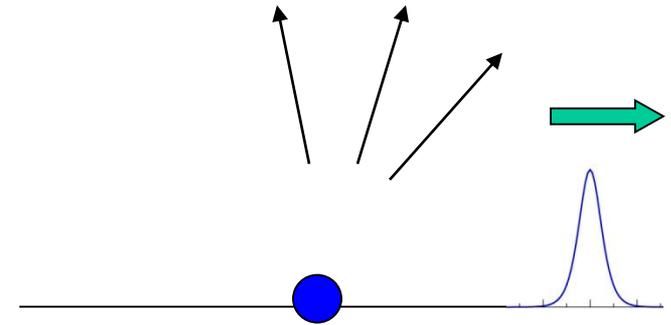
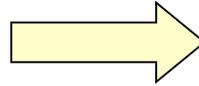


Absorption cross sections

Reaction processes

- Elastic scatt.
- Inelastic scatt.
- Transfer reaction
- Compound nucleus formation (fusion)



Loss of incident flux
(absorption)

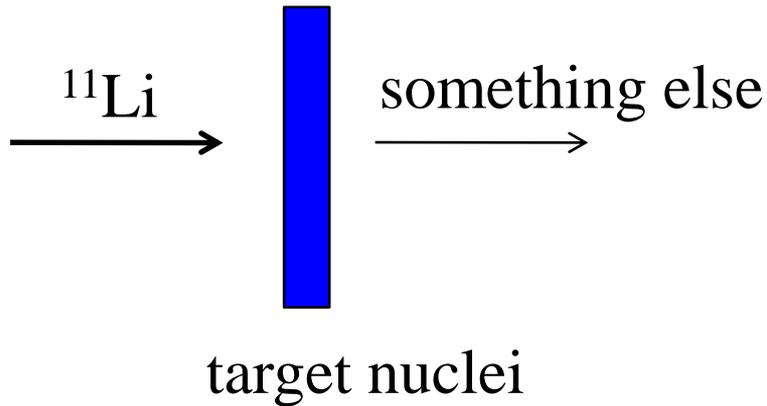
reaction cross sections

total scattering cross section - elastic cross section

$$\sigma_R = \sigma_{\text{tot}} - \sigma_{\text{el}}$$

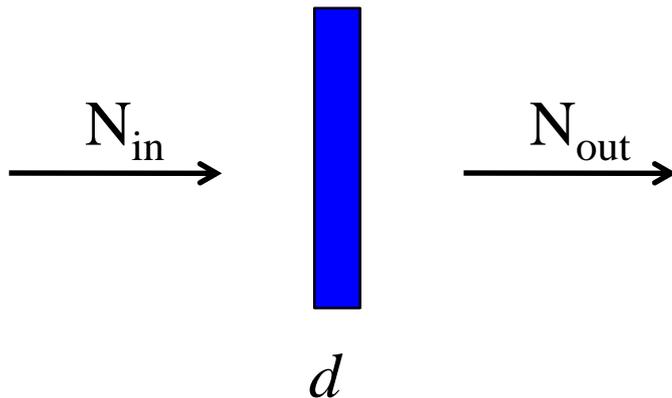
- fusion
- inelastic
- transfer

Interaction cross sections and halo nuclei



interaction cross section σ_I
= cross section for the change
of Z a/o N in the incident nucleus

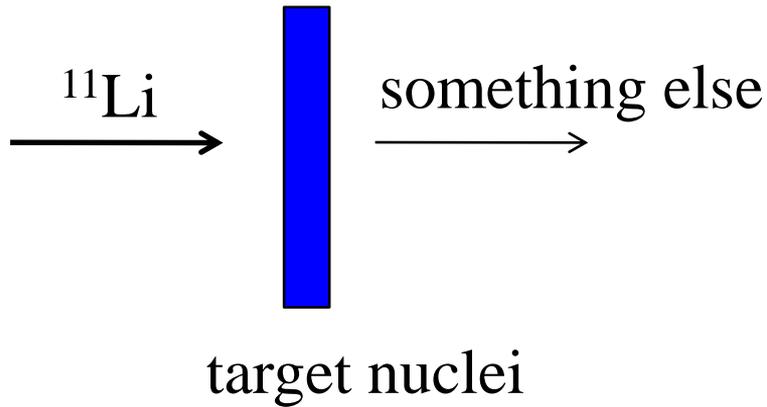
transmission method



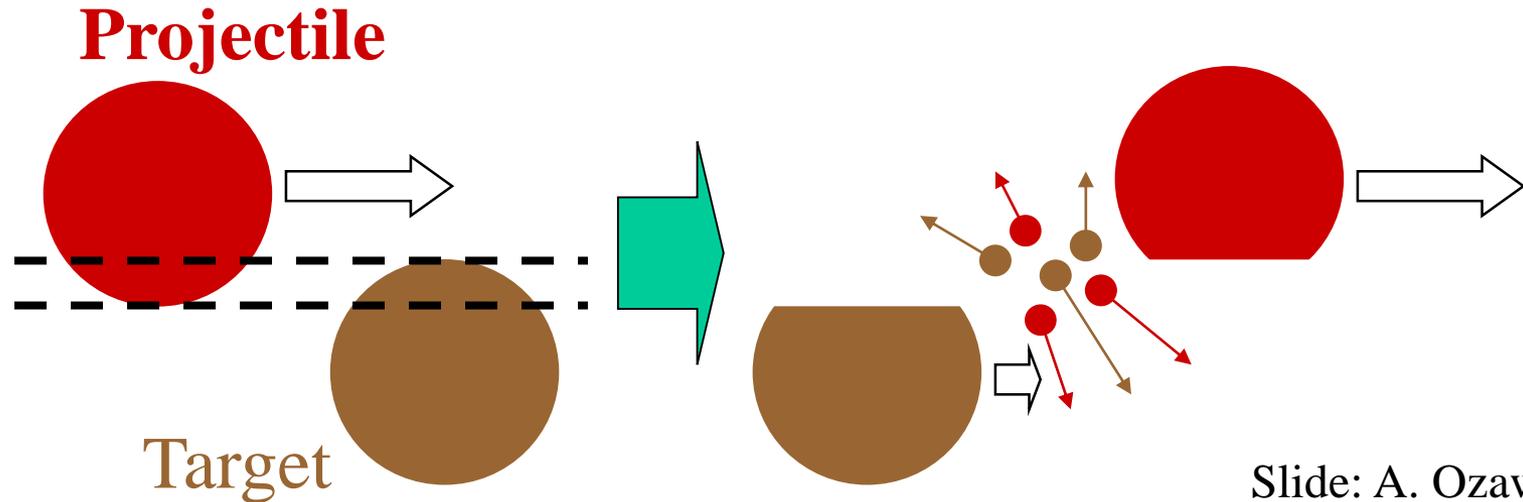
$$\sigma_R = -\frac{1}{t} \ln \left(\frac{N_{\text{out}}}{N_{\text{in}}} \right)$$

$$t = \rho_T \cdot d \cdot \epsilon$$

Interaction cross sections and halo nuclei



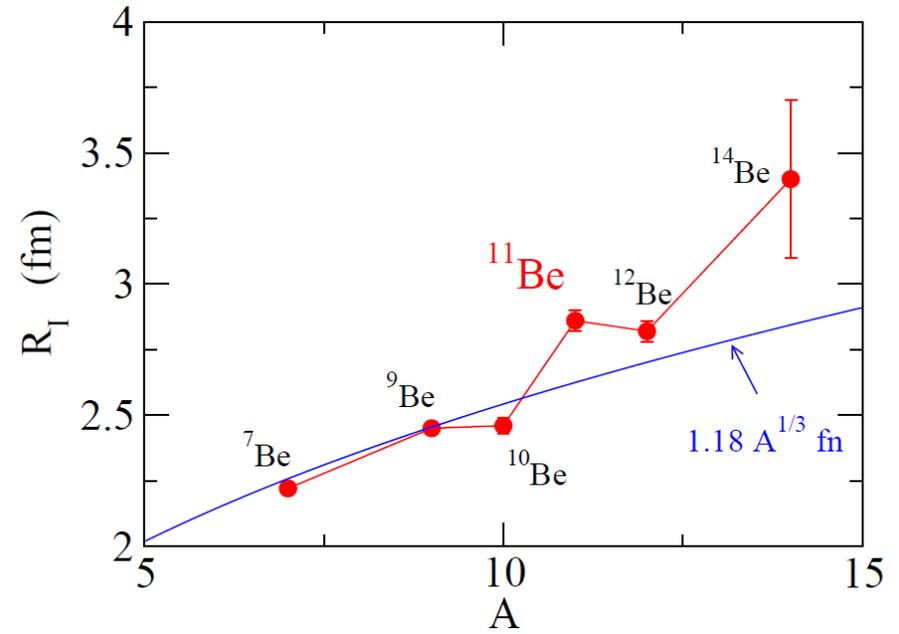
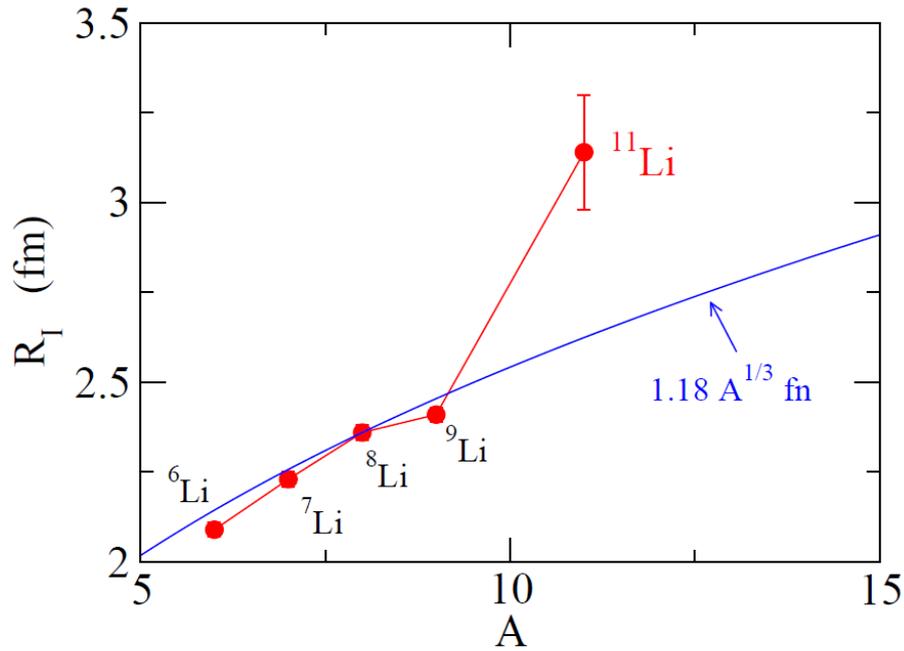
interaction cross section σ_I
= cross section for the change
of Z a/o N in the incident nucleus



Slide: A. Ozawa

$$\sigma_I \sim \pi [R_I(P) + R_I(T)]^2 \longrightarrow R_I(P)$$

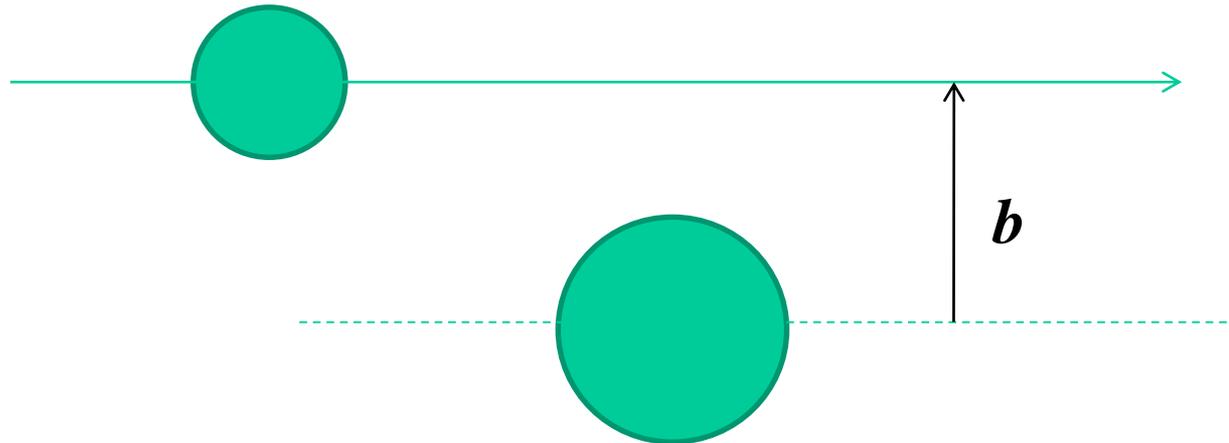
Discovery of halo nuclei



I. Tanihata, T. Kobayashi, O. Hashimoto et al., PRL55('85)2676; PLB206('88)592



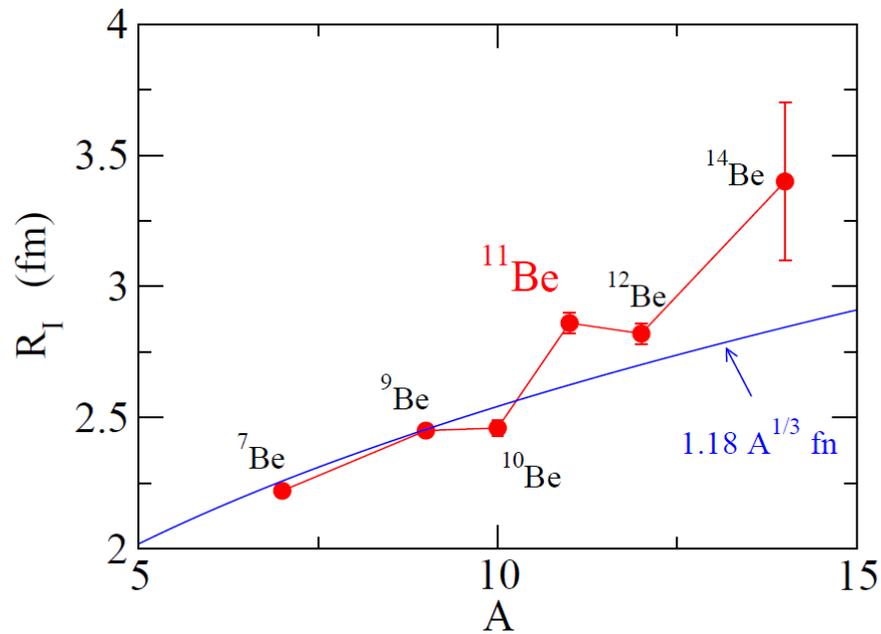
Reaction cross sections



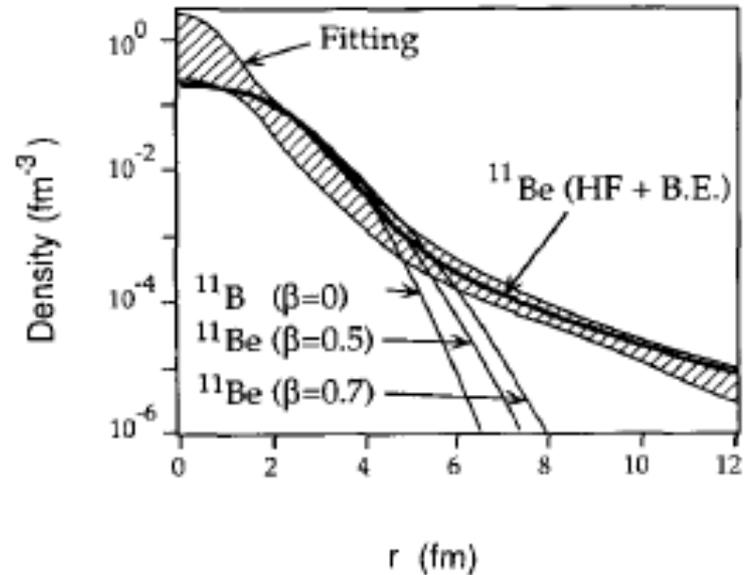
Glauber theory (optical limit approximation : OLA)

$$\sigma_R \sim 2\pi \int_0^\infty b db \left[1 - \exp \left(-\sigma_{NN} \int d^2s \rho_P^{(z)}(\mathbf{s}) \rho_T^{(z)}(\mathbf{s} - \mathbf{b}) \right) \right]$$

