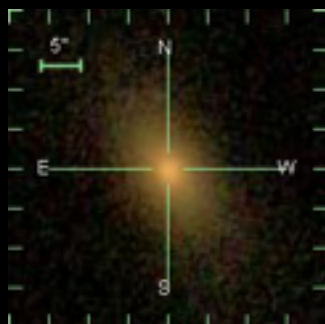
Redshift parameter  $z = \Delta\lambda / \lambda$

Hubble law

recession 'velocity'  $v = H d$  distance to galaxy  
(=  $c z$ ) Hubble parameter



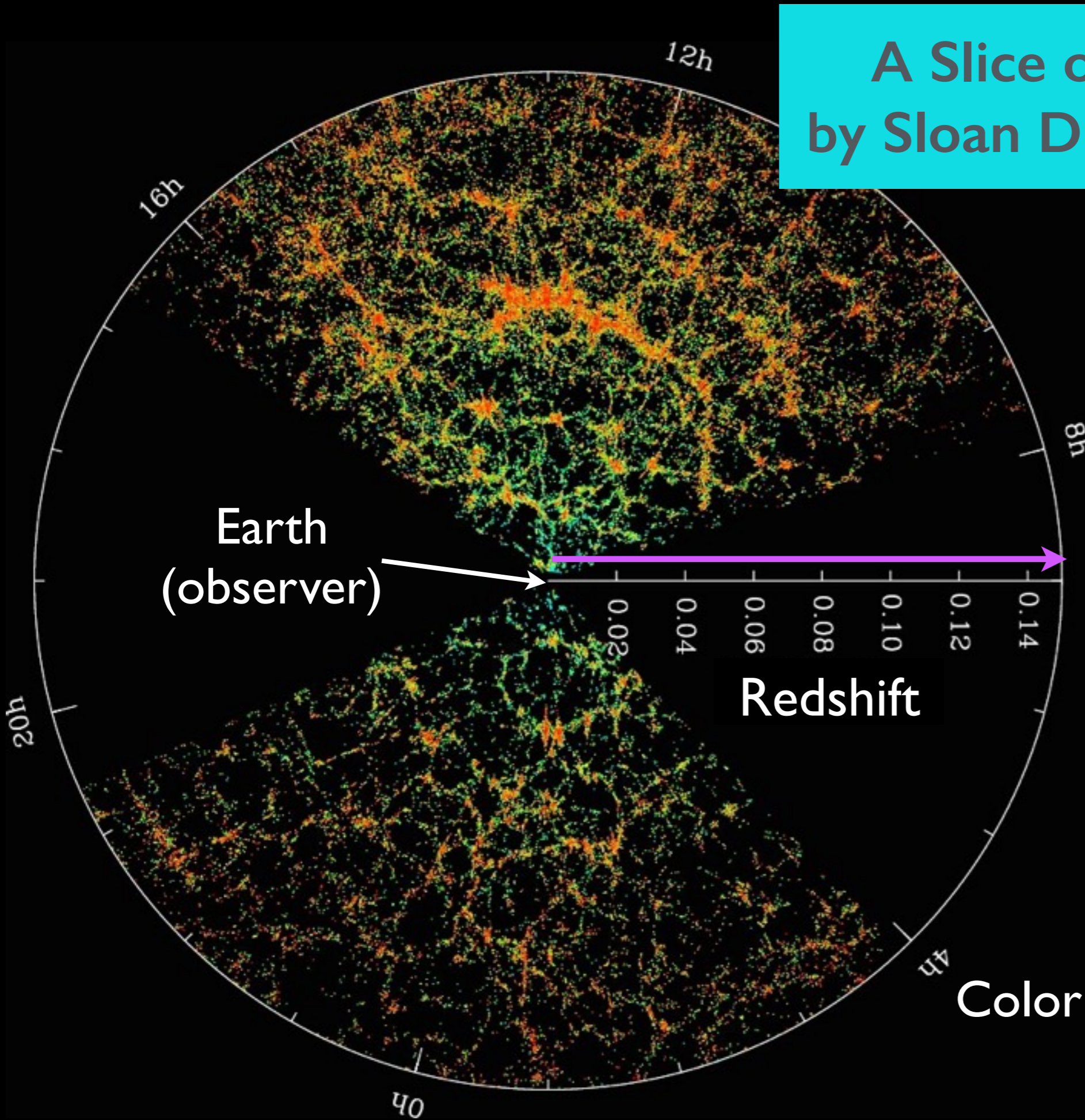
Distant galaxy



SDSS SkyServer

