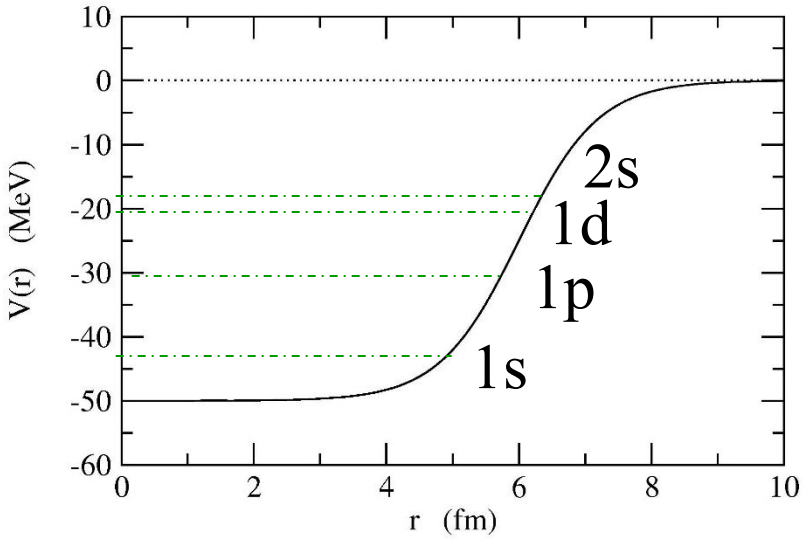
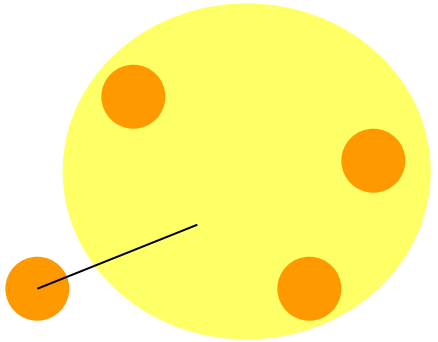
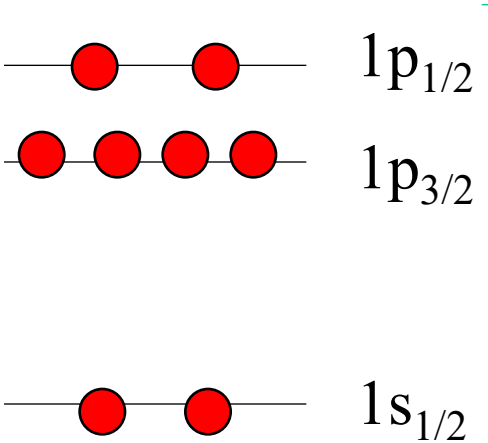
Mean-field (Hartree-Fock) Theory



平均場





shell model

naively speaking,

$$V(\boldsymbol{r}) \sim \int v(\boldsymbol{r}-\boldsymbol{r}')\rho(\boldsymbol{r}')d\boldsymbol{r}'$$

independent motion

$$\rho(\boldsymbol{r}) = \sum_i |\psi_i(\boldsymbol{r})|^2$$

Mean-field (Hartree-Fock) Theory

$$0 = \left[-\frac{\hbar^2}{2m}\nabla^2 + \int v(\boldsymbol{r}-\boldsymbol{r}') \left(\sum_j |\psi_j(\boldsymbol{r}')|^2 \right) d\boldsymbol{r}' - \epsilon_i \right] \psi_i(\boldsymbol{r})$$

the potential depends on the solutions

→ self-consistent solutions

Iteration: $\{\psi_i\} \to \rho \to V \to \{\psi_i\} \to \cdots$

自己無撞着（むどうちゃく）な解

anti-symmetrization

nucleon: fermion

$$\Psi(x_1, x_2, x_3 \cdots) = -\Psi(x_2, x_1, x_3 \cdots)$$

$$\psi_1(x_1)\psi_2(x_2) \rightarrow \frac{1}{\sqrt{2}}[\psi_1(x_1)\psi_2(x_2) - \psi_2(x_1)\psi_1(x_2)]$$

Slater determinat

$$0 = \left[-\frac{\hbar^2}{2m}\nabla^2 + \int v(\boldsymbol{r}-\boldsymbol{r}')\left(\sum_j |\psi_j(\boldsymbol{r}')|^2\right) d\boldsymbol{r}' - \epsilon_i\right]\psi_i(\boldsymbol{r})$$

$$\psi_j^*(\boldsymbol{r}')\psi_j(\boldsymbol{r}')\psi_i(\boldsymbol{r}) \rightarrow \psi_j^*(\boldsymbol{r}')\psi_i(\boldsymbol{r}')\psi_j(\boldsymbol{r})$$

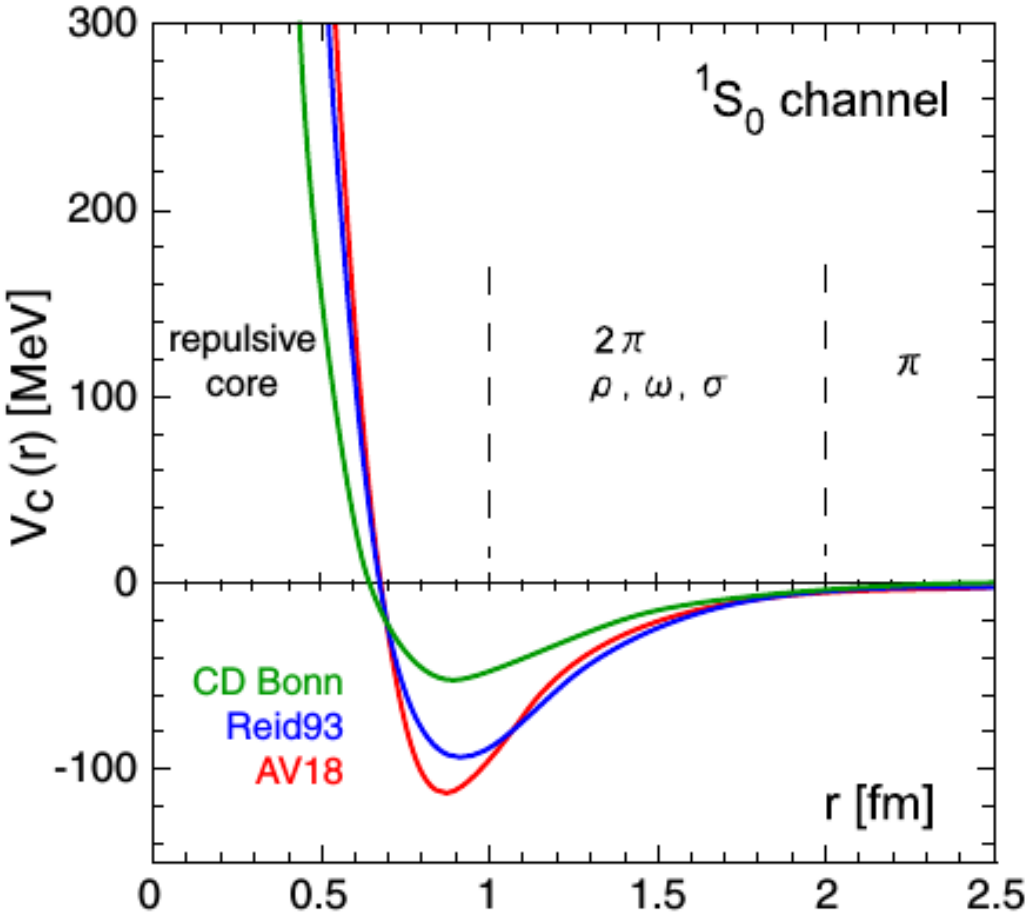
$$\rightarrow \left[-\frac{\hbar^2}{2m}\nabla^2 + \int v(\boldsymbol{r}-\boldsymbol{r}')\left(\sum_j |\psi_j(\boldsymbol{r}')|^2\right) d\boldsymbol{r}' - \epsilon_i\right]\psi_i(\boldsymbol{r})$$

$$- \int v(\boldsymbol{r}-\boldsymbol{r}')\left(\sum_j \psi_j^*(\boldsymbol{r}')\psi_i(\boldsymbol{r}')\right) d\boldsymbol{r}'\psi_j(\boldsymbol{r})$$

exchange term

Hartree-Fock theory

Bare nucleon-nucleon interaction



N. Ishii, S. Aoki, and T. Hatsuda,
PRL99, 022001 (2007)

Bruckner's G-matrix

Nucleon-nucleon interaction *in medium*

Nucleon-nucleon interaction with a hard core

⟹ HF method: does not work

⟵ Matrix elements: diverge

…..but the HF picture seems to work in nuclear systems

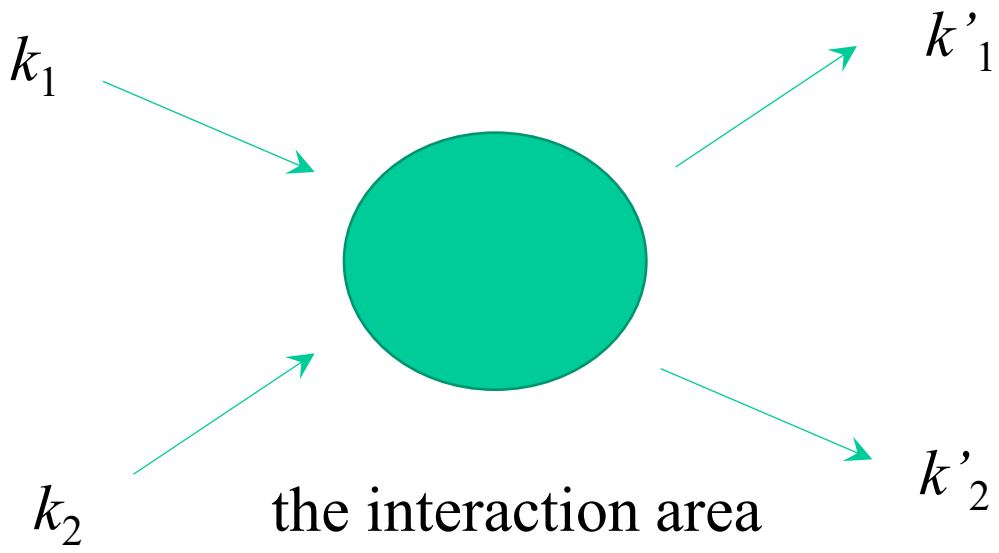
cf. magic numbers

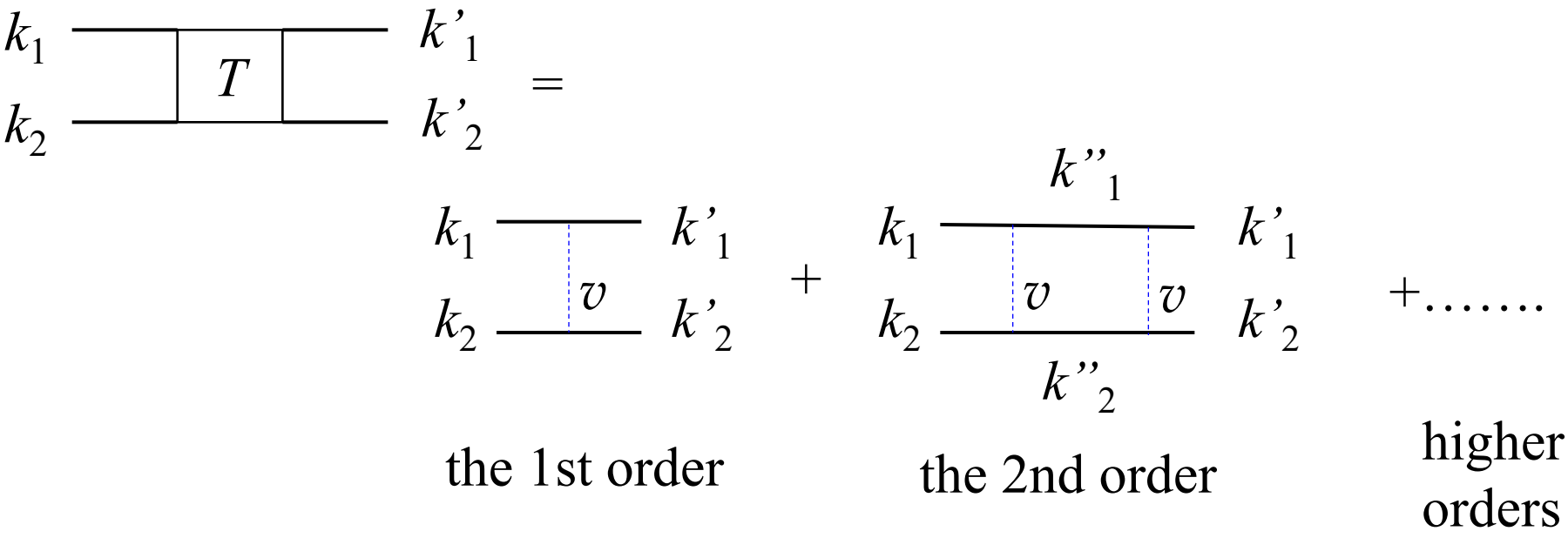
Solution: a nucleon-nucleon interaction *in medium* (effective interaction) rather than a bare interaction

