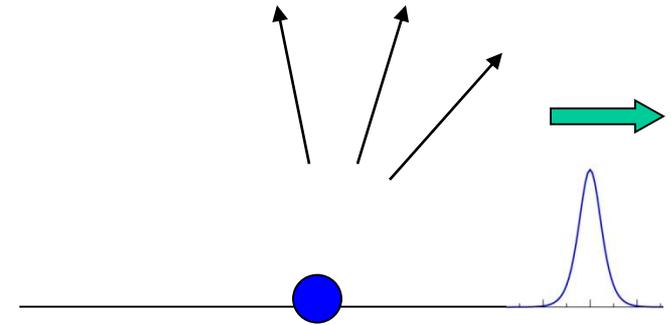
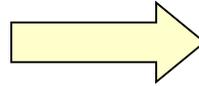


# Absorption cross sections

## Reaction processes

- Elastic scatt.
- Inelastic scatt.
- Transfer reaction
- Compound nucleus formation (fusion)



Loss of incident flux  
(absorption)

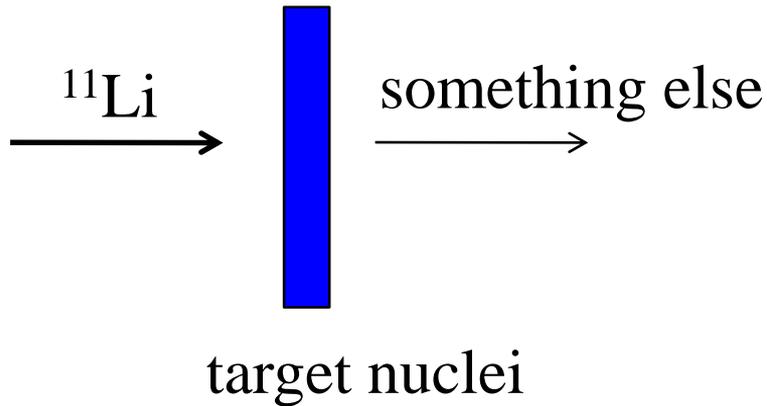
## reaction cross sections

total scattering cross section - elastic cross section

$$\sigma_R = \sigma_{\text{tot}} - \sigma_{\text{el}}$$

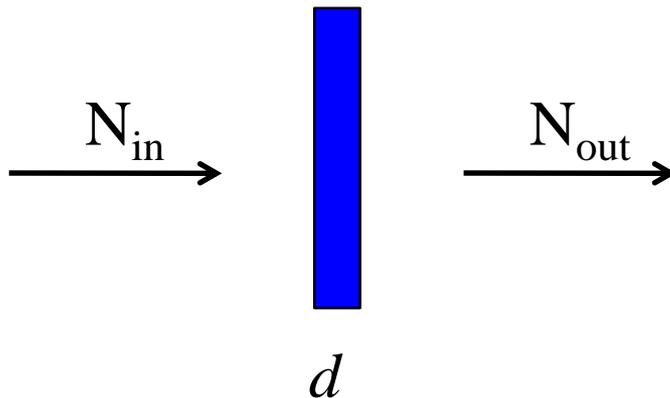
- fusion
- inelastic
- transfer

# Interaction cross sections and halo nuclei



interaction cross section  $\sigma_I$   
= cross section for the change  
of Z a/o N in the incident nucleus

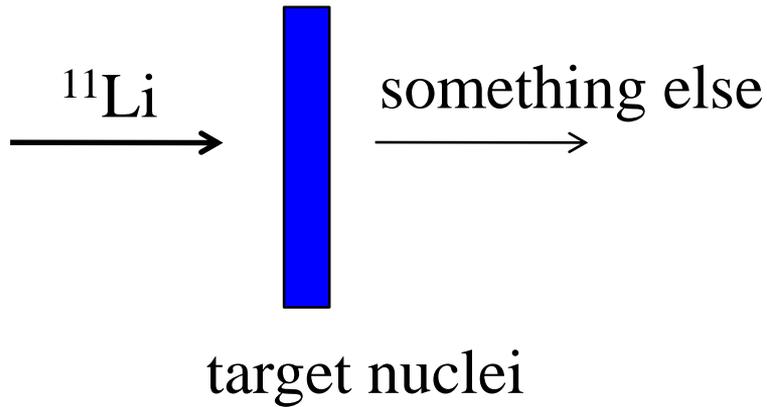
transmission method



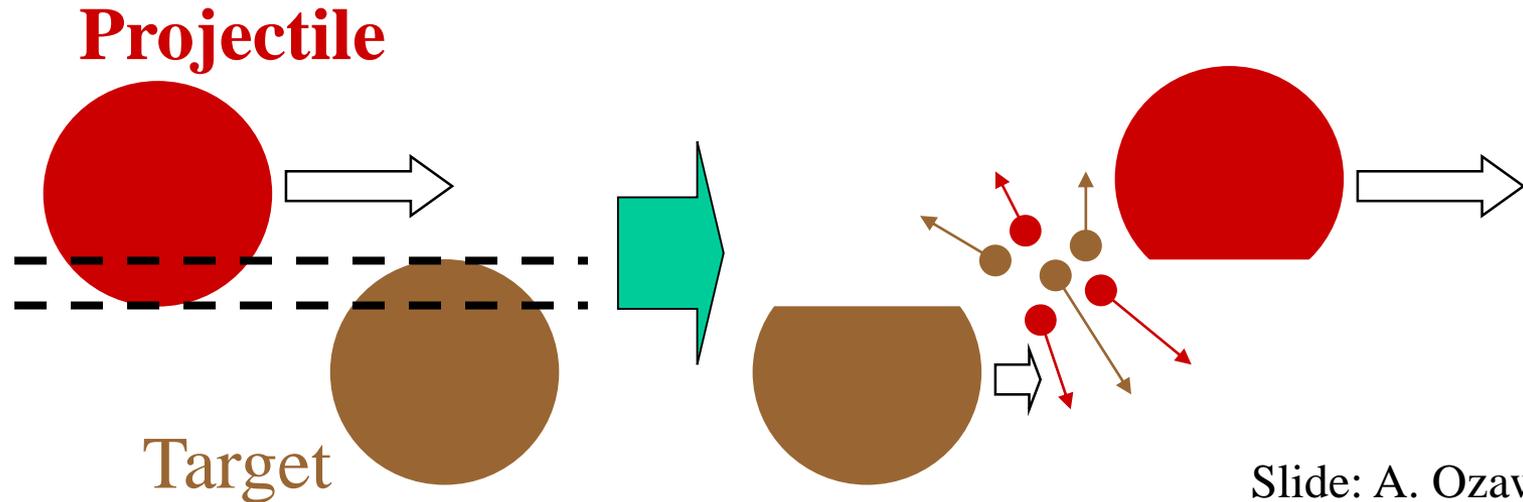
$$\sigma_R = -\frac{1}{t} \ln \left( \frac{N_{\text{out}}}{N_{\text{in}}} \right)$$

$$t = \rho_T \cdot d \cdot \epsilon$$

# Interaction cross sections and halo nuclei

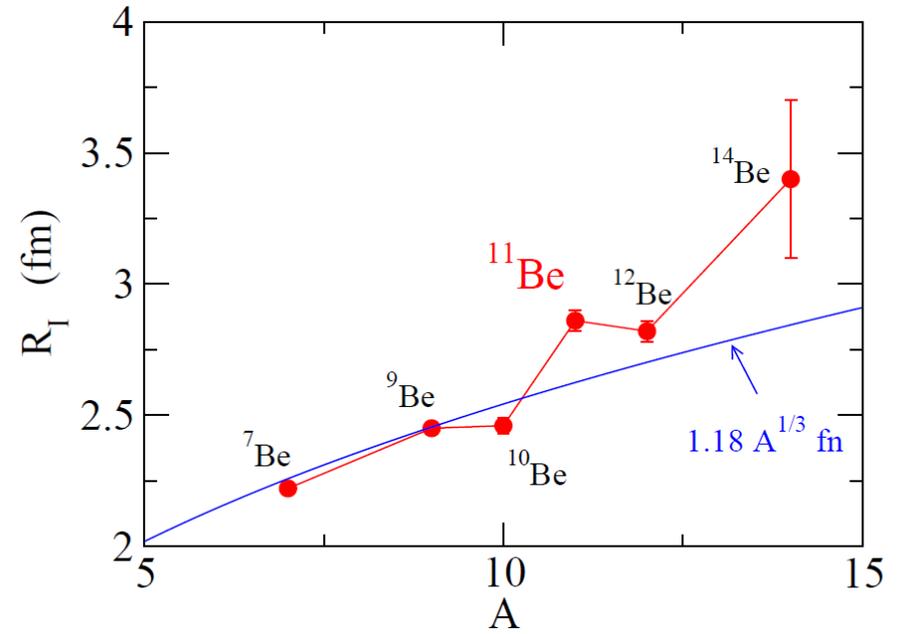
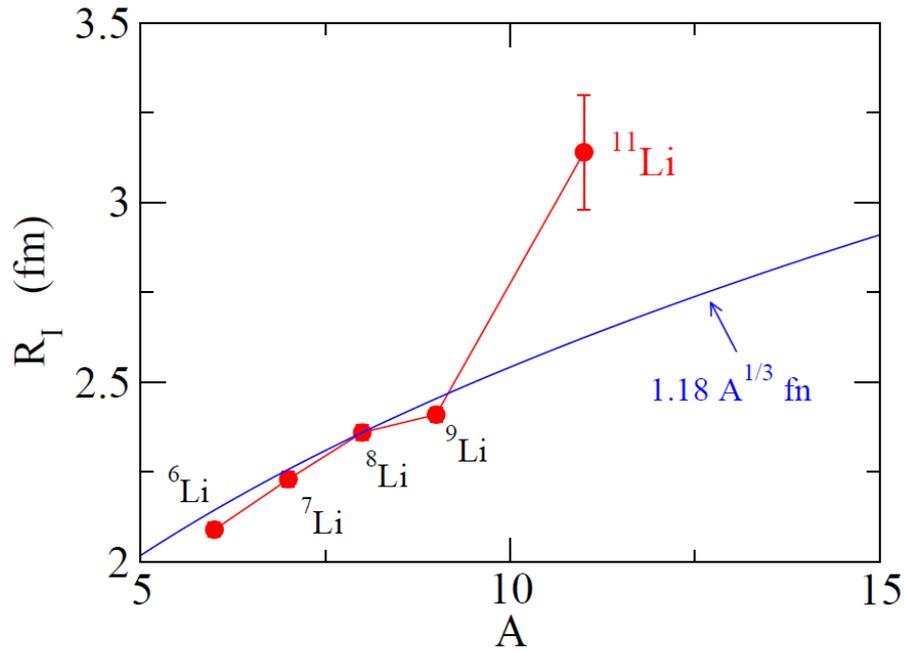


interaction cross section  $\sigma_I$   
= cross section for the change  
of Z a/o N in the incident nucleus



$$\sigma_I \sim \pi [R_I(P) + R_I(T)]^2 \longrightarrow R_I(P)$$

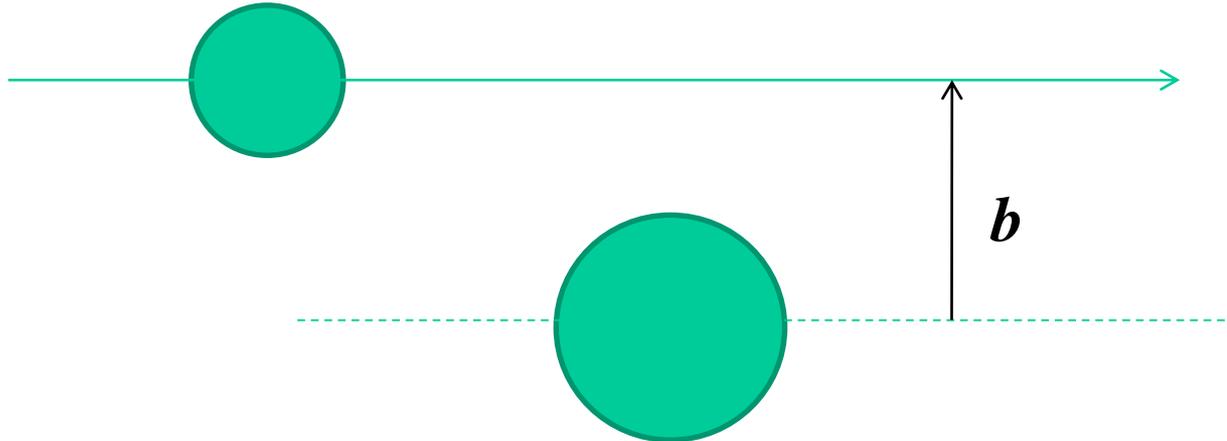
# Discovery of halo nuclei



I. Tanihata, T. Kobayashi, O. Hashimoto et al., PRL55('85)2676; PLB206('88)592



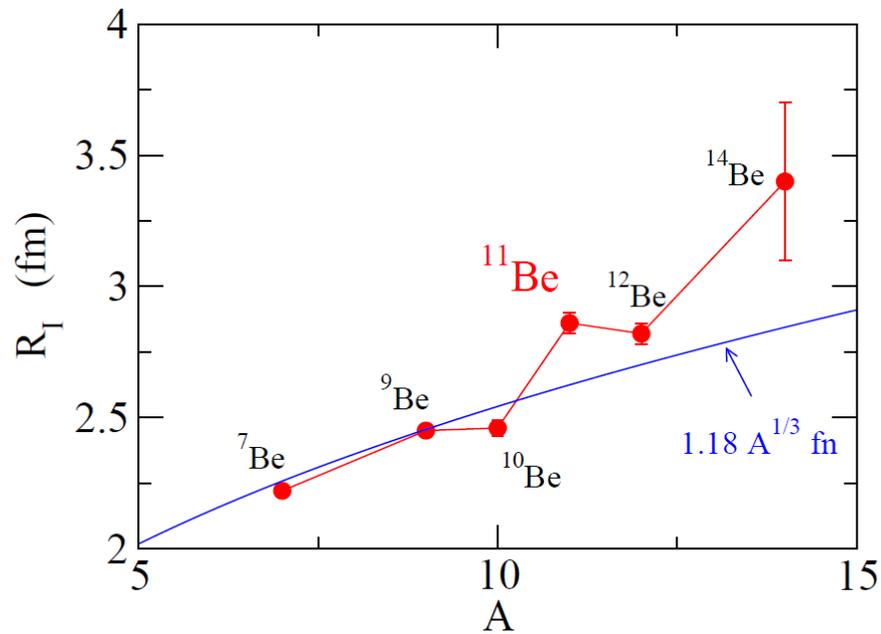
## Reaction cross sections



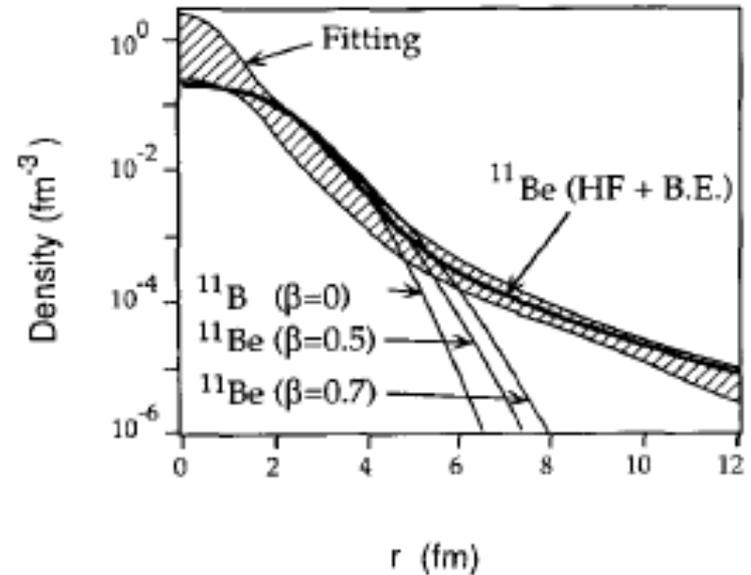
**Glauber theory** (optical limit approximation : OLA)

$$\sigma_R \sim 2\pi \int_0^\infty b db \left[ 1 - \exp \left( -\sigma_{NN} \int d^2s \rho_P^{(z)}(\mathbf{s}) \rho_T^{(z)}(\mathbf{s} - \mathbf{b}) \right) \right]$$