- straight-line trajectory (high energy scattering)
- adiabatic approximation
- simplified treatment for multiple scattering: $(1 - x)^N \rightarrow e^{-Nx}$



Density distribution which explains the experimental σ_R

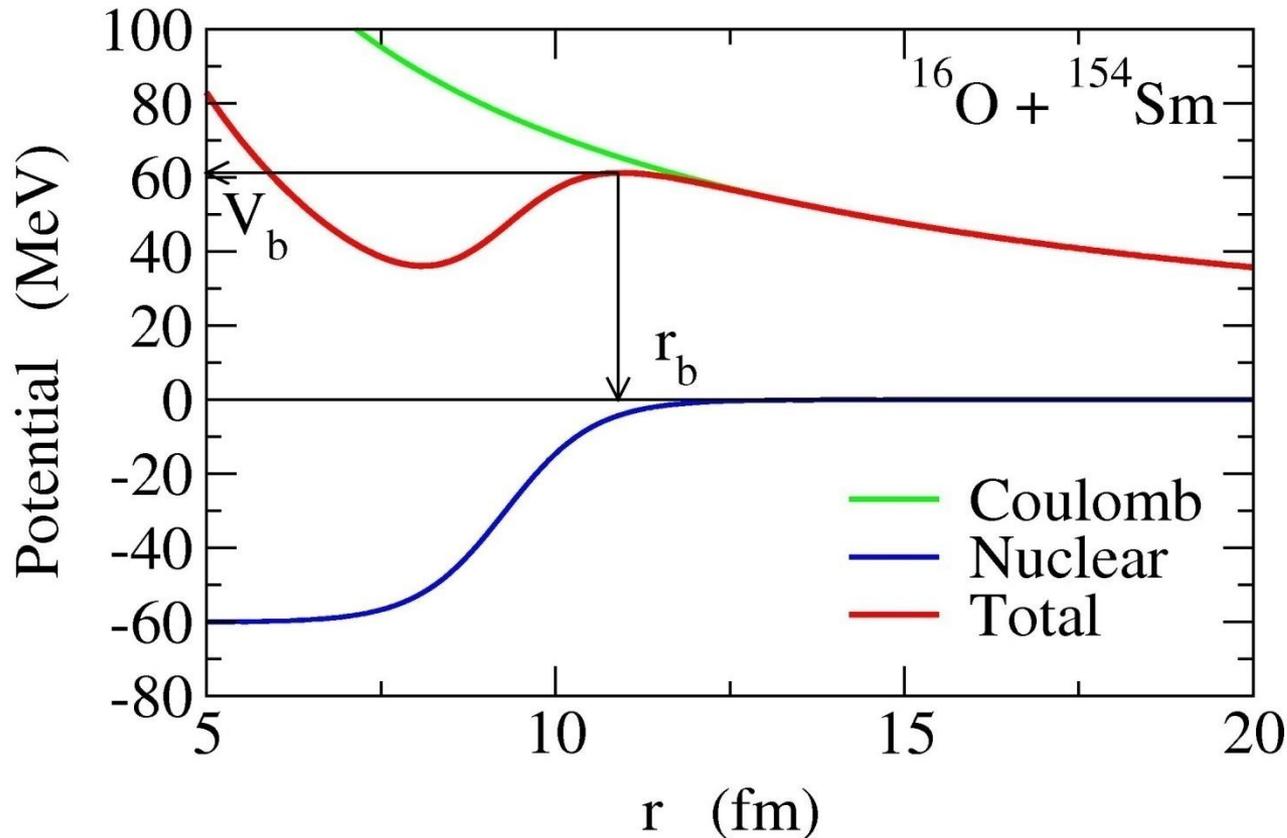


M. Fukuda et al., PLB268('91)339

$$\sigma_R \sim 2\pi \int_0^\infty b db \left[1 - \exp \left(-\sigma_{NN} \int d^2s \rho_P^{(z)}(s) \rho_T^{(z)}(s-b) \right) \right]$$

Heavy-ion subbarrier fusion reactions

Inter-nucleus potential



- above barrier
- sub-barrier
- deep subbarrier

Two forces:

1. Coulomb force

Long range,
repulsive

2. Nuclear force

Short range,
attractive



Potential barrier due to the compensation between the two
(Coulomb barrier)

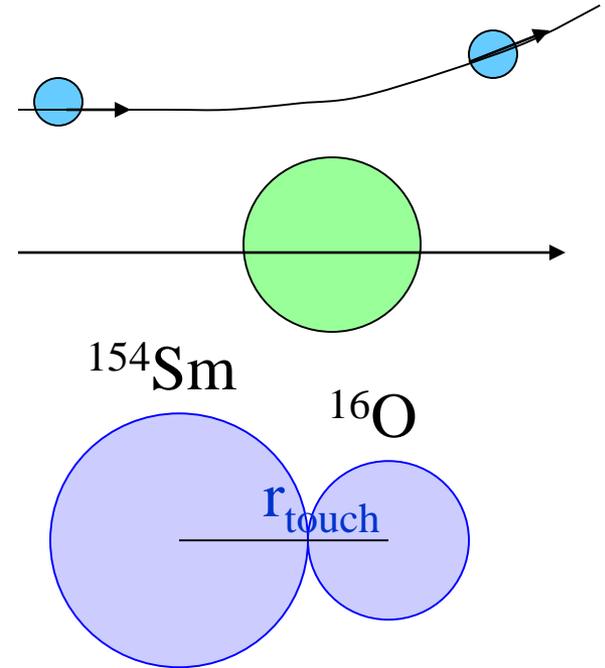
Three important features of heavy-ion reactions

1. Coulomb interaction: important

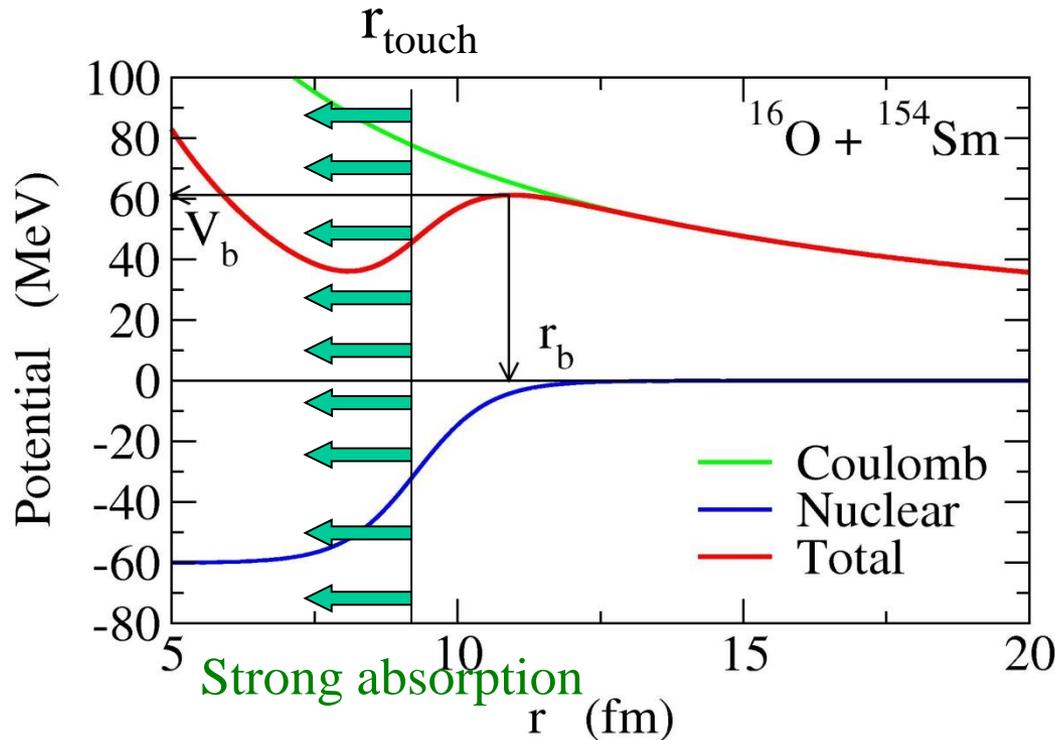
2. Reduced mass: large \longrightarrow

$$\mu = \frac{m_T m_P}{m_T + m_P}$$

(semi-) classical picture
concept of trajectory

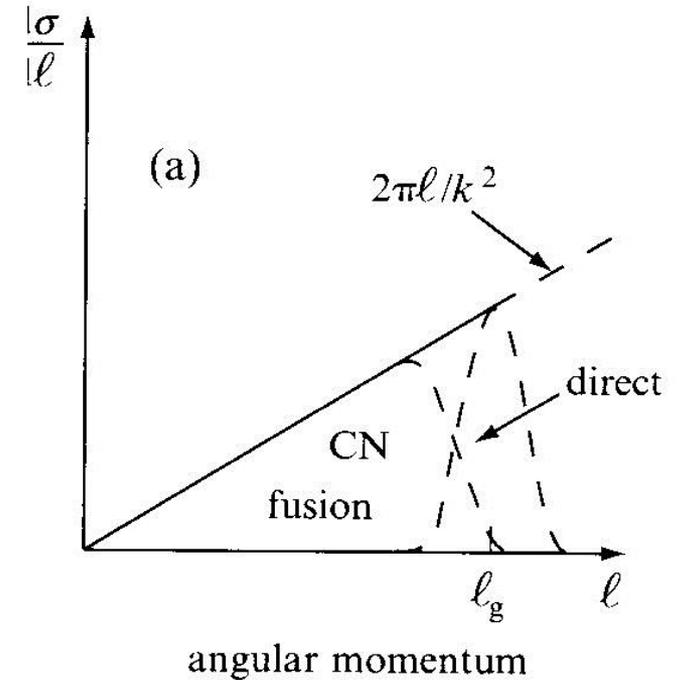
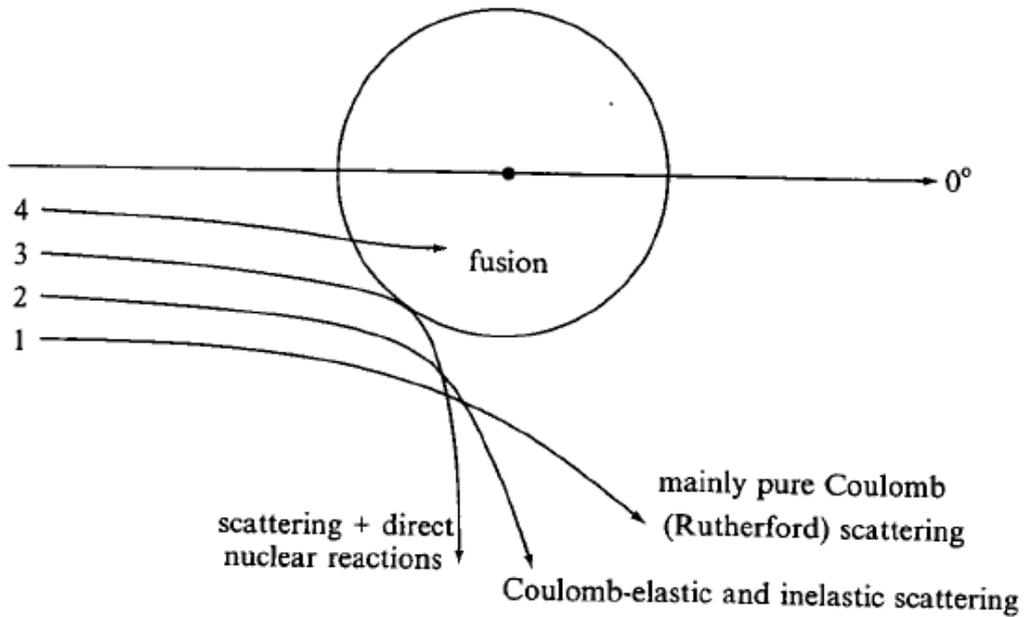


3. Strong absorption inside the Coul. barrier



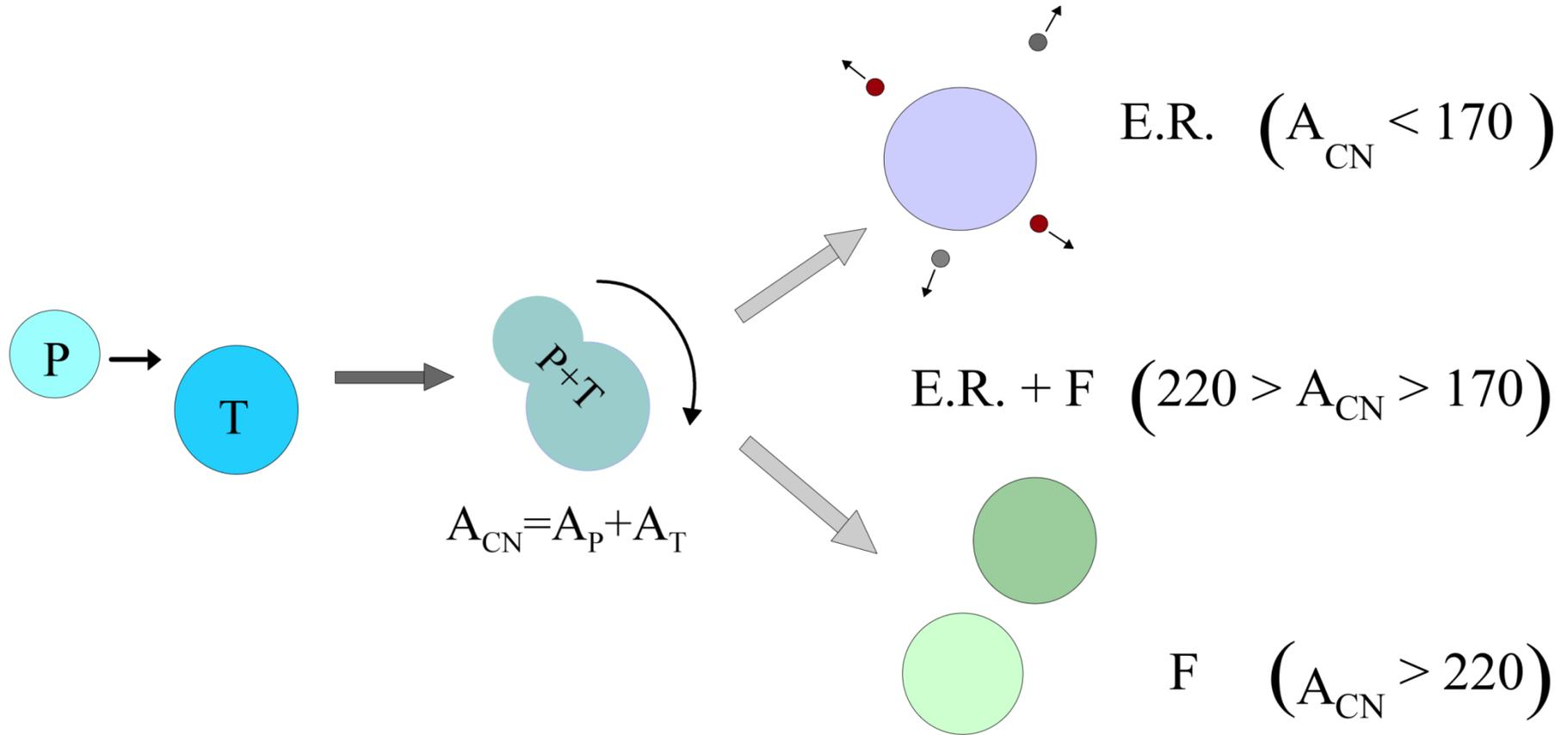
Automatic compound nucleus formation once touched (assumption of strong absorption)

