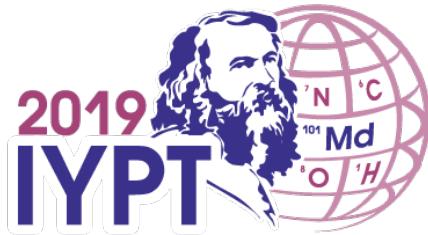


Theoretical issues in physics of SHE : nuclear reaction perspectives



International Year
of the Periodic Table
of Chemical Elements

Kouichi Hagino
Kyoto University

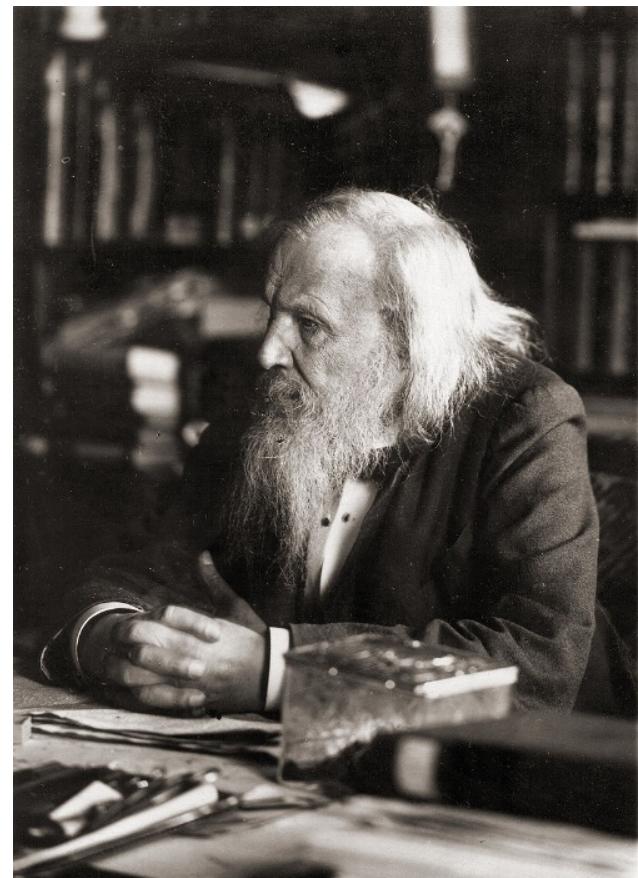


1. Introduction-1: periodic tables
2. Introduction-2: superheavy elements
3. Formation reactions
4. Theoretical issues
5. Physics of neutron-rich nuclei
6. Summary

Periodic table of elemetns (1869)

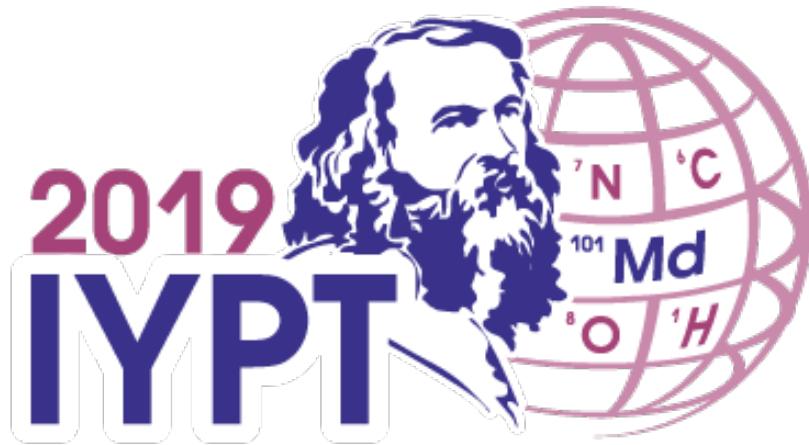


International Year
of the Periodic Table
of Chemical Elements



Mendereev
(1834-1907)

Periodic table of elements (1869)



International Year
of the Periodic Table
of Chemical Elements



closing ceremony (2019/12/5,
Tokyo)

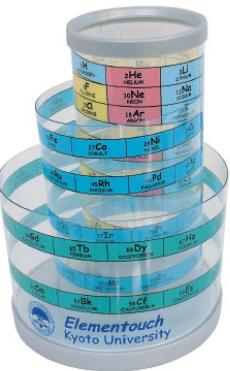
periodic table of elements

noble
gas

	Group →	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	
↓	Period	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	
1		1 H																		
2		3 Li	4 Be											5 B	6 C	7 N	8 O	9 F	2 He	
3		11 Na	12 Mg											13 Al	14 Si	15 P	16 S	17 Cl	10 Ne	
4		19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	18 Ar	
5		37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe	
6		55 Cs	56 Ba	57 La	*	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn
7		87 Fr	88 Ra	89 Ac	*	104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Nh	114 Fl	115 Mc	116 Lv	117 Ts	118 Og
		*	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu				
		*	90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr				

periodic table of elements

Group →	1	2																	
↓ Period	1	2																	
1	1 H																		
2	3 Li	4 Be																	
3	11 Na	12 Mg																	
4	19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr	
5	37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe	
6	55 Cs	56 Ba	57 La	*	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn
7	87 Fr	88 Ra	89 Ac	*	104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Nh	114 Fl	115 Mc	116 Lv	117 Ts	118 Og



“elementouch”
(Y. Maeno, 2001)

5 B	6 C	7 N	8 O	9 F	10 Ne	11 Ar	12 Kr	13 Cl	14 Si	15 P	16 S	17 Br	18 Xe	19 At	20 Og	
21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr	
37 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe	
55 La	*	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn
87 Ac	*	104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Nh	114 Fl	115 Mc	116 Lv	117 Ts	118 Og
*	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu		
*	90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr		

Prof. Yoshiteru Maeno (Kyoto U., cond. matt. expt.)

noble
gas

periodic table of elements-nuclei?

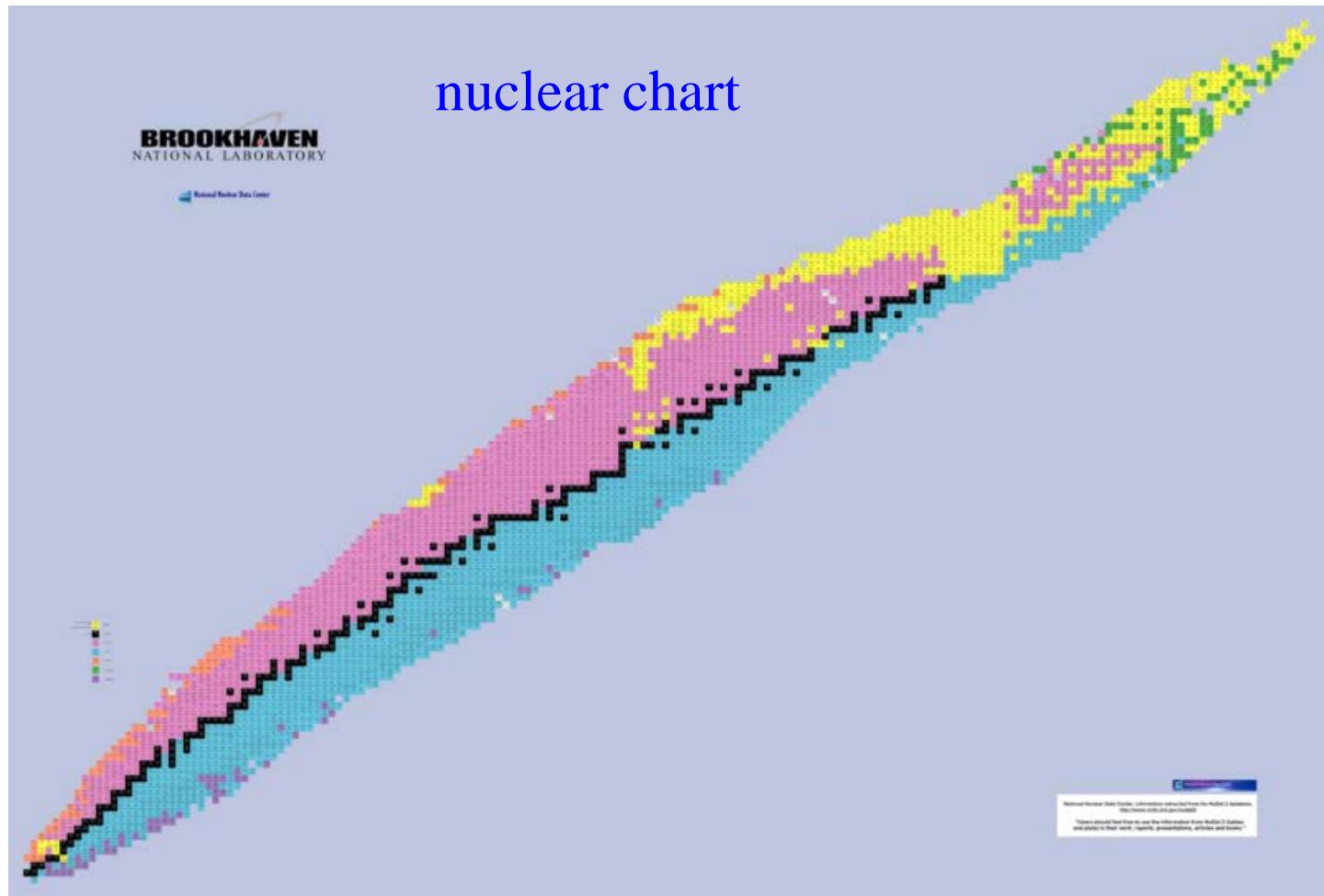
proton
magic # ← noble
gas

Group → 1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	
↓ Period																		
1	1 H																2 He	
2	3 Li	4 Be										5 B	6 C	7 N	8 O	9 F	10 Ne	
3	11 Na	12 Mg										13 Al	14 Si	15 P	16 S	17 Cl	18 Ar	
4	19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	
5	37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	
6	55 Cs	56 Ba	57 La	*	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At
7	87 Fr	88 Ra	89 Ac	*	104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Nh	114 Fl	115 Mc	116 Lv	117 Ts
	*	*	*	*	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu
	*	*	*	*	90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr



Prof. Yoshiteru Maeno (Kyoto U., cond. matt. expt.)

Nuclear periodic table

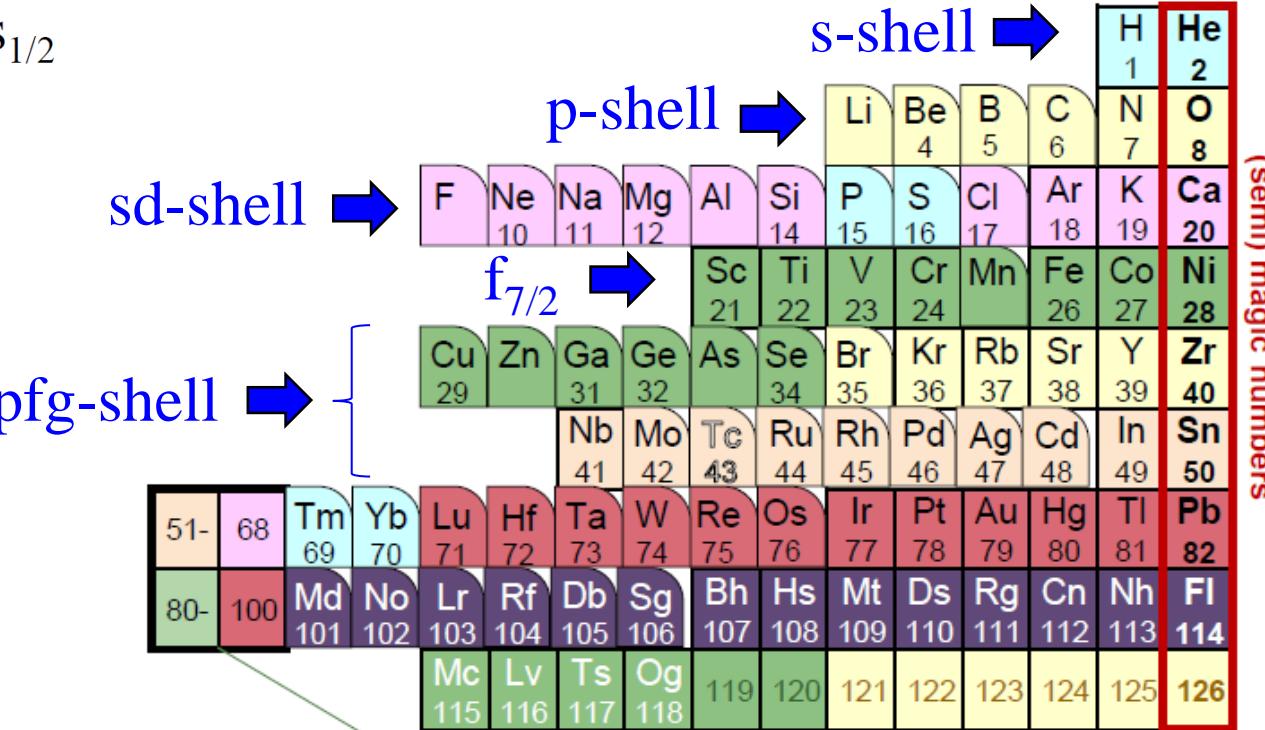
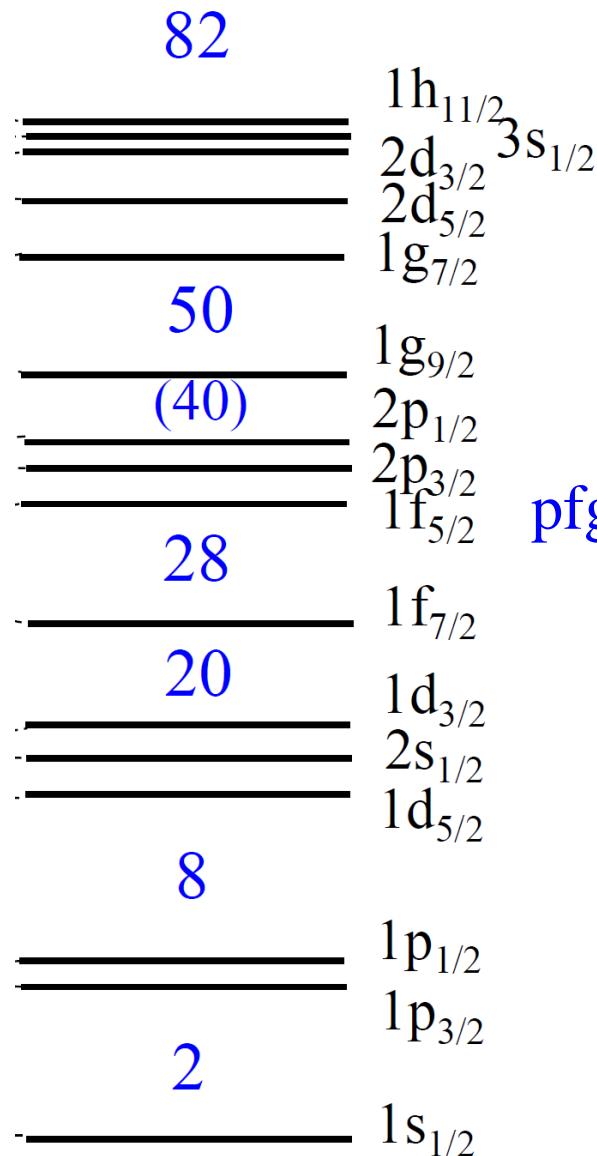


Yet, a pedagogical significance
(to familiarize nuclear physics)

Nuclear periodic table

K. Hagino and Y. Maeno,
Found. of Chem. 22, 267 (2020).

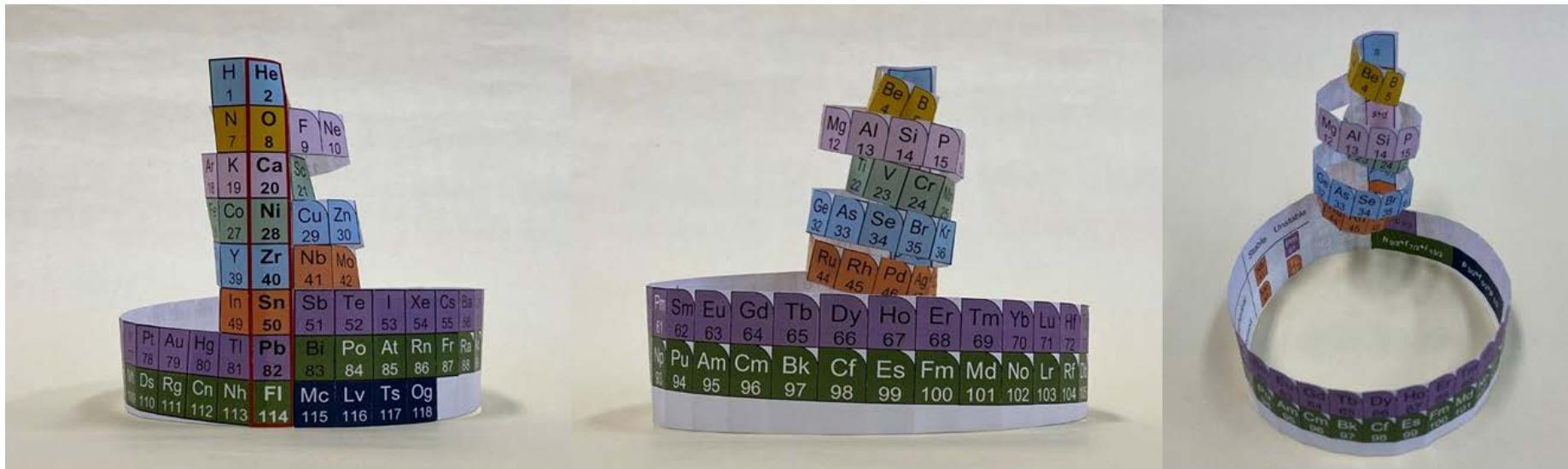
proton
magic



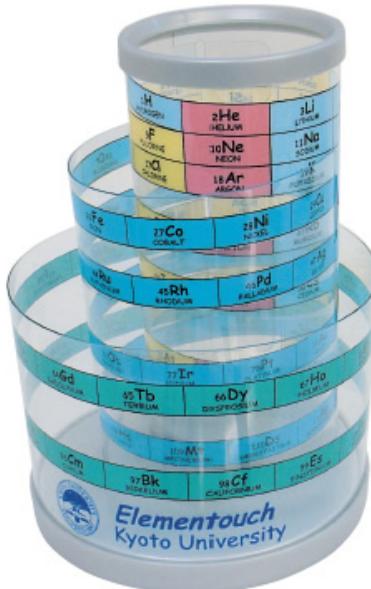
I	Xe	Cs	Ba	La	Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er
53	54	55	56	57	58	59	60	61	62	63	64	65	66	67	68
At	Rn	Fr	Ra	Ac	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm
85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100

Nuclear periodic table

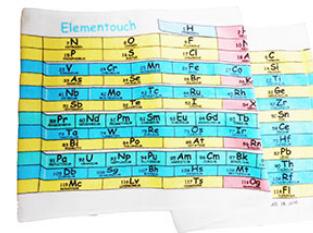
“nucletouch” (a 3D model)



cf. “elementouch”
(Y. Maeno, 2001)



mug cup



towel



T-shirt
(Kyoto-U. coop)

購読はこちら [1週間無料]

オンライン

ログイン 新規登録

ニュース > 科学・IT

「すいへーりーべ」でおなじみの元素周期表、新パターン提案…京大が原子核の状態着目

2020/04/22 23:41

日本経済新聞

朝刊



ストーリー



Myニュース 日

トップ 速報 マネー 経済・金融 政治 ビジネス マーケット テクノロジー

新しい周期表を考案 京大、原子核の性質を表現

2020/5/3付 | 日本経済新聞 朝刊

保存 共有

京都大学の前野悦輝教授と萩野浩一教授は、原子核の性質をわかりやすく見やすい周期表を考案した。従来の周期表が元素の化学的な性質を知るのに役立つて、新しい周期表は元素の原子核の性質を知るのに使える。

NEWS RELEASE 27-MAY-2020

A special elemental magic

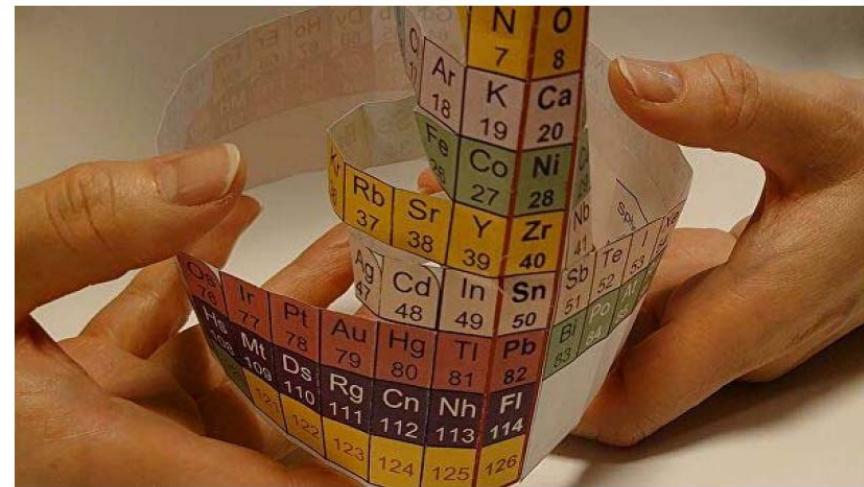
Kyoto scientists announce a 'nuclear' periodic table

KYOTO UNIVERSITY

Японские физики представили новую периодическую таблицу элементов

19:34 27.05.2020

58122



© Фото : Kyoto University/Yoshiteru Maeno/Kouichi Hagino

МОСКВА, 27 мая — РИА Новости. Ученые из Кюотского университета представили периодическую таблицу элементов, которая в отличие от таблицы Менделеева, где за основу взяты электроны в атоме, основана на

a magical coincidence

Y. Maeno, K. Hagino, and T. Ishiguro,
Found. of Chem., in press.

periodic table of elements

the nuclear magic numbers are there
in the same column!
(the same elements around them)

nuclear periodic table

			H 1	He 2	Li	Be	B		
		C	N 7	O 8	F	Ne 10	Na 11	Mg 12	
P 15	S 16	Cl 17	Ar 18	K 19	Ca 20	Sc 21	Ti 22	V 23	Cr 24
	Mn	Fe 26	Co 27	Ni 28	Cu 29	Zn 30	Ga 31	Ge 32	
Br 35	Kr 36	Rb 37	Sr 38	Y 39	Zr 40	Nb 41	Mo 42	Tc 43	Ru 44
Pd	Ag	Cd	In 49	Sn 50	Sb 51	Te 52	I 53	Xe 54	
Ir 77	Pt 78	Au 79	Hg 80	Tl 81	Pb 82	Bi 83	Po 84	At 85	Rn 86
M 101	Ds 110	Rg 111	Cn 112	Nh 113	Fl 114	Mc 115	Lv 116	Ts 117	Og 118

