

# Heavy-ion fusion reactions and superheavy elements

Kouichi Hagino

Tohoku University, Sendai, Japan



1. H.I. fusion reactions: why are they interesting?
2. Coupled-channels approach
3. Future perspectives: superheavy elements

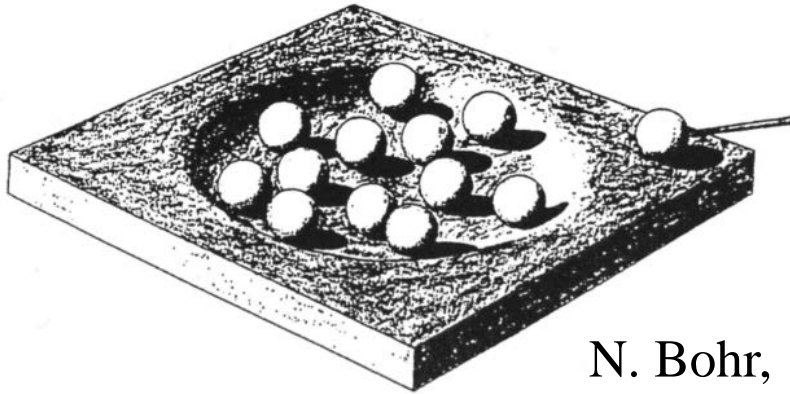
Recent review article:

K. Hagino and N. Takigawa, Prog. Theo. Phys.128 ('12)1061.

# Fusion reactions: compound nucleus formation

Niels Bohr (1936)

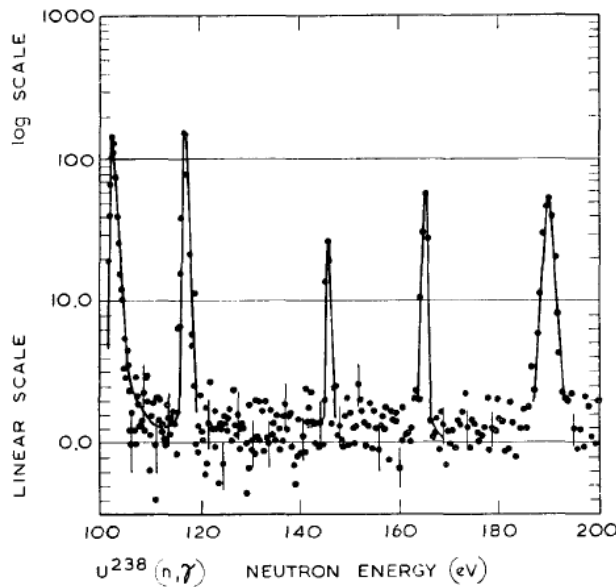
Neutron capture of nuclei  $\rightarrow$  compound nucleus



N. Bohr,  
Nature 137 ('36) 351



Wikipedia



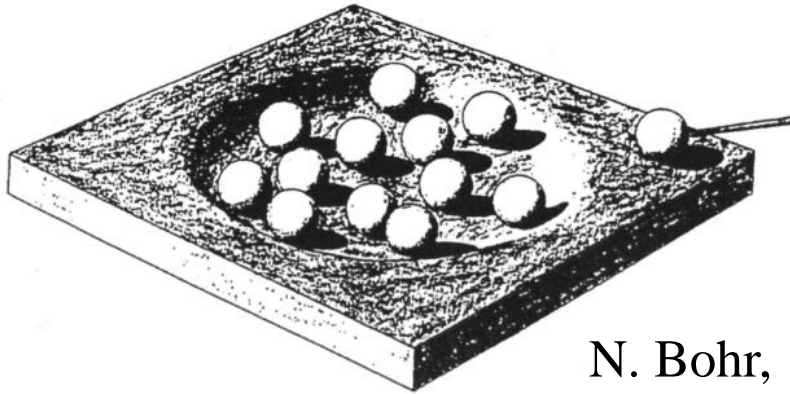
cf. Experiment of Enrico Fermi (1935)  
many very narrow (=long life-time)  
resonances (width  $\sim$  eV)

M. Asghar et al., Nucl. Phys. 85 ('66) 305

# Fusion reactions: compound nucleus formation

Niels Bohr (1936)

Neutron capture of nuclei  $\rightarrow$  compound nucleus

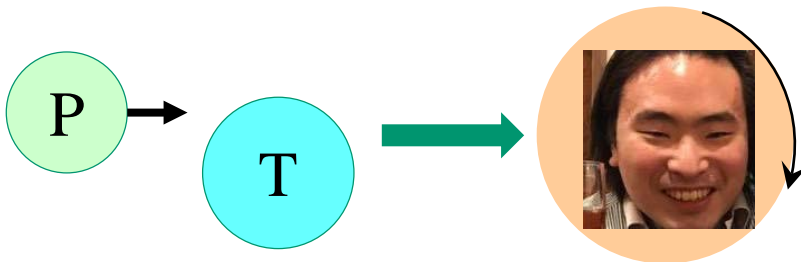


N. Bohr,  
Nature 137 ('36) 351



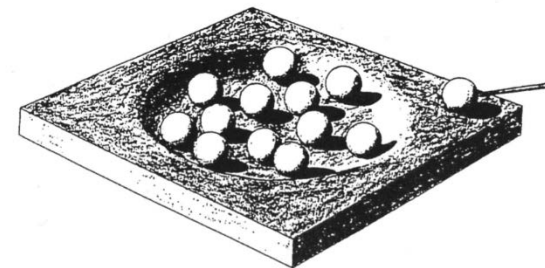
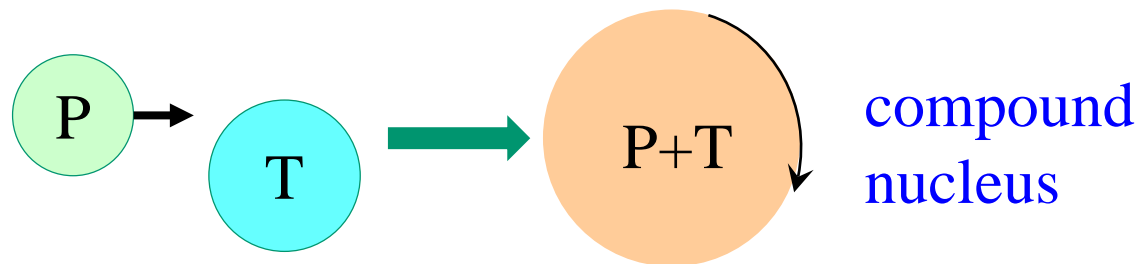
Wikipedia

forming a compound nucleus with heavy-ion reactions = H.I. fusion

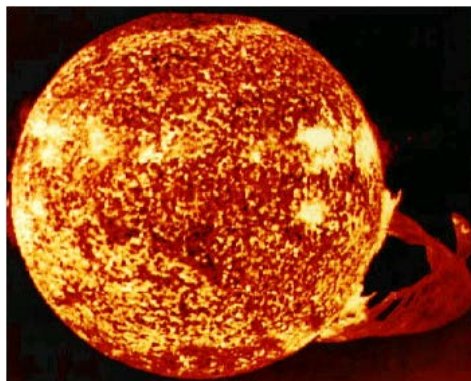


compound nucleus

# Fusion reactions: compound nucleus formation

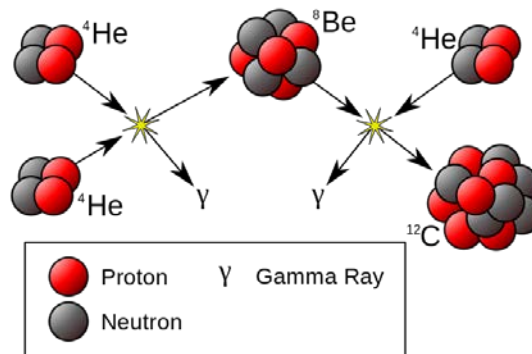


cf. Bohr '36

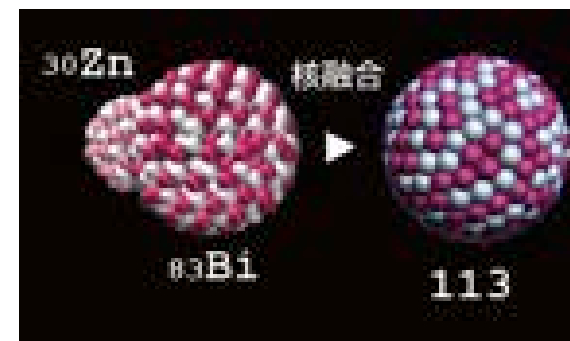


NASA, Skylab space station on December 19, 1973, solar flare reaching 589 000 km off solar surface

energy production  
in stars (Bethe '39)



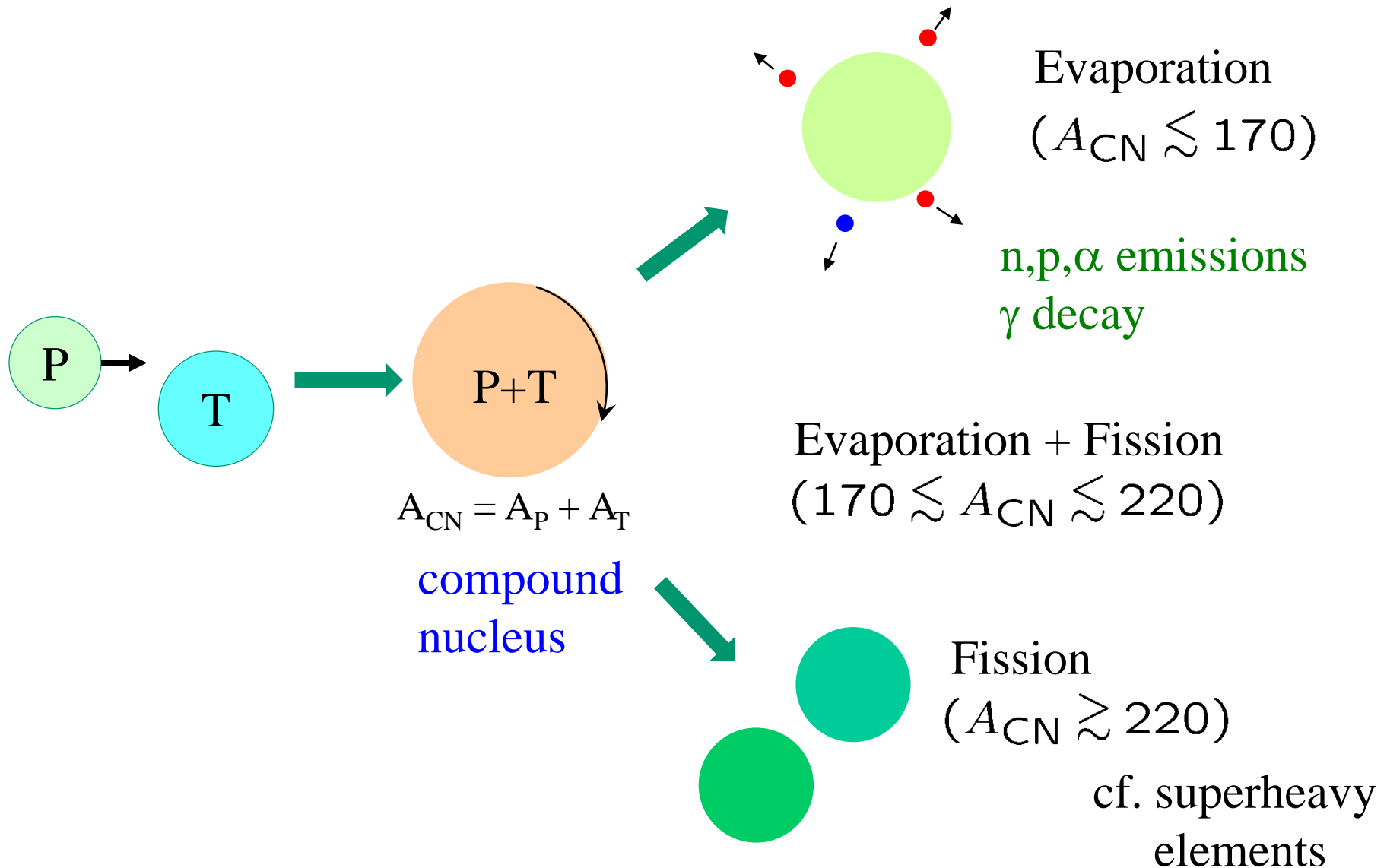
nucleosynthesis



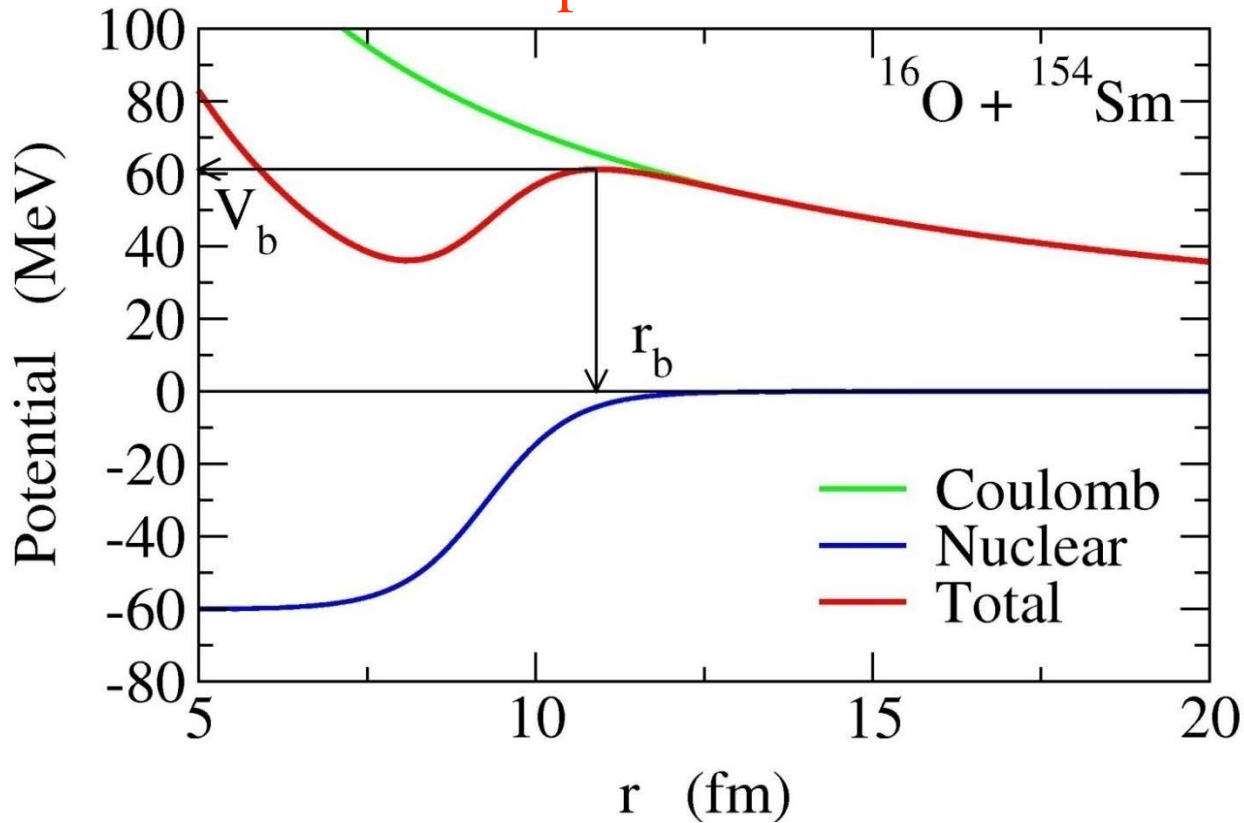
superheavy elements

**Fusion and fission:** large amplitude motions of quantum many-body systems with strong interaction  
← microscopic understanding: an ultimate goal of nuclear physics