⇑

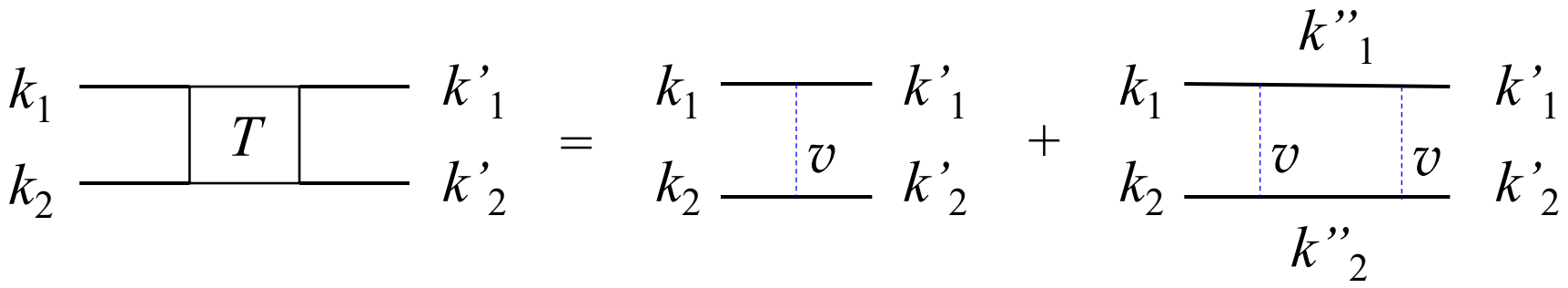
Bruckner's G-matrix

➢ Firstly, two-body (multiple) scattering *in vacuum*





➢ Firstly, two-body (multiple) scattering *in vacuum*



+.....

Lippmann-Schwinger equation

$$T = v + v\frac{1}{E - H_0}T$$

$$\left(-\frac{\hbar^2}{2m}\nabla^2 + V - E\right)\psi = 0$$

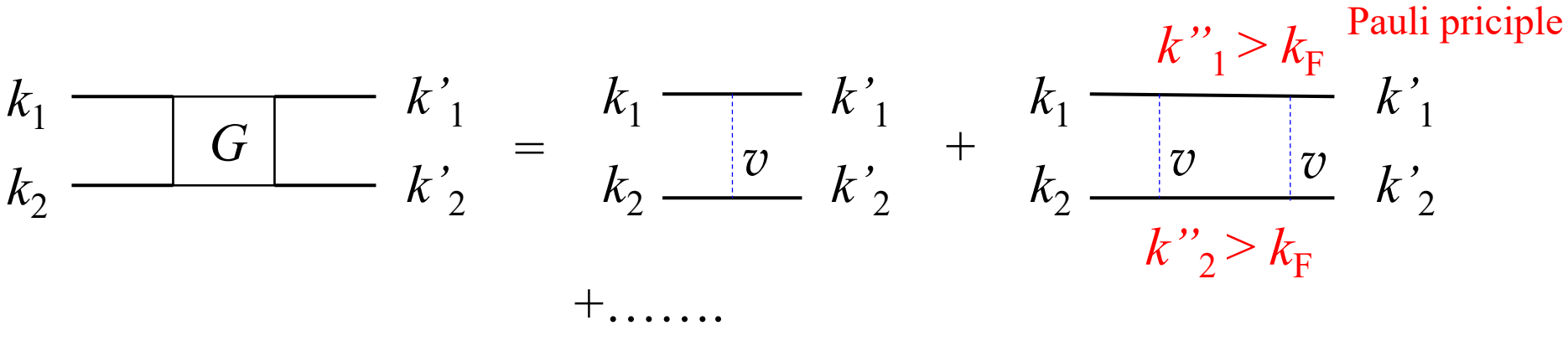
$$\Rightarrow \left(-\frac{\hbar^2}{2m}\nabla^2 - E\right)\psi = -V\psi$$

$$\Rightarrow \psi = \phi - \frac{1}{H_0 - E}V\psi \qquad H_0 = -\frac{\hbar^2}{2m}\nabla^2, \quad (H_0 - E)\phi = 0$$

$$\Rightarrow V\psi = V\phi - V\frac{1}{H_0 - E}V\psi \quad \Rightarrow \quad T = V - V\frac{1}{H_0 - E}T \qquad (V\psi = T\phi)$$

核内における核子間相互作用（媒質効果）

➢ two-body (multiple) scattering *in medium*



Bethe-Goldstone equation

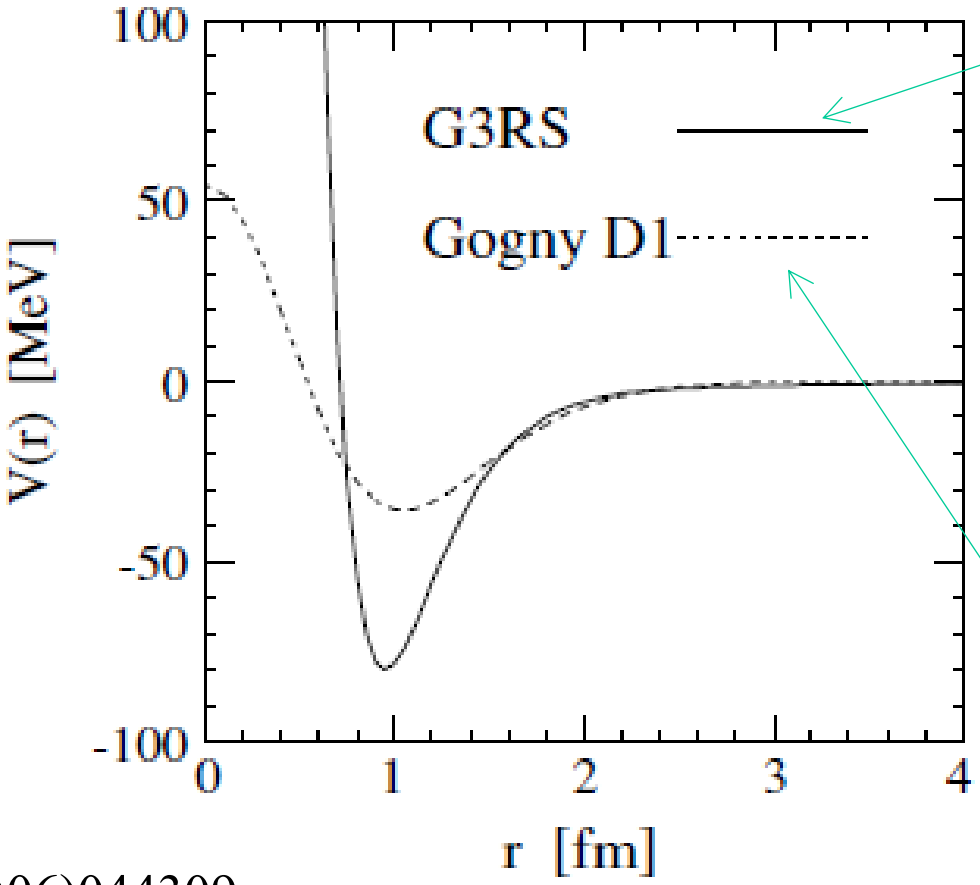
$$G = v + v\frac{Q_F}{E - H_0}G$$

*scattering: suppressed
because intermediate states have to have
$k > k_F$ → independent particle picture

◆Hard core

$$G = v + v\frac{Q_F}{E - H_0}G \quad \Longleftrightarrow \quad G = \frac{v}{1 - v\frac{Q_F}{E - H_0}}$$

Even if v tends to infinity, G may stay finite.



bare interaction

effective interaction

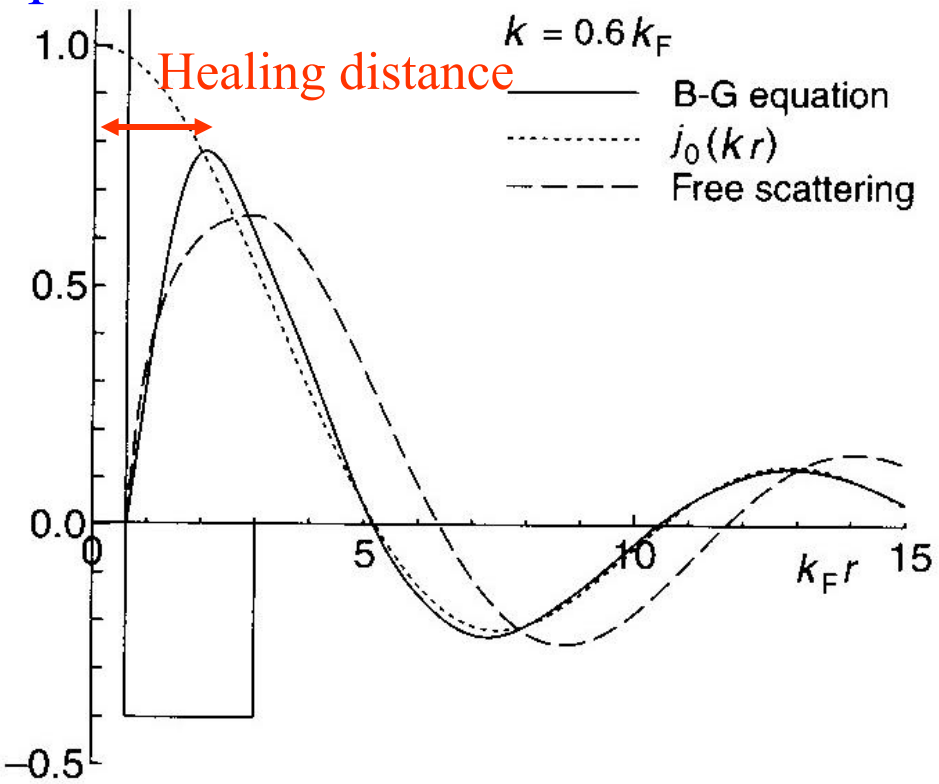
figure from
M. Matsuo,
Phys. Rev. C73(‘06)044309

◆Hard core

$$G = v + v\frac{Q_F}{E - H_0}G \quad \Longleftrightarrow \quad G = \frac{v}{1 - v\frac{Q_F}{E - H_0}}$$

Even if v tends to infinity, G may stay finite.

