



Deep learning to diagnose the neutron star



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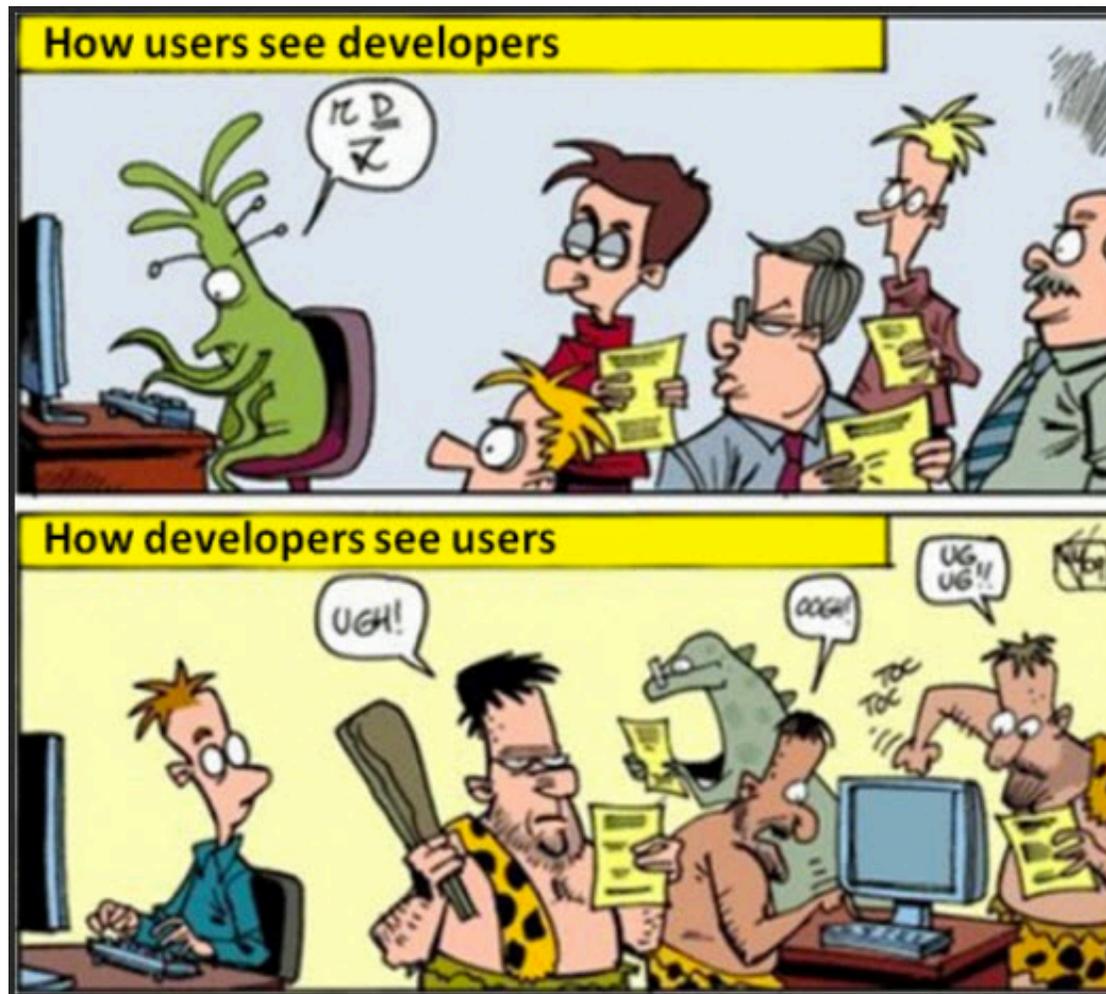
Based on work with Yuki Fujimoto, Koichi Murase

— Deep Learning and Physics 2018 —

Disclaimer



I am a “user” of the deep learning...