# Fusion reactions: compound nucleus formation



## Inter-nucleus potential



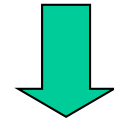
Two interactions:

1. Coulomb force

long range repulsion

2. Nuclear force

short range attraction



potential barrier  
due to a cancellation  
between the two  
(Coulomb barrier)

• Above-barrier energies



• Sub-barrier energies

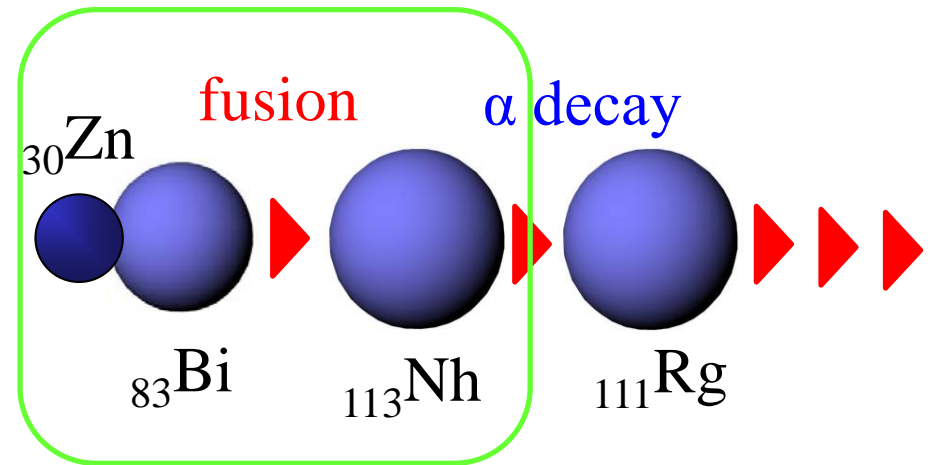
(energies around the Coulomb barrier)

• Deep sub-barrier energies

# Why sub-barrier fusion?

two obvious reasons:

113 <b>Nh</b> nihonium	115 <b>Mc</b> moscovium
117 <b>Ts</b> tennessine	118 <b>Og</b> oganesson



superheavy elements

cf.  $^{209}\text{Bi} (^{70}\text{Zn}, n) ^{278}\text{Nh}$

$V_B \sim 260 \text{ MeV}$

$E_{\text{cm}}^{(\text{exp})} \sim 262 \text{ MeV}$

# Why sub-barrier fusion?

two obvious reasons:

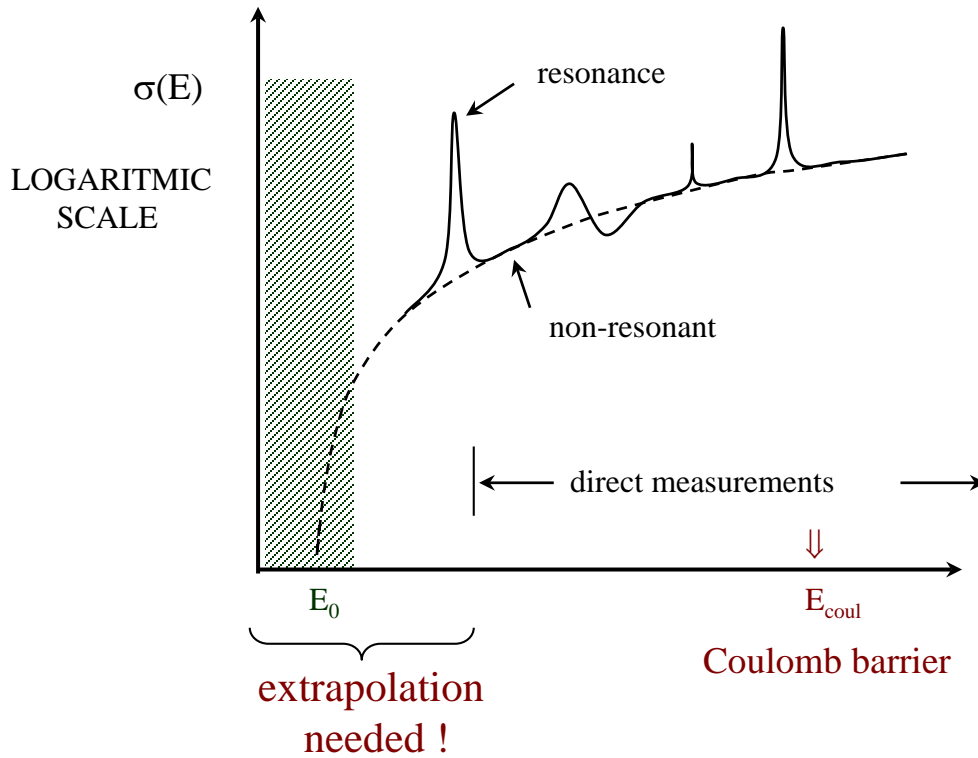
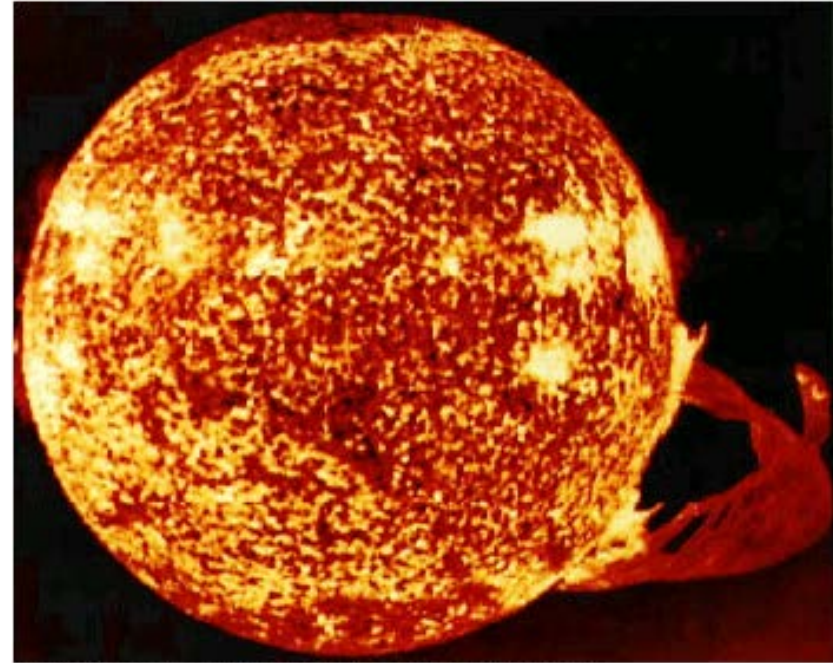


figure: M. Aliotta



NASA, Skylab space station December 19, 1973, solar flare reaching 588 000 km off solar surface

nuclear astrophysics  
(nuclear fusion in stars)

cf. extrapolation of data



# Why sub-barrier fusion?

two obvious reasons:

- ✓ superheavy elements
- ✓ nuclear astrophysics

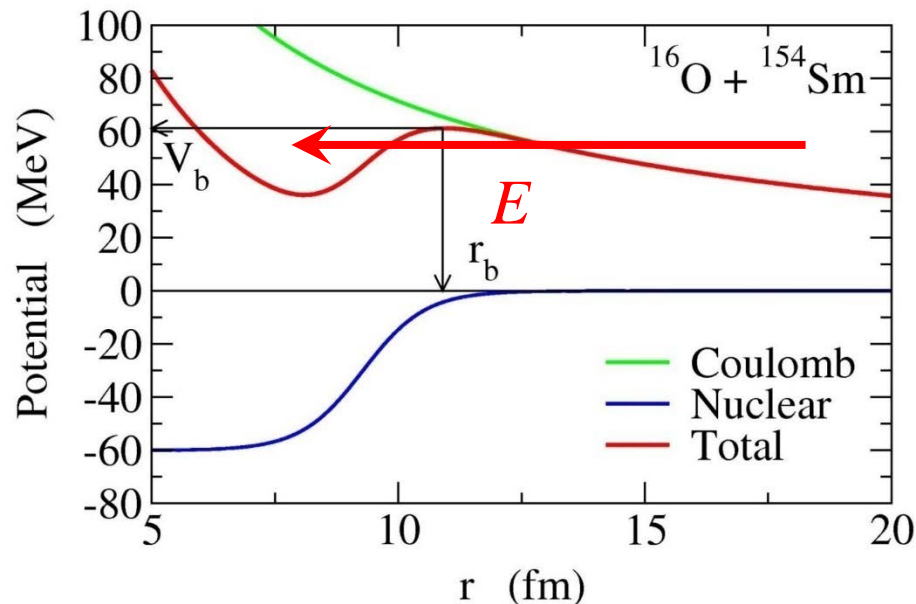
other reasons:

- ✓ reaction dynamics

**strong interplay between reaction and structure**

cf. high  $E$  reactions: much simpler reaction mechanisms

- ✓ many-particle tunneling



# Why sub-barrier fusion?

two obvious reasons:

- ✓ superheavy elements
- ✓ nuclear astrophysics

other reasons:

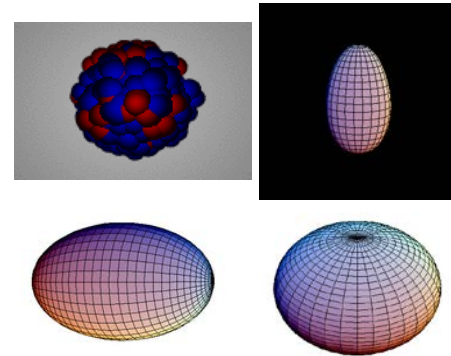
- ✓ reaction dynamics

**strong interplay between reaction and structure**

cf. high  $E$  reactions: much simpler reaction mechanisms

- ✓ many-particle tunneling

- many types of intrinsic degrees of freedom (several types of collective vibrations, deformation with several multipolarities)
- energy dependence of tunneling probability  
cf. alpha decay: fixed energy



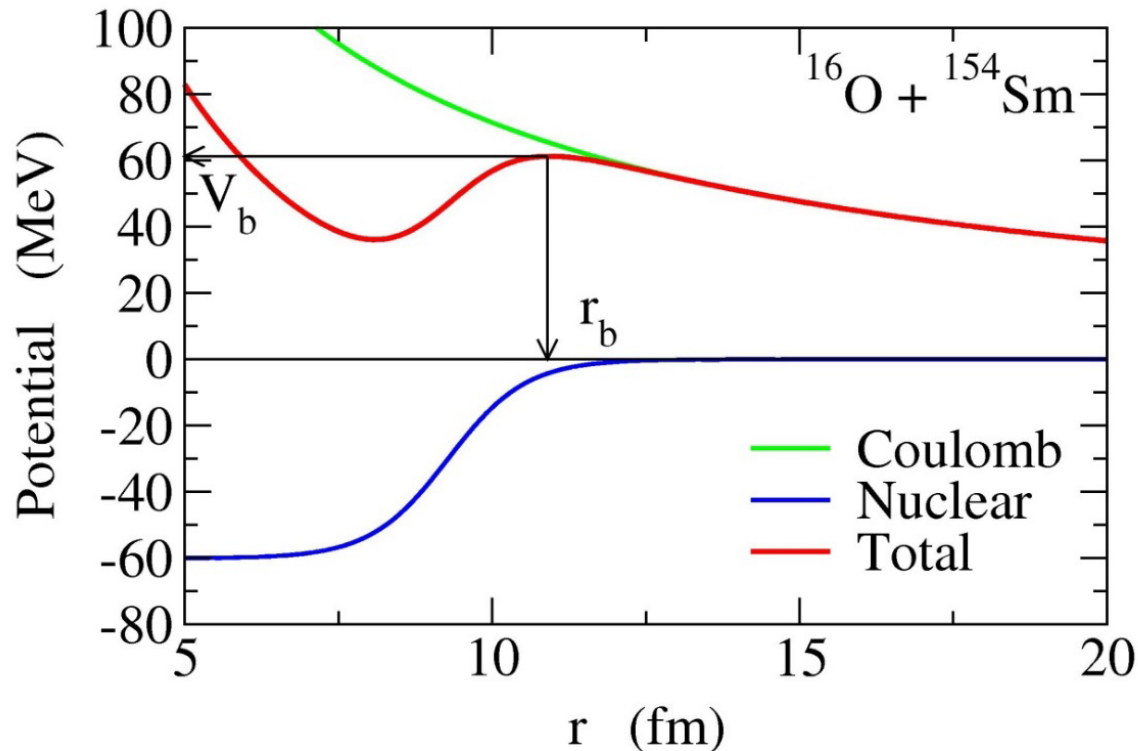
H.I. fusion reaction = an ideal playground to study quantum tunneling with many degrees of freedom