◆Independent particle motion

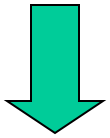


→ use G instead of v in mean-field calculations

Phenomenological effective interactions

G-matrix

- ab initio
- but, cumbersome to compute (especially for finite nuclei)
- qualitatively good, but quantitatively not successful



HF calculations with a phenomenological effective interaction

Philosophy: take the functional form of *G*, but determine the parameters phenomenologically

- Skyrme interaction (non-rel., zero range)
- Gogny interaction (non-rel., finite range)
- Relativistic mean-field model (relativistic, "meson exchanges")

Skyrme interaction density dependent zero-range interaction

$$
\begin{aligned}
v(\boldsymbol{r},\boldsymbol{r}') \;=\; & t_0(1+x_0\hat{P}_\sigma)\delta(\boldsymbol{r}-\boldsymbol{r}') \\
& +\frac{1}{2}t_1(1+x_1\hat{P}_\sigma)(\boldsymbol{k}^2\delta(\boldsymbol{r}-\boldsymbol{r}')+\delta(\boldsymbol{r}-\boldsymbol{r}')\boldsymbol{k}^2) \\
& +t_2(1+x_2\hat{P}_\sigma)\boldsymbol{k}\delta(\boldsymbol{r}-\boldsymbol{r}')\boldsymbol{k} \\
& +\frac{1}{6}t_3(1+x_3\hat{P}_\sigma)\delta(\boldsymbol{r}-\boldsymbol{r}')\rho^\alpha((\boldsymbol{r}+\boldsymbol{r}')/2) \\
& +iW_0(\boldsymbol{\sigma}_1+\boldsymbol{\sigma}_2)\cdot\boldsymbol{k}\times\delta(\boldsymbol{r}-\boldsymbol{r}')\boldsymbol{k}
\end{aligned}
$$

$$\boldsymbol{k}=(\boldsymbol{\nabla}_1-\boldsymbol{\nabla}_2)/2i$$

if x_i=0, t_1=t_2=0:

$$v(\boldsymbol{r},\boldsymbol{r}') \;=\; \underbrace{t_0\delta(\boldsymbol{r}-\boldsymbol{r}')}_{\text{short-range attraction}}+\underbrace{\frac{1}{6}t_3\delta(\boldsymbol{r}-\boldsymbol{r}')\rho^\alpha(\boldsymbol{r})}_{\text{repulsion to avoid collapse}}$$

$$+\underbrace{iW_0(\boldsymbol{\sigma}_1+\boldsymbol{\sigma}_2)\cdot\boldsymbol{k}\times\delta(\boldsymbol{r}-\boldsymbol{r}')\boldsymbol{k}}_{\text{spin-orbit interaction}}$$

Skyrme interaction density dependent zero-range interaction

$$
\begin{aligned}
v(\boldsymbol{r},\boldsymbol{r}') \;=\; & t_0(1+x_0\hat{P}_\sigma)\delta(\boldsymbol{r}-\boldsymbol{r}') \\
& +\frac{1}{2}t_1(1+x_1\hat{P}_\sigma)(\boldsymbol{k}^2\delta(\boldsymbol{r}-\boldsymbol{r}')+\delta(\boldsymbol{r}-\boldsymbol{r}')\boldsymbol{k}^2) \\
& +t_2(1+x_2\hat{P}_\sigma)\boldsymbol{k}\delta(\boldsymbol{r}-\boldsymbol{r}')\boldsymbol{k} \\
& +\frac{1}{6}t_3(1+x_3\hat{P}_\sigma)\delta(\boldsymbol{r}-\boldsymbol{r}')\rho^\alpha((\boldsymbol{r}+\boldsymbol{r}')/2) \\
& +iW_0(\boldsymbol{\sigma}_1+\boldsymbol{\sigma}_2)\cdot\boldsymbol{k}\times\delta(\boldsymbol{r}-\boldsymbol{r}')\boldsymbol{k}
\end{aligned}
$$

$$\boldsymbol{k}=(\boldsymbol{\nabla}_1-\boldsymbol{\nabla}_2)/2i$$

(note) finite range effect $\Longleftrightarrow$ momentum dependence

$$
\begin{aligned}
\langle \boldsymbol{p}|V|\boldsymbol{p}'\rangle \;=\; & \frac{1}{(2\pi\hbar)^3}\int d\boldsymbol{r}\, e^{-i(\boldsymbol{p}-\boldsymbol{p}')\cdot\boldsymbol{r}/\hbar}V(\boldsymbol{r}) \\
\sim\; & V_0+V_1(\boldsymbol{p}^2+\boldsymbol{p}'^2)+V_2\boldsymbol{p}\boldsymbol{p}'+\cdots \\
\rightarrow\; & V_0\delta(\boldsymbol{r})+V_1(\hat{\boldsymbol{p}}^2\delta(\boldsymbol{r})+\delta(\boldsymbol{r})\hat{\boldsymbol{p}}^2)+V_2\,\hat{\boldsymbol{p}}\delta(\boldsymbol{r})\hat{\boldsymbol{p}}
\end{aligned}
$$

Skyrme interaction density dependent zero-range interaction

$$
\begin{aligned}
v(\boldsymbol{r},\boldsymbol{r}') \;=\; & t_0(1+x_0\hat{P}_\sigma)\delta(\boldsymbol{r}-\boldsymbol{r}') \\
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& +iW_0(\boldsymbol{\sigma}_1+\boldsymbol{\sigma}_2)\cdot\boldsymbol{k}\times\delta(\boldsymbol{r}-\boldsymbol{r}')\boldsymbol{k}
\end{aligned}
$$

$$\boldsymbol{k}=(\boldsymbol{\nabla}_1-\boldsymbol{\nabla}_2)/2i$$

the exchange potential ⟶ local

$$
\begin{aligned}
0 \;=\; & \left[-\frac{\hbar^2}{2m}\boldsymbol{\nabla}^2+\int v(\boldsymbol{r}-\boldsymbol{r}')\left(\sum_j|\psi_j(\boldsymbol{r}')|^2\right)d\boldsymbol{r}'-\epsilon_i\right]\psi_i(\boldsymbol{r}) \\
& -\int v(\boldsymbol{r}-\boldsymbol{r}')\left(\sum_j\psi_j^*(\boldsymbol{r}')\psi_i(\boldsymbol{r}')\right)d\boldsymbol{r}'\psi_j(\boldsymbol{r})
\end{aligned}
$$

Skyrme interactions: 10 adjustable parameters

$$
\begin{aligned}
v(\boldsymbol{r},\boldsymbol{r}') \;=\; & t_0(1+x_0\hat{P}_\sigma)\delta(\boldsymbol{r}-\boldsymbol{r}') \\
& +\frac{1}{2}t_1(1+x_1\hat{P}_\sigma)(\boldsymbol{k}^2\delta(\boldsymbol{r}-\boldsymbol{r}')+\delta(\boldsymbol{r}-\boldsymbol{r}')\boldsymbol{k}^2) \\
& +t_2(1+x_2\hat{P}_\sigma)\boldsymbol{k}\delta(\boldsymbol{r}-\boldsymbol{r}')\boldsymbol{k} \\
& +\frac{1}{6}t_3(1+x_3\hat{P}_\sigma)\delta(\boldsymbol{r}-\boldsymbol{r}')\rho^\alpha((\boldsymbol{r}+\boldsymbol{r}')/2) \\
& +iW_0(\boldsymbol{\sigma}_1+\boldsymbol{\sigma}_2)\cdot\boldsymbol{k}\times\delta(\boldsymbol{r}-\boldsymbol{r}')\boldsymbol{k}
\end{aligned}
$$

A fitting strategy:

B.E. and r_{rms}: ^{16}O, ^{40}Ca, ^{48}Ca, ^{56}Ni, ^{90}Zr, ^{208}Pb,…..
Infinite nuclear matter: E/A, ρ_{eq} ,….

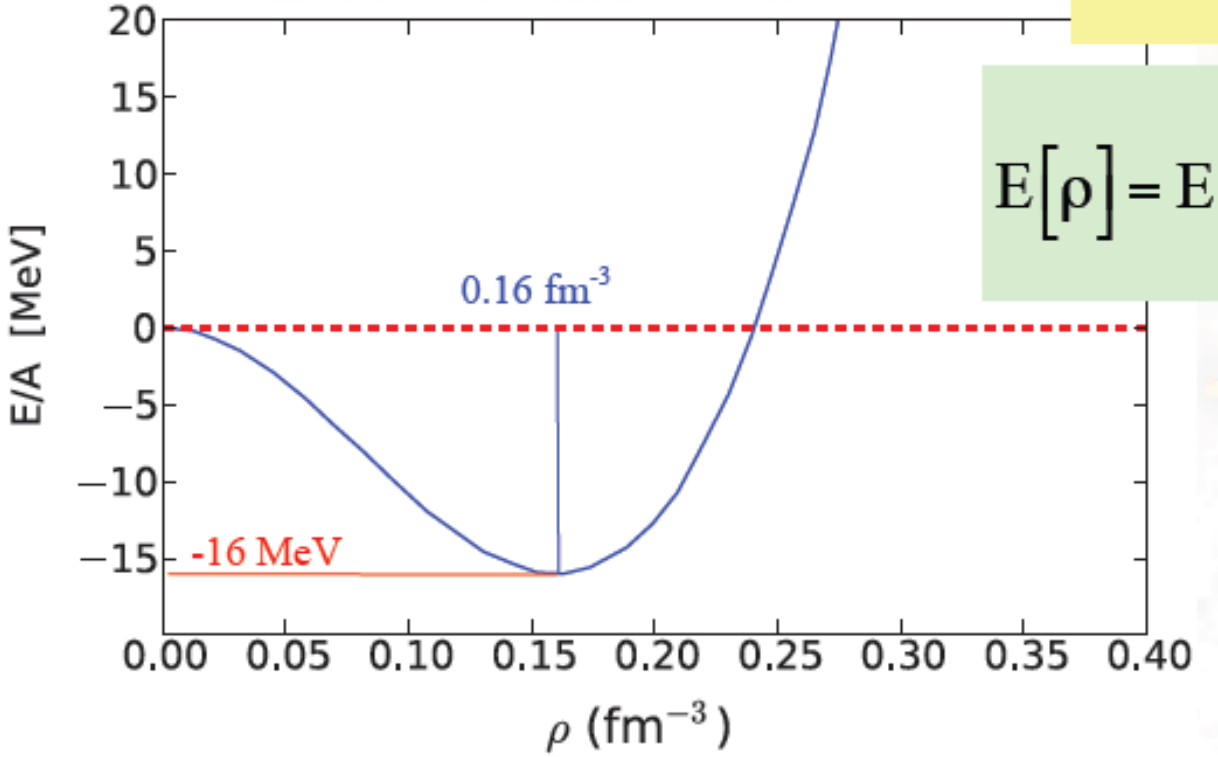
Parameter sets:

SIII, SkM*, SGII, SLy4,…….

EOS of infinite nuclear matter

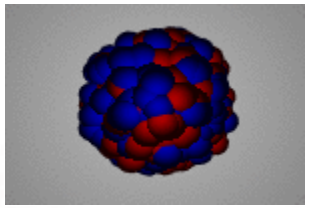
$$K_\infty = 9\rho^2 \left.\frac{d^2\left[E(\rho)/\rho\right]}{d\rho^2}\right|_{\rho_0}$$

$$E[\rho] = E[\rho_0] + \frac{1}{18} K_\infty \left(\frac{\rho-\rho_0}{\rho_0}\right)^2$$



slide: Carlos Bertulani

EOS of infinite nuclear matter



原子核がどのくらい硬いのか

$$K_\infty = 9\rho^2 \left.\frac{d^2\left[E(\rho)/\rho\right]}{d\rho^2}\right|_{\rho_0}$$