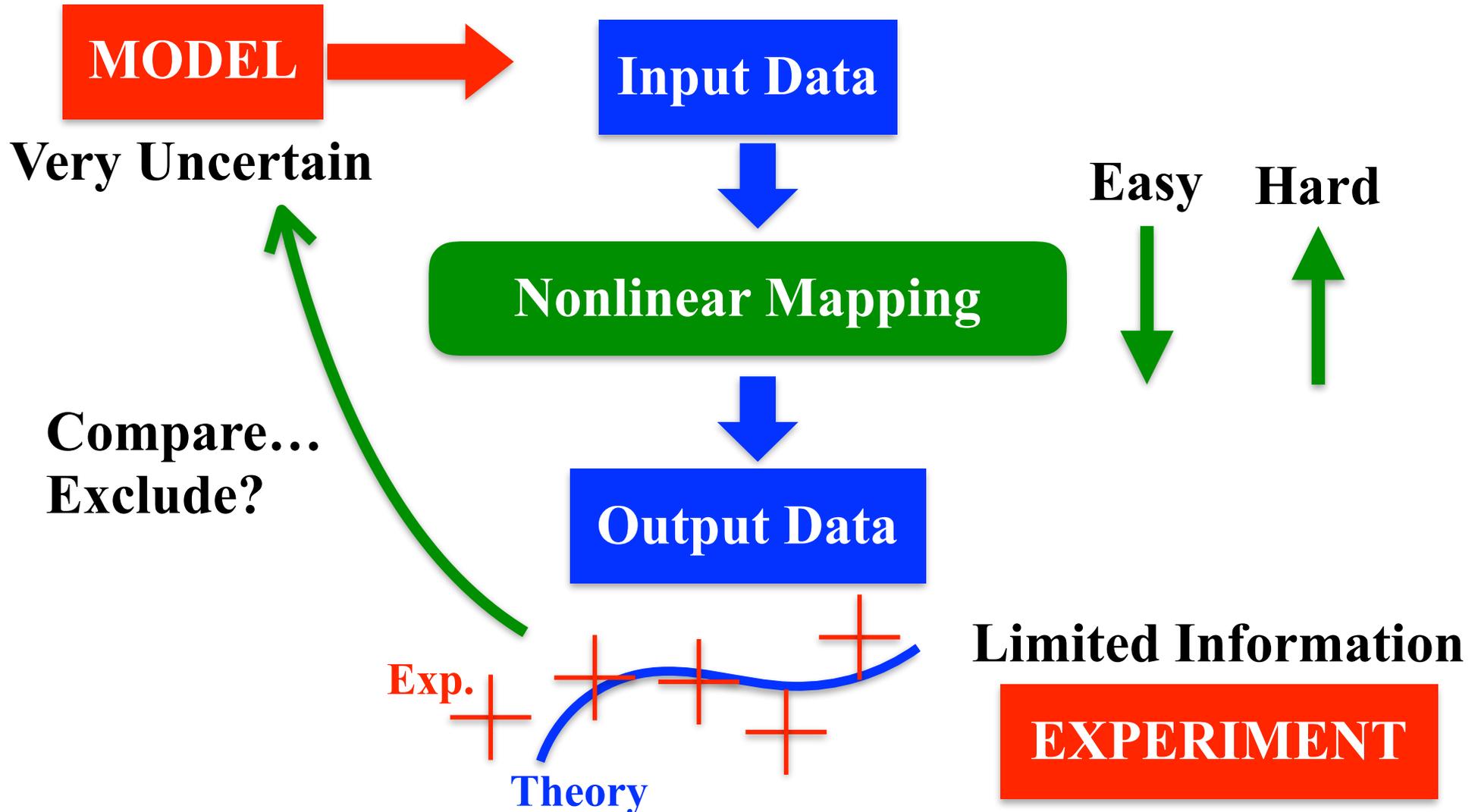
From the point of view of **physics users...**

Sounds fancy is not enough...

Useful ?

Advantageous?