# The simplest approach to fusion: potential model

Potential model:  $V(r) + \text{absorption}$

$$\sigma_{\text{fus}}(E) = \frac{\pi}{k^2} \sum_l (2l + 1) P_l(E)$$

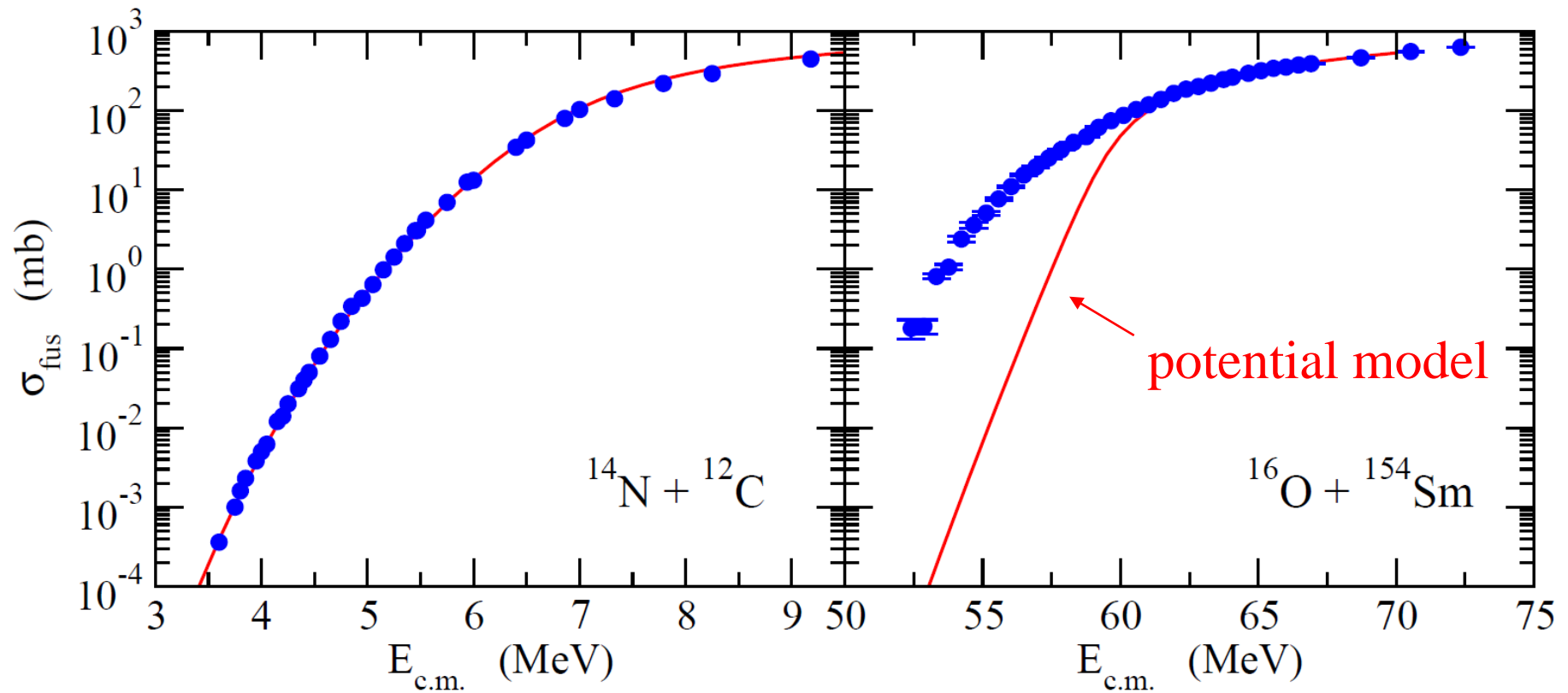
$P_l(E)$ : barrier penetrability



# Comparison with experimental data: large enhancement of $\sigma_{\text{fus}}$

Potential model:  $V(r) + \text{absorption}$

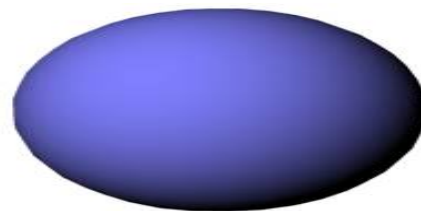
$$\sigma_{\text{fus}} = \frac{\pi}{k^2} \sum_l (2l + 1)(1 - |S_l|^2)$$



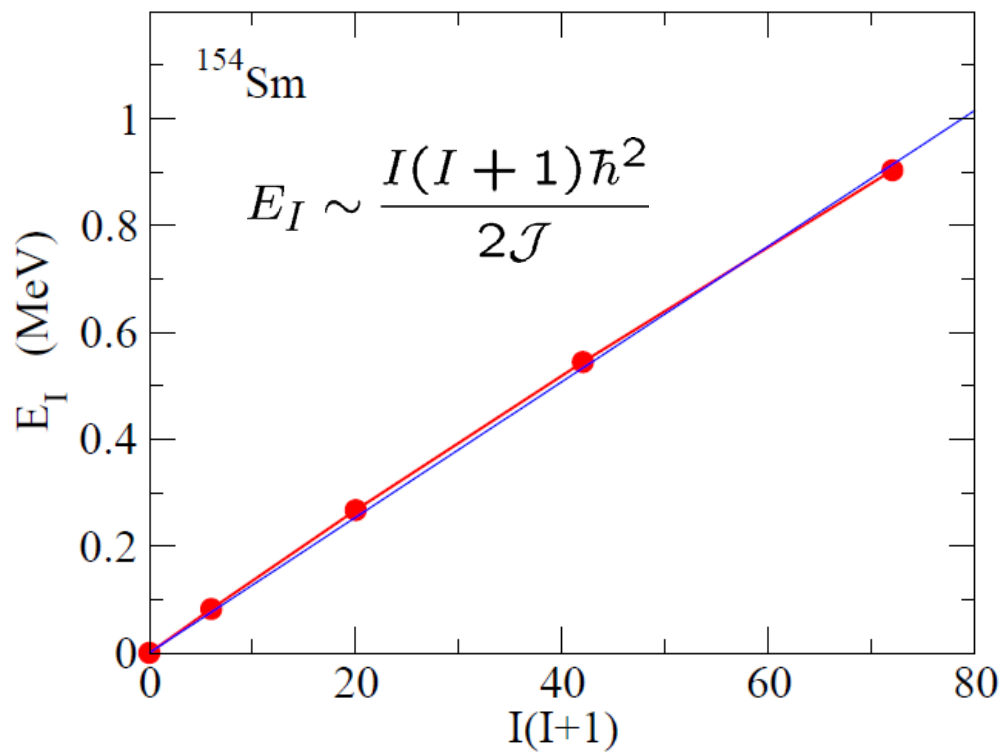
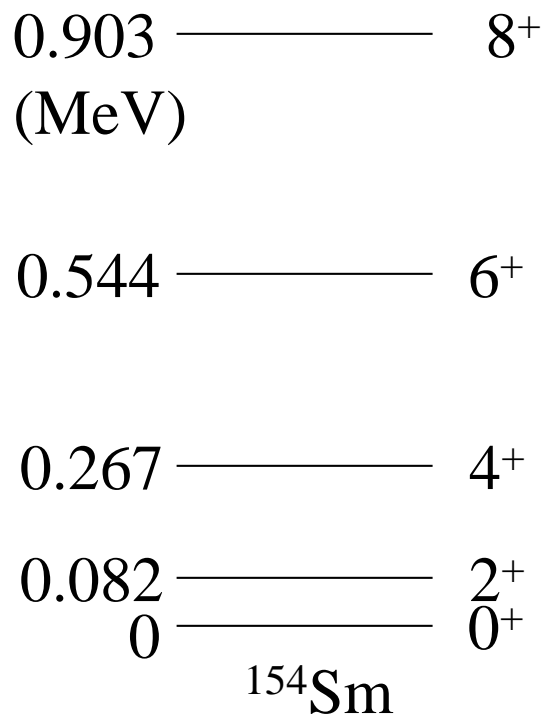
cf. seminal work:

R.G. Stokstad et al., PRL41('78) 465

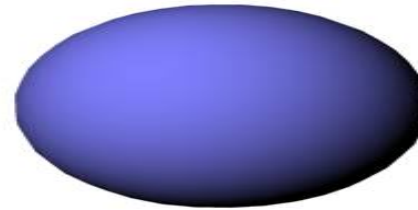
$^{154}\text{Sm}$  : a typical deformed nucleus



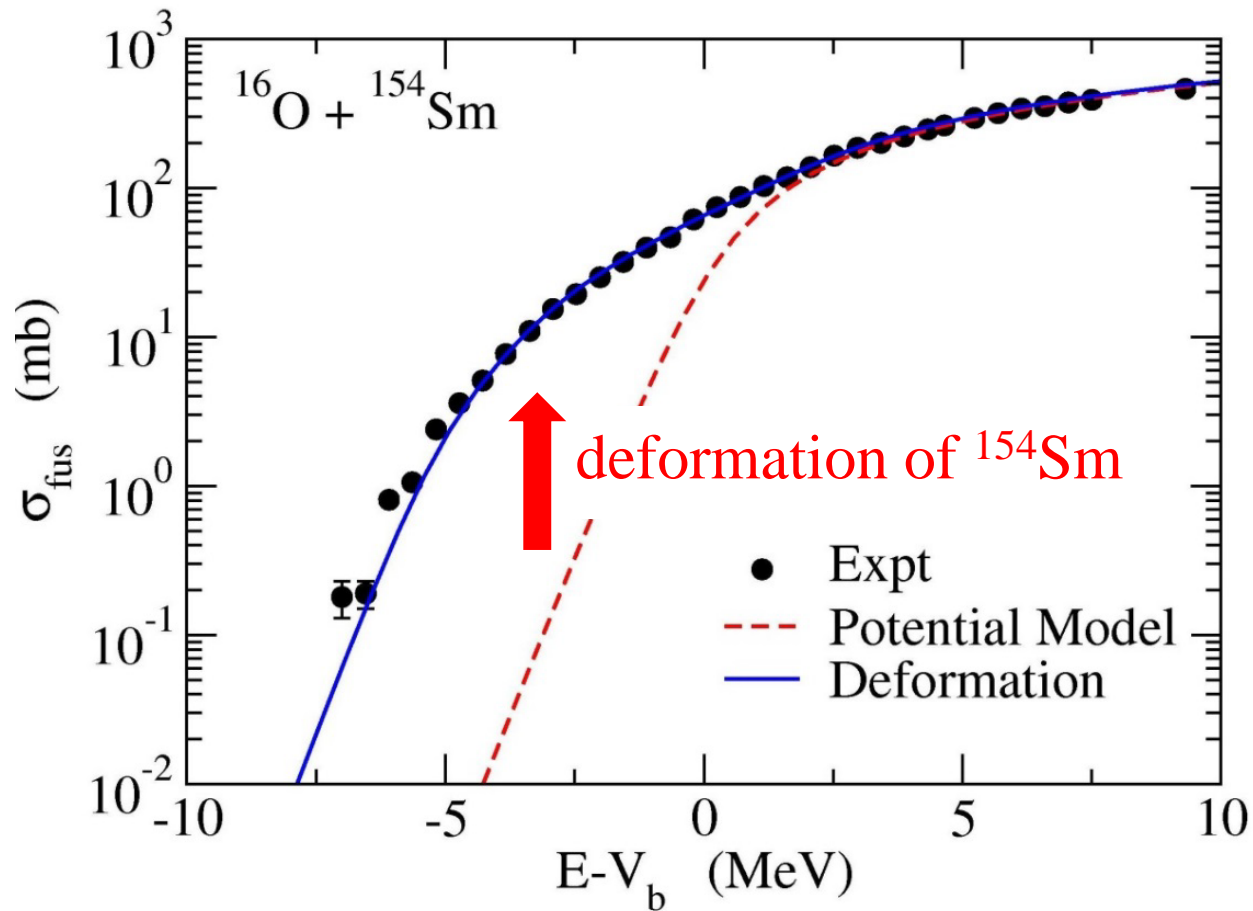
$^{154}\text{Sm}$



$^{154}\text{Sm}$  : a typical deformed nucleus

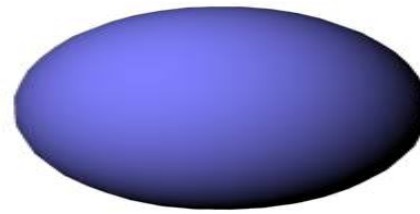


$^{154}\text{Sm}$

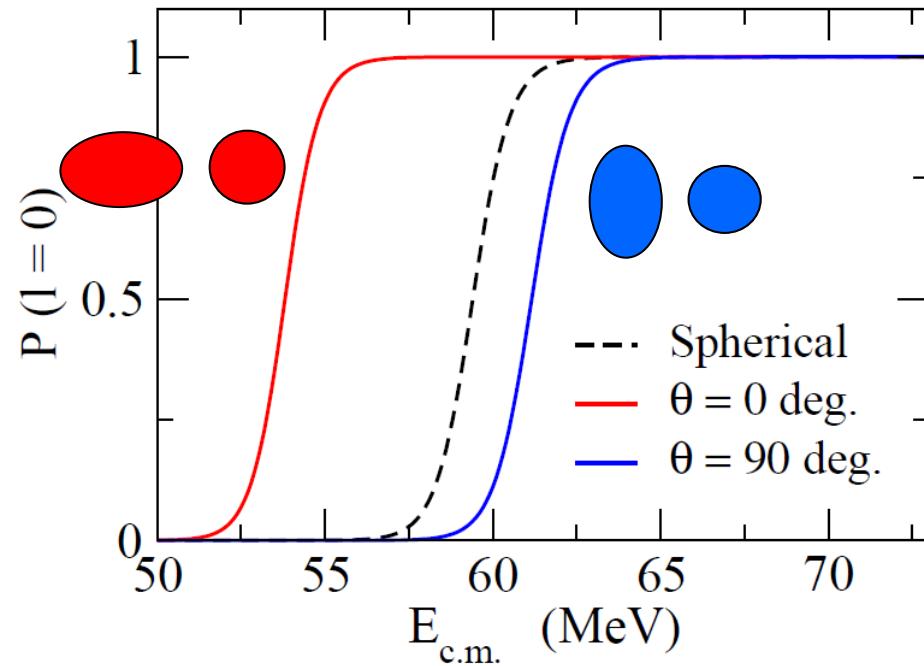
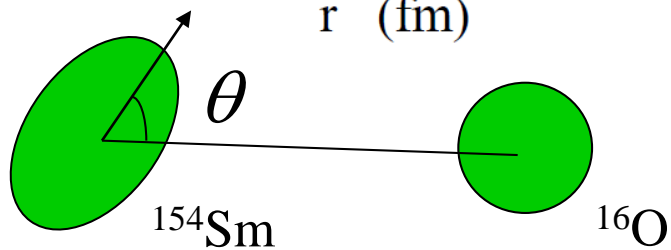
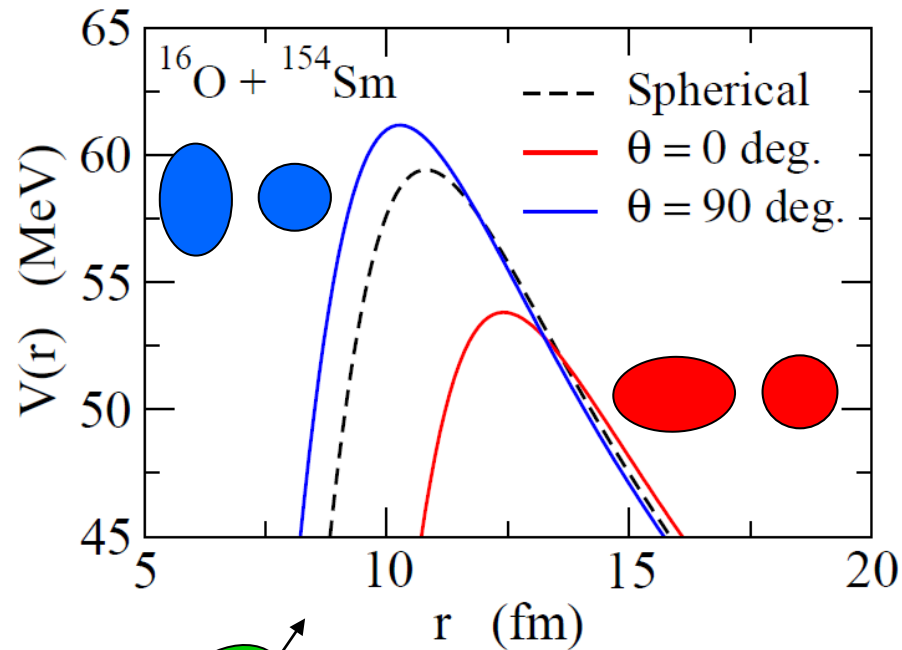


