

AdS_5 with two boundaries and holography of N=4 SYM theory.

Masafumi Ishihara

Tohoku U. AIMR

Collaborators: Kazuo Ghoroku Fukuoka Inst. Tech.

Akihiro Nakamura Kagoshima U.

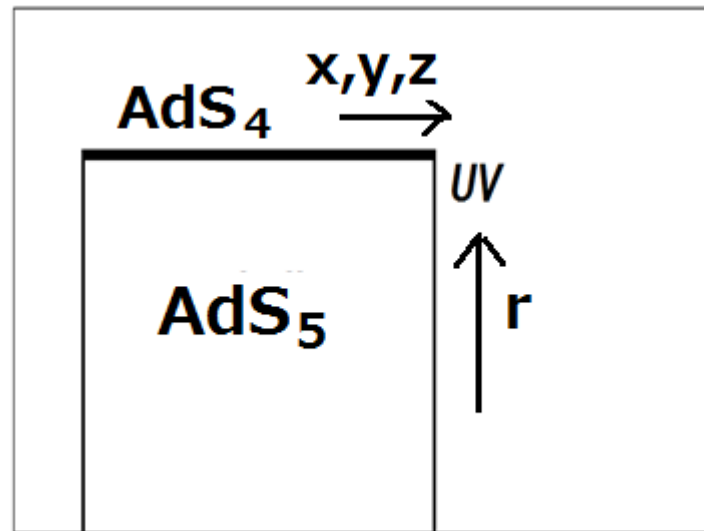
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Introduction

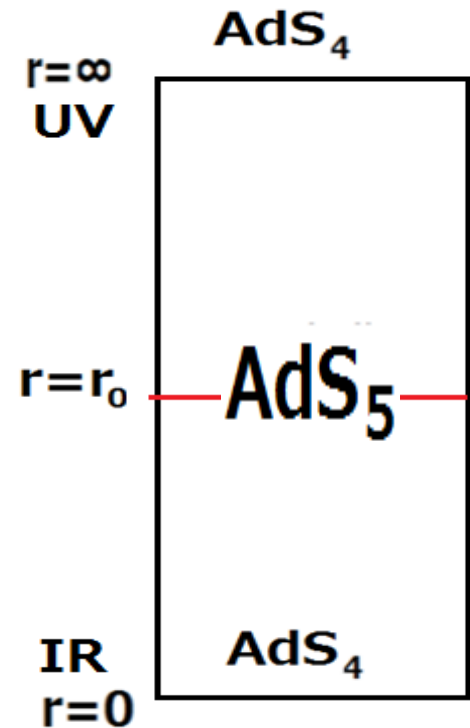
We consider AdS_5 metric whose 4D boundary at $r \rightarrow \infty$ (UV) has (AdS_4) metric.



r : 5-dimensional direction

Introduction

It is found that there is another 4D boundary at $r \rightarrow 0$ (IR).
UV and IR boundary field theories are described by common gravity dual but there is a domain wall in the 5D-bulk at some specific $r = r_0$



Contents

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Set Up

We consider the 10D metric in the following form

$$ds_{10}^2 = g_{MN} dx^M dx^N + R^2 d\Omega_5^2 \quad (M, N = 0 \cdots 5)$$

where the equation of motion for the non-compact 5-dimensional space-time is given as

$$R_{MN} = -\Lambda g_{MN} \quad (M, N = 0 \cdots 5)$$

Set up

Solution is obtained in the following form of the metric,

$$ds_5^2 = \frac{r^2}{R^2} A(r)^2 (-dt^2 + a_0^2(t) \gamma_{ij}(x) dx^i dx^j) + \frac{R^2}{r^2} dr^2$$

such that 4-dimensional UV boundary ($r \rightarrow \infty$) metric becomes AdS₄ metric.