- straight-line trajectory (high energy scattering)
- adiabatic approximation
- simplified treatment for multiple scattering:  $(1 - x)^N \rightarrow e^{-Nx}$



Density distribution which explains the experimental  $\sigma_R$

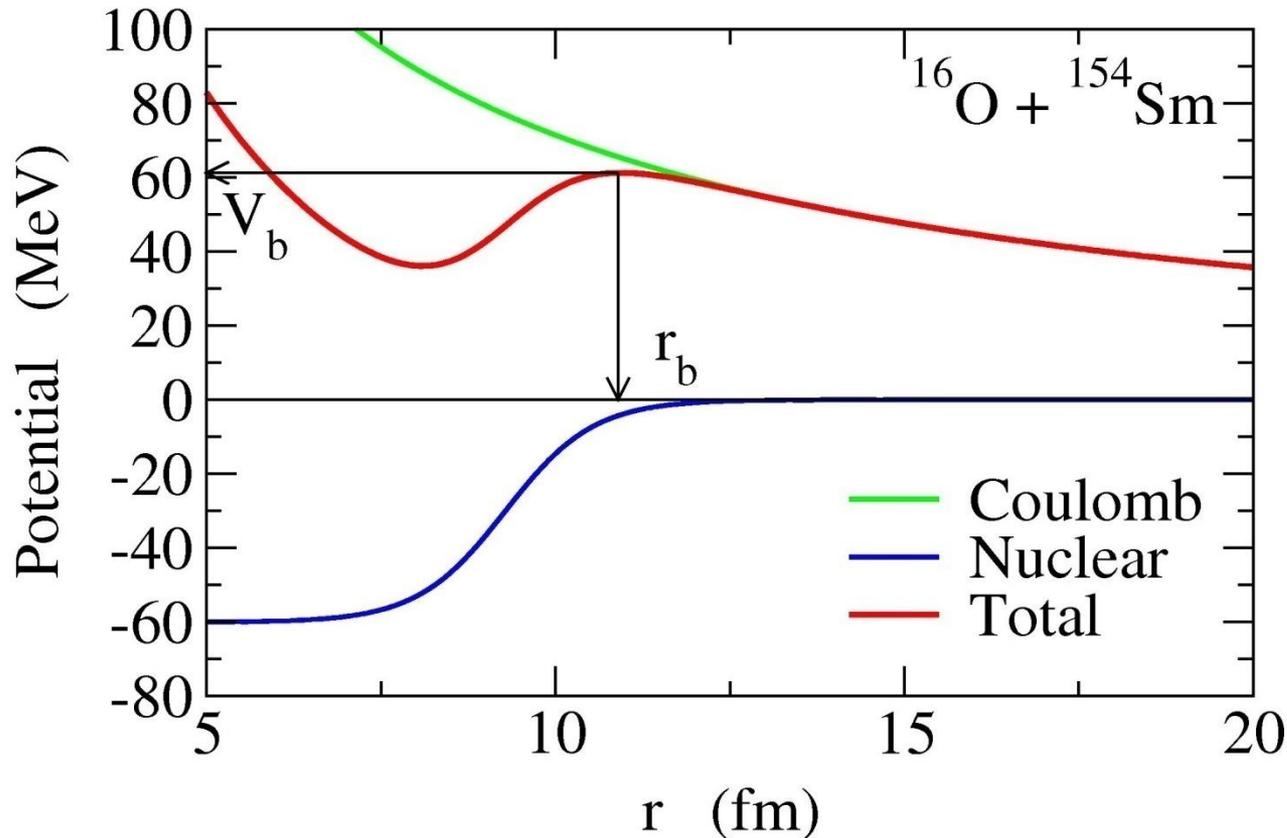


M. Fukuda et al., PLB268('91)339

$$\sigma_R \sim 2\pi \int_0^\infty b db \left[ 1 - \exp \left( -\sigma_{NN} \int d^2s \rho_P^{(z)}(s) \rho_T^{(z)}(s-b) \right) \right]$$

# Heavy-ion subbarrier fusion reactions

## Inter-nucleus potential



- above barrier
- sub-barrier
- deep subbarrier

Two forces:

### 1. Coulomb force

Long range,  
repulsive

### 2. Nuclear force

Short range,  
attractive



Potential barrier due to the compensation between the two  
(Coulomb barrier)

# Three important features of heavy-ion reactions

1. Coulomb interaction: important

2. Reduced mass: large  $\longrightarrow$

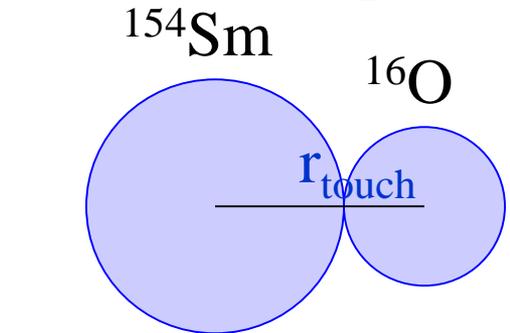
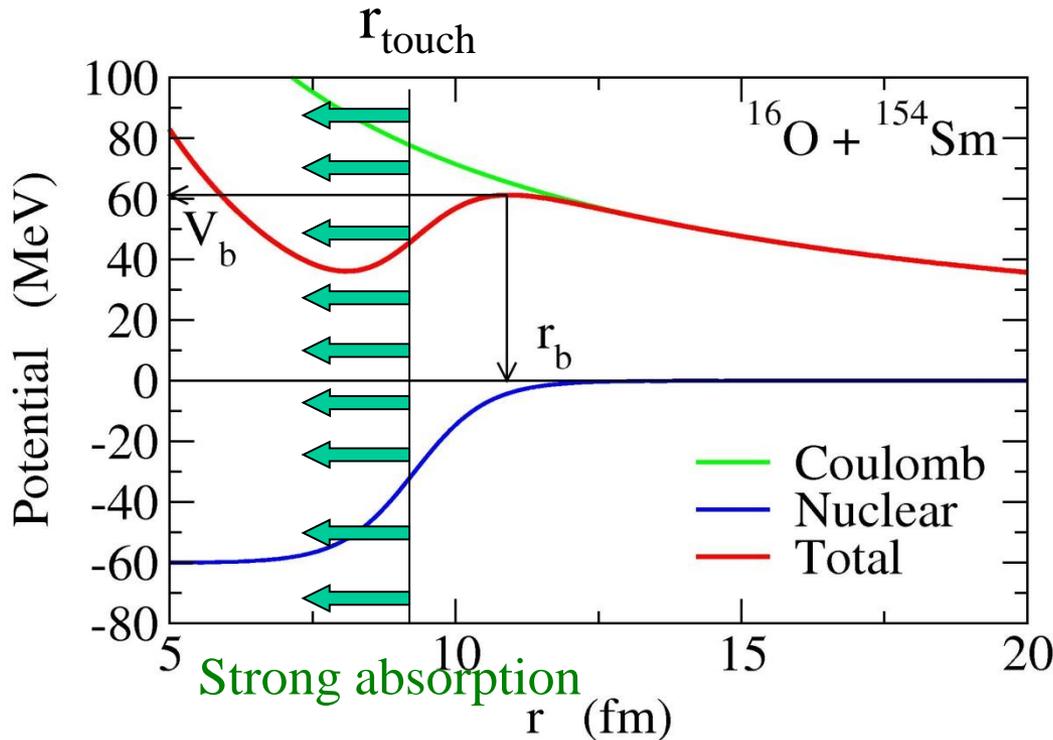
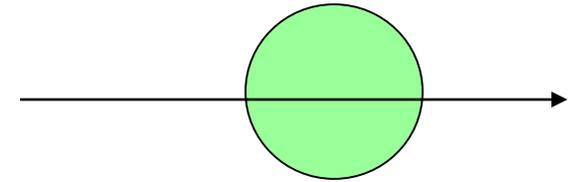
$$\mu = \frac{m_T m_P}{m_T + m_P}$$

(semi-) classical picture

concept of trajectory

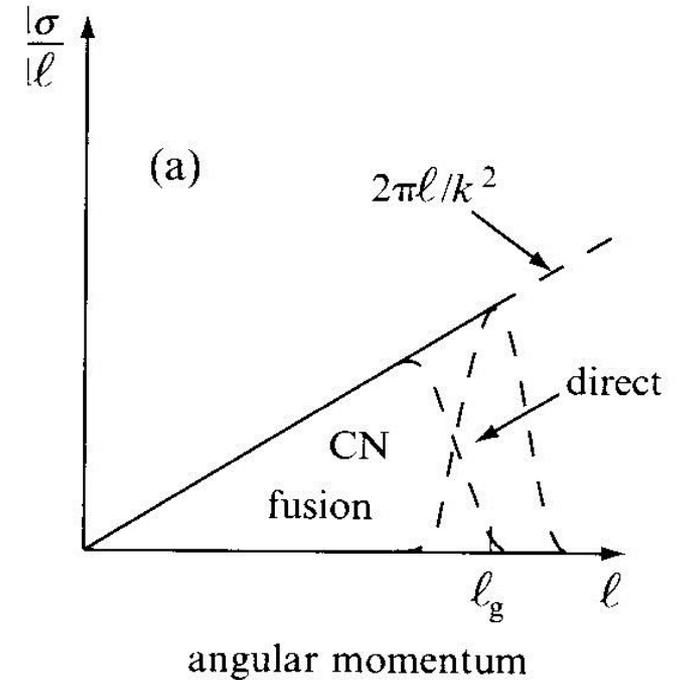
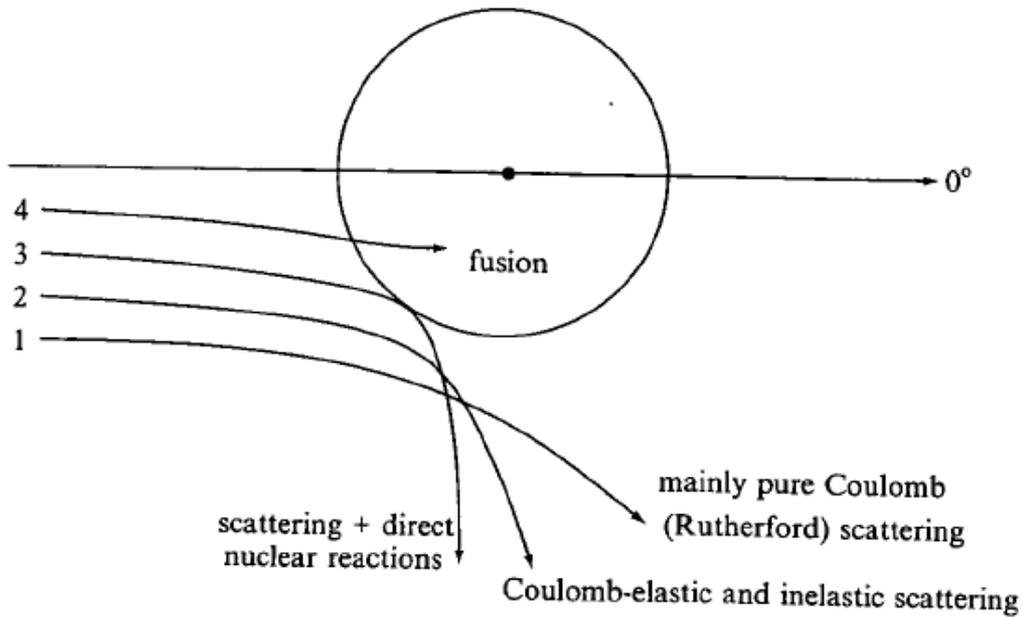


3. Strong absorption inside the Coul. barrier



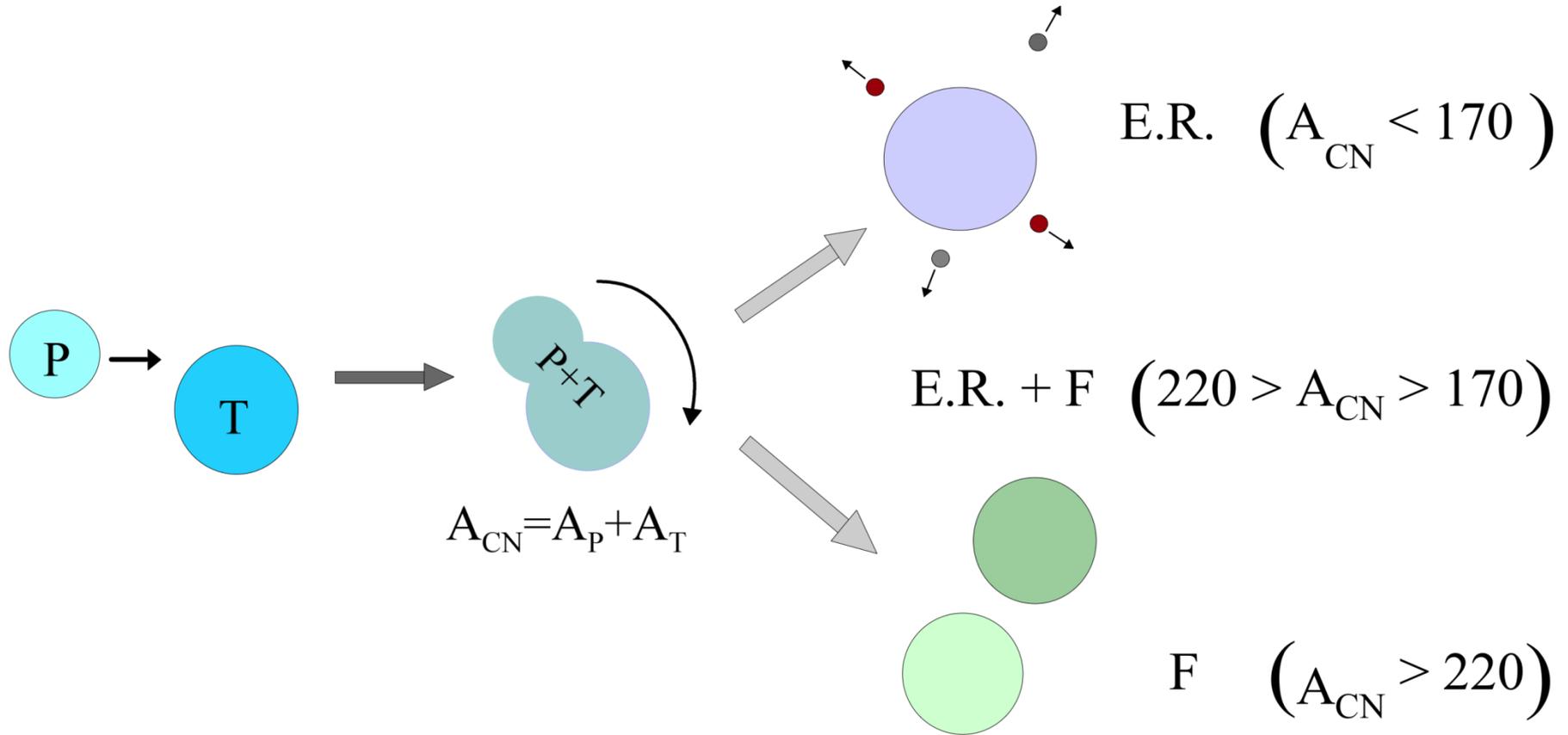
*Automatic* compound nucleus formation once touched (assumption of strong absorption)