# Effects of nuclear deformation

$^{154}\text{Sm}$  : a typical deformed nucleus

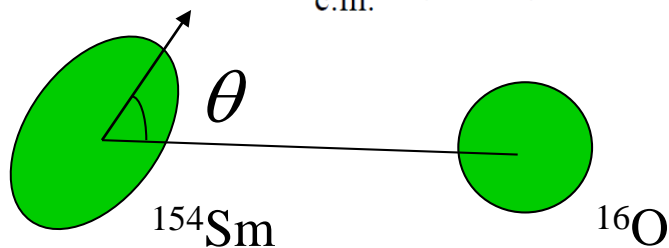
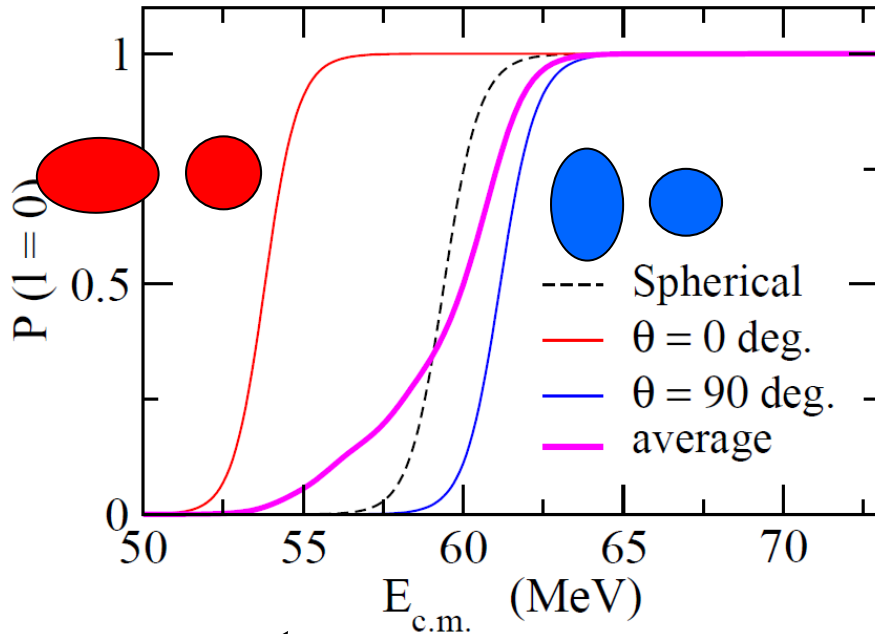


$^{154}\text{Sm}$

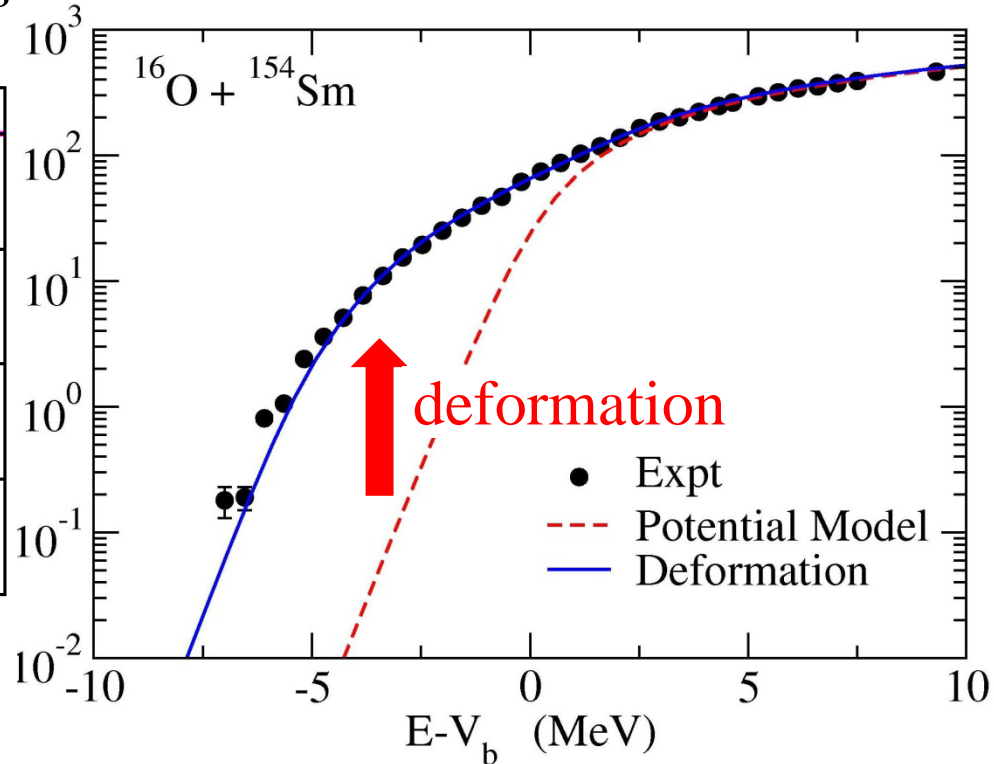


# Effects of nuclear deformation

$^{154}\text{Sm}$  : a typical deformed nucleus



$$\sigma_{\text{fus}}(E) = \int_0^1 d(\cos \theta) \sigma_{\text{fus}}(E; \theta)$$

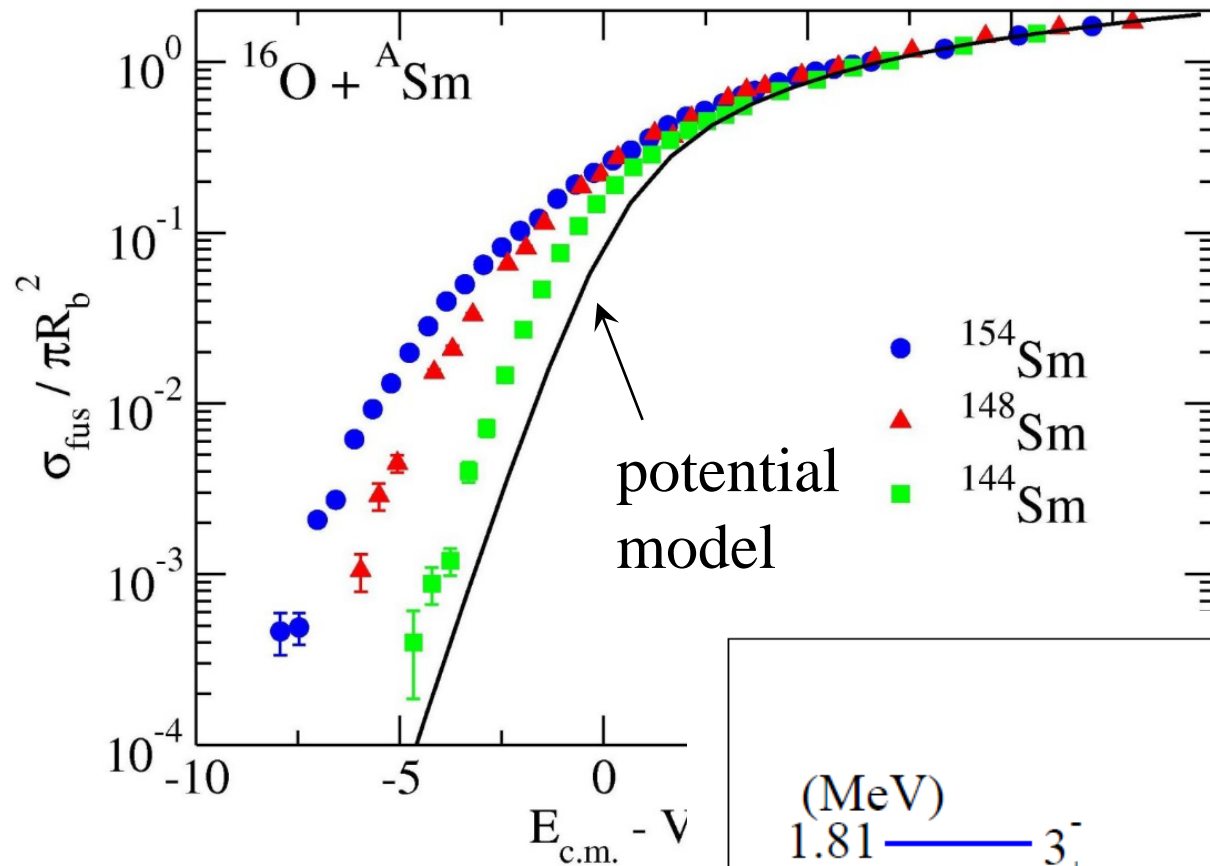


**Fusion: strong interplay between nuclear structure and reaction**

\* Sub-barrier enhancement also for non-deformed targets:

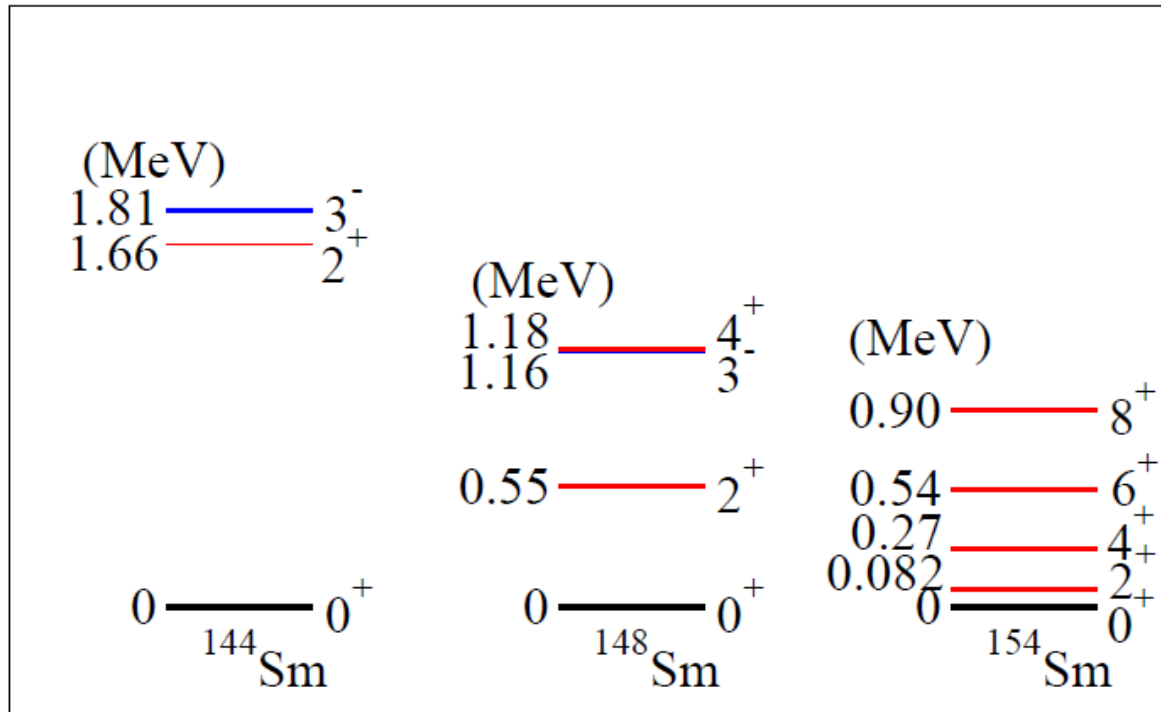
couplings to low-lying collective excitations  $\rightarrow$  coupling assisted tunneling