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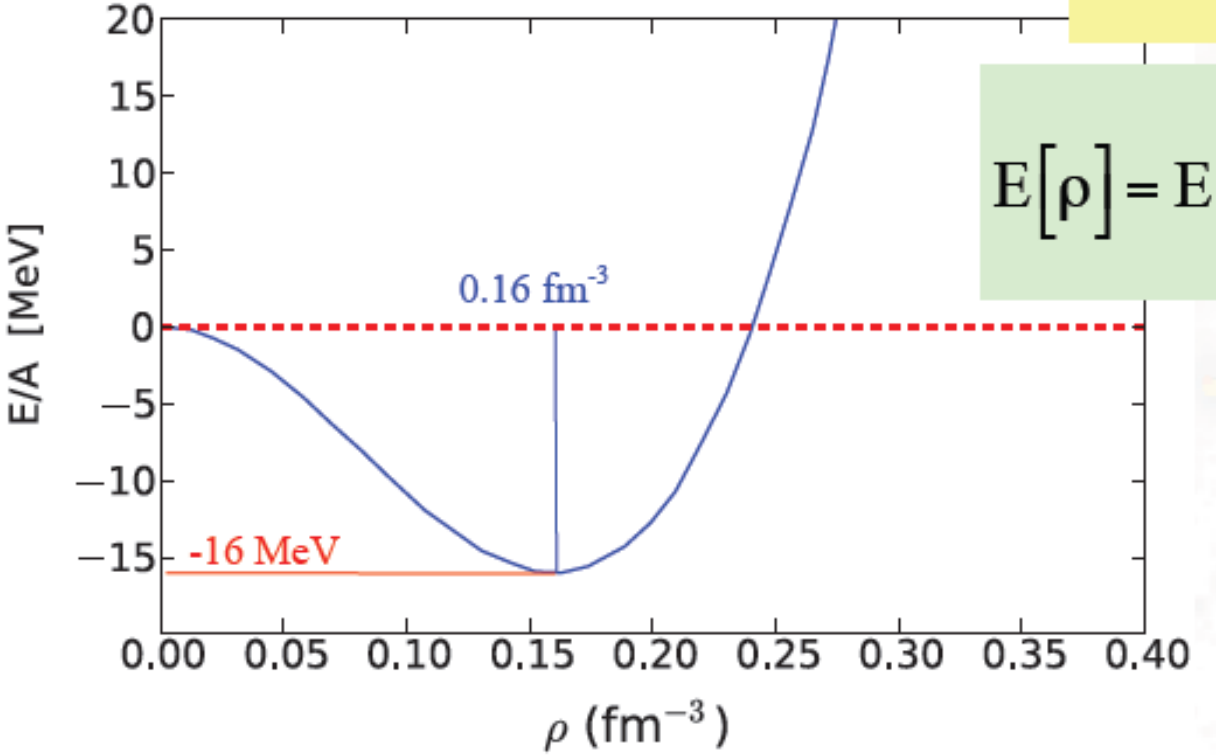
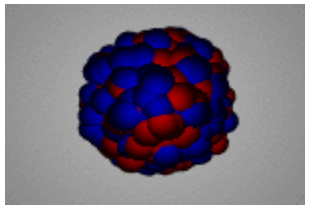


cf. 中性子星の大きさや重さ（MR曲線）

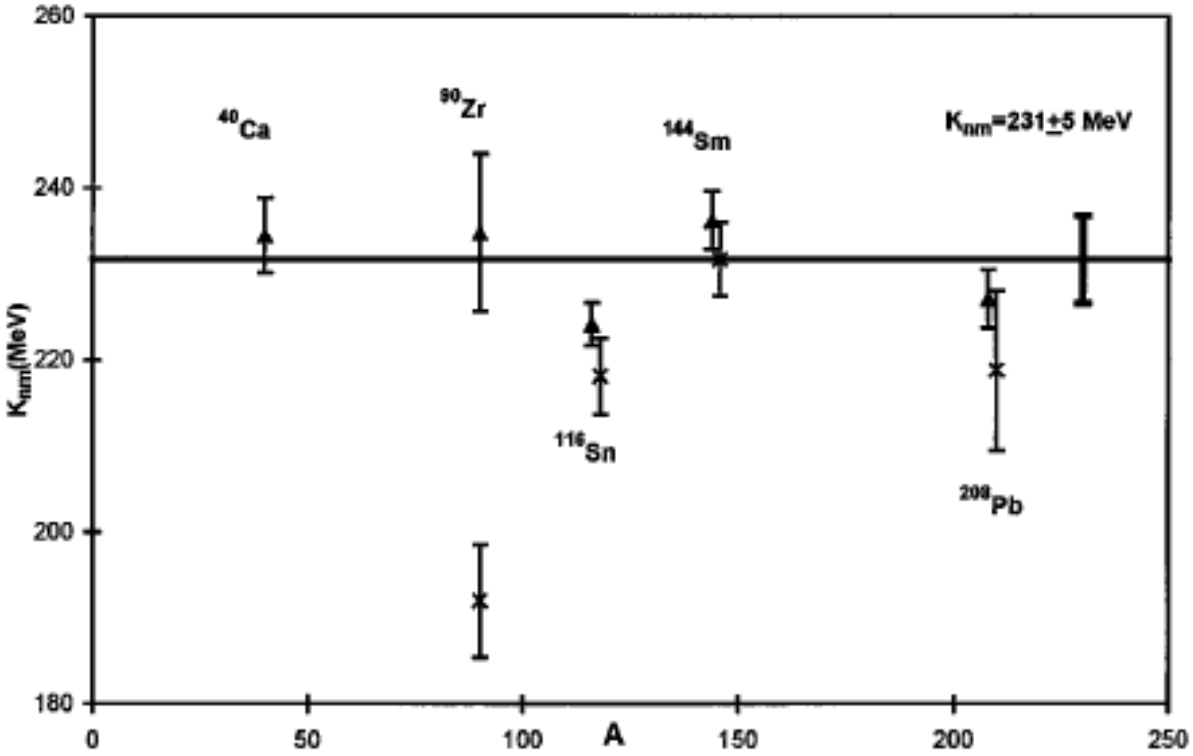
slide: Carlos Bertulani



Isoscalar giant monopole resonances (breathing mode)

$$E_{\mathrm{ISGMR}} \sim \sqrt{\frac{\hbar^2 K}{m\langle r^2 \rangle}}$$

J.P. Blaizot,
Phys. Rep. 64 (‘80) 171



$K \sim 231$ +/- 5 MeV

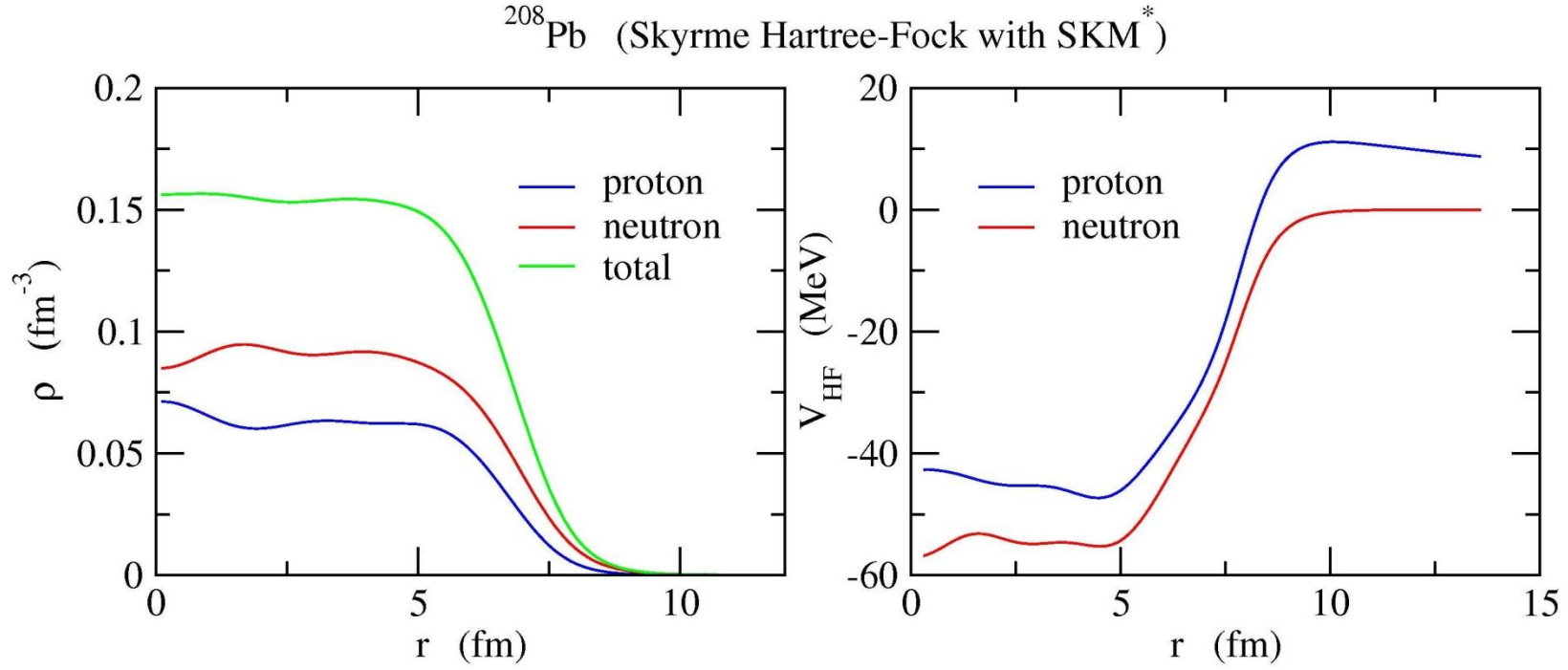
D.H. Youngblood, H.L. Clark, and Y.-W. Lui, PRL82 (‘99) 691

$$-\frac{\hbar^2}{2m}\nabla^2\psi_i(\boldsymbol{r}) + \int v(\boldsymbol{r},\boldsymbol{r}')\rho_{\mathsf{HF}}(\boldsymbol{r}')d\boldsymbol{r}'\psi_i(\boldsymbol{r}) - \int \rho_{\mathsf{HF}}(\boldsymbol{r},\boldsymbol{r}')v(\boldsymbol{r},\boldsymbol{r}')\psi_i(\boldsymbol{r}')\,d\boldsymbol{r}' = \epsilon_i\,\psi_i(\boldsymbol{r})$$