# Partial decomposition of reaction cross section



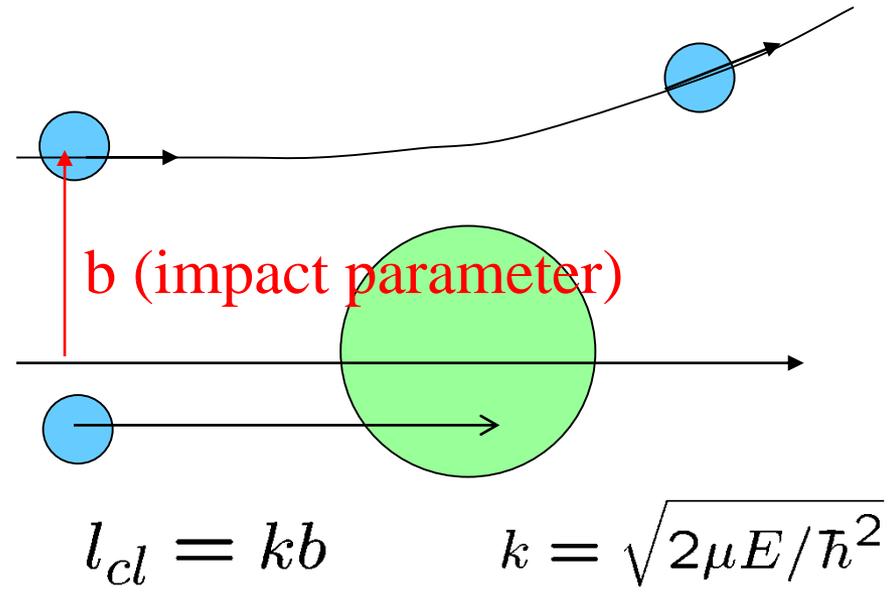
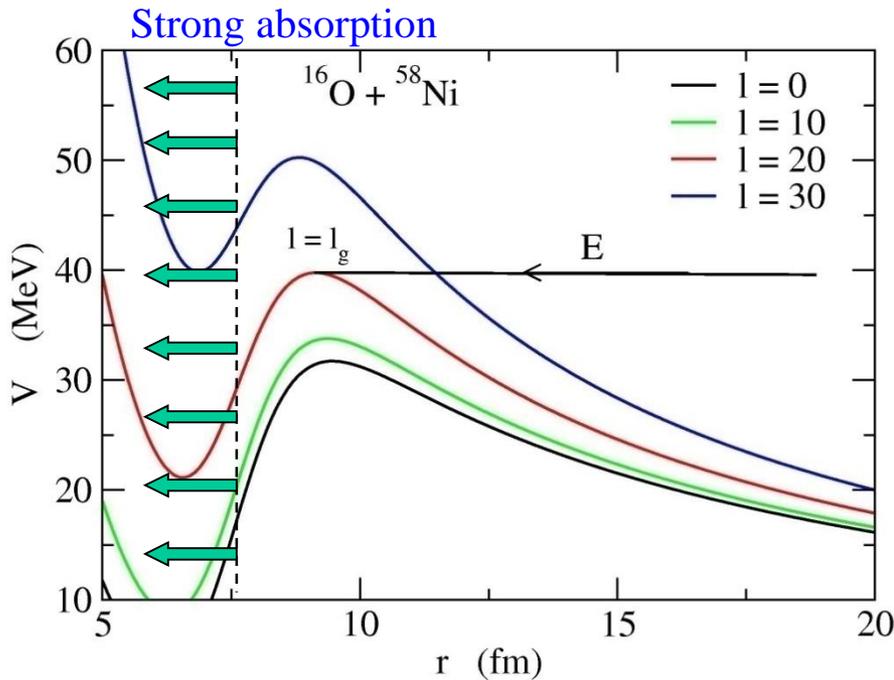
Taken from J.S. Lilley,  
"Nuclear Physics"

# Fusion: compound nucleus formation



courtesy: Felipe Canto

# classical fusion cross sections

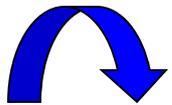


$l < l_g$  : can access to the strong absorption region classically

$\Rightarrow b_g = l_g/k$

$$\sigma^{cl} = 2\pi \int_0^{b_g} b db = \pi b_g^2$$

$$V_b + \frac{(kb_g)^2 \hbar^2}{2\mu R_b^2} = E$$

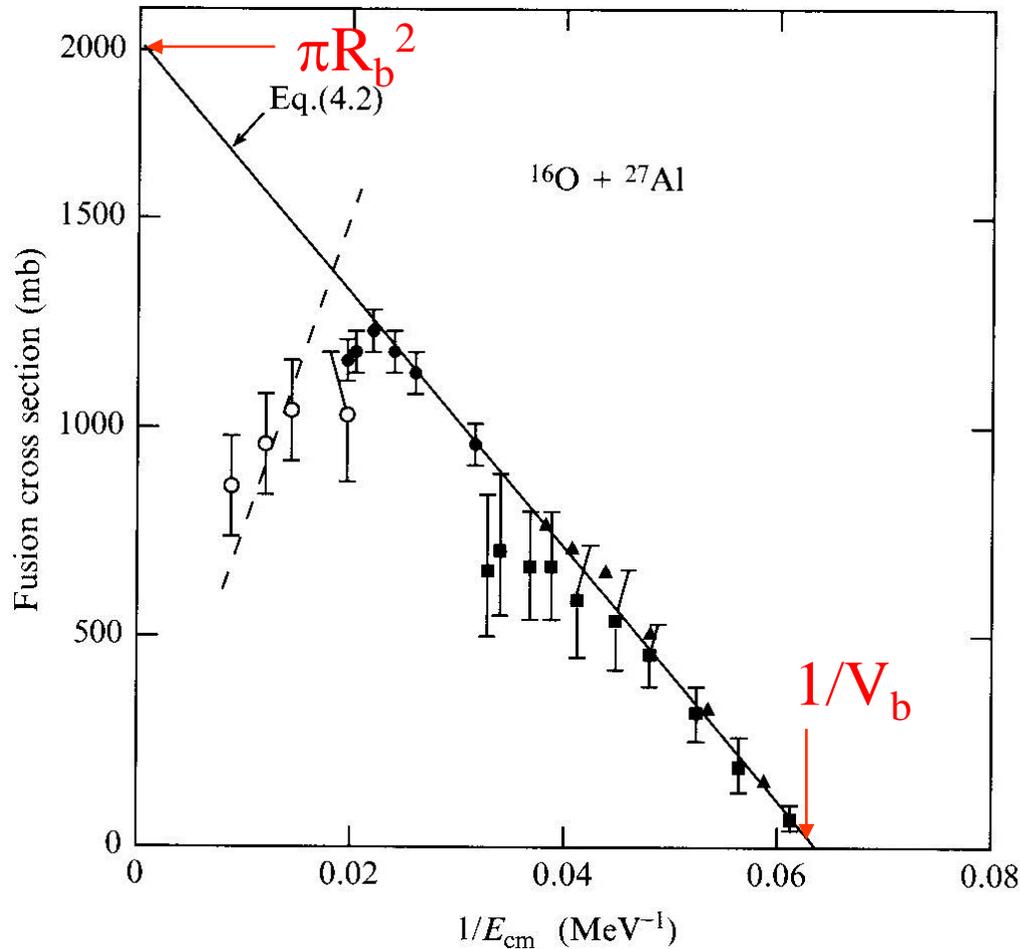


$$\sigma_{fus}^{cl}(E) = \pi R_b^2 \left( 1 - \frac{V_b}{E} \right)$$

## $\sigma_{\text{fus}}$ vs $1/E$ (~70's)

Classical fusion cross section is proportional to  $1/E$  :

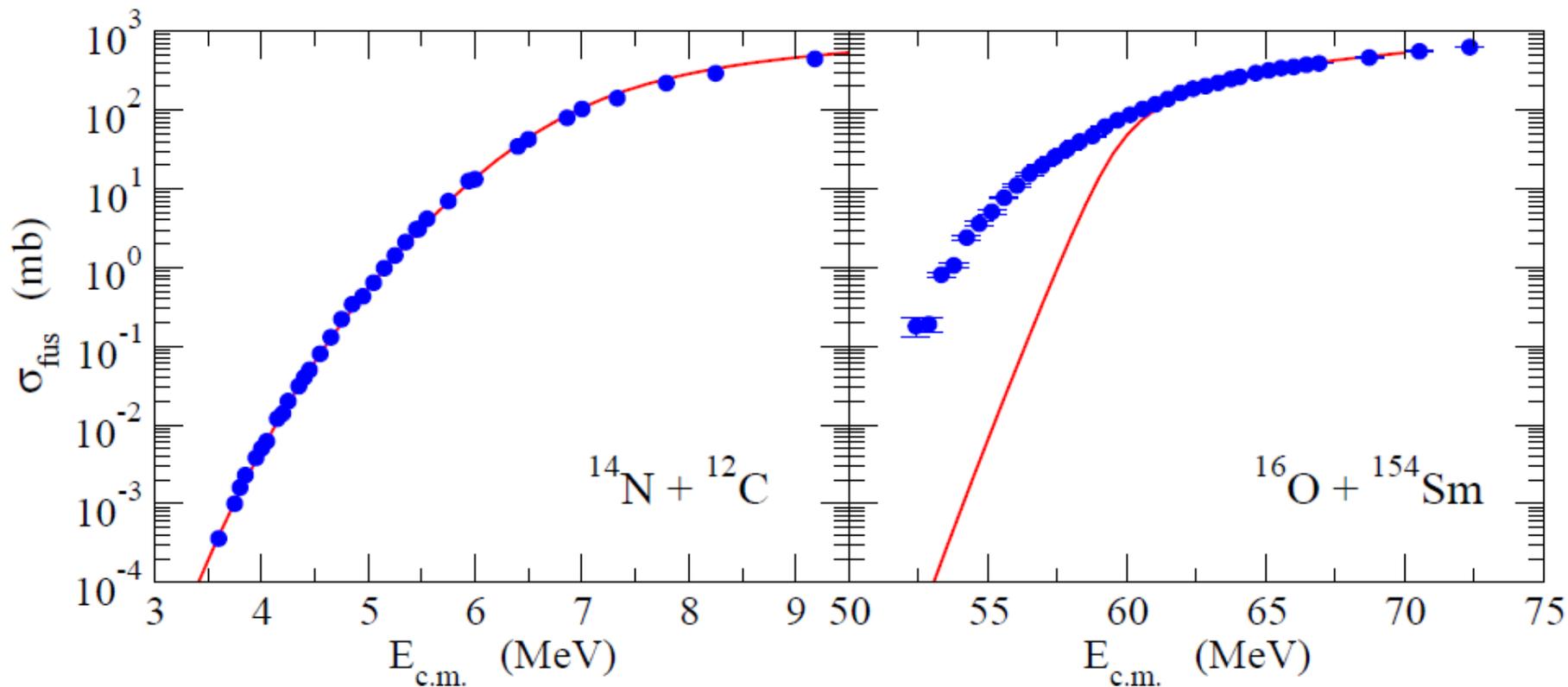
$$\sigma_{\text{fus}}^{\text{cl}}(E) = \pi R_b^2 \left(1 - \frac{V_b}{E}\right)$$



Taken from J.S. Lilley,  
"Nuclear Physics"

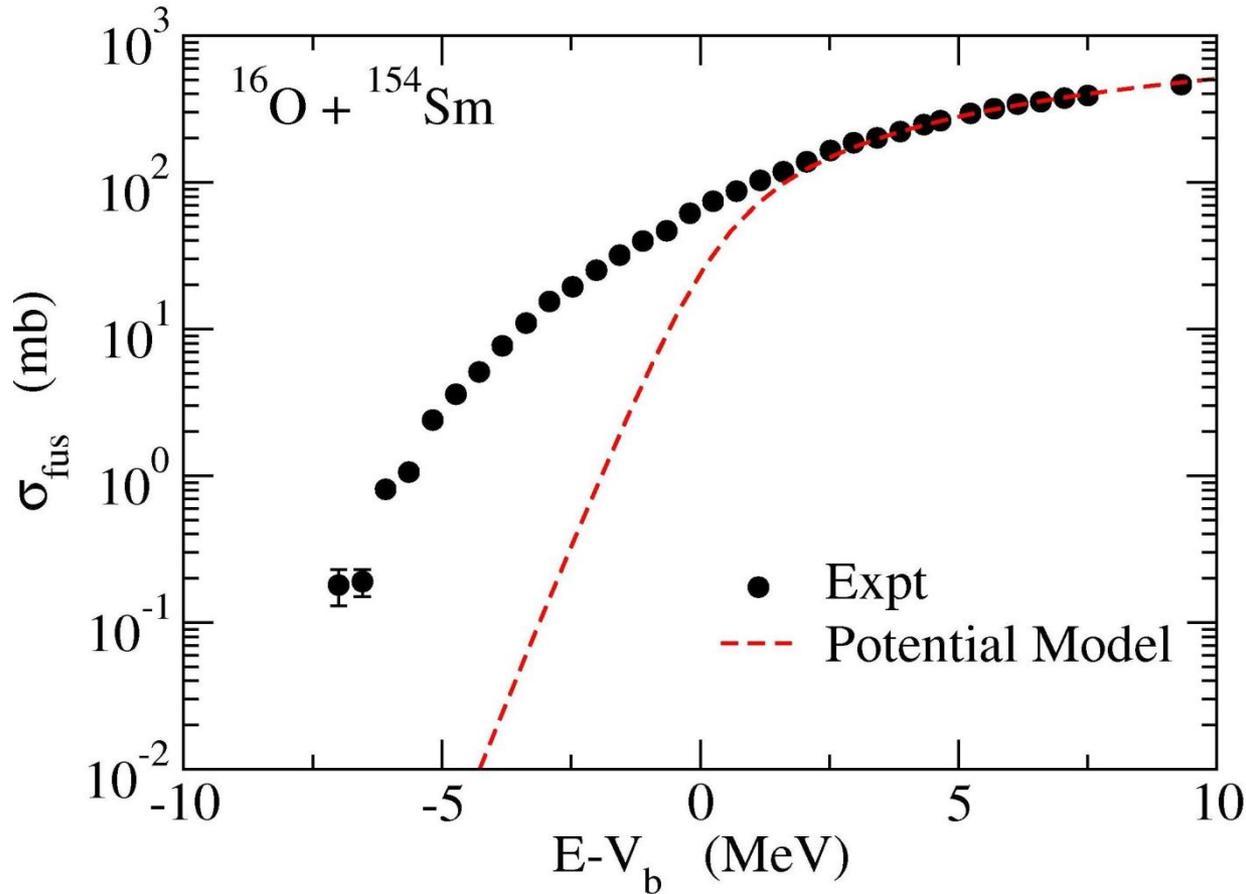
## Fusion cross sections at subbarrier energies