# A Slice of galaxy catalog by Sloan Digital Sky Survey II

finished in 2008



2 G yrs  
(look back time)

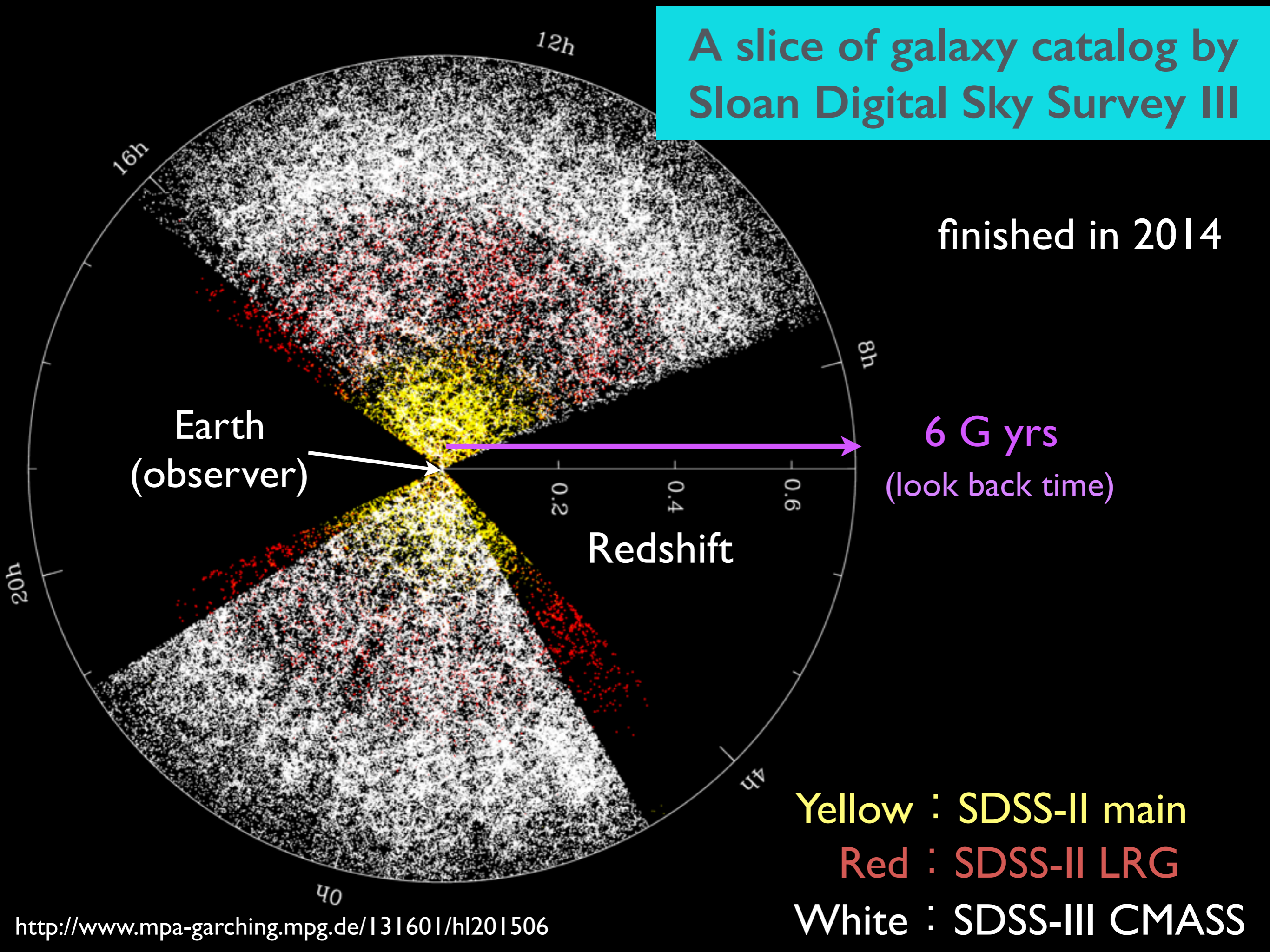
Color indicates age of galaxy

Blue : young

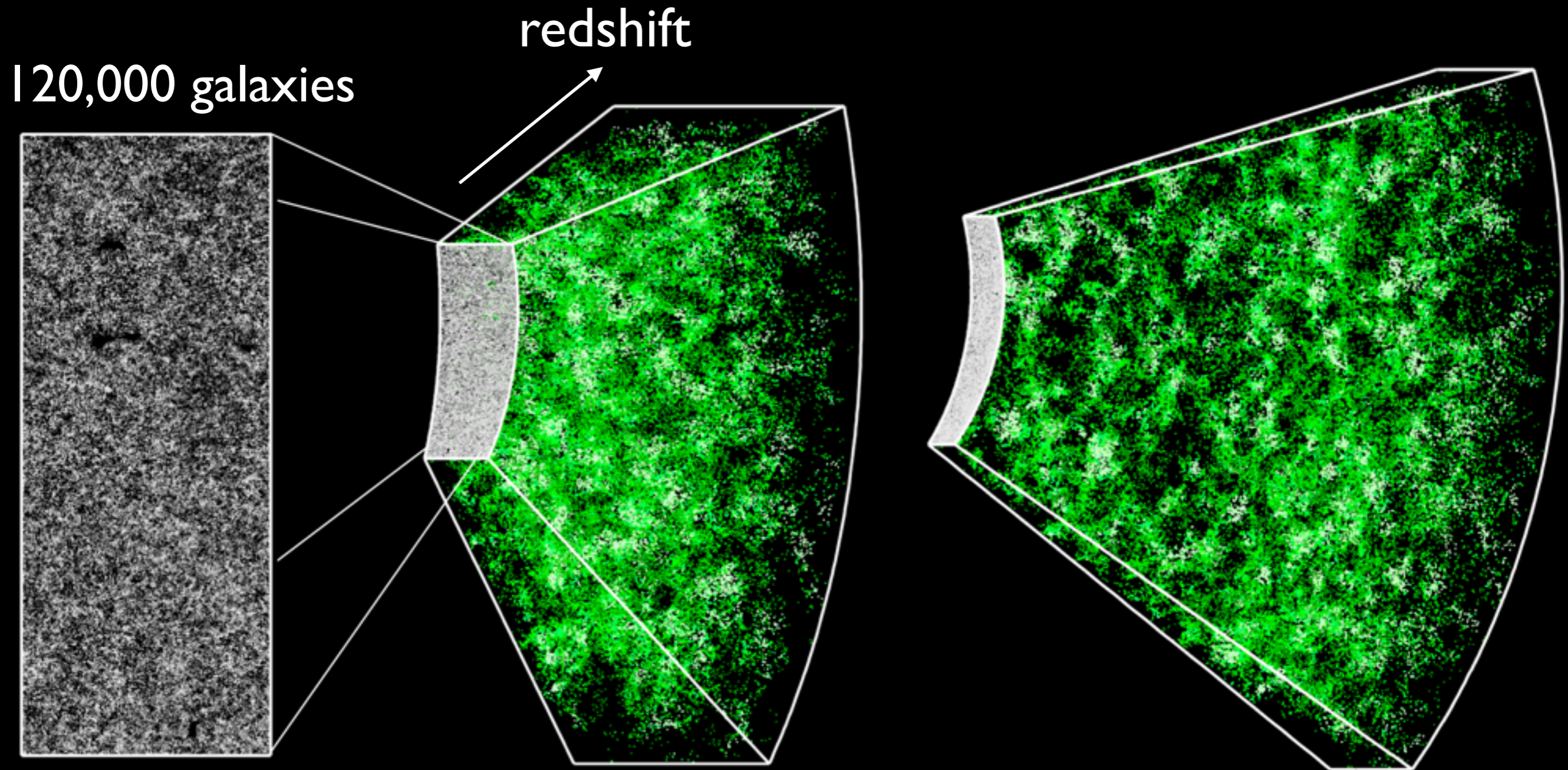
Red : old

# A slice of galaxy catalog by Sloan Digital Sky Survey III

finished in 2014



# A section of 3D map





# Cosmology with galaxy 3D map

## Statistical properties

- Initial conditions for primordial fluctuations (cosmic inflation)
- Growth of structure
- Matter contents of the Universe