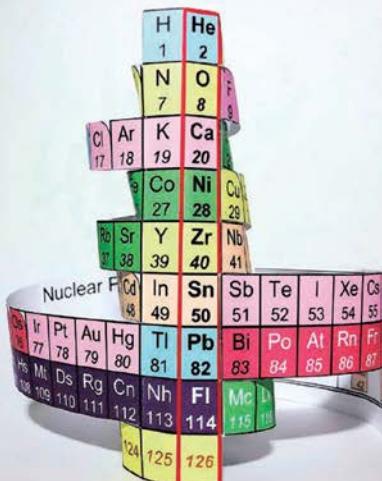
a magical coincidence which can be recognized only after making a nuclear periodic table

•磁気記録の材料と物理
•原子核の周期表

NO. 12

2020 | VOL. 75

B U T S U R I
日本物理学会誌



日本物理学会 | www.jps.or.jp



原子核の周期表——Magicな関係

前野 悅輝

〈京都大学大学院理学研究科 maeno.yoshiteru.2e@kyoto-u.ac.jp〉

萩野 浩一

〈京都大学大学院理学研究科 hagino.kouichi.5m@kyoto-u.ac.jp〉

Superheavy elements and Neutron-rich Nuclei



International Year of the Periodic Table of Chemical Elements

The figure shows a periodic table with several elements highlighted:

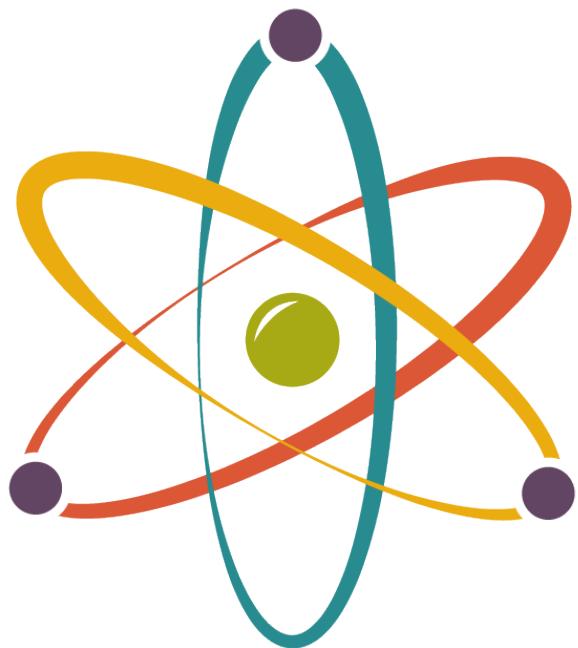
- Hydrogen (H):** Located at the top left.
- Heavier elements than Z=104:** A red box highlights the elements Rf through Nh.
- Nihonium (Nh):** A large black block highlights element 113, Nihonium, with its symbol and name.
- Actinides and Lanthanides:** The entire series of elements from Ce to Lu is highlighted with pink boxes.
- Other specific elements:** Th, Pa, U, Np, Pu, Am, Cm, Bk, Cf, Es, Fm, Md, No, and Lr are also highlighted with pink boxes.

1	1 H																2 He	
2	3 Li	4 Be															10 Ne	
3	11 Na	12 Mg															18 Ar	
4	19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	36 Kr	
5	37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	53 Te	54 Xe	
6	55 Cs	56 Ba	57 La	*	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	86 Rn	
7	87 Fr	88 Ra	89 Ac	104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Nh	114 Fl	115 Mc	116 Lv	117 Ts	118 Og
	*	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu			
	*	90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr			

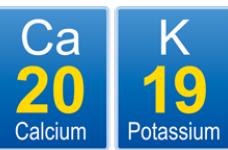
Interests in physics and chemistry of superheavy elements

- what is the heaviest element? → atomic property
- what is the double magic nucleus next to ^{208}Pb ? → nuclear property
- should the periodic table be changed or not? → chemical property
- how do superheavy elements influence the r-process nucleosynthesis?
→ astrophysics

what determines the limit of existence of elements?



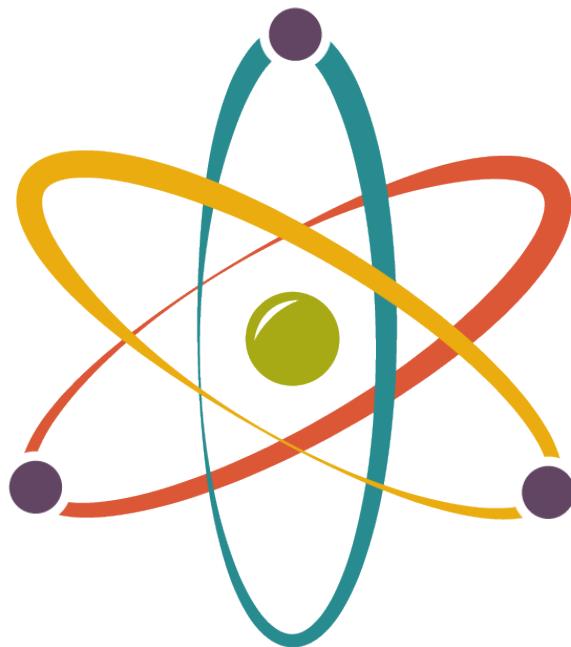
INTERNATIONAL YEAR OF THE PERIODIC TABLE **2019**



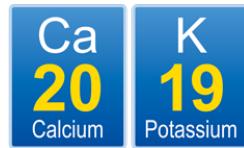
possibilities to be considered:

- ✓ electron orbitals in atom
- ✓ stability of nucleus in atom (← magic number)

what determines the limit of existence of elements?



INTERNATIONAL YEAR OF THE PERIODIC TABLE **2019**



possibilities to be considered:

let's first discuss

- ✓ electron orbitals in atom
- ✓ stability of nucleus in atom (← magic number)

what determines the limit of existence of elements? (i) electron orbits

hydrogen-like atom

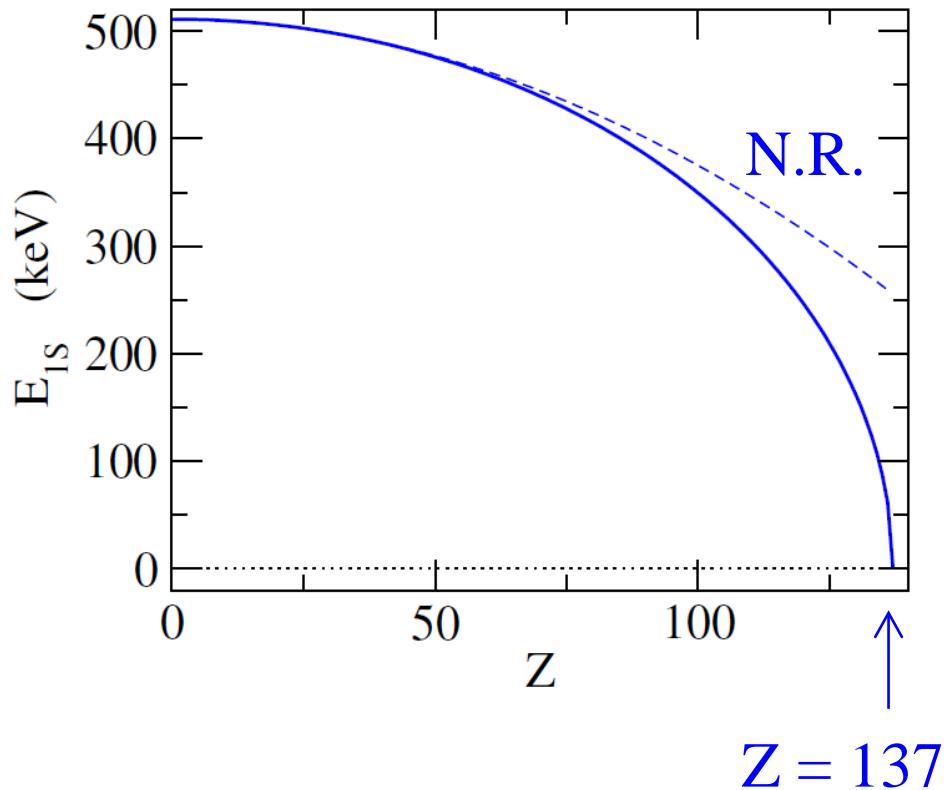
$$V(r) = -\frac{Ze^2}{r}$$

1S state (Dirac equation)

$$E_{1S} = mc^2 \sqrt{1 - (Z\alpha)^2}$$

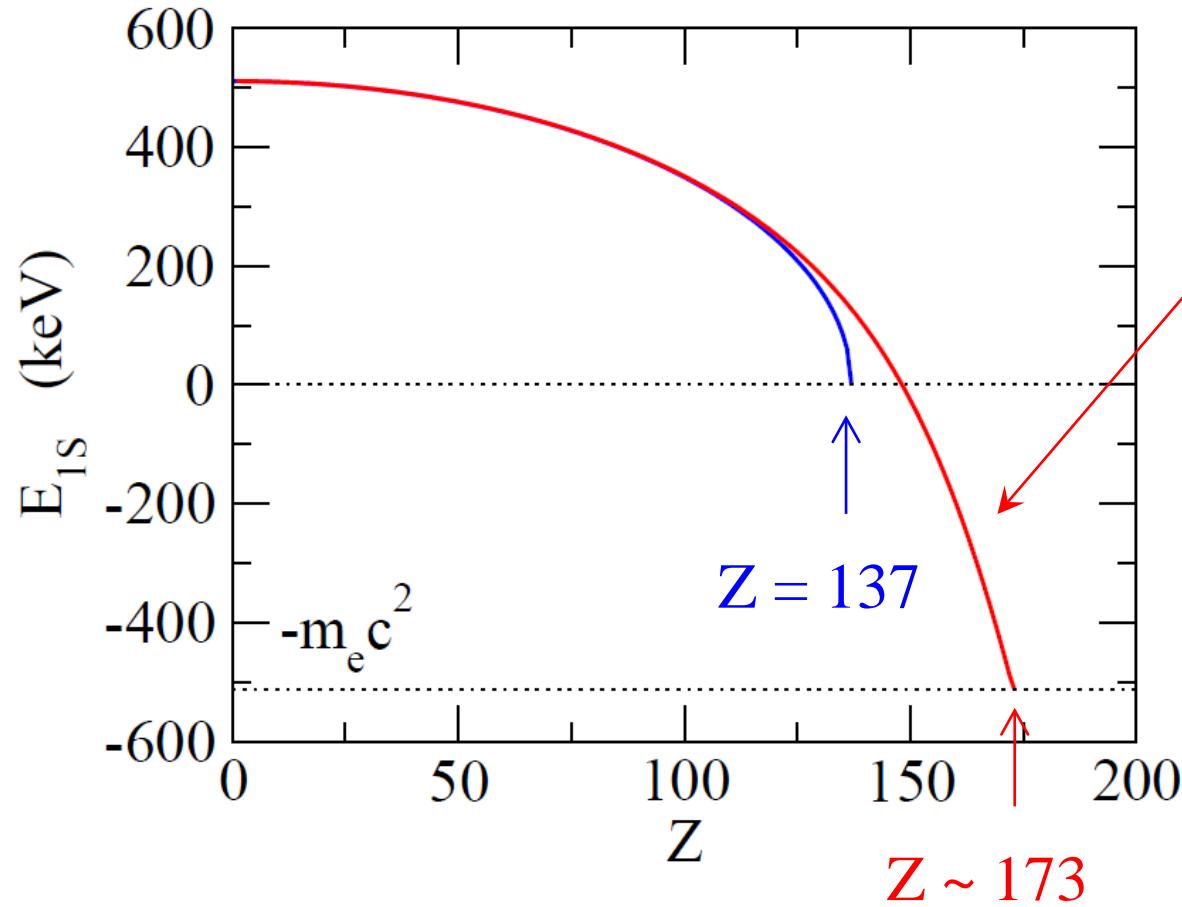
$$\alpha = \frac{e^2}{\hbar c} \sim \frac{1}{137}$$

$Z > 137 \rightarrow$ no solution



what determines the limit of existence of elements? (i) electron orbits

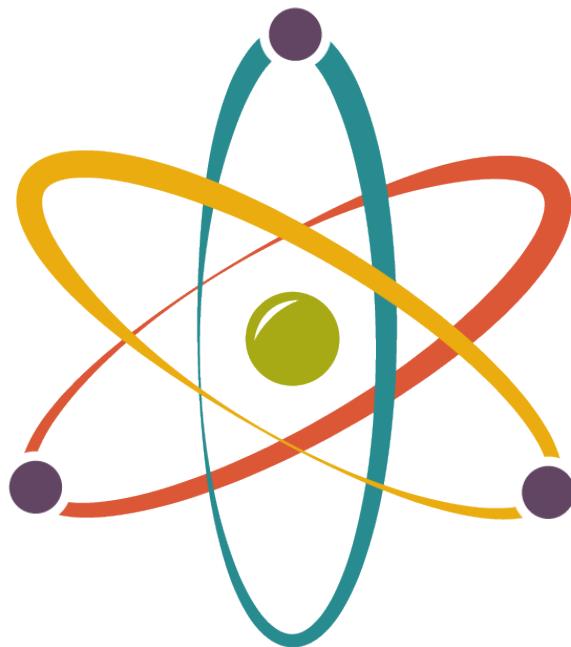
hydrogen-like atom



finite size of a
nucleus
(uniform charge)

$$R_N \sim 10^{-4} \text{ \AA}$$
$$\text{rms} \sim 10^{-3} \text{ \AA}$$
$$(A=173)$$

what determines the limit of existence of elements?



INTERNATIONAL YEAR OF THE PERIODIC TABLE **2019**

I 53 Iodine	Y 39 Yttrium	Pt 78 Platinum	Ca 20 Calcium	K 19 Potassium
-------------------	--------------------	----------------------	---------------------	----------------------

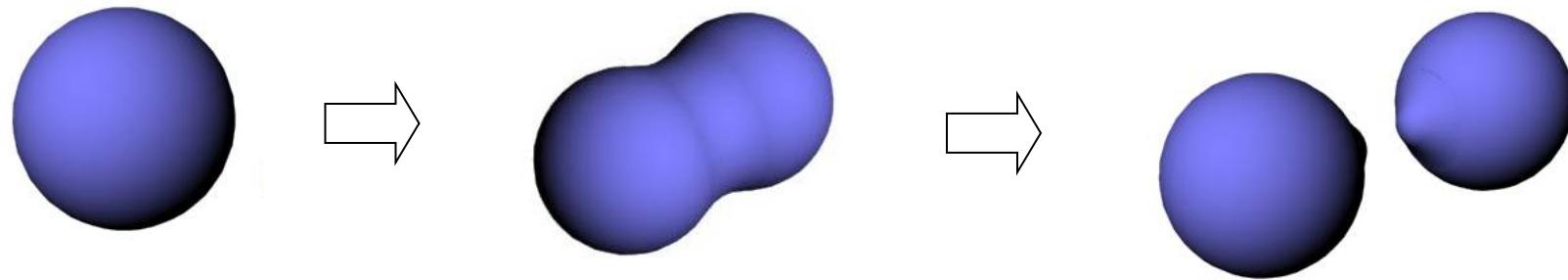
possibilities to be considered:

let's next discuss

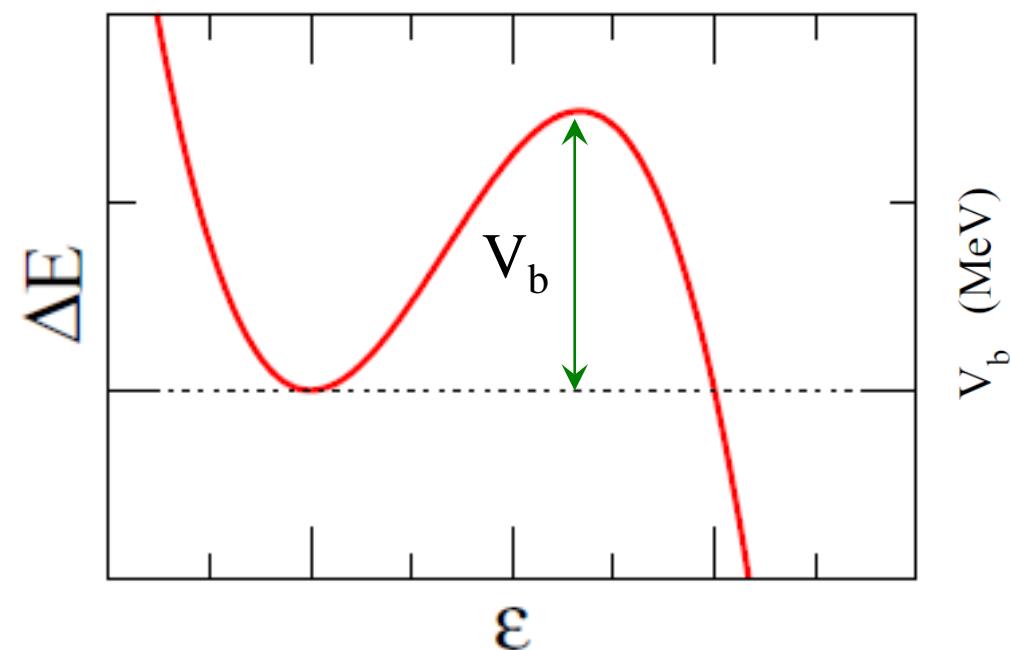
- ✓ electronic orbitals in atom
- ✓ stability of nucleus in atom