Fusion cross sections of structure-less nuclei (a potential model)



Simple potential model:

- OK for relatively light systems
- underestimates  $\sigma_{\text{fus}}$  for heavier systems at subbarrier energies



Potential model:

Reproduces the data reasonably well for

$$E > V_b$$

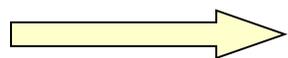
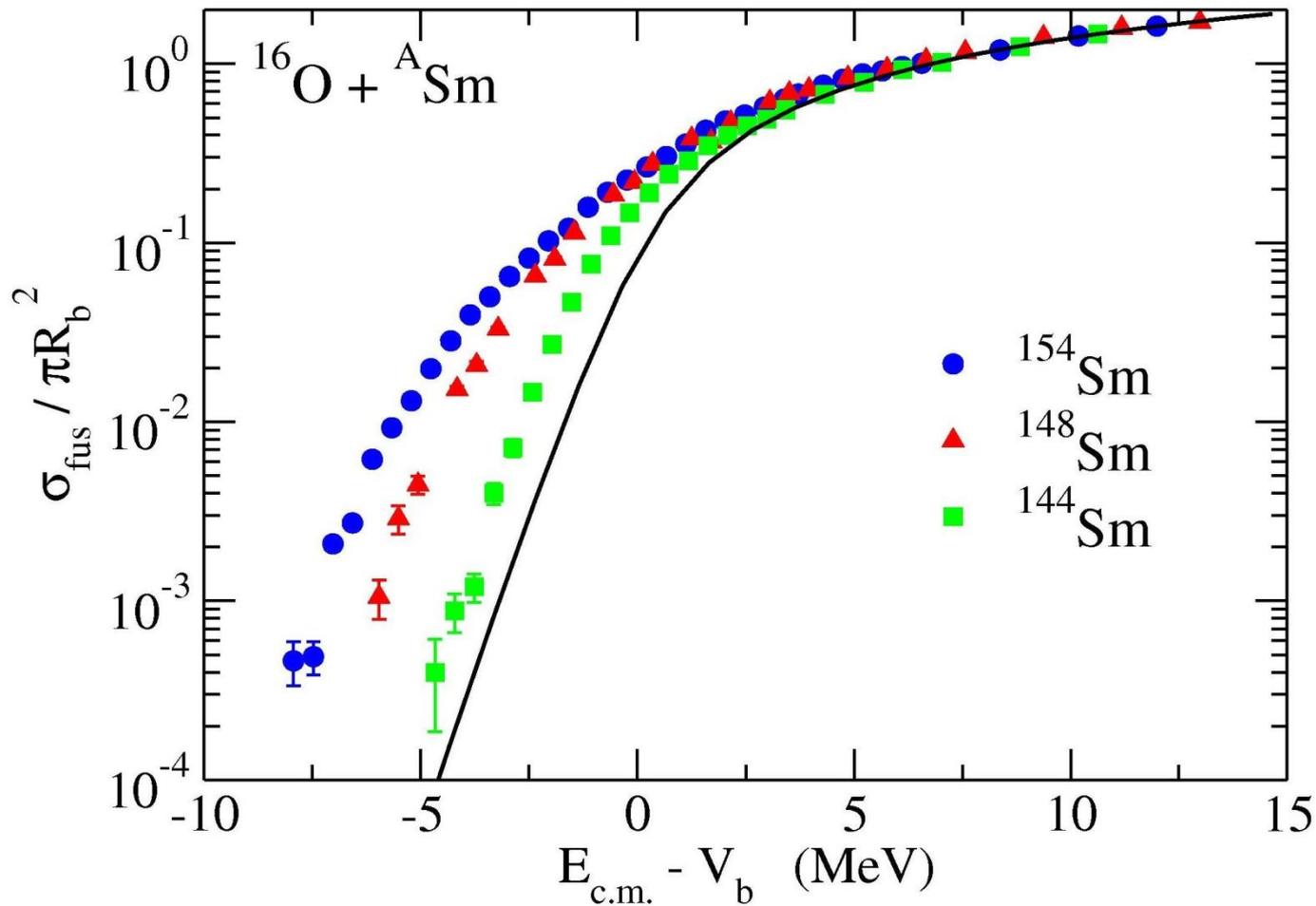
Underpredicts  $\sigma_{\text{fus}}$  for

$$E < V_b$$

cf. seminal work:

R.G. Stokstad et al., PRL41('78)465

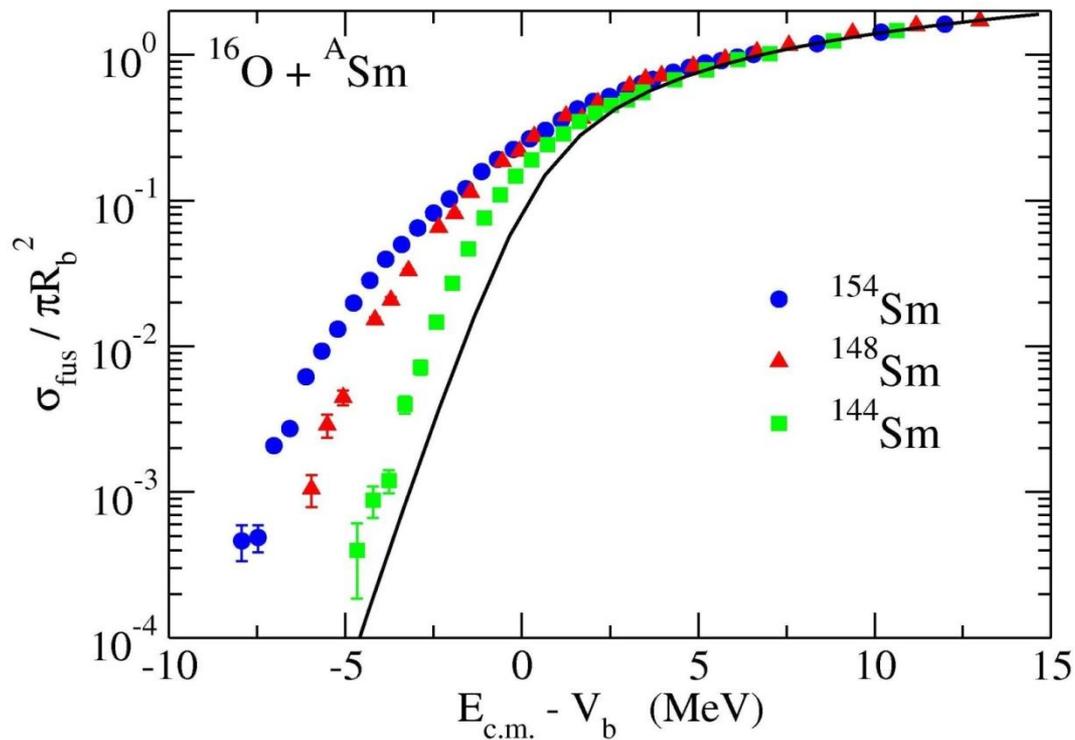
PRC21('80)2427



Strong target dependence at  $E < V_b$



low-lying collective excitations?



(MeV)

1.81 —  $3^-$   
1.66 —  $2^+$

(MeV)

1.18 —  $4^+$   
1.16 —  $3^-$

(MeV)

0.90 —  $8^+$   
0.54 —  $6^+$   
0.27 —  $4^+$   
0.082 —  $2^+$   
0 —  $0^+$

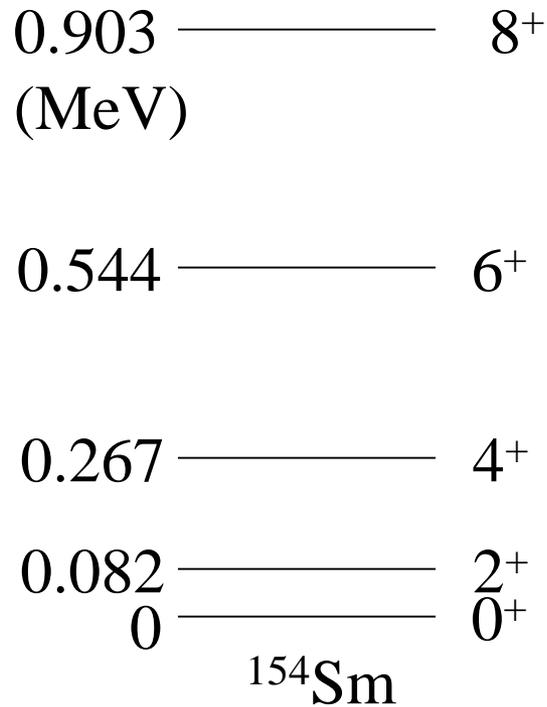
0 —  $0^+$   
 $^{144}\text{Sm}$

0 —  $0^+$   
 $^{148}\text{Sm}$

$^{154}\text{Sm}$

# Effect of deformation on subbarrier fusion

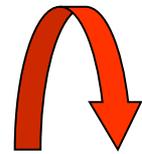
## Excitation spectra of $^{154}\text{Sm}$



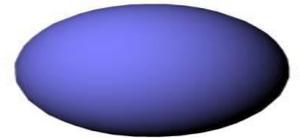
cf. Rotational energy of a rigid body  
(Classical mechanics)

$$E = \frac{1}{2} \mathcal{J} \omega^2 = \frac{I^2}{2\mathcal{J}}$$

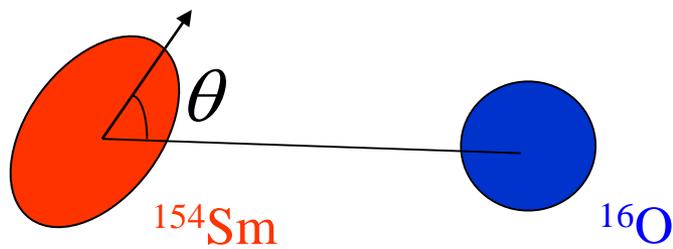
$$(I = \mathcal{J}\omega, \omega = \dot{\theta})$$



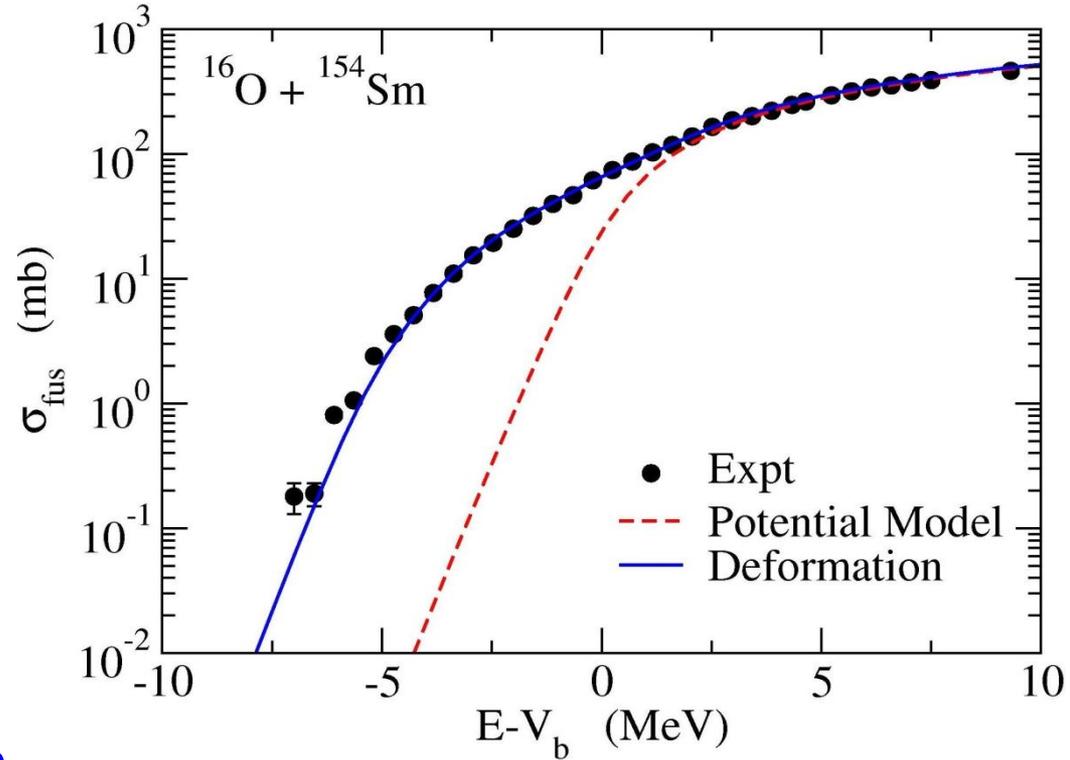
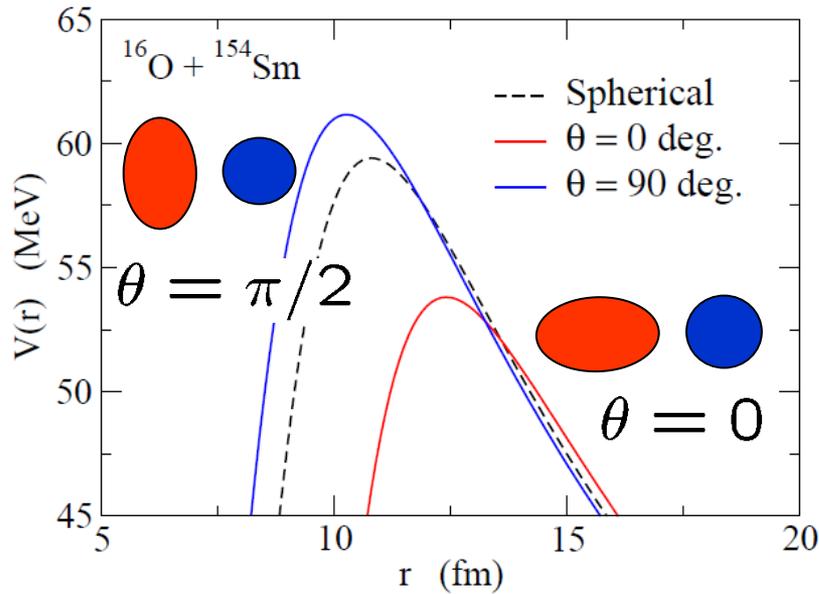
$^{154}\text{Sm}$  is deformed



$$E_I \sim \frac{I(I+1)\hbar^2}{2\mathcal{J}}$$



$$\sigma_{\text{fus}}(E) = \int_0^1 d(\cos \theta) \sigma_{\text{fus}}(E; \theta)$$



The barrier is lowered for  $\theta=0$  because an attraction works from large distances.