$$ds_4^2 = -dt^2 + a_0^2(t) \gamma_{ij}(x) dx^i dx^j$$

where $\gamma_{ij}(x) = \delta_{ij} \left(1 - \frac{1}{4} \sum_{i=1}^3 (x^i)^2\right)^{-2}$

$$a_0(t) = \sin(\sqrt{-\lambda} t) / \sqrt{-\lambda}$$

$\lambda (< 0)$: 4D negative cosmological constant

We will find $A(r)$ which satisfies

$$A(r) \rightarrow 1 \quad \text{for } r \rightarrow \infty.$$

Set up

From the Einstein equation, $A(r)$ is determined as

$$A(r) = 1 + \left(\frac{r_0}{r}\right)^2 \text{ where } r_0^2 = -\frac{\lambda R^4}{4}$$

We can get the metric as follows

$$ds^2 = \frac{r^2}{R^2} \left(1 + \frac{r_0^2}{r^2}\right)^2 (-dt^2 + a_0^2(t) \gamma_{ij}(x) dx^i dx^j) + \frac{R^2}{r^2} dr^2 + R^2 d\Omega_5$$

$$\gamma_{ij}(x) = \delta_{ij} \left(1 - \frac{1}{4} \sum_{i=1}^3 (x^i)^2\right)^{-2}$$

The boundary ($r \rightarrow \infty$) 4D metric becomes AdS_4

Domain wall

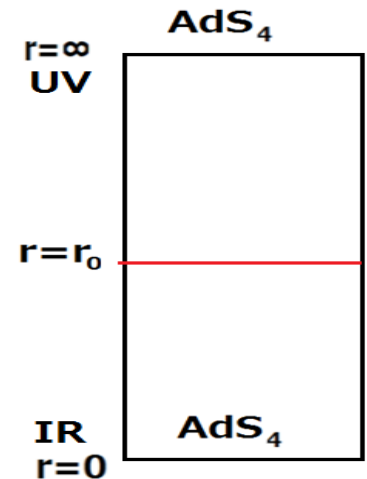
In this metric, there are two boundaries for $(r \rightarrow 0)$ and $(r \rightarrow \infty)$

By changing $r = r_0^2/z$, the metric becomes

$$ds^2 = \frac{z^2}{R^2} \left(1 + \frac{r_0^2}{z^2} \right)^2 (-dt^2 + a_0^2(t) \gamma_{ij}(x) dx^i dx^j) + \frac{R^2}{z^2} dz^2$$

Thus, the IR boundary metric at $z \rightarrow \infty (r \rightarrow 0)$ is also AdS_4 which is the same form as the UV boundary metric at $r \rightarrow \infty$.

The bulk metric are separated at $r = r_0$.



Trace anomaly by the AdS_4 field theory

The trace anomaly of the field theory on the 4D curved space-time with n_s scalars, n_f fermions and n_v vectors is given as

$$\langle T_{\mu}^{\mu} \rangle = -\frac{n_s + 11n_f + 62n_v}{90\pi^2} E_4 - \frac{n_s + 6n_f + 12n_v}{30\pi^2} I_{(4)}$$

$$E_{(4)} = \frac{1}{64} (R^{\mu\nu\lambda\sigma} R_{\mu\nu\lambda\sigma} - 4R^{\mu\nu} R_{\mu\nu} + R^2)$$

$$I_4 = -\frac{1}{64} (R^{\mu\nu\lambda\sigma} R_{\mu\nu\lambda\sigma} - 2R^{\mu\nu} R_{\mu\nu} + \frac{1}{3} R^2) \quad (\mu, \nu = t, x, y, z)$$

(N.J.Duff 1994) (M. Henningson and K. Skenderis, 1998)

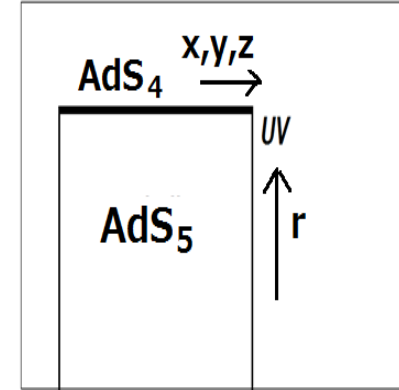
For 4D N=4 SYM theory on AdS_4 with cosmological constant $\lambda (< 0)$, it becomes

$$\langle T_{\mu}^{\mu} \rangle = -\frac{3\lambda^2}{8\pi^2} N^2 \quad (n_s = 6, n_f = 2, n_v = 1)$$