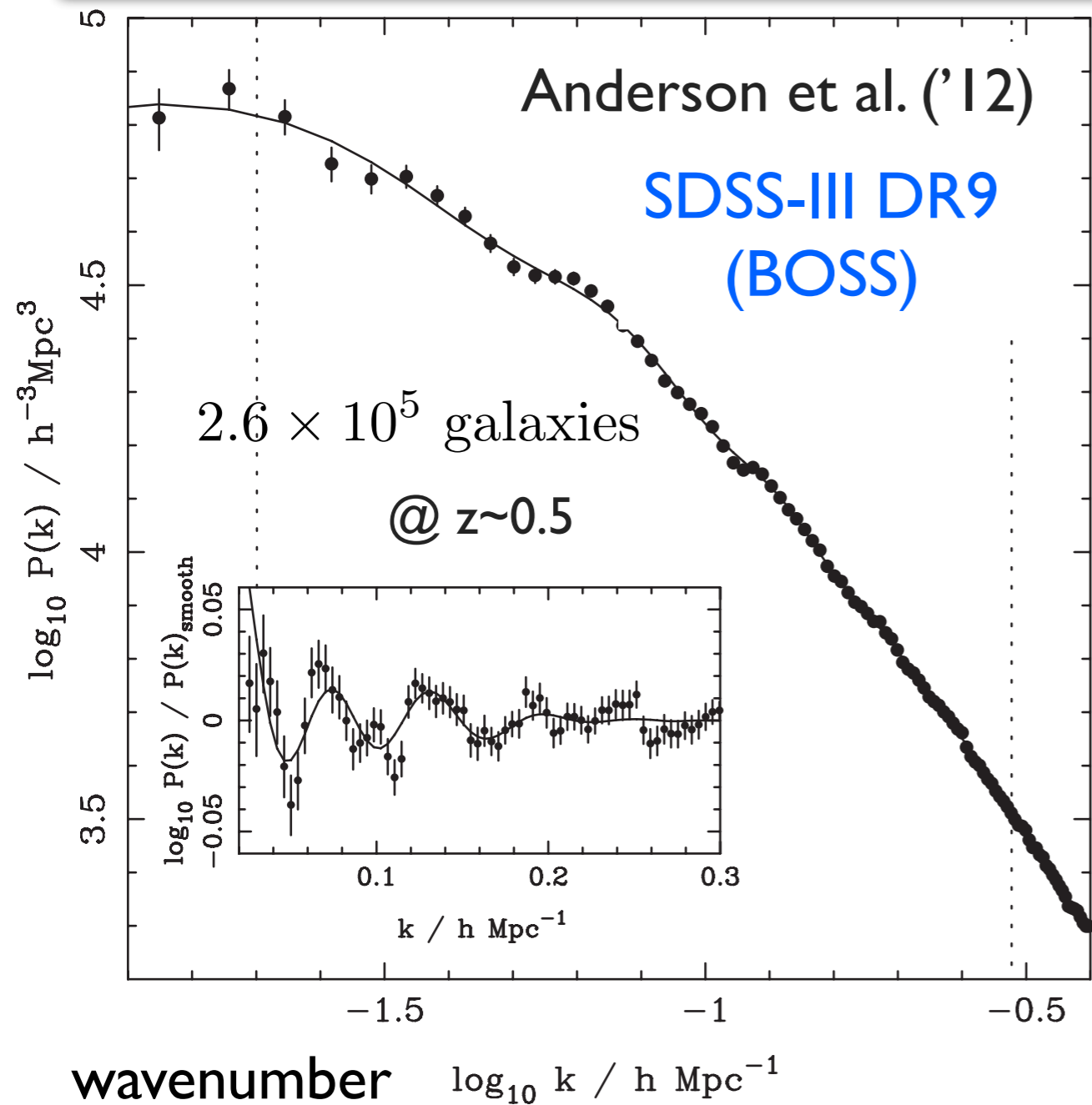
## Gauging the scales of large-scale structure



