

The background features a deep space scene with a galaxy in the upper left, the Earth's horizon in the lower right, and a bright star in the lower left. A prominent blue and purple beam of light streaks from the top left towards the bottom right. In the upper right, there are several blue, glowing, spiral-like particle tracks emanating from a central point.

# Connecting high-energy astroparticle physics for origins of cosmic rays and future perspectives

**December 7 - 10, 2020**

**Kyoto University, Kyoto, Japan**

**<http://www2.yukawa.kyoto-u.ac.jp/~crphys2020>**

<http://www2.yukawa.kyoto-u.ac.jp/~crphys2020/>

# What should we do in the workshop



Avoid crowded places and limit time in enclosed spaces



Maintain at least 1m distance from others



When possible, open windows and doors for ventilation



Keep hands clean and cover coughs and sneezes



Wear a mask if requested or if physical distancing is not possible

from WHO website

 **Masking, Washing and Social distance**

 **If you are unwell, please switch on-line at any time**

# What should we do in the workshop

**Avoid the Three Cs**

World Health Organization  
Western Pacific Region

Be aware of different levels of risk in different settings.

There are certain places where COVID-19 spreads more easily:



**Crowded places**

*with many people nearby*



**Close-contact settings**

*Especially where people have close-range conversations*



**Confined and enclosed spaces**

*with poor ventilation*

from WHO website

# Connecting high-energy astroparticle physics for origins of cosmic rays and future perspectives

December 7 - 10, 2020, Yukawa Institute for Theoretical Physics, Kyoto University, Kyoto, Japan

Invited talk (25 min + 5 min), Oral contribution (10 min + 5 min), Poster talk (5 min + on-site display)  
Time-zone: Japan Standard Time

148 registrations  
51 contributions

## 12/7 (Mon.)

13:00-14:30 Registration

14:30-15:20 Welcome coffee, Opening, Self-introduction (1 min/person) (optional)

15:30-19:00 Session1

Chairperson: Kohta Murase

15:30-16:00 **Kazumasa Kawata**

**100 TeV Gamma-Ray Observation with Extensive Air Shower Arrays**

16:00-16:15 Sei Kato

VHE gamma-ray astronomy using the prototype array of a new extensive air-shower-array experiment ALPACA in the southern hemisphere

16:15-16:30 Kimura Shigeo

Gamma-ray and neutrino emission from radiatively inefficient accretion flows

16:30-17:00 **Foteini Oikonomou**

**High-energy neutrino emission from blazars**

[30 minutes break]

17:30-18:00 **Olivier Deligny**

**The UHECR science after 15 years of operation of the Pierre Auger Observatory**

18:00-18:30 **Yana Zhezher**

**Overview of the Telescope Array experiment**

18:30-19:00 **Anatoli Fedynitch**

**Hadronic interactions in cosmic ray physics**

## 12/8 (Tue.)

09:00-12:30 Session2

Chairperson: Aya Ishihara

09:00-09:30 **Takatomi Yano**

**Neutrino astrophysics prospect at Super-Kamiokande and Hyper-Kamiokande**

09:30-10:00 **Ignacio Taboada**

**Recent Astrophysical results of IceCube**

10:00-10:30 **Stephanie Wissel**

**Radio detection of ultrahigh-energy neutrinos: present and future**

[30 minutes break]

11:00-11:30 **Nepomuk Otte**

**Trinity: An Air-Shower Imaging Instrument to detect Ultrahigh-Energy Neutrinos**

11:30-12:00 **Ruoyu Liu**

**The Giant Radio Array for Neutrino Detection**

12:00-12:15 Mahdi Bagheri

The UHE-Neutrino Cherenkov telescope onboard EUSO-SPB2

12:15-12:30 Susumu Inoue

High-energy neutrino and gamma-ray emission from AGN-driven winds

14:30 - 15:00 Poster session (5 min / person) and coffee, Chairperson: Wataru Ishizaki

Satoshi Takashima

GRAMS project: A MeV gamma-ray large area telescope using liquid argon and its concept study

Susumu Inoue

The Blazar Hadronic Code Comparison Project

Ken Matsuno

Particle acceleration by ion-acoustic solitons in plasma in a magnetic field

Ken Ohashi

Effects of diffractive collisions on predictions of the number of muons in the air shower

Tomohiko Oka

The time-evolution measurement of a diffusive shock acceleration using supernova remnants and local molecular clouds

15:30-19:15 Session3

Chairperson: Hiroaki Menjo

15:30-16:00 **Teruaki Enoto**

**High-Energy Atmospheric Physics of Lightning and Thunderstorms Observed along the Sea of Japan**

16:00-16:30 **Markus Alhers**

**Cosmic-Ray Anisotropy**

16:30-17:00 **Roberta Colalillo**

**The Pierre Auger Observatory and the study of atmospheric electricity phenomena**

[30 minutes break]

17:30-18:00 **Ioana Maris**

**Future Detectors for Measuring Ultra High Energy Cosmic Rays from the Ground**

18:00-18:30 **Maria Petropoulou**

**Blazar neutrinos: implications of recent IceCube observations**

18:30-19:00 **Walter Winter**

**Gamma-ray bursts and tidal disruption events as the sources of UHECRs and neutrinos**

19:00-19:15 Norita Kawanaka

Origin of Spectral Hardening of Secondary Cosmic-Ray Nuclei

## 12/9 (Wed.)

09:00-12:10 Session4

Chairperson: Kunihito Ioka

09:00-09:30 **Yoshihiro Ueda**

**The origin of the cosmic X-ray background**

09:30-10:00 **Yoshiyuki Inoue**

**Future Prospects of MeV Gamma-ray Astronomy**

10:00-10:30 **Keith Bechtol**

**Using optical surveys to explore the origin of cosmic rays**

[25 minutes break]

10:55-11:25 **Andreas Zoglauer**

**Future missions in the MeV domain: COSI & AMEGO**

11:25-11:55 **Atsushi Takada**

**MeV gamma-ray observations utilizing electron-tracking Compton cameras loaded on balloons**

11:55-12:10 Nagisa Hiroshima

Dark matter search in extended dwarf spheroidal galaxies with CTA

14:30 - 15:00 Poster session (5 min /person) and coffee, Chairperson: Wataru Ishizaki

Yutaka Fujita

Intrusion of Cosmic-Rays into Molecular Clouds Studied by Ionization, the Neutral Iron Line, and Gamma-Rays

Yugo Omura

NICHE detector and analysis results

Ryo Sawada

A Consistent Modeling of Neutrino-driven Wind with Accretion Flow onto a Protoneutron Star and its Implications for  $^{56}\text{Ni}$  Production

Kenta Terauchi

The Fluorescence detector Array of Single-pixel Telescopes: The next-generation cosmic ray observatory

15:30 - 19:00 Session5

Chairperson: Yudai Suwa

15:30-16:00 **Yutaka Ohira**

**Cosmic-ray acceleration in supernova remnants**

16:00-16:15 Naomi Tsuji

Systematic study of acceleration efficiency in young supernova remnants

16:15-16:30 Hiromasa Suzuki

Observational gamma-ray and X-ray study on cosmic-ray escape from supernova remnants