Energy momentum tensor (UV) by holography

The five dimensional metric is rewritten as

$$\begin{aligned}
 ds_5^2 &= \frac{r^2}{R^2} A(r)^2 (-dt^2 + a_0^2(t) \gamma_{ij}(x) dx^i dx^j) + \frac{R^2}{r^2} dr^2 \\
 &= \frac{1}{\rho} A(r)^2 (-dt^2 + a_0^2(t) \gamma_{ij}(x) dx^i dx^j) + \frac{d\rho^2}{4\rho^2} \\
 &\equiv \frac{1}{\rho} \hat{g}_{\mu\nu} dx^\mu dx^\nu + \frac{d\rho^2}{4\rho^2} \quad \text{where } \rho \equiv \frac{r_0^2}{r^2} \quad (\text{and } R = 1)
 \end{aligned}$$



By expanding 4D metric by powers of ρ ($= \frac{r_0^2}{r^2}$) as

$$\hat{g}_{\mu\nu} = g_{(0)\mu\nu} + g_{(2)\mu\nu} \rho + \rho^2 (g_{(4)\mu\nu} + h_{1(4)\mu\nu} \log \rho + h_{2(4)\mu\nu} (\log \rho)^2) + \dots$$

the Energy momentum tensor of the **UV** boundary ($r \rightarrow \infty$) AdS₄ field theory are given by following formula.

$$\begin{aligned}
 \langle T_{\mu\nu} \rangle &= \frac{4R^3}{16\pi G_N} (g_{(4)\mu\nu} - \frac{1}{8} g_{(0)\mu\nu} \left((\text{Tr} g_{(2)})^2 - \text{Tr} g_{(2)}^2 \right) - \frac{1}{2} (g_{(2)}^2)_{\mu\nu} \\
 &\quad + \frac{1}{4} g_{(2)\mu\nu} \text{Tr} g_{(2)})
 \end{aligned}$$

(S.de Haro, et al. 2000)

Trace anomaly by Holography

Then, we can get

$$\langle T_{\mu\nu} \rangle = \frac{4R^3}{16\pi G_N^5} \left(\frac{3\lambda^2}{16} (-\mathbf{1}, g_{0ij}) \right)$$

$g_{(0)ij} \equiv a_0(t)^2 \gamma_{ij}(x)$: boundary AdS_4 space metric

Trace anomaly is obtained as

$$\langle T^\mu{}_\mu \rangle = -\frac{3\lambda^2}{8\pi^2} N^2$$

which is the same as the result of the field theory in AdS_4 space-time .

quark-antiquark potential in AdS_4

Quark-Antiquark potential in the AdS_4 field theory is obtained by the energy of **U-shaped string** in the 10D bulk.

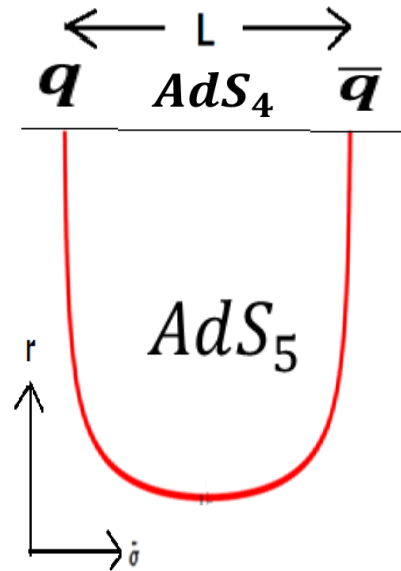
$$ds^2 = \frac{r^2}{R^2} \left(1 + \frac{r_0^2}{r^2} \right)^2 (-dt^2 + a_0^2(t) \gamma_{ij}(x) dx^i dx^j) + \frac{R^2}{r^2} dr^2 + R^2 d\Omega_5^2$$

The Nambu-Goto action of the U-shaped string :

$$L_{NG} = -\frac{1}{2\pi\alpha'} \int d\sigma \left(\frac{rA}{R} \right)^2 \sqrt{r'^2 + \left(\frac{r}{R} \right)^4 (A(r) a_0(t) \gamma(x))^2}$$