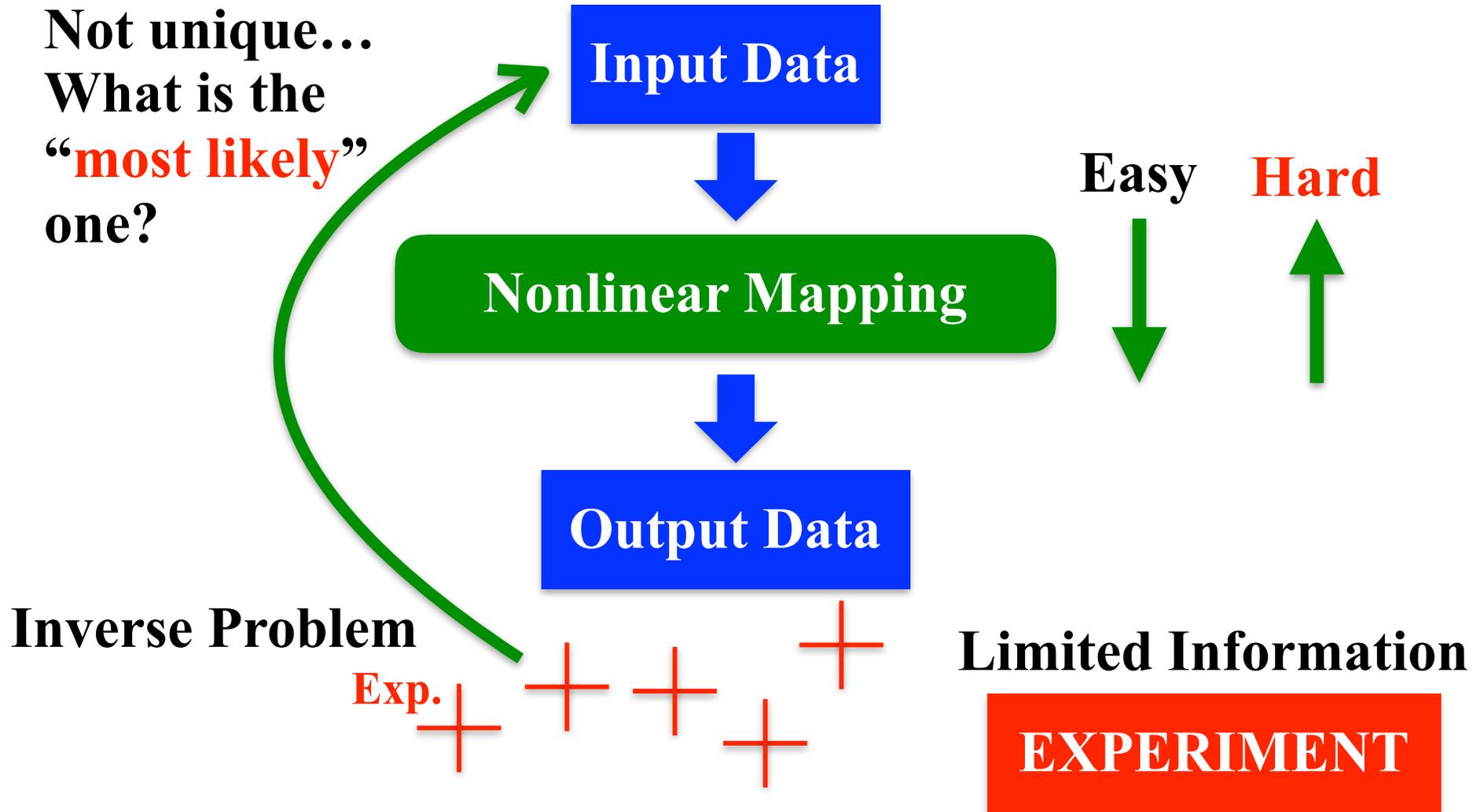
Conventional Physics Approach



One another to infinity, so what ?



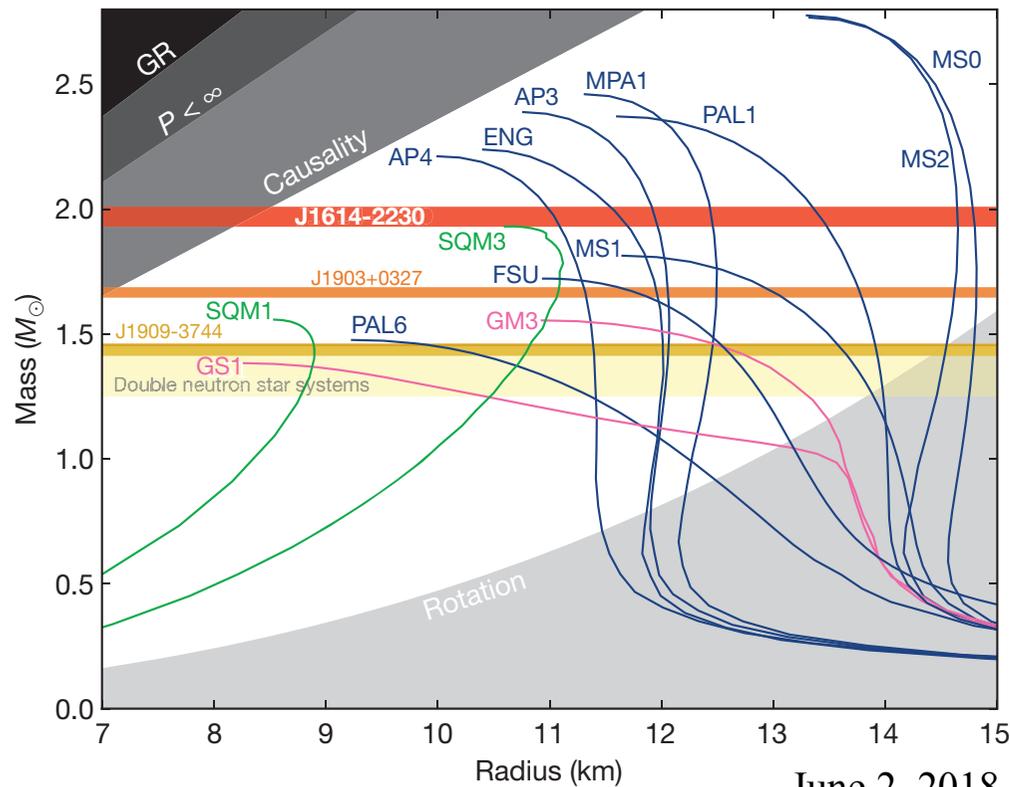
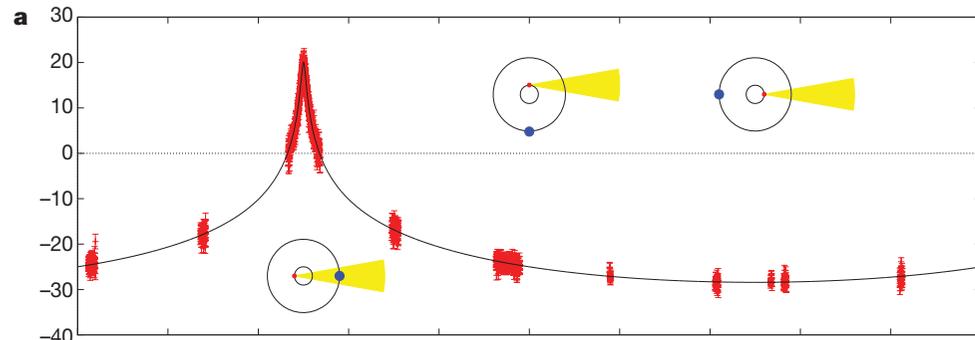
“Model Independent” Analysis