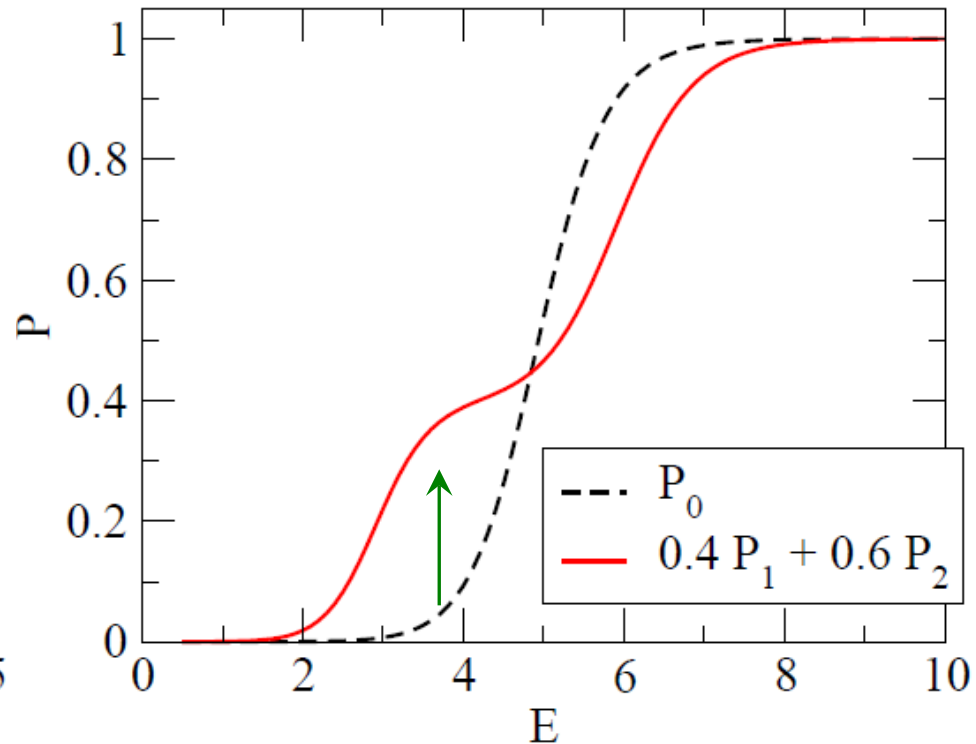
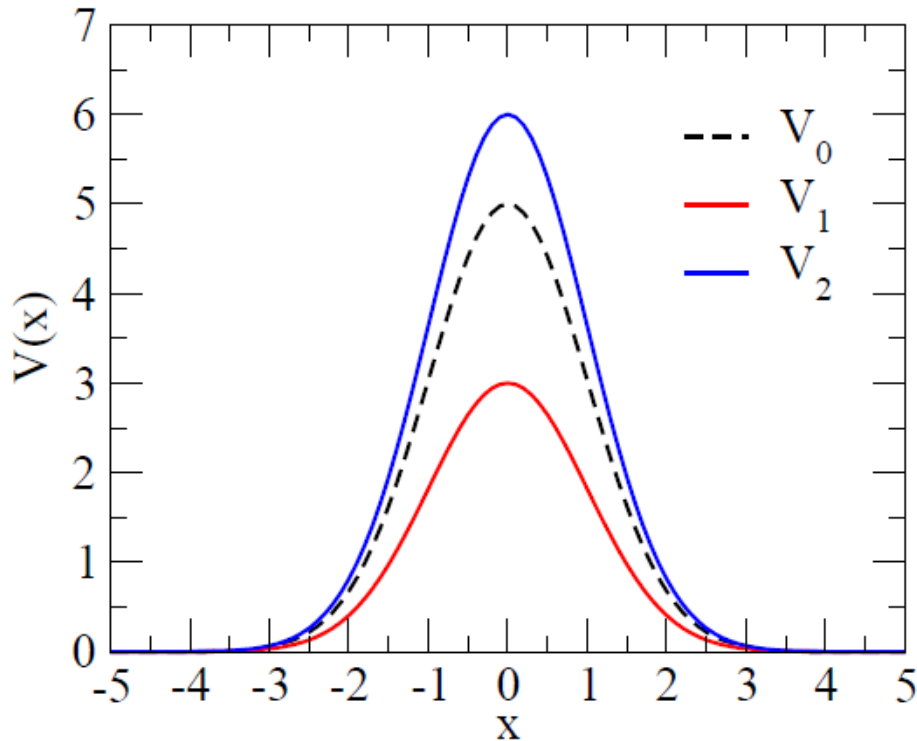
enhancement of fusion cross sections  
: a general phenomenon

strong correlation with nuclear spectrum  
→ coupling assisted tunneling



# Enhancement of tunneling probability : a problem of two potential barriers

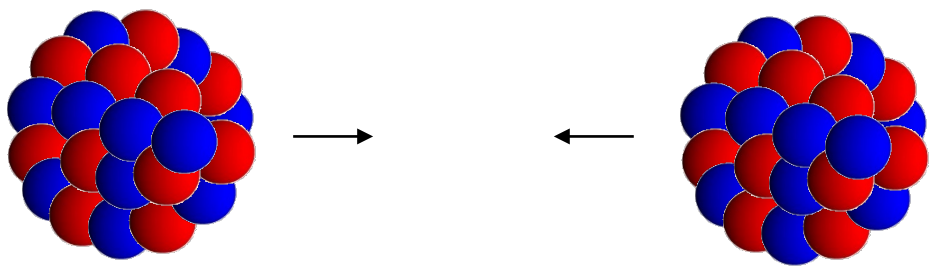
$$P(E) = P(E; V_0) \rightarrow w_1 P(E; V_1) + w_2 P(E; V_2)$$



“barrier distribution” due to couplings to excited states  
in projectile/target nuclei

# Coupled-channels method: a quantal scattering theory with excitations

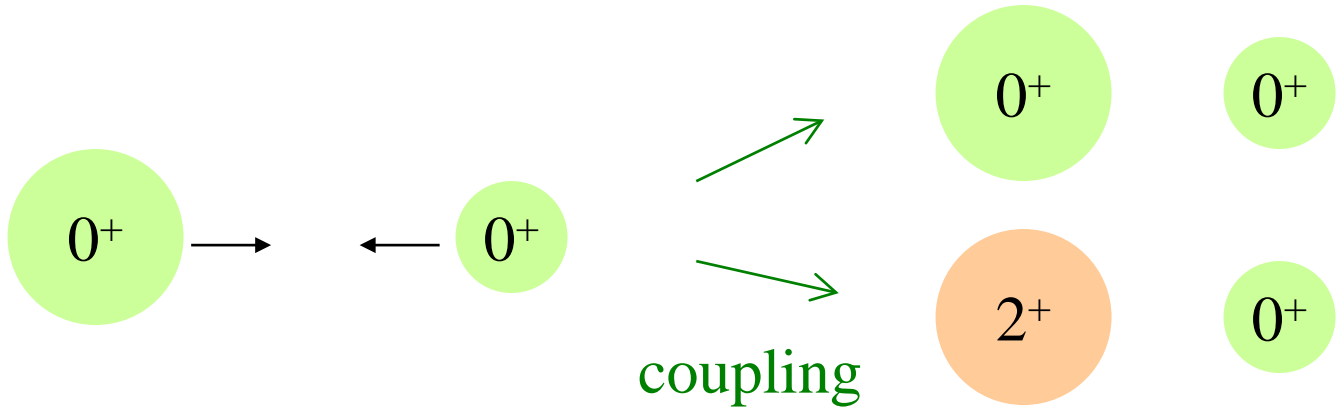
many-body problem



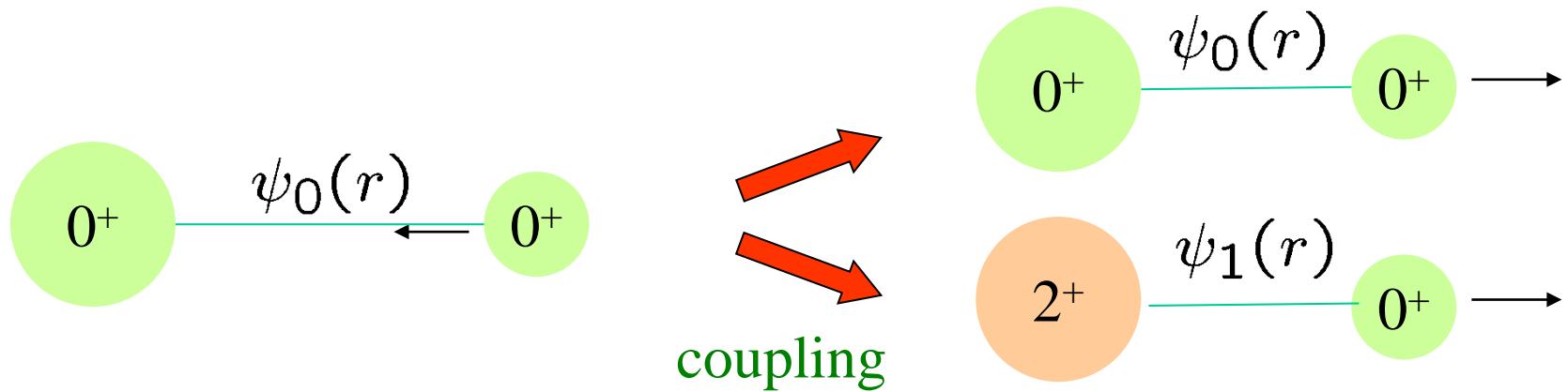
still very challenging



two-body problem, but with excitations  
(coupled-channels approach)



# Coupled-channels method: a quantal scattering theory with excitations



$$\left[ -\frac{\hbar^2}{2\mu} \nabla^2 + \overleftarrow{V}(r) - \overleftarrow{E} \right] \overrightarrow{\psi}(r) = 0$$

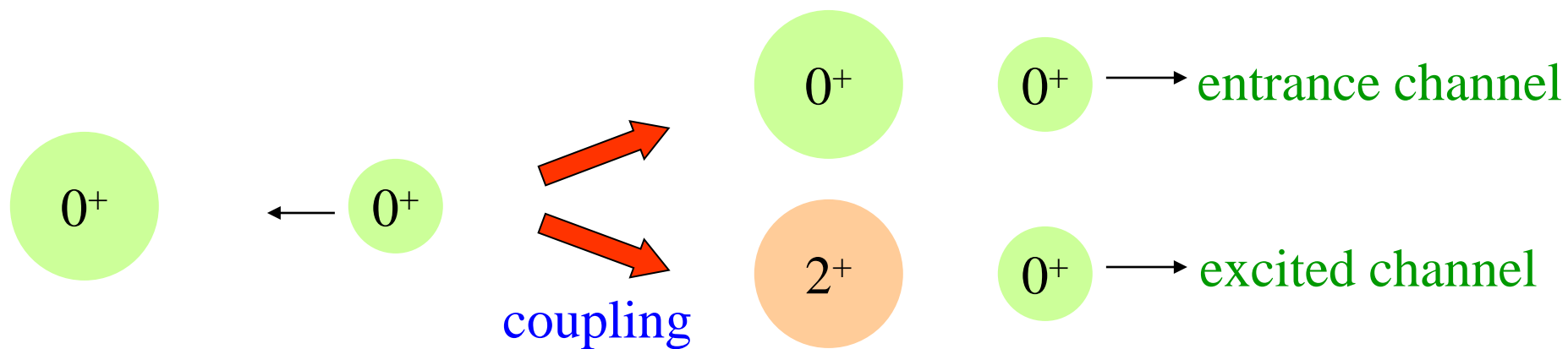
if written down more explicitly:

$$\left[ -\frac{\hbar^2}{2\mu} \nabla^2 + V_0(r) + \epsilon_k - E \right] \psi_k(r) + \sum_{k'} \langle \phi_k | V_{\text{coup}} | \phi_{k'} \rangle \psi_{k'}(r) = 0$$

↑  
excitation energy

↑  
excitation operator

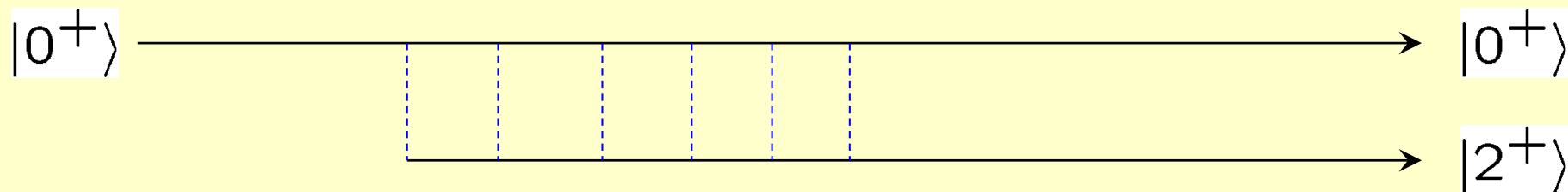
# Coupled-channels method: a quantal scattering theory with excitations



$$\left[ -\frac{\hbar^2}{2\mu} \nabla^2 + V_0(r) + \epsilon_k - E \right] \psi_k(\mathbf{r}) + \sum_{k'} \langle \phi_k | V_{\text{coup}} | \phi_{k'} \rangle \psi_{k'}(\mathbf{r}) = 0$$

excitation energy

excitation operator



full order treatment of excitation/de-excitation dynamics during reaction

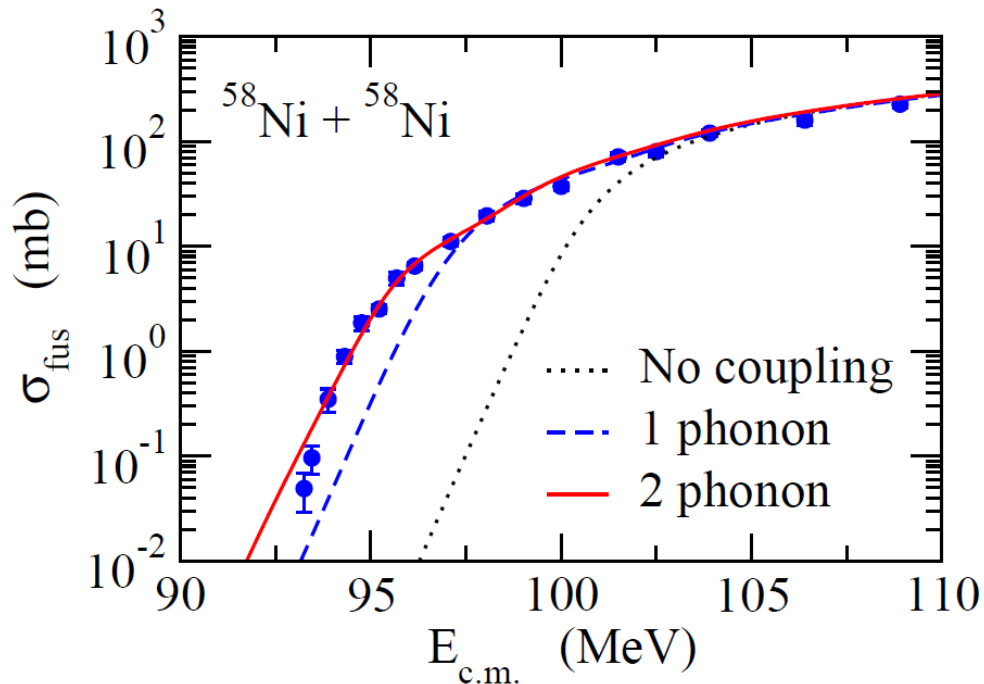
## Inputs for C.C. calculations

### i) Inter-nuclear potential

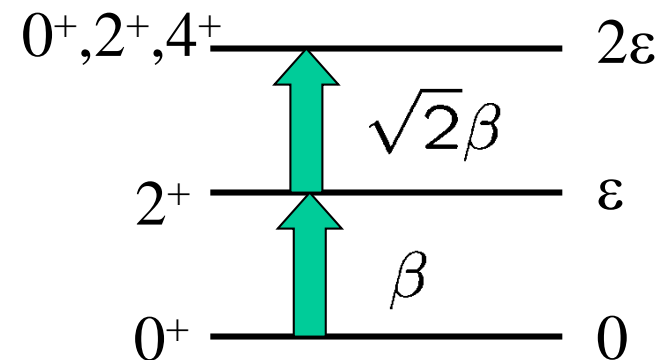
a fit to experimental data at above barrier energies

### ii) Intrinsic degrees of freedom

in most of cases, (macroscopic) collective model  
(rigid rotor / harmonic oscillator)



simple harmonic oscillator



# Further development: semi-microscopic modelling

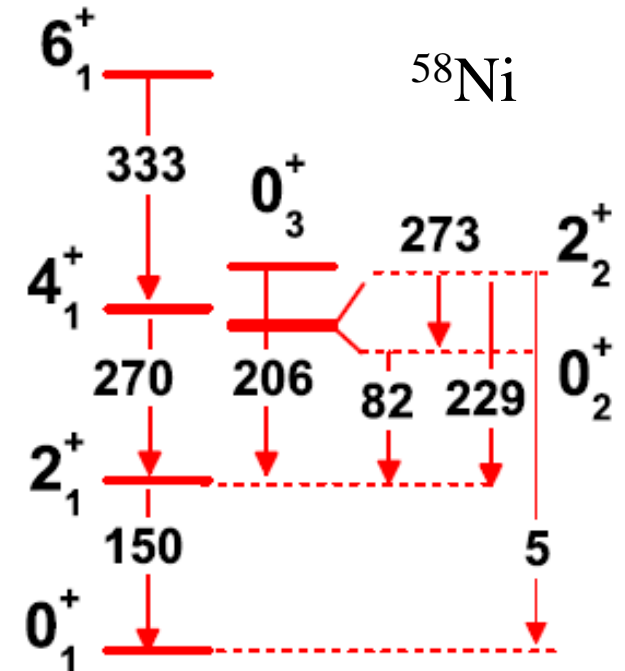
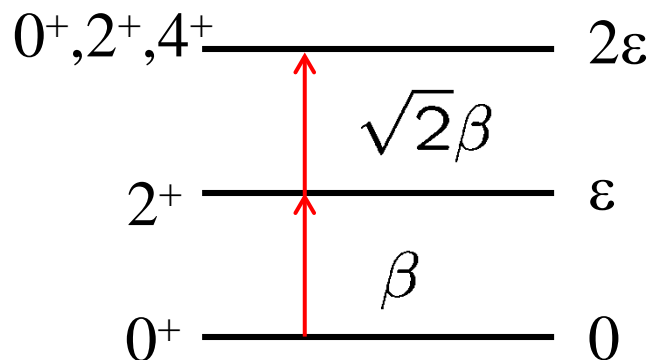
K.H. and J.M. Yao, PRC91('15) 064606

CCFULL

+ microscopic nuclear structure calculations  
(GCM, Shell Model, IBM.....)