# Power spectrum of matter fluctuations

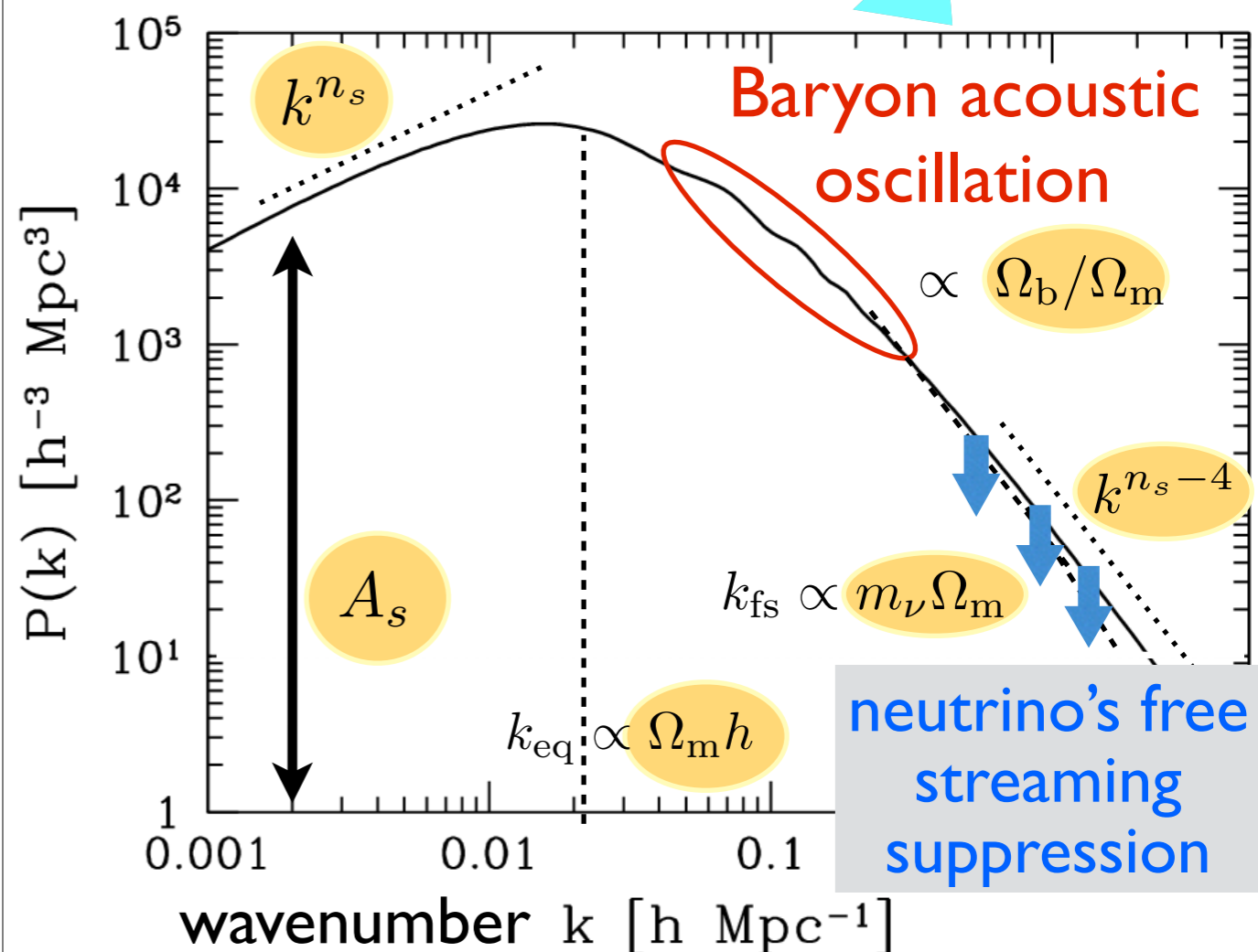
$$\delta(\vec{x}) \equiv \frac{\delta\rho_m(\vec{x})}{\bar{\rho}_m} = \frac{1}{\sqrt{V}} \sum_{\vec{k}} \delta(\vec{k}) e^{i\vec{k}\cdot\vec{x}}$$