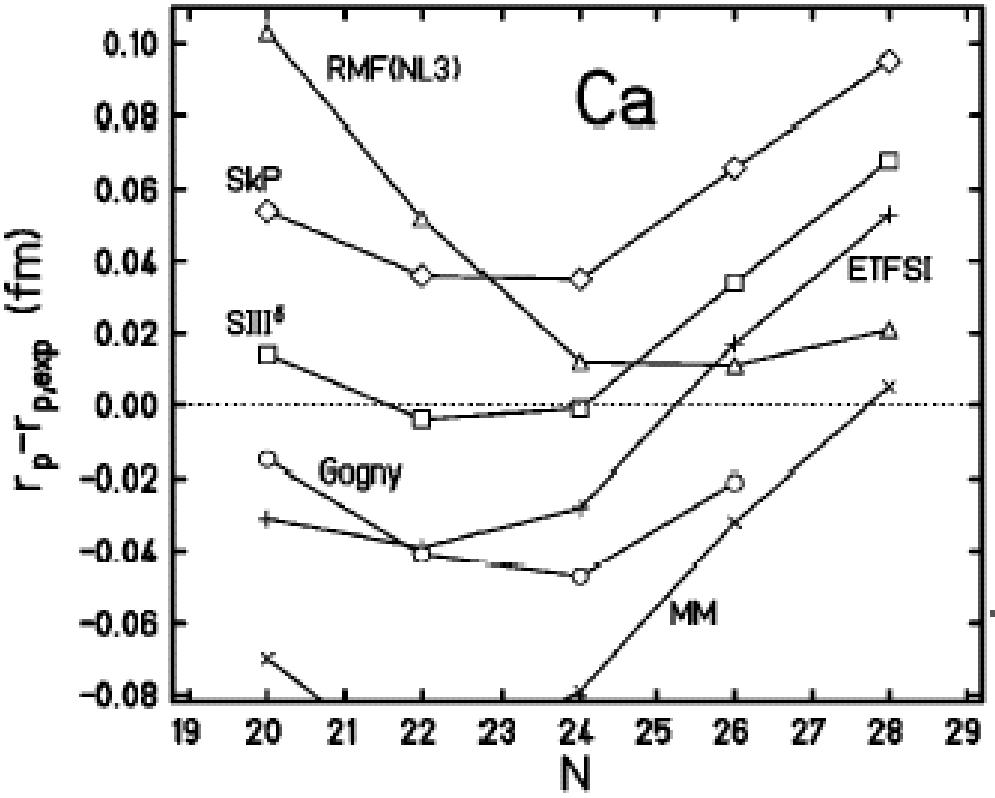
Iteration

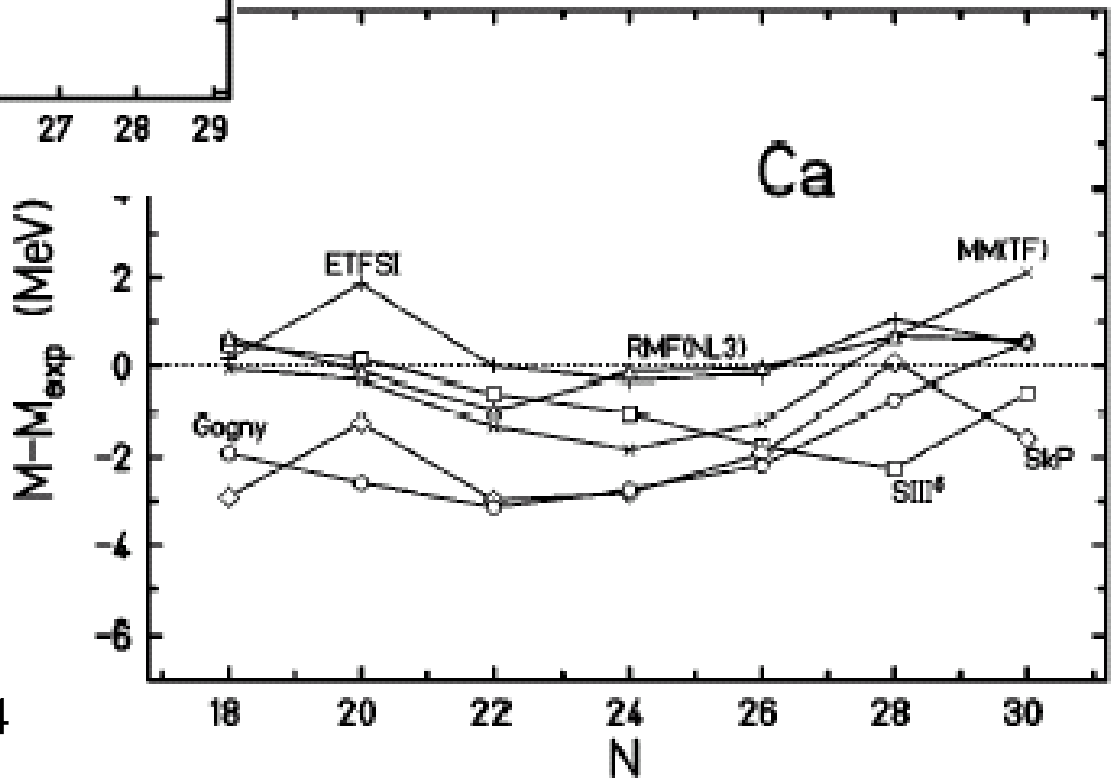
V_{HF}: depends on ψ_i ⟵ non-linear problem

Iteration: $\{\psi_i\} \rightarrow \rho_{\mathsf{HF}} \rightarrow V_{\mathsf{HF}} \rightarrow \{\psi_i\} \rightarrow \cdots$



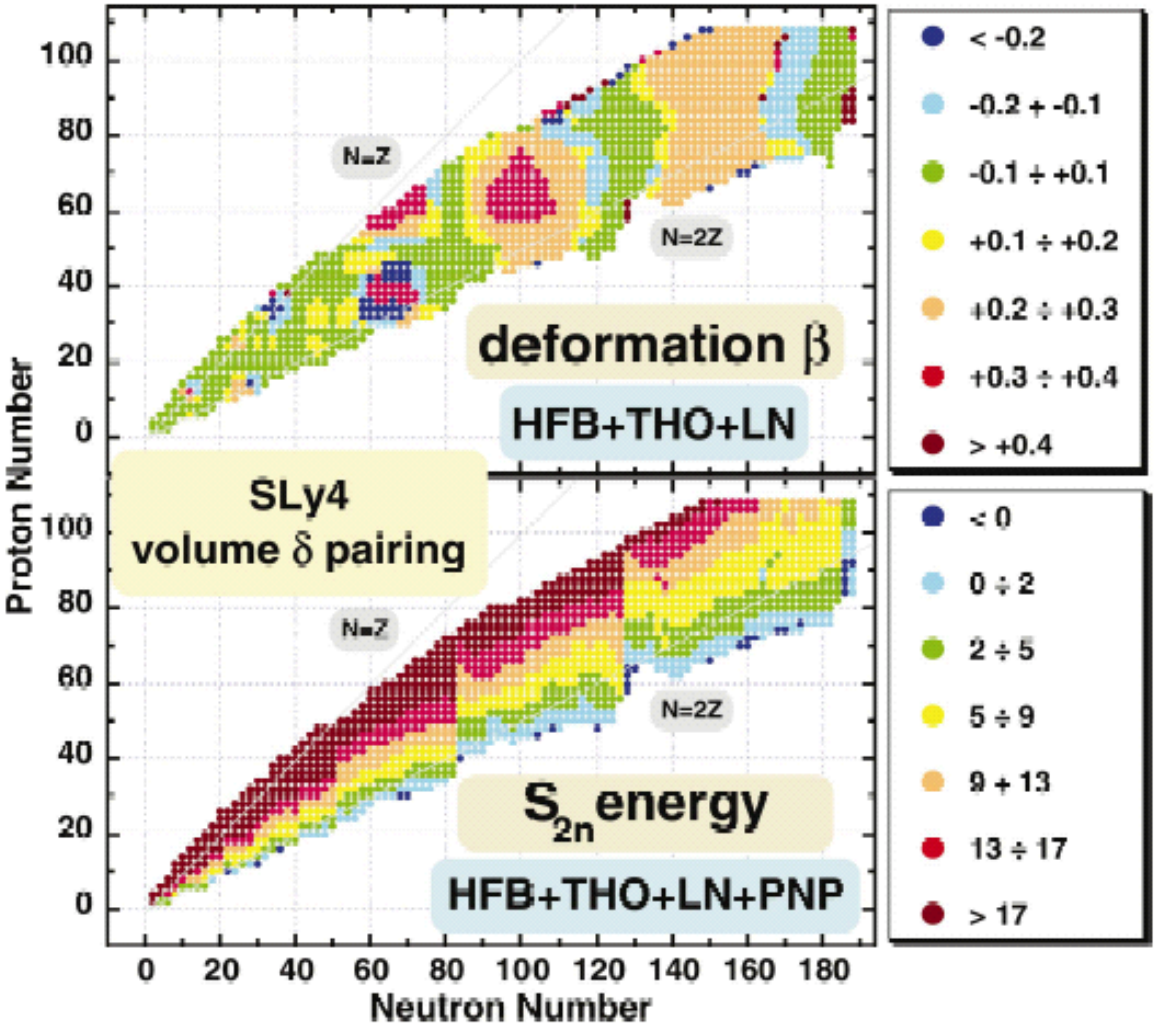
Examples of HF calculations for masses and radii





Z. Patyk et al.,
PRC59(’99)704

deformation and two-neutron separation energy



M.V. Stoitsov et al., PRC68('03)054312

Mean-field approximation and deformation

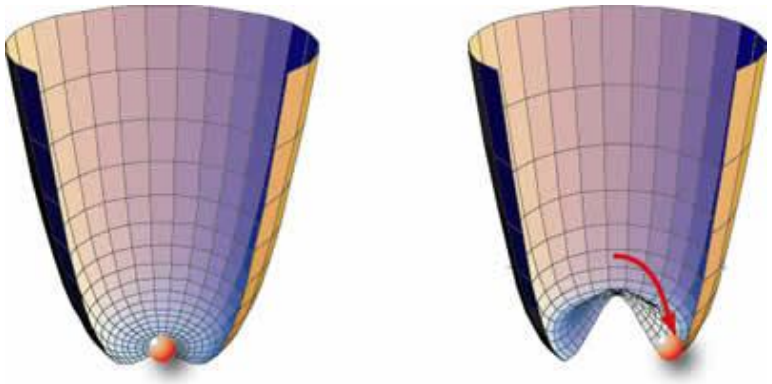
平均場近似＝2体場→1体場に近似

$$H = \sum_{i=1}^{A} -\frac{\hbar^2}{2m}\nabla_i^2 + \frac{1}{2}\sum_{i,j}^{A} v(\boldsymbol{r}_i, \boldsymbol{r}_j)$$

$$= \sum_{i=1}^{A}\left(-\frac{\hbar^2}{2m}\nabla_i^2 + V_{\mathrm{MF}}(\boldsymbol{r}_i)\right) + \frac{1}{2}\sum_{i,j}^{A} v(\boldsymbol{r}_i, \boldsymbol{r}_j) - \sum_i V_{\mathrm{MF}}(\boldsymbol{r}_i)$$

→ Ψ_{MF}：ハミルトニアン H が持っている対称性を持たなくてもいい

“対称性の自発的破れ”



Ψ_{MF} : ハミルトニアン H が持っている対称性を持たなくてもいい

典型的な例

- 並進対称性: 原子核の平均場近似（DFT）では常に破れる

$$H = -\sum_{i=1}^{A} \frac{\hbar^2}{2m}\nabla_i^2 + \frac{1}{2}\sum_{i,j}^{A} v(\boldsymbol{r}_i - \boldsymbol{r}_j) \rightarrow \sum_{i=1}^{A}\left(-\frac{\hbar^2}{2m}\nabla_i^2 + V_{\mathrm{MF}}(\boldsymbol{r}_i)\right)$$

Ψ_{MF} : ハミルトニアン H が持っている対称性を持たなくてもいい

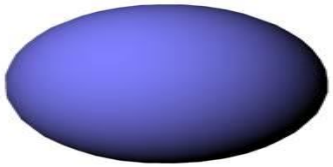
典型的な例

- **並進対称性:** 原子核の平均場近似（DFT）では常に破れる

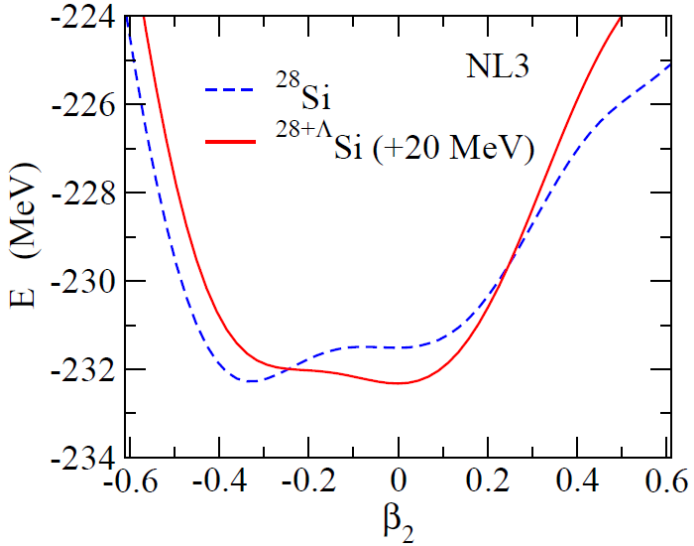
$$H = -\sum_{i=1}^{A}\frac{\hbar^2}{2m}\nabla_i^2 + \frac{1}{2}\sum_{i,j}^{A} v(\boldsymbol{r}_i - \boldsymbol{r}_j) \rightarrow \sum_{i=1}^{A}\left(-\frac{\hbar^2}{2m}\nabla_i^2 + V_{\mathrm{MF}}(\boldsymbol{r}_i)\right)$$