ともがく
具体的に
動いていこう
具体的に動けば
具体的な
答が出る
がら

あつた


Neutron Star EoS



Demorest et al. (2010)

Precise determination of
NS mass using Shapiro delay

1.928(17) M_{sun} (J1614-2230)
(slightly changed in 2016)

Antoniadis et al. (2013)

2.01(4) M_{sun} (PSRJ0348+0432)

Neutron Star EoS



Equation of State

Pressure : p

Mass density : ρ

(Energy density : $\varepsilon = \rho c^2$)

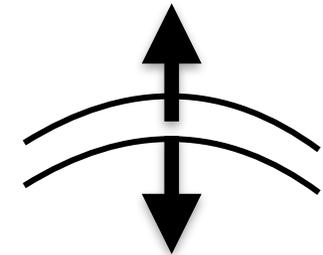
Input Data

$$p = p(\rho)$$

Nonlinear Mapping

Tolman-Oppenheimer-Volkoff (TOV) Eqs

pressure diff



gravity

M-R Relation

NS mass : M

NS radius : R

Output Data

$$M = M(\rho_{\max})$$

$$R = R(\rho_{\max})$$

Mathematically one-to-one correspondence

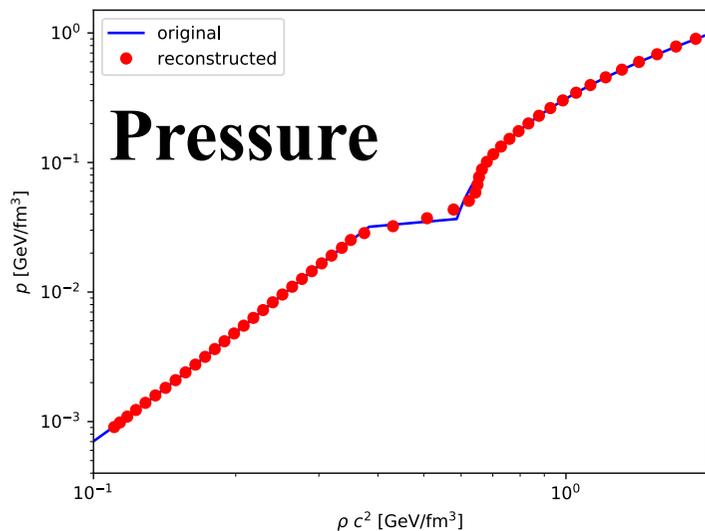
Neutron Star EoS



Lindblom (1992)

Brute-force solution of the inverse problem

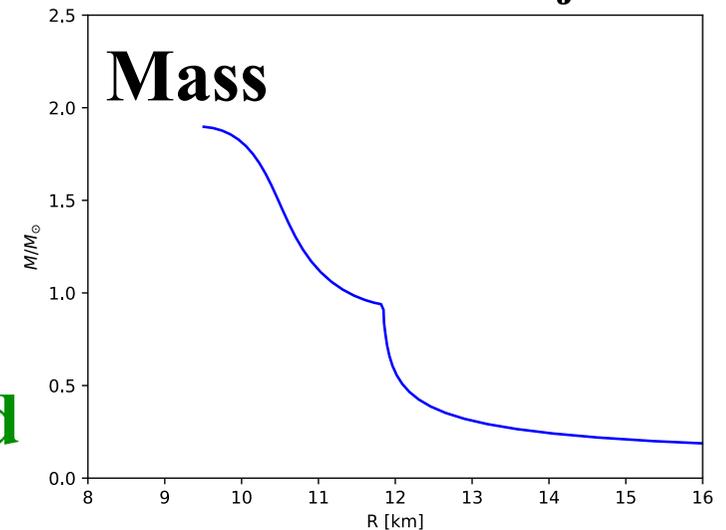
Test data put by hand



Density

Solve TOV
→
←
Reconstructed

Thanks to Y. Fujimoto



Radius

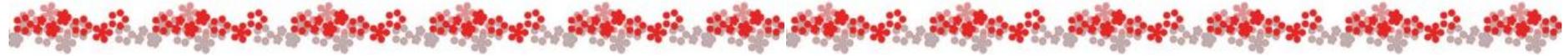
The answer exists!