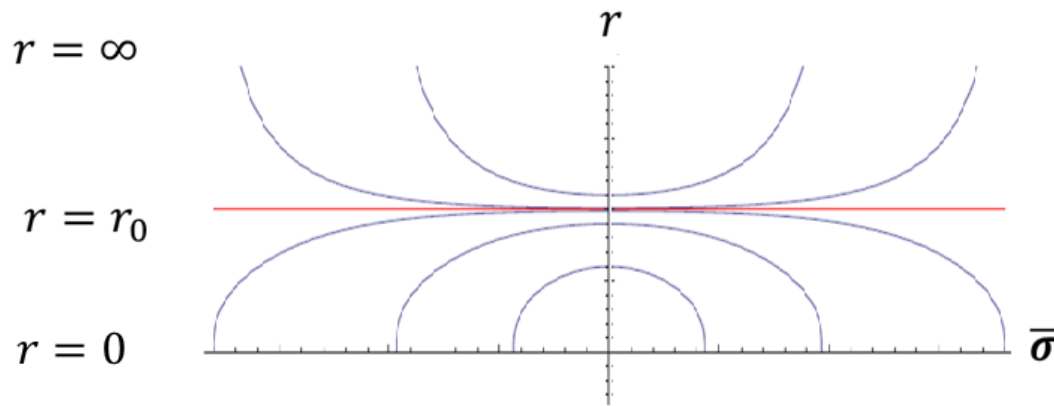
where we take the string world-volume coordinates (τ, σ) as

$$\tau = t, \quad \sigma = x$$



Numerical solution and domain wall

U-shaped strings given by solving the equation of motion numerically.



There are two classes of U-shaped strings.

- U-shaped strings whose endpoints are on the UV boundary ($r \rightarrow \infty$)
- U-shaped strings whose endpoints are on the IR boundary ($r \rightarrow 0$)

They are separated by domain walls (**red line**: $r = r_0$).

The $q\bar{q}$ potential and confinement

We consider the quark-antiquark ($q\bar{q}$) potential on UV-boundary

By introducing the proper distance as $\bar{\sigma} \equiv a_0(t) \int d\sigma \frac{1}{1-\frac{\sigma^2}{4}}$

energy of U-shaped string becomes

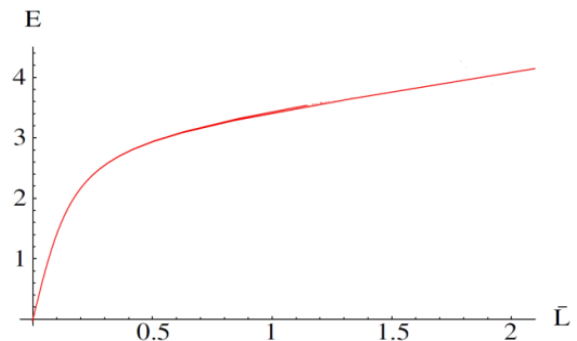
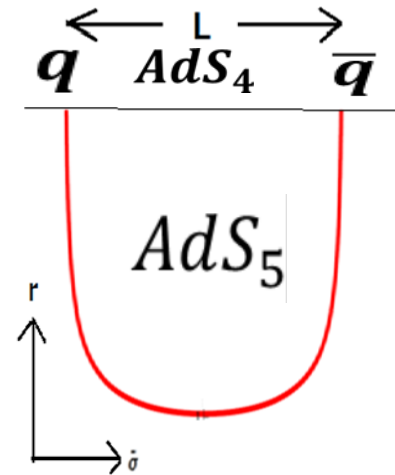
$$E = -L_{NG} = \frac{1}{2\pi\alpha'} \int d\bar{\sigma} \left(\frac{rA}{R}\right)^2 \sqrt{1 + \left(\frac{R^2}{r^2 A} \partial_{\bar{\sigma}} r\right)^2}$$

We can get the linear potential and $q\bar{q}$ are confined.

$$E \sim \frac{(r_* A(r_*))^2}{2\pi\alpha' R^2} \int d\bar{\sigma} = \tau_{q\bar{q}} L \quad (\partial_{\bar{\sigma}} r|_{r=r_*} = 0)$$

$$L = \int_{\bar{\sigma}_{min}}^{\bar{\sigma}_{max}} d\bar{\sigma} \quad \text{the distance between quark ant antiquark}$$

$$\tau_{q\bar{q}} \equiv \frac{(r_* A(r_*))^2}{2\pi\alpha' R^2} \quad \tau_{q\bar{q}} \text{ tension}$$