what determines the limit of existence of elements? (ii) atomic nucleus

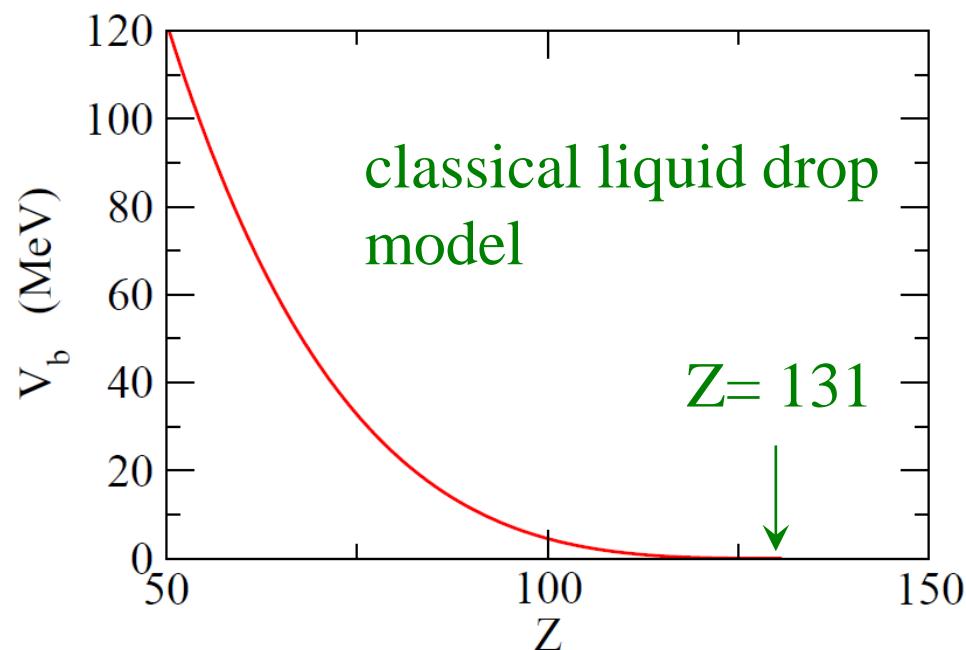
heavy nuclei → unstable against fission



fission barrier

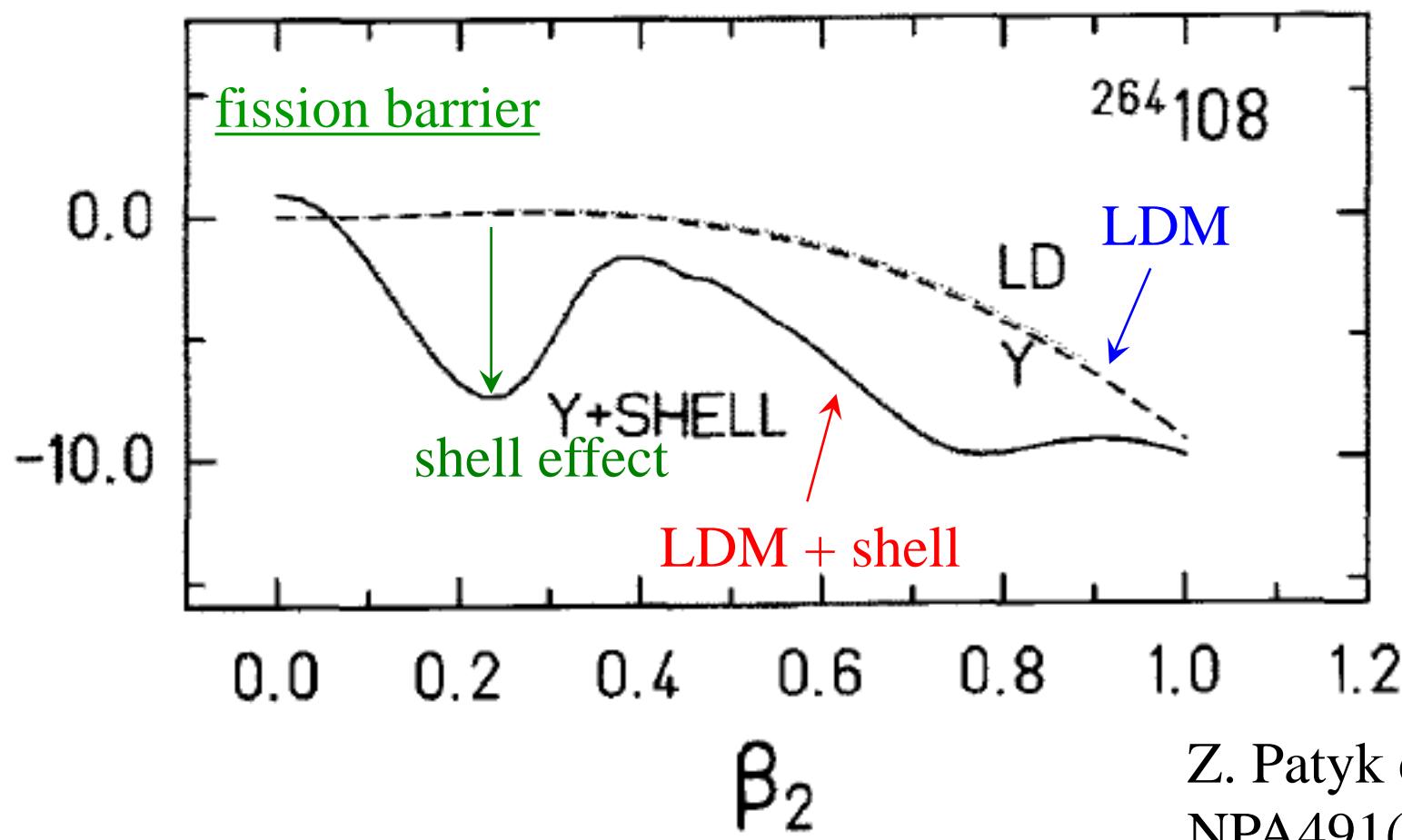


barrier height



classical liquid drop model

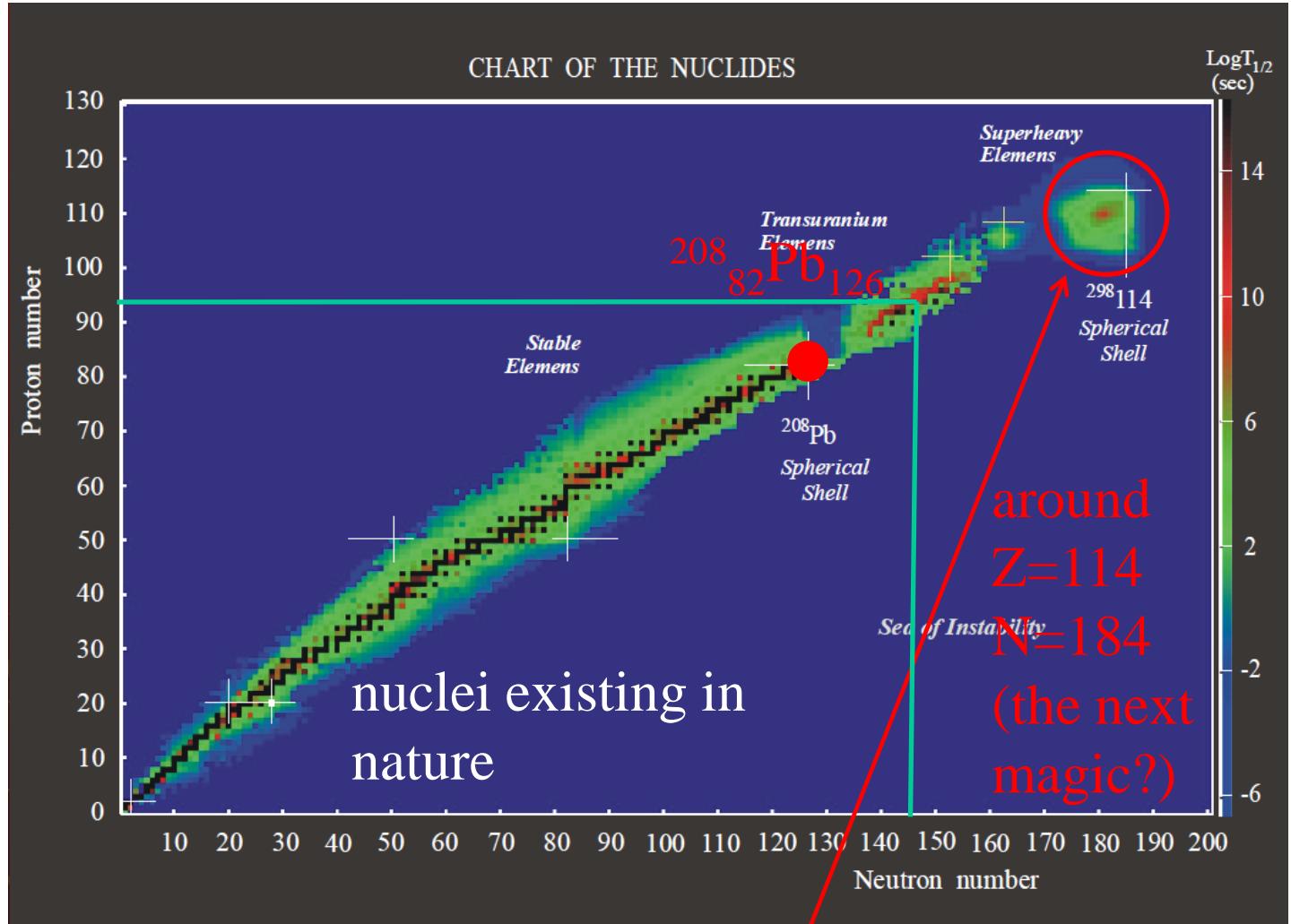
what determines the limit of existence of elements? (ii) atomic nucleus



Z. Patyk et al.,
NPA491('89) 267

QM shell effect (magic numbers) raises B_{fiss} and stabilizes a nucleus

Superheavy elements (the island of stability)

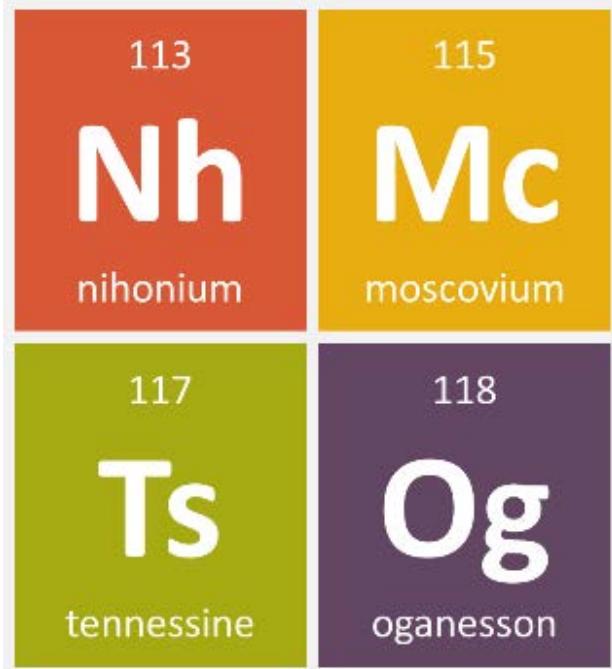


Yuri Oganessian

long-lived with 10^{3-5} years

Fusion reactions for SHE

the element 113: Nh

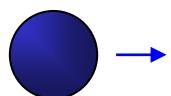
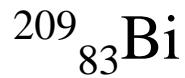


November, 2016



A detailed periodic table showing elements from hydrogen (H) to oganesson (Og). The element Nh (Nihonium) is highlighted with a red box in its group 13 position. Other elements are color-coded by group: groups 14-17 are green, group 18 is light blue, groups 1-3 are pink, and transition metals are grey. An asterisk (*) is placed below several elements, likely indicating they are synthetic or radioactive.

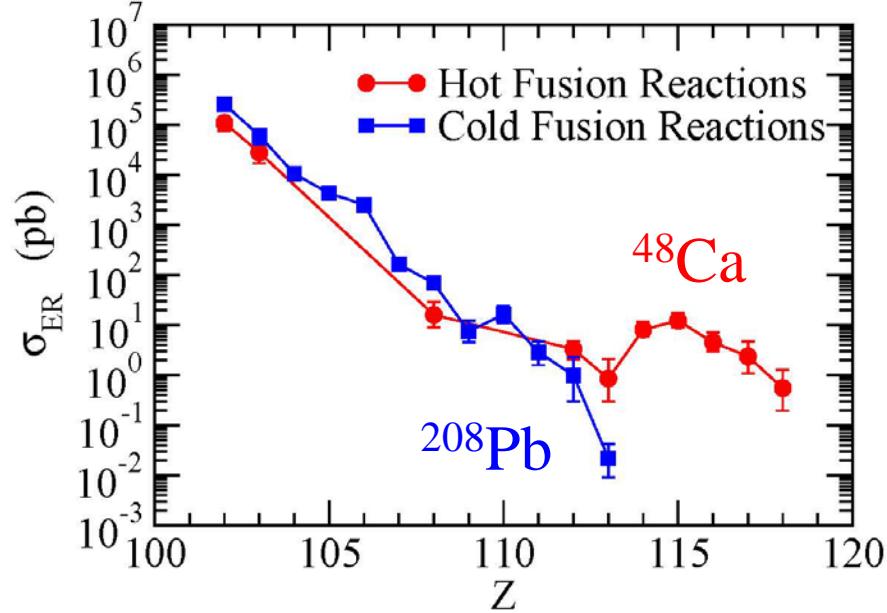
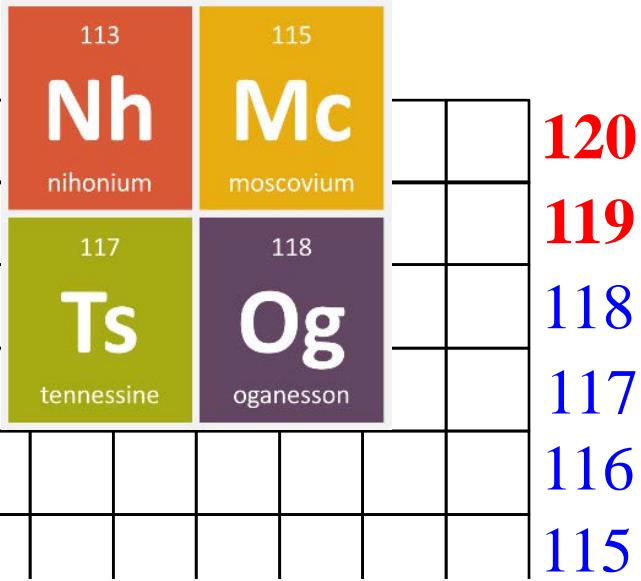
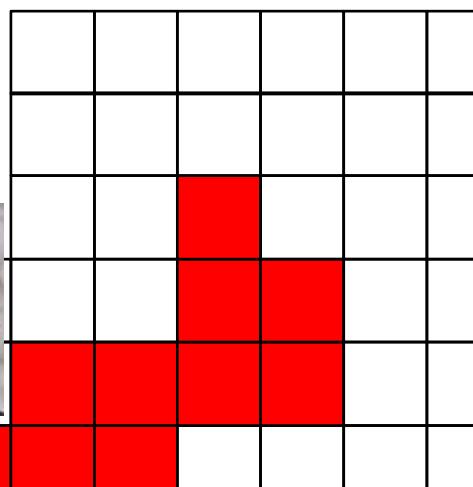
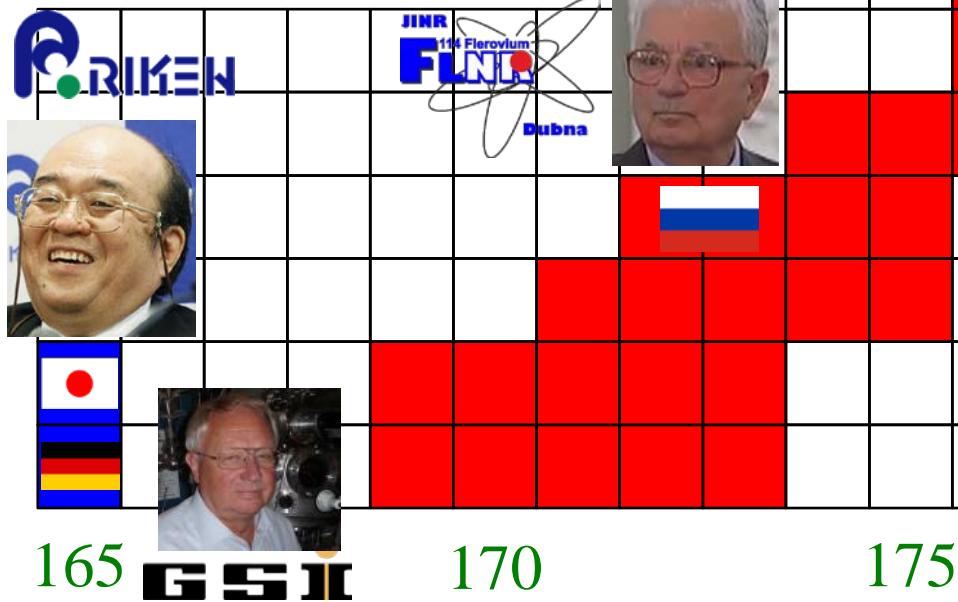
Wikipedia