The barrier increases for  $\theta=\pi/2$  because the rel. distance has to get small for the attraction to work

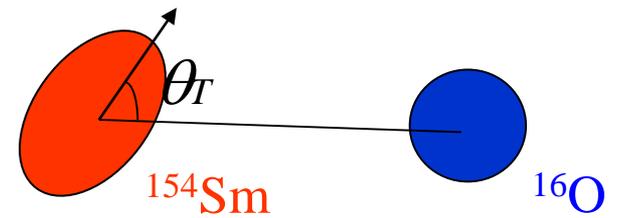
**Def. Effect:** enhances  $\sigma_{\text{fus}}$  by a factor of 10 ~ 100



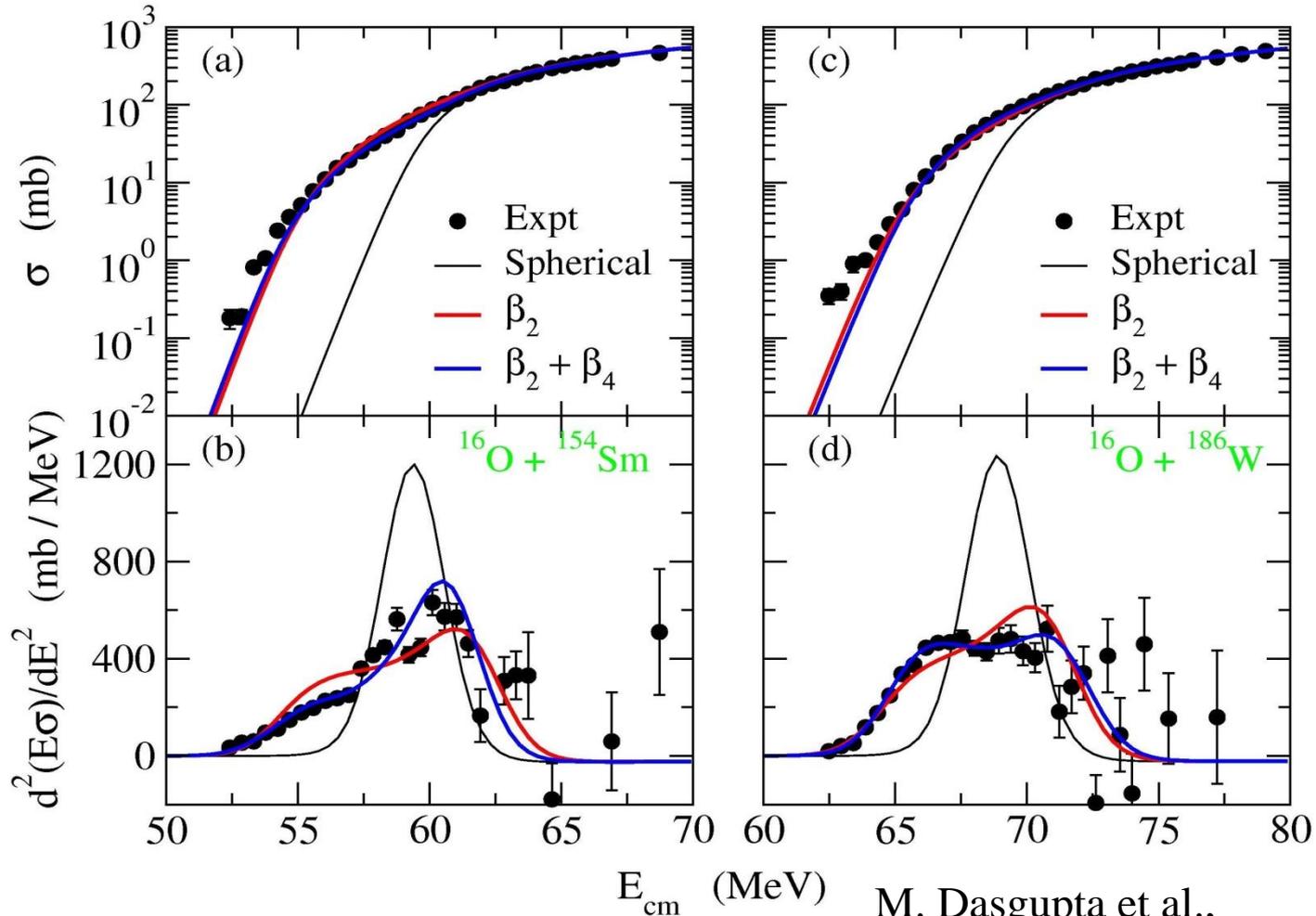
**Fusion:** interesting probe for nuclear structure

# Fusion barrier distributions

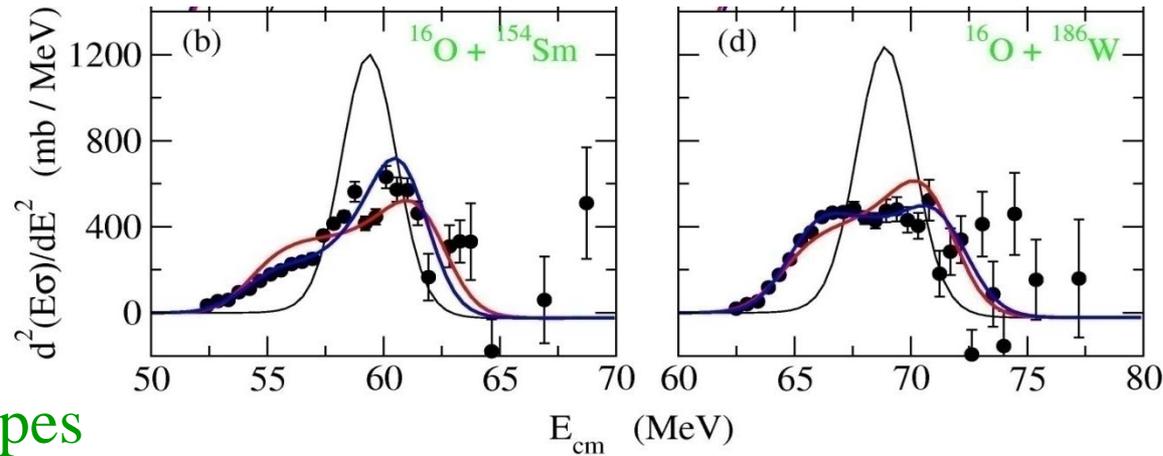
$$D_{\text{fus}}(E) = \frac{d^2(E\sigma_{\text{fus}})}{dE^2}$$



$$\sigma_{\text{fus}}(E) = \int_0^1 d(\cos \theta_T) \sigma_{\text{fus}}(E; \theta_T)$$



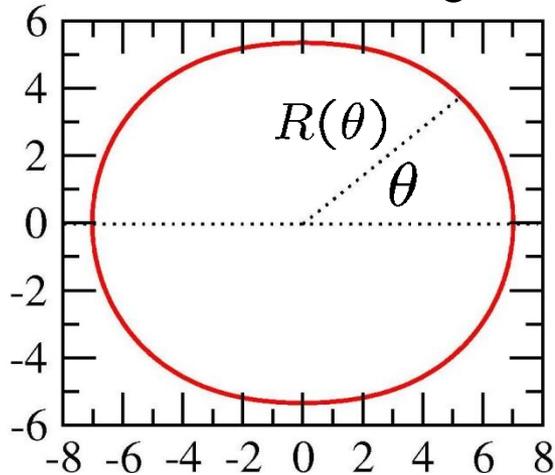
# Investigate nuclear shape through barrier distribution



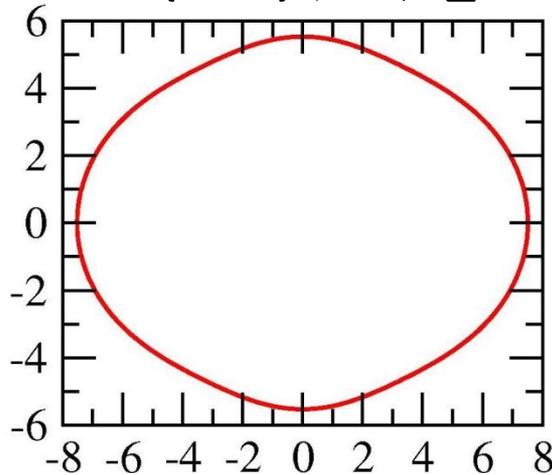
## Nuclear shapes

$$R(\theta) = R_0(1 + \beta_2 Y_{20}(\theta) + \beta_4 Y_{40}(\theta) + \dots)$$

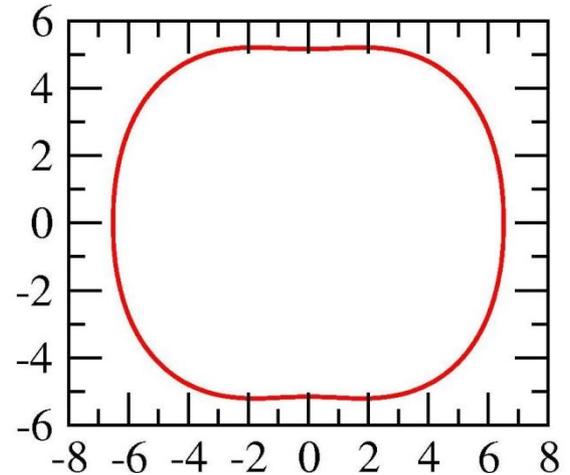
$$R_0 = 5.9 \text{ (fm)}, \quad \beta_2 = 0.3$$



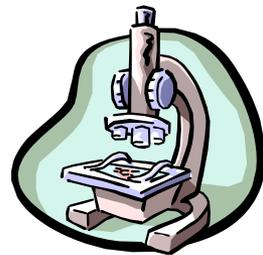
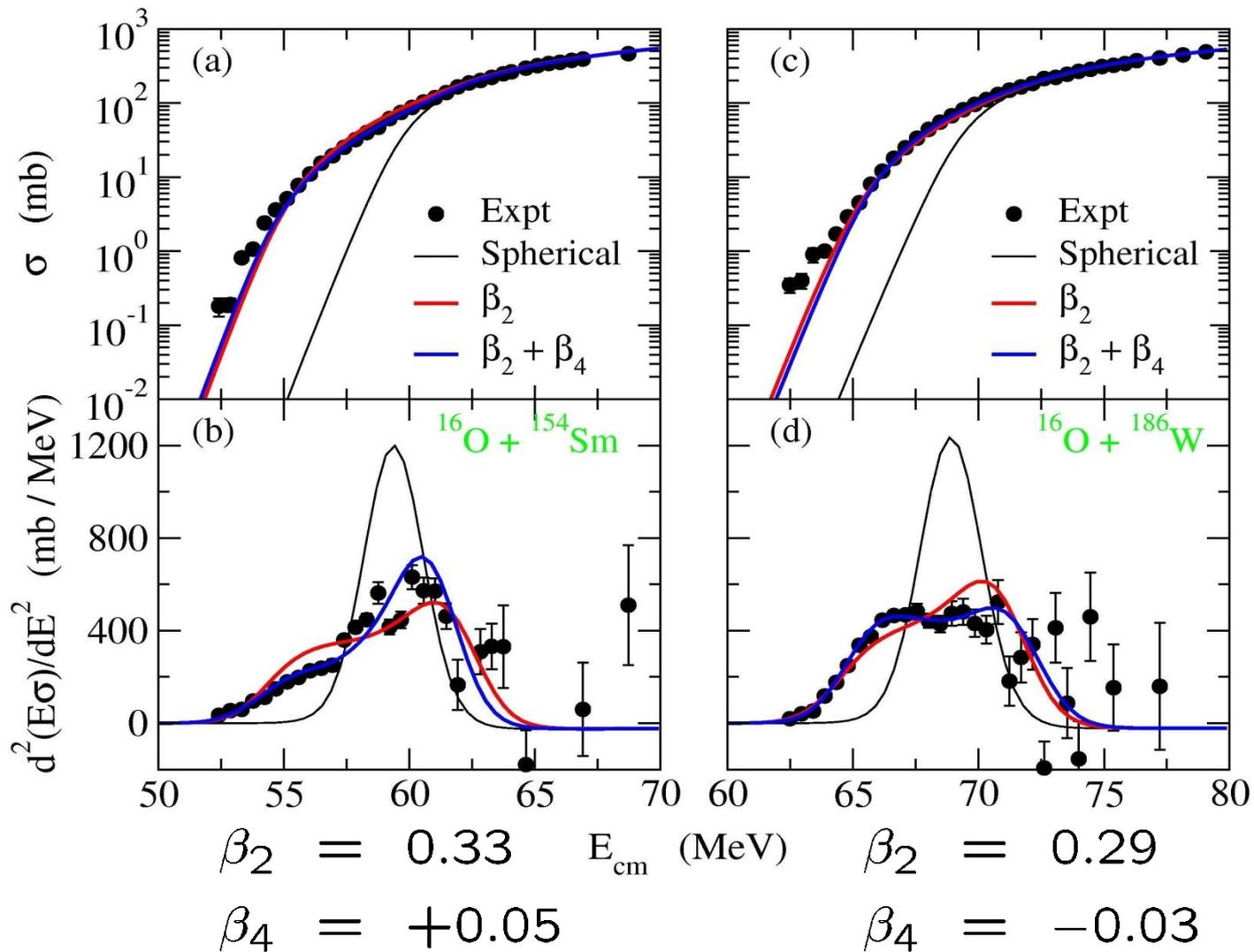
$$\beta_4 = 0$$



$$\beta_4 = 0.1$$



$$\beta_4 = -0.1$$

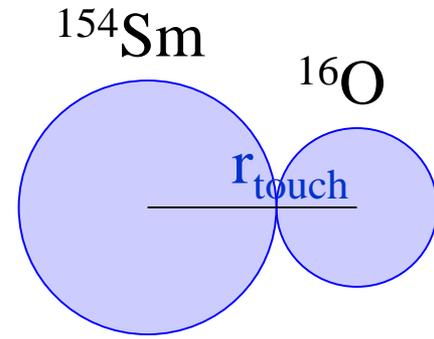
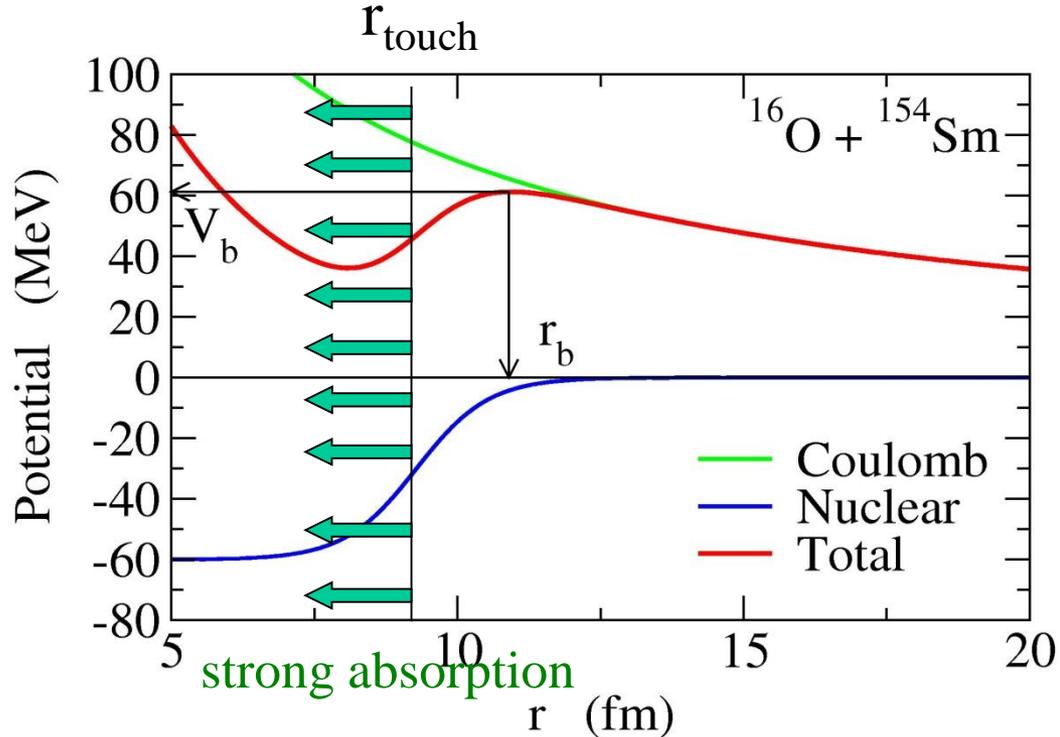


By taking the barrier distribution, one can very clearly see the difference due to  $\beta_4$ !