16:30-16:45 Tomoaki Kasuga

cipher: a CubeSat-Based Hard X-ray Imaging Polarimetry Mission

[30 minutes break]

17:15-17:45 **Kumiko Kotera**

**Pulsars and magnetars as high-energy cosmic particle sources**

17:45-18:15 **Andrew Taylor**

**Particles Acceleration in the Jets of Centaurus A**

18:15-18:30 Merten Lukas

Ultra-high Energy Cosmic Rays Acceleration in FR 0 Radio Galaxies

18:30-19:00 **Yoshiyuki Takizawa**

**Observation of ultra high energy cosmic rays from space (K-EUSO and POEMMA)**

## 12/10 (Thu.)

09:00-12:45 Session6

Chairperson: Tsuyoshi Nakaya

09:00-09:30 **Kazumi Kashiyama**

**Fast Radio Bursts: A Mystery Being Solved?**

09:30-09:45 Lin Haoxiang

Afterglows of neutron star mergers and fast radio bursts

09:45-10:15 **Imre Bartos**

**Compact object mergers as high-energy multi-messenger sources**

10:15-10:30 Shuta Tanaka

Stochastic acceleration model of very young pulsar wind nebula associated with SN 1986J

[30 minutes break]

11:00-11:30 **Ke Fang**

**High-energy Cosmic Particles by Black-hole Jets in Galaxy Clusters**

11:30-12:00 **Ali Kheirandish**

**High-Energy Neutrinos as Probes of New Physics**

12:00-12:15 Ryo Higuchi

Effects of Galactic magnetic field on the UHECR anisotropy studies

12:15-12:30 On Alvina Yee Lian

Diagnosing the invisible: cosmic magnetism and the radio sky

12:30-12:45 Takahiro Sudoh

Millisecond Pulsars Modify the Radio-SFR Correlation

14:30 - 16:00 Summary

Chairperson: Toshihiro Fujii

14:30-15:30 Overview Discussion and Summary

15:30-16:00 Workshop Photo and Closing

Workshop photo!!

# What should we do in the workshop

- 📌 To unentangle origins of cosmic rays
  - 📌 Connect multi-wavelength and multi-particle observations
  - 📌 Connect active theorists and experimentalists
- 📌 Coffee break
  - 📌 [on-line]
    - 📌 we will try "**breakout room**" with a **random allocation** of on-line participants (~8 persons / room)
  - 📌 [on-site]
    - 📌 please remember **masking and social distance**
    - 📌 please eat sweets **only outside of YITP building**
    - 📌 **Only water** acceptable in the Panasonic hall and Y206 (poster room)
    - 📌 **Only drink** acceptable in Y105 (at front of Panasonic hall)
      - 📌 Please understand any epidemic possibilities on your own responsibility

# Trash

Others

Pet bottles

CANS



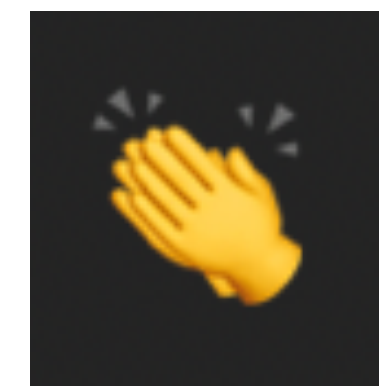
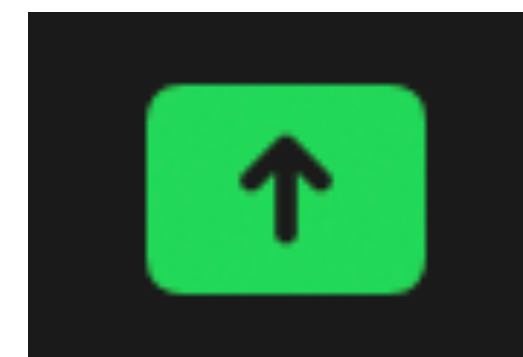
# Workshop location



↓ vending machine

# To contributors and audiences

- 📌 **Please keep on time**
- 📌 Invited talk (25 min + 5 min)
- 📌 Oral (10 min + 5 min)
- 📌 Poster (5 min + on-site display)
- 📌 If possible, please activate your video during your talk
- 📌 **[Both] Join Zoom and share your screen**
- 📌 **[Onsite] Please be seated with a social distance**
- 📌 **[Online] Please "clap" via reactions**



## Do boring speakers really talk for longer?

Robert M. Ewers ✉

**2.5 $\sigma$  (in tension)**

Dull talks at conferences can feel interminable. Or could it be that they really do go on for longer?

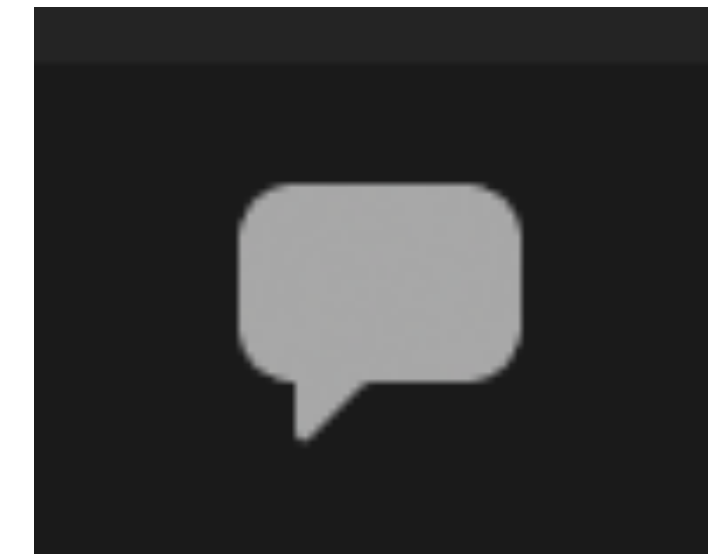
I investigated this idea at a meeting where speakers were given 12-minute slots. I sat in on 50 talks for which I recorded the start and end time. I decided whether the talk was boring after 4 minutes, long before it became apparent whether the speaker would run overtime. The 34 interesting talks lasted, on average, a punctual 11 minutes and 42 seconds. The 16 boring ones dragged on for 13 minutes and 12 seconds (thereby wasting a statistically significant 1.5 min;  $t$ -test,  $t = 2.91$ ,  $P = 0.007$ ). For every 70 seconds that a speaker droned on, the odds that their talk had been boring doubled. For the audience, this is exciting news. Boring talks that seem interminable actually do go on for longer.

To avoid banality, speakers should introduce their objectives early on and focus on pertinent information. They should avoid trite explanations, repetition, getting bogged down by irrelevant minutiae and passing off common knowledge as fresh insight.