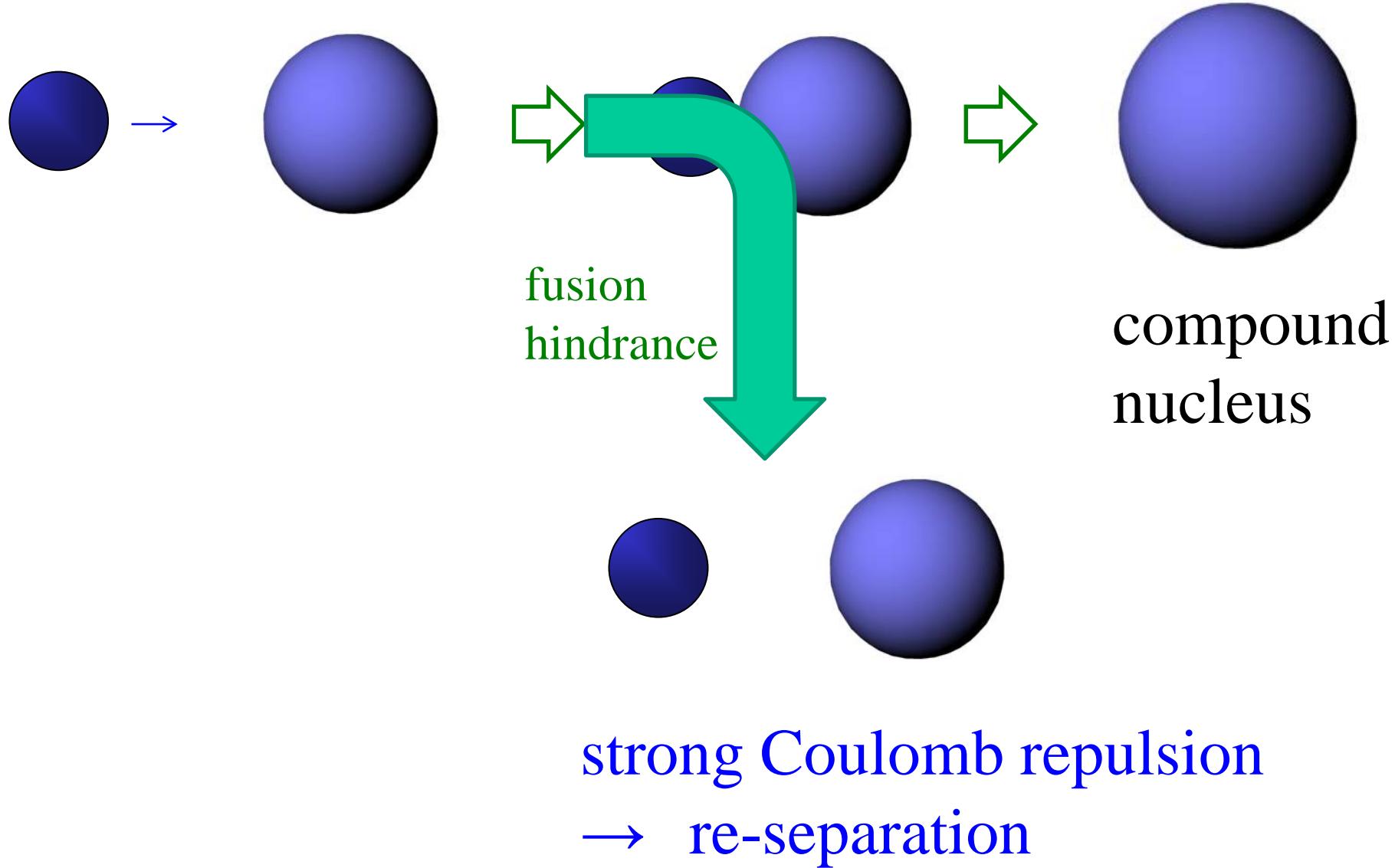
Heavy-ion fusion reaction

Fusion for superheavy elements

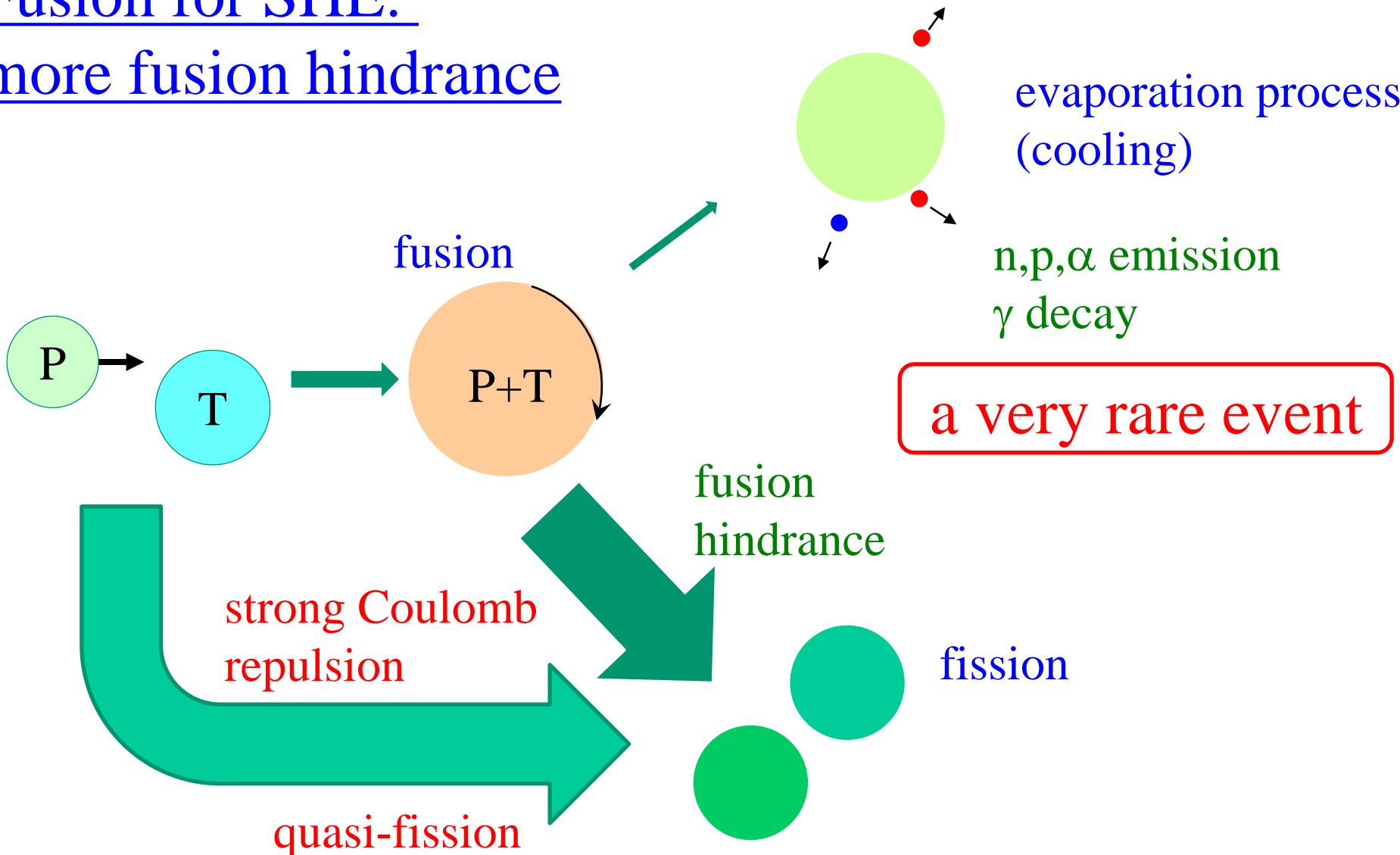
Superheavy elements
synthesized so far



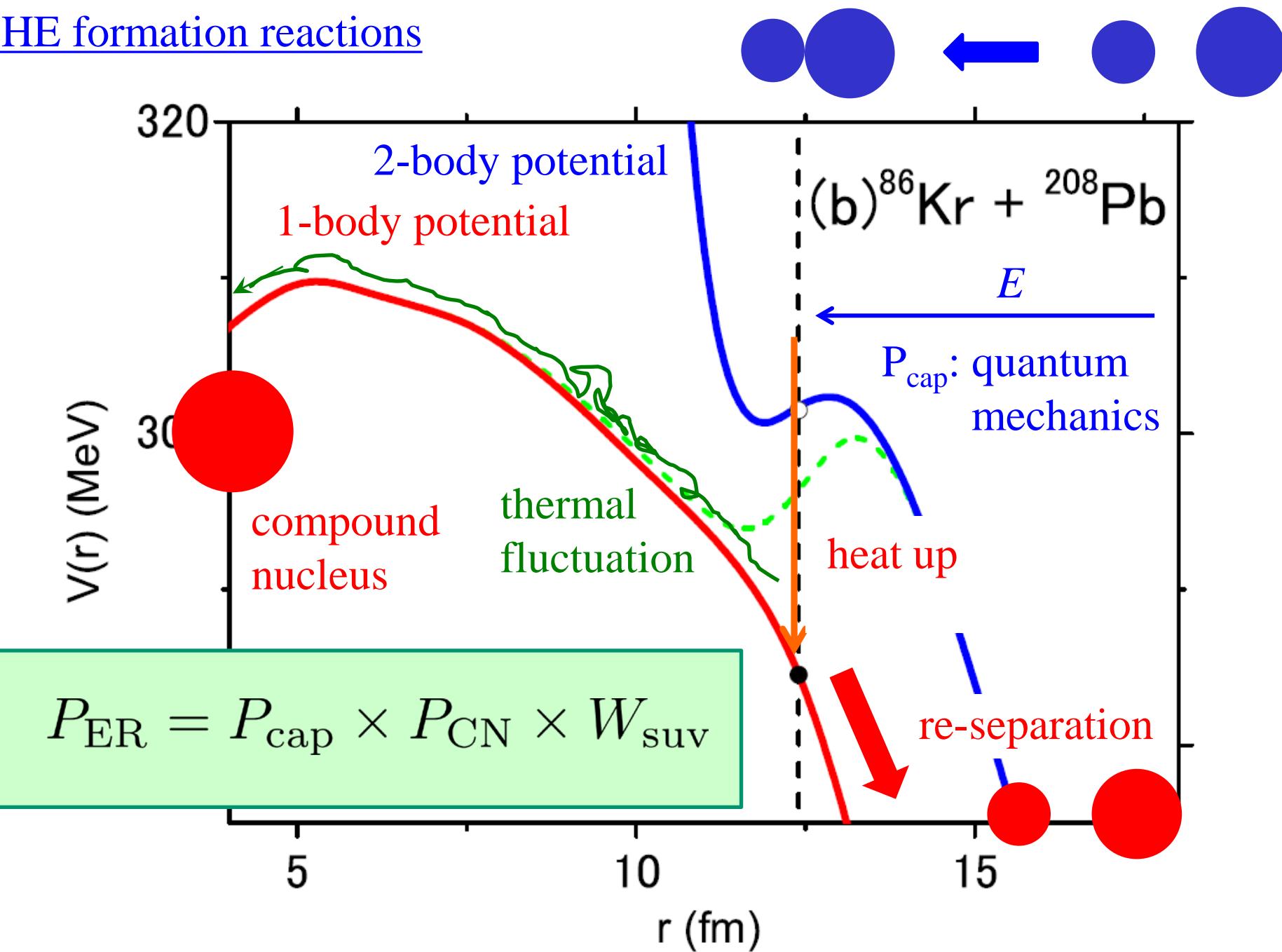
Fusion for SHE: fusion hindrance



Fusion for SHE: more fusion hindrance



SHE formation reactions



Theoretical challenges

formation of SHE: very rare
→ a large theoretical uncertainty

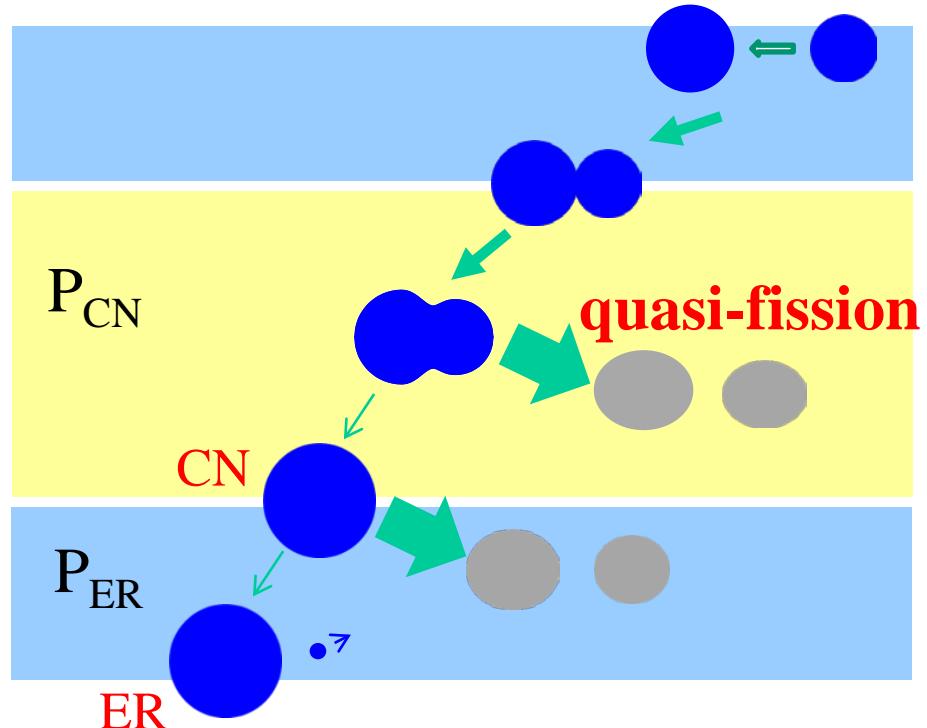
$$P_{\text{ER}} = P_{\text{cap}} \cdot P_{\text{CN}} \cdot W_{\text{suv}}$$

- ✓ no exp. data for P_{CN}
- ✓ exp. data: P_{ER} only

CN=複合核、ER=蒸発残留核

theoretical challenges:
to reduce the uncertainties and
make reliable predictions

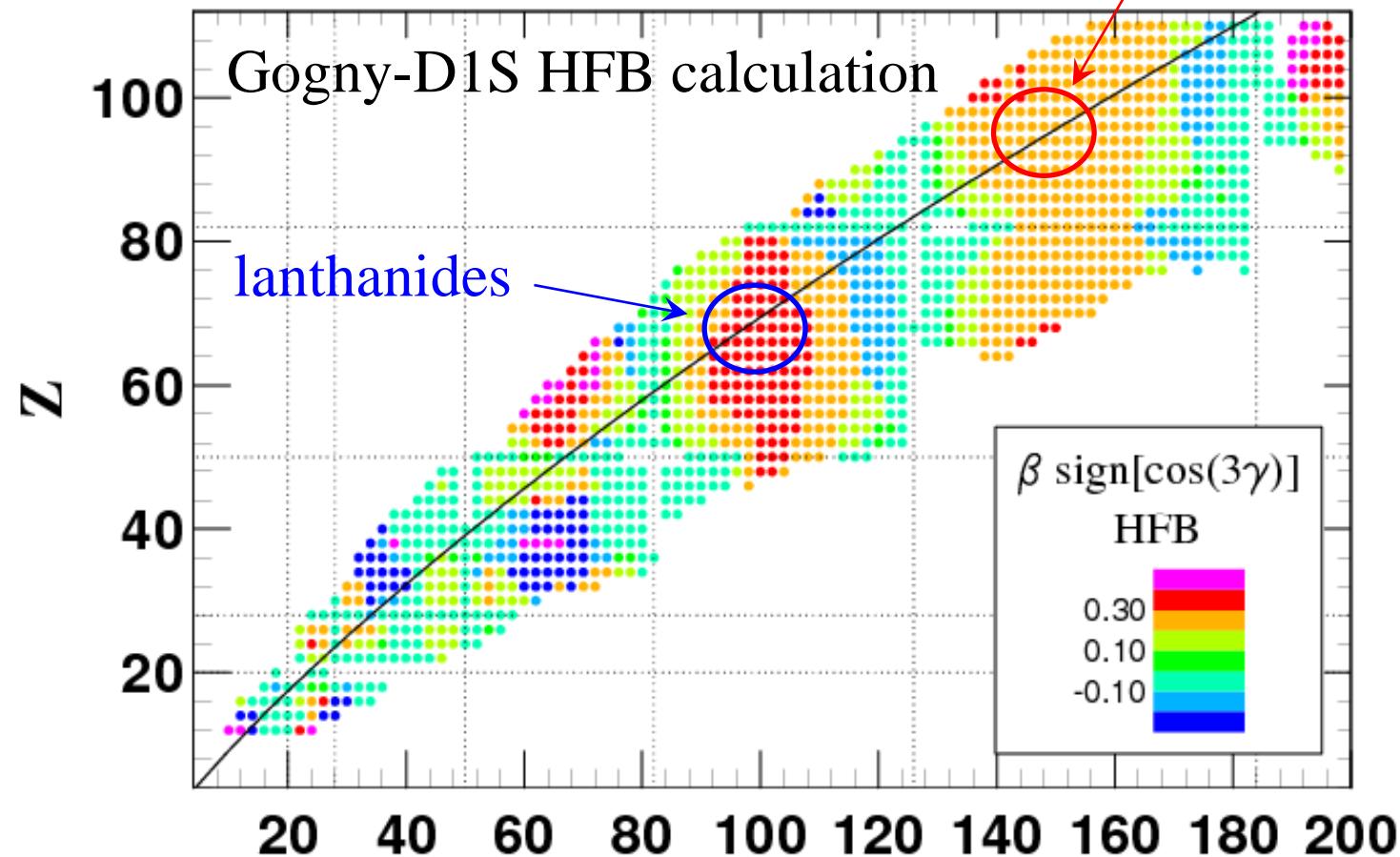
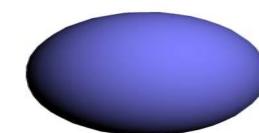
fusion hindrance



Physics of open quantum systems
量子開放系の物理

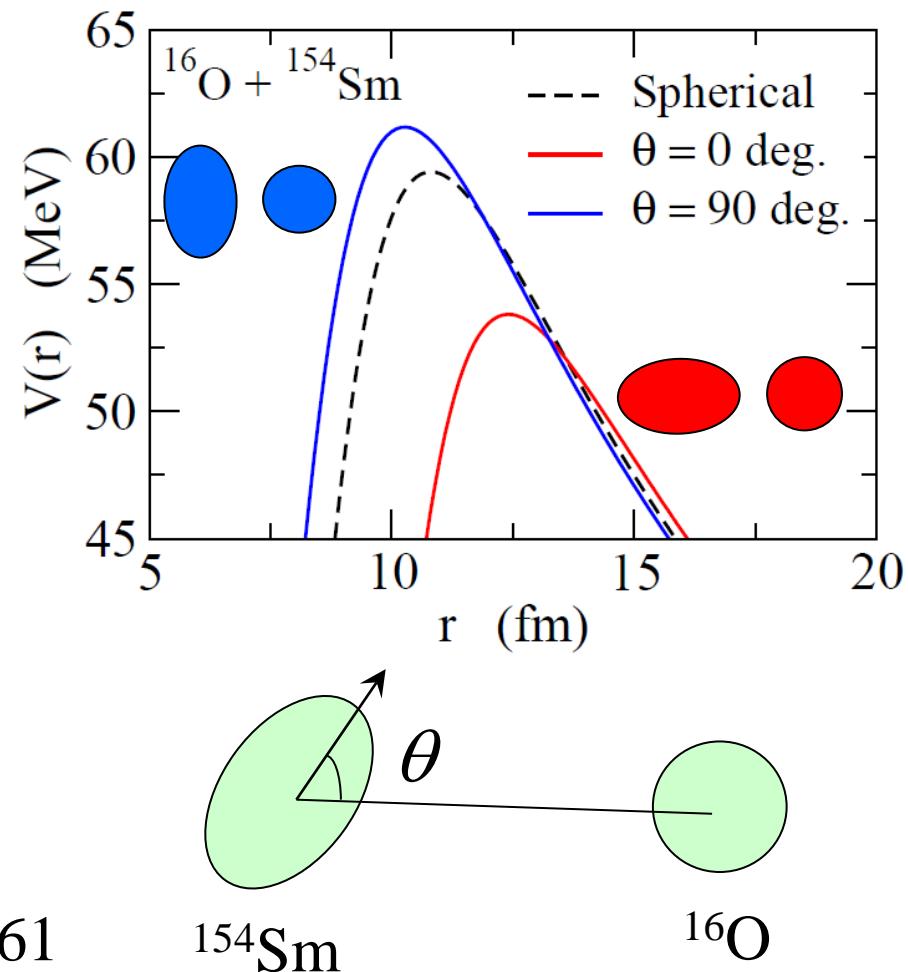
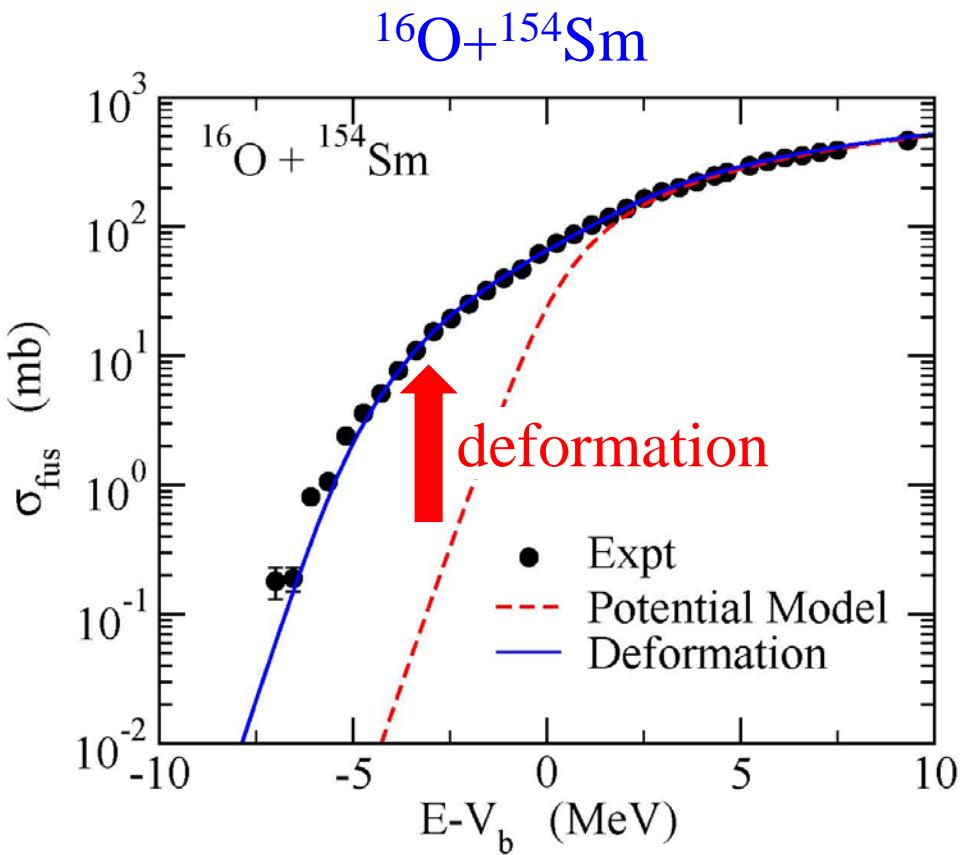
hot fusion: Nuclear Deformation

hot fusion: ^{48}Ca + deformed target



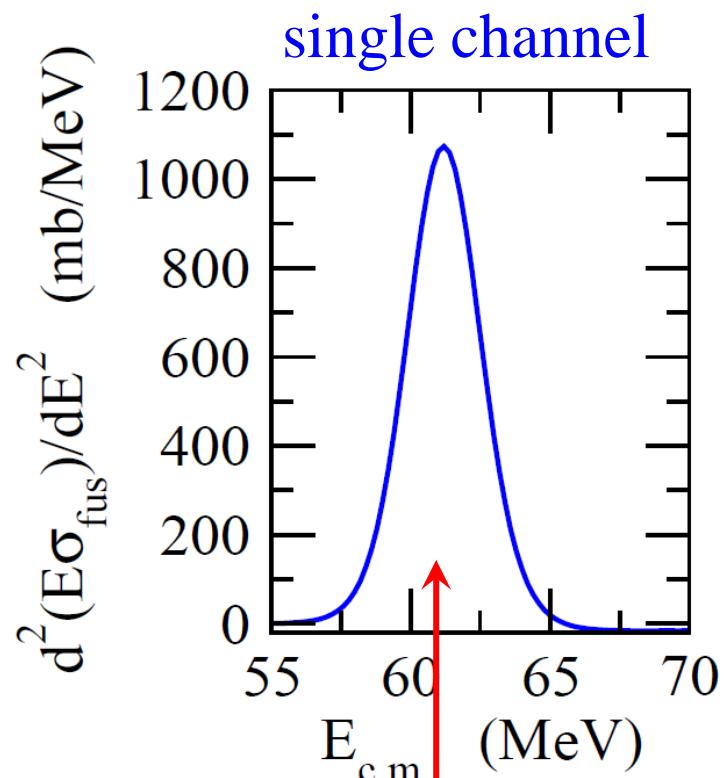
Nuclear deformation and barrier distribution

Nuclear deformation → a large sub-barrier enhancement of fusion cross sections



Fusion barrier distribution [Rowley, Satchler, Stelson, PLB254('91)]

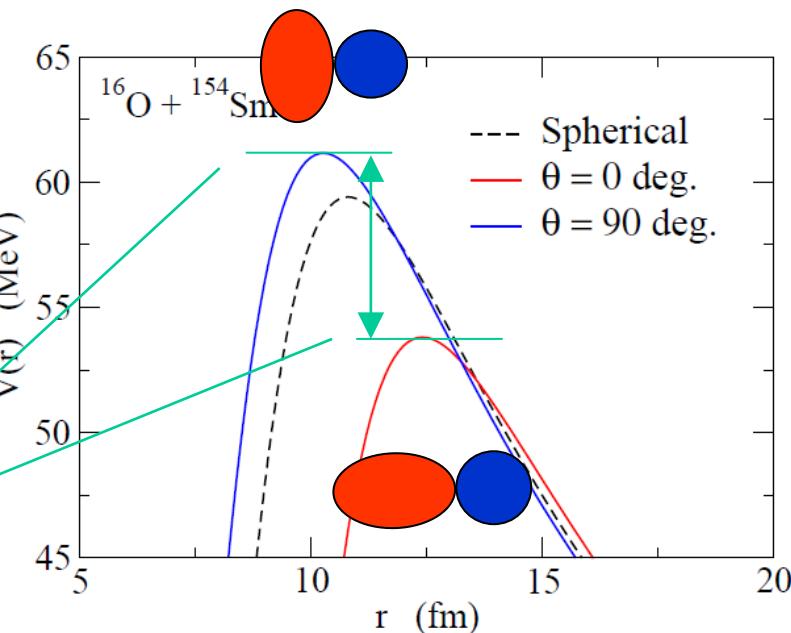
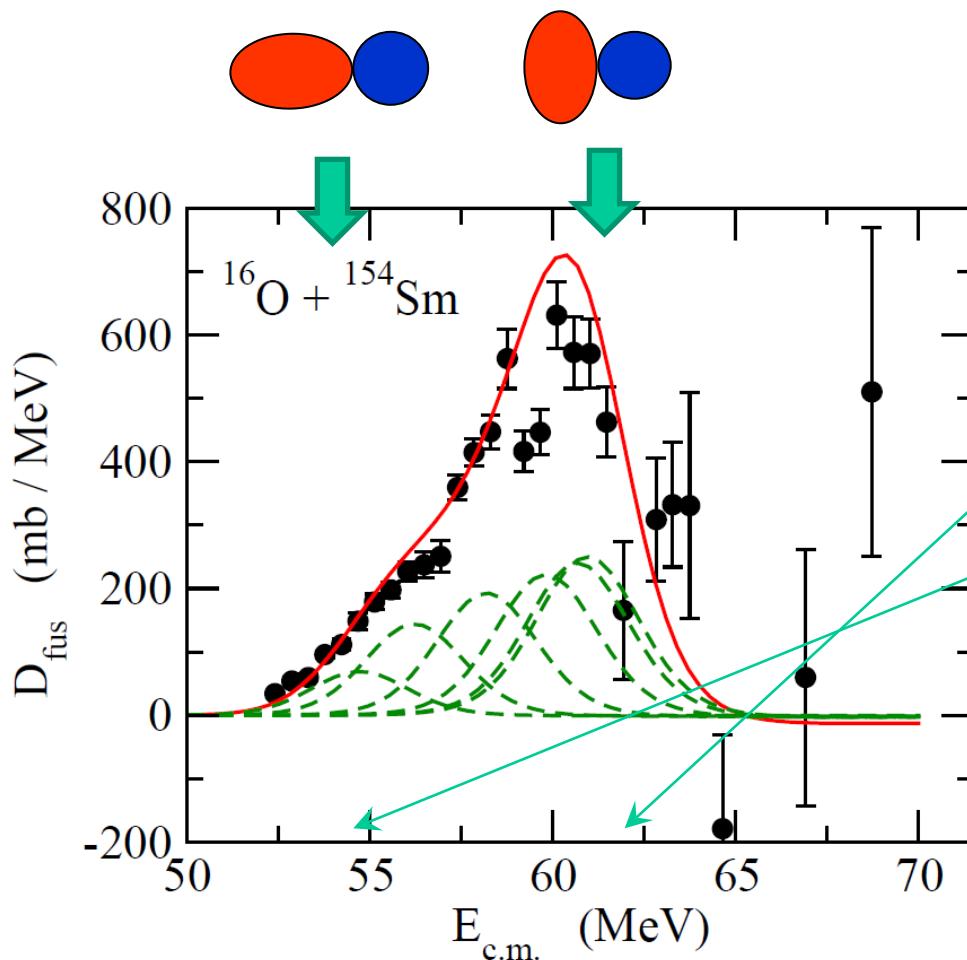
$$D_{\text{fus}}(E) = \frac{d^2(E\sigma_{\text{fus}})}{dE^2} \propto \frac{dP_{l=0}}{dE}$$



V_b

✓ Fusion barrier distribution (Rowley, Satchler, Stelson, PLB254('91))