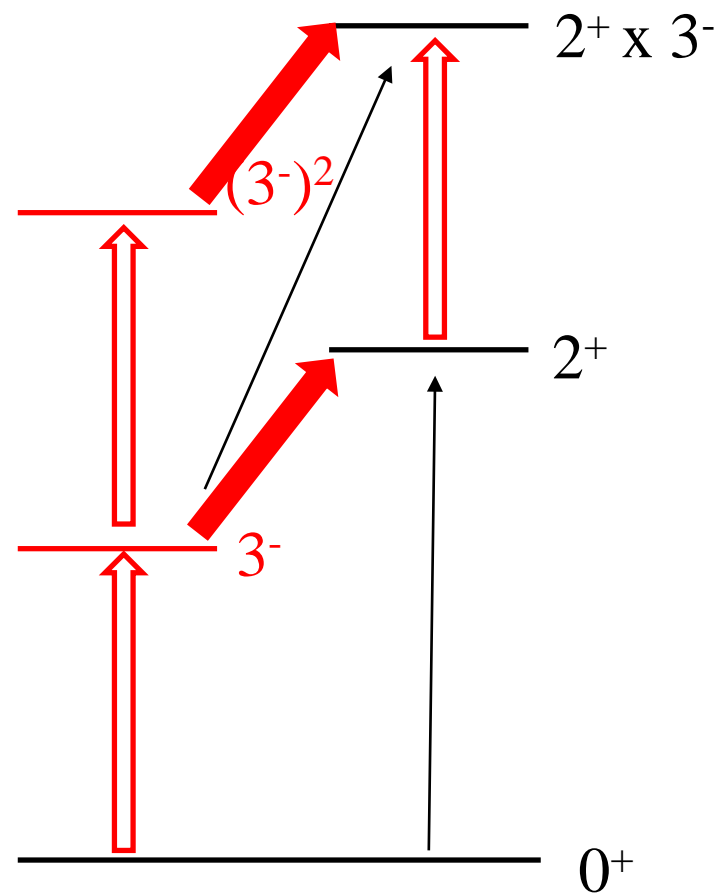
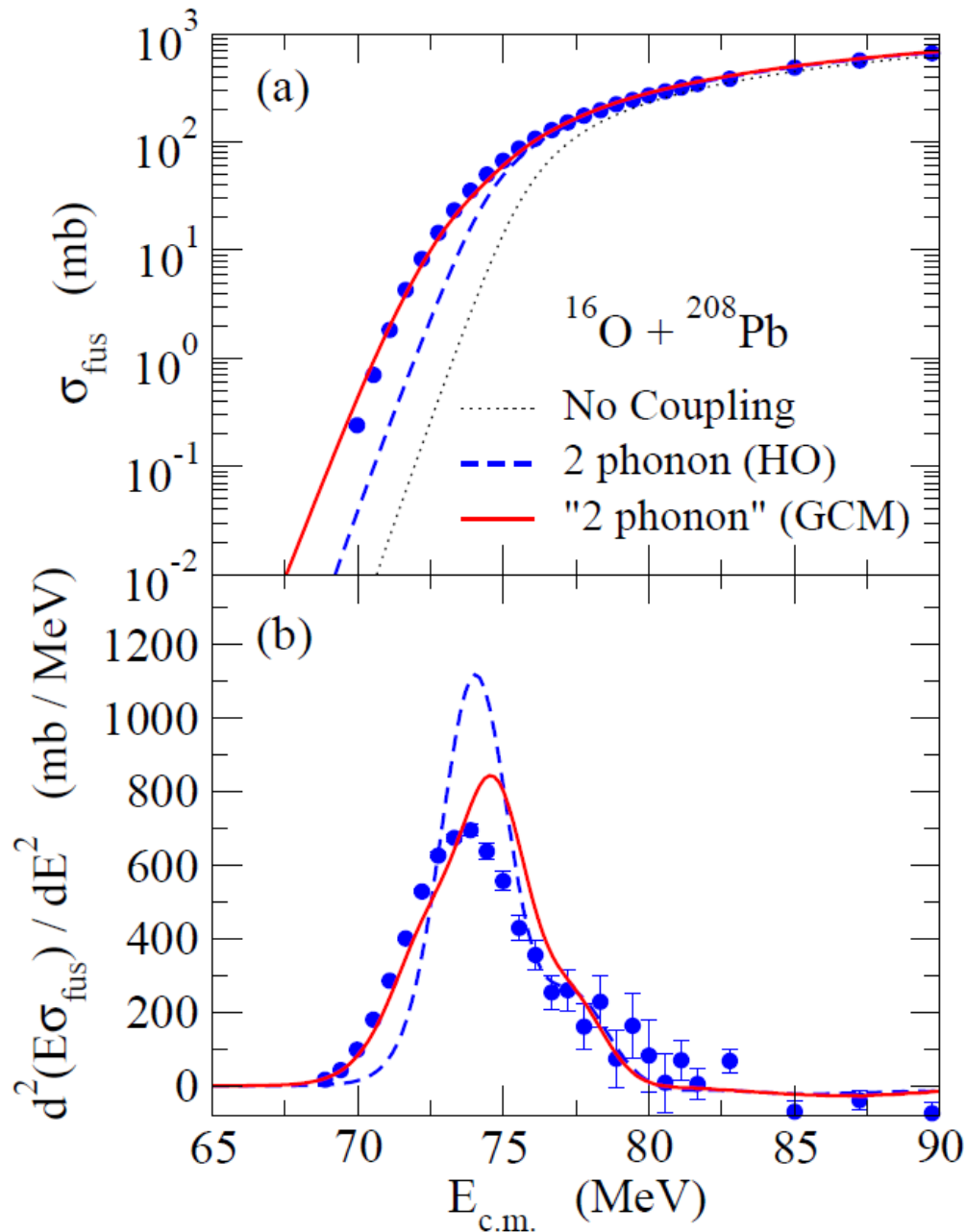
simple harmonic oscillator



relativistic MF + GCM

anharmonicity of phonon spectra

# CCFULL with RMF+GCM



J.M. Yao and K.H.,  
PRC94 ('16) 11303(R)



# From phenomenological approach to microscopic approach

Macroscopic (phenomenological)

C.C. with collective model

C.C. with inputs from  
microscopic nuclear  
structure calculations

- \* Hagino-Yao
- \* Ichikawa-Matsuyanagi

C.C. with inputs based  
on TDHF

- \* Umar et al. (DC-TDHF)
- \* Washiyama-Lacroix
- \* Simenel et al.

TDHF simulations

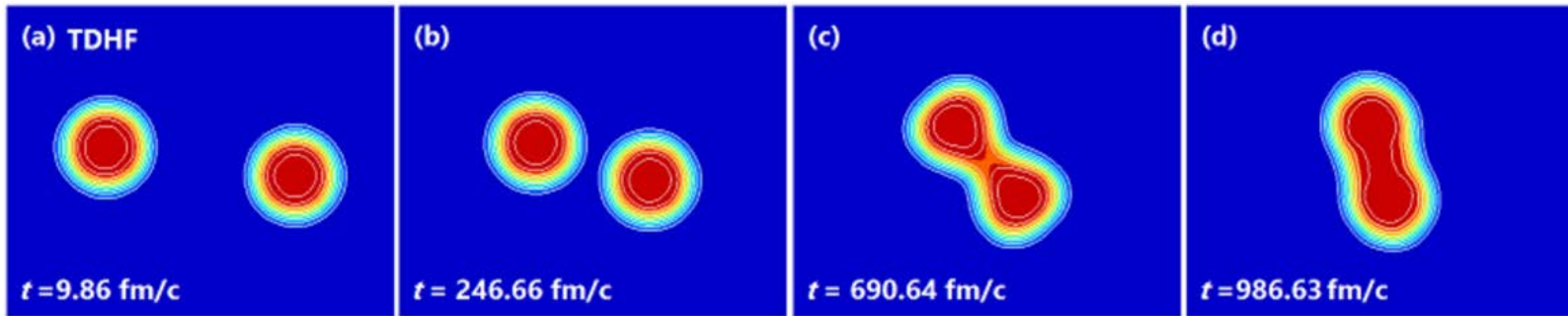
- \* Simenel
- \* Sekizawa
- \* Washiyama
- \* Iwata-Otsuka
- \* Maruhn, Stevenson

Microscopic

“ab initio”, but no tunneling

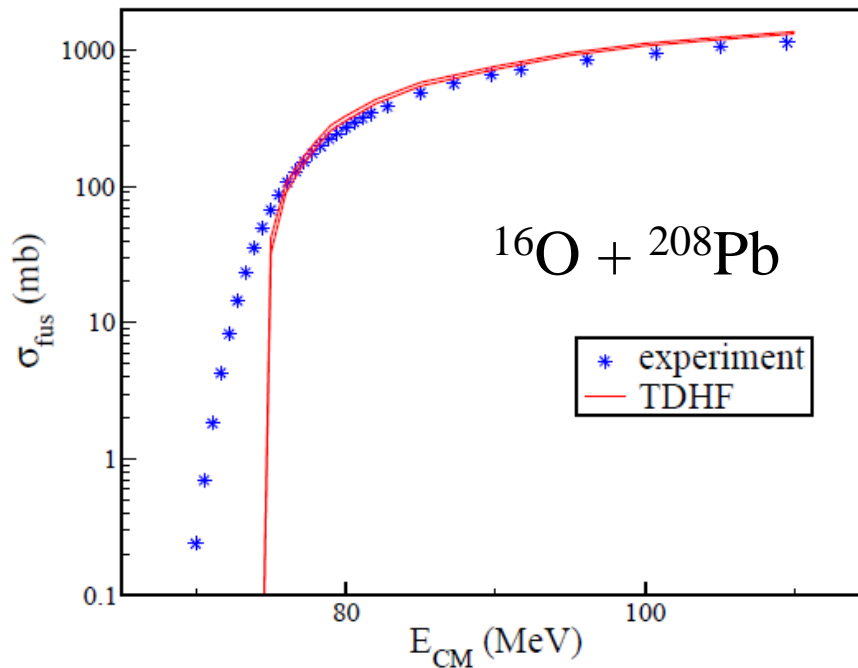
# TDHF simulation

TDHF = Time Dependent Hartree-Fock



S. Ebata, T. Nakatsukasa, JPC Conf. Proc. 6 ('15) 020056

“ab-initio”, but no tunneling



C. Simenel,  
EPJA48 ('12) 152

## TDHF simulation

“ab-initio”, but no tunneling

✓ DC-TDHF (Umar, Oberacker, Maruhn)

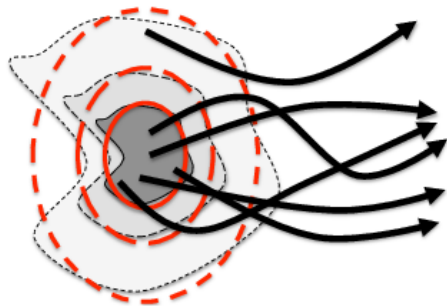
➤ Beyond mean-field approximation

✓ Collective Hamiltonian and requantization

K. Wen and T. Nakatsukasa, PRC96 ('17) 014610

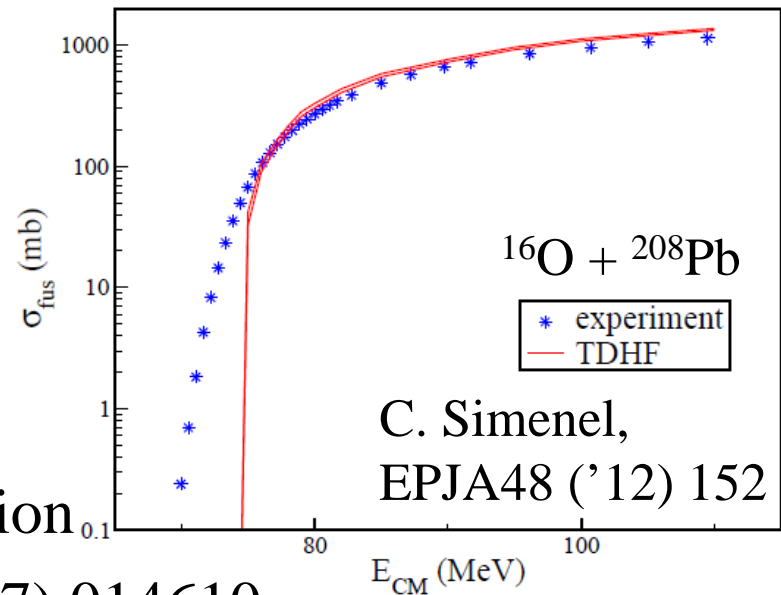
✓ Time-dependent Generator Coordinate Method (TD-GCM)

$$|\Psi(t)\rangle = \int dq f(q, t) |\Phi_q(t)\rangle$$



a linear superposition of many  
TDHF trajectories (Slater determinants)

an important future direction



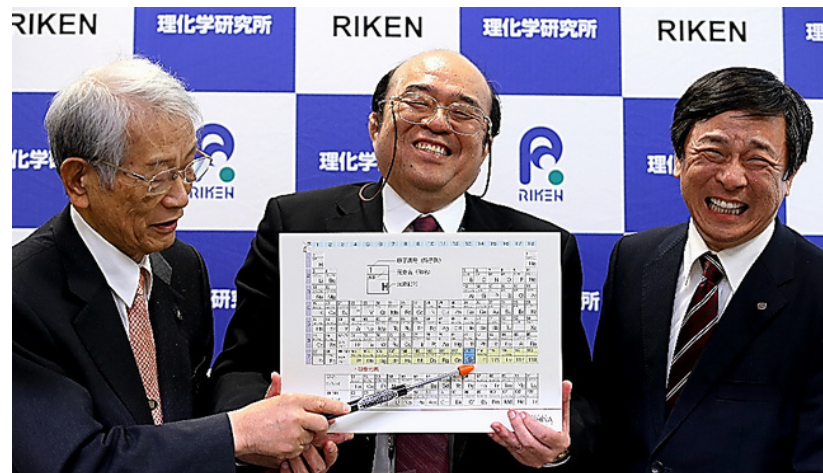
cf. Stochastic mean-field method  
B. Yilmaz et al.,  
PRC90 ('14) 054617