Partial decomposition of reaction cross section



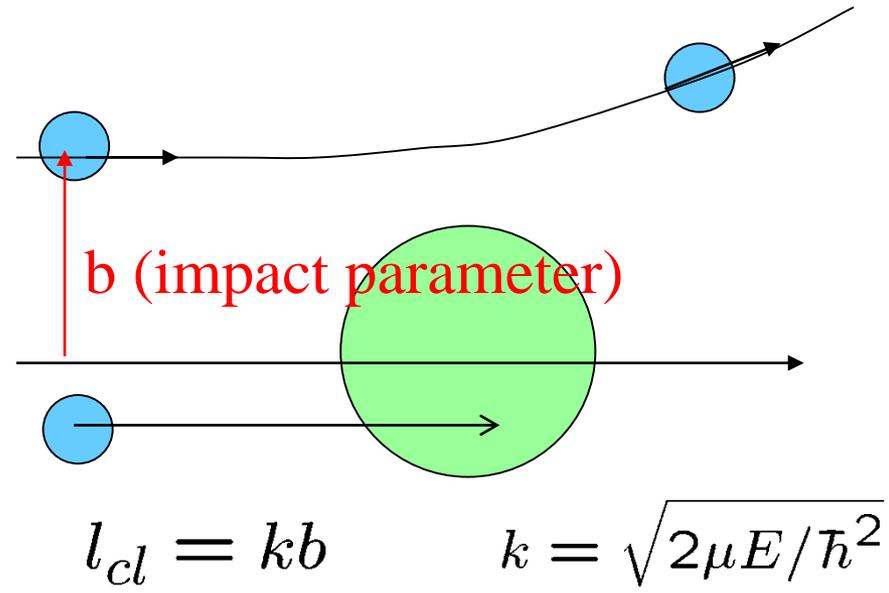
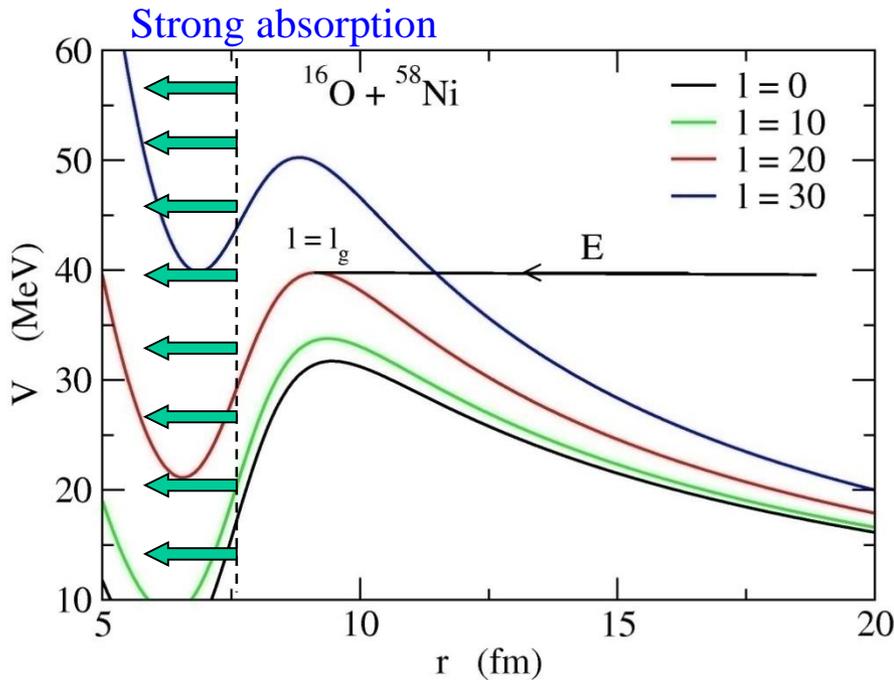
Taken from J.S. Lilley,
"Nuclear Physics"

Fusion: compound nucleus formation



courtesy: Felipe Canto

classical fusion cross sections

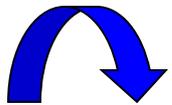


$l < l_g$: can access to the strong absorption region classically

$\Rightarrow b_g = l_g/k$

$$\sigma^{cl} = 2\pi \int_0^{b_g} b db = \pi b_g^2$$

$$V_b + \frac{(kb_g)^2 \hbar^2}{2\mu R_b^2} = E$$

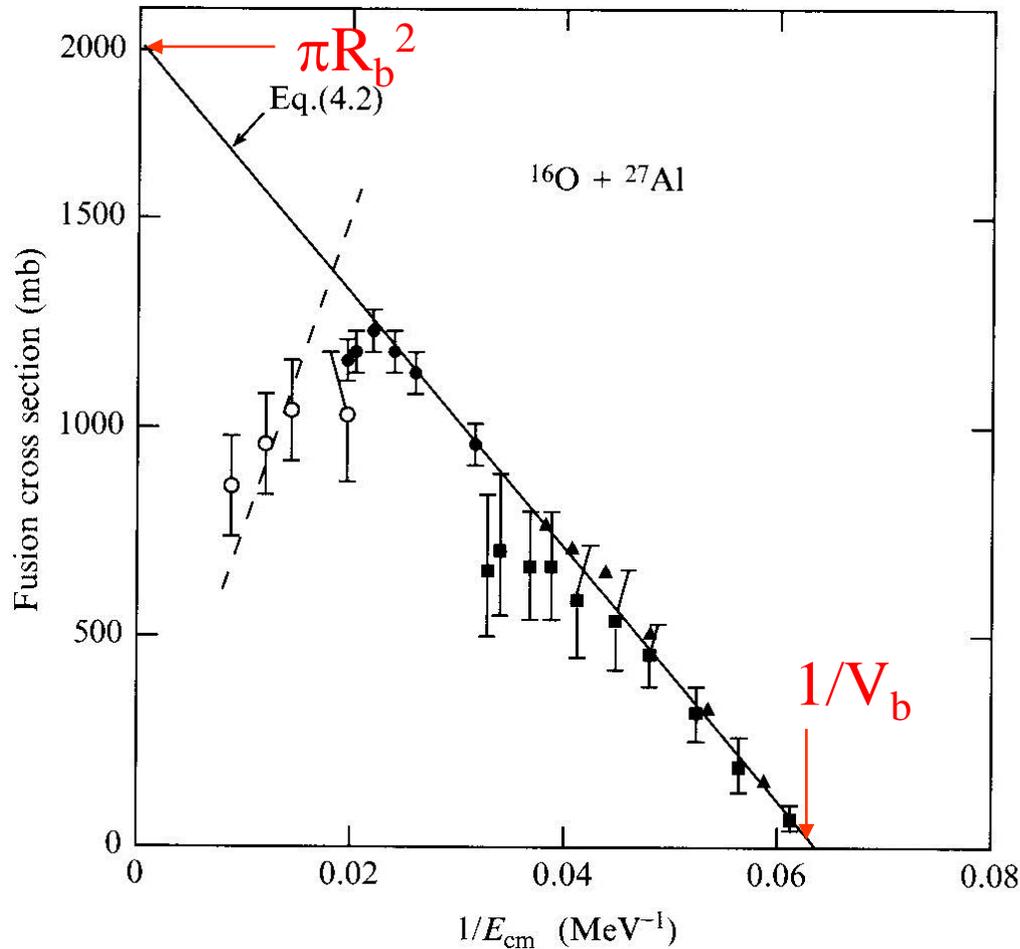


$$\sigma_{fus}^{cl}(E) = \pi R_b^2 \left(1 - \frac{V_b}{E} \right)$$

σ_{fus} vs $1/E$ (~70's)

Classical fusion cross section is proportional to $1/E$:

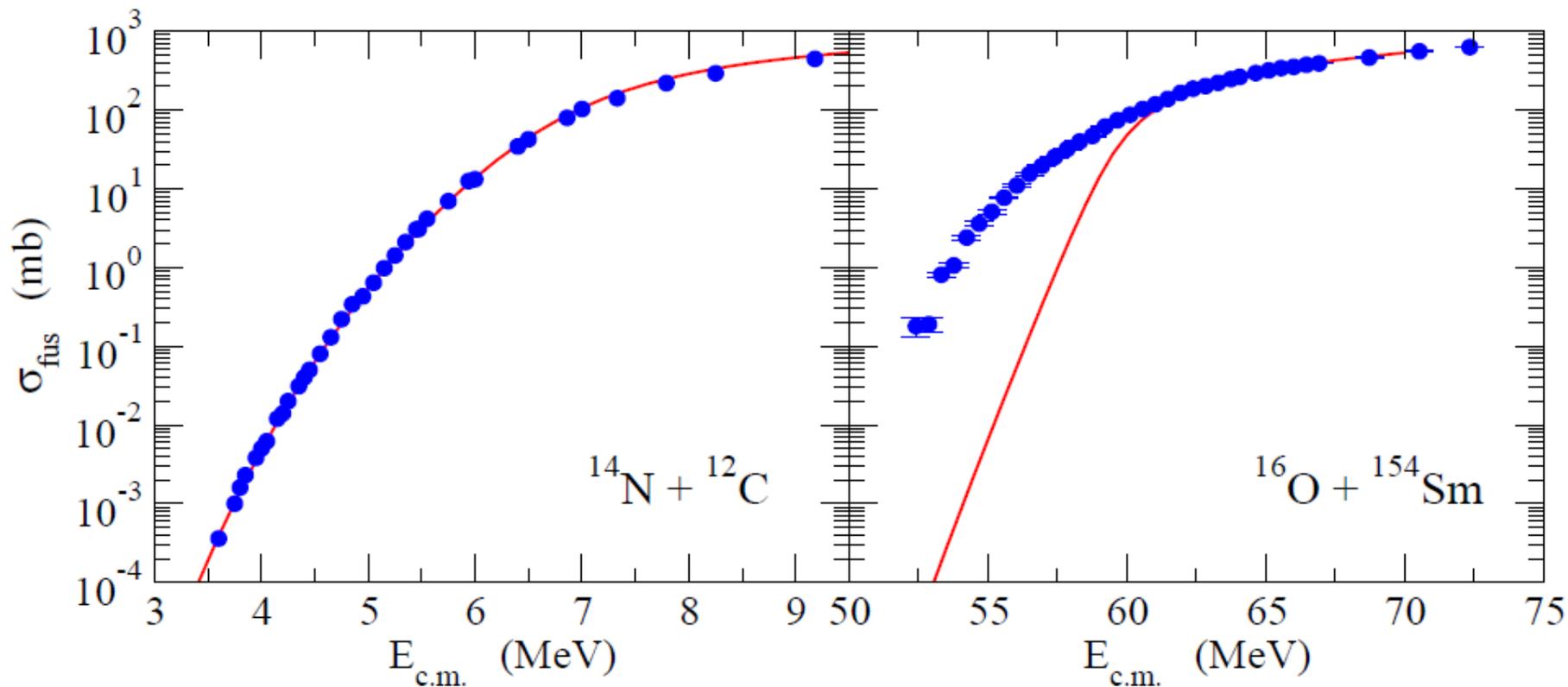
$$\sigma_{\text{fus}}^{\text{cl}}(E) = \pi R_b^2 \left(1 - \frac{V_b}{E}\right)$$



Taken from J.S. Lilley,
"Nuclear Physics"

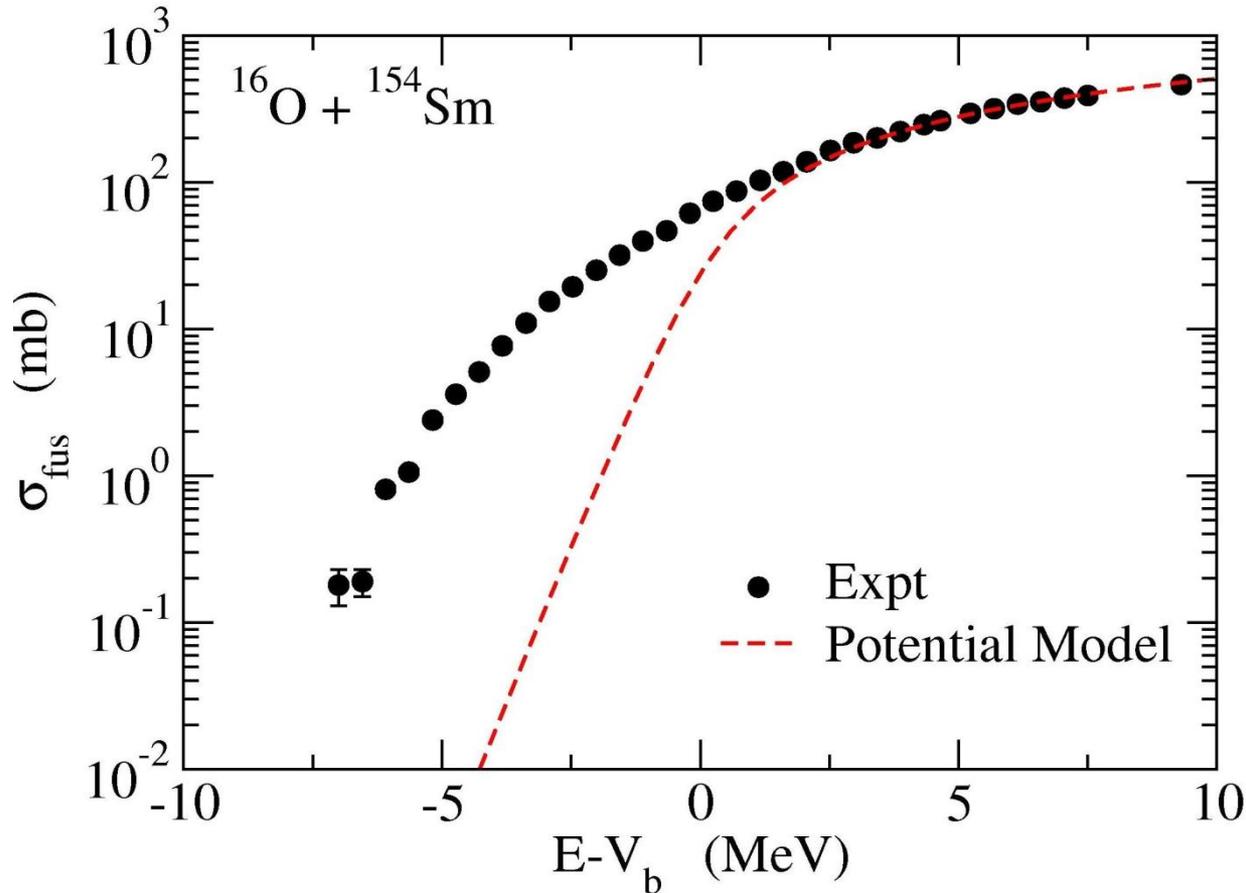
Fusion cross sections at subbarrier energies

Fusion cross sections of structure-less nuclei (a potential model)



Simple potential model:

- OK for relatively light systems
- underestimates σ_{fus} for heavier systems at subbarrier energies



Potential model:

Reproduces the data reasonably well for

$$E > V_b$$

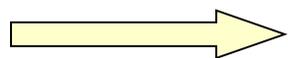
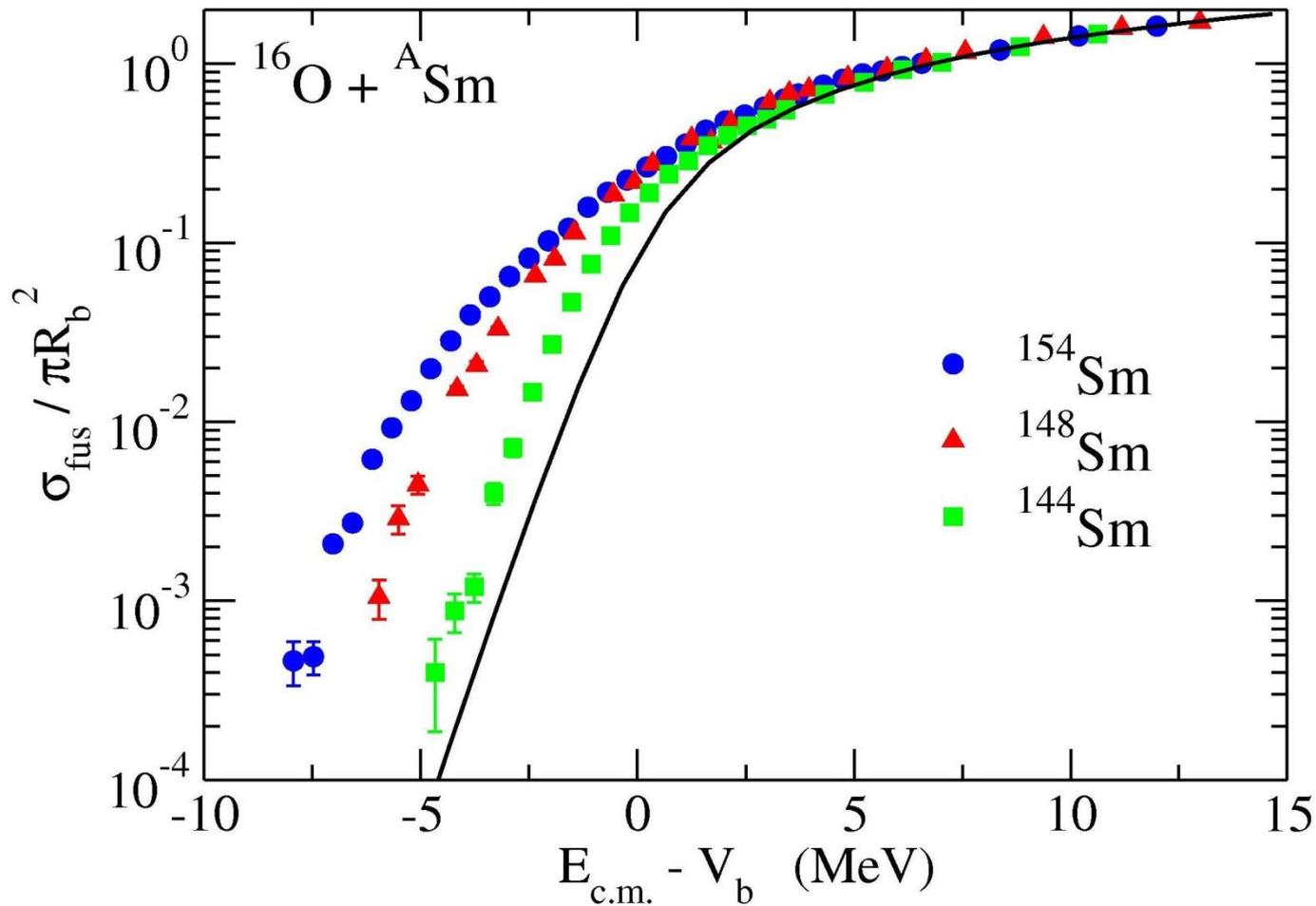
Underpredicts σ_{fus} for

$$E < V_b$$

cf. seminal work:

R.G. Stokstad et al., PRL41('78)465

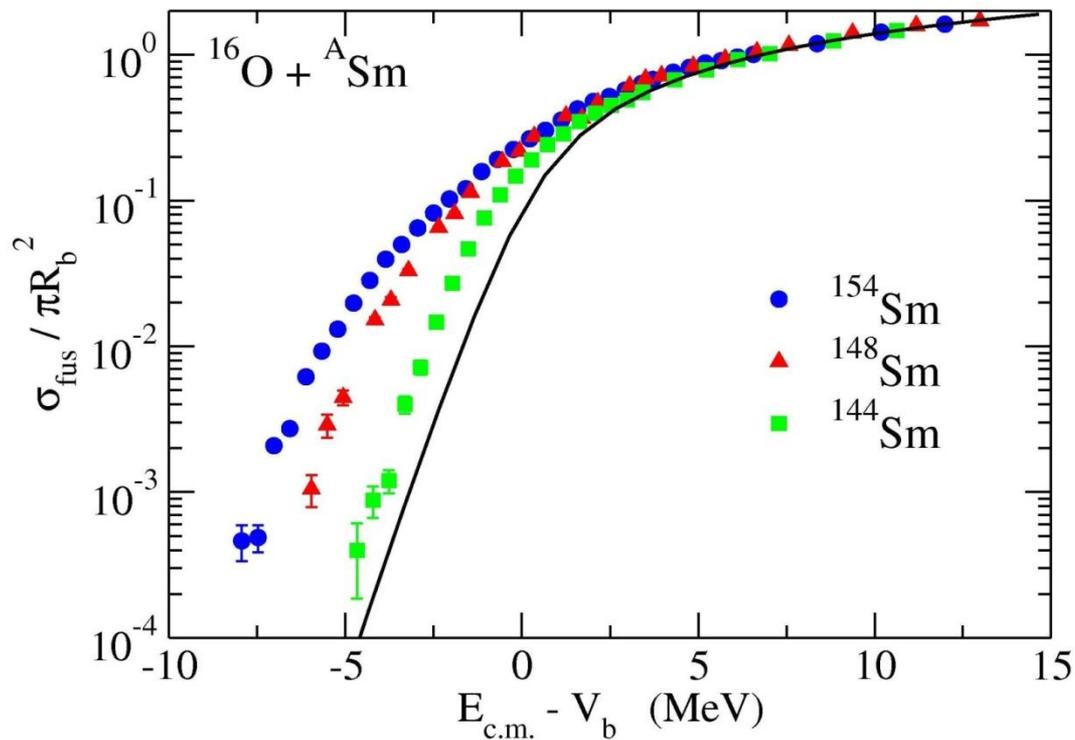
PRC21('80)2427



Strong target dependence at $E < V_b$



low-lying collective excitations?



(MeV)

1.81 — 3^-
1.66 — 2^+

(MeV)

1.18 — 4^+
1.16 — 3^-

(MeV)

0.90 — 8^+
0.54 — 6^+
0.27 — 4^+
0.082 — 2^+
0 — 0^+

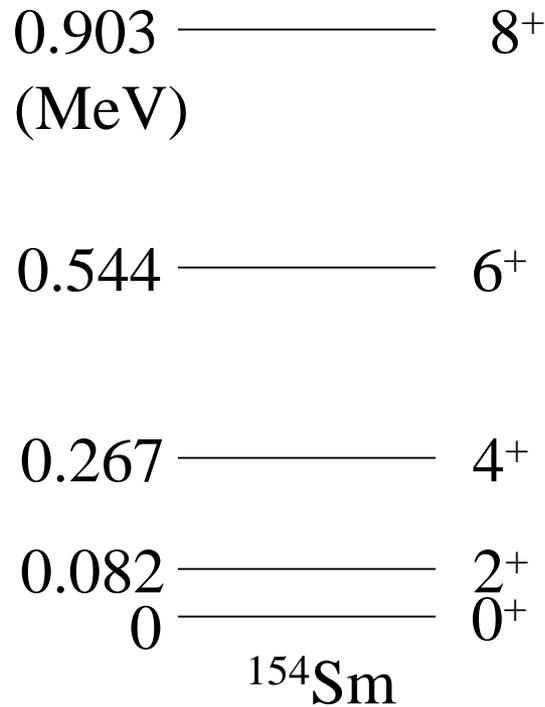
0 — 0^+
 ^{144}Sm

0 — 0^+
 ^{148}Sm

^{154}Sm

Effect of deformation on subbarrier fusion

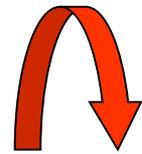
Excitation spectra of ^{154}Sm



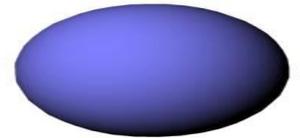
cf. Rotational energy of a rigid body
(Classical mechanics)

$$E = \frac{1}{2} \mathcal{J} \omega^2 = \frac{I^2}{2\mathcal{J}}$$

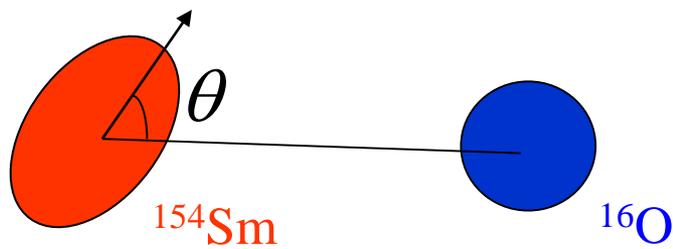
$$(I = \mathcal{J}\omega, \omega = \dot{\theta})$$



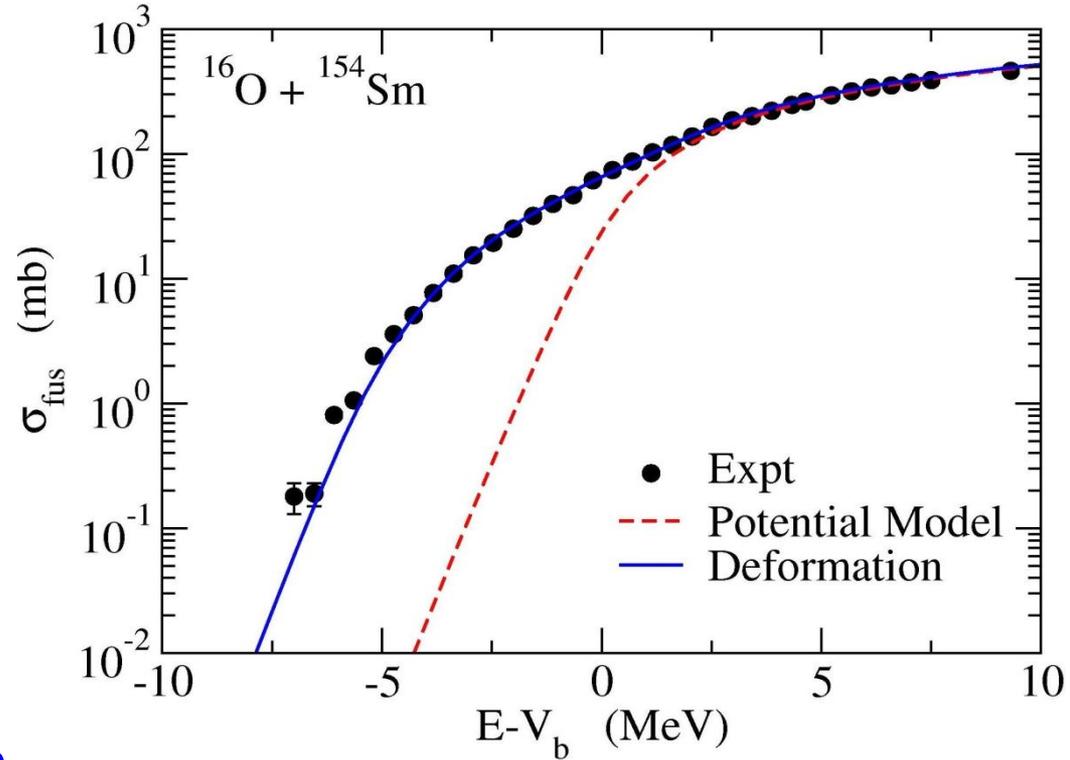
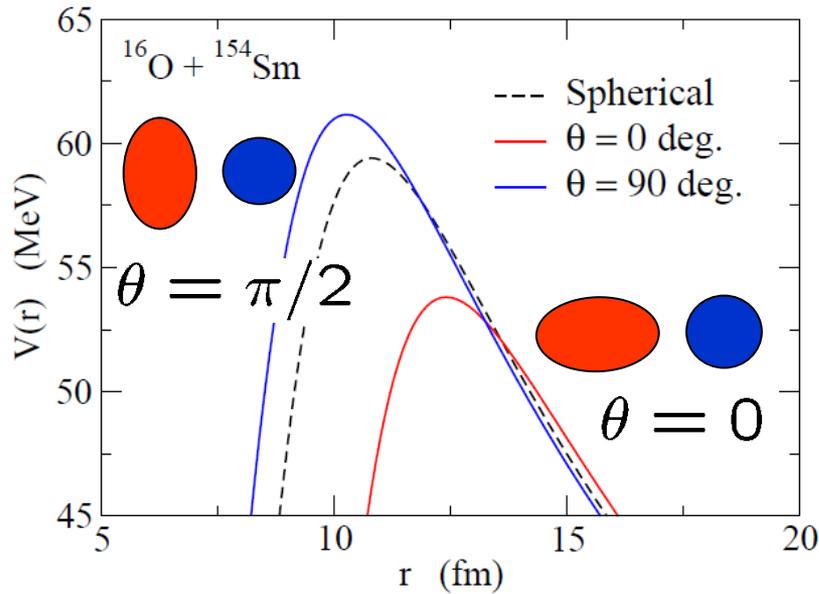
^{154}Sm is deformed



$$E_I \sim \frac{I(I+1)\hbar^2}{2\mathcal{J}}$$



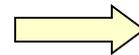
$$\sigma_{\text{fus}}(E) = \int_0^1 d(\cos \theta) \sigma_{\text{fus}}(E; \theta)$$



The barrier is lowered for $\theta=0$ because an attraction works from large distances.

The barrier increases for $\theta=\pi/2$ because the rel. distance has to get small for the attraction to work

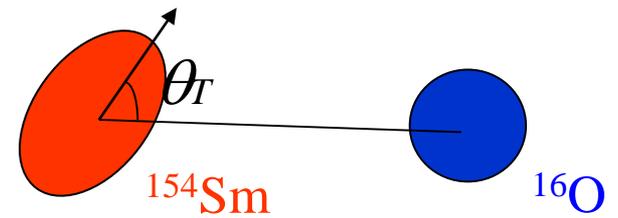
Def. Effect: enhances σ_{fus} by a factor of 10 ~ 100



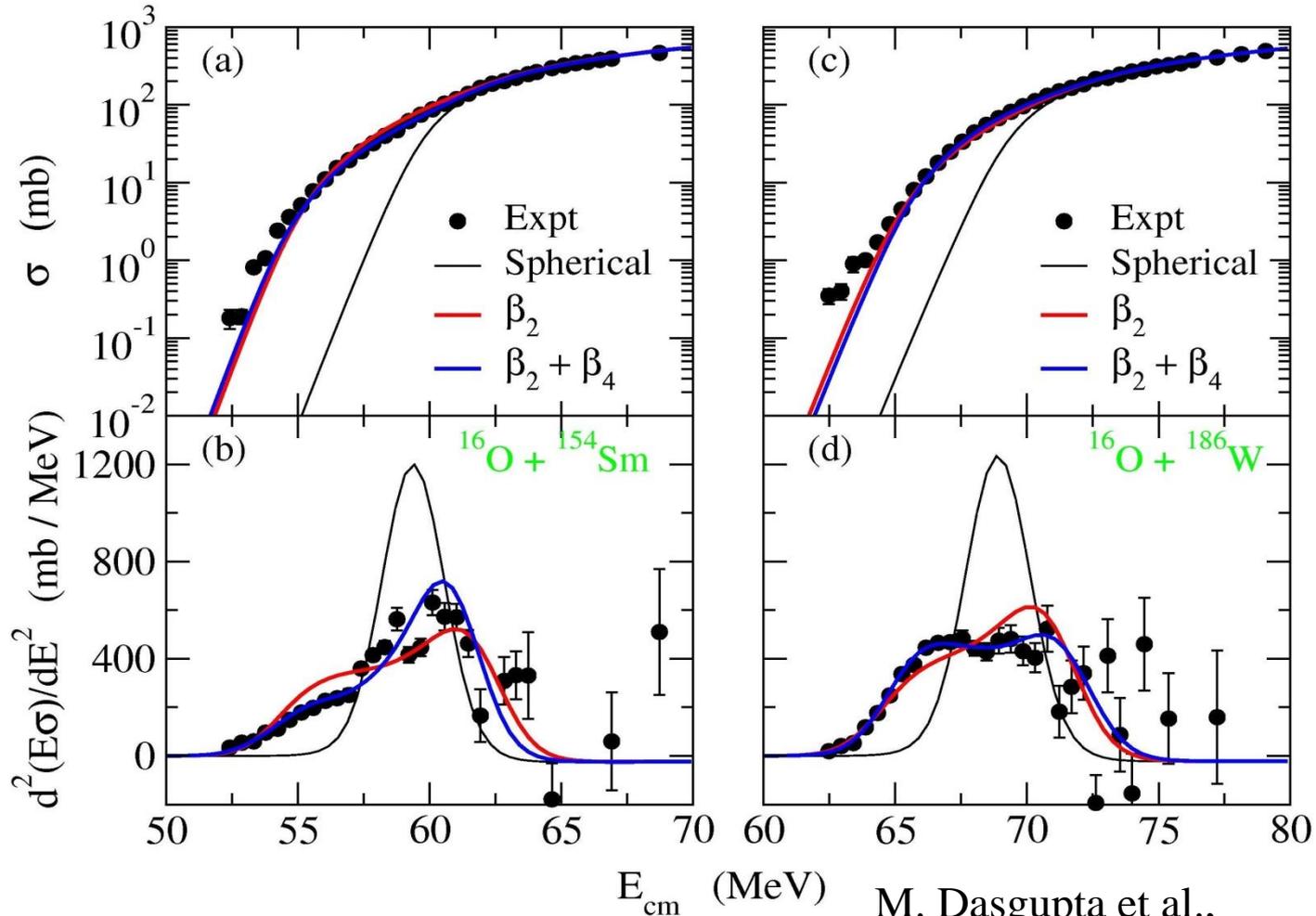
Fusion: interesting probe for nuclear structure