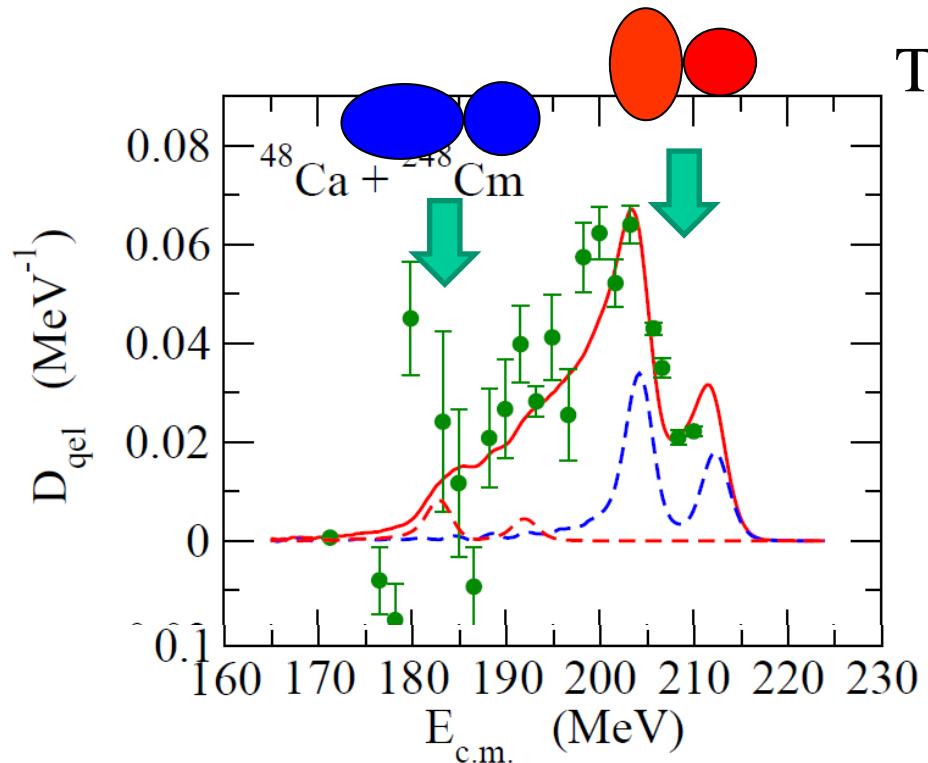
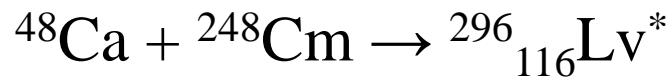
$$D_{\text{fus}}(E) = \frac{d^2(E\sigma_{\text{fus}})}{dE^2}$$



Data: J.R. Leigh et al.,
PRC52 ('95) 3151

can be used to identify
the side/tip collisions

Application to hot fusion reactions

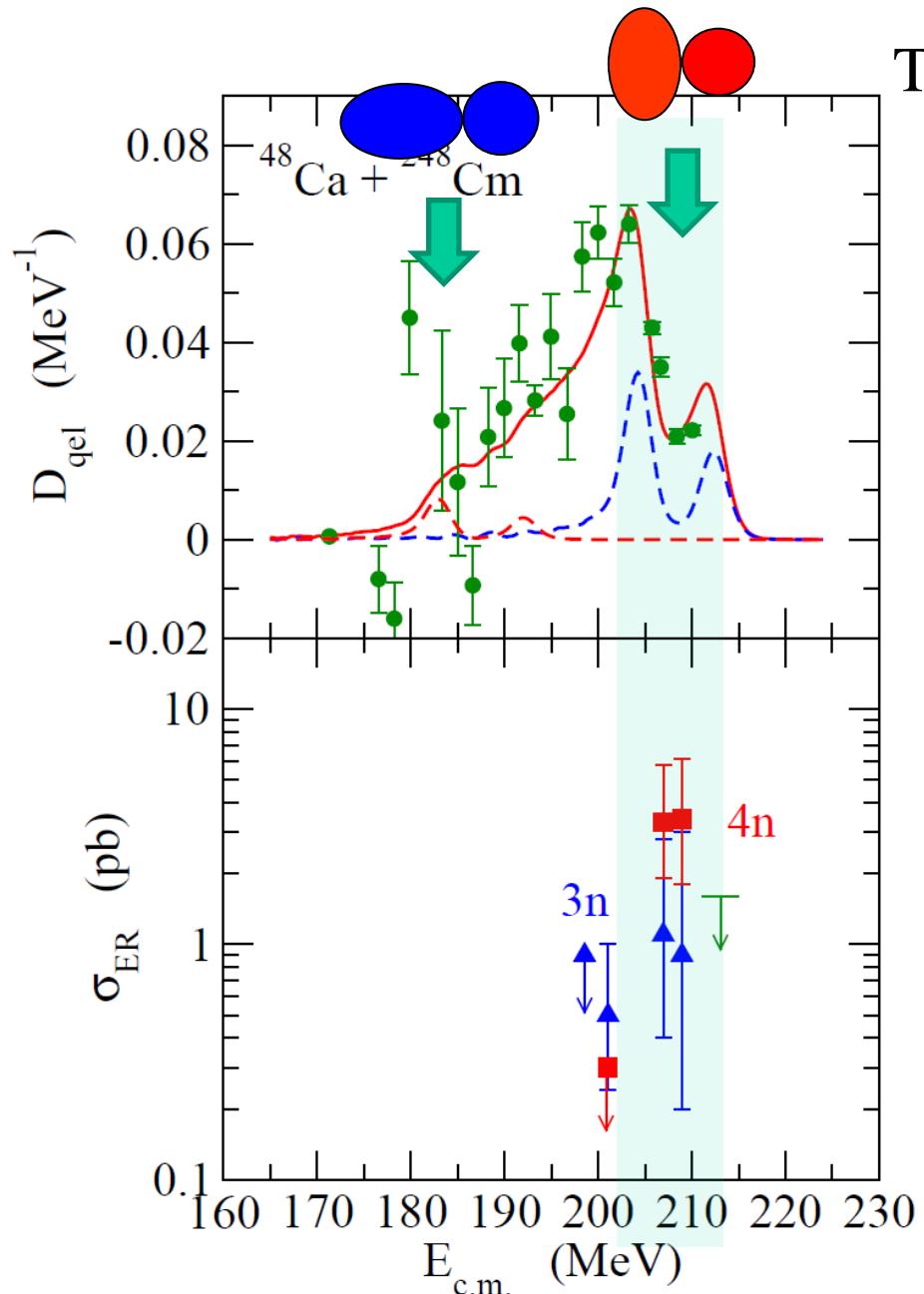
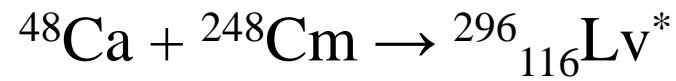


T. Tanaka,..., K.H., et al.,
JPSJ 87 ('18) 014201
PRL124 ('20) 052502



capture barrier distribution

Application to hot fusion reactions



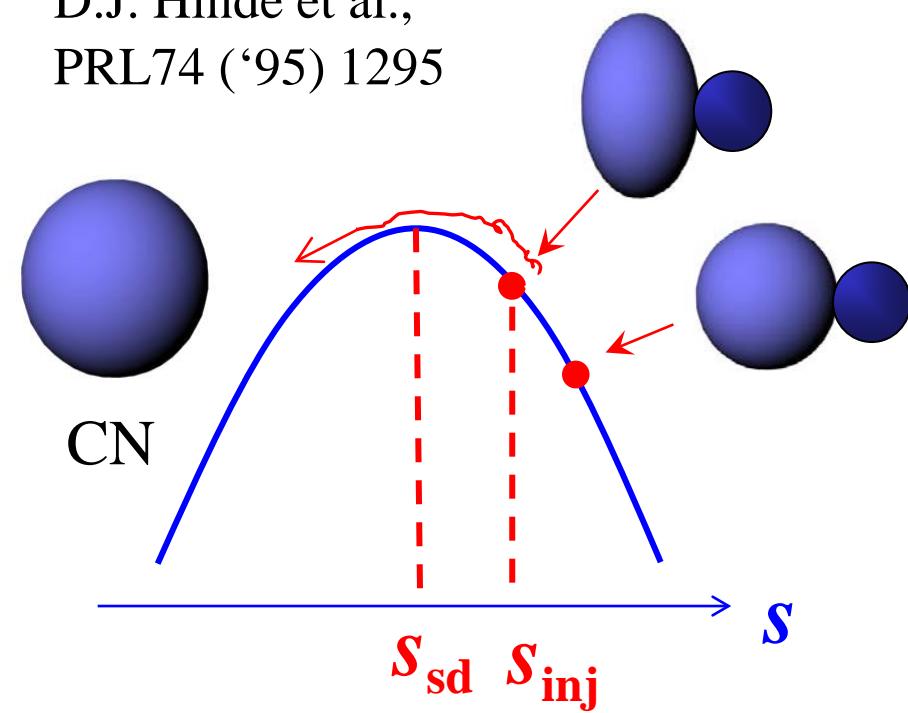
T. Tanaka,..., K.H., et al.,
JPSJ 87 ('18) 014201
PRL124 ('20) 052502



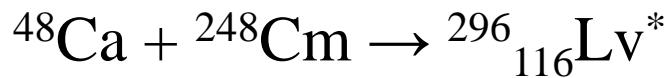
capture barrier distribution

cf. notion of compactness:

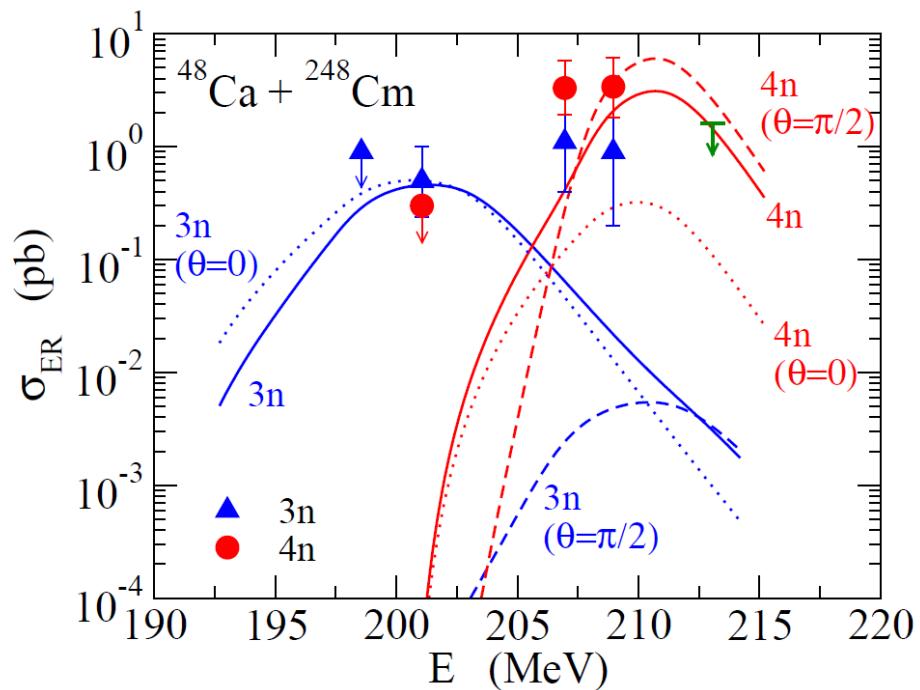
D.J. Hinde et al.,
PRL74 ('95) 1295



Application to hot fusion reactions



T. Tanaka,..., K.H., et al.,
JPSJ 87 ('18) 014201
PRL124 ('20) 052502

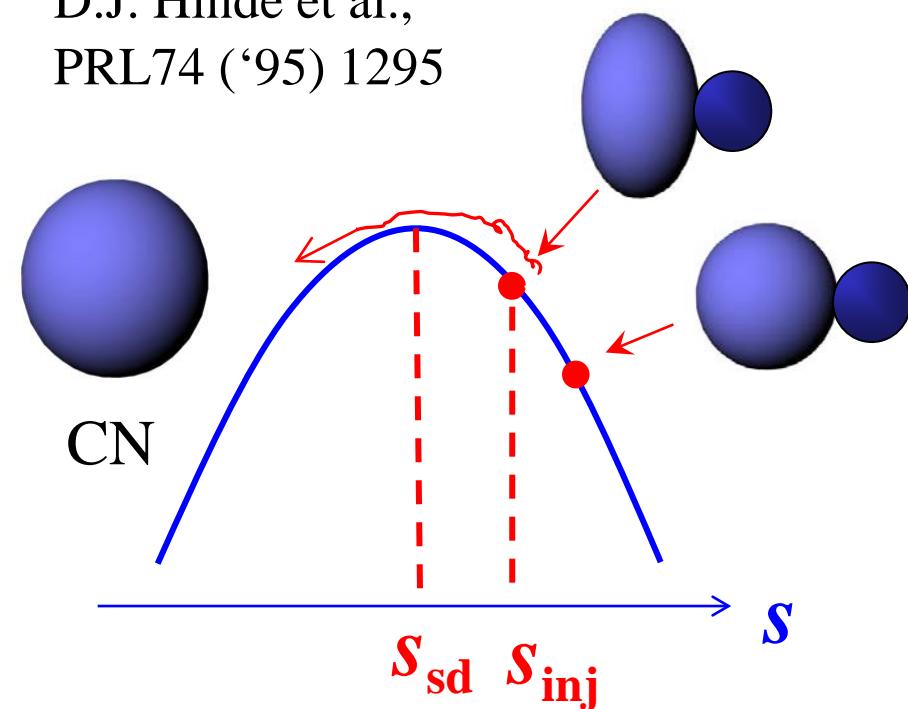


K. Hagino, PRC98 ('18) 014607

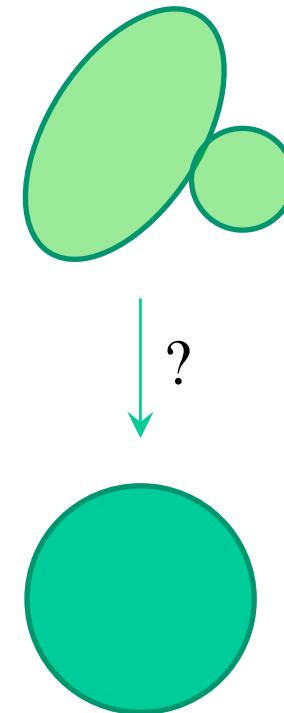
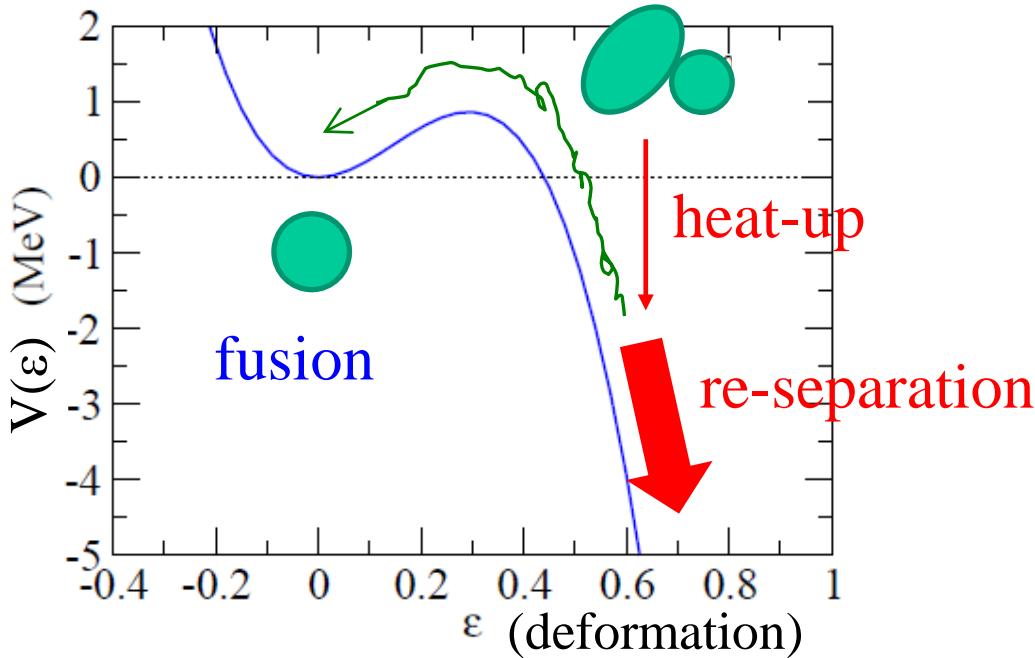
capture barrier distribution

cf. notion of compactness:

D.J. Hinde et al.,
PRL74 ('95) 1295



open problems



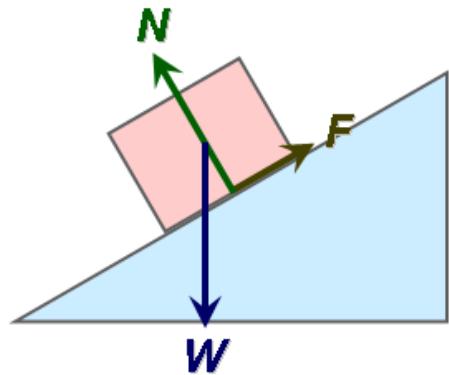
- how is the shape evolved to a compound nucleus?
- Deformation: a quantum effect
how does the deformation disappear during heat-up?

quantum friction/open quantum systems

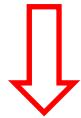
M. Tokieda and K.H., Ann. of Phys. 412 ('20) 168005.
Front. in Phys. 8 ('20) 8.

quantum friction

in quantum mechanics:

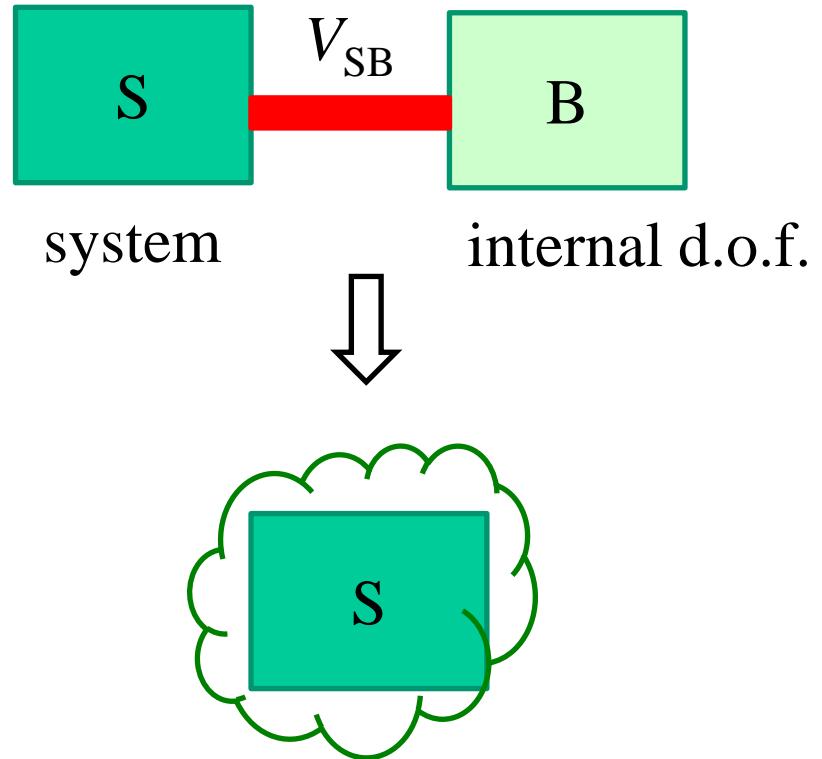


heat generation when a rigid body stops



the energy conversion from the rigid body to intrinsic d.o.f. (atoms)

quantum Langevin?



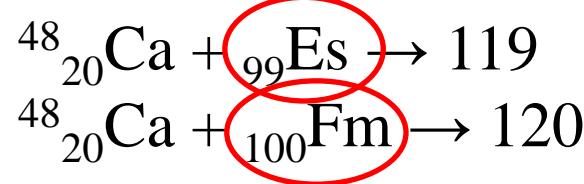
“quasi” particle

solve the whole H without introducing the quasi-particle

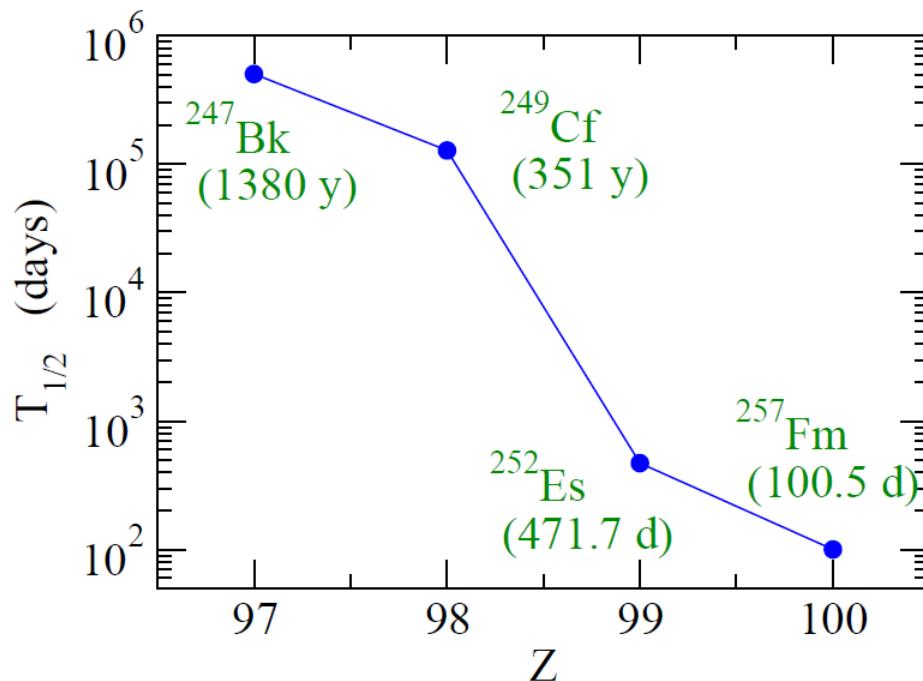
M. Tokieda and K.H. (2020)