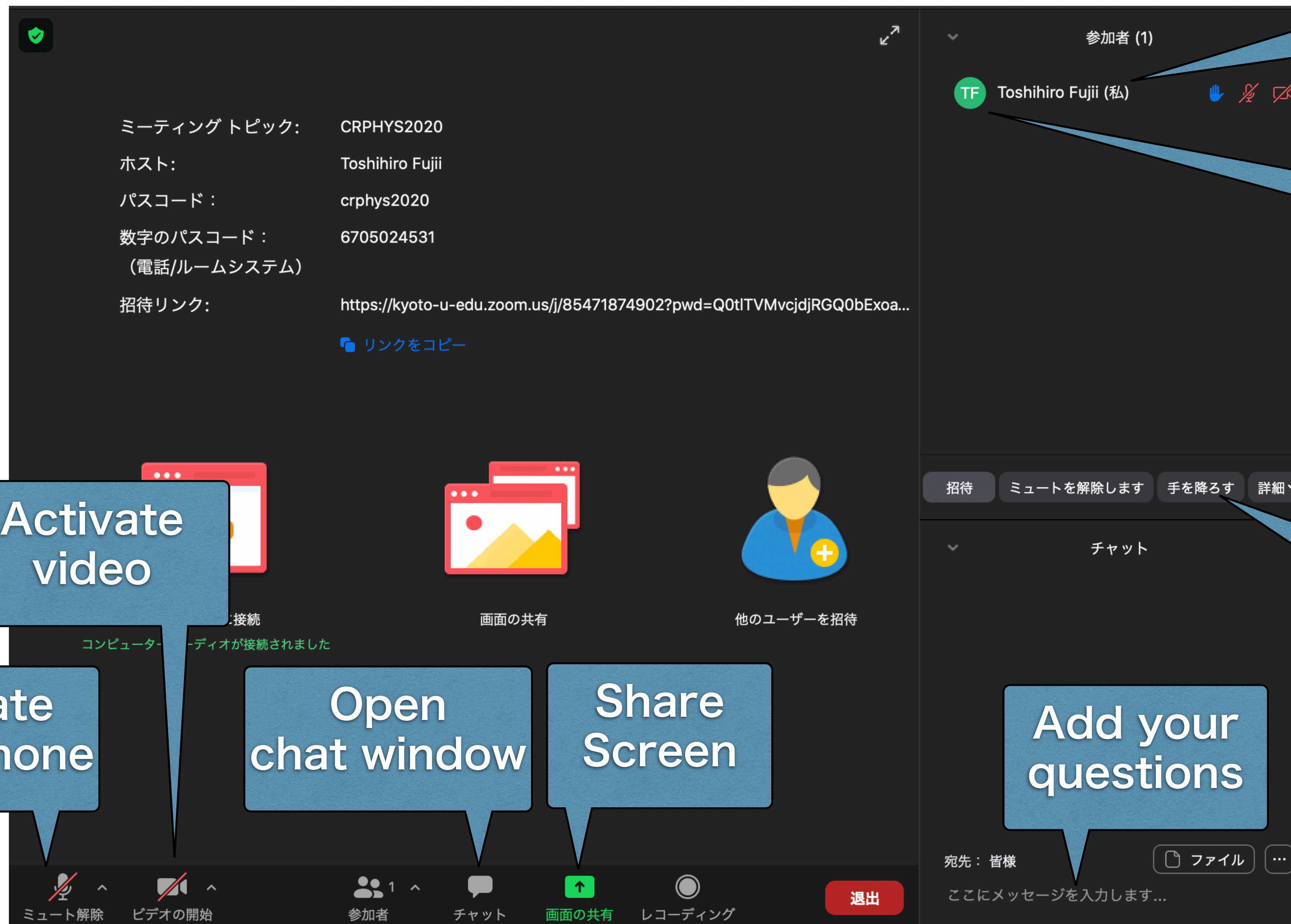


# Question and discussion

- 📌 During talk,
  - 📌 [Both] please use the chat window of Zoom
  - 📌 Short question acceptable
- 📌 After talk,
  - 📌 [Online] Please **raise your hand by Zoom**, and wait for call from chairperson
  - 📌 [Onsite] Please **step to microphone stand**, and wait for call from chairperson
- 📌 Please take over additional questions and discussions via **Slack**



# Zoom tips



Change your name and institute

Please add your photogenic photo

Raise your hand

Activate video

Activate microphone

Open chat window

Share Screen

Add your questions

# Join Slack

- 📌 Login Slack from the invitation link
- 📌 **LINK: participants only**
- 📌 Unlimited discussion during workshop
  - 📌 English or Japanese (日本語)
- 📌 The important discussions will be addressed in the Summary session
- 📌 **Please upload your slide via Slack** (after removing your confidential slides)
- 📌 Please compress your PDF **below 100 MB**
  - 📌 Alternatively, just send slide by an email to [crphys2020@yukawa.kyoto-u.ac.jp](mailto:crphys2020@yukawa.kyoto-u.ac.jp)
  - 📌 The slides will be shared among participants

# Slack tips

Channels for sessions and posters

CRPHYS2020

#general ☆  
Company-wide announcements and work-based matters

35

CRPHYS2020\_program\_ver4.pdf  
102 kB PDF

Connecting high-energy astroparticle physics  
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15:30-16:00 Kazumasa Kawata 100 TeV Gamma-Ray Observation with Extensive Air Shower Arrays  
16:00-16:15 Sei Kato VHE gamma-ray astronomy using the prototype array of a new extensive air-shower-array

Toshiiro Fujii 8:22 PM  
Abstract book  
PDF

CRPHYS2020\_abstract\_ver1.pdf  
175 kB PDF

Connecting high-energy astroparticle physics  
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15:30-16:00 100 TeV Gamma-Ray Observation with Extensive Air Shower Arrays  
Kazumasa Kawata (Institute for Cosmic Ray Research, the University of Tokyo)

Message #general

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Aa @ 😊 📎 ➤

Upload your slide (<100 MB)  
Add questions and discussions

# To invited speakers

- 📌 Please describe your personal opinion at your final slide
- 📌 **What's your targeted physics in next decades?**
- 📌 **What we need to accomplish?**
- 📌 and take-home messages (optional)

# For poster contributors

-  Please upload your poster in **Slack**
-  The poster room is Y206 at the second floor, please attach your poster
-  Please login Slack and check any comments and questions in your own channel (# poster\_your\_name)