No magic box...

Only “solvable” problem can be solved...

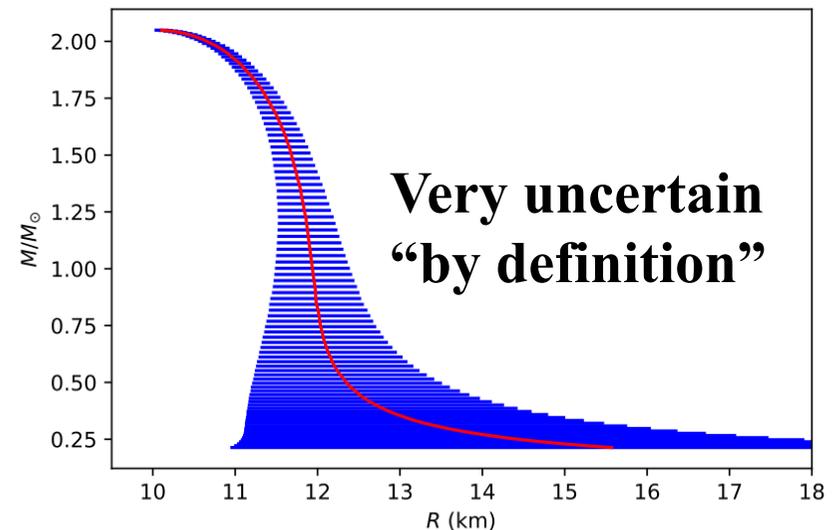


Neutron Star EoS (Side Remark)



R is fixed by TOV with $p(R)=0$ (“surface” condition)

$$\begin{aligned} dp/dr(r = R) &= 0 \\ d^2p/dr^2(r = R) &\propto M^2/R^2 \end{aligned}$$



Determination of R is very uncertain, and on top of that, R itself is anyway very uncertain...

People do not care assuming that NS mass $> 1.2 M_{\text{sun}}$

Model Independent Analysis



Bayesian Analysis

B : M - R Observation

A : EoS Parameters

(Bayes' theorem) Normalization

$$\underline{P(A|B)} \cancel{P(B)} = \underline{P(B|A)} \underline{P(A)}$$

Want to know

Likelihood

prior

Model must be assumed.

Calculable by TOV

Model

EoS parametrization must be introduced.

Integration in parameter space must be defined.

If infinite observations, prior dependence should be gone.

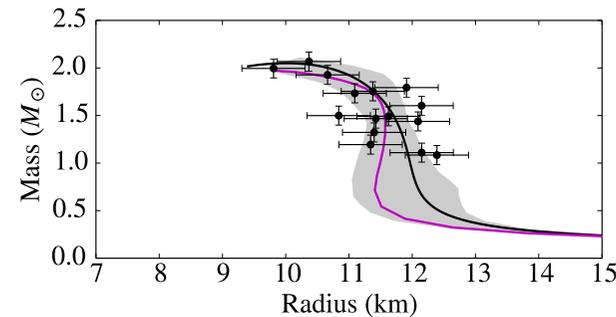
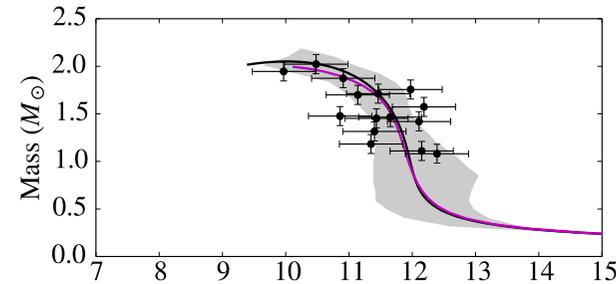
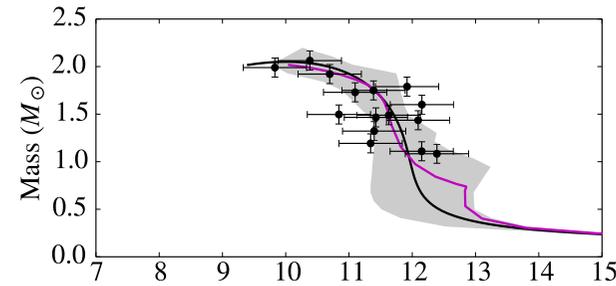
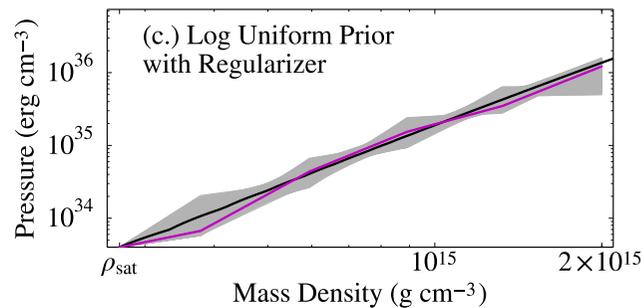
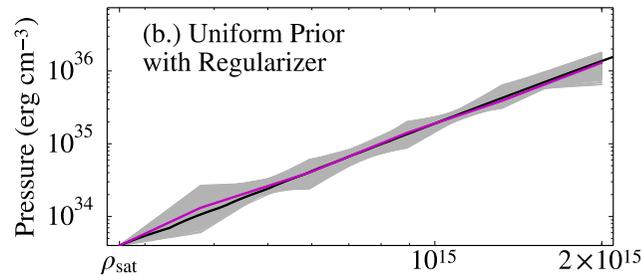
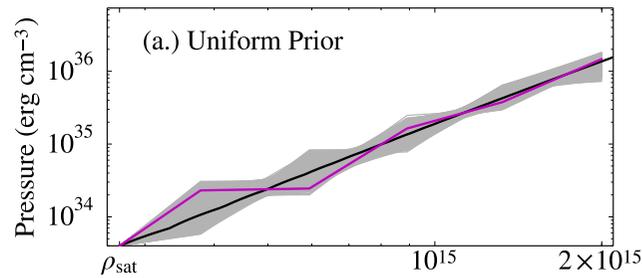
Model Independent Analysis



