# Analysis of the Einstein equation in the Large D limit

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Holographic vistas on Gravity and Strings@YITP, May 26-28, 2014

## Motivation

Large D limit seems to be a successful analytical method in the linear analysis of Black holes

- Gregory-Laflamme instability
- QNMs of BH (AdS/rotating/brane)
- Inst. of MPBH (bar/axissym. mode)

etc...

#### How about beyond the linear analysis?

Let us solve the Einstein Eq with 1/D expansion.

#### No interaction of BHs



Horizon can be in the arbitrary shape
→ Tractable beyond the linear regime ?

## Hierarchy at Large D



## Setup

#### **D=n+4 Static Cylindrical Ansatz**

original var.  $B = K^{n+1}$ 

Written by Two variables : A(r,v), B(r,v) (HarmarkObers'02)



## Large D limit

#### As the linear analysis,

$$\mathsf{R} = r^n / r_0^n$$

#### and

$$A = \frac{1}{r_0^2} \sum_{k \ge 0} \frac{A^{(k)}(\mathsf{R}, v)}{n^k}, \quad B = \frac{1}{r_0^{2n+2}} \sum_{k \ge 0} \frac{B^{(k)}(\mathsf{R}, v)}{n^k}.$$

Here after  $r_0 = 1$ 

Also assume  $\partial_v \sim \mathcal{O}(1)$ 

so that  $\partial_{\bar{v}} = \sqrt{n} \partial_v \ll n \partial_R$  here  $v = \sqrt{n} \bar{v}$ 

## Leading order equation

#### **Master Equation**

from  $R_{rr} = 0, \ R_{zz} = 0, \ R_{\Omega\Omega} = 0$ 



#### 2nd order ODE for $\partial_{\mathsf{R}} \ln B^{(0)}$

A is expressed by B from  $R_{\Omega\Omega} = 0$ 

$$A^{(0)} = 1 - \frac{(\mathsf{R} - 1)\mathsf{R}(\partial_{\mathsf{R}}B^{(0)})^2}{2B^{(0)2}} + \frac{(2\mathsf{R} - 1)(\partial_{\mathsf{R}}B^{(0)})}{2B^{(0)}} + \frac{(\mathsf{R} - 1)\mathsf{R}(\partial_{\mathsf{R}}^2B^{(0)})}{2B^{(0)}}$$

## Leading order Solution

LO Solution with appropriate BCs

$$A^{(0)} = 1, \quad B^{(0)} = S(v)^2$$

 $S(v) \approx$  Horizon Cross Section

#### S(v) cannot be determined in the leading order !



## Next to Leading Order

$$A = A^{(0)} + \frac{A^{(1)}}{n}, \quad B = B^{(0)} + \frac{B^{(1)}}{n}$$

#### - Ordinary linear analysis with unknown function S(v)

$$R_{rr} = 0, \ R_{zz} = 0, \ R_{\Omega\Omega} = 0$$

$$A^{(1)} = a_1(v) = 2\ln S(v) + S'(v)^2 + 2S(v)S''(v),$$
$$B^{(1)} = b^{(1)}(v) + 2S(v)^3S''(v)\ln \mathsf{R}$$

regularity@R=1

## Horizon Equation

$$A^{(1)} = a_1(v) = 2 \ln S(v) + S'(v)^2 + 2S(v)S''(v),$$
  

$$B^{(1)} = b^{(1)}(v) + 2S(v)^3 S''(v) \ln \mathbb{R}$$
  
**ODE for S(v)**  

$$R_{rz} = 0$$
  
**ODE for S(v)**

$$S'(v)(1+2S(v)S''(v)) + S(v)^2S^{(3)}(v) = 0.$$

Integration

$$a_1(v) = 2a = 2\ln S(v) + S'(v) + 2S(v)S''(v).$$

Integration



$$S(v)(-2 - 2a + 2\ln S(v) + S'(v)^2) = 2b \ge -e^a$$

#### 2 parameters a,b

#### **Potential Problem**

a can be set a = -1 ( a is a scaling)

$$S(v) \to e^{\Delta a} S(e^{\Delta a} v), \quad v \to e^{\Delta a} v, \quad b \to e^{\Delta a} b$$

$$\frac{1}{2}S'(v)^2 = -\ln S(v) + \frac{b}{S(v)} + a + 1$$

Solving  $S(v) \rightarrow Potential Problem$ 

#### Large D limit of Cylinder Spacetime

## b = 0

If b=0, S(v) has an analytic form



#### Large D limit of Cylinder Spacetime

#### Potential

#### No analytic form has found for any $b \neq 0$ But, can be understood by the Potential

$$\frac{1}{2}S'(v)^2 = -\ln S(v) + \frac{b}{S(v)}$$

b≠0 -> rescaling by |b|

$$y = S/|b| \ge 0.$$
  $\frac{|b|^2}{2}y'(v)^2 = \frac{\text{``Energy''}}{-\ln|b|} - V(y)$ 

$$V(y) = \ln y \pm \frac{1}{y}$$

#### b > 0



- One zero for every b>0
- collapse into y=0 ? ( assumption not valid, though )

 $\rightarrow$  Caged BH or BH with a waist ?

