RESPONSE FUNCTION OF THE COSMIC DENSITY POWER SPECTRUM AND Takahiro Nishimichi (Kavli IPMU) w/ Francis Bernardeau (IAP) Atsushi Taruya (YITP) Based on PLB 762 (2016) 247 and arXiv:1708.08946 See also RESPRESSO webpage: http://www-utap.phys.s.u-tokyo.ac.jp/~nishimichi/public_codes/respresso/

OBJECTIVES

- To extract information from large-scale clustering signals, we need
 - an accurate theory/model
 - to meet the low statistical error level of big surveys
 - quick evaluation
 - to explore multi-D cosmological parameter space
- Different approaches available
 - Perturbation theory (and its variant)
 - Accuracy: large scale good (?), small scale no
 - Speed: high loop order (2 or 3 loops) takes time
 - (N-body) simulations
 - Accuracy: good (w/ sufficient volume/realizations, only after a careful convergence study)
 - Speed: takes much more time
 - RESPRESSO !
 - A possible integration of the 2 approaches

PERTURBATION THEORY IS IN CRISIS z = 0.375





PERTURBATION THEORY IS IN CRISIS z = 0.375N-body 1.4 Next to leading The success of the next-to-next-toleading (2 loop) order calculation just an illusion? P(k)/P order--linear theory 1.0 Next-to-next-to-next-toleading order 0.9 0.15 0.20 0.25 0.10 0.30 0.05 0.0 k [h/Mpc] Blas, Garny, Konstandin '14

WHA5

- Perturbation theory (PT) is fine only when the quantity of interest is small and the series expansion is convergent
 - Overdensity can reach >> 1 at present
- The fluid is assumed to follow an irrotational single-streaming flow
 - Enters to multi-streaming phase after shell crossing (even if the initial condition is cold)
- These 2 things happen together (small scale, late time)
 - The breakdown can propagate to large scales due to mode coupling
 - The way how PT breaks down is totally non-trivial —> simulations!

WHAT TO MEASURE?

Basic things that determine the nonlinear mode-coupling structure:

$$\{F_n(\mathbf{q}_1,\ldots,\mathbf{q}_n),G_n(\mathbf{q}_1,\ldots,\mathbf{q}_n)\}$$

 $\left\{\Gamma_{\delta}^{(n)}(\mathbf{q}_{1},\ldots,\mathbf{q}_{n}),\Gamma_{\theta}^{(n)}(\mathbf{q}_{1},\ldots,\mathbf{q}_{n})\right\}$

$$K(k,q) = q \frac{\delta P_{nl}(k)}{\delta P_{lin}(q)}$$

STANDARD PT KERNEL

• Very, very basic things (not observable from sims, though...)

$$\{F_n(\mathbf{q}_1,\ldots,\mathbf{q}_n),G_n(\mathbf{q}_1,\ldots,\mathbf{q}_n)\}$$

1

continuity + Euler + Poisson eqs.

GAMMA EXPANSION (eg., RegPT) Bernardeau + '09

$$\left\{\Gamma_{\delta}^{(n)}(\mathbf{q}_1,\ldots,\mathbf{q}_n),\Gamma_{\theta}^{(n)}(\mathbf{q}_1,\ldots,\mathbf{q}_n)\right\}$$

final state (density or velocity)multi-point propagator $\frac{1}{p!} \left\langle \frac{\delta^p \Psi_a(\mathbf{k}, \eta)}{\delta \phi_{b_1}(\mathbf{k}_1) \dots \delta \phi_{b_p}(\mathbf{k}_p)} \right\rangle = \delta_D(\mathbf{k} - \mathbf{k}_1 - \dots - \mathbf{k}_p) \Gamma_{ab_1 \dots b_p}^{(p)}(\mathbf{k}_1, \dots, \mathbf{k}_p; \eta)$ initial state

- Can calibrate against sims
- Good ansatz known
- Simpler expression for the spectra
- Clear physical interpretation (Crocce & Scoccimarro '06 for the 2pt propagator)

VIRTUE OF RESUMMATION

- Better convergence: everything is positive definite
- BAO almost done at the lowest order

large scale structure gravitational evolution

Input $(\Omega_m, h, ...; z)$ $P_{nl}(k)$ Output

large scale structure gravitational evolution

Input $(\Omega_m, h, \ldots; z)$ $P_{lin}(k)$ $P_{nl}(k)$

To a very good approximation

Output

large scale structure gravitational evolution

Input $P_{lin}(k)$ $P_{nl}(k)$ Output $K(k,q) = q \, \frac{\delta P_{nl}(k)}{\delta P_{lin}(q)}$

large scale structure gravitational evolution

Input $P_{lin}(k)$ $+\delta P_{lin}(k)$ $P_{nl}(k)$ Output $K(k,q) = q \, \frac{\delta P_{nl}(k)}{\delta P_{lin}(q)}$

large scale structure gravitational evolution

Input $P_{lin}(k)$ $+\delta P_{lin}(k)$ $P_{nl}(k)$ Output $K(k,q) = q \frac{\delta P_{nl}(k)}{\delta P_{lin}(q)}$ $+\delta P_{nl}(k)$

RESPONSE FUNCTION

$$\left| K(k,q) = q \, \frac{\delta P_{nl}(k)}{\delta P_{lin}(q)} \right|$$

I want to study this mode at some late time t k **q**₃ 2 what is the impact from wave mode q at the initial time to?