➡ Fusion as a quantum tunneling microscope for nuclei

# Heavy-ion subbarrier fusion reactions in the SHE region

## ➤ Fusion of medium-heavy systems:

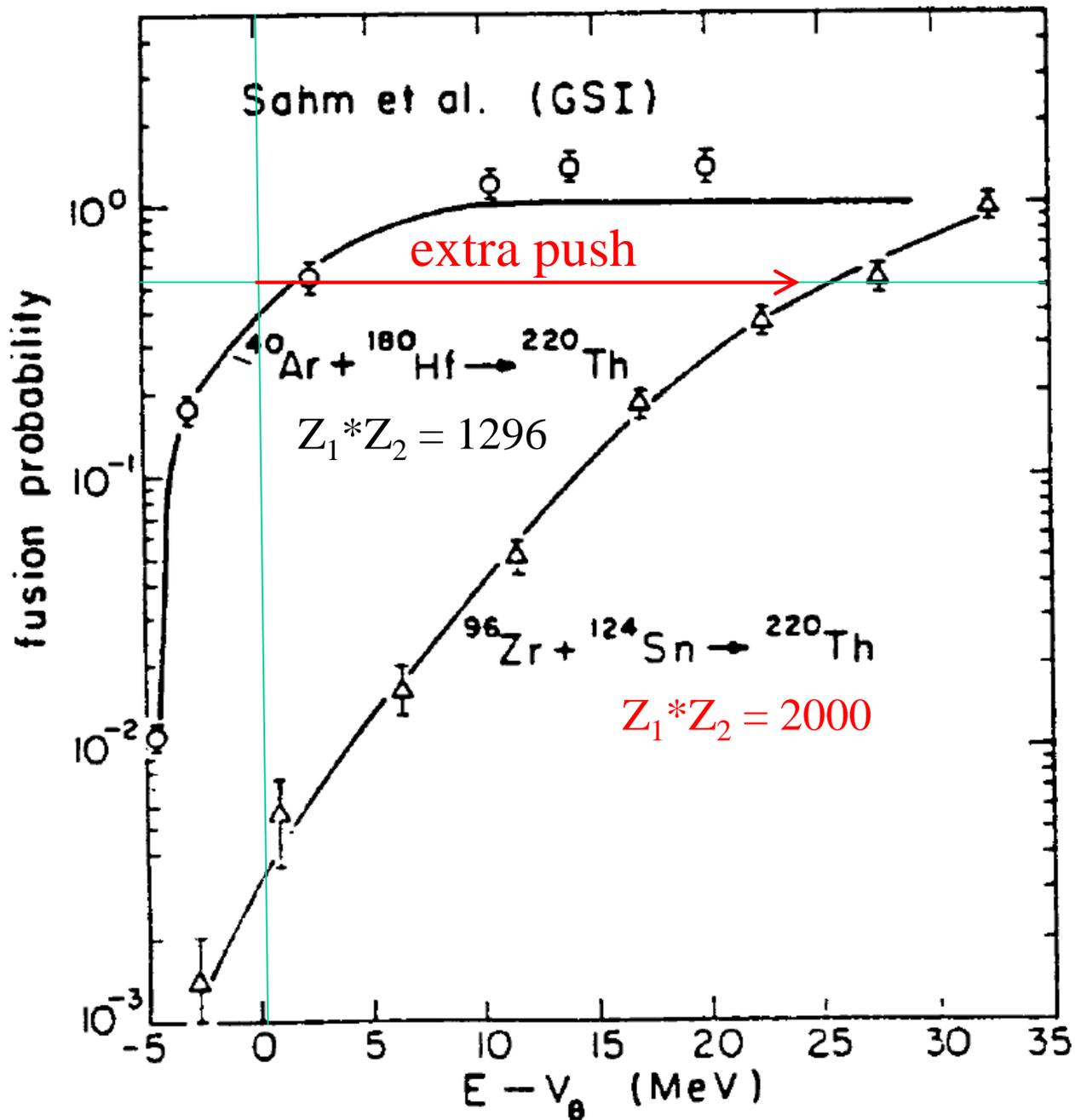


Compound nucleus:  
automatically formed  
once touched  
(strong absorption)

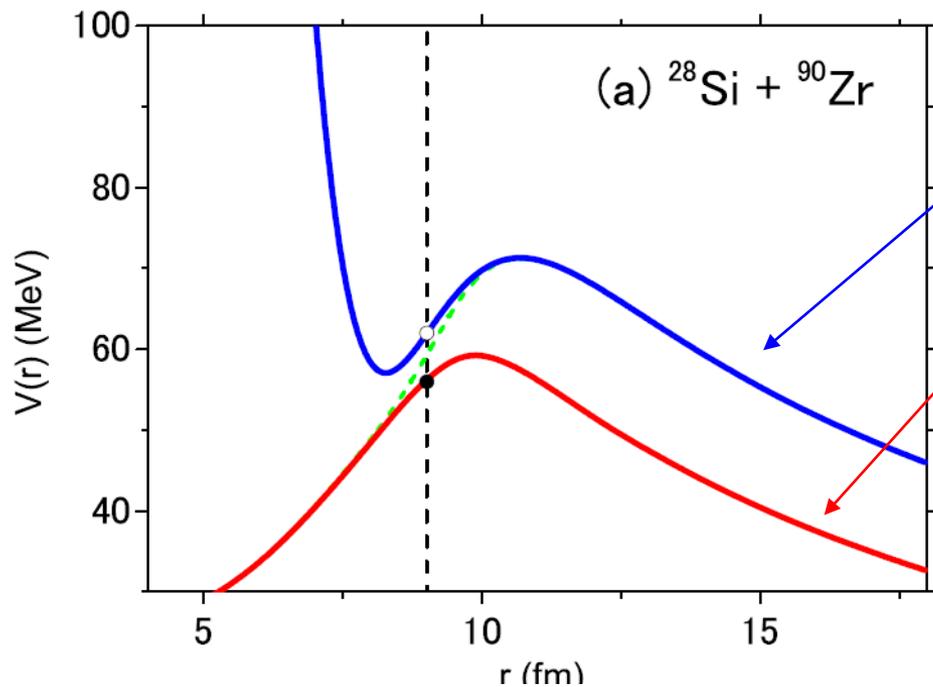
## ➤ Fusion of heavy and super-heavy systems

Large probability of re-separation (due to the strong Coulomb repulsion)

[This happens for  $Z_1 * Z_2 > 1600 \sim 1800$ .]

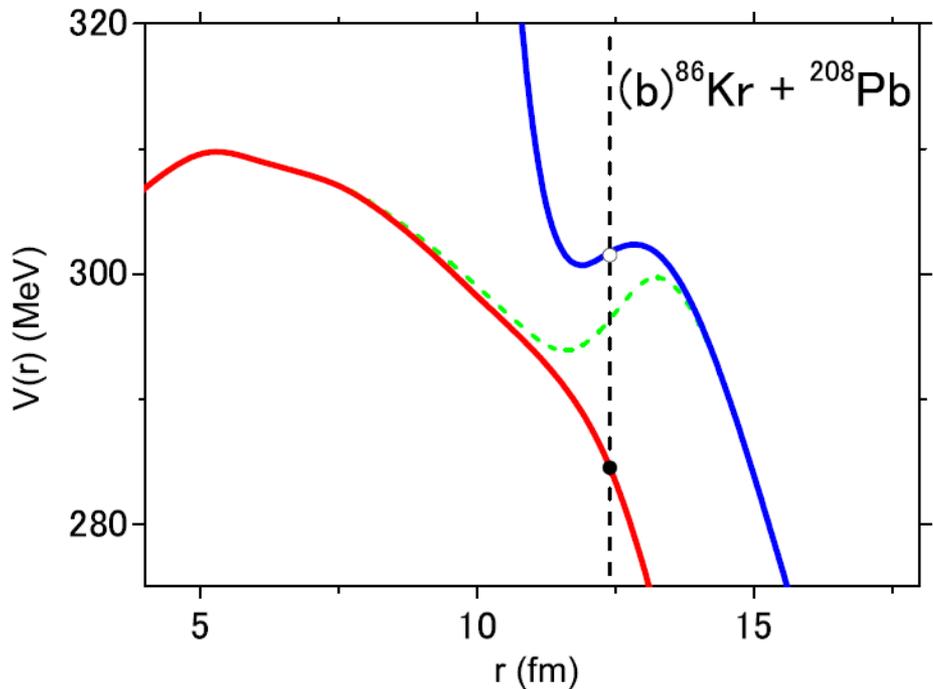


C.-C. Sahm et al.,  
 Z. Phys. A319('84)113



2-body potential before touching

1-body potential after touching

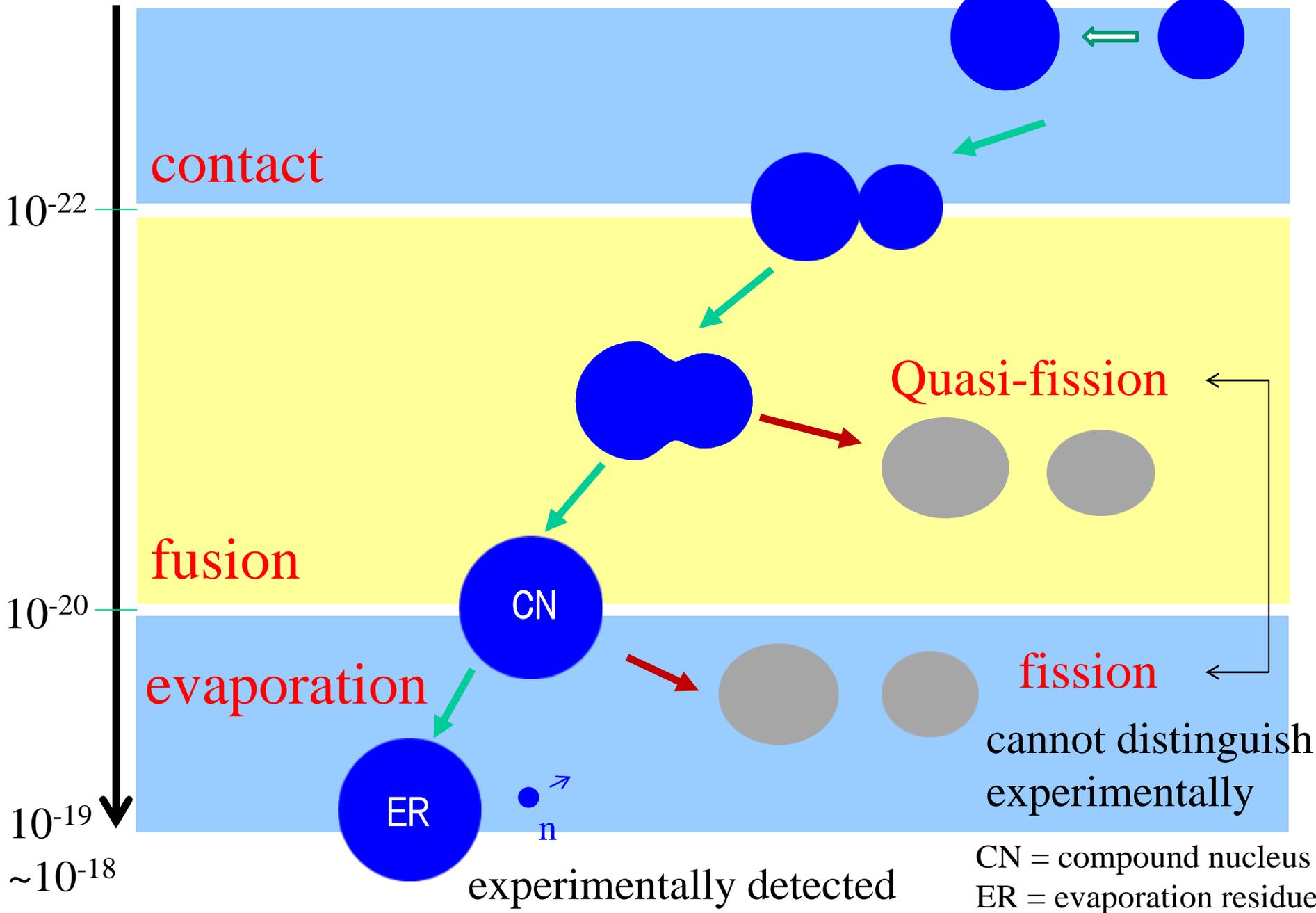


The red potential has to be overcome even if the blue potential has been overcome.

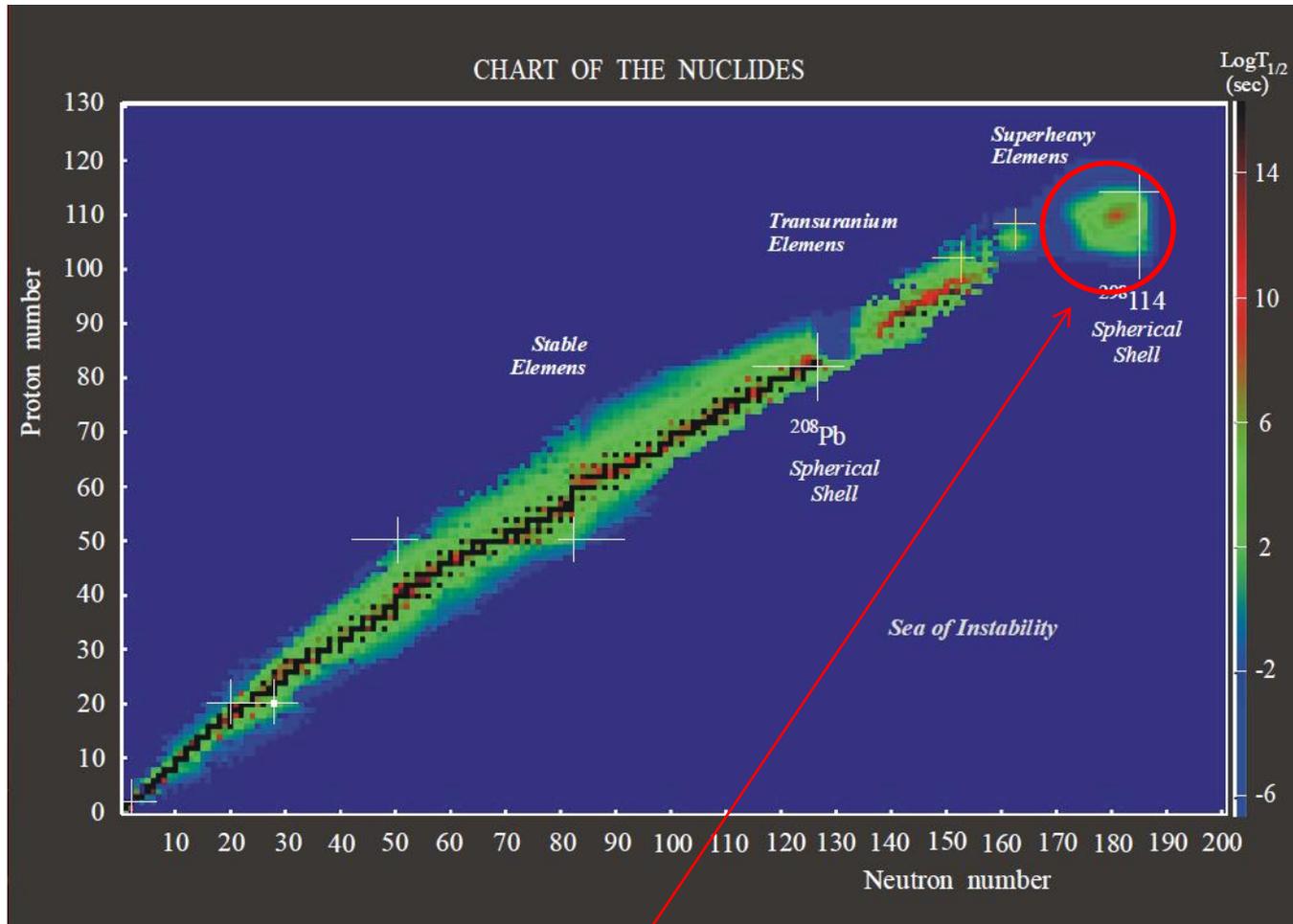


Re-separation if failed  
(quasi-fission)

typical time-scale (sec.)