# Fusion reactions for SHE

the element 113: Nh

113 <b>Nh</b> nihonium	115 <b>Mc</b> moscovium
117 <b>Ts</b> tennessine	118 <b>Og</b> oganesson

November, 2016



Group	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	
1	1 H																		2 He
2	3 Li	4 Be											5 B	6 C	7 N	8 O	9 F	10 Ne	
3	11 Na	12 Mg											13 Al	14 Si	15 P	16 S	17 Cl	18 Ar	
4	19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr	
5	37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe	
6	55 Cs	56 Ba	57 La	* 72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn	
7	87 Fr	88 Ra	89 Ac	* 104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Nh	114 Fl	115 Mc	116 Lv	117 Ts	118 Og	
	* 58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu					
	* 90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr					

Wikipedia

# Fusion reactions for SHE

the element 113: Nh

113 钿 nihonium	115 镆 moscovium
117 砹 tennessine	118 镏 oganesson

核データニュース, No.118 (2017)

話題・解説(I)

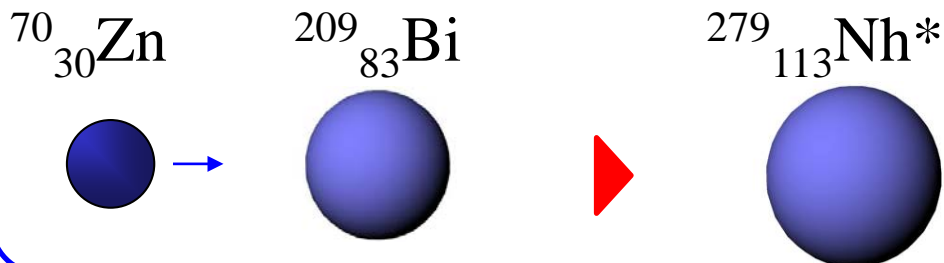
Chinese Names of New Elements with  $Z = 113, 115, 117$  &  $118$

Shan-Gui Zhou (周善贵/周善貴)<sup>1</sup>

Institute of Theoretical Physics, Chinese Academy of Sciences, Beijing 100190, China

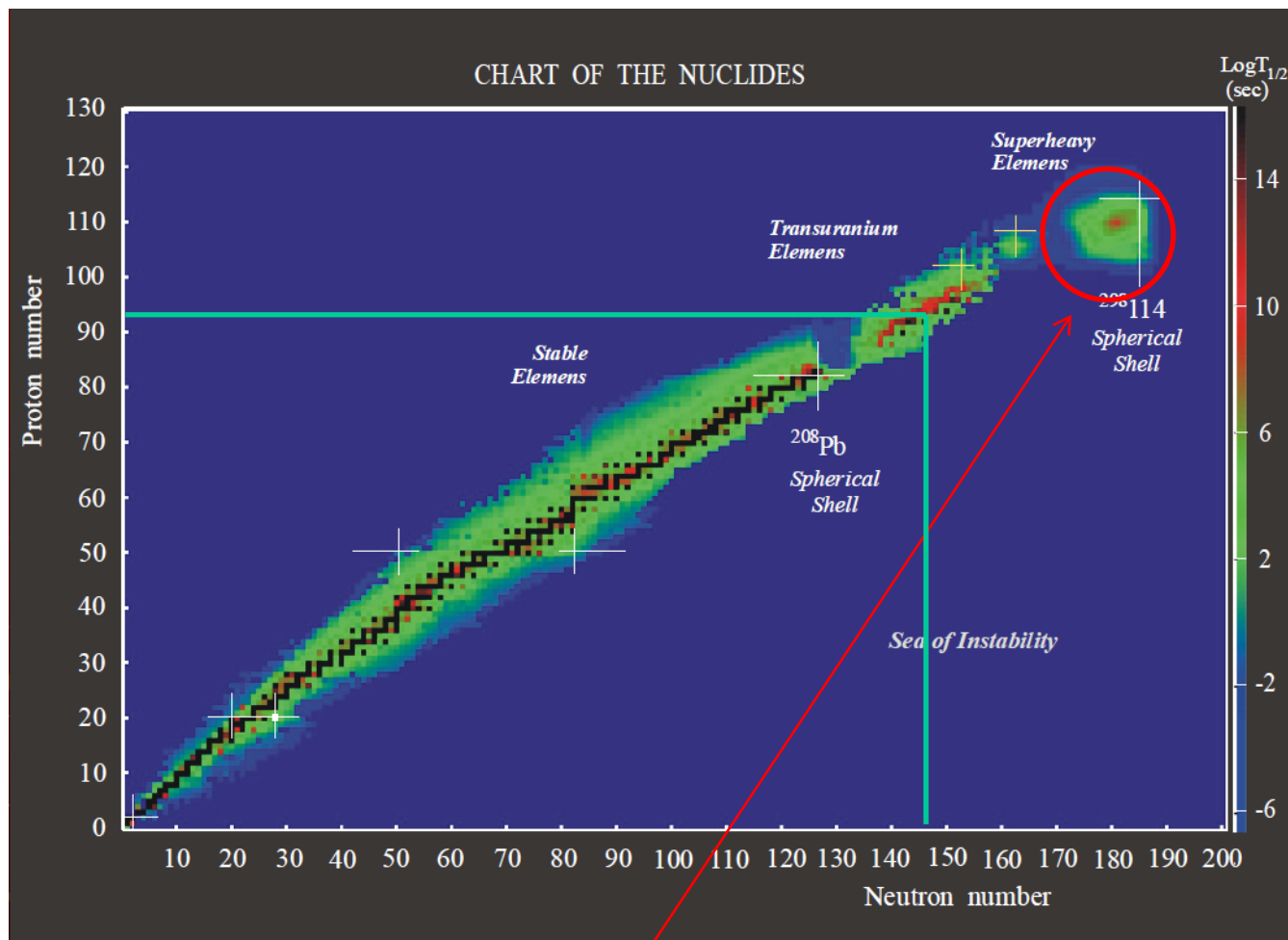
(中国科学院理论物理研究所/中國科學院理論物理研究所)

November, 2016



Heavy-ion fusion reaction

# Future perspectives: Superheavy elements



**island of stability around  $Z=114$ ,  $N=184$**

Yuri Oganessian

W.D. Myers and W.J. Swiatecki (1966), A. Sobiczewski et al. (1966)

# who is she?

7

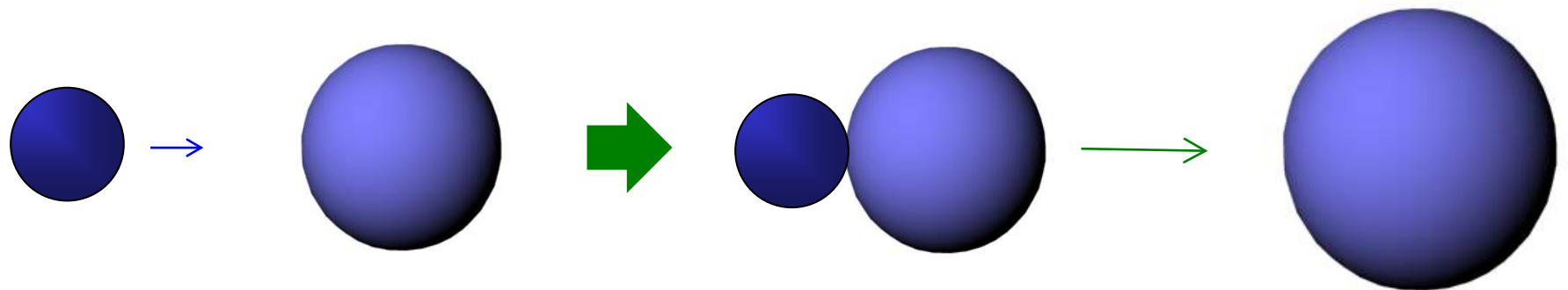
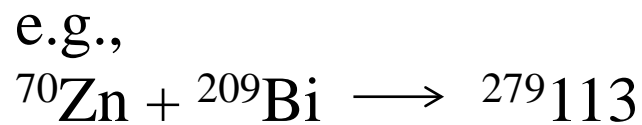
87	88		104	105	106	107	108	109	110	111	112	113	114	115	116	117	118
Fr	Ra		Rf	Db	Sg	Bh	Hs	Mt	Ds	Rg	Cn	Uut	Fl	Uup	Lv	Uus	Uuo

Z=110	Darmstadtium (Ds)	1994	Germany
Z=111	Roentgenium (Rg)	1994	Germany
Z=112	Copernicium (Cn)	1996	Germany
Z=113	<b>Nihonium (Nh)</b>	2003	Russia / 2004 Japan
Z=114	Flerovium (Fl)	1999	Russia
Z=115	<b>Moscovium (Mc)</b>	2003	Russia
Z=116	Livermorium (Lv)	2000	Russia
Z=117	<b>Tennessine (Ts)</b>	2010	Russia
Z=118	<b>Oganesson (Og)</b>	2002	Russia

113 <b>Nh</b> nihonium	115 <b>Mc</b> moscovium
117 <b>Ts</b> tennessine	118 <b>Og</b> oganesson

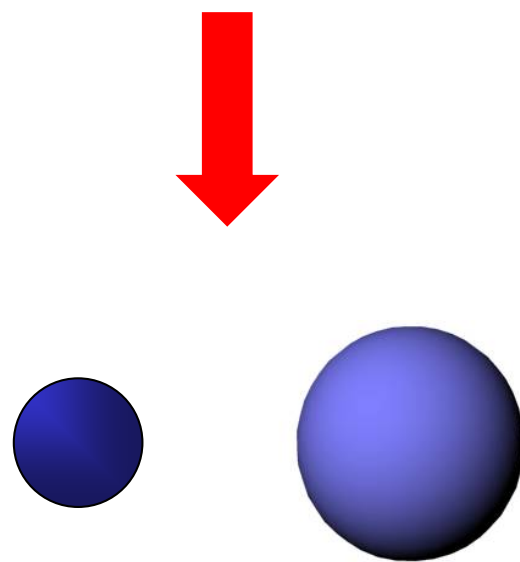
# How to synthesize SHE?