# Connecting multi-wavelength and multi-particle observations for Cosmic Ray Ground Unified Theory (CR-GUT)

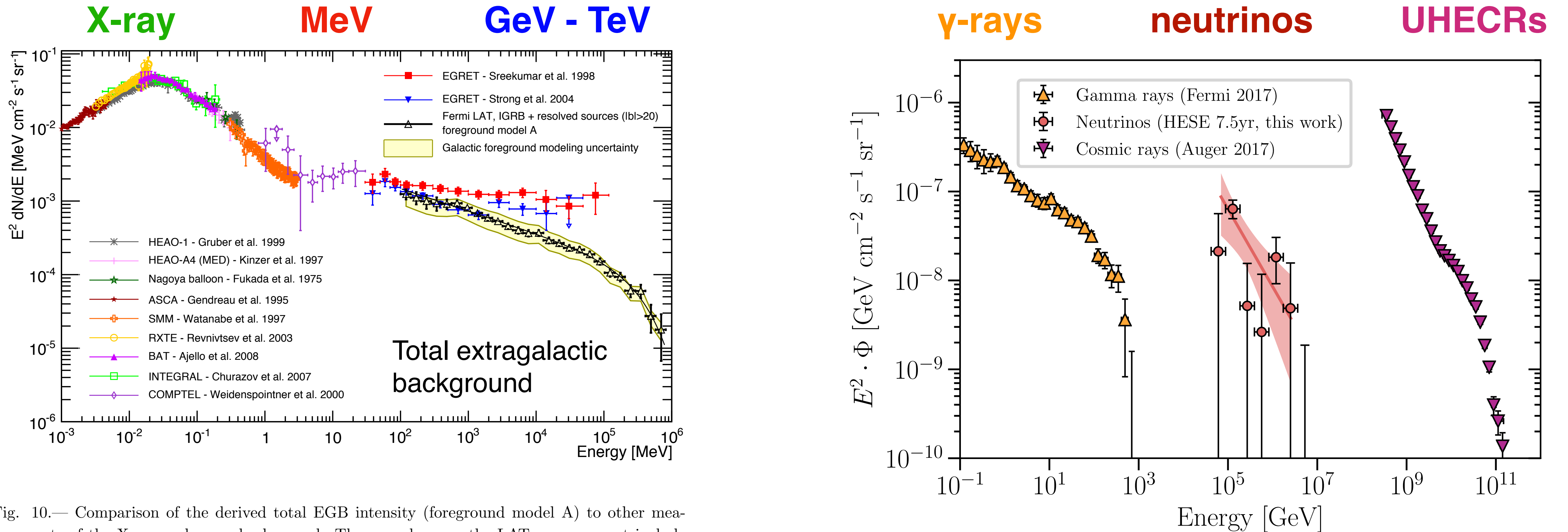
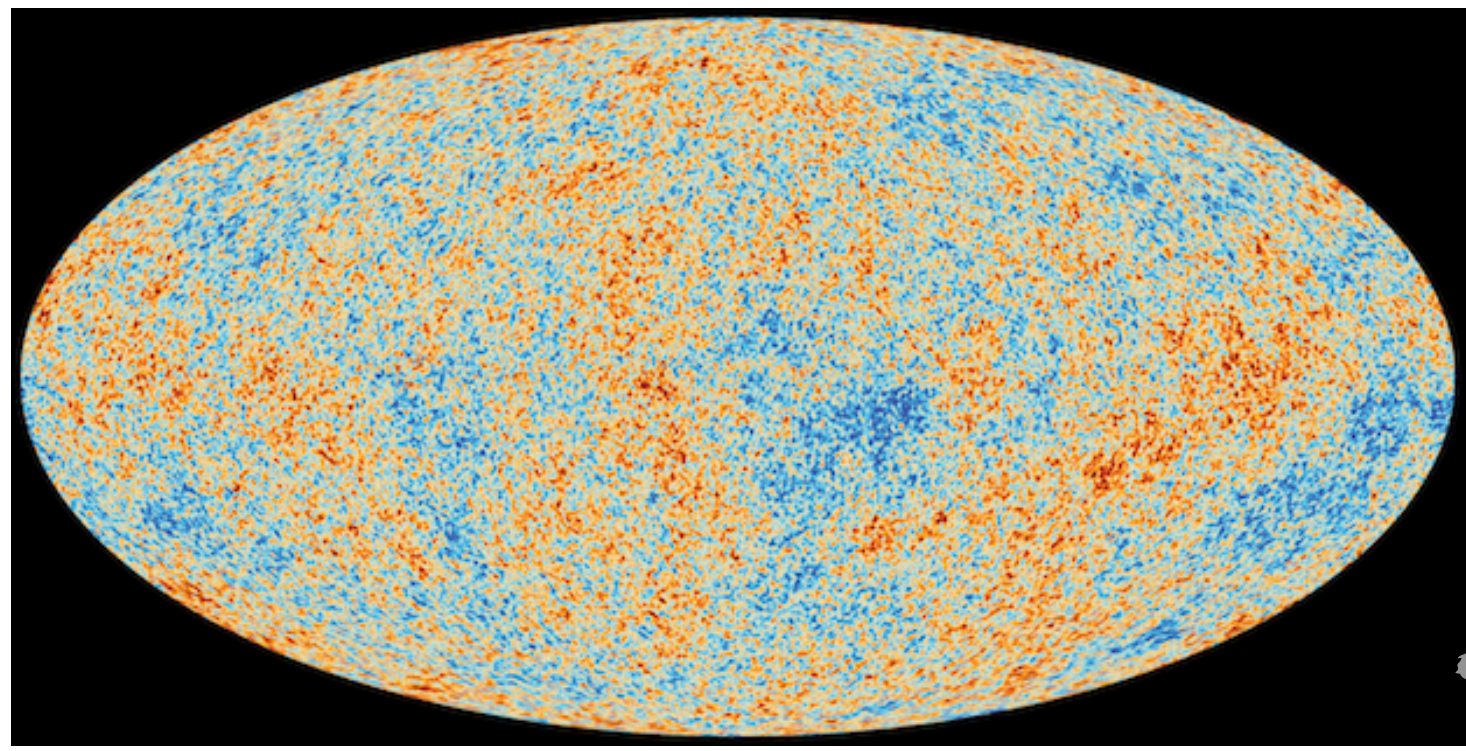
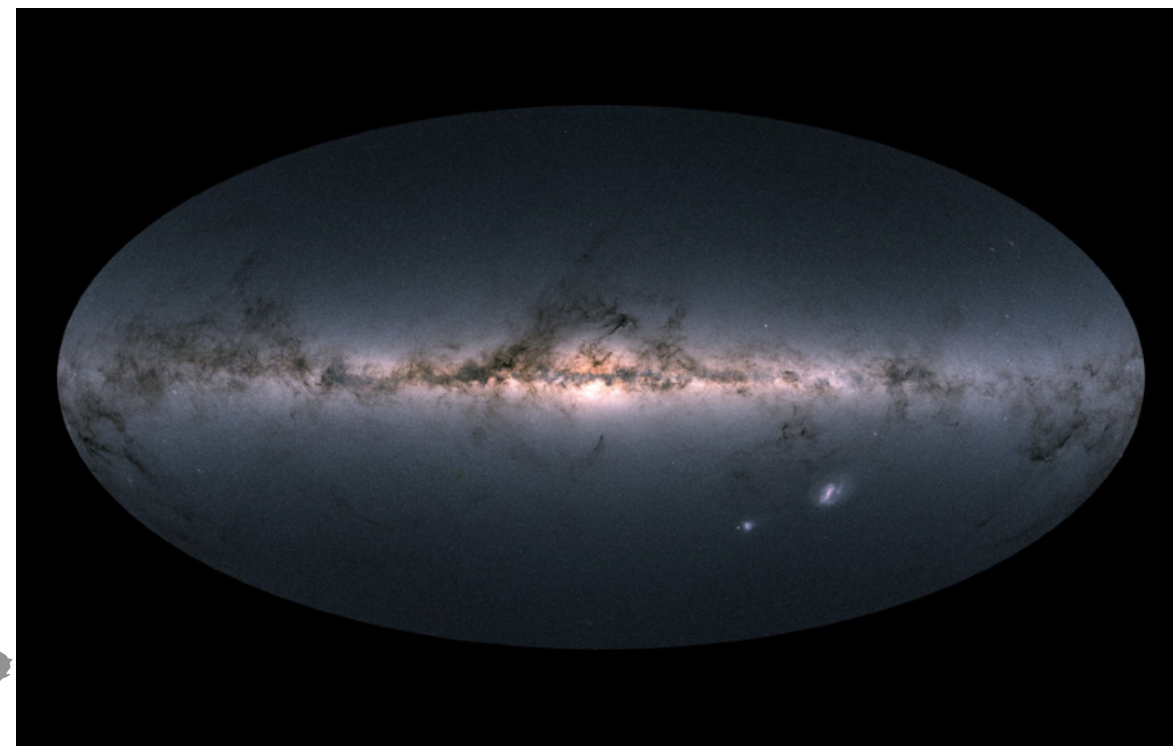