# Heavy-ion fusion for SHE



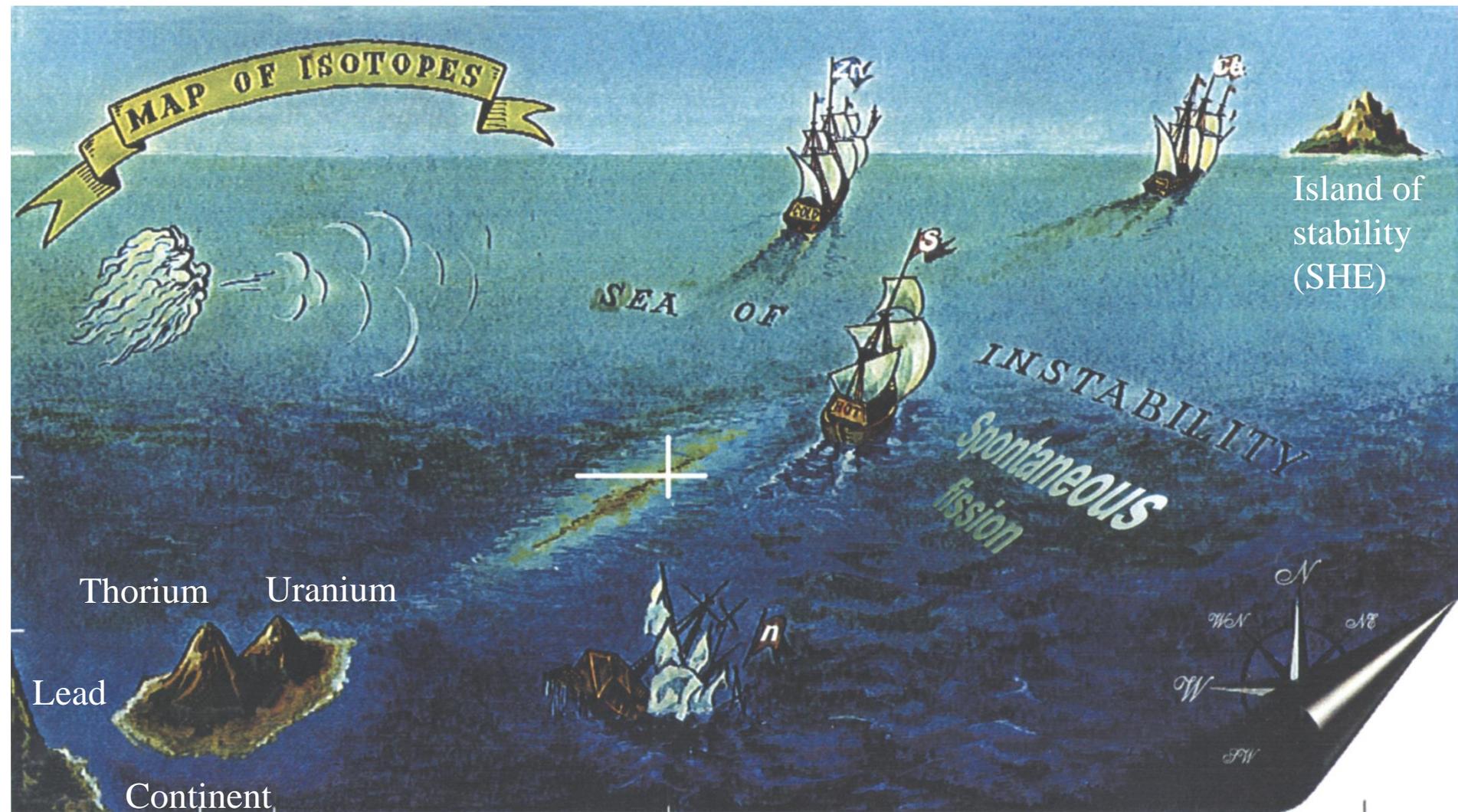
**island of stability around  $Z=114, N=184$**

Yuri Oganessian

W.D. Myers and W.J. Swiatecki (1966), A. Sobiczewski et al. (1966)

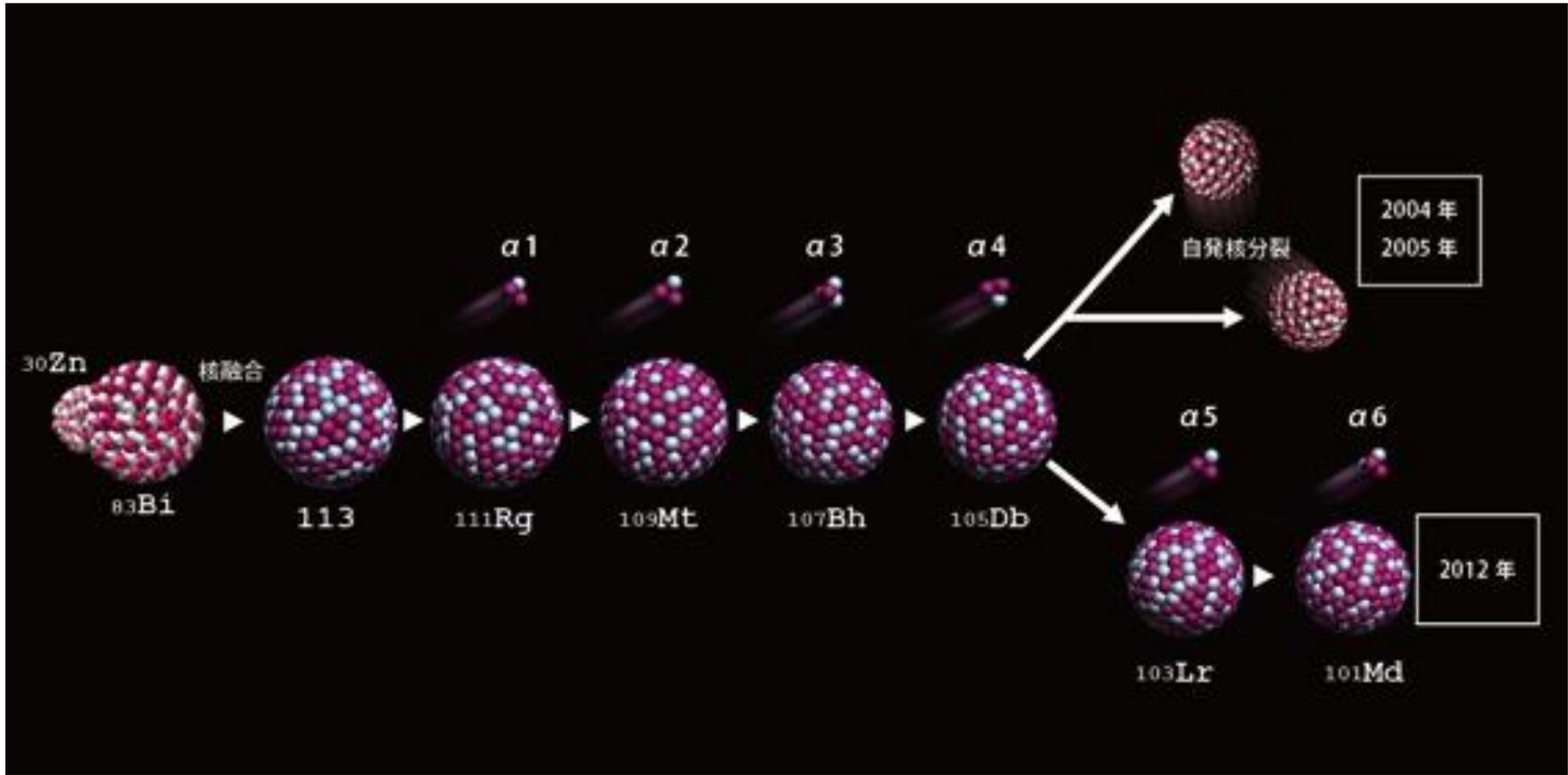
→ modern calculations:  $Z=114, 120, \text{ or } 126, N=184$

e.g., H. Koura et al. (2005)



Yuri Oganessian

# Element 113 (RIKEN, K. Morita et al.)

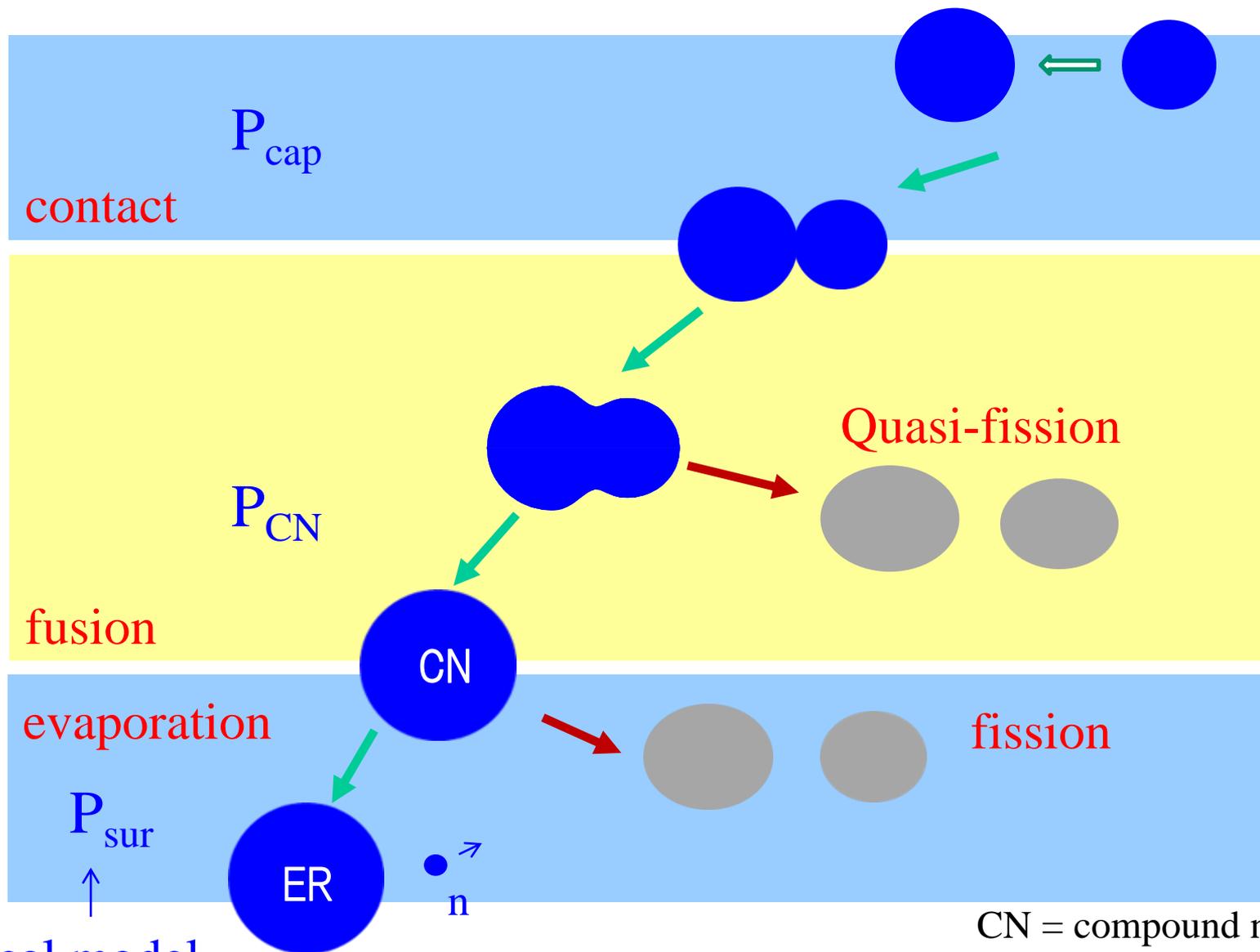


K. Morita et al., J. Phys. Soc. Jpn. 81('12)103201

$$\sigma_{\text{ER}} = 22^{+20}_{-13} \text{ fb} \quad \text{only 3 events for 553 days experiment}$$