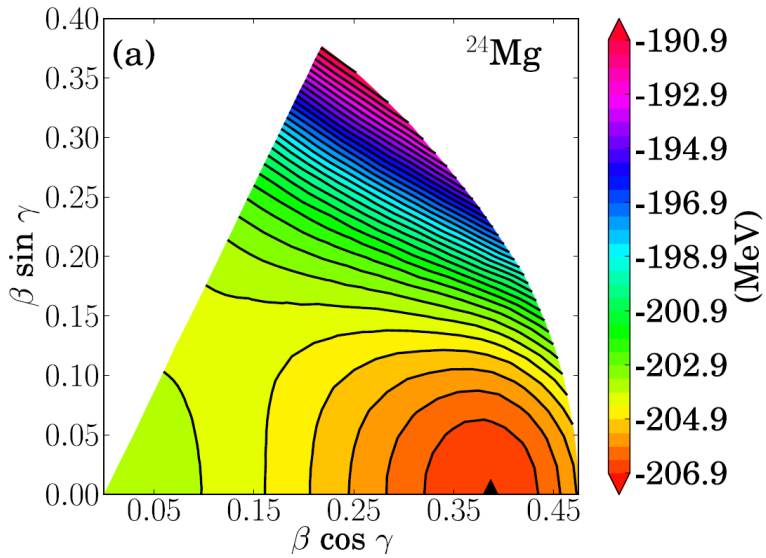
- **回転対称性**

変形した基底状態



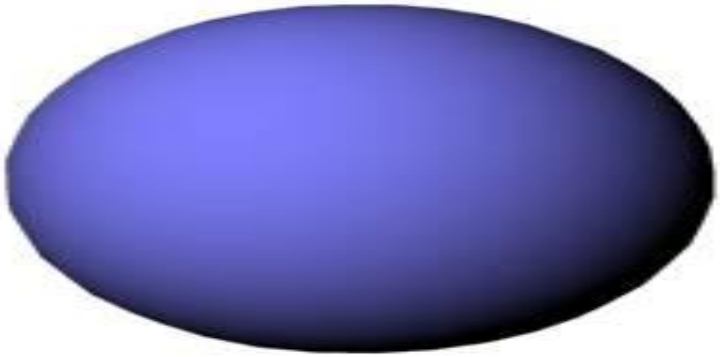
$\rightarrow V(r,\theta)$





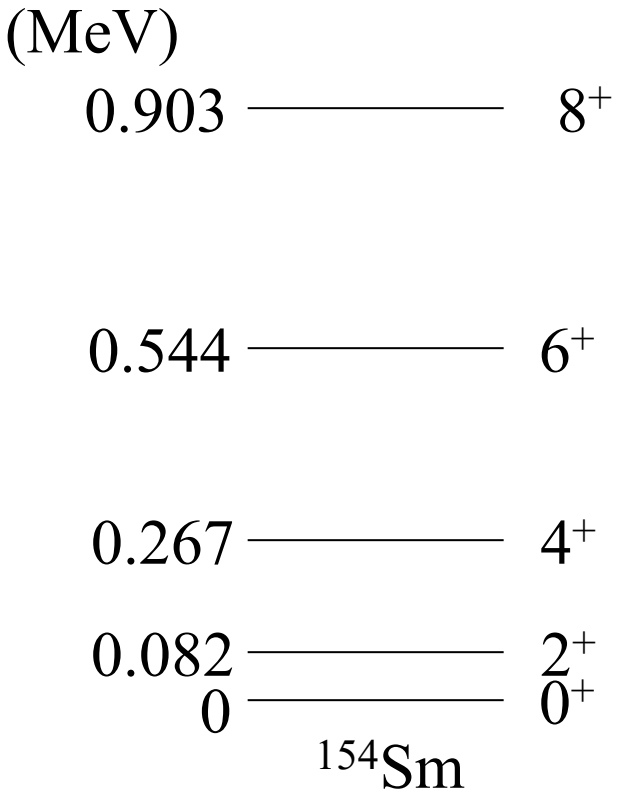
Nuclear Deformation

実験的な証拠

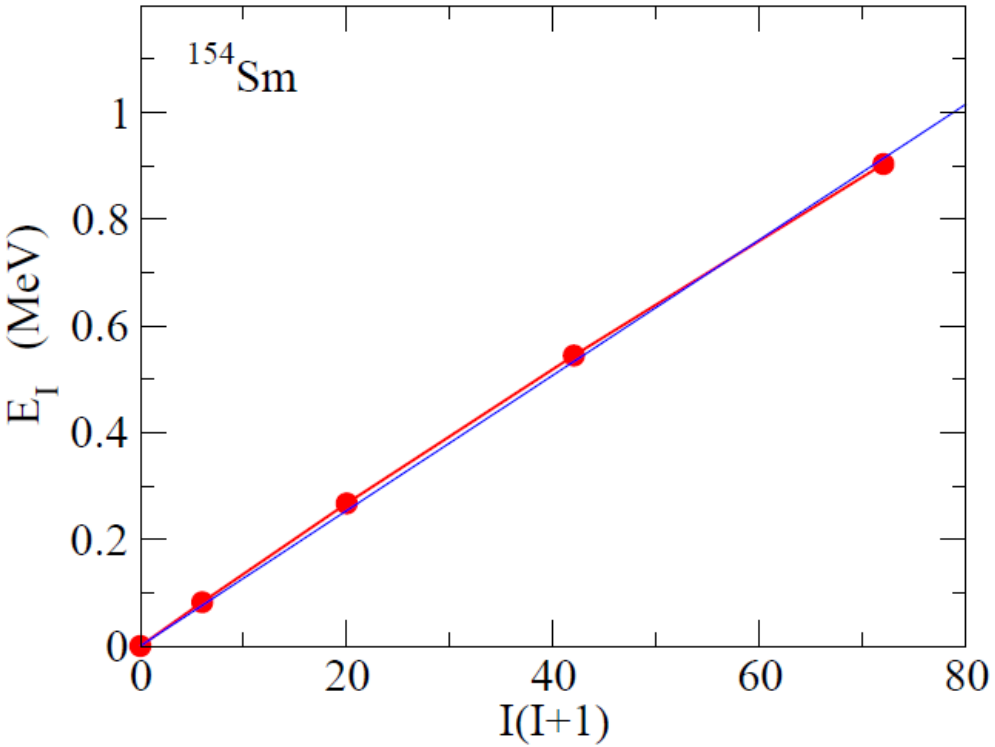


Nuclear Deformation

Excitation spectra of ^{154}Sm



$$E_I \sim \frac{I(I+1)\hbar^2}{2\mathcal{J}}$$



Nuclear Deformation

Excitation spectra of ^{154}Sm

(MeV)

0.903 ——— 8^+

0.544 ——— 6^+

0.267 ——— 4^+

0.082 ——— 2^+

0 ——— 0^+

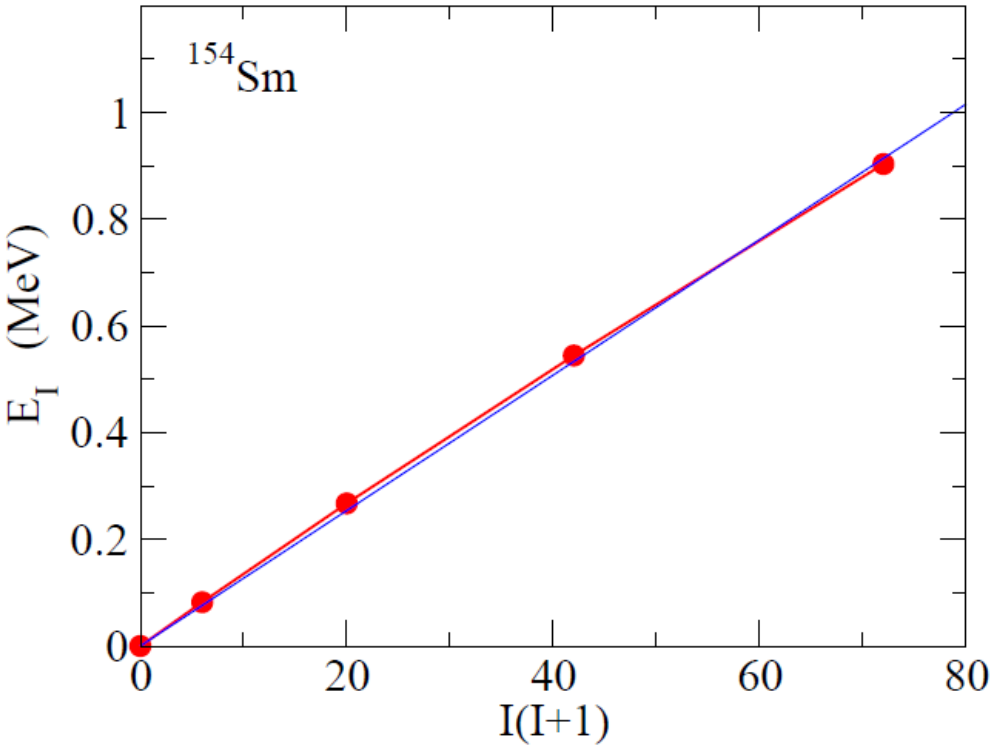
^{154}Sm

$$E_I \sim \frac{I(I+1)\hbar^2}{2\mathcal{J}}$$

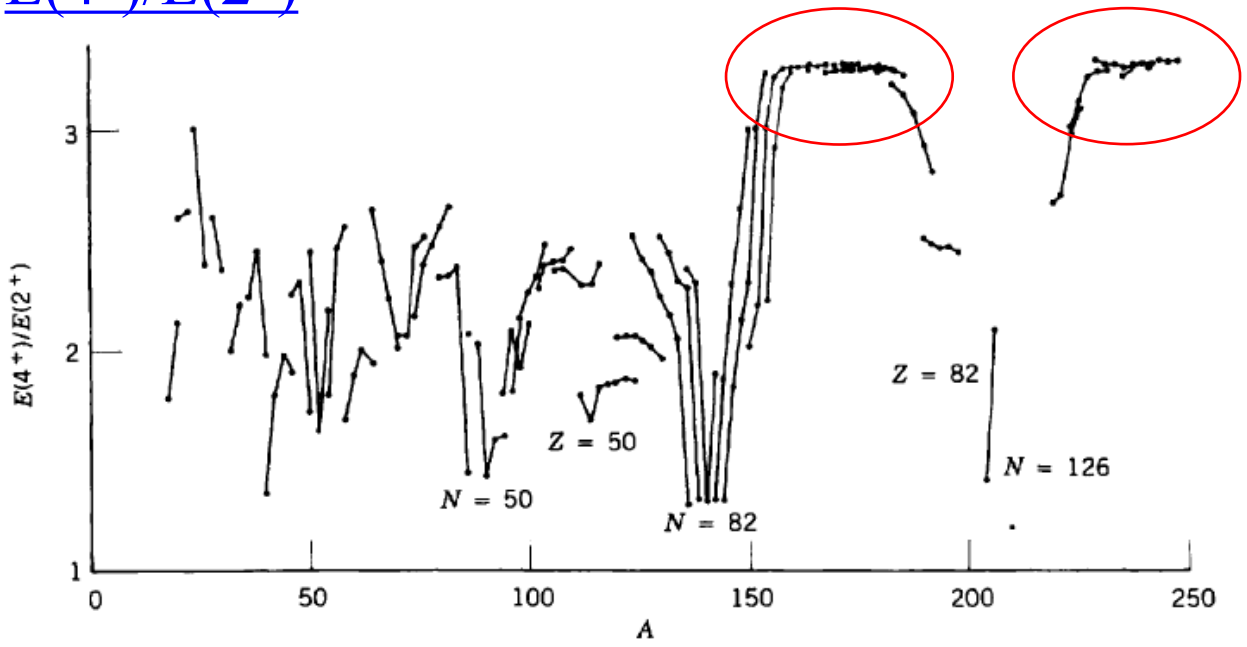
cf. a classical rigid rotor

$$E = \frac{1}{2}\mathcal{J}\omega^2 = \frac{I^2}{2\mathcal{J}}$$

$$(I = \mathcal{J}\omega,\ \omega = \dot{\theta})$$



$E(4^+)/E(2^+)$



deformed nuclei:
$E(4^+)/E(2^+) \sim 3.3$

spherical nuclei:
$E(4^+)/E(2^+) \sim 2$

K.S. Krane, "Introductory Nuclear Physics"

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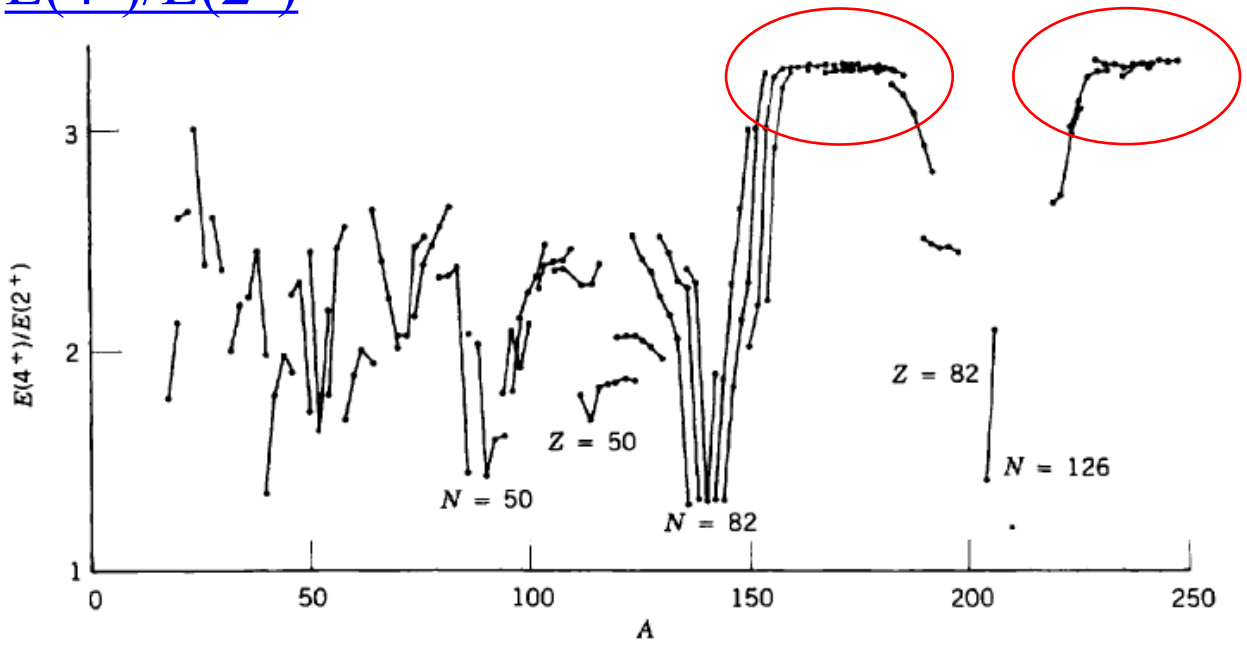