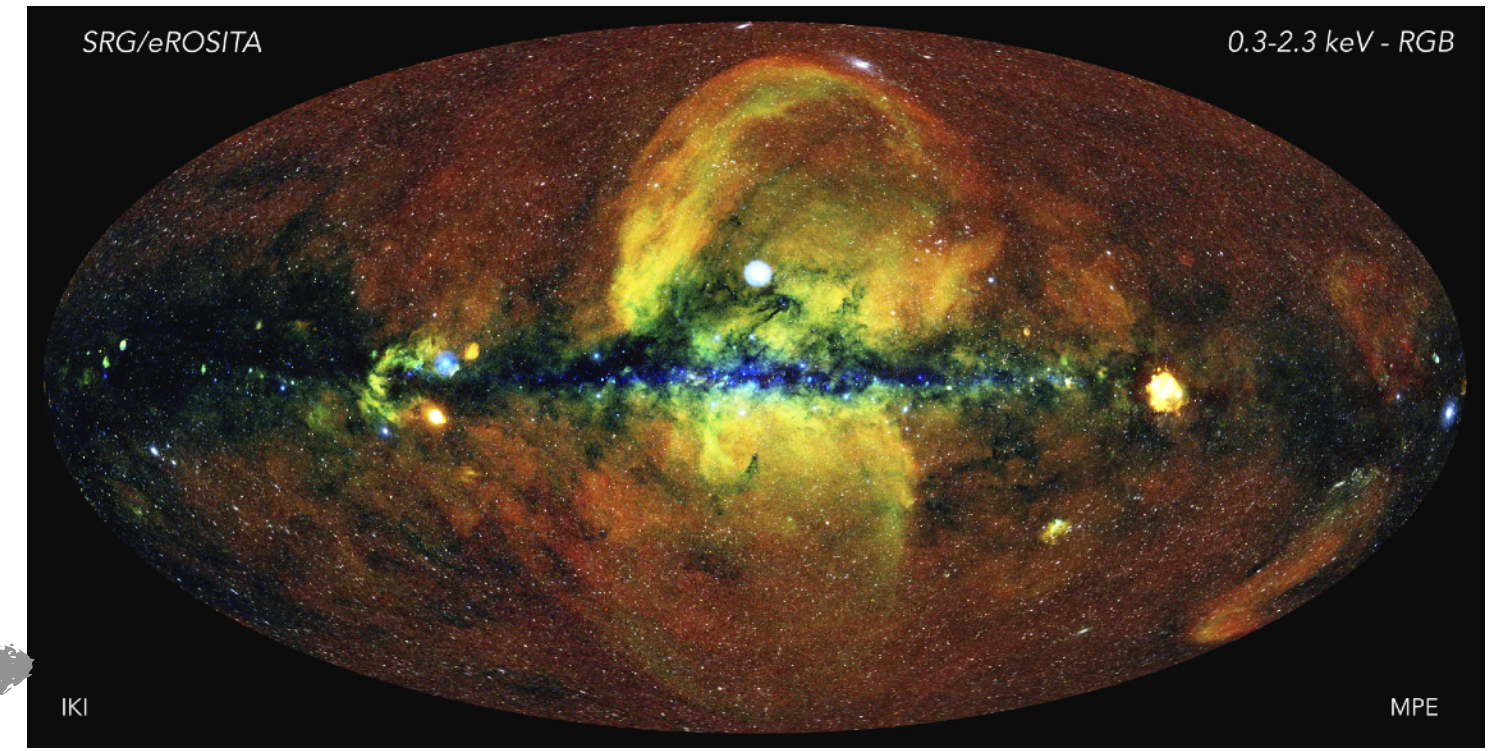
Fig. 10.— Comparison of the derived total EGB intensity (foreground model A) to other measurements of the X-ray and  $\gamma$ -ray background. The error bars on the LAT measurement include the statistical uncertainty and systematic uncertainties from the effective area parametrization, as well as the CR background subtraction. Statistical and systematic uncertainties have been added in quadrature. The shaded band indicates the systematic uncertainty arising from uncertainties in the Galactic foreground. (Note that the EGRET measurements shown are measurements of the IGRB. However, EGRET was more than an order of magnitude less sensitive to resolve individual sources on the sky than the *Fermi*-LAT.)



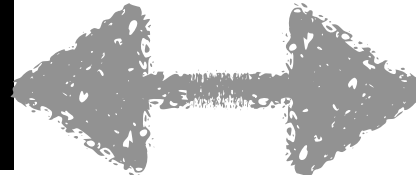
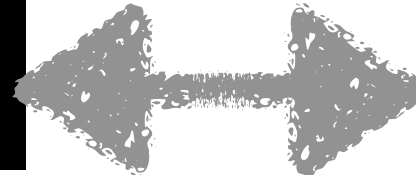
Planck Collaboration