$$P(k) = \frac{1}{N_k} \sum_{|\vec{k}|=k} |\delta(\vec{k})|^2$$



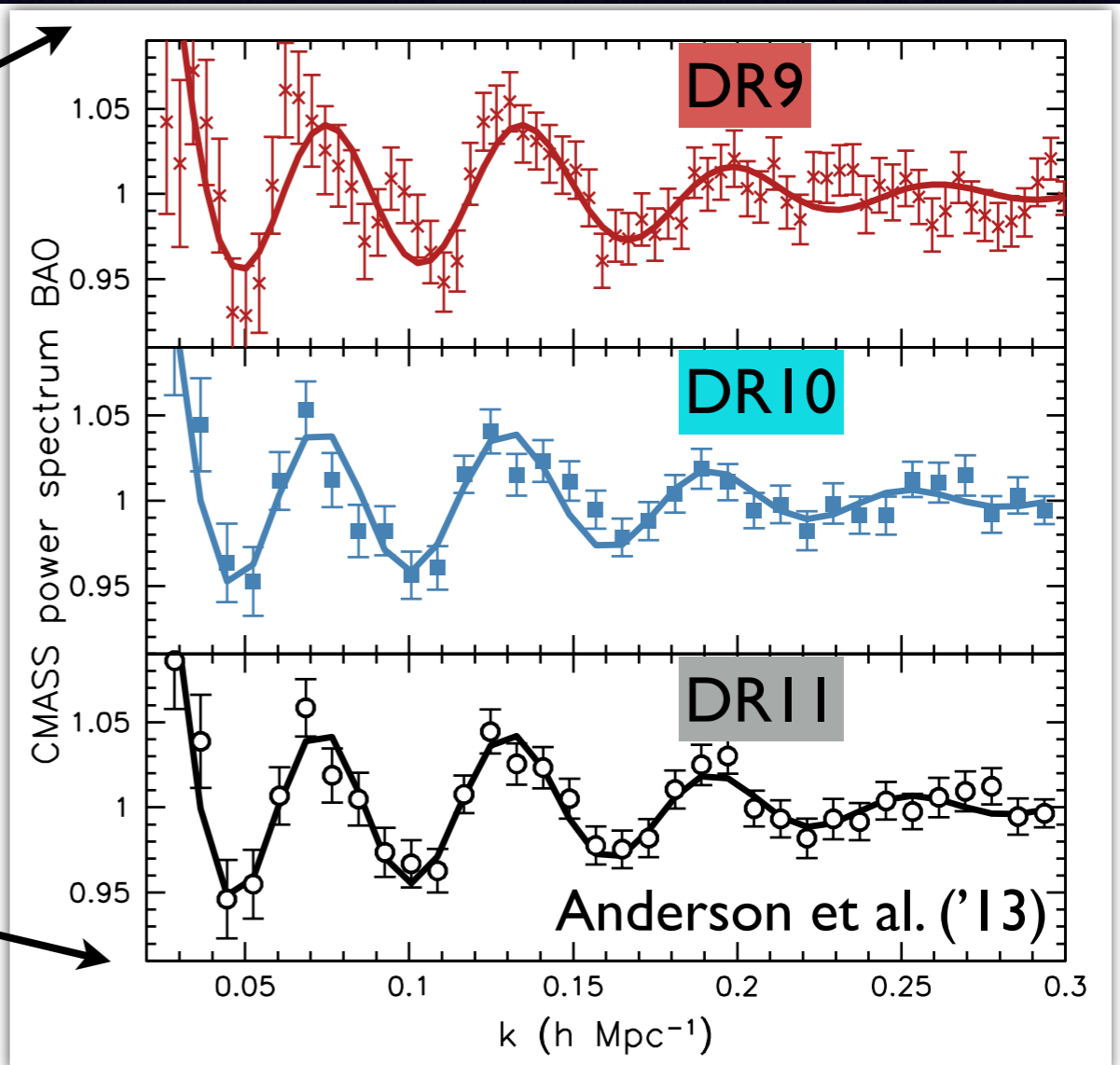
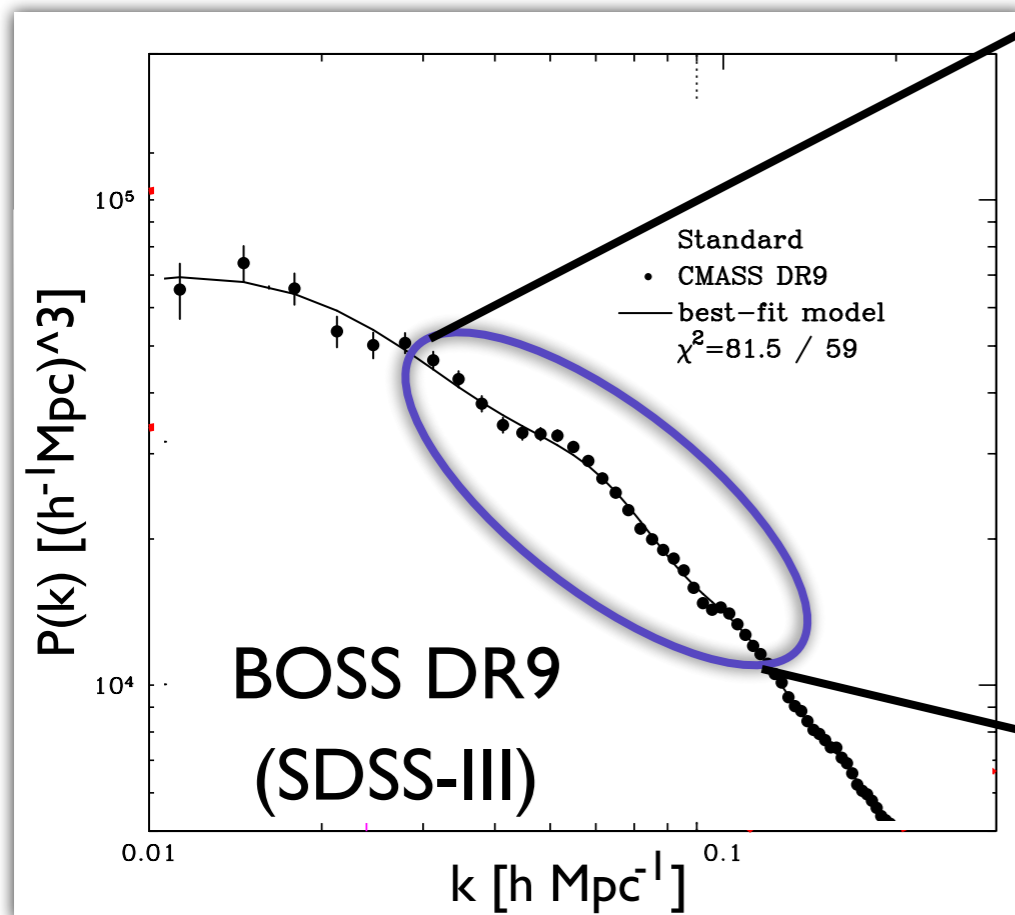
Observation