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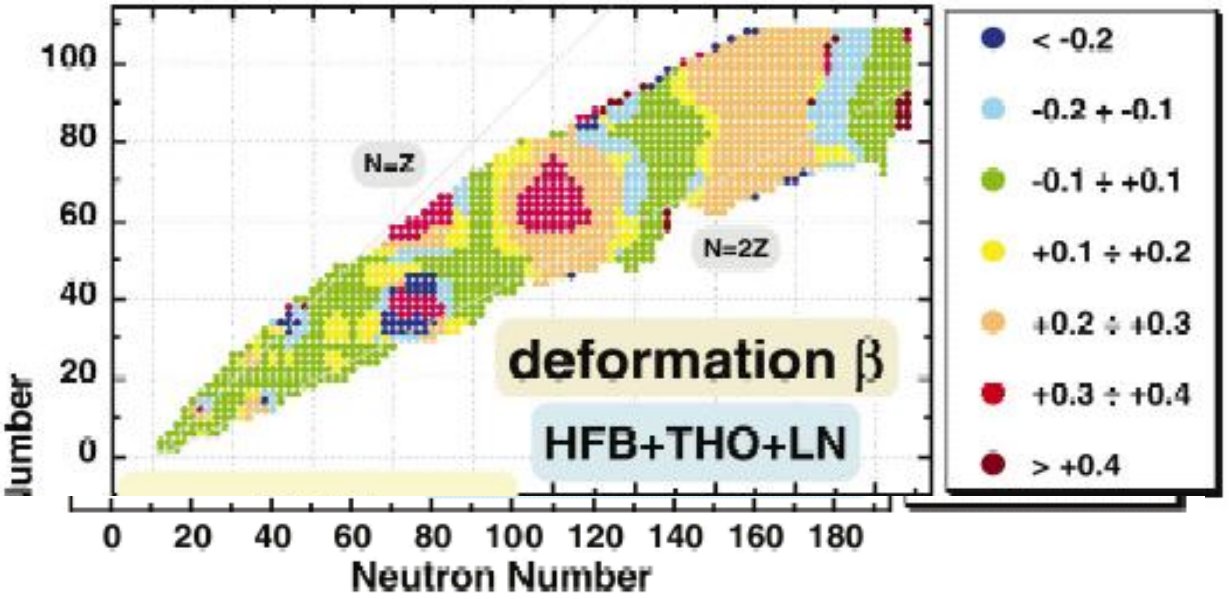
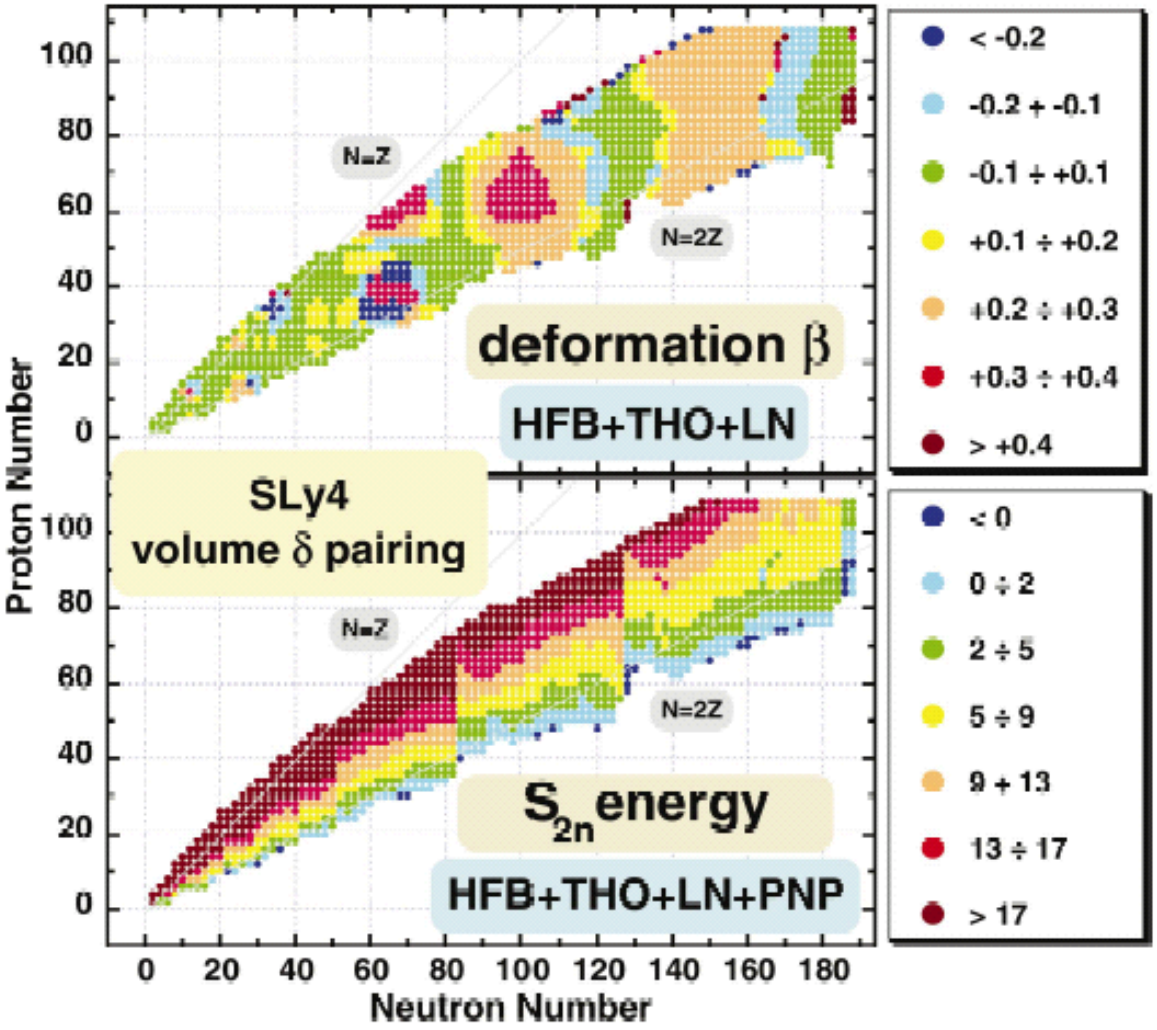
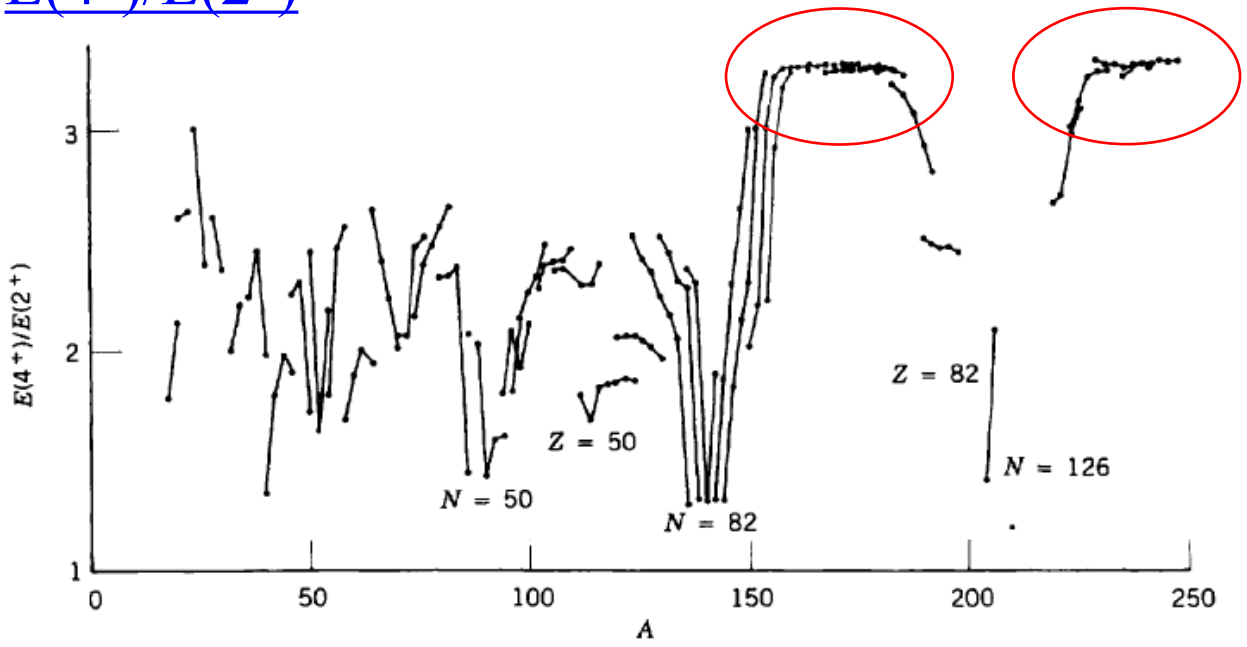


deformation and two-neutron separation energy



M.V. Stoitsov et al., PRC68('03)054312

$E(4^+)/E(2^+)$



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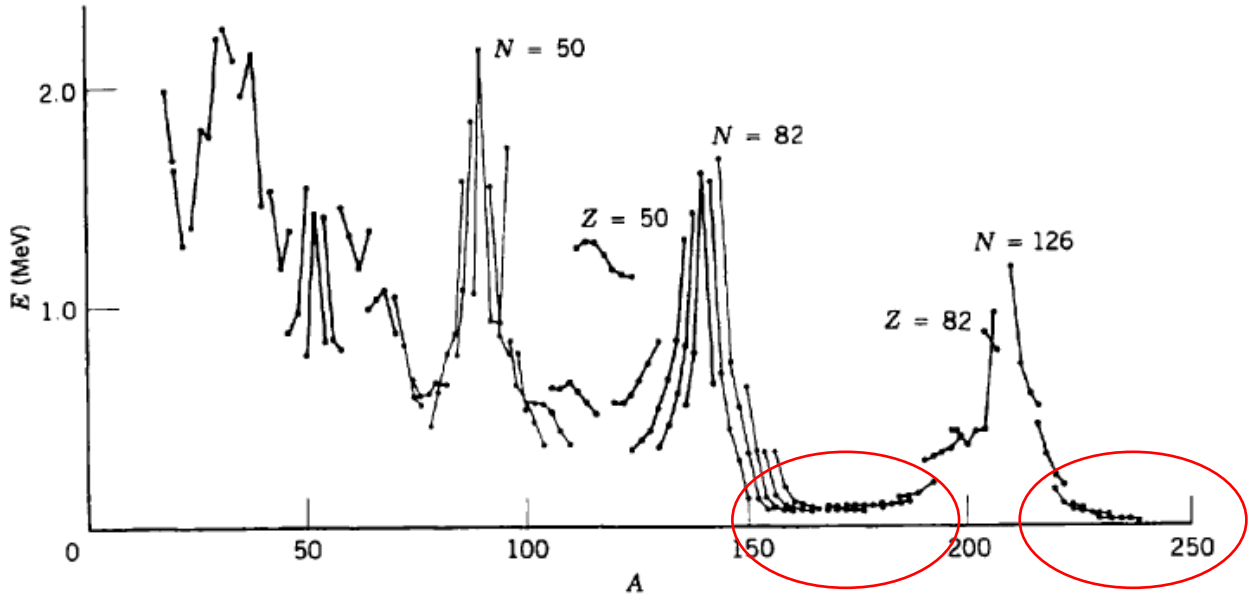
K.S. Krane. "Introductory Nuclear Physics"