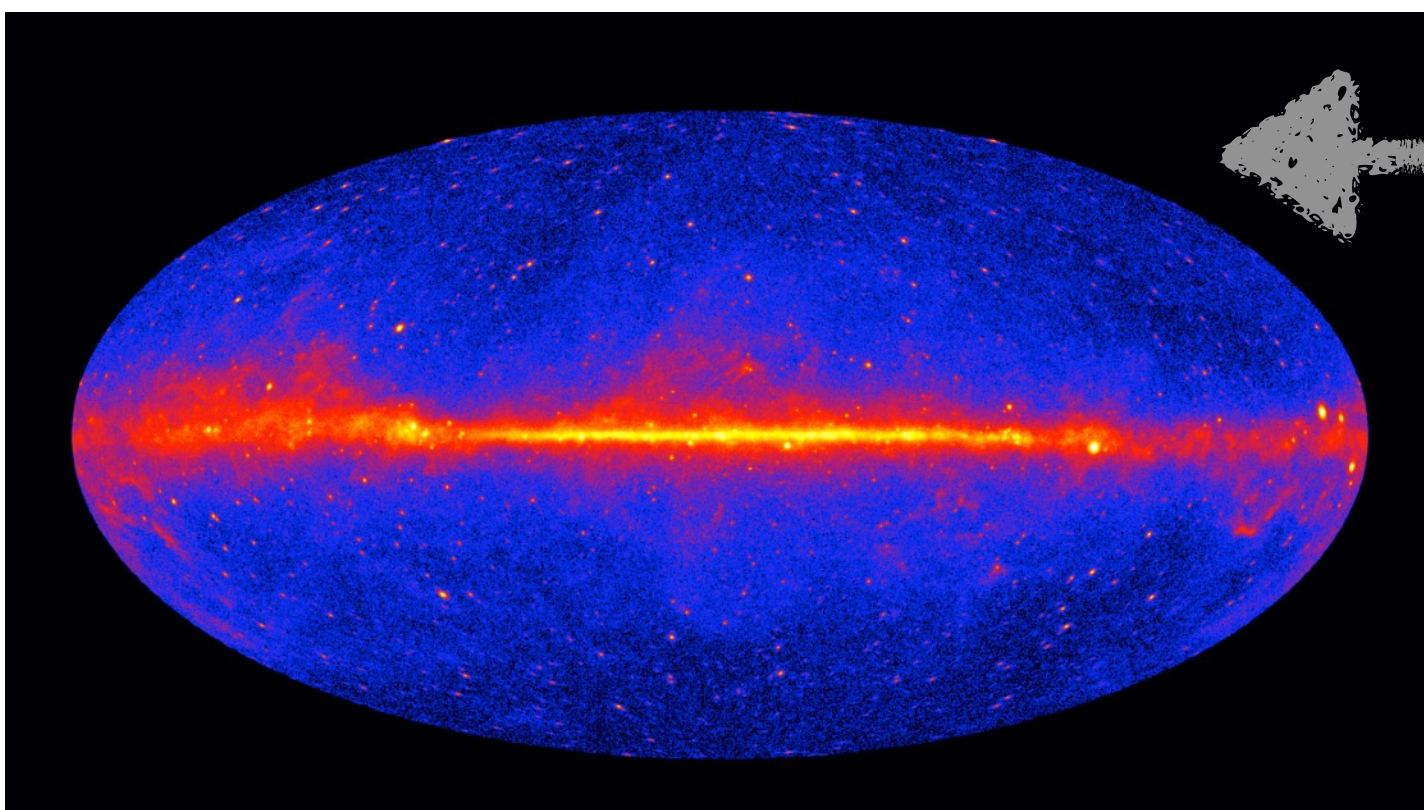
GAIA Collaboration



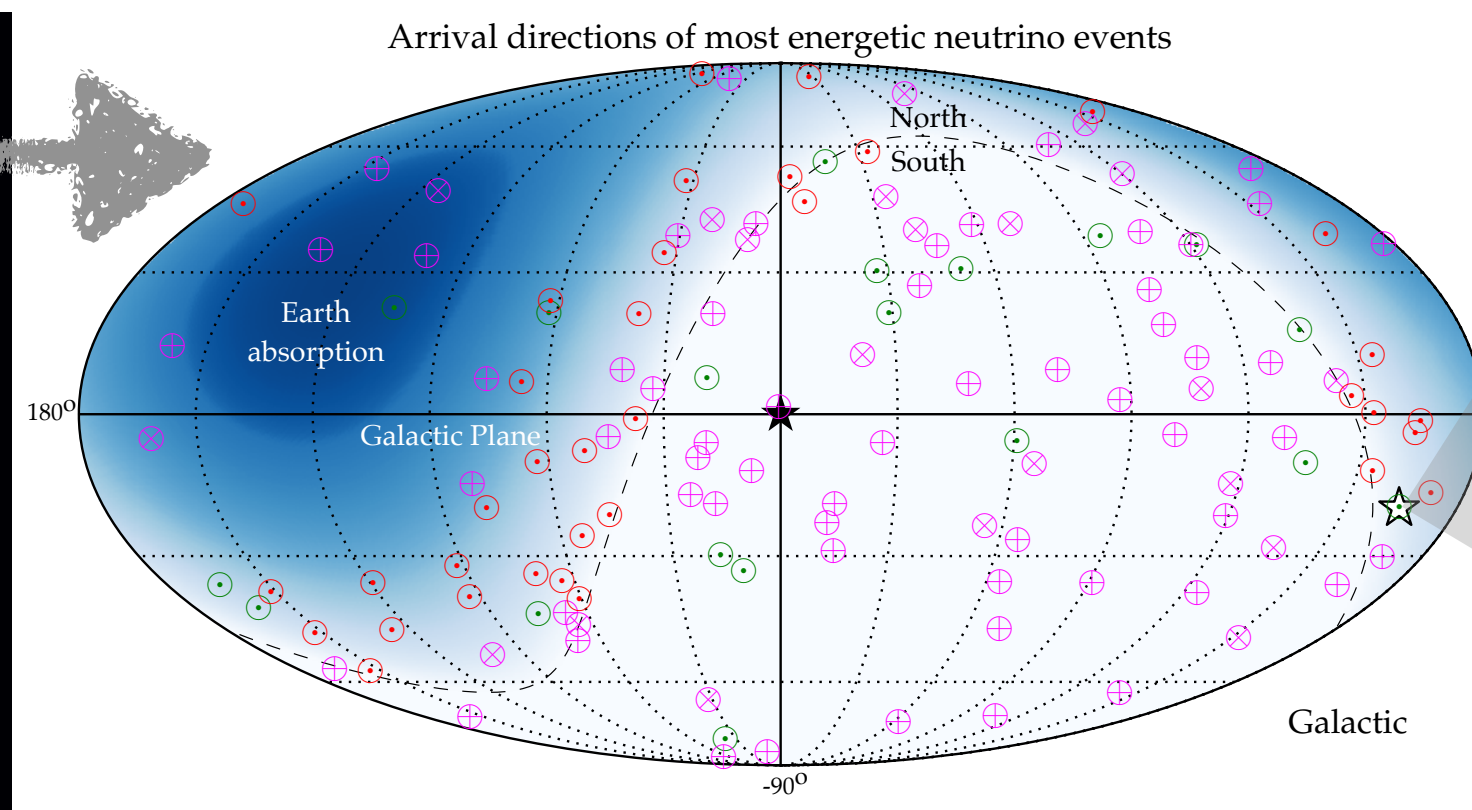
eROSITA Collaboration