D7 brane embedding

D7-brane: mesons in dual field theory on AdS_4 (*A. Karch and E.Katz, 2002*)
(*M.Kruczenski et al. 2003*)

We consider the D7-brane embedding in the 10D bulk with AdS_4 boundary.

$$ds^2 = \frac{r^2}{R^2} \left(1 + \frac{r_0^2}{r^2} \right)^2 (-dt^2 + a_0^2(t) \gamma_{ij}(x) dx^i dx^j) + \frac{R^2}{r^2} dr^2 + R^2 d\Omega_5^2$$

X^0 X^1 X^2 X^3 X^4 X^5 X^6 X^7 X^8 X^9

D3-brane : ○ ○ ○ ○

D7-brane : ○ ○ ○ ○ ○ ○ ○ ○

D7 brane action

The induced metric for the D7-brane

$$ds_{D7}^2 = \frac{r^2}{R^2} \left(\mathbf{1} + \frac{r_0^2}{r^2} \right)^2 (-dt^2 + a_0^2(t) \gamma_{ij}(x) dx^i dx^j) \\ + \frac{R^2}{r^2} \left(\left(\mathbf{1} + (\partial_\rho w)^2 \right) d\rho^2 + \rho^2 d\Omega_3^2 \right)$$

where $\rho^2 = \sum_{i=4}^7 (X^i)^2$, $r^2 = \rho^2 + w(\rho)^2$

and we set $(X^8, X^9) \equiv (w(\rho), \mathbf{0})$ due to the rotational invariance in the $X^8 - X^9$ plane

the DBI action of the D7-brane :

$$S_{D7} = -T_7 \int d^8 \xi \sqrt{-\det(g_{ab})} \\ = -T_7 \int d^8 \xi \rho^3 \gamma(x)^3 a_0(t)^3 \left(\mathbf{1} + \left(\frac{r_0}{r} \right)^2 \right)^4 \sqrt{\mathbf{1} + w'(\rho)^2}$$

$w(\rho)$ determines the shape of the D7 brane.

Equation of motion for D7

The equation of motion for $w(\rho)$

$$\frac{w}{\rho + ww'} \left(\phi' - \sqrt{1 + w'^2} (\phi + 4 \log A)' \right) + \left(w' \left(\frac{3}{\rho} + (\phi + 4 \log A)' \right) + \frac{w''}{1 + w'^2} \right) = 0$$

From the profile solution $w(\rho)$, we can read the quark mass m_q and the VEV of chiral condensate $c \equiv \langle \bar{\psi} \psi \rangle$

$$w(\rho) = m_q + \frac{c}{\rho^2} + \dots (\text{at large } \rho)$$

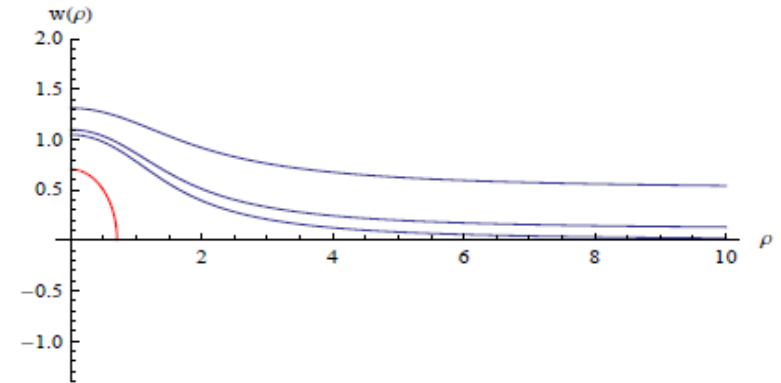
D7 Solution and chiral symmetry breaking

Solution is given numerically

$$w(\rho) = m_q + \frac{c}{\rho^2} + \dots$$

$c \equiv \langle \bar{\psi}\psi \rangle$: VEV of chiral condensates

m_q : mass of a quark



Solutions $w(\rho)$ with $m_q = 0$ $c > 0$ mean that Chiral symmetry is spontaneously broken

Red circle ($r = r_0$) is the domain wall.