Fusion barrier distributions

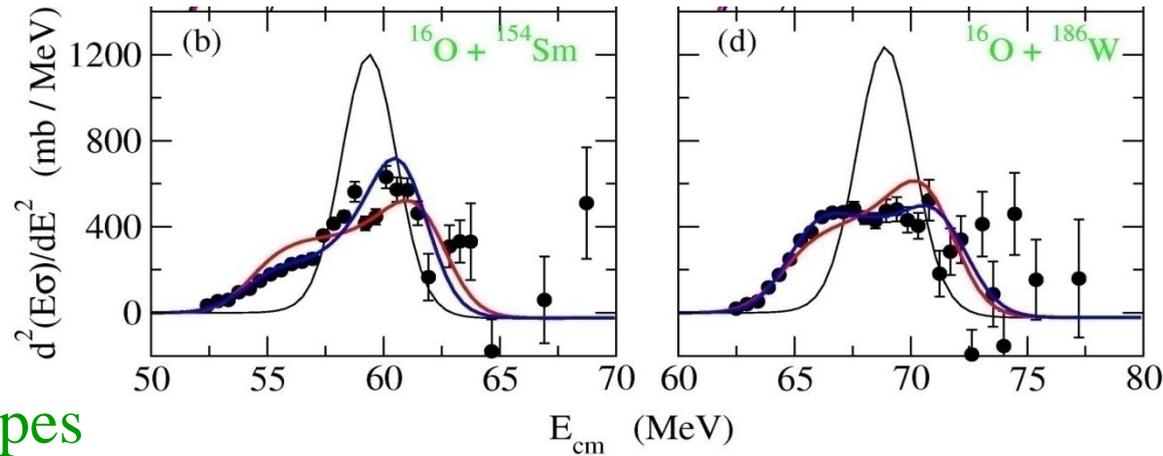
$$D_{\text{fus}}(E) = \frac{d^2(E\sigma_{\text{fus}})}{dE^2}$$



$$\sigma_{\text{fus}}(E) = \int_0^1 d(\cos \theta_T) \sigma_{\text{fus}}(E; \theta_T)$$



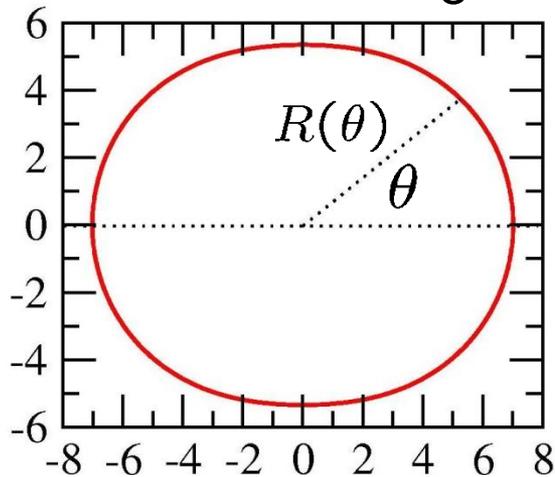
Investigate nuclear shape through barrier distribution



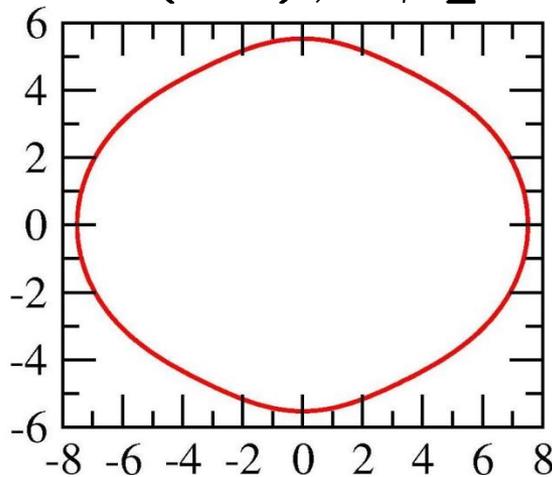
Nuclear shapes

$$R(\theta) = R_0(1 + \beta_2 Y_{20}(\theta) + \beta_4 Y_{40}(\theta) + \dots)$$

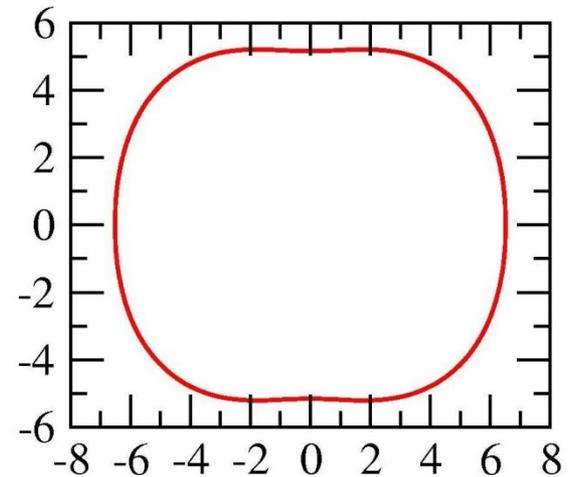
$$R_0 = 5.9 \text{ (fm)}, \quad \beta_2 = 0.3$$



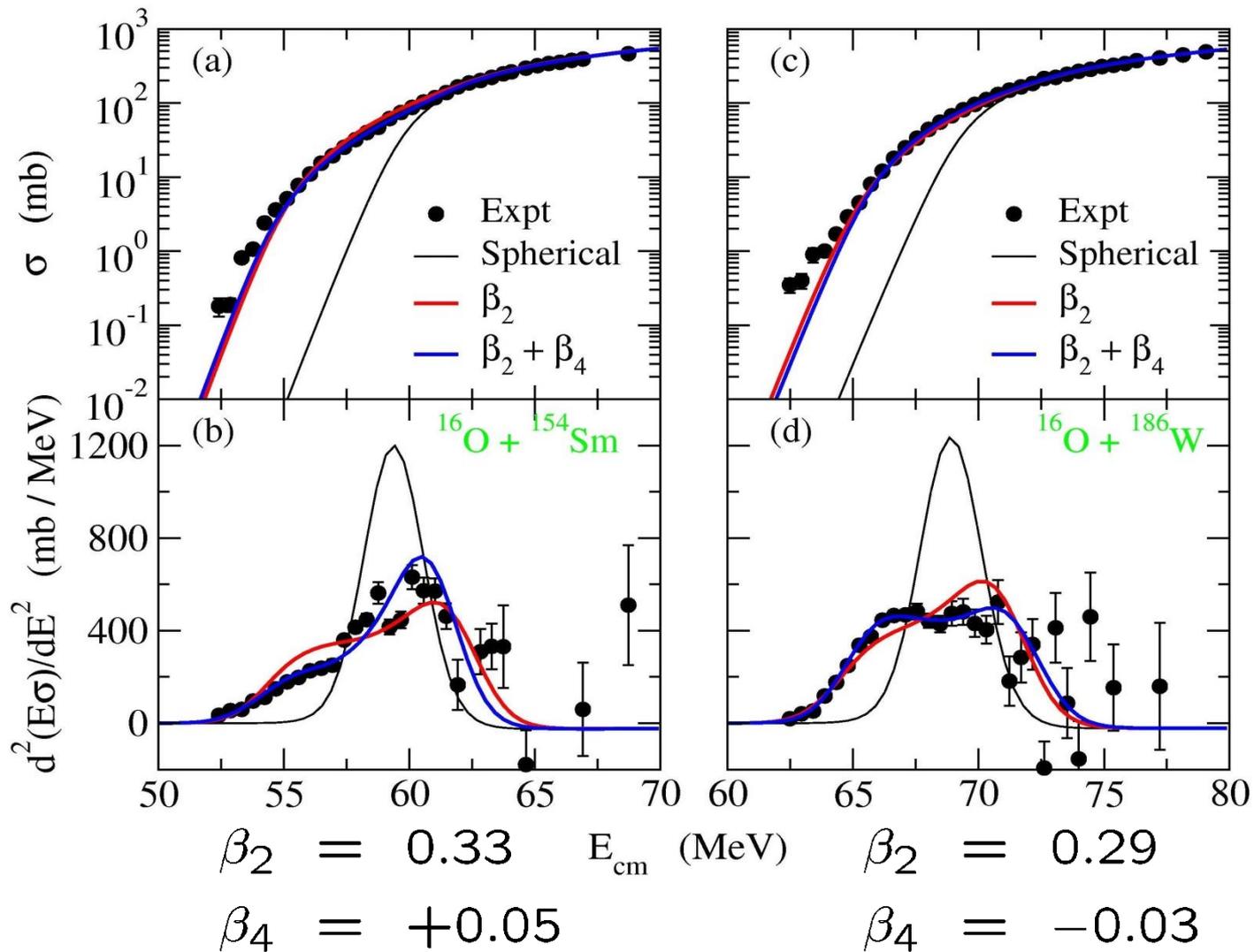
$$\beta_4 = 0$$



$$\beta_4 = 0.1$$



$$\beta_4 = -0.1$$

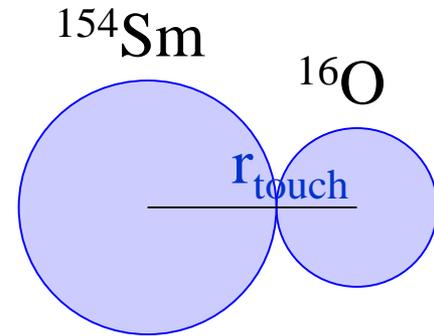
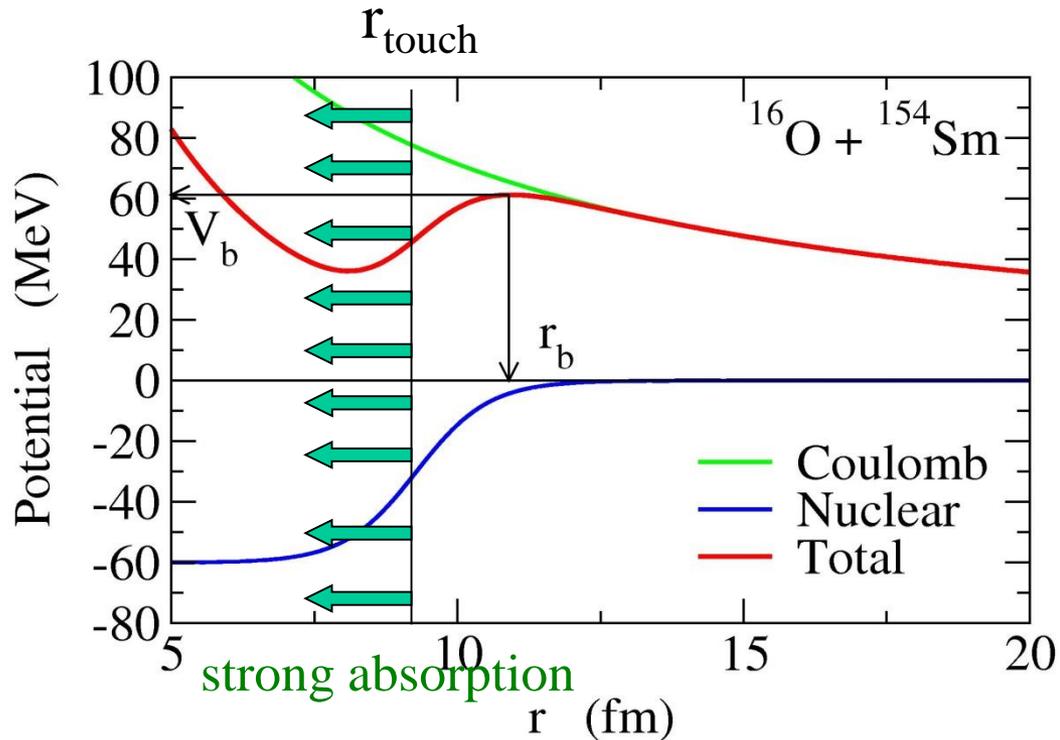


By taking the barrier distribution, one can very clearly see the difference due to β_4 !