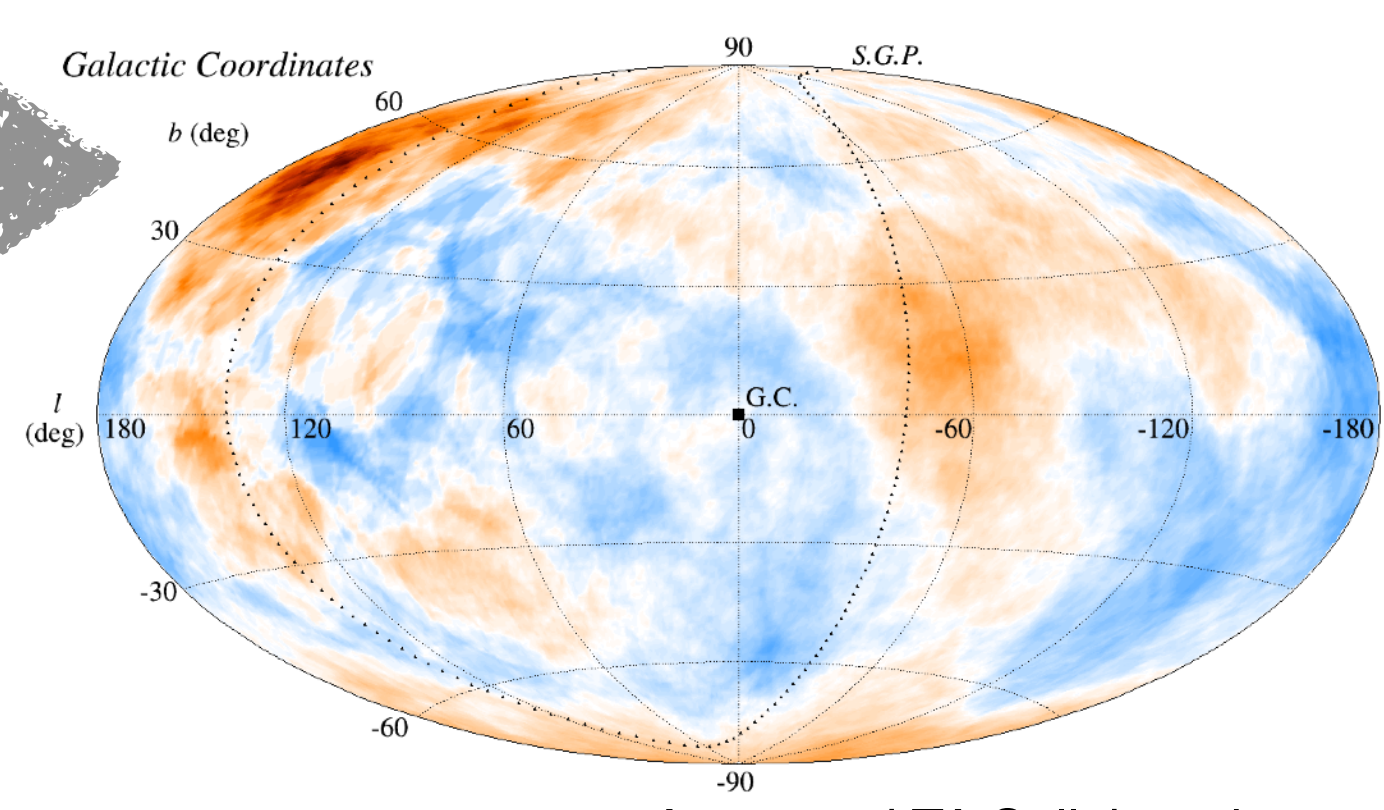
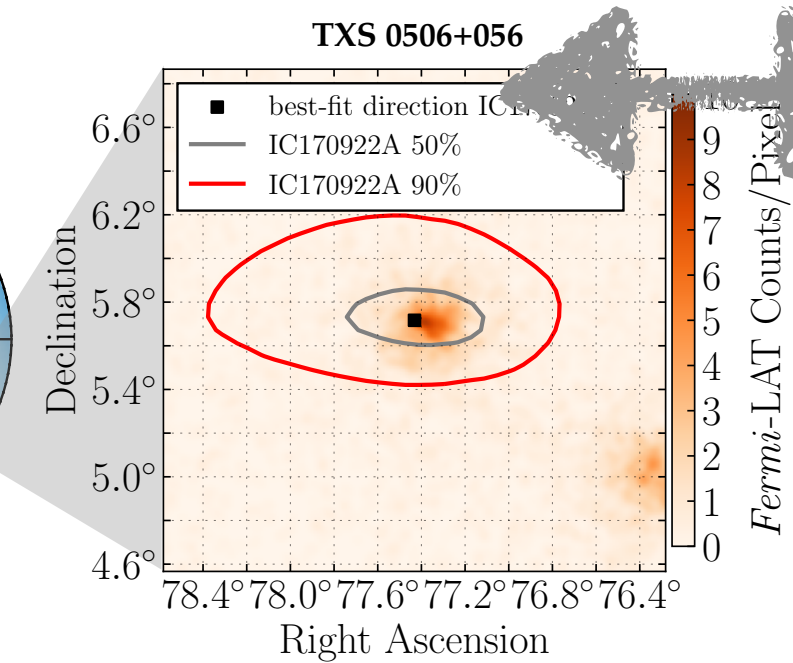
# Cosmic Ray Ground Unified Theory (CR-GUT)