$E(2^+)$

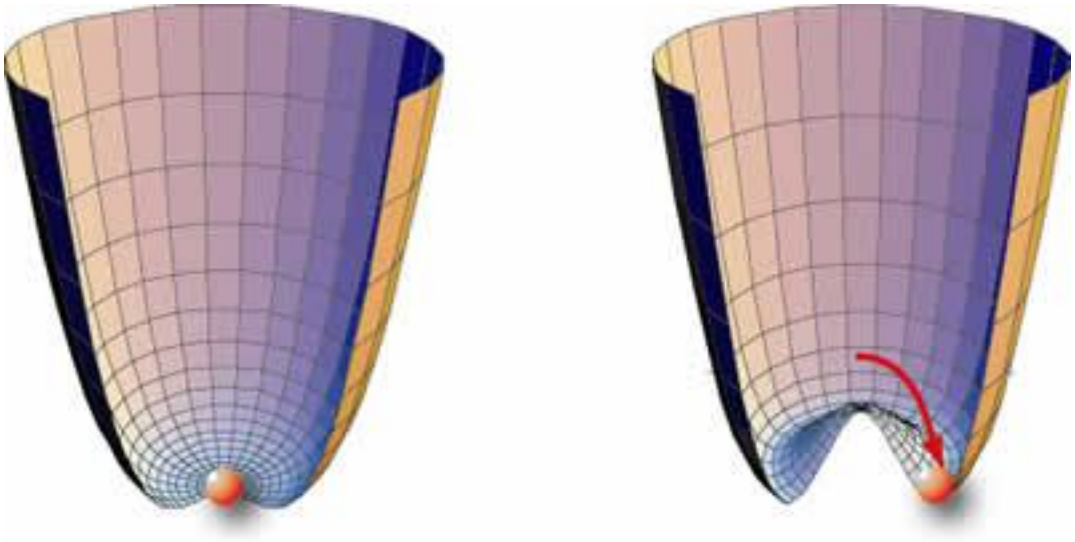


a small energy
→ spontaneously
symm. breaking

deformed nuclei

Spontaneous symmetry breaking

The vacuum state does not have (i.e, the vacuum state violates) the symmetry which the Hamiltonian has.



Nambu-Goldstone mode (zero energy mode) to restore the symmetry

Quiz: spontaneous symmetry breaking

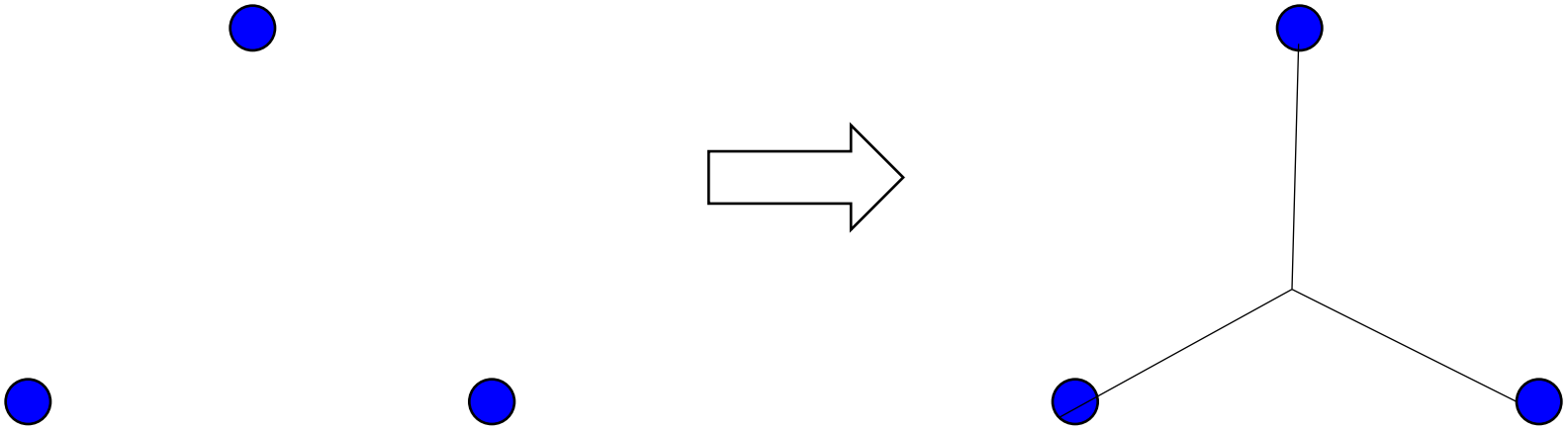
There are a few dots.

- Connect the dots.
- The number of lines is not limited.
- Two lines can cross.
- Connect the dots so that one can go from one dot to all the other dots.

How do you connect the lines if you want to make the total length of lines the shortest?

e.g.) Equilateral triangle

Connect symmetrically



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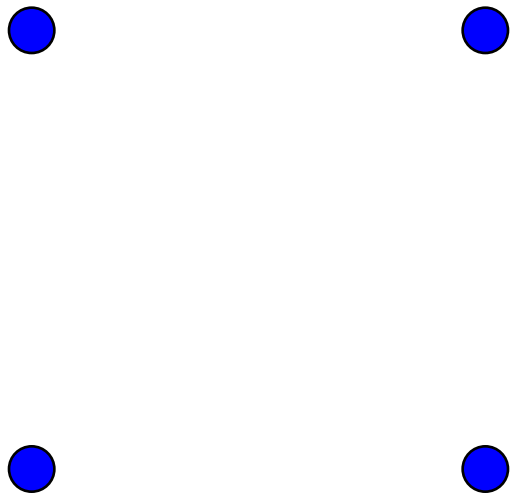
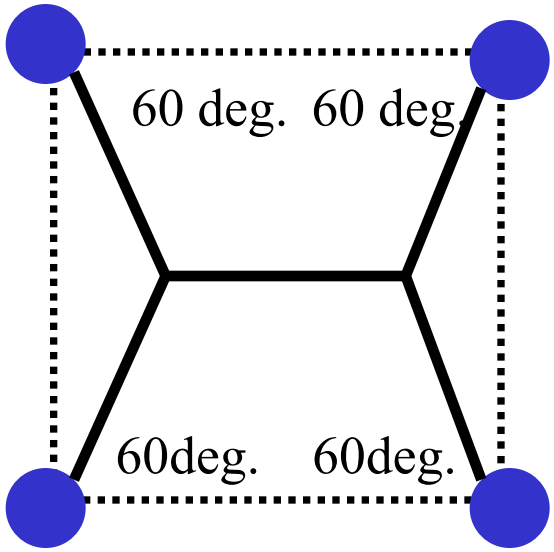
(question) how about the case for a square?



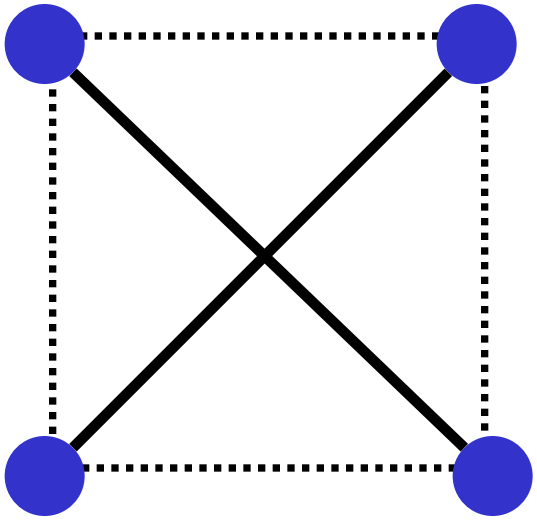
(the answer)



Length

$$4 \times \frac{1}{\sqrt{3}} + \left(1 - 2 \times \frac{1}{2\sqrt{3}}\right)$$
$$= 1 + \sqrt{3}$$
$$= 2.732\cdots$$

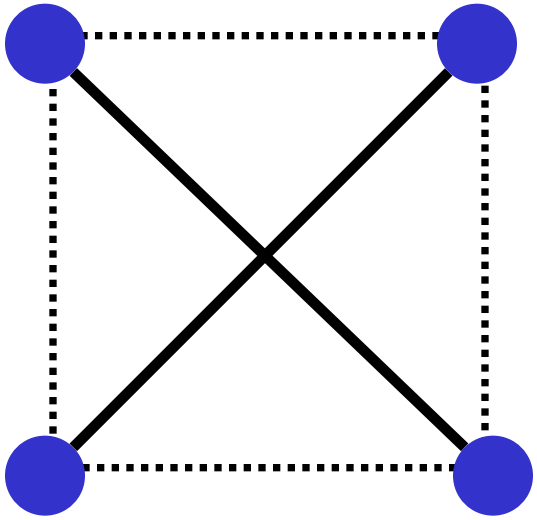
cf.



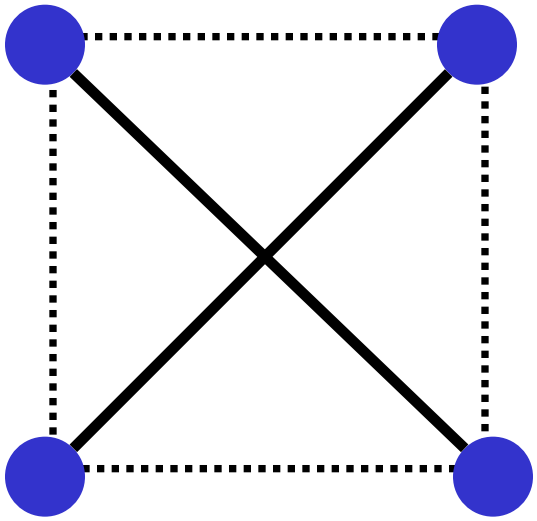
Length

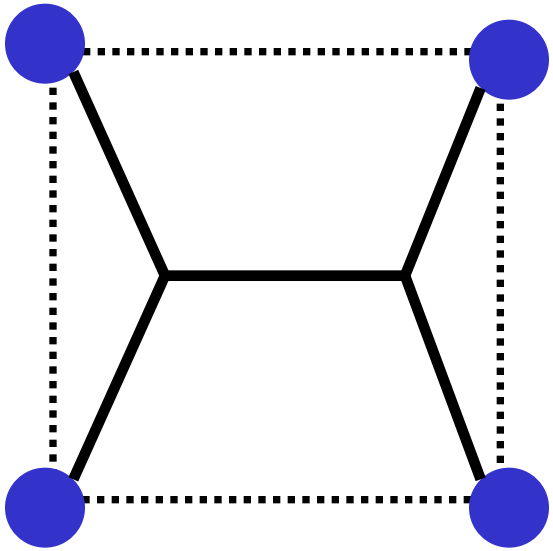
$$2 \times \sqrt{2} = 2.828\cdots$$

Ref. Takeshi Koike,
"Genshikaku Kenkyu" Vol. 52 No. 2, p. 14



invariant with
rotation by 90 deg.





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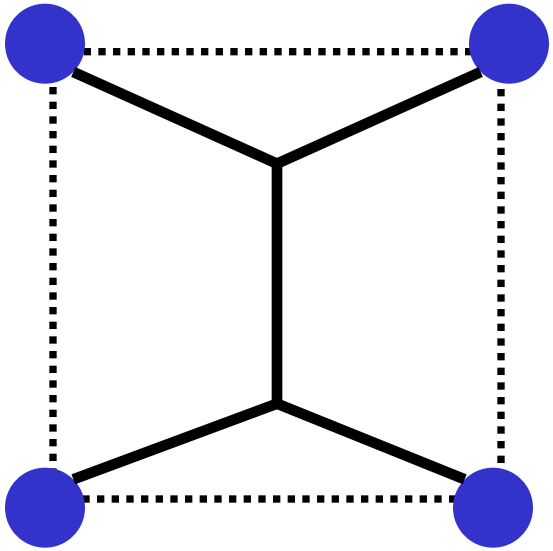

a good example of spontaneous symm. breaking

Courtesy: Takeshi Koike