➡ Fusion as a quantum tunneling microscope for nuclei

Heavy-ion subbarrier fusion reactions in the SHE region

➤ Fusion of medium-heavy systems:

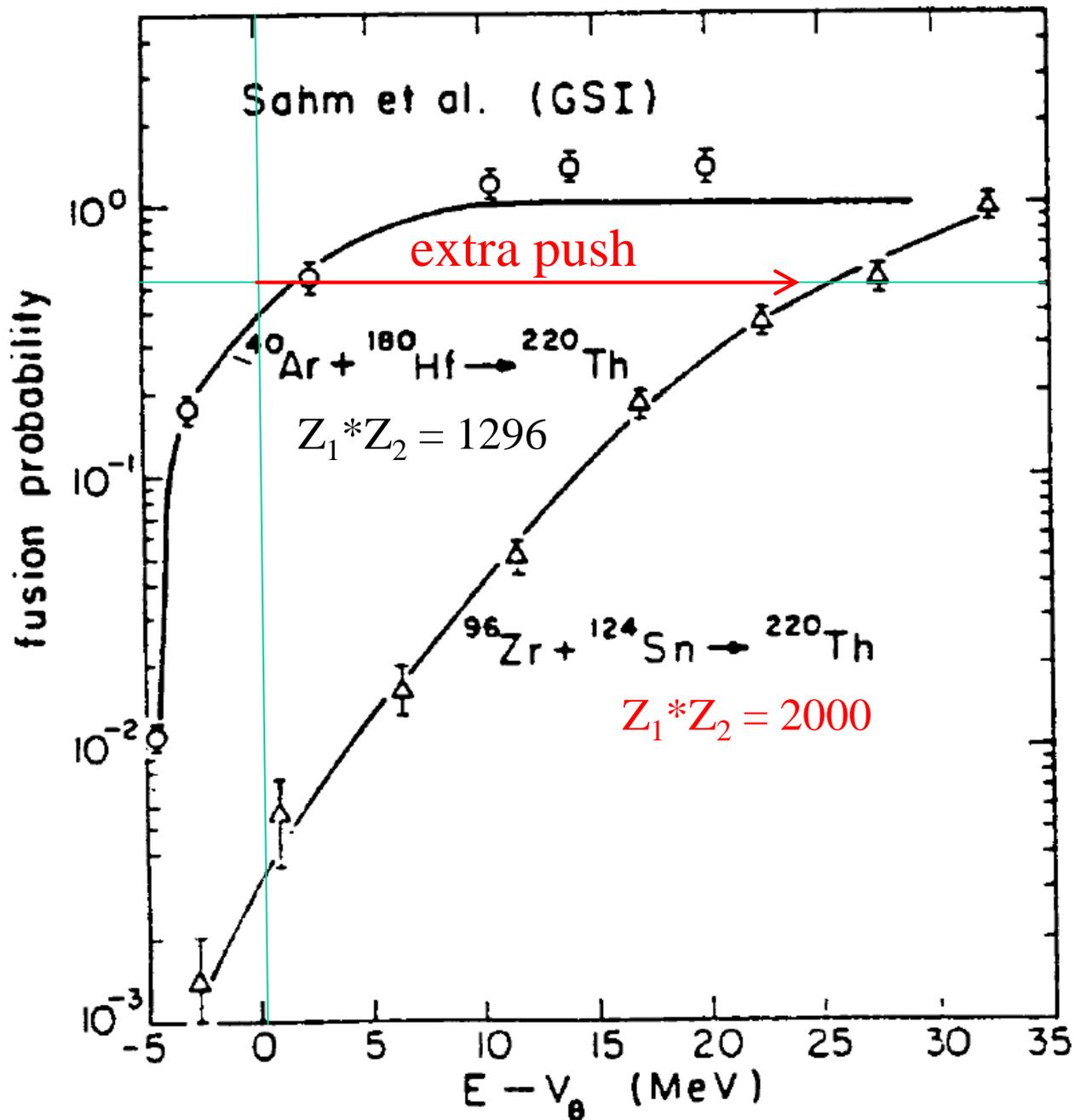


Compound nucleus:
automatically formed
once touched
(strong absorption)

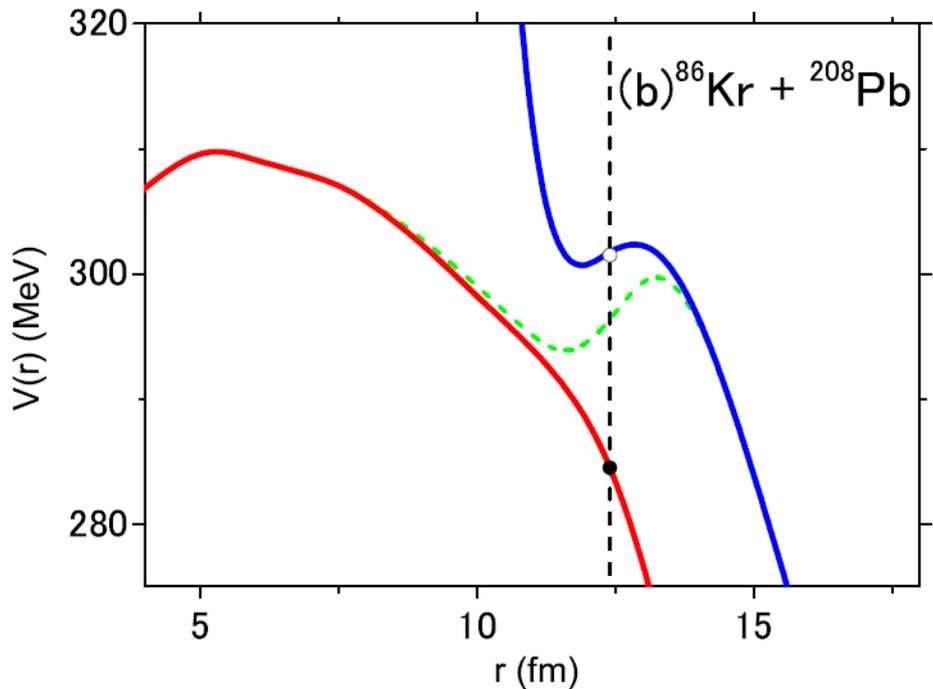
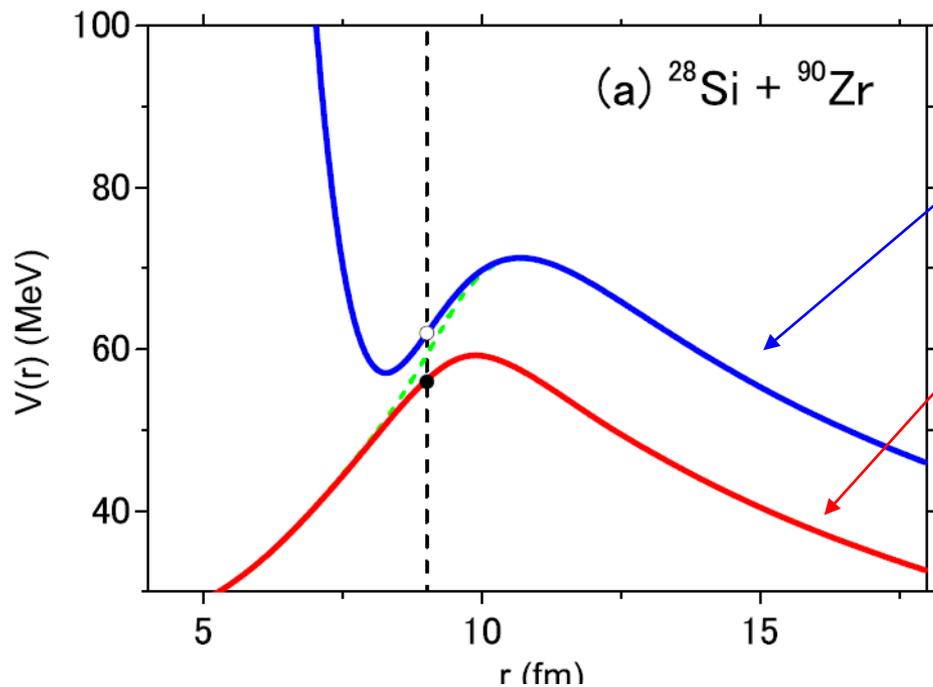
➤ Fusion of heavy and super-heavy systems

Large probability of re-separation (due to the strong Coulomb repulsion)

[This happens for $Z_1 * Z_2 > 1600 \sim 1800$.]



C.-C. Sahm et al.,
 Z. Phys. A319('84)113



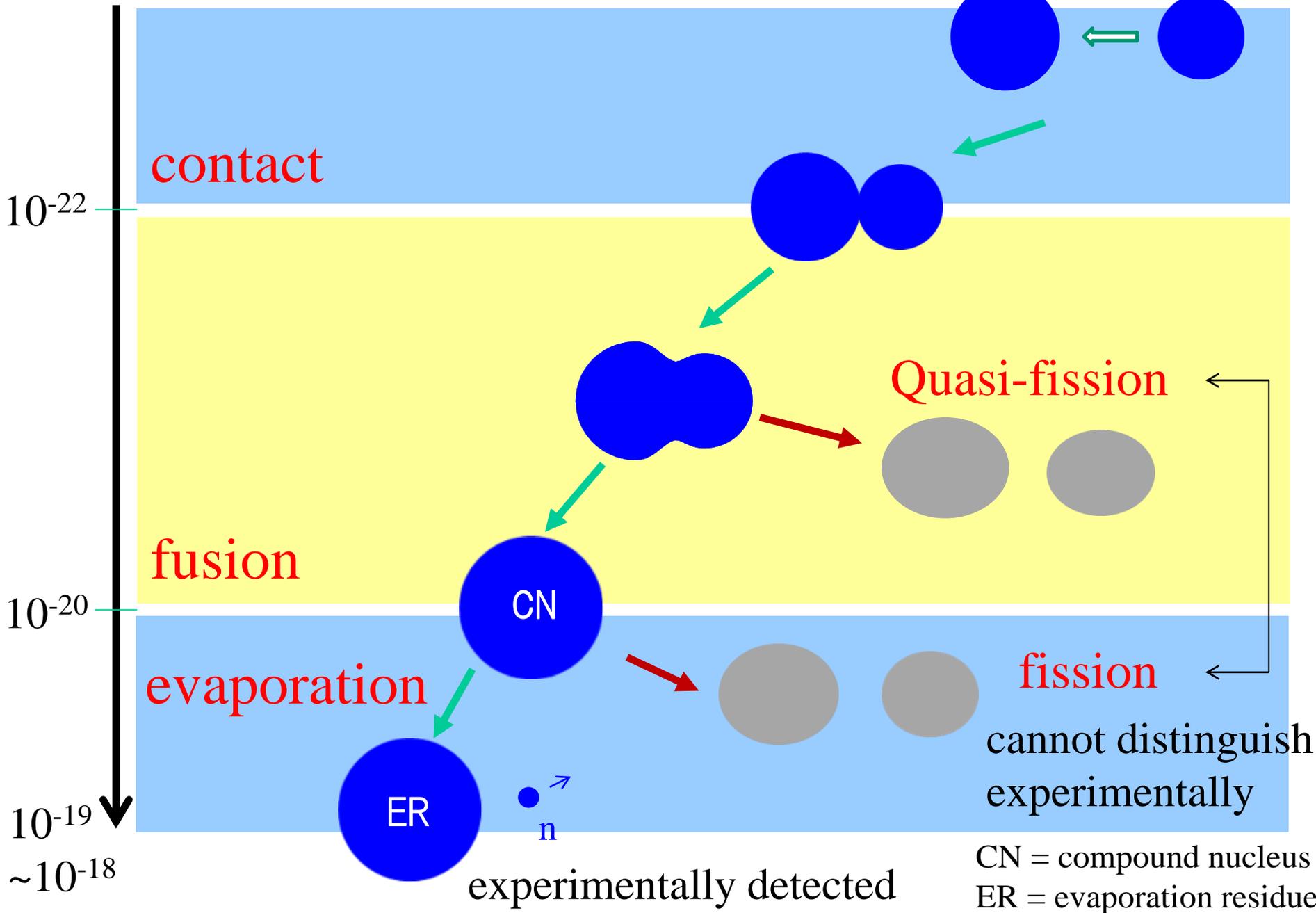
2-body potential before touching

1-body potential after touching

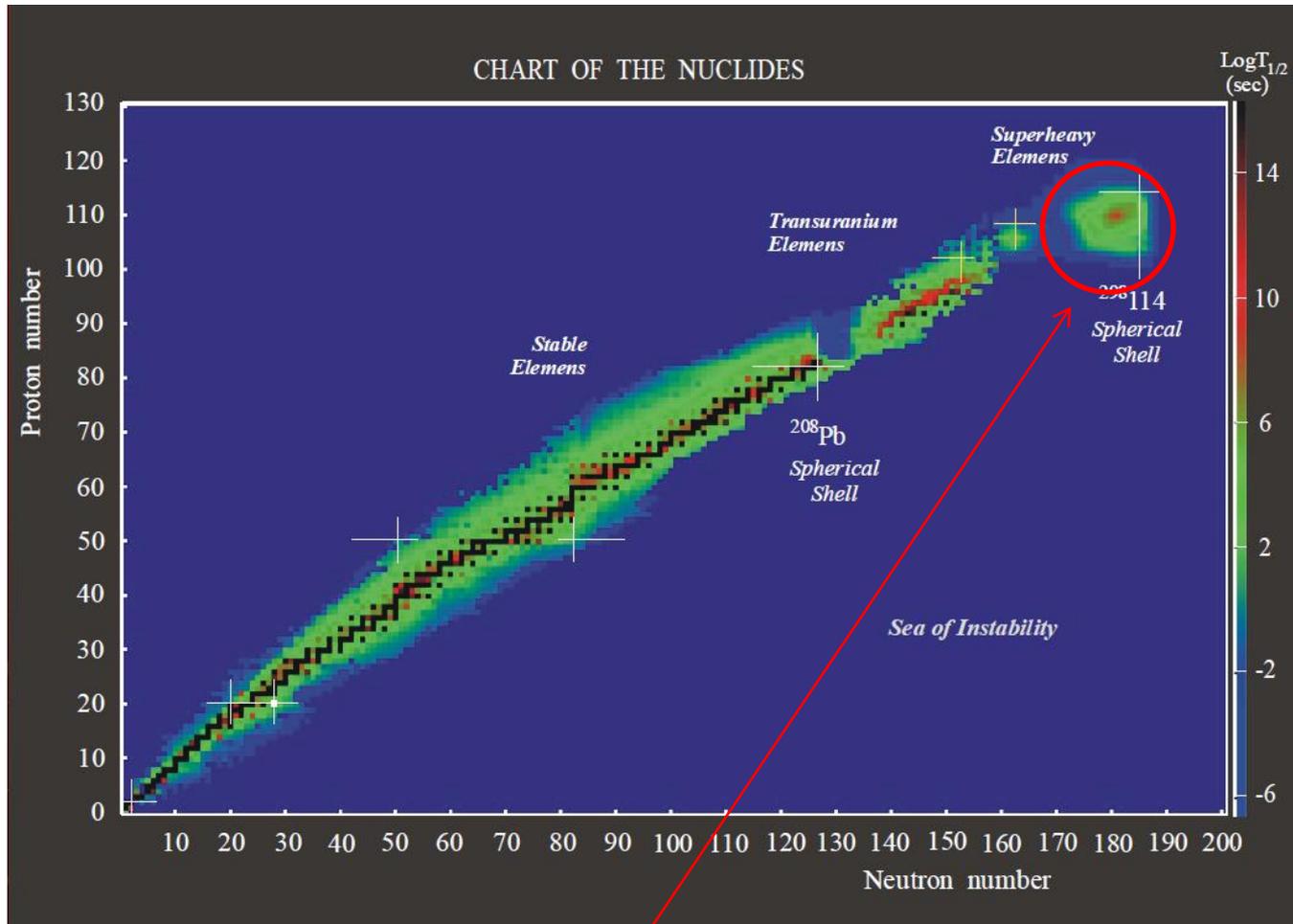
← The red potential has to be overcome even if the blue potential has been overcome.

↓
Re-separation if failed
(quasi-fission)

typical time-scale (sec.)