Raithel-Ozel-Psaltis (2017)

Mock data (SLy + Noises)

**Prior
Dep.**



**Black curve
True EoS**

**Magenta curve
Guessed EoS**

**Gray band
68% credibility**

Model Independent Analysis

Bayesian Analysis Supervised Learning



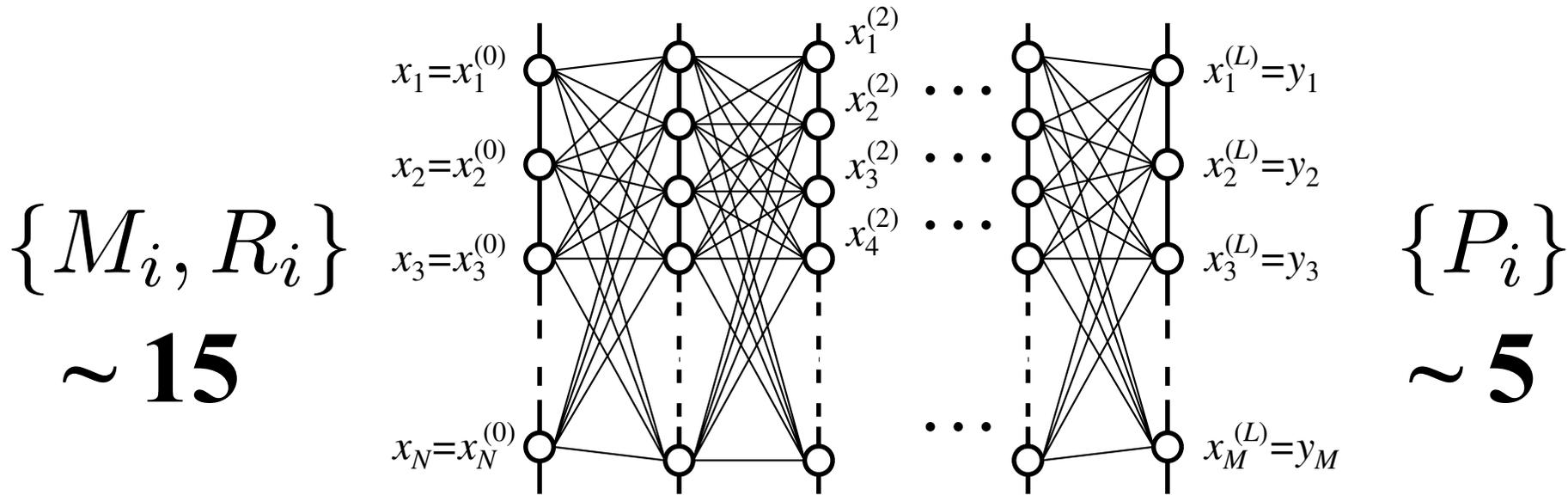
$$\{M_i, R_i\} \quad \{P_i\} = F(\{M_i, R_i\}) \quad \{P_i\}$$

~ **15** Points (observations)

~ **5** Points

**Too precise parametrization of EoS is useless
(beyond the uncertainty from observations)**

Deep Learning



$$x_i^{(k+1)} = \sigma^{(k+1)} \left(\sum_{j=1}^{N_k} \underline{W_{ij}^{(k+1)}} x_j^{(k)} + \underline{a_i^{(k+1)}} \right)$$

Backpropagation

Parameters to be tuned

~~sigmoid func.~~
 $\sigma(x) = 1/(e^x + 1)$

ReLU
 $\sigma(x) = \max\{0, x\}$

tanh
 $\sigma(x) = \tanh(x)$

Deep Learning



Our Neural Network Design

Layer index	Nodes	Activation
1	30	N/A
2	60	ReLU
3	40	ReLU
4	40	ReLU
5	5	tanh

Probably we don't need such many hidden layers and such many nodes... anyway, this is one working example...