Hot fusion towards Z=119 and 120 nuclei

hot fusion reactions with ^{48}Ca :



short lived → not available with sufficient amounts



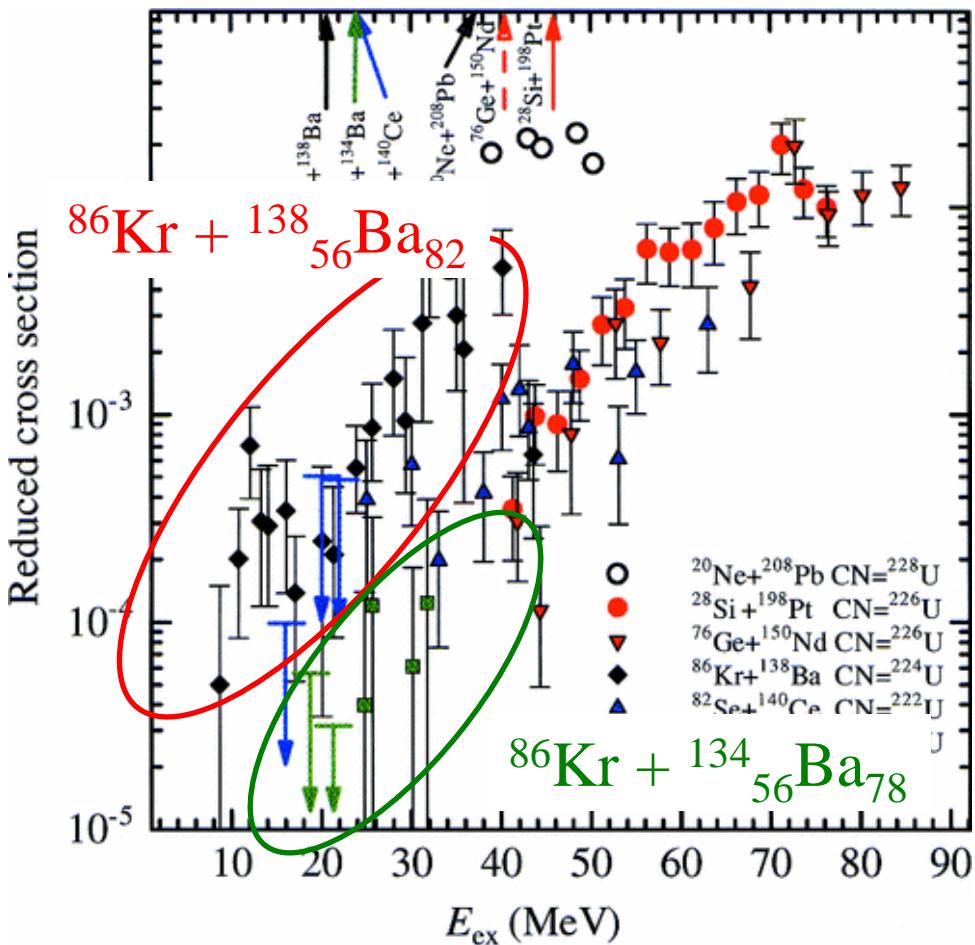
$^{48}\text{Ca} \rightarrow {}_{22}^{50}\text{Ti}, {}_{23}^{51}\text{V}, {}_{24}^{54}\text{Cr}$ projectiles

closed shell → open shells

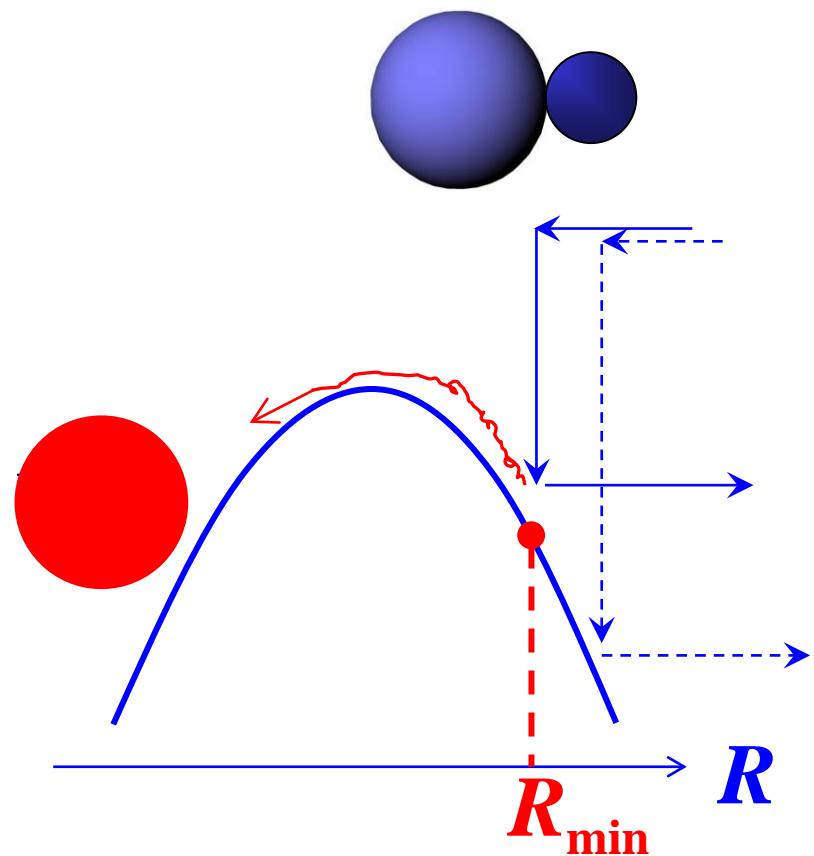
how much will cross sections be affected?

Role of magicity

can proceed deeper
with less friction



K. Satou, H. Ikezoe et al.,
PRC73 ('06) 034609

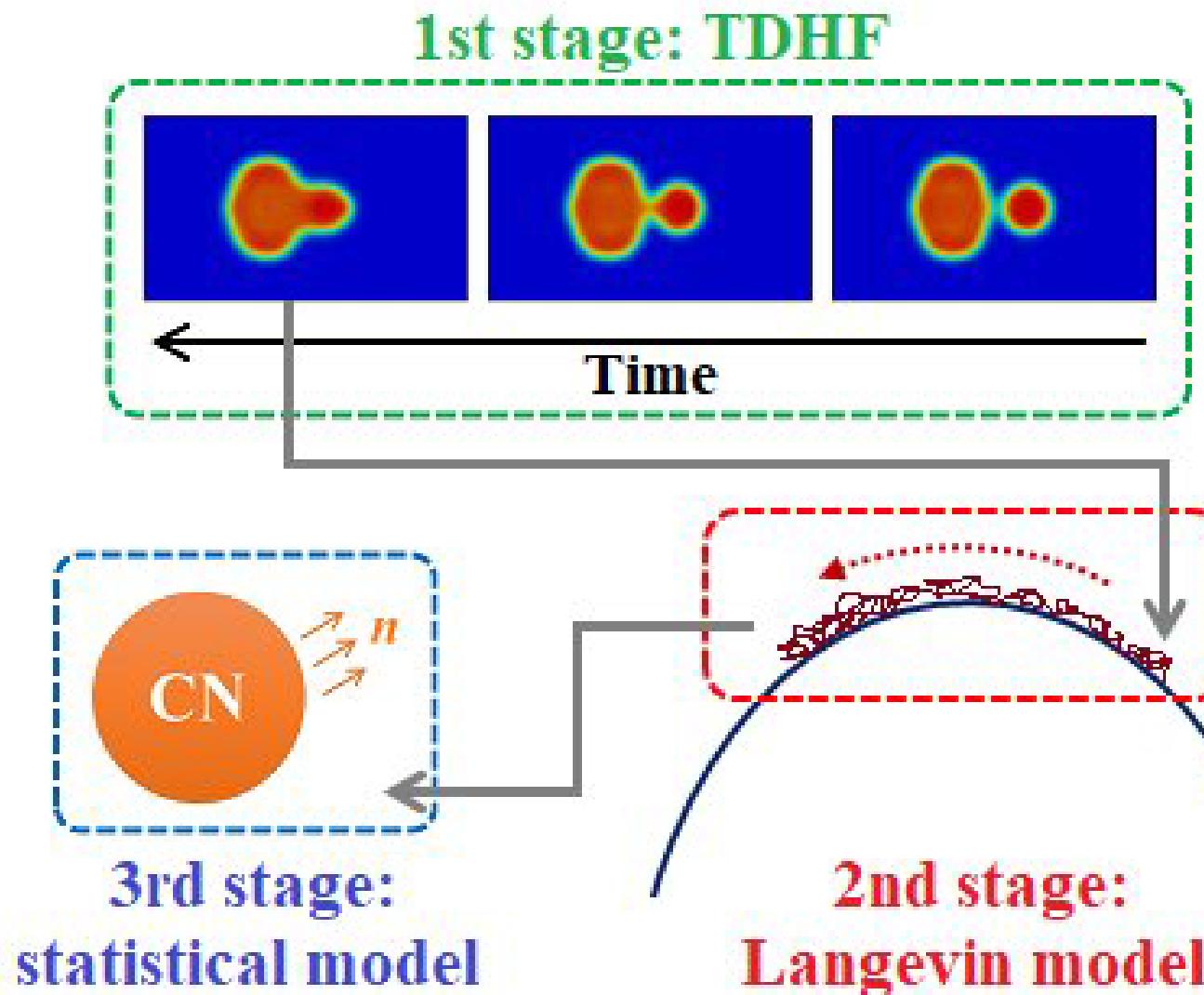


cf. P. Moller et al.,
Z. Phys. A359 ('97) 251.

similar effect for ^{48}Ca ?

New hybrid model: TDHF + Langevin approach

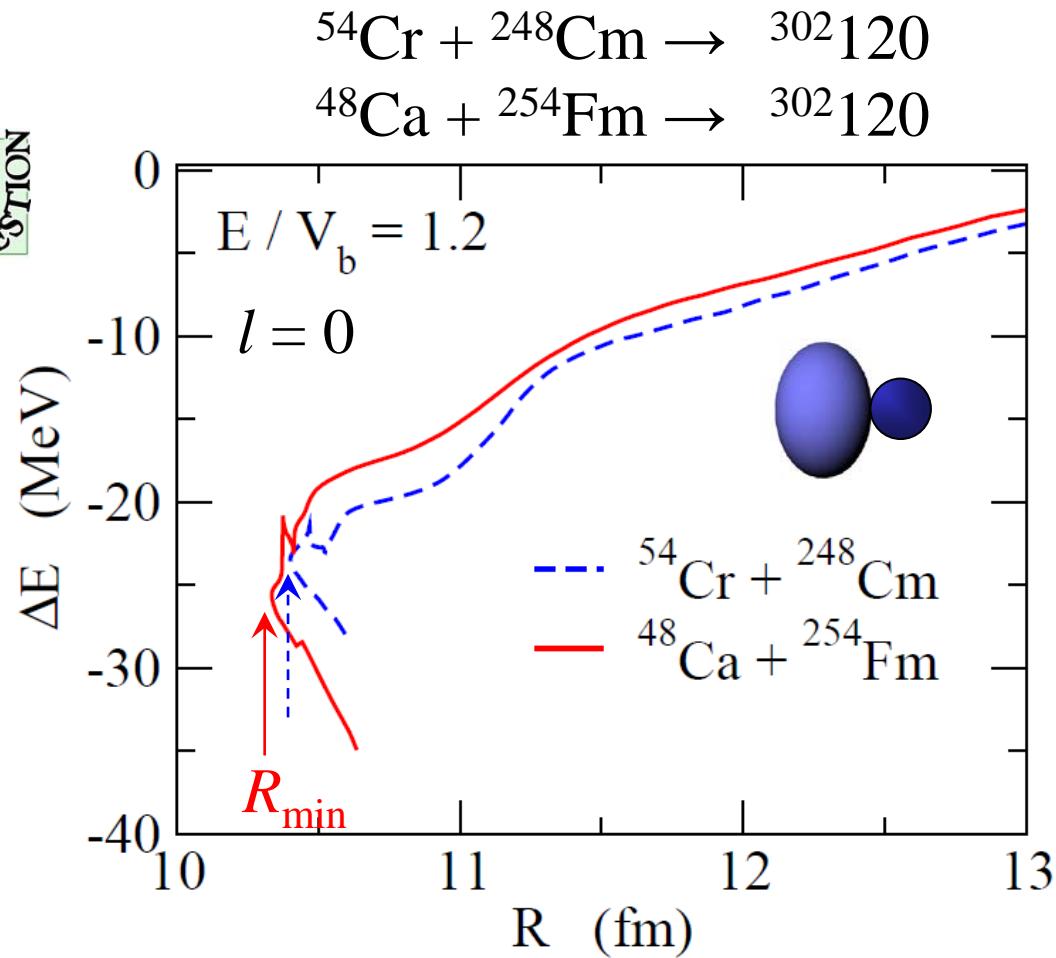
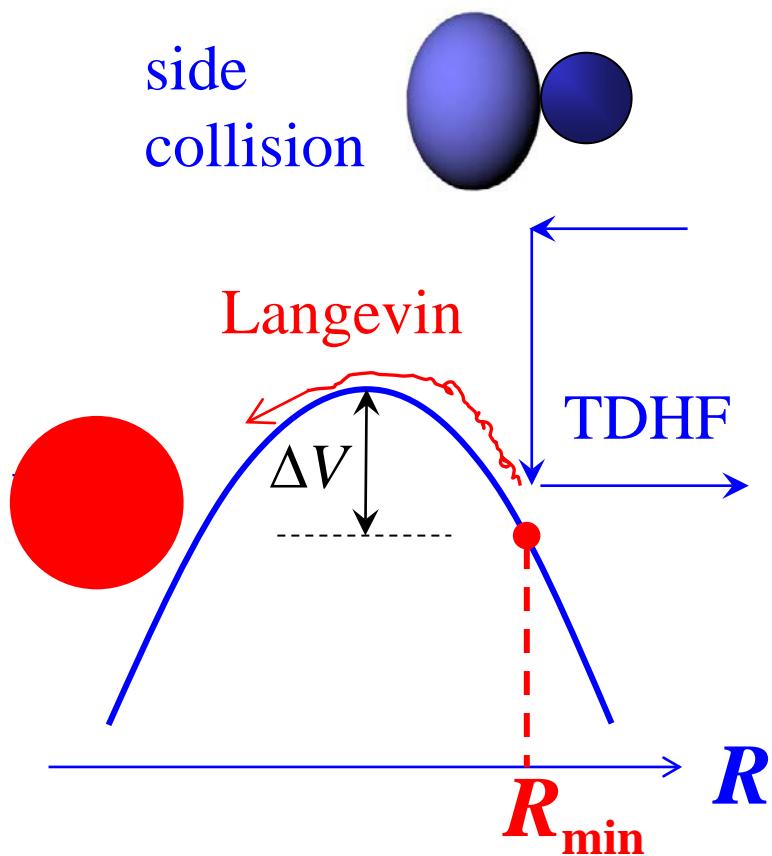
K. Sekizawa and K.H., PRC99 (2019) 051602(R)



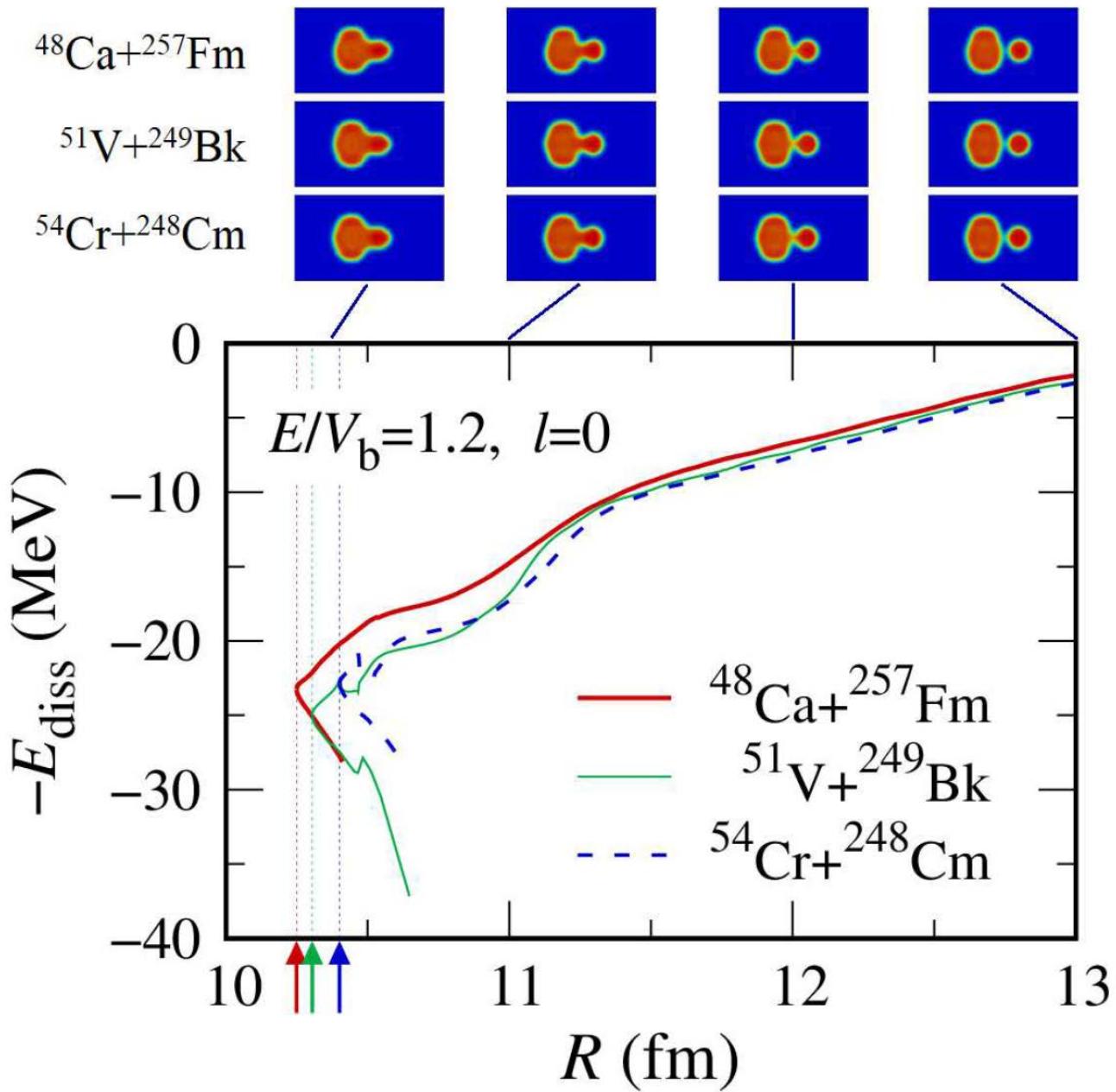
Furthermore : TDHF + Langevin approach and quantum friction

TDHF + Langevin approach :

K. Sekizawa and K. H.,
PRC99 (2019) 051602(R)



→ Langevin calculation



New model for fusion for SHE: TDHF + Langevin approach

K. Sekizawa and K.H., PRC99 (2019) 051602(R)



how special is ^{48}Ca ?

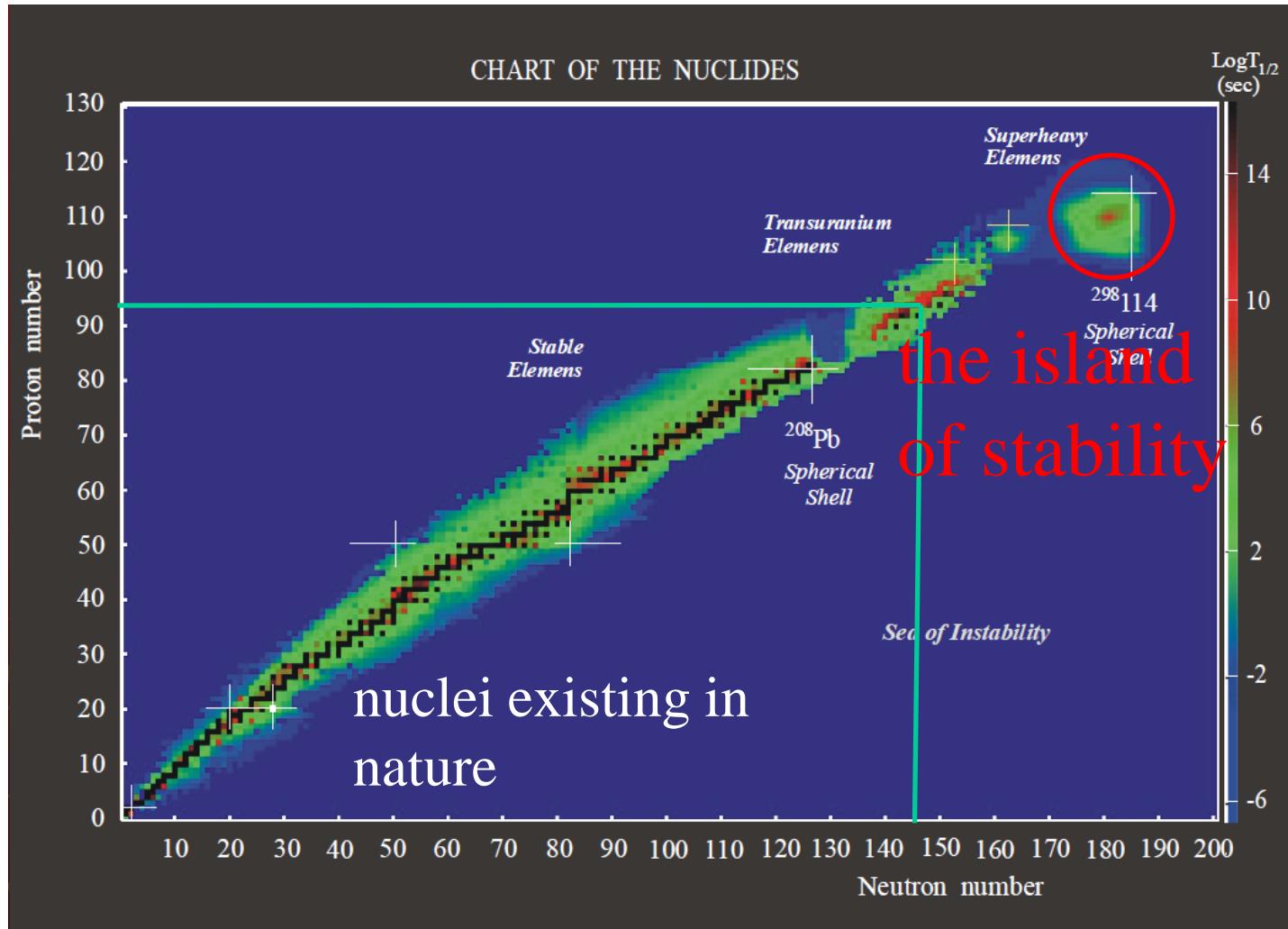
System	CN	E^* (MeV)	R_{\min} (fm)	P_{CN} ($\times 10^4$)	W_{sur} ($\times 10^9$)	$P_{\text{CN}} W_{\text{sur}}$ ($\times 10^{13}$)
$^{48}\text{Ca} + ^{254}\text{Fm}$	$^{302}120$	29.0	12.93	1.72	176	302
$^{54}\text{Cr} + ^{248}\text{Cm}$	$^{302}120$	33.2	13.09	1.89	1.31	2.47
$^{51}\text{V} + ^{249}\text{Bk}$	$^{300}120$	37.0	12.94	3.95	0.117	0.461
$^{48}\text{Ca} + ^{257}\text{Fm}$	$^{305}120$	30.5	12.94	2.49	0.729	1.82

$$P_{\text{ER}} = P_{\text{cap}} \cdot P_{\text{CN}} \cdot W_{\text{suv}}$$

similar P_{CN}

- ✓ no special role of ^{48}Ca in the entrance channel
- ✓ non- ^{48}Ca proj.: about 2 order of magnitude smaller
due mainly to W_{sur}