The trivial solution $w(\rho) = 0$ ($m_q = 0, c = 0$) also exists. However, its energy is larger than nontrivial solution $w(\rho)$ with $m_q = 0, c > 0$

Meson spectrum

Meson: fluctuation of the D7-brane solution $(X^8, X^9) = (\mathbf{w}(\boldsymbol{\rho}), \mathbf{0})$

$$X^8 = \mathbf{w}(\boldsymbol{\rho}) + \tilde{\phi}^8(t, \mathbf{x}, \boldsymbol{\rho}, S^3) \quad X^9 = \mathbf{0} + \tilde{\phi}^9(t, \mathbf{x}, \boldsymbol{\rho}, S^3)$$

We write the functions in the following factorized form.

$$\tilde{\phi}^k = \psi^k(t, \mathbf{x}) \phi_l(\boldsymbol{\rho}) Y_l(S^3) \quad (k = 8, 9 \quad l: \text{angular momentum})$$

$Y_l(S^3)$: spherical harmonic function on S^3

$$-\square_4 \psi^k(t, \mathbf{x}) = -m_k^2 \psi^k(t, \mathbf{x}) \quad (k=8, 9)$$

m_k ($k = 8, 9$) : The meson mass spectrum

Meson spectrum

Equation of motion for $\phi_l^9(\rho)$

$$\partial_\rho^2 \phi_l^9 + \frac{1}{L_0} \partial_\rho(L_0) \partial_\rho \phi_l^9 + (1 + w'^2) \left(\frac{\left(\frac{R}{r}\right)^4 m_9^2}{A^2} - \frac{l(l+2)}{\rho^2} - 2K_{(1)} \right) \phi_l^9 + \frac{1}{r} \frac{\partial \Phi}{\partial r} \phi_l^9 = 0$$

$$L_0 = \frac{\rho^3 A^4}{\sqrt{1+w'^2}} \quad K_{(1)} = \frac{\partial_r A^4}{A^4}$$

Equation of motion for $\phi_l^8(\rho)$

$$\begin{aligned} \partial_\rho^2 \phi_l^8 + \frac{1}{L_1} \partial_\rho(L_1) \partial_\rho \phi_l^8 + (1 + w'^2) \left(\frac{\left(\frac{R}{r}\right)^4 m_8^2}{A^2} - \frac{l(l+2)}{\rho^2} - 2(1 + w'^2)(K_{(1)} + 2w'^2 K_{(2)}) \right) \phi_l^8 \\ + (1 + w'^2)^{\frac{3}{2}} \left(\left(2rK_{(1)} \frac{\partial \phi}{\partial r} + \frac{\partial^2 \phi}{\partial r^2} \right) \frac{w^2}{r^2} + \frac{\partial \phi}{\partial r} \frac{\rho^2}{r^3} \right) \phi_l^8 = -2 \frac{1}{L_1} \partial_\rho(L_0 w w' K_{(1)}) \phi_l^8 \end{aligned}$$

$$L_{(1)} = \frac{L_0}{1+w'^2} \quad K_{(2)} = \frac{1}{A^4} \partial_r^2 A^4$$

Meson mass spectrum (simple case)

For simple case: Trivial solution $w(\rho) = 0$

$$\phi_l^8(\rho) = \phi_l^9(\rho) \equiv \phi_l(\rho) \quad \text{and} \quad m_8 = m_9 \equiv m$$

Equation of motion for $\phi_l(\rho)$

$$\partial_\rho^2 \phi_l + \left(\frac{3}{\rho} - \frac{3r_0^2}{A\rho^3} \right) \partial_\rho \phi_l + \left(\frac{m^2 R^4}{A^2 \rho^4} - \frac{l(l+2)}{\rho^2} + \frac{3r_0^2}{A\rho^4} \right) \phi_l = 0$$

Solution is obtained analytically in this case.