Linear theory



# Baryon acoustic oscillations (BAO)

- Characteristic scale of primeval baryon-photon fluid ( $\sim 150\text{Mpc}$ )  
( $\Leftrightarrow$  acoustic signal in CMB anisotropies)
- Can be used as standard ruler to measure cosmic expansion  
(theoretical prior)



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Angular diameter distance

$$D_A = \frac{r_s}{\theta}$$



$\theta$

Redshift  $z (=0\sim 3)$

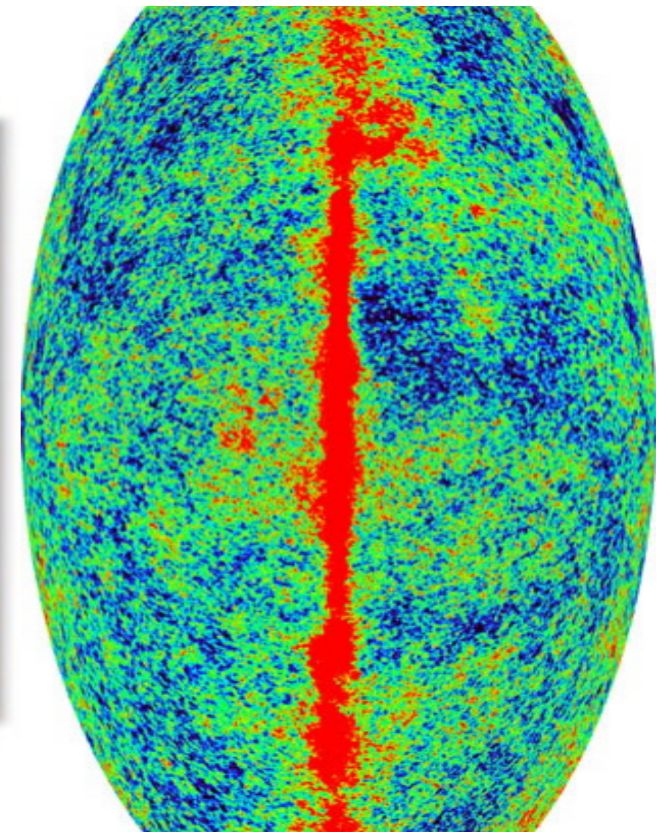
$r_s$



Acoustic oscillation scales

$r_s$

Redshift  $z=1100$



cosmic expansion

changes the distances !!

distant galaxies

cosmic microwave background

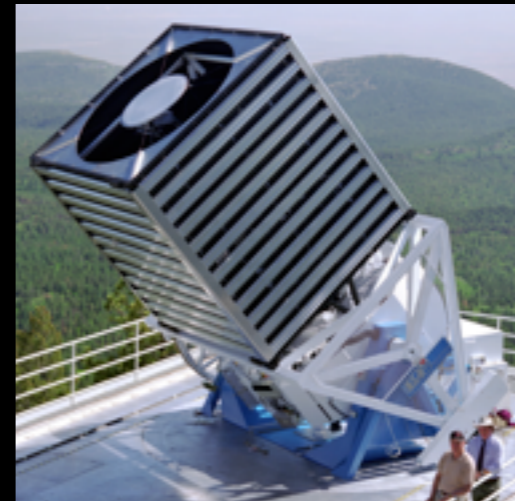
# Upcoming/on-going projects

Multi-purpose ground- & space-based experiments

DES (2013~)