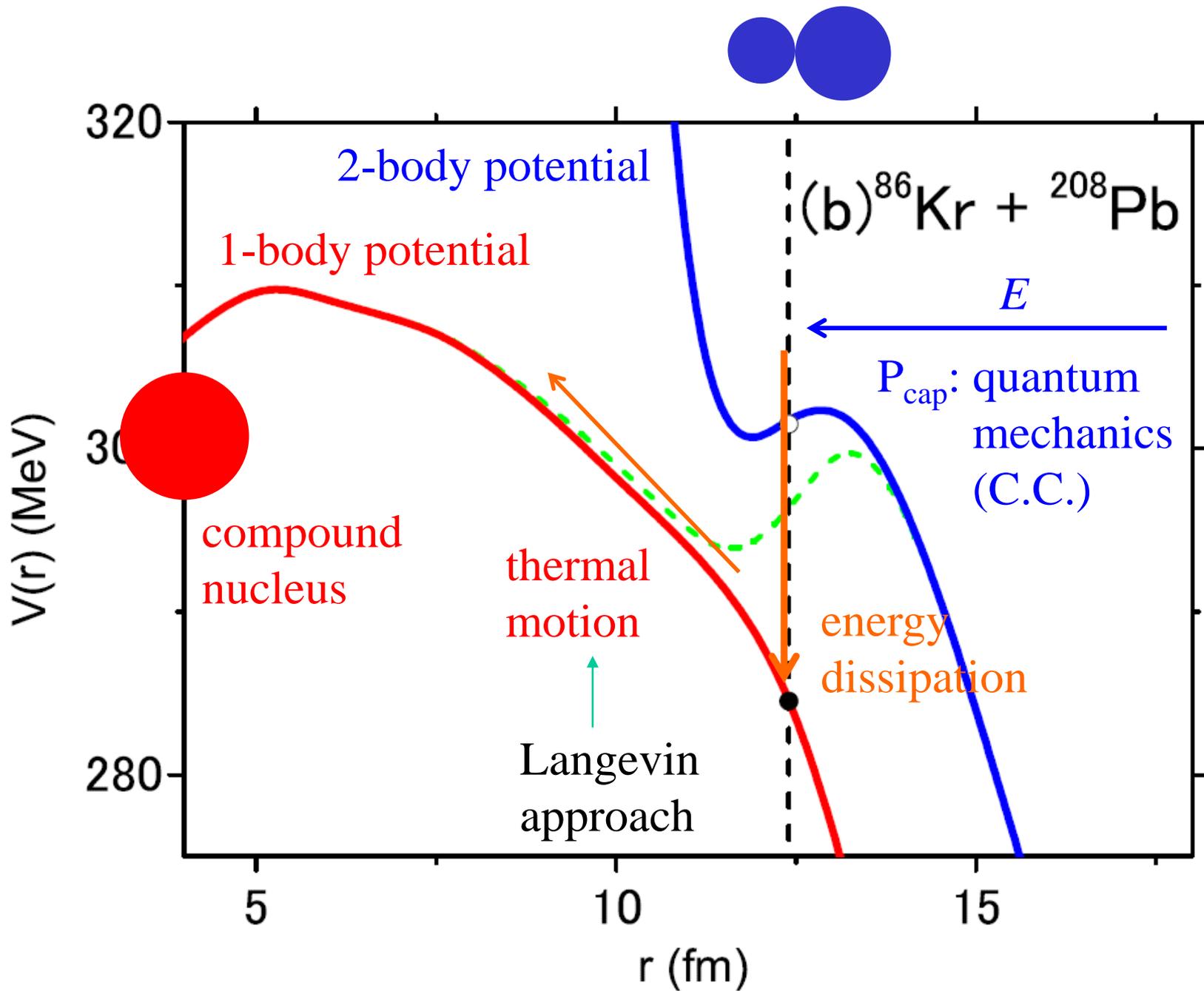
Theoretical treatment

$$P_{ER} = P_{cap} \cdot P_{CN} \cdot P_{sur}$$



statistical model

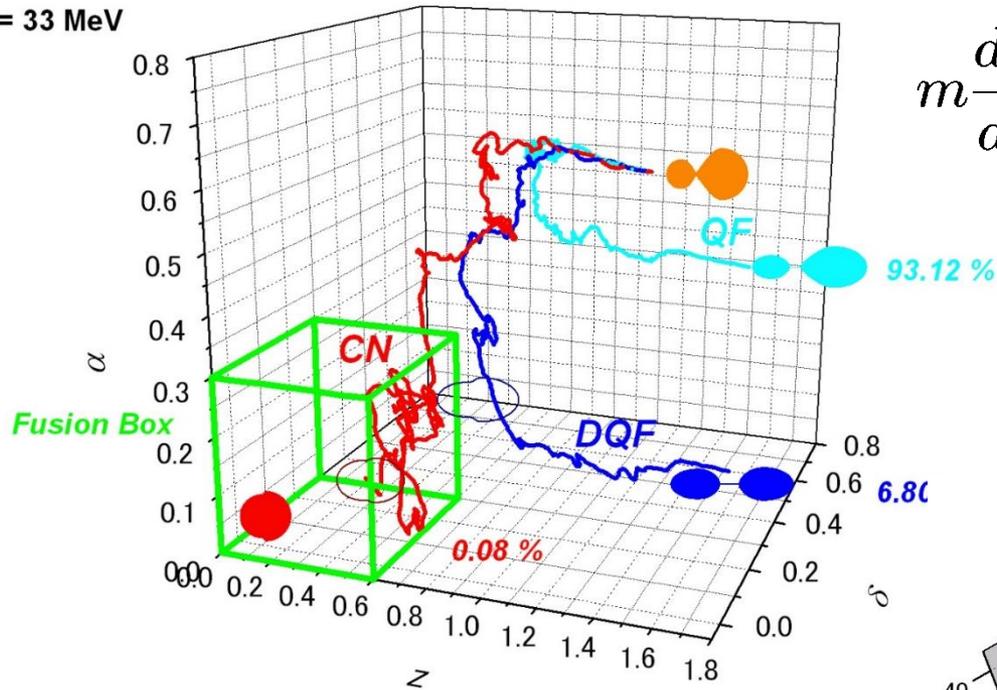
CN = compound nucleus  
ER = evaporation residue



# Theory: Lagenvin approach

$^{48}\text{Ca} + ^{244}\text{Pu} \rightarrow ^{292}\text{114}$

$E^* = 33 \text{ MeV}$

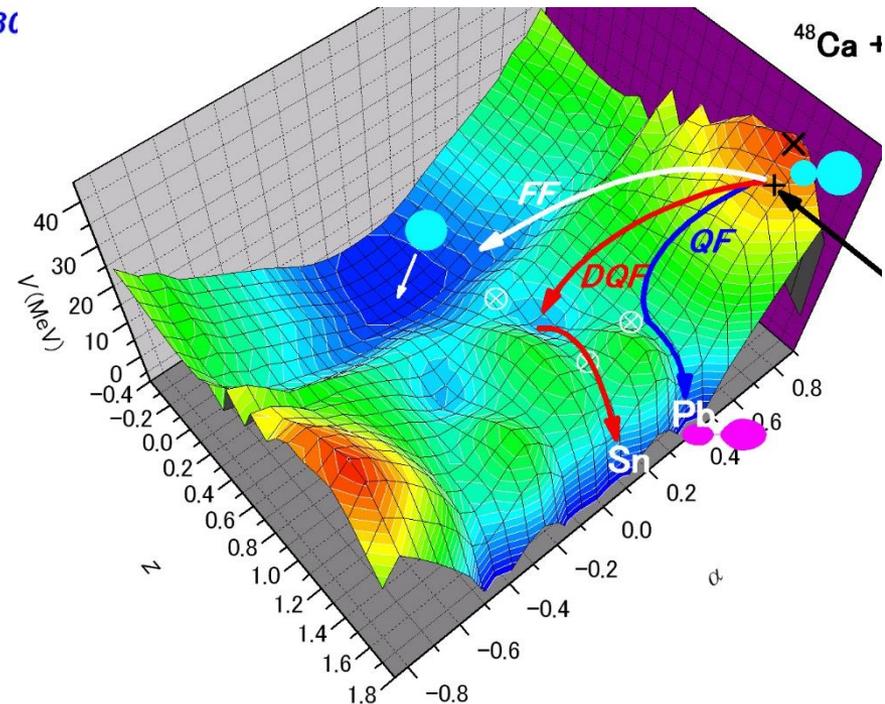


multi-dimensional extension of:

$$m \frac{d^2 q}{dt^2} = - \frac{dV(q)}{dq} - \gamma \frac{dq}{dt} + R(t)$$

$\gamma$ : friction coefficient  
 $R(t)$ : random force

- $q$ : ▪ internuclear separation ( $z$ ),
- deformation ( $\delta$ ),
- asymmetry of the two fragments ( $\alpha$ )



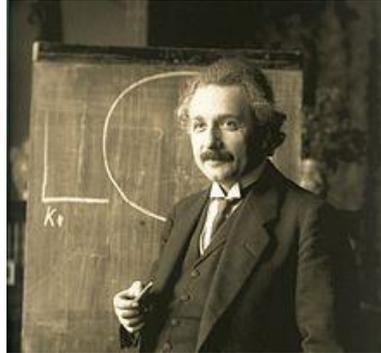
# Chemistry of superheavy elements

Group → ↓ Period	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18
1	1 H																	2 He
2	3 Li	4 Be											5 B	6 C	7 N	8 O	9 F	10 Ne
3	11 Na	12 Mg											13 Al	14 Si	15 P	16 S	17 Cl	18 Ar
4	19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr
5	37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe
6	55 Cs	56 Ba		72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn
7	87 Fr	88 Ra		104 Rf	105 Db	106 Sq	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Uut	114 Fl	115 Uup	116 Lv	117 Uus	118 Uuo
Lanthanides	57 La	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu			
Actinides	89 Ac	90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr			

- Are they here in the periodic table?
- That is, does e.g., Lv show the same chemical properties as O, S, Se, Te, and Po?

**relativistic effect** : important for large  $Z$

$$E = mc^2$$

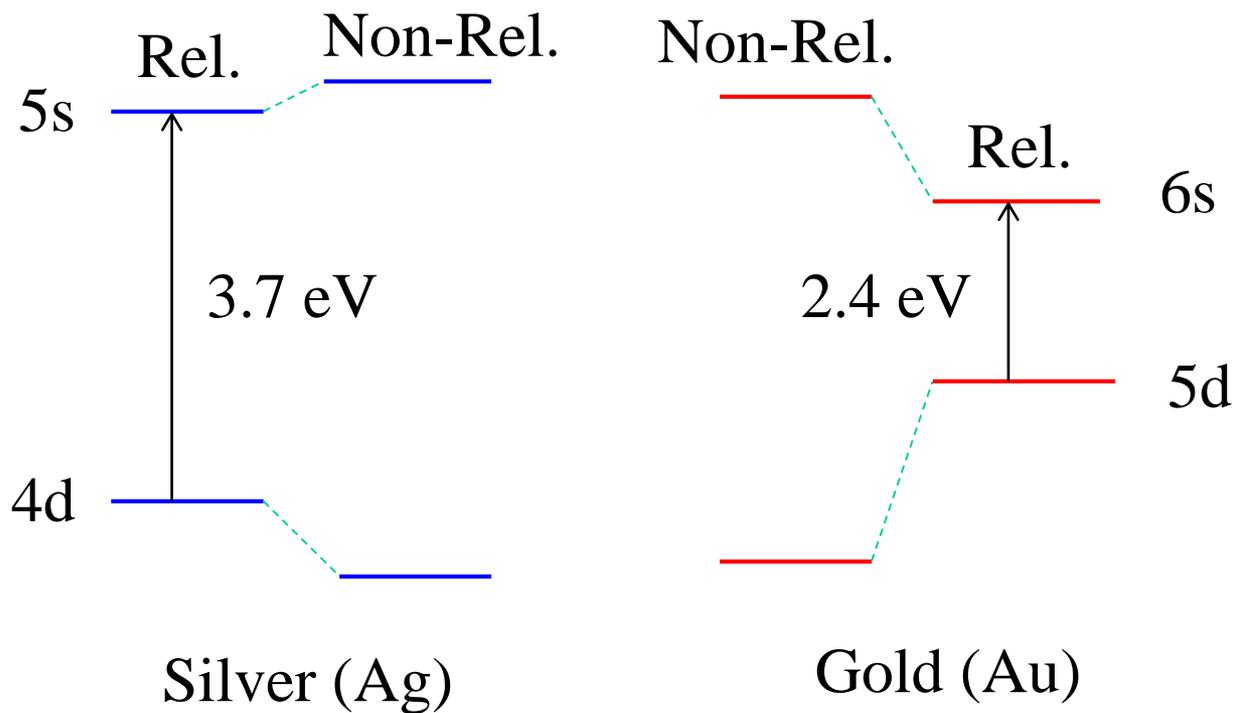


Solution of the Dirac equation (relativistic quantum mechanics)  
for a hydrogen-like atom:

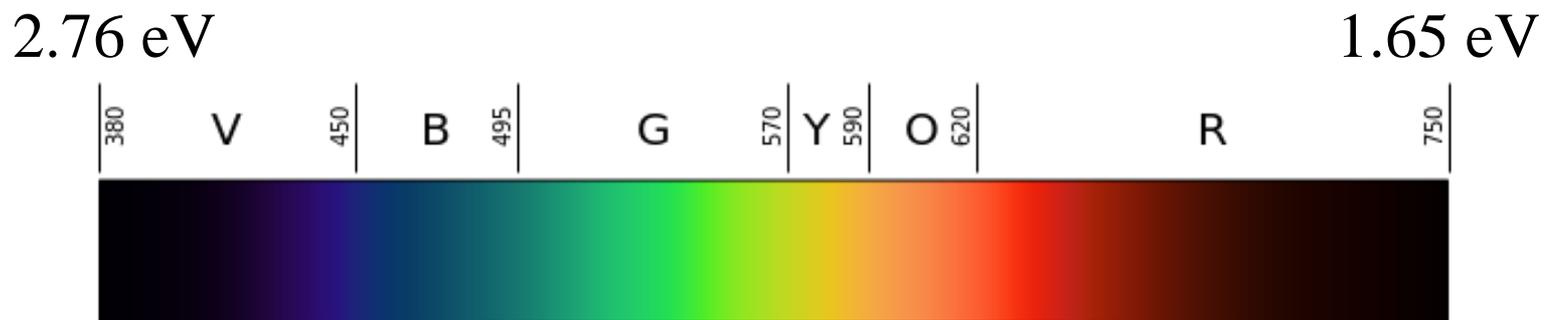
$$E_{1S} = mc^2 \sqrt{1 - (Z\alpha)^2} \sim mc^2 \left( 1 - \frac{(Z\alpha)^2}{2} - \underbrace{\frac{(Z\alpha)^4}{8} + \dots}_{\text{relativistic effect}} \right)$$

relativistic effect



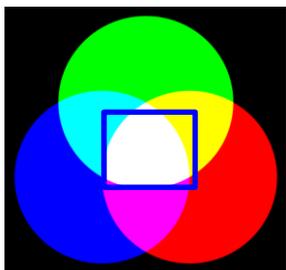


cf. visible spectrum

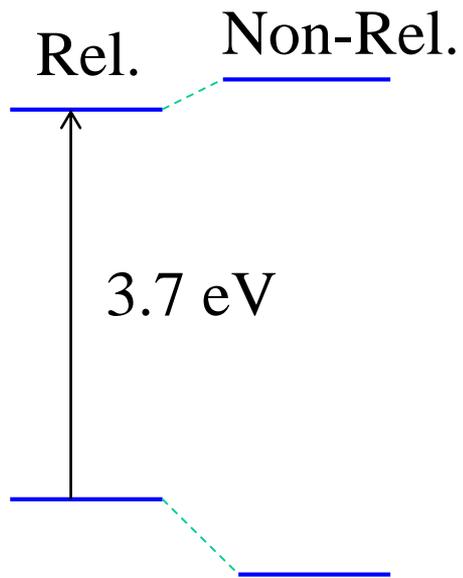


↑  
3.7 eV

↑  
2.4 eV

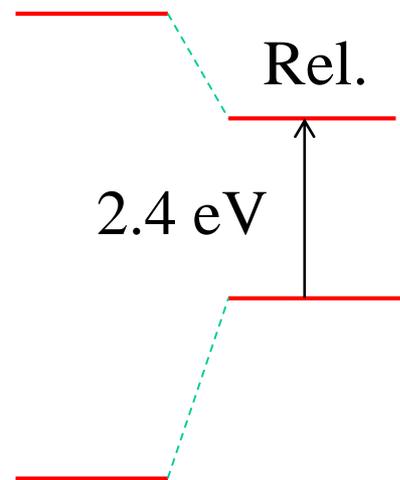


no color  
absorbed

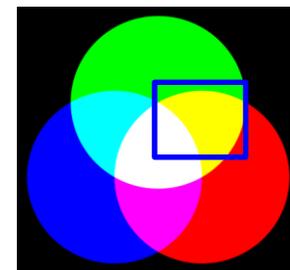


Silver (Ag)

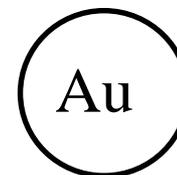
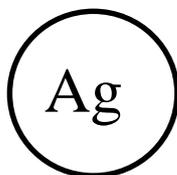
Non-Rel.



Gold (Au)



blue: absorbed



# Chemistry of superheavy elements

Group → ↓ Period	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	
1	1 H																		2 He
2	3 Li	4 Be											5 B	6 C	7 N	8 O	9 F	10 Ne	
3	11 Na	12 Mg											13 Al	14 Si	15 P	16 S	17 Cl	18 Ar	
4	19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr	
5	37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe	
6	55 Cs	56 Ba		72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn	
7	87 Fr	88 Ra		104 Rf	105 Db	106 Sq	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Uut	114 Fl	115 Uup	116 Lv	117 Uus	118 Uuo	
Lanthanides			57 La	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu		
Actinides			89 Ac	90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr		

How do the relativistic effects alter the periodic table for SHE?

→ a big open question