$$\phi_l = (r_0^2 + \rho^2)^{\frac{-3 + \sqrt{9 + \tilde{m}^2}}{2}} \times \left(c_1 \rho^{4+l} F \left(\alpha, \alpha + l + 1, l + 2, -\frac{\rho^2}{r_0^2} \right) + c_2 \rho^{2-l} F \left(\alpha, \alpha - l - 1, -l, -\frac{\rho^2}{r_0^2} \right) \right)$$

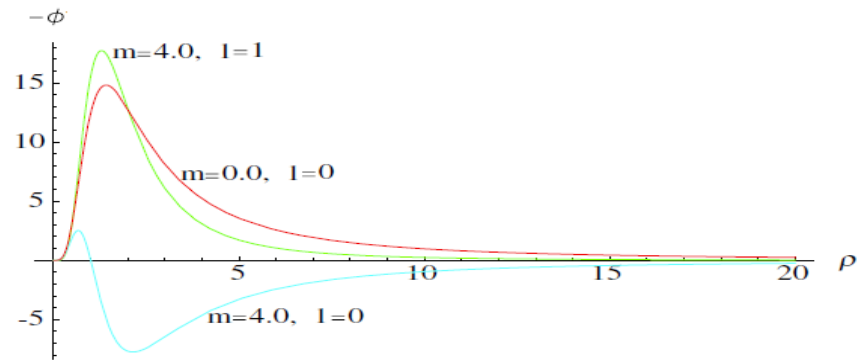
$$\alpha = \frac{1}{2} \left(1 - \sqrt{9 + \tilde{m}^2} \right) \quad \tilde{m} = \frac{R^2 m}{r_0}$$

c_1 c_2 : arbitrary constants F : hypergeometric function

Meson mass spectrum (simple case)

From the regularity of $\phi_l(\rho)$,

$$c_2 = 0 \text{ and } \alpha + l + 1 = -n \quad (n = 0, 1, 2 \dots)$$



Then, the meson mass spectrum becomes

$$m_s^2 = -\lambda I(I - 3) \quad I \equiv n + l + 3 \quad (n, l = 0, 1, 2 \dots)$$

This is the same spectrum as the discrete mass spectrum of the scalar fields obtained by the field theory in AdS_4

$$(m_s^2 = -\lambda I(I - 3) \quad I = 3, 4 \dots \text{ *Avis, Isham and Storey 1978*}).$$

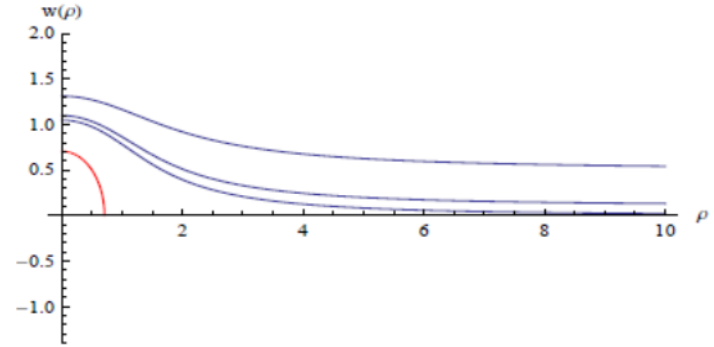
Meson mass spectrum

Meson spectrums m_ρ for the non-trivial solutions

$$w(\rho) = m_q + \frac{c}{\rho^2} + \dots$$

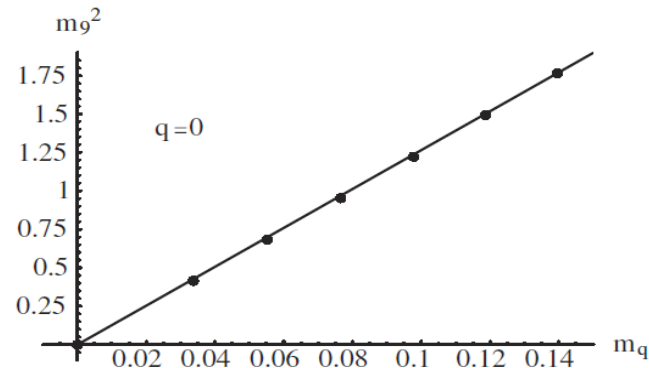
$$(c \equiv \langle \bar{\psi}\psi \rangle \neq 0)$$

m_q : quark mass



m_ρ ($l = n = 0$): mass of the NG boson by spontaneous chiral symmetry breaking

We can get the Gellmann-Oaks Renar relation ($m_\rho^2 \propto m_q$) for small m_q