HETDEX (2016+)



eBOSS (2014~)

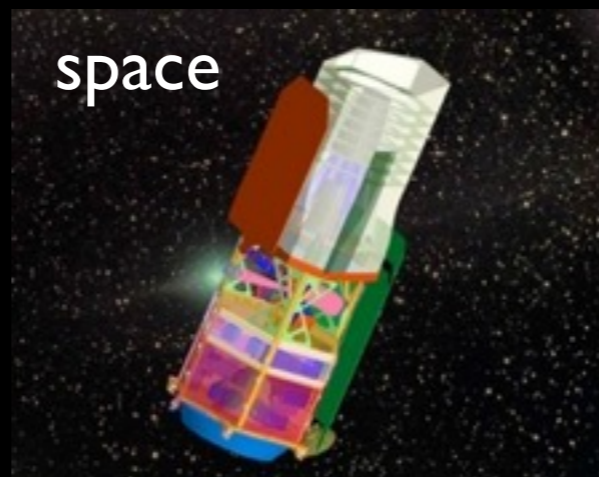


DESI  
(2018+)

WFIRST  
(2024++)



Euclid (2020)



LSST  
(2022++)

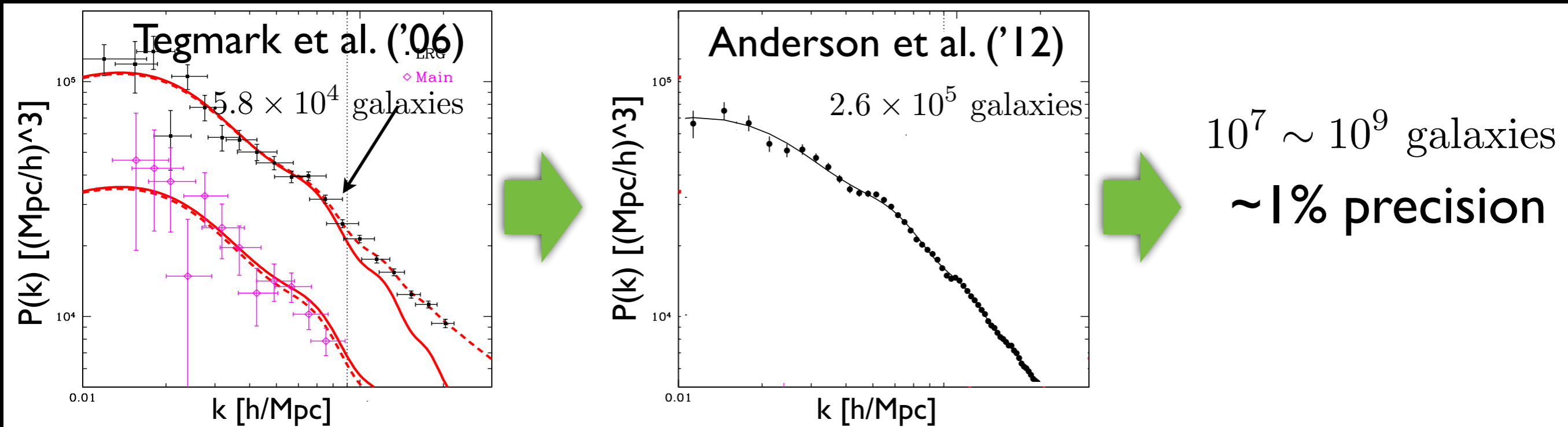


SuMIRe  
(2014~)



# LLS as precision cosmological tools

Large data set will reveal statistical properties of LSS at an unprecedented precision level ( $\rightarrow$  **precision cosmology**)



*New opportunity & scientific synergy :*

- Clarifying nature of dark energy (cosmic acceleration)
- Testing general relativity on cosmological scales
- Weighing total mass of neutrinos

Accurate theoretical description for LSS needs to be developed

# Goal of this lecture

Understanding of large-scale structure (LSS)  
as cosmology probe

- Theoretical basis of formation & evolution of LSS
  - Standard model of structure formation:  $\Lambda$ CDM
  - Cosmological information imprinted on LSS
- Theoretical tools to confront with precision observations of LSS  
(mainly focusing on galaxy surveys)
  - precision theoretical calculations of LSS

# Plan

- Friedmann cosmology

15th Nov.

- Linear theory of structure formation

- Observational effects:

Redshift-space & geometric distortions

16th Nov.

17th Nov.

- Analytic approaches to nonlinear structure formation