Deep Learning



For good learning, the “textbook” choice is important...

Training data (200000 sets in total)

Randomly generate **5** sound velocities \rightarrow EoS \times 2000 sets

Solve TOV to identify the corresponding M - R curve

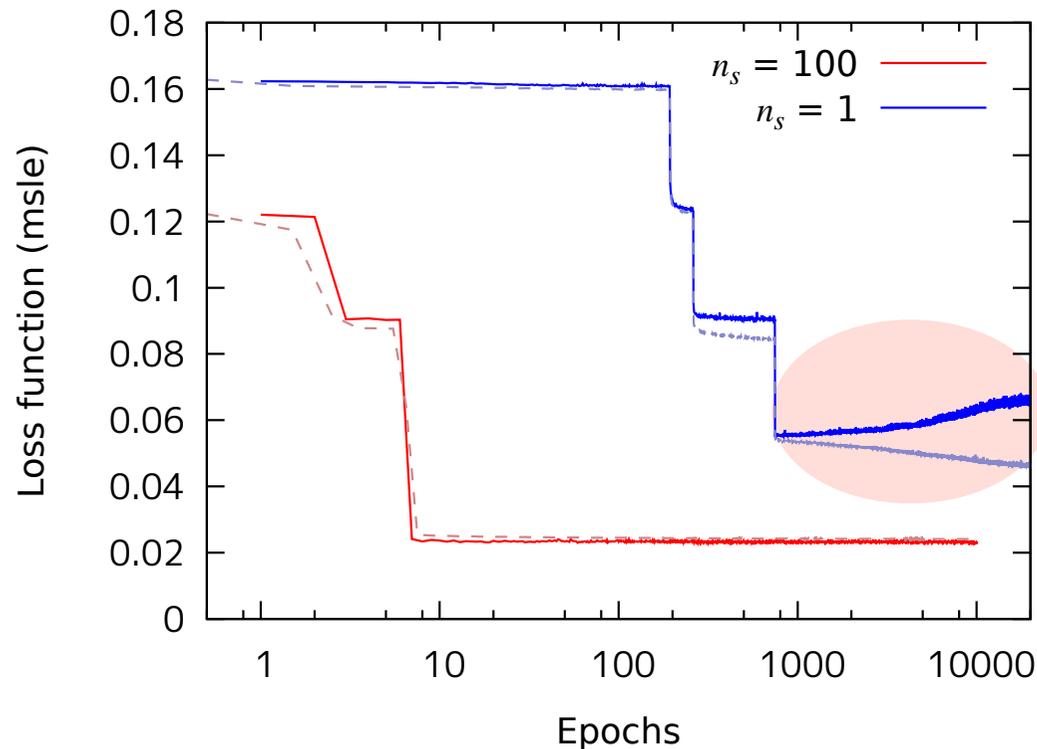
Randomly pick up **15** observation points \times ($n_s = 100$) sets
(with $\Delta M = 0.1M_{\odot}$, $\Delta R = 0.5$ km)

The machine learns the M - R data have error fluctuations

Validation data (200 sets)

Generate independently of the training data

Deep Learning



“Loss Function”
= deviation from the
true answers (msle)

Monotonically decrease
for the training data, but
not necessarily so for
the validation data