Glueball mass spectrum

Glueball spectrum can be obtained by the 5D bulk metric fluctuation $\mathbf{h}_{ij}(\mathbf{t}, \mathbf{x}^i, \mathbf{r})$ *(R.C. Brower, S.D.Mathur and C.I. Tan. 2003)*

$$\frac{1}{\sqrt{-g}} \partial_M (\sqrt{-g} g^{MN} \partial_N \mathbf{h}_{ij}) = \mathbf{0}$$

By decomposing $\mathbf{h}_{ij}(\mathbf{x}^\mu, \mathbf{r}) = \mathbf{p}_{ij} \chi(\mathbf{x}^\mu) \phi(\mathbf{r})$

The equation of 4D part $\chi(\mathbf{x}^\mu)$ is given by

$$\frac{1}{g_4} \partial_\mu \sqrt{g_4} g^{\mu\nu} \partial_\nu \chi(\mathbf{x}^\mu) = m^2 \chi(\mathbf{x}^\mu)$$

m : Glueball mass

Glueball mass spectrum

By defining $x \equiv \frac{r}{r_0}$, Equation for $\phi(r)$ is given by

$$\partial_x^2 \phi + g_2(x) \partial_x \phi + \frac{R^4 m^2}{r_0^2 x^4 A^2(x)} \phi = 0$$

$$\text{Where } g_2(x) = \frac{1}{x} \left(5 - \frac{8}{x^2 A(x)} \right) \quad A(x) = 1 + \frac{1}{x^2}$$

ϕ becomes normalizable by choosing m as

$$m^2 = -\lambda(n+1)(n+4) \quad -\lambda = \frac{4r_0^2}{R^4} \quad n = 0, 1, 2 \dots$$

This is the similar spectrum to the discrete mass spectrum of the scalar fields obtained by the field theory in AdS_4

($m_s^2 = -\lambda I(I-3)$ $I = 3, 4 \dots$ *Avis, Isham and Storey 1978*).

However, the lowest glueball mass is different as $m^2 = -4\lambda (> 0)$.

Summary

We consider 10D bulk which has AdS_4 boundaries.

There is IR boundary at $r \rightarrow 0$ and UV and IR boundaries are separated by a domain wall in the bulk at $r = r_0$

From the holographic energy momentum tensor, we find the Weyl anomaly which is the same as the result of field theory in AdS_4 space-time.

By calculating the embedding solutions of strings and D7-branes in the 10D bulk, We found the quark confinement, spontaneous chiral symmetry breaking, meson mass spectrum which shows the Gellmann-Oaks Renar relation.

Glueball spectrums are also calculated by the bulk gravity fluctuations.

Current and future works

We consider the effect both **4D cosmological constant** and **temperature** by solving the 5D Einstein equation in the following ansatz

$$ds_5^2 = \frac{r^2}{R^2} \left(-n(r) dt^2 + A^2(r) a_0^2(t) \gamma_{ij}(x) dx^i dx^j \right) + \frac{R^2}{r^2} dr^2$$

$n(r)$ and $A(r)$ as determined as

$$n(r) = \frac{\left(\left(1 + \left(\frac{r_0}{r} \right)^2 \right)^2 - c_0 \left(\frac{R}{r} \right)^4 \right)}{A} \quad A(r) = \left(\left(1 + \left(\frac{r_0}{r} \right)^2 \right)^2 + c_0 \left(\frac{R}{r} \right)^4 \right)^{1/2}$$

c_0 : temperature of dual Yang-Mills theory (integration constant)

$r_0 \equiv \frac{R^2}{2} \sqrt{-\lambda}$: cosmological constant of boundary 4D space-time.

(K.Ghoroku and A. Nakamura 2012) (J. Erdmenger K.Ghoroku and R. Meyer 2011)

In this case, the field theory on the UV boundary ($r \rightarrow \infty$) are different from that of the IR boundary ($r \rightarrow \mathbf{0}$).