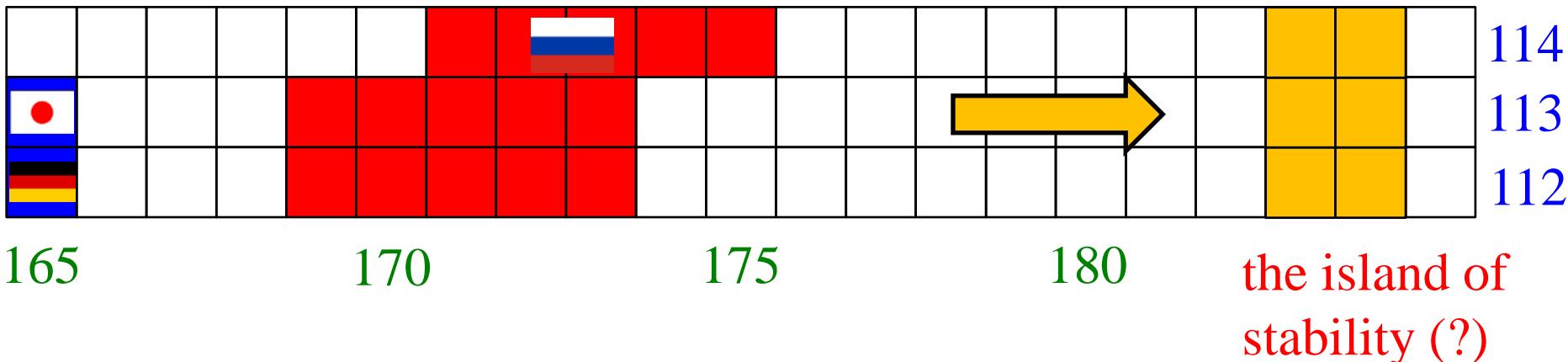
Another important issue: physics of neutron-rich nuclei



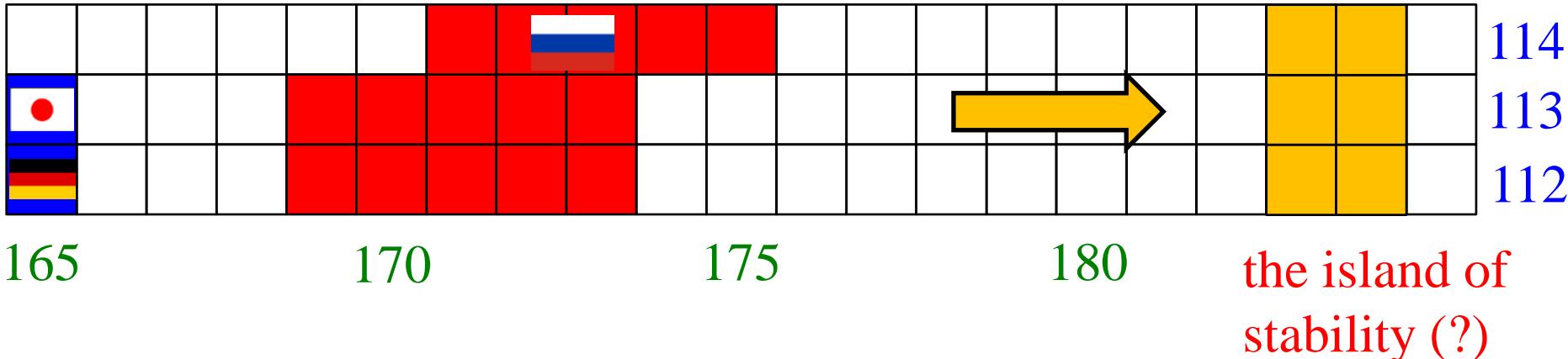
Yuri Oganessian

how to reach the island of stability?

Fusion of unstable nuclei



Fusion of unstable nuclei



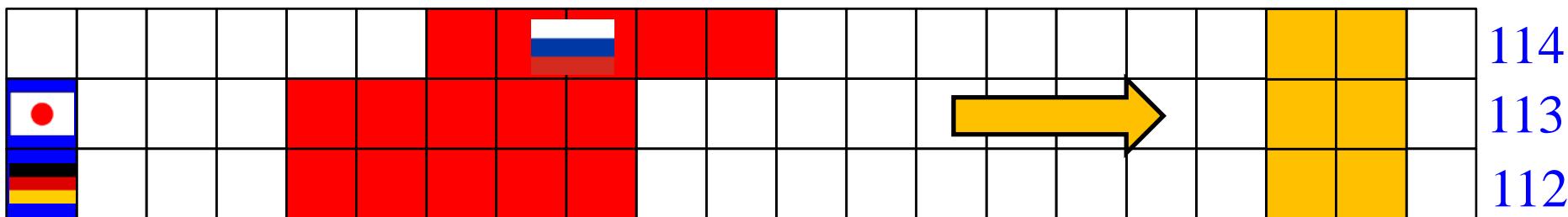
neutron-rich beams: indispensable

- how to deal with low beam intensity?
- reaction dynamics of neutron-rich beams?
 - ✓ capture: role of breakup and (multi-neutron) transfer?
 - ✓ diffusion: neutron emission during a shape evolution?
 - ✓ survival: validity of the statistical model?

structure of exotic nuclei

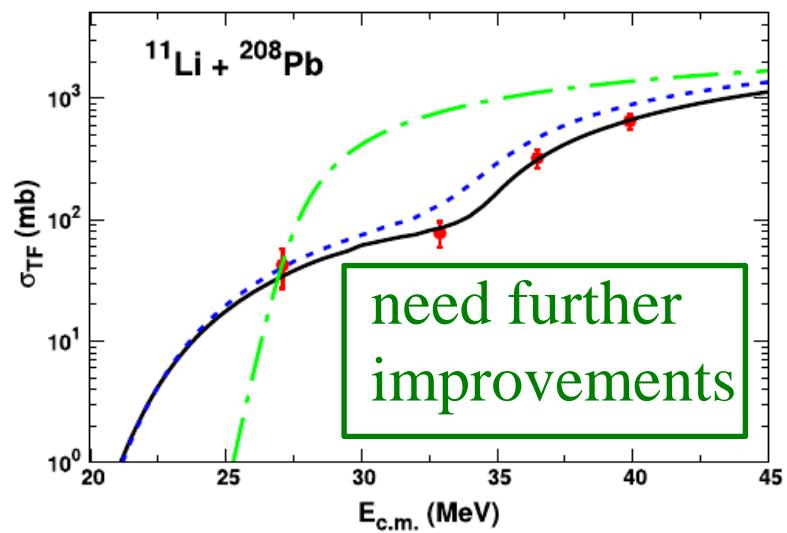
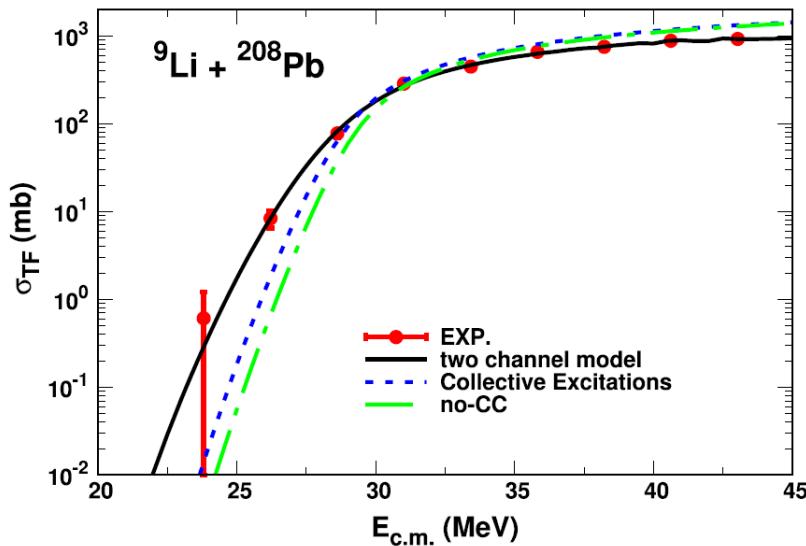
more studies are required

Fusion of unstable nuclei



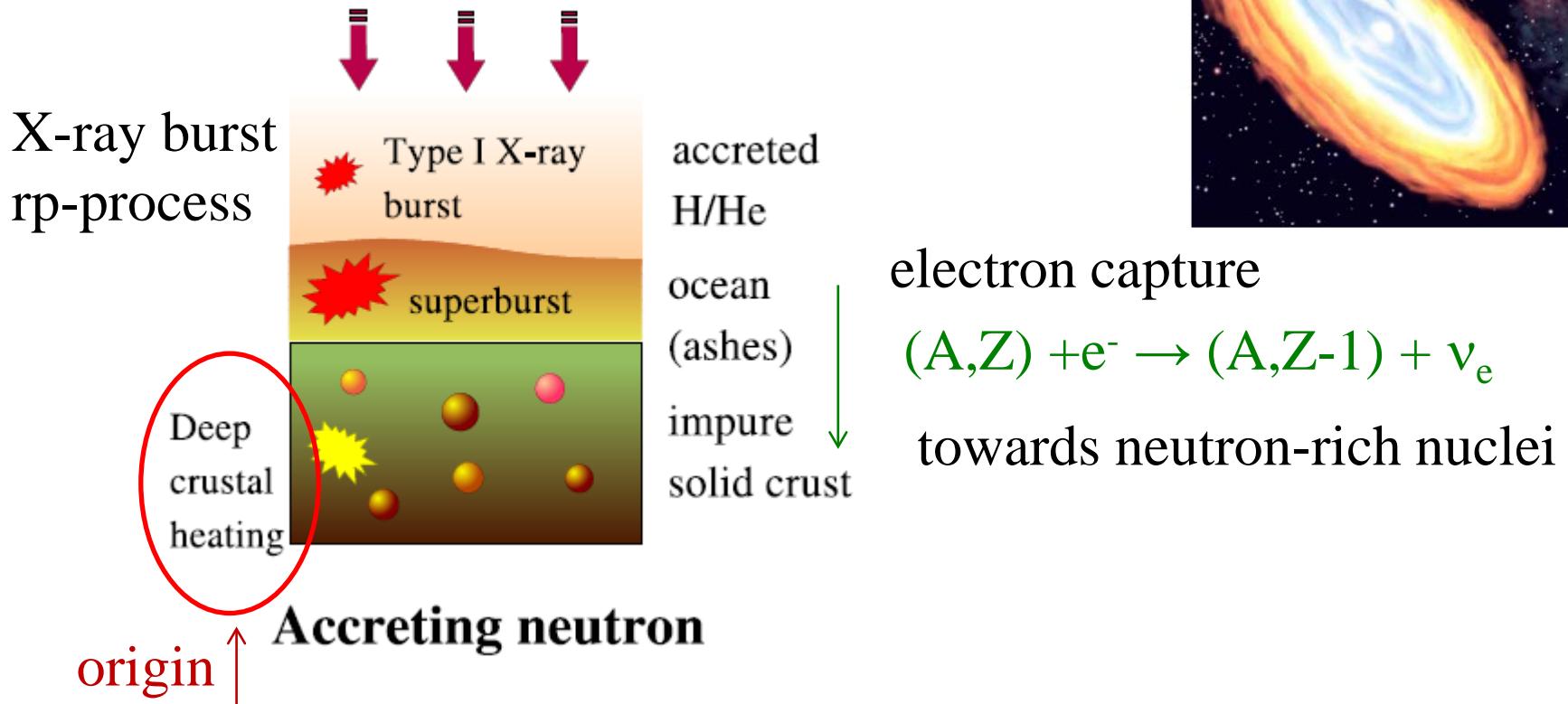
neutron-rich beams: indispensable
→ reaction dynamics?

K.-S. Choi, K. Hagino et al.,
Phys. Lett. B780 ('18) 455



simultaneous explanation for ${}^9\text{Li} + {}^{208}\text{Pb}$ and ${}^{11}\text{Li} + {}^{208}\text{Pb}$ with:
 ${}^{11}\text{Li} + {}^{208}\text{Pb} \longleftrightarrow {}^9\text{Li} + {}^{210}\text{Pb} \longleftrightarrow {}^7\text{Li} + {}^{212}\text{Pb}$ transfer couplings

fusion of neutron-rich nuclei in accreting (質量降着) neutron stars

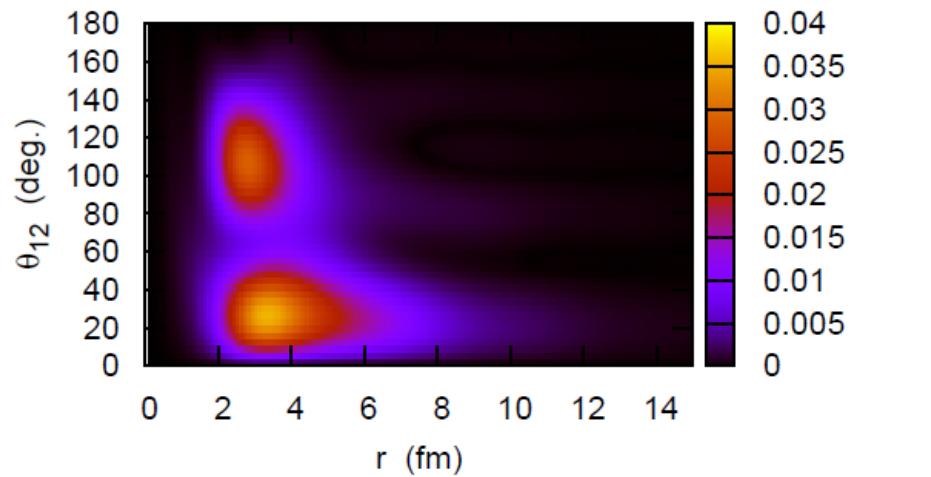
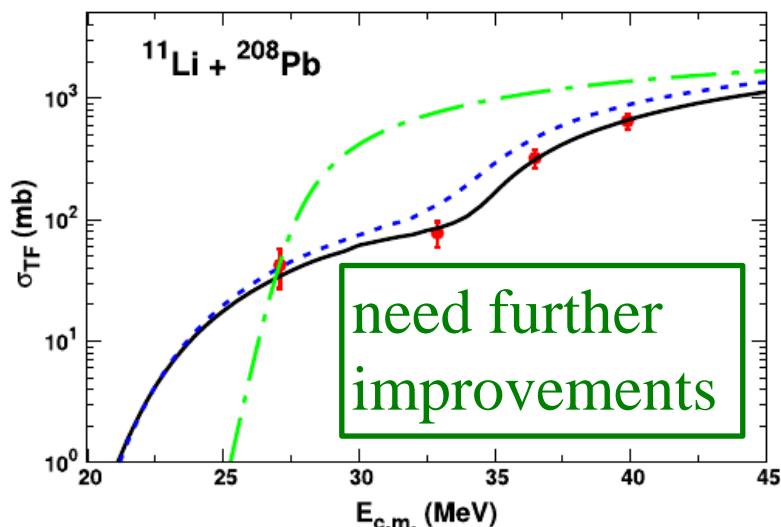
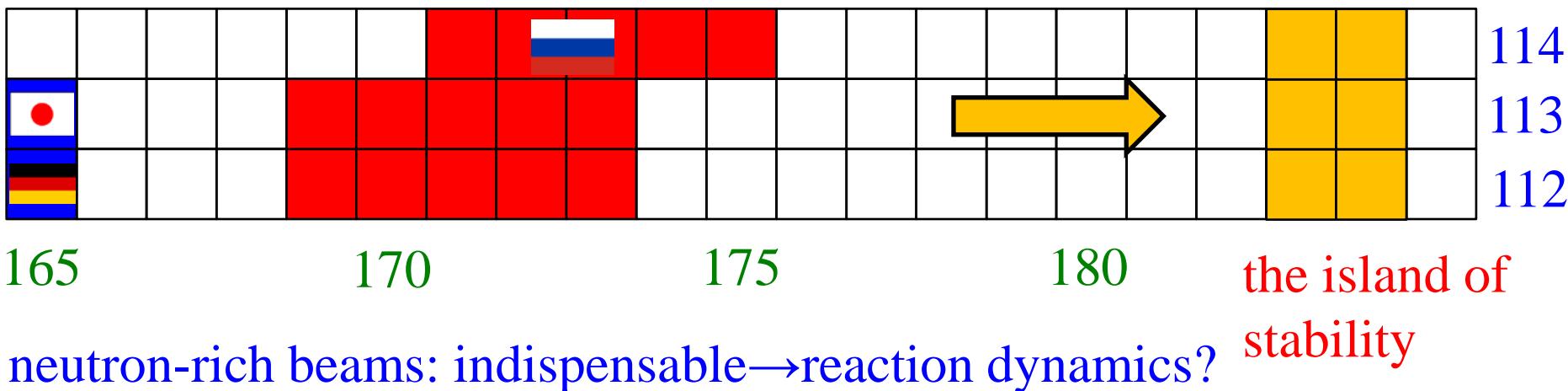


fusion of neutron-rich nuclei
when Z becomes small enough



N. Chamel and P. Haensel,
Living Rev. Relativity, 11 ('08) 10.

Fusion of unstable nuclei



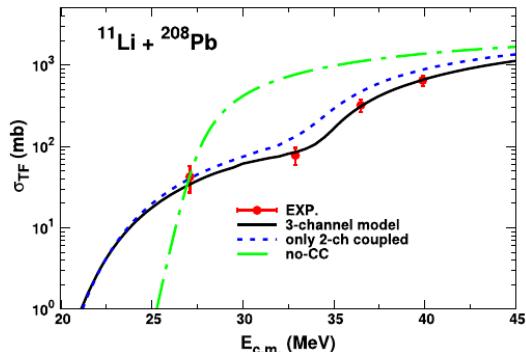
K.-S. Choi, K. Hagino et al.,
Phys. Lett. B780 ('18) 455

good understandings of the structure
of neutron-rich nuclei is also important

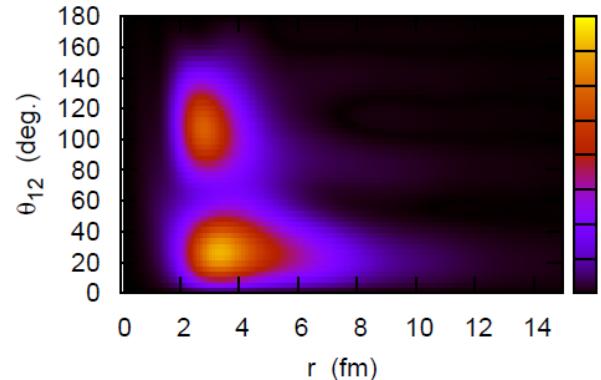
Summary

Physics of SHE

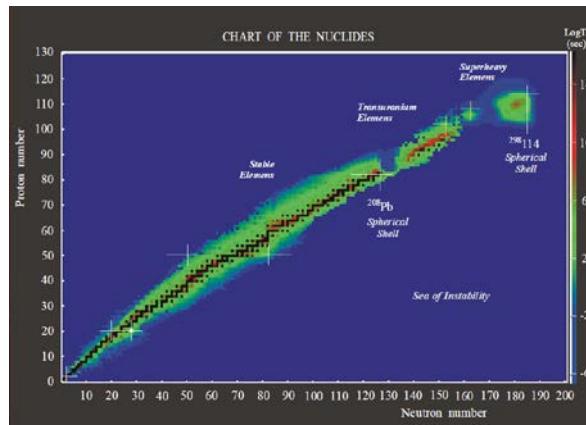
Reactions of n-rich nucl.



Structure of n-rich nucl.