Heavy-ion fusion for SHE



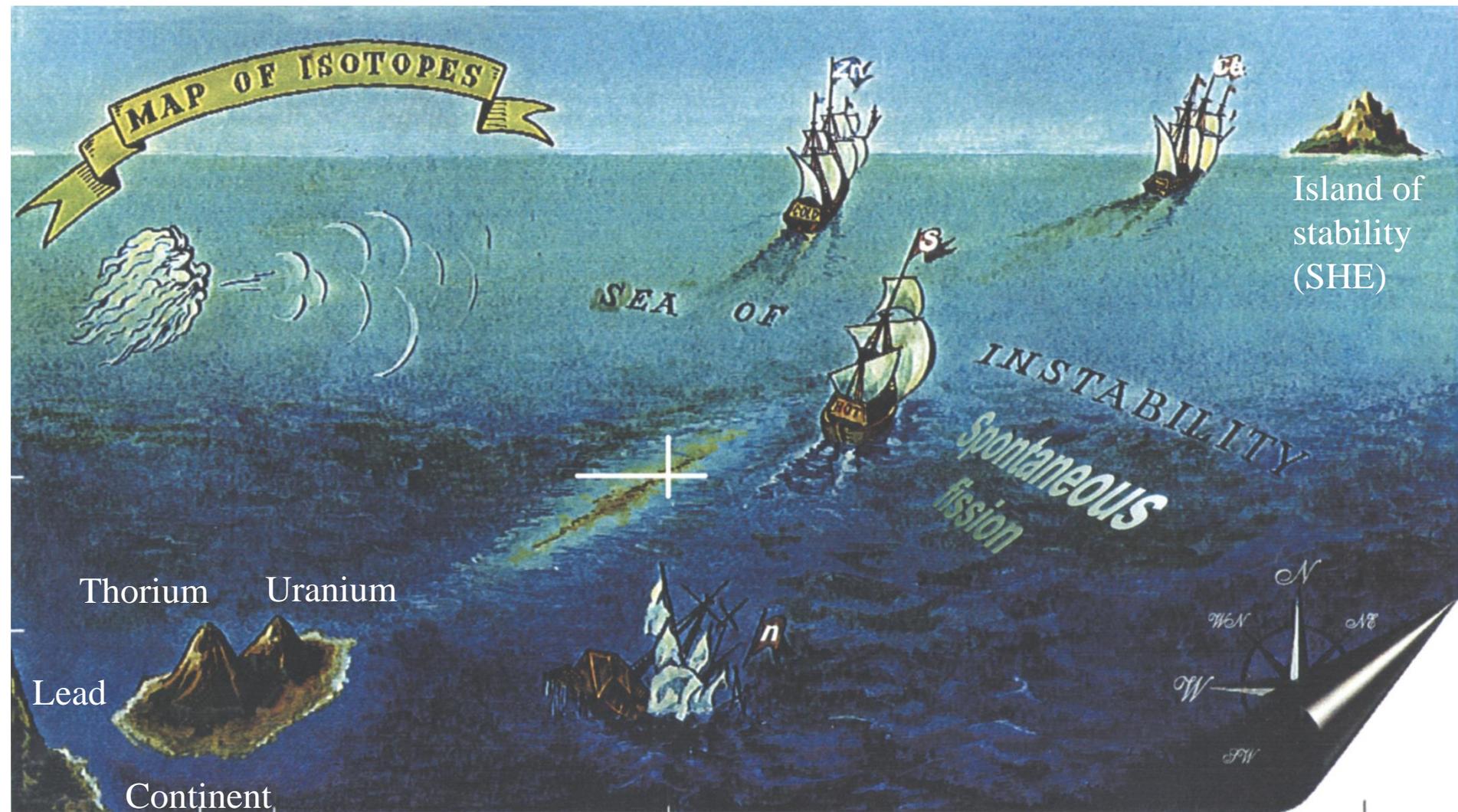
island of stability around $Z=114, N=184$

Yuri Oganessian

W.D. Myers and W.J. Swiatecki (1966), A. Sobiczewski et al. (1966)

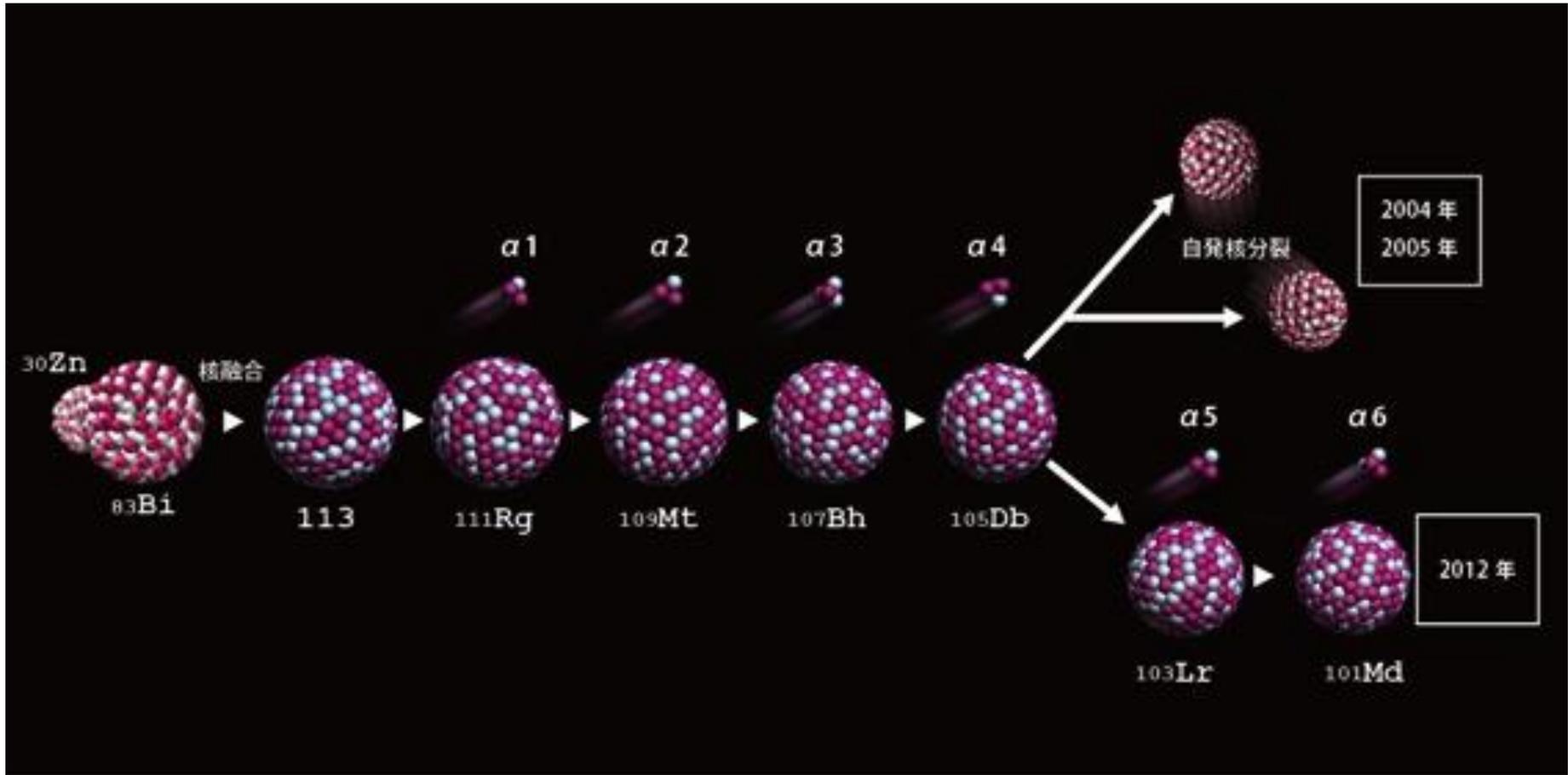
→ modern calculations: $Z=114, 120, \text{ or } 126, N=184$

e.g., H. Koura et al. (2005)



Yuri Oganessian

Element 113 (RIKEN, K. Morita et al.)

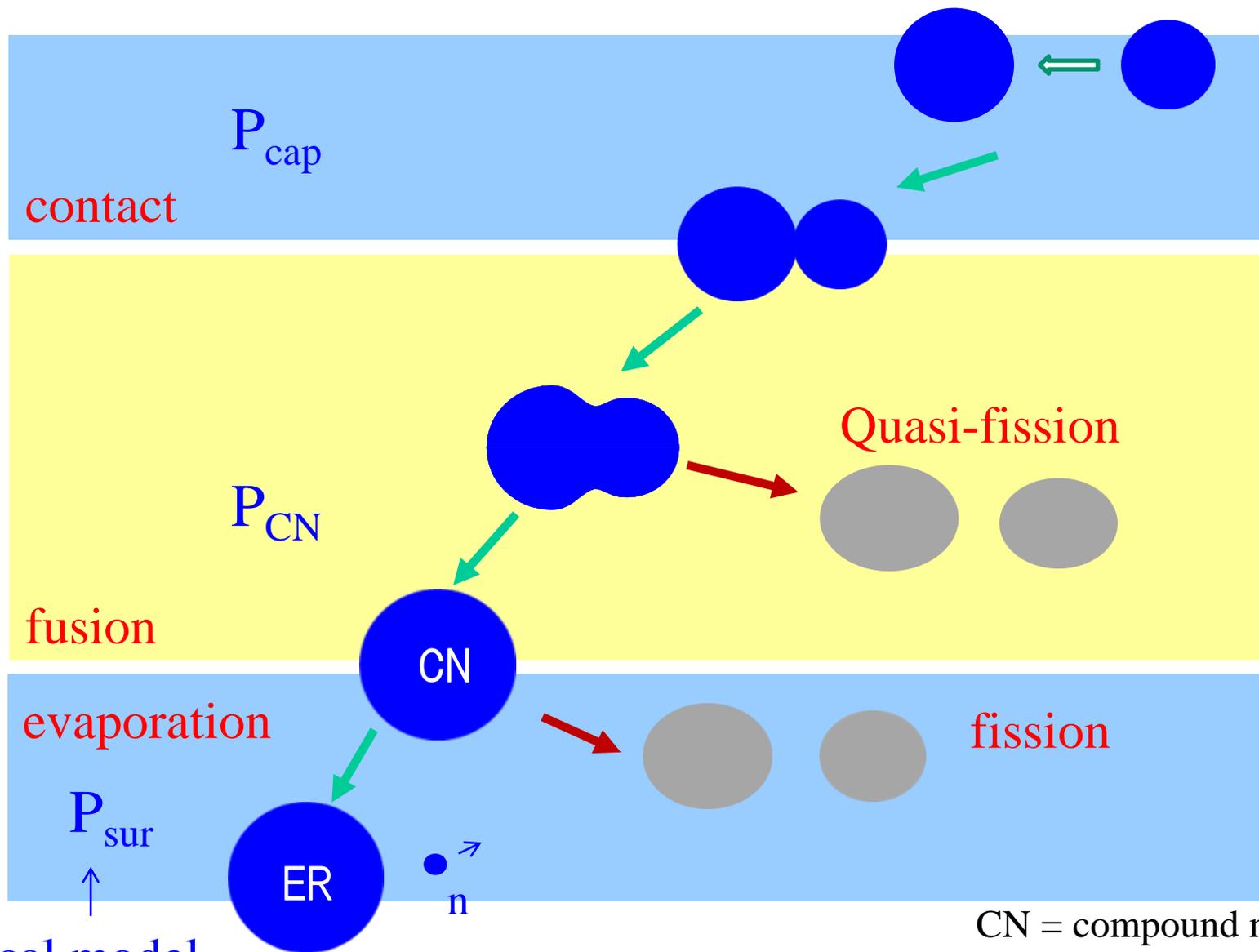


K. Morita et al., J. Phys. Soc. Jpn. 81('12)103201

$$\sigma_{\text{ER}} = 22_{-13}^{+20} \text{ fb} \quad \text{only 3 events for 553 days experiment}$$

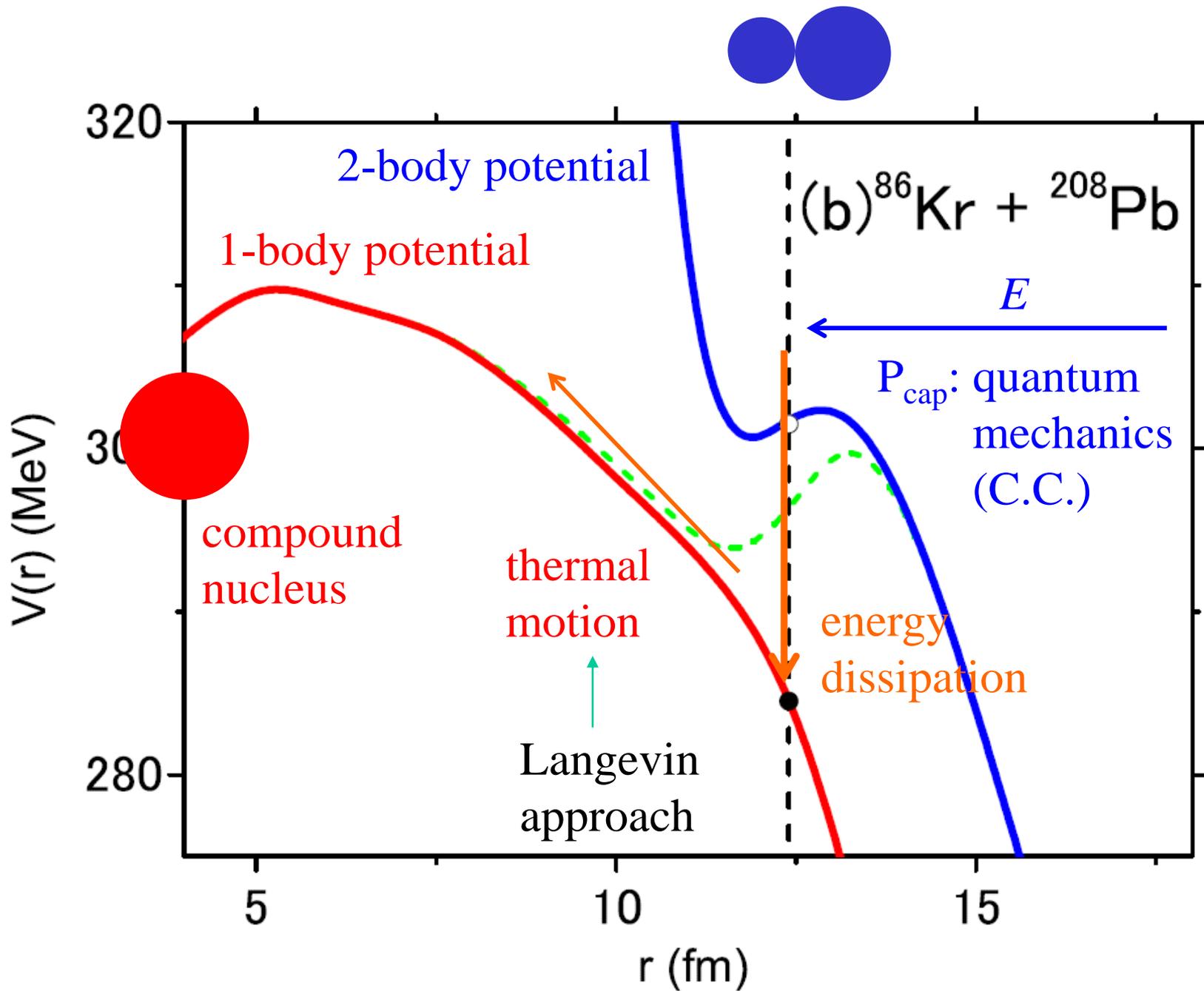
Theoretical treatment

$$P_{ER} = P_{cap} \cdot P_{CN} \cdot P_{sur}$$



statistical model

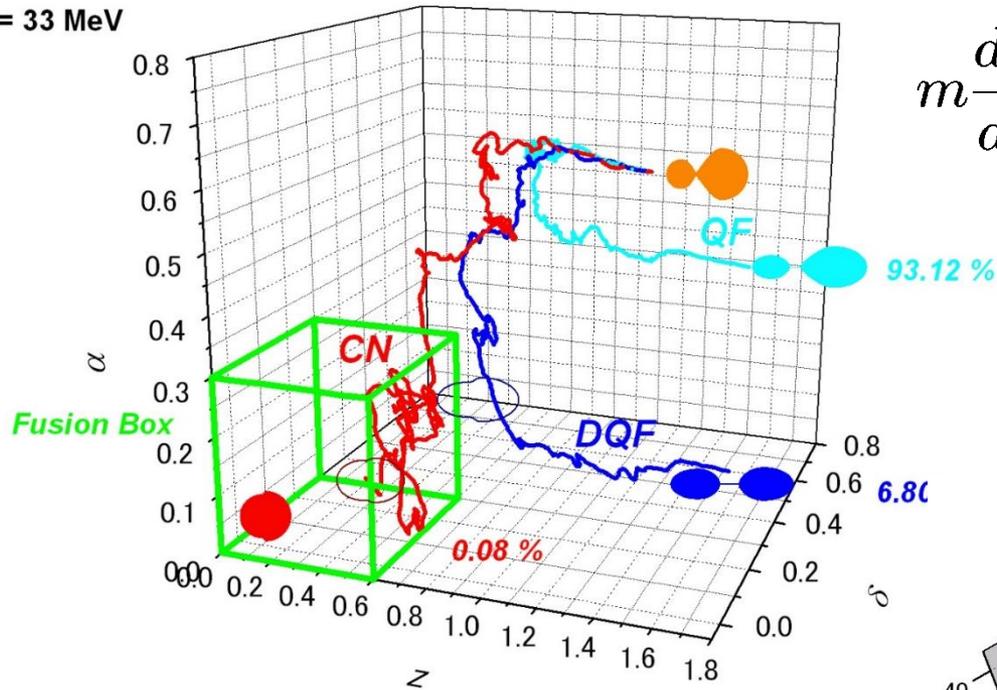
CN = compound nucleus
ER = evaporation residue