RESPONSE FUNCTION: THE FIRST TRIAL

RESPONSE FUNCTION: SIM VS PT

- SPT (2-loop) >> N-body @ high q
- This is exactly where PT breaks down
- What N-body tells us is:

"Physics at strongly nonlinear regime does not propagate to large scales"

Rescaled quantity:

$$\Gamma(k,q) \equiv [K(k,q) - K^{\text{lin}}(k,q)]/[qP^{\text{lin}}(k)]$$

HIGH RES RESPONSE FUNCTION

2

- 1400 N=512³ simulations to study fine structures of the response function
- Vs 2-loop calculation based on different schemes (SPT and RegPT)
- New phenomenological model introduced

TN, Bernardeau, Taruya '17 (arXiv:1708.08946)

RESPONSE FUNCTION AT q << k

- Response function goes to zero from simulations
 - Extended galilean invariance
- This is nicely explained by SPT
- Not the case for RegPT

RESPONSE FUNCTION AT q~k

- Peaky structure decay as time goes by
- SPT behaves weirdly at late time
- **RegPT** has its strength in this regime
 - Efficiently captures mode transfer between nearby modes

RESPONSE FUNCTION AT k << q

• We need phenomenology here anyway!

OUR MODEL

$$\begin{split} K_{\text{model}}(k,q) &= \left[\left(1 + \beta_{k,q} + \frac{1}{2} \beta_{k,q}^2 \right) K_{\text{tree}}^{\text{SPT}}(k,q) \right. \\ &+ \left(1 + \beta_{k,q} \right) K_{1\text{-loop}}^{\text{SPT}}(k,q) + K_{2\text{-loop}}^{\text{SPT}}(k,q) \right] D(\beta_{k,q}) \end{split}$$

$$D(x) = \begin{cases} \exp(-x), \text{ if } K_{\text{model}}(k,q) > 0, \\ \frac{1}{1+x}, & \text{ if } K_{\text{model}}(k,q) < 0. \end{cases}$$

 $\beta_{k,q} = \alpha_k + \alpha_q$ Regularize both in k and q

$$\alpha_k = \frac{1}{2}k^2\sigma_d^2; \quad \sigma_d^2 = \int \frac{dq}{6\pi^2} P_{\text{lin}}(q)$$

- Well-behaved over all q
- Eventually fails at high k

PRACTICAL USAGE? RECONSTRUCTION

• From the definition of a functional derivative

$$P_{\mathrm{nl}}(k;\boldsymbol{p}_{1}) \approx P_{\mathrm{nl}}(k;\boldsymbol{p}_{0}) + \int \mathrm{d}\ln q \, K(k,q)$$
$$\times \left[P_{\mathrm{lin}}(q;\boldsymbol{p}_{1}) - P_{\mathrm{lin}}(q;\boldsymbol{p}_{0})\right],$$

- Use this to predict P_{nl} for cosmological model \textbf{p}_1 given P_{nl} for another model \textbf{p}_0

STARTING POINT: SIMULATION DATABASE

- P_{nl} database for the fiducial Planck 2015 cosmology from 10 x 2048^3 sims
 - Cosmic variance suppressed with Angulo-Pontzen technique
 - Fractional error < 0.1%
- Can smoothly interpolate over k and time

A SIMPLE IMPLEMENTATION

• Double the reliable k range from the pure RegPT prediction

MORE EXTREME MODELS

Employ multi-steps

RESPRESSO PYTHON PACKAGE AVAILABLE!

(Rapid and Efficient SPectrum calculation based on RESponSe functiOn)

In [1]: %pylab inline import respresso

In [3]: respresso_obj = respresso.respresso_core()

Hello. This is RESPRESSO. Load precomputed data files... RESPRESSO ready.

In [6]: respresso_obj.set_target(plin_target_spl)

In [7]: respresso_obj.find_path()

In [9]: kwave = respresso_obj.get_kinternal()
pnl_rec = respresso_obj.reconstruct()

http://www-utap.phys.s.u-tokyo.ac.jp/~nishimichi/public_codes/respresso/

SUMMARY

- P(k) to a 2D quantity K(k,q): more physical insight
- Difficulty in perturbative approaches
 - Suppress small to large scale mode transfer!
 - SPT and RegPT have good and bad behavior in different regimes
- RESPRESSO package available
 - Response function is a natural interpolator over the cosmological parameter space
 - Can go to k ~ 0.44 (0.35) h/Mpc at z=1 (0.5) within 1%
 - You can put your own simulation data if you do not like mine ;)

http://www-utap.phys.s.u-tokyo.ac.jp/~nishimichi/public_codes/respresso/