- SHE
- the island of stability

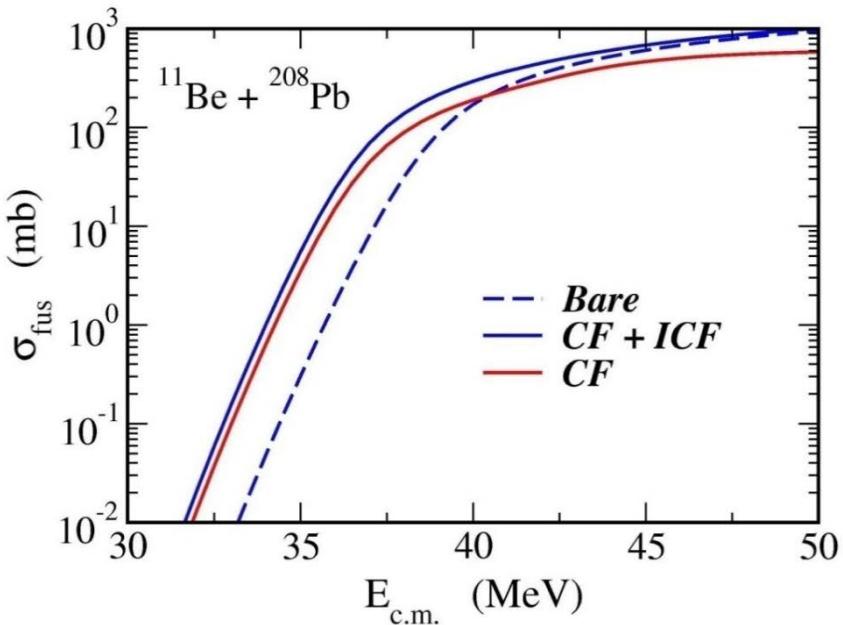


open quantum systems (OQS)

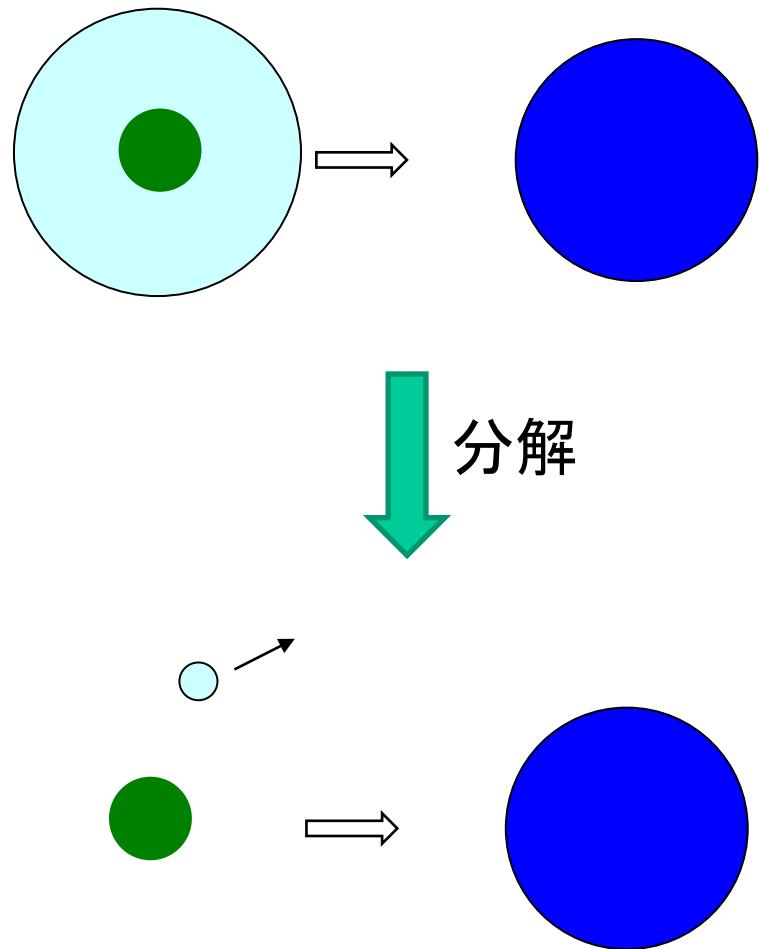
SHE + neutron-rich nuclei + OQS → new direction

➤ 安定の島に向けて

中性子過剰核ビームを用いた
実験が必要不可欠



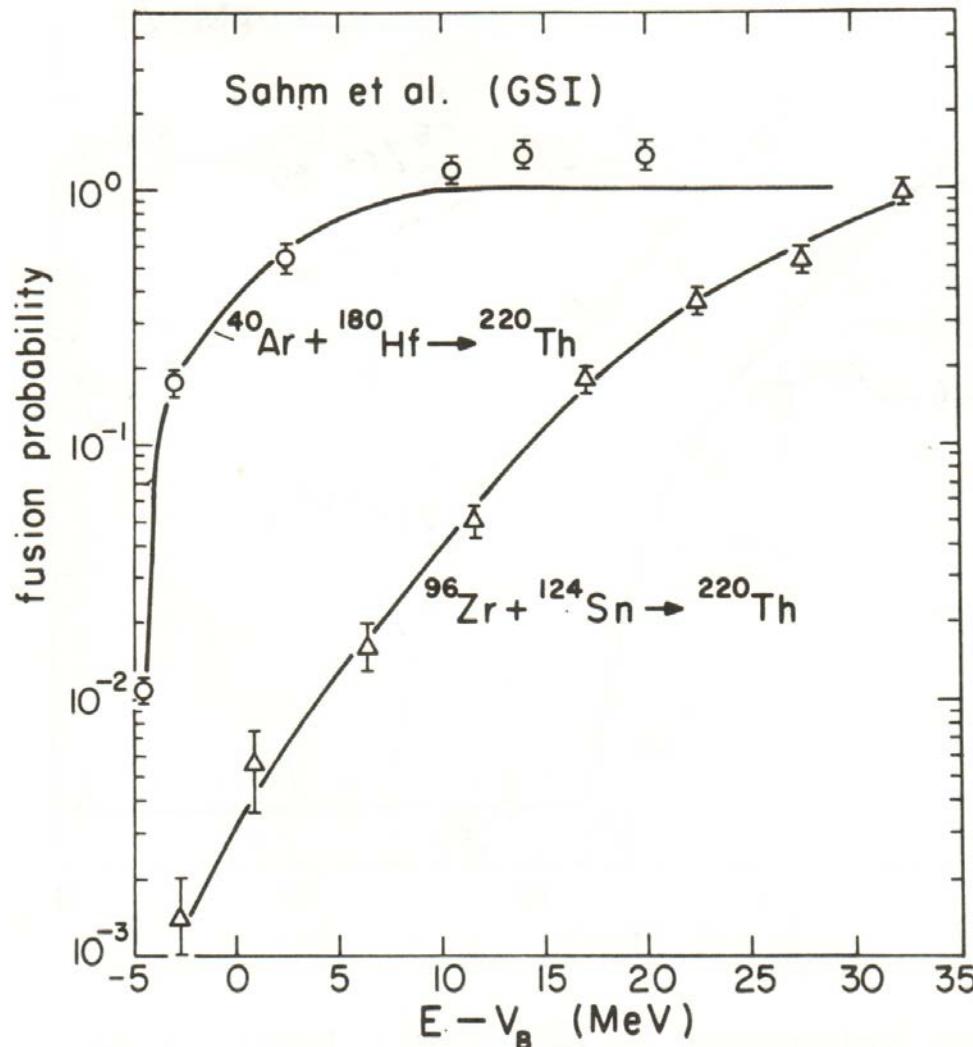
中性子過剰核 = 弱束縛



K. Hagino, A. Vitturi, C.H. Dasso,
and S.M. Lenzi, Phys. Rev. C61 ('00) 037602

反応機構の理解(分解、核子移行、融合)

fusion hindrance



C.C. Sahm et al.,
Z. Phys. A319 ('84) 113

New method for open quantum systems

M. Tokieda and K.H., Ann. of Phys., in press ('19)

$$H_{\text{tot}} = H_S + \sum_i \hbar \omega_i a_i^\dagger a_i + h(q) \sum_i d_i (a_i^\dagger + a_i)$$

system environment coupling

naive coupled-channels equations:

$$\Psi_{\text{tot}}(q, t) = \sum_{\{n_i\}} \psi_{\{n_i\}}(q, t) |\{n_i\}\rangle; \quad |\{n_i\}\rangle = \prod_i \frac{1}{\sqrt{n_i!}} (a_i^\dagger)^{n_i} |0\rangle$$

$$\rightarrow \langle \{n_i\} | i\hbar \frac{\partial}{\partial t} | \Psi_{\text{tot}} \rangle = \langle \{n_i\} | H_{\text{tot}} | \Psi_{\text{tot}} \rangle$$

→ coupled-channels eqs. for $\psi_{\{n_i\}}(q, t)$

difficult when the number of environmental osc. modes is large

$$H_{\text{tot}} = H_S + \sum_i \hbar \omega_i a_i^\dagger a_i + h(q) \sum_i d_i (a_i^\dagger + a_i)$$

→ introduce more efficient basis

$$e^{-i\omega t} \sim \sum_{k=1}^K \eta_k(\omega) u_k(t)$$

exp. basis
coef.

cf. correlation function

$$L(t) = \int_{-\infty}^{\infty} d\omega \frac{J(\omega)}{1 - e^{-\beta \hbar \omega}} e^{-i\omega t}$$

cf. 階層型運動方程式

Y. Tanimura and R. Kubo,
J. Phys. Soc. Jpn 58 ('89)101

$$b_k^\dagger = \sum_i \frac{d_i}{\hbar} \eta_k(\omega_i) a_i^\dagger$$

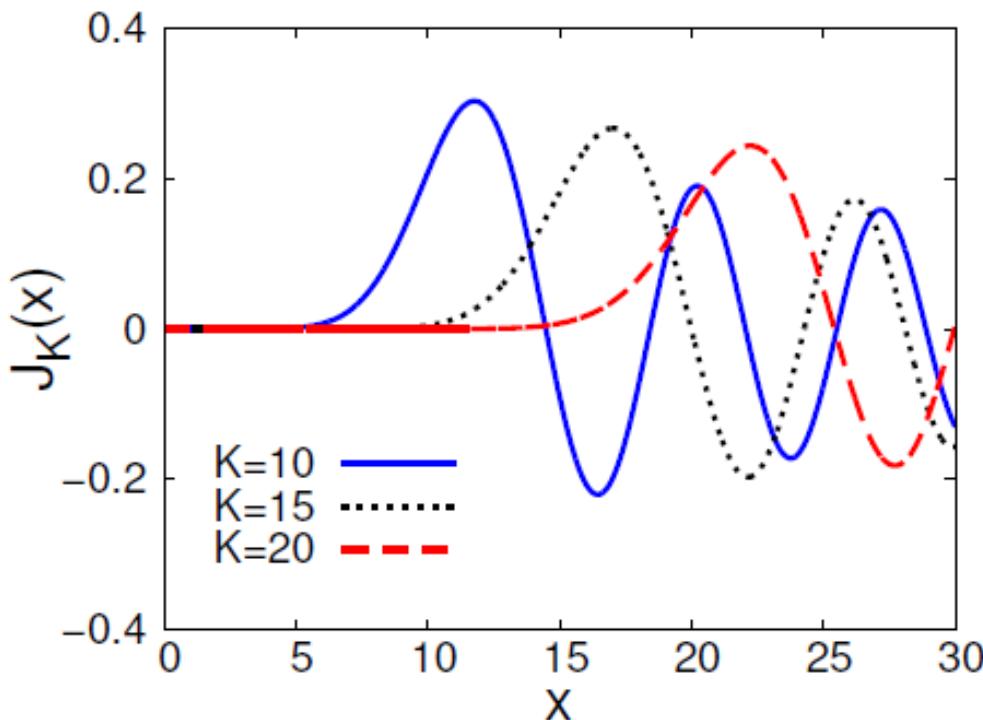
cf. Lanczos method

$$\Psi_{\text{tot}}(q, t) = \sum_{\{\tilde{n}_k\}} \tilde{\psi}_{\{\tilde{n}_k\}}(q, t) |\{\tilde{n}_k\}\rangle; \quad |\{\tilde{n}_k\}\rangle = \prod_{k=1}^K \frac{1}{\sqrt{\tilde{n}_k!}} (b_k^\dagger)^{\tilde{n}_k} |0\rangle$$

$$e^{-i\omega t} \sim \sum_{k=1}^K \eta_k(\omega) u_k(t) \quad \longrightarrow \quad b_k^\dagger = \sum_i \frac{d_i}{\hbar} \eta_k(\omega_i) a_i^\dagger$$

in actual calc.: expansion with Bessel function (Jacobi-Anger identity):

$$e^{-i\omega t} = J_0(\Omega t) + 2 \sum_{k=1}^{\infty} (-i)^k T_k \left(\frac{\omega}{\Omega} \right) J_k(\Omega t)$$



Chebyshev polynomials

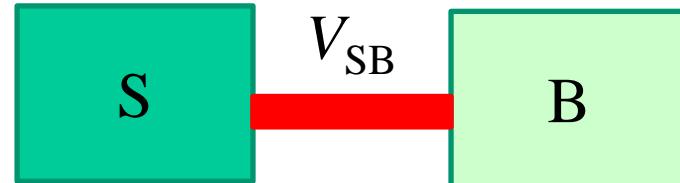
large k : does not contribute
when t is small

$$J_k(x) = \sum_{s=0}^{\infty} \frac{(-1)^s}{s!(s+k)!} \left(\frac{x}{2} \right)^{k+2s}$$

modelling of open quantum systems

i) system + bath

$$H = H_S + H_B + V_{SB}$$



- ✓ Caldeira-Leggett
- ✓ Feynmann-Vernon

solution:

a) eliminate B (bath)
→ eff. action for S
(influence functional)

b) $\rho_S = Tr_B[\rho]$
→ $i\hbar\dot{\rho}_S = \dots$

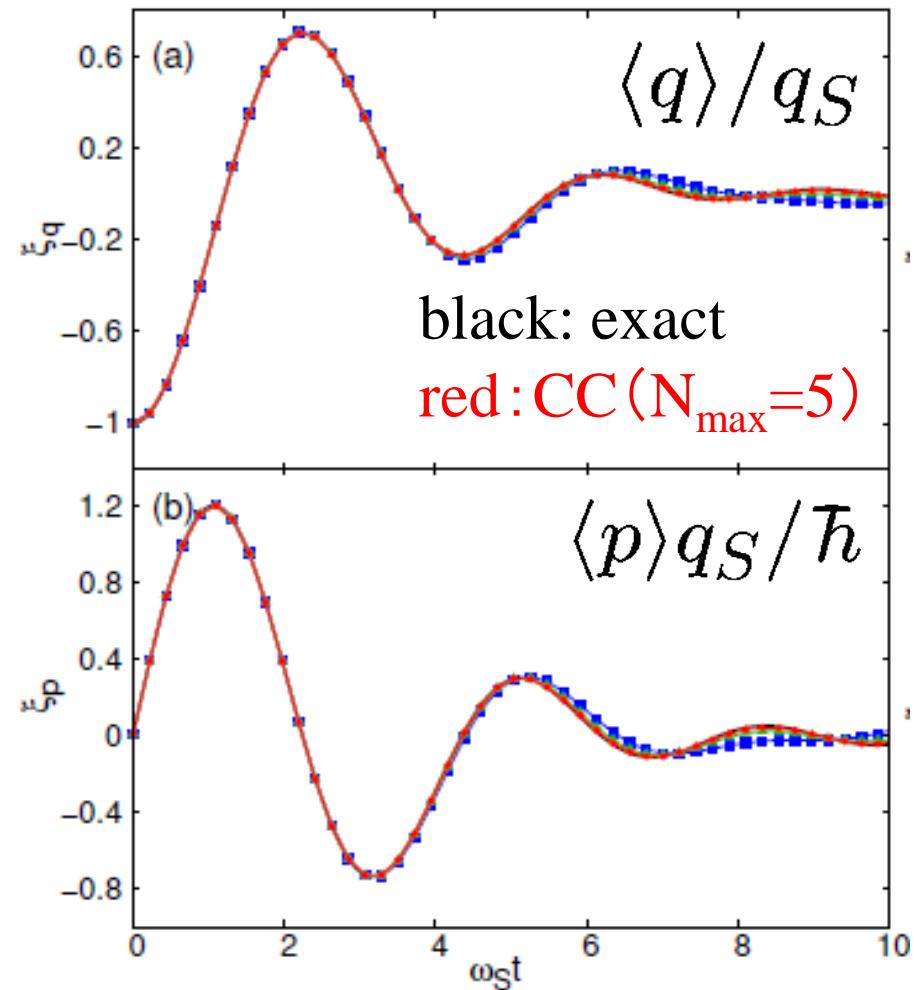
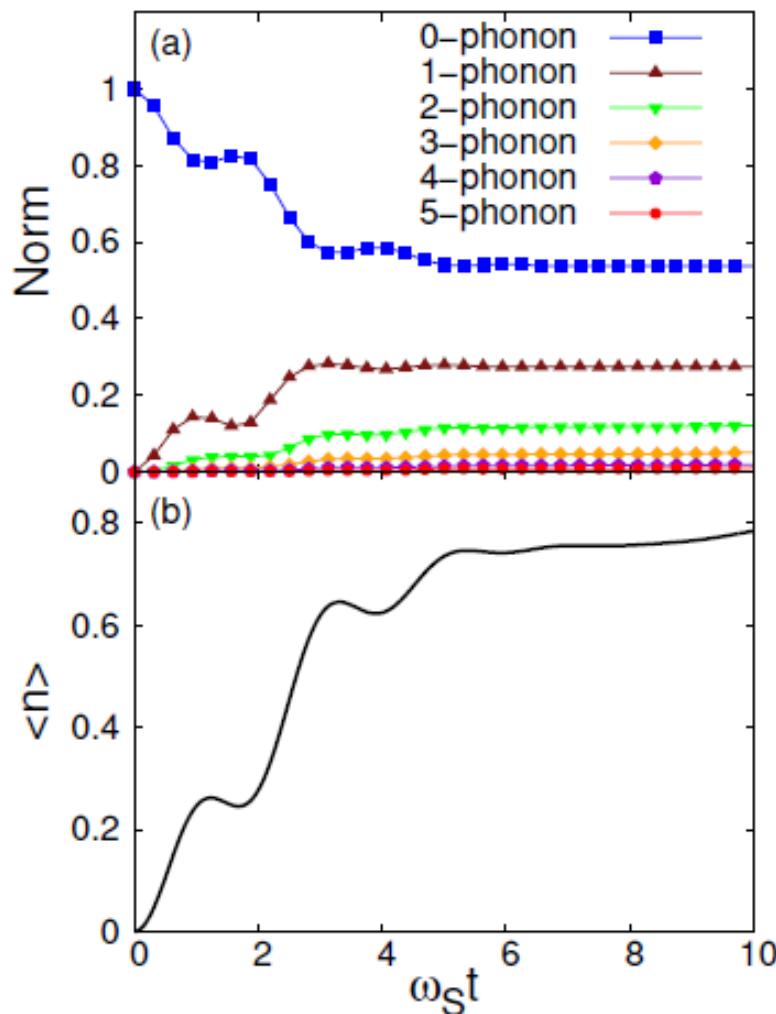
Markovian approximation
→ Lindblad equation

c) expand the tot. wf with the eigen states of H_B (coupled-channels eq.)

M. Tokieda and K.H.,
Ann. of Phys. 412 ('20) 168005.
Front. in Phys. 8 ('20) 8.

application to damped oscillation: HO + environment (conti. spectrum)

expansion with Bessel functions up to K=20



modelling of open quantum systems

i) system + bath

$$H = H_S + H_B + V_{SB}$$

- ✓ Caldeira-Leggett
- ✓ Feynmann-Vernon

solution:

a) eff. action for S
(influence functional)

b) Lindlad eq.

c) coupled-channels eqs.

- microscopic
- but, hard to solve

ii) quantum friction model

construct a Hamiltonian which
leads to

$$\frac{d}{dt}\langle p \rangle = -\langle V'(x) \rangle - \gamma \langle p \rangle$$

- ✓ E. Kanai, PTP 3 ('48)
- ✓ M.D. Kostin, JCP 57 ('72)
- ✓ K. Albrecht, PLB56 ('75)
- ✓ K.-K. Kan & J.-J. Griffin,
PLB50 ('74)
- ✓ A. Bulgac, S. Jin, and I. Stetcu,
PRC100, 014615 (2019)

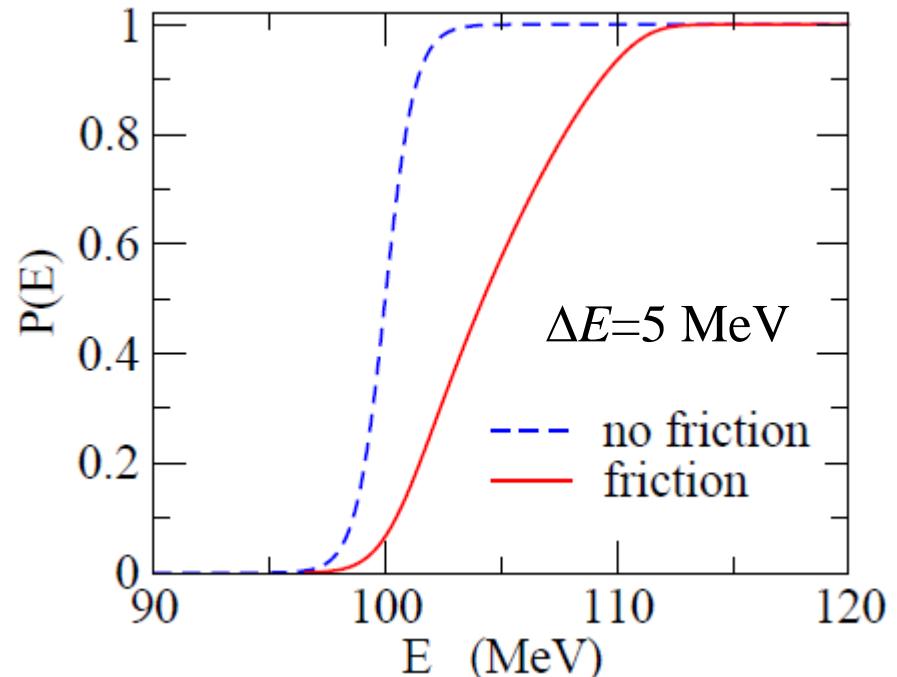
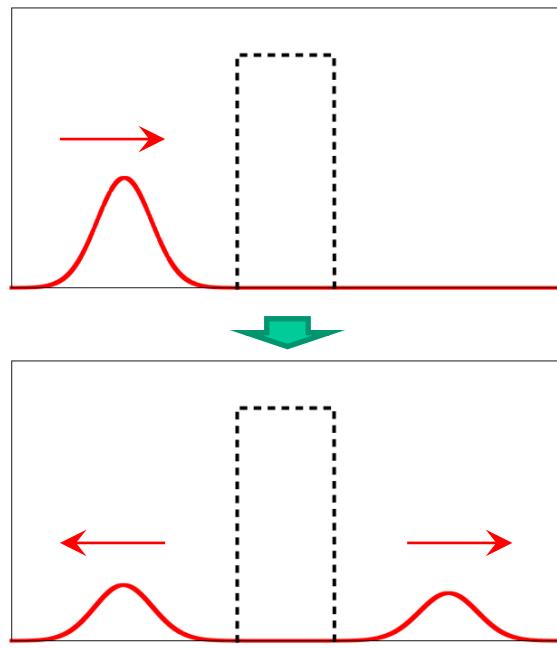
Quantum friction classical eq. of motion $\dot{p} = -V'(x) - \gamma p$

a quantization: Kanai model E. Kanai, PTP 3 (1948) 440

$$H = \frac{p^2}{2m} + V(x) \rightarrow \frac{\pi^2}{2m} e^{-\gamma t} + e^{\gamma t} V(x) \quad (\pi = e^{\gamma t} p)$$

$$\longrightarrow \frac{d}{dt} \langle p \rangle = -\langle V'(x) \rangle - \gamma \langle p \rangle$$

time-dep. wave packet approach



modelling of open quantum systems

i) system + bath

$$H = H_S + H_B + V_{SB}$$

- ✓ Caldeira-Leggett
- ✓ Feynmann-Vernon
- microscopic
- but, hard to solve

ii) quantum friction model

construct Hamiltonian which
leads to

$$\frac{d}{dt} \langle p \rangle = -\langle V'(x) \rangle - \gamma \langle p \rangle$$

- phenomenological
- easy to solve