With fluctuations in the training data, the learning goes quickly

**Once the over-fitting occurs,
the model becomes more stupid...**

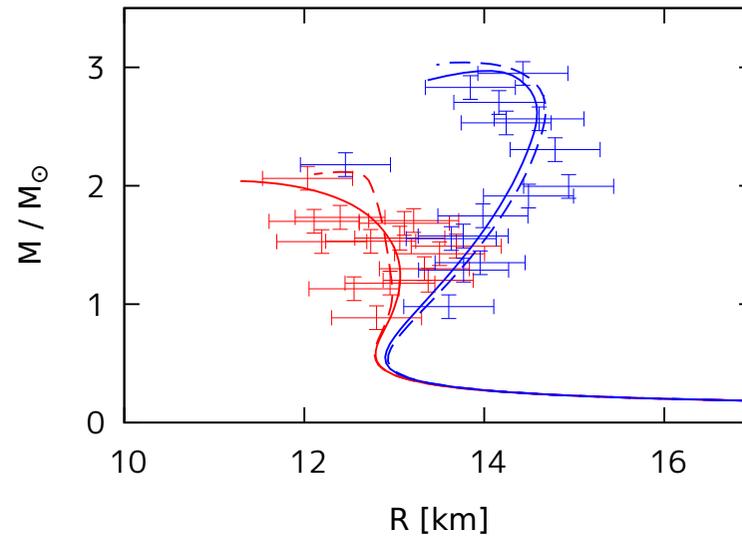
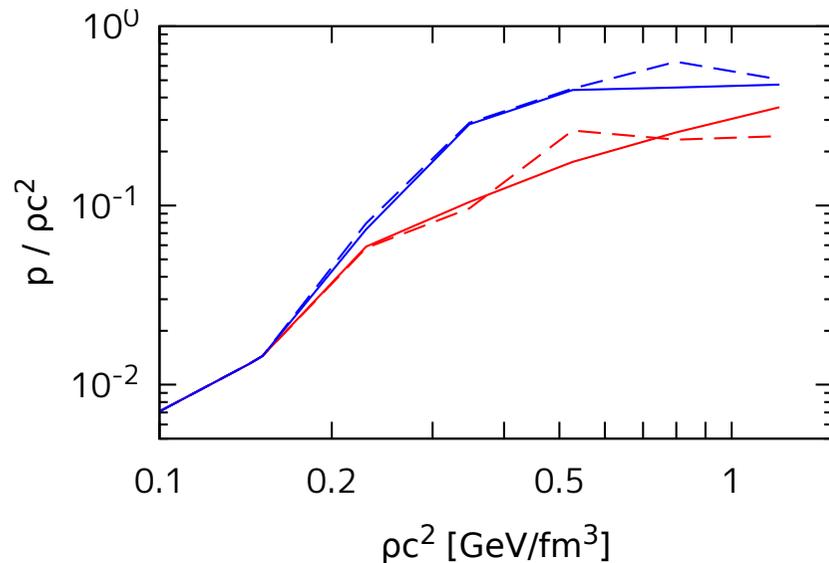


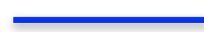
Deep Learning



Test with the validation data Fujimoto-Fukushima-Murase (2017)
(parameters not optimized to fit the validation data)

Two Typical Examples (not biased choice)



-  : randomly generated original EoS
-  : reconstructed EoS and associated $M-R$

Deep Learning



Overall performance test

Mass (M_{\odot})	0.6	0.8	1.0	1.2	1.4	1.6	1.8
RMS (km)	0.16	0.12	0.10	0.099	0.11	0.11	0.12

(with $\Delta M = 0.1M_{\odot}$, $\Delta R = 0.5$ km)

Very promising!

Credibility estimate has not been done for simplicity, but it can be included in the learning process.

**Usefulness confirmed,
easy implementation
but
advantageous ?**

**Bayesian or NN,
which to choose?**



Bayesian vs NN



EoS $\theta := \{c_{s,i}^2\}$ **Obs.** $\mathcal{D} = \{(M_i, R_i)\}$

Bayesian $f_{\text{MAP}}(\mathcal{D}) = \arg \max_{\theta} \underbrace{\text{Pr}(\theta) \text{Pr}(\mathcal{D}|\theta)}_{\text{Pr}(\theta|\mathcal{D})}$

NN minimizes $\langle \ell[f] \rangle = \int \underbrace{d\theta}_{\text{Approximated}} \underbrace{d\mathcal{D} \text{Pr}(\theta) \text{Pr}(\mathcal{D}|\theta)}_{\text{estimated}} \ell(\theta, f(\mathcal{D}))$
Approximated estimated \rightarrow Bayesian

**NN allows for more general choice of loss functions.
Bayesian assumes parametrized likelihood functions.**

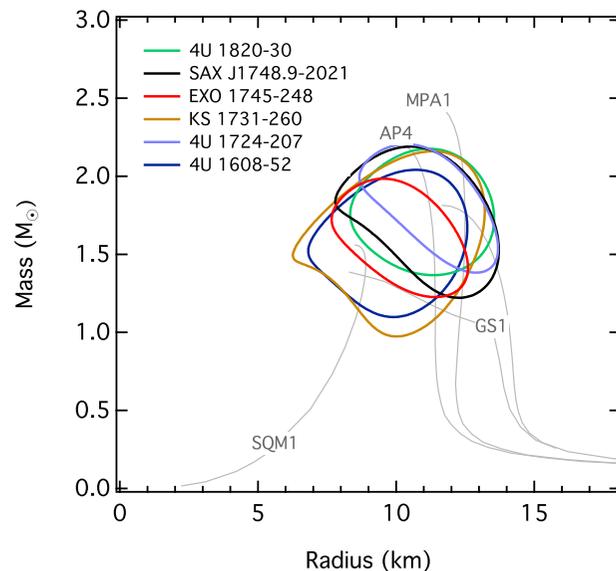
Conclusion



Yes, useful!

Maybe, less biased ?

Developing a toolkit for real data like



not discrete data with error,
but regions of credibility

Error analysis (credibility estimate)
in the output side