Theory: Lagenvin approach

$^{48}\text{Ca} + ^{244}\text{Pu} \rightarrow ^{292}\text{114}$

$E^* = 33 \text{ MeV}$

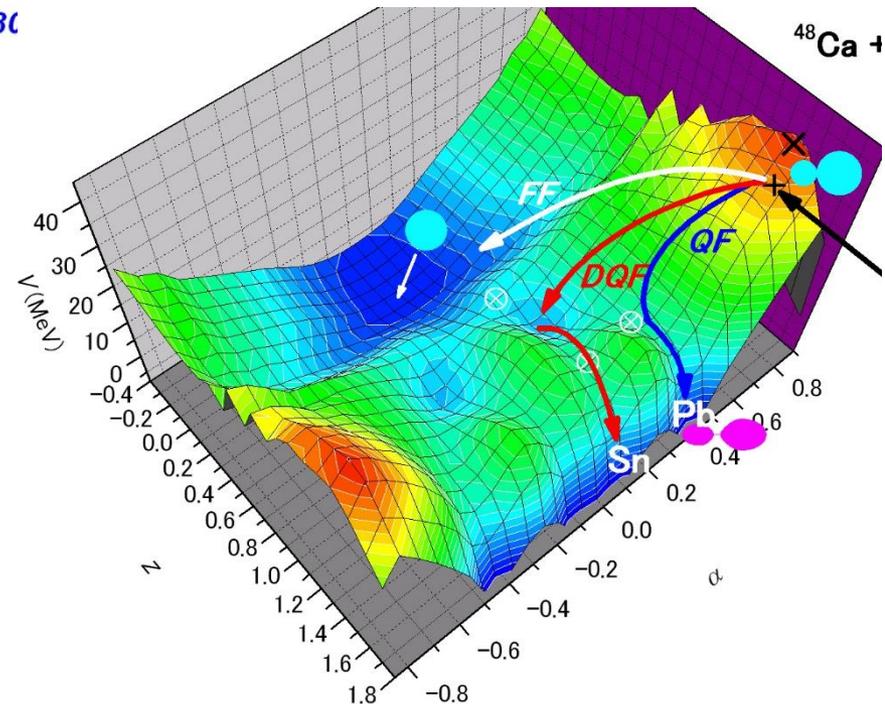


- q :
- internuclear separation (z),
 - deformation (δ),
 - asymmetry of the two fragments (α)

multi-dimensional extension of:

$$m \frac{d^2 q}{dt^2} = - \frac{dV(q)}{dq} - \gamma \frac{dq}{dt} + R(t)$$

γ : friction coefficient
 $R(t)$: random force



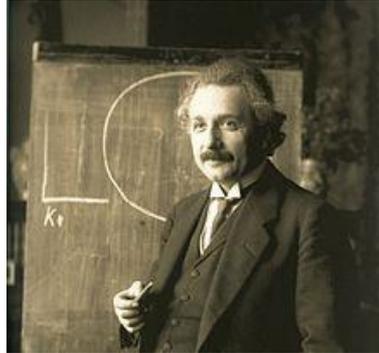
Chemistry of superheavy elements

| Group → ↓ Period | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | 10 | 11 | 12 | 13 | 14 | 15 | 16 | 17 | 18 |
|---------------------|----------|----------|----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|------------|-----------|------------|-----------|------------|------------|
| 1 | 1 H | | | | | | | | | | | | | | | | | 2 He |
| 2 | 3 Li | 4 Be | | | | | | | | | | | 5 B | 6 C | 7 N | 8 O | 9 F | 10 Ne |
| 3 | 11 Na | 12 Mg | | | | | | | | | | | 13 Al | 14 Si | 15 P | 16 S | 17 Cl | 18 Ar |
| 4 | 19 K | 20 Ca | 21 Sc | 22 Ti | 23 V | 24 Cr | 25 Mn | 26 Fe | 27 Co | 28 Ni | 29 Cu | 30 Zn | 31 Ga | 32 Ge | 33 As | 34 Se | 35 Br | 36 Kr |
| 5 | 37 Rb | 38 Sr | 39 Y | 40 Zr | 41 Nb | 42 Mo | 43 Tc | 44 Ru | 45 Rh | 46 Pd | 47 Ag | 48 Cd | 49 In | 50 Sn | 51 Sb | 52 Te | 53 I | 54 Xe |
| 6 | 55 Cs | 56 Ba | | 72 Hf | 73 Ta | 74 W | 75 Re | 76 Os | 77 Ir | 78 Pt | 79 Au | 80 Hg | 81 Tl | 82 Pb | 83 Bi | 84 Po | 85 At | 86 Rn |
| 7 | 87 Fr | 88 Ra | | 104 Rf | 105 Db | 106 Sq | 107 Bh | 108 Hs | 109 Mt | 110 Ds | 111 Rg | 112 Cn | 113 Uut | 114 Fl | 115 Uup | 116 Lv | 117 Uus | 118 Uuo |
| Lanthanides | 57 La | 58 Ce | 59 Pr | 60 Nd | 61 Pm | 62 Sm | 63 Eu | 64 Gd | 65 Tb | 66 Dy | 67 Ho | 68 Er | 69 Tm | 70 Yb | 71 Lu | | | |
| Actinides | 89 Ac | 90 Th | 91 Pa | 92 U | 93 Np | 94 Pu | 95 Am | 96 Cm | 97 Bk | 98 Cf | 99 Es | 100 Fm | 101 Md | 102 No | 103 Lr | | | |

- Are they here in the periodic table?
- That is, does e.g., Lv show the same chemical properties as O, S, Se, Te, and Po?

relativistic effect : important for large Z

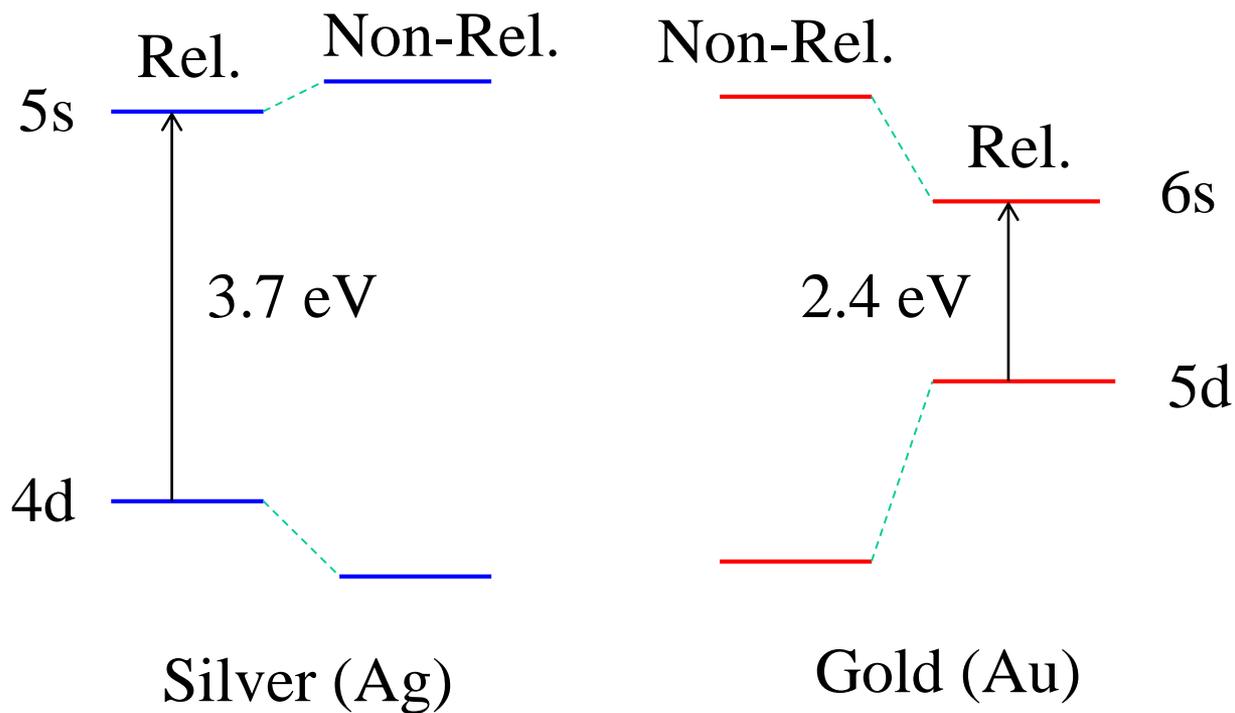
$$E = mc^2$$



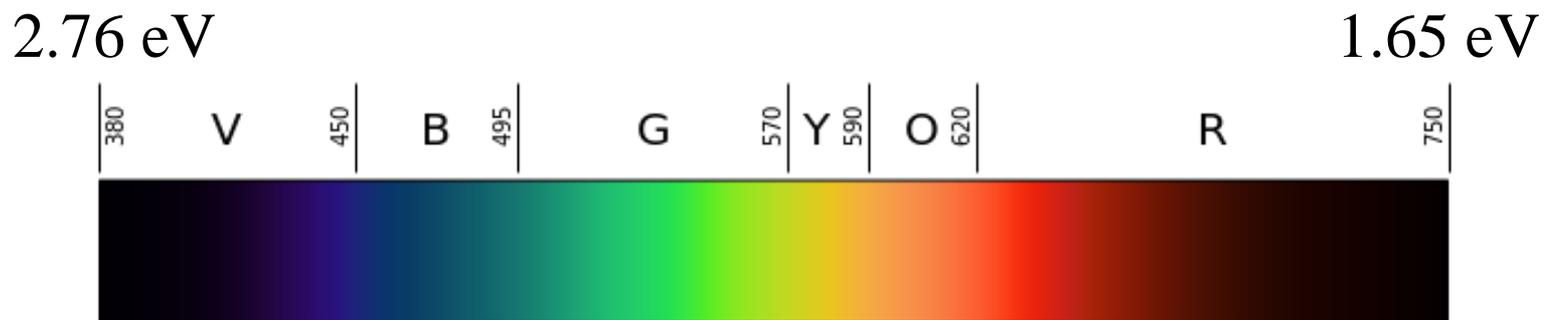
Solution of the Dirac equation (relativistic quantum mechanics)
for a hydrogen-like atom:

$$E_{1S} = mc^2 \sqrt{1 - (Z\alpha)^2} \sim mc^2 \left(1 - \frac{(Z\alpha)^2}{2} - \underbrace{\frac{(Z\alpha)^4}{8} + \dots}_{\text{relativistic effect}} \right)$$

relativistic effect

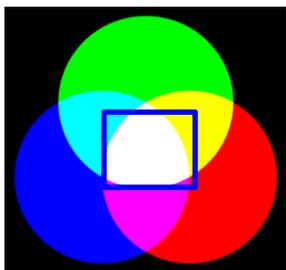


cf. visible spectrum

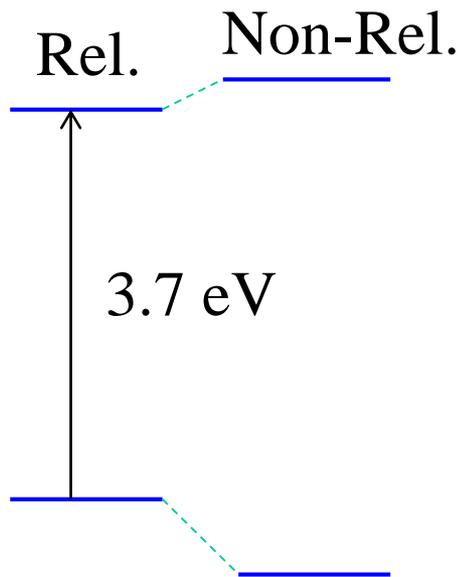


↑
3.7 eV

↑
2.4 eV

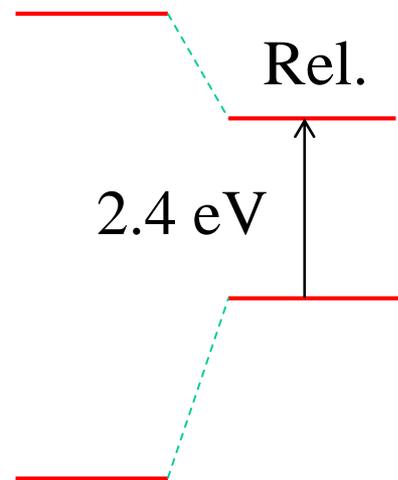


no color
absorbed

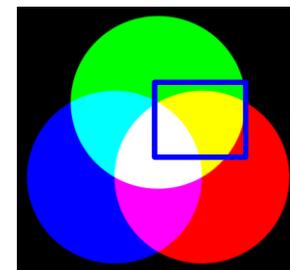


Silver (Ag)

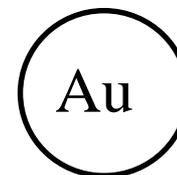
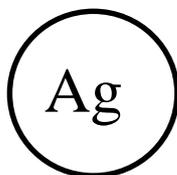
Non-Rel.



Gold (Au)



blue: absorbed



Chemistry of superheavy elements

| Group → ↓ Period | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | 10 | 11 | 12 | 13 | 14 | 15 | 16 | 17 | 18 | |
|---------------------|----------|----------|----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|------------|-----------|------------|-----------|------------|------------|---------|
| 1 | 1 H | | | | | | | | | | | | | | | | | | 2 He |
| 2 | 3 Li | 4 Be | | | | | | | | | | | 5 B | 6 C | 7 N | 8 O | 9 F | 10 Ne | |
| 3 | 11 Na | 12 Mg | | | | | | | | | | | 13 Al | 14 Si | 15 P | 16 S | 17 Cl | 18 Ar | |
| 4 | 19 K | 20 Ca | 21 Sc | 22 Ti | 23 V | 24 Cr | 25 Mn | 26 Fe | 27 Co | 28 Ni | 29 Cu | 30 Zn | 31 Ga | 32 Ge | 33 As | 34 Se | 35 Br | 36 Kr | |
| 5 | 37 Rb | 38 Sr | 39 Y | 40 Zr | 41 Nb | 42 Mo | 43 Tc | 44 Ru | 45 Rh | 46 Pd | 47 Ag | 48 Cd | 49 In | 50 Sn | 51 Sb | 52 Te | 53 I | 54 Xe | |
| 6 | 55 Cs | 56 Ba | | 72 Hf | 73 Ta | 74 W | 75 Re | 76 Os | 77 Ir | 78 Pt | 79 Au | 80 Hg | 81 Tl | 82 Pb | 83 Bi | 84 Po | 85 At | 86 Rn | |
| 7 | 87 Fr | 88 Ra | | 104 Rf | 105 Db | 106 Sq | 107 Bh | 108 Hs | 109 Mt | 110 Ds | 111 Rg | 112 Cn | 113 Uut | 114 Fl | 115 Uup | 116 Lv | 117 Uus | 118 Uuo | |
| Lanthanides | | | 57 La | 58 Ce | 59 Pr | 60 Nd | 61 Pm | 62 Sm | 63 Eu | 64 Gd | 65 Tb | 66 Dy | 67 Ho | 68 Er | 69 Tm | 70 Yb | 71 Lu | | |
| Actinides | | | 89 Ac | 90 Th | 91 Pa | 92 U | 93 Np | 94 Pu | 95 Am | 96 Cm | 97 Bk | 98 Cf | 99 Es | 100 Fm | 101 Md | 102 No | 103 Lr | | |

How do the relativistic effects alter the periodic table for SHE?

→ a big open question