

Minimal Supergravity and Inflation with a Large Cutoff

Izawa K.-I.

with

M. Ibe, Y. Shinbara, T. Yanagida

Low-Energy Supersymmetry (sparticle spectrum < 5TeV)

**Various Candidate Models
of Elementary Particles
(beyond the standard model)**

To be minimal

MSSM + mInflation
+ ~~mSUSY~~ in mSUGRA

**speculations on orders of
characteristic parameters
and their origins**

Minimal SUGRA Inflation

ϕ inflaton superfield

$$W = v^2 \phi - \frac{\lambda}{M^2} \phi^5$$

$$K = |\phi|^2 + \frac{k}{M^2} |\phi|^4 + \dots$$

$$V=e^{\frac{K}{M_G^2}}\left\{\left(\frac{\partial^2 K}{\partial \phi \partial \phi^*}\right)^{-1}|DW|^2-3\left|\frac{W}{M_G}\right|^2\right\}$$

$$DW=\frac{\partial W}{\partial \phi}+\frac{\partial K}{\partial \phi}\frac{W}{M_G^2}$$

$$\textcolor{red}{M_G=2.4\times 10^{18} \mathrm{GeV}}$$

$$\mathcal{L}_\mathrm{kin} = \frac{1}{2} \dot{r}^2 + \frac{1}{2} r^2 \dot{\theta}^2 + \frac{1}{2} r^2 \sin^2 \theta \dot{\varphi}^2 - \frac{1}{2} \frac{1}{r^2} \partial_\theta V(r)$$

$$V(\varphi) \simeq v^4-\frac{2k}{M^2}v^4\varphi^2-\frac{5\lambda}{2M^2}v^2\varphi^4$$

$$\varphi=\sqrt{2}\mathrm{Re}~\phi$$

$$\text{spectral index} \quad n_s \simeq 1 - 8 \frac{k M_G^2}{M^2}$$

$$|n_s-1|\lesssim 0.1$$

$$\frac{k M_G^2}{M^2}\lesssim 10^{-2}$$

$$\text{loop effects} \quad \frac{k}{M^2} \; \gtrsim \; \frac{1}{16\pi^2 M_G^2}$$

mSUGRA mediation

$$K = |Z|^2 + |q_i|^2 + \frac{c^{ij}}{M^2} |Z|^2 q_i q_j^* + \dots$$

loop effects $\frac{|c^{ij}|}{M^2} \gtrsim \frac{1}{16\pi^2 M_G^2}$

$$W_{eff} = \Lambda^2 Z$$

$$\delta m_0^2 \simeq \frac{c^{ij}}{M^2} m_0^2$$

K^0 - \bar{K}^0 mixing suppressed

$$\left| \frac{\delta m_0^2}{m_0^2} \right| \lesssim 0.6 \times 10^{-2} \left(\frac{m}{\text{TeV}} \right)$$

$m_0 < 1 \text{ TeV} < m_{1/2}$ or
 $1 \text{ TeV} < m_0 < 5 \text{ TeV}$

Dark Matter Candidate

Gaugino-Higgsino mixture

$$\frac{1}{2}m_{Z^0}^2 \simeq \frac{m_{H_d}^2}{\tan^2 \beta} - m_{H_u}^2 - |\mu|^2$$

$$\tan \beta = \langle H_u \rangle / \langle H_d \rangle$$

$$\delta U = \frac{c_{1/2}}{M} Z \mathcal{W}^\alpha \mathcal{W}_\alpha$$

$$\delta K=\frac{c}{M}Z^*H\bar H$$

$$m_{1/2}\sim \frac{c_{1/2} M_G}{M} m_0 \qquad m_{\tilde H}\sim \frac{c M_G}{M} m_0$$

$$\mathcal{O}(100 {\mathrm{GeV}})$$

$$\frac{kM_G^2}{M^2} \quad \frac{c^{ij}M_G^2}{M^2} \quad \frac{c_{1/2}M_G}{M} \quad \frac{cM_G}{M}$$

$$n_s\qquad \delta m_0^2\qquad m_{1/2}\qquad m_{\tilde H}$$

$$M_G\simeq 10^{-1} M$$

Slow-Roll Parameters

$$\epsilon = M_G^2 \left(\frac{V'(\varphi)}{V(\varphi)} \right)^2 \quad \eta = M_G^2 \frac{V''(\varphi)}{V(\varphi)}$$

Inflaton Candidates

Multiple Inflations

Larger Macroscopic Universe

“Inflationary Dynamics Suppress FCNC”

**Toward a Minimal Model
of the Supersymmetric
Universe**

Baryogenesis, CP violations,

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