Large D limit of Cylinder Spacetime





- Minimum@y=1,(b=-1/e) : UBS
- 2 zeros for 0>b>-1/e (y\_min, y\_max)
   Oscillate between y\_min and y\_max : NUBS

Large D limit of Cylinder Spacetime

#### Horizon Cross section

Numerically Integrating from  $S_{\max}(b)$  to  $S_{\min}(b)$ 





## Summary

done

- Solved Einstein Eq. in static cylindrical ansatz
  - in the large D limit (only near horizon)
- Obtained an equation of horizon
- Solution may describe NUBS to BH

work in progress

- Check 1st Law
- Match with Asymptotics ( $\rightarrow$  Mass, Tension)

near axis (for BH solution) – Comparison with Numerics

## Future Work

- Other Ansatz : Spherical Collapse( r,z $\rightarrow$ r,t ), (A)dS, Rotating BH
- Generalized formulation
- Application to Holography

# Appendix

## Leading order solution : General

#### **General solution**

 $b_1(v), \ b_2(v), \ b_3(v)$  as arbitrary functions

$$A^{(0)} = -\frac{b_1(v)^2 b_2(v) (\mathsf{R} - 1)^{b_1(v) - 1} \mathsf{R}^{b_1(v) - 1}}{(\mathsf{R}^{b_1(v)} + b_2(v) (\mathsf{R} - 1)^{b_1(v)})^2},$$
  

$$B^{(0)} = \frac{b_2(v) b_3(v) (\mathsf{R} - 1)^{b_1(v) - 1 - \sqrt{b_1(v)^2 - 1}} \mathsf{R}^{b_1(v) - 1 + \sqrt{b_1(v)^2 - 1}}}{(\mathsf{R}^{b_1(v)} + b_2(v) (\mathsf{R} - 1)^{b_1(v)})^2}.$$

 $b_1, b_2, b_3$  should satisfy  $(R_{rz} = 0)$ 

$$\frac{b_2(z)\left(2b_3(z)b_1'(z) + b_1(z)\left(b_1(z)^2 - 1\right)b_3'(z)\right) + \sqrt{b_1(z)^2 - 1}b_1(z)^2b_3(z)b_2'(z)}{(\mathsf{R} - 1)\mathsf{R}b_1(z)b_2(z)b_3(z)\sqrt{b_1(z)^2 - 1}} = 0$$

#### Leading order : Asymptotics

★ cannot be solved directly. Instead, we focus on the asymptotics

$$A^{(0)} = -\frac{b_1(z)^2 b_2(z)}{(1+b_2(z)^2)\mathsf{R}^2} + \mathcal{O}(\mathsf{R}^{-3}) \quad (b_2(z) \neq -1)$$
$$A^{(0)} = 1 + \frac{1-b_1(z)}{\mathsf{R}} + \mathcal{O}(\mathsf{R}^{-2}) \quad (b_2(z) = -1).$$

looks better ?

$$b_2(z) = -1$$

#### Leading order Solution

Substituting 
$$b_2(z) = -1$$
 into  $\bigstar$ 

$$2b_3(z)b_1'(z) + b_1(z)(b_1(z)^2 - 1)b_3'(z) = 0$$

$$b_3(z) = \frac{b_1(z)^2 C}{b_1(z)^2 - 1} \quad \text{or} \quad b_1(z) = 1.$$

For now, we take  $b_1(z) = 1$ .

$$A^{(0)} = 1, \ B^{(0)} = b_3(z) = S(z)^2$$

## **NUBS from UBS**

#### **Expansion from UBS (b = -1/e)** $\overline{b} = eb + 1 \ll 1$

 $eS(v) = 1 + \bar{b} - \frac{\bar{b}^2}{12} + \left(1 - \frac{55\bar{b}}{144} + \frac{2347\bar{b}^2}{20736}\right) 2^{1/2}\bar{b}^{1/2}\cos(v/L)$  $- \left(\frac{2\bar{b}}{3} - \frac{5\bar{b}^2}{9}\right)\cos(2v/L) + \frac{17\bar{b}^{3/2}}{24\sqrt{2}}\cos(3v/L) - \frac{247\bar{b}^2}{540}\cos(4v/L) + \mathcal{O}(\bar{b}^{5/2})$  $\textbf{non-linear effect} \sim \bar{b}^{m/2}\cos(mv/L)$ Period $eL = 1 + \frac{\bar{b}}{12} + \frac{\bar{b}^2}{576} + \mathcal{O}(\bar{b}^{5/2})$ 

#### **Recover the scaling**

$$S(v) \to (e/\lambda) S_{eb/\lambda}(ev/\lambda), \quad a \to -\ln\lambda$$