Fermi Collaboration



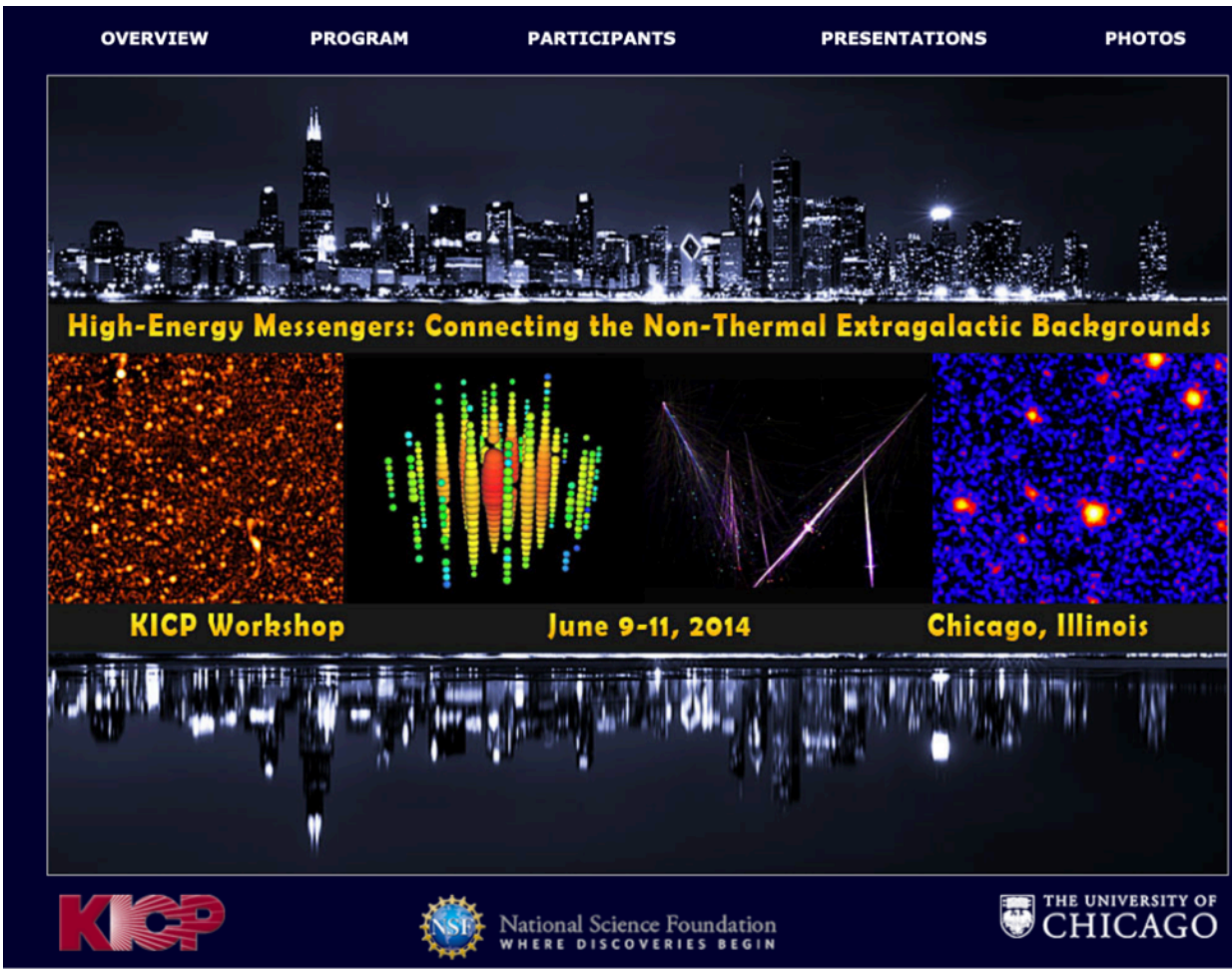
IceCube Collaboration



Auger and TA Collaborations

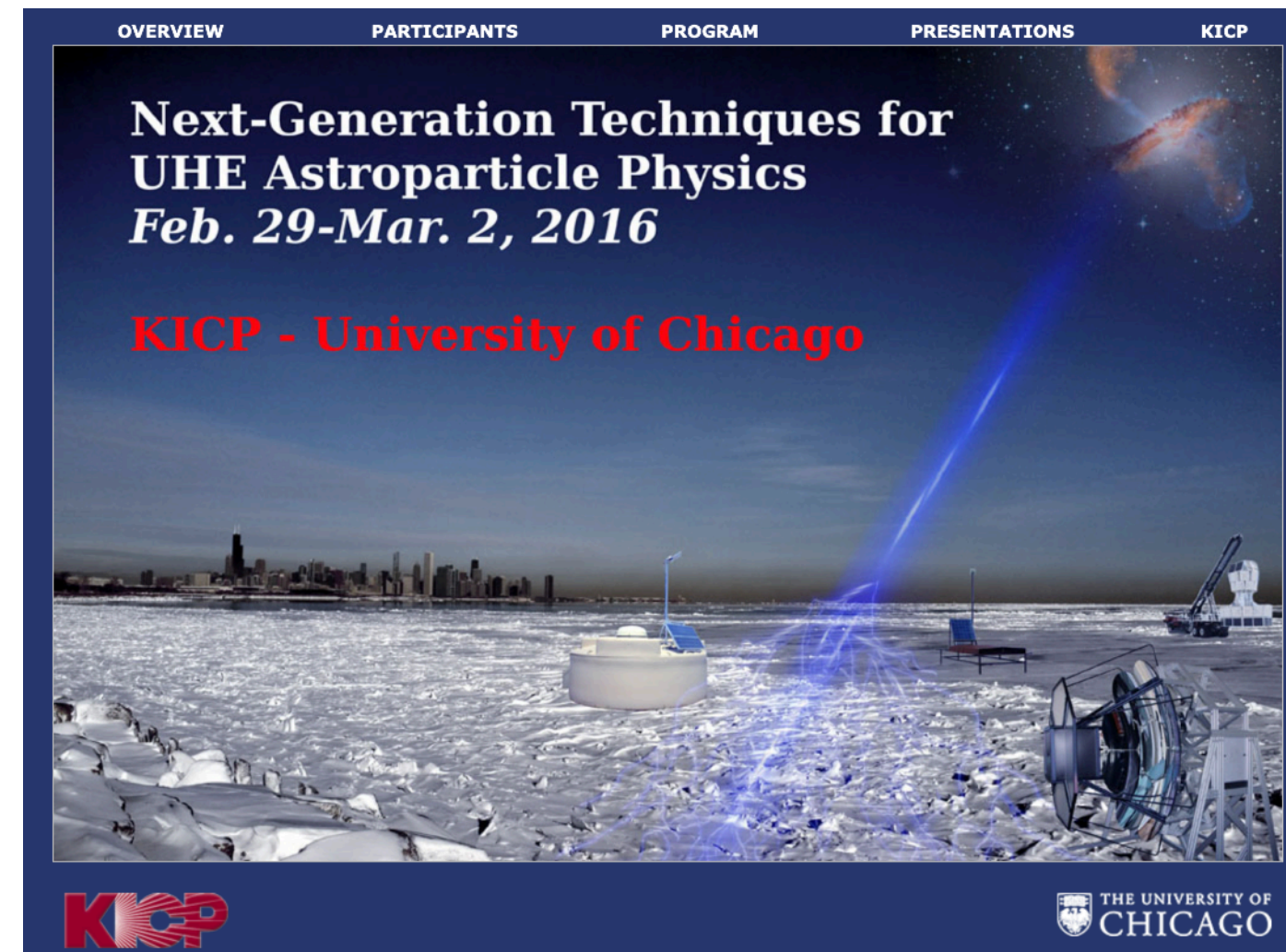


# A series of workshops... (just my personally)



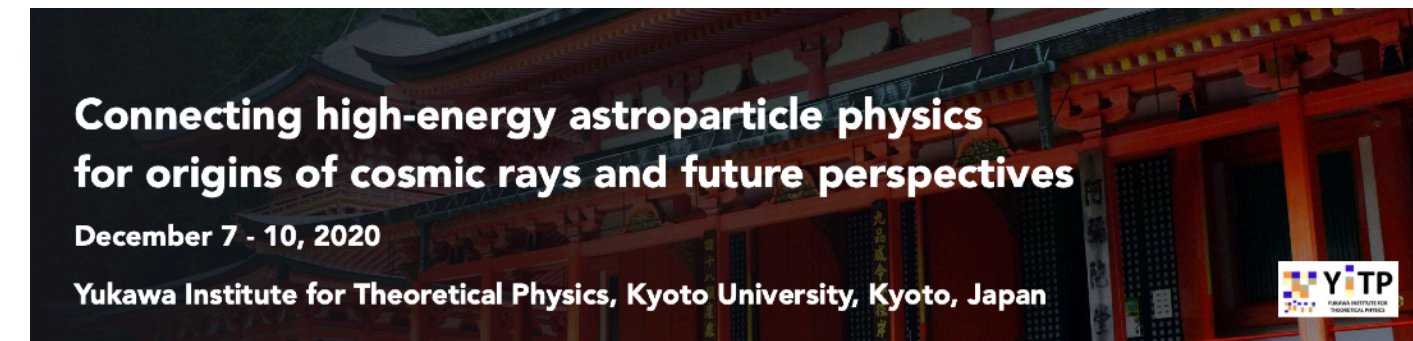
<https://kicp-workshops.uchicago.edu/hem2014/>

2014



2016

<https://kicp-workshops.uchicago.edu/uheap2016/>



2020

<http://www2.yukawa.kyoto-u.ac.jp/~crphys2020/>



# Let's start self introduction

- 📌 ~1 min / person
- 📌 Your name
- 📌 Your research interests
- 📌 Your current ongoing tasks
- 📌 etc...
- 📌 To on-line participants
- 📌 **Please raise your hand if you would like to do self-introduction**
- 📌 If possible, please activate your video
- 📌 Important for sound-check of your system