Nuclear fusion reactions



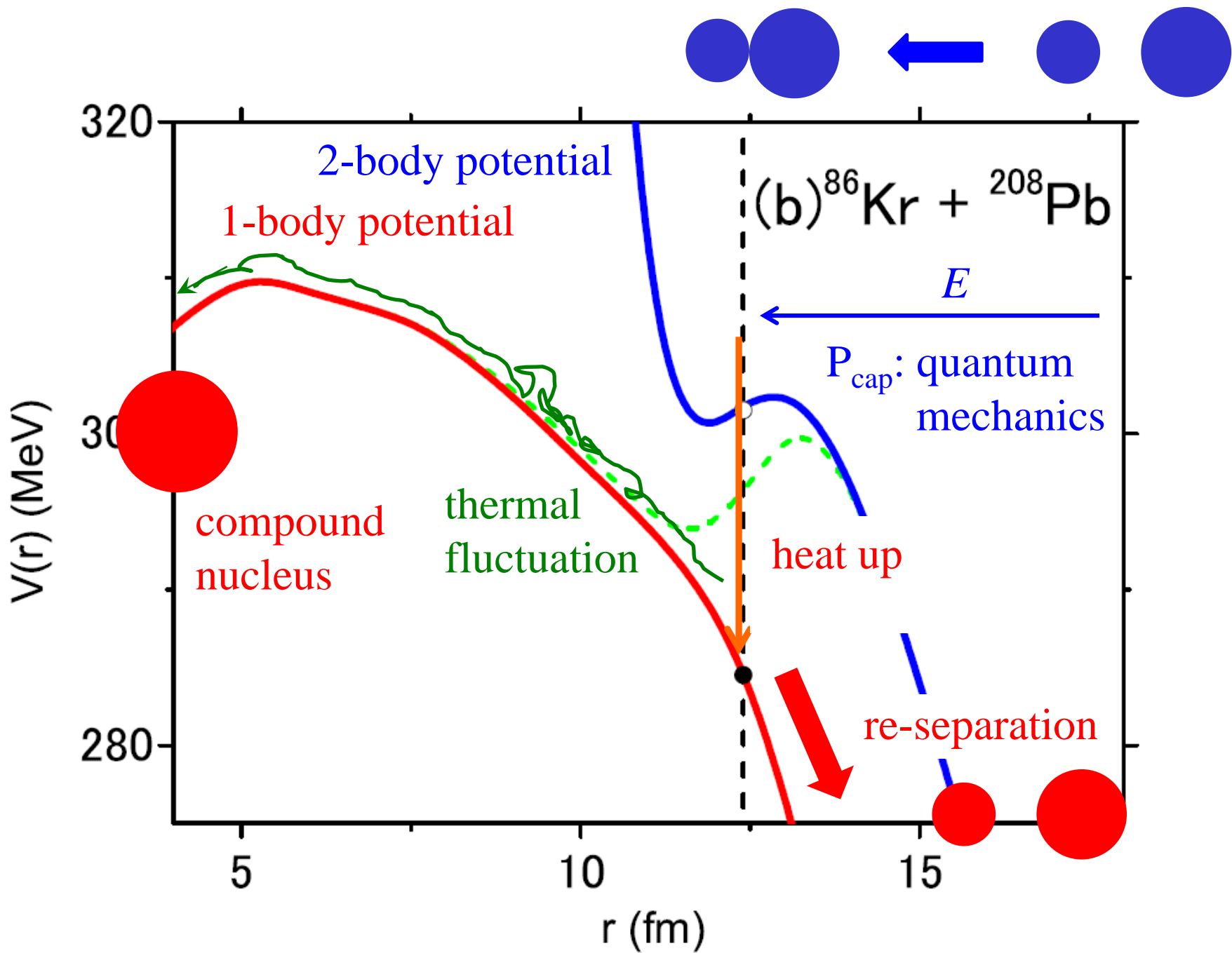
two positive charges  
repel each other

compound  
nucleus

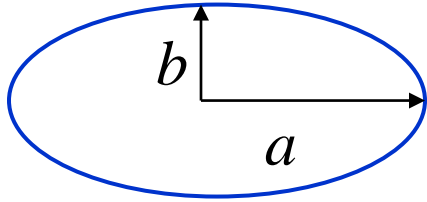


re-separation





## (note) fission barrier in the liquid drop model

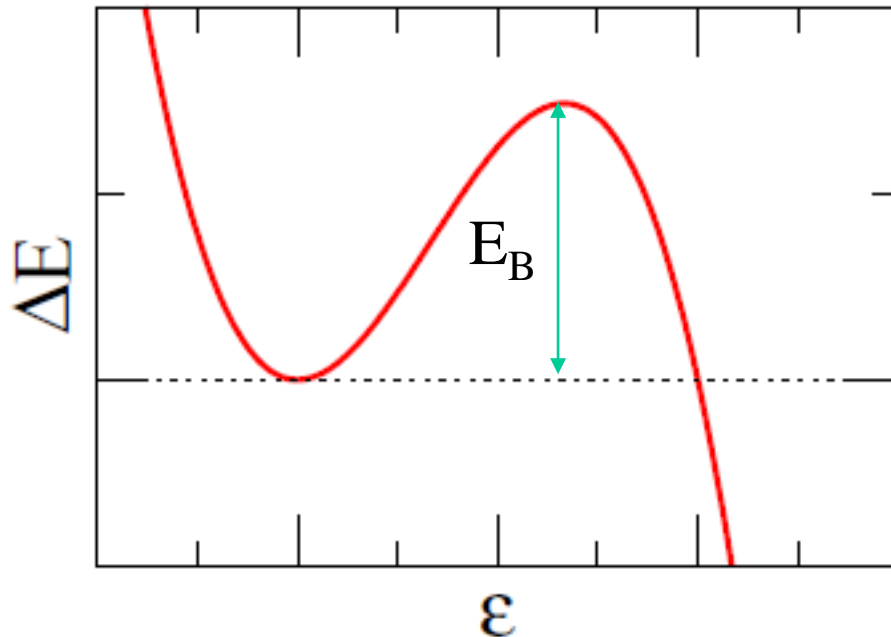


$$a = R \cdot (1 + \epsilon)$$

$$b = R \cdot (1 + \epsilon)^{-1/2}$$

$$ab^2 = R^3 = \text{constant}$$

$$\begin{aligned} \Delta E &= \Delta E_{\text{surf}} + \Delta E_{\text{coul}} \\ &= E_S^{(0)} \left\{ \frac{2}{5}(1 - x)\epsilon^2 - \frac{4}{105}(1 + 2x)\epsilon^3 + \dots \right\} \end{aligned}$$

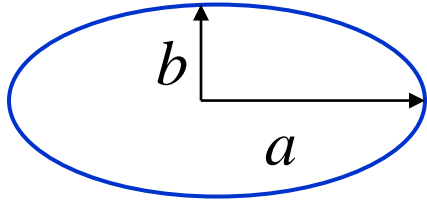


$$E_S^{(0)} = +a_S A^{2/3}$$

$$x \equiv \frac{E_C^{(0)}}{2E_S^{(0)}} = \frac{a_C}{2a_S} \cdot \frac{Z^2}{A} \sim \frac{1}{53.3} \cdot \frac{Z^2}{A}$$

$$E_C^{(0)} = a_C Z(Z - 1)/A^{1/3}$$

## (note) fission barrier in the liquid drop model

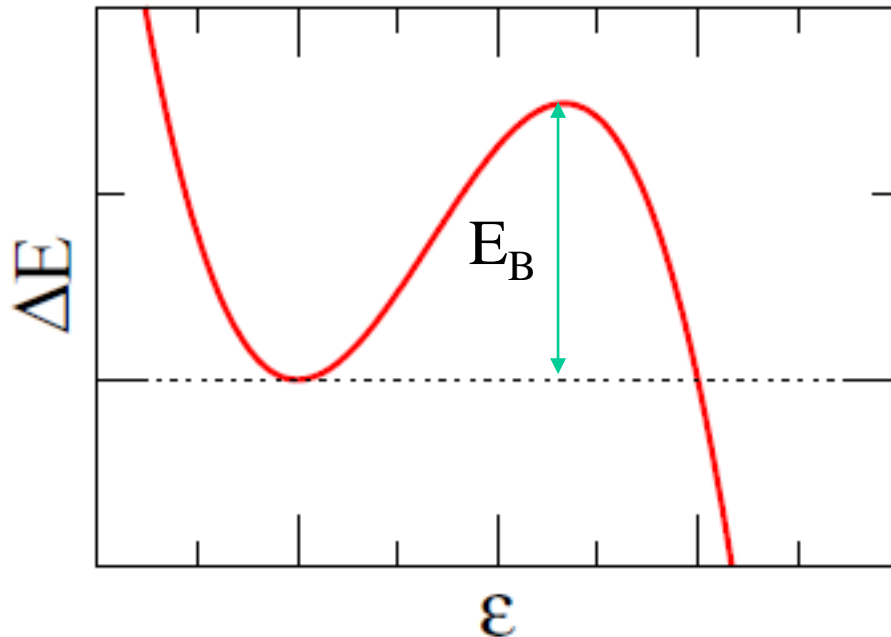


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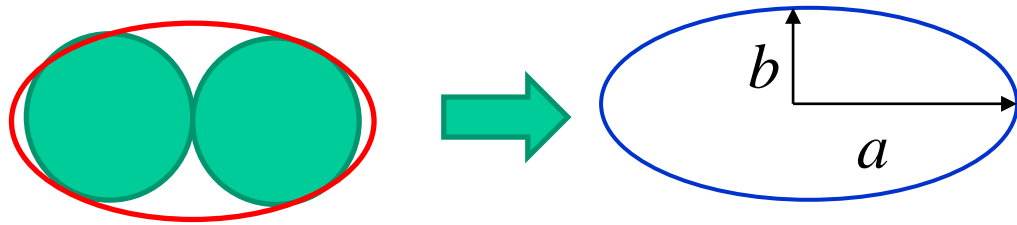


fission barrier:

$$\epsilon_B = \frac{21(1-x)}{3(1+2x)}$$

$$E_B = \frac{98}{15} \cdot \frac{(1-x)^3}{(1+2x)^2} \cdot E_S^{(0)}$$

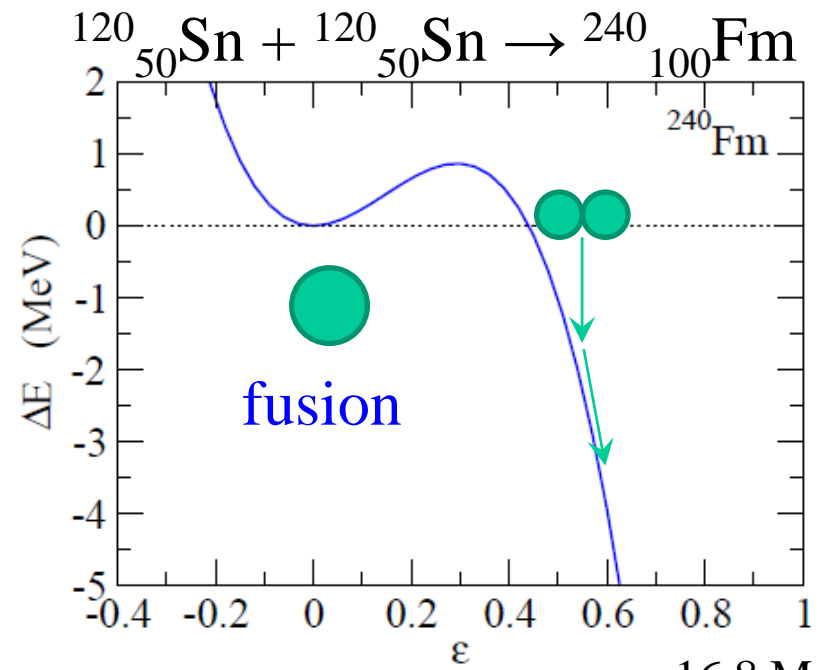
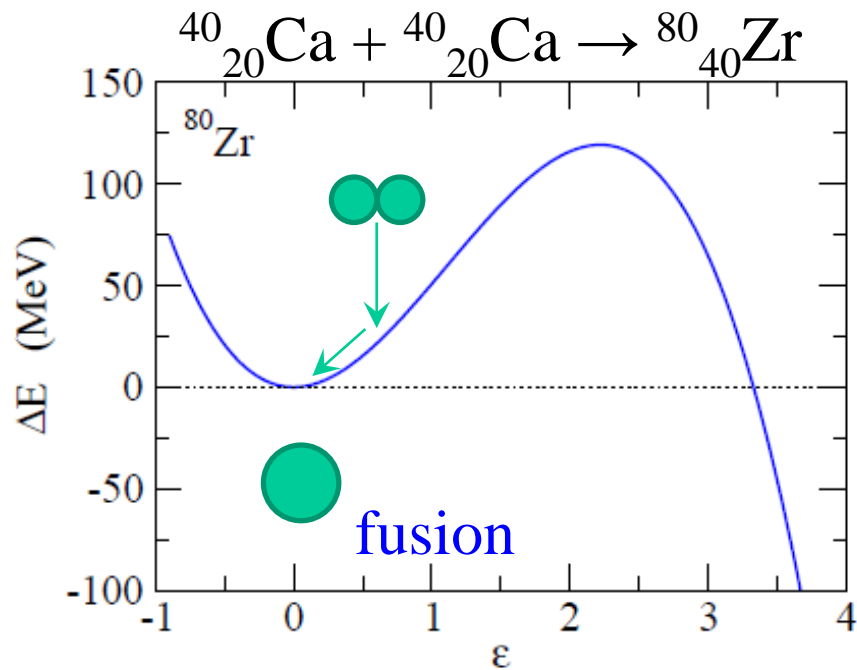
if two identical nuclei contact:



$$a = R_0 \cdot (1 + \epsilon)$$

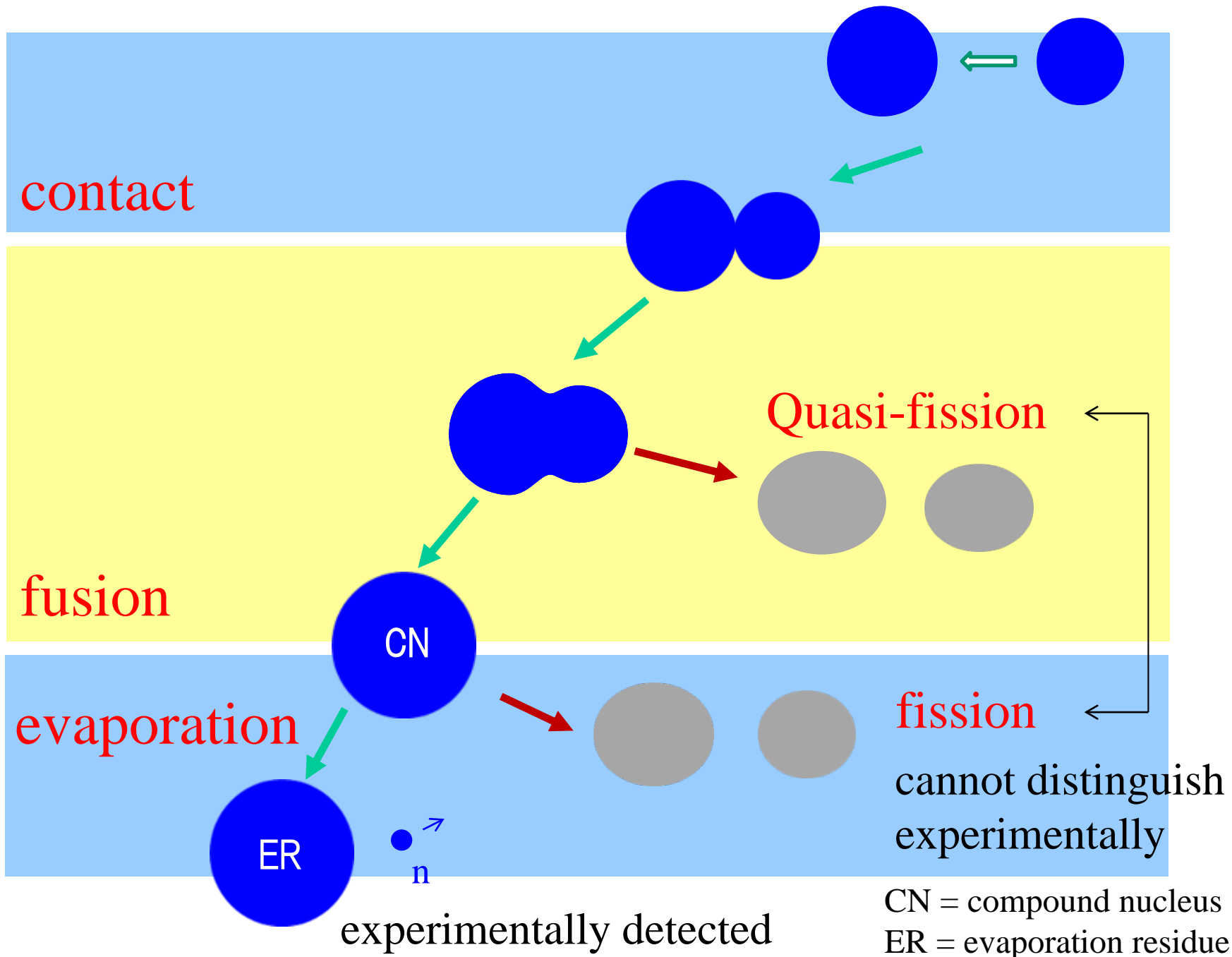
$$b = R_0 \cdot (1 + \epsilon)^{-1/2}$$

$$\frac{a}{b} \sim \frac{2R}{R} = 2 \rightarrow \epsilon \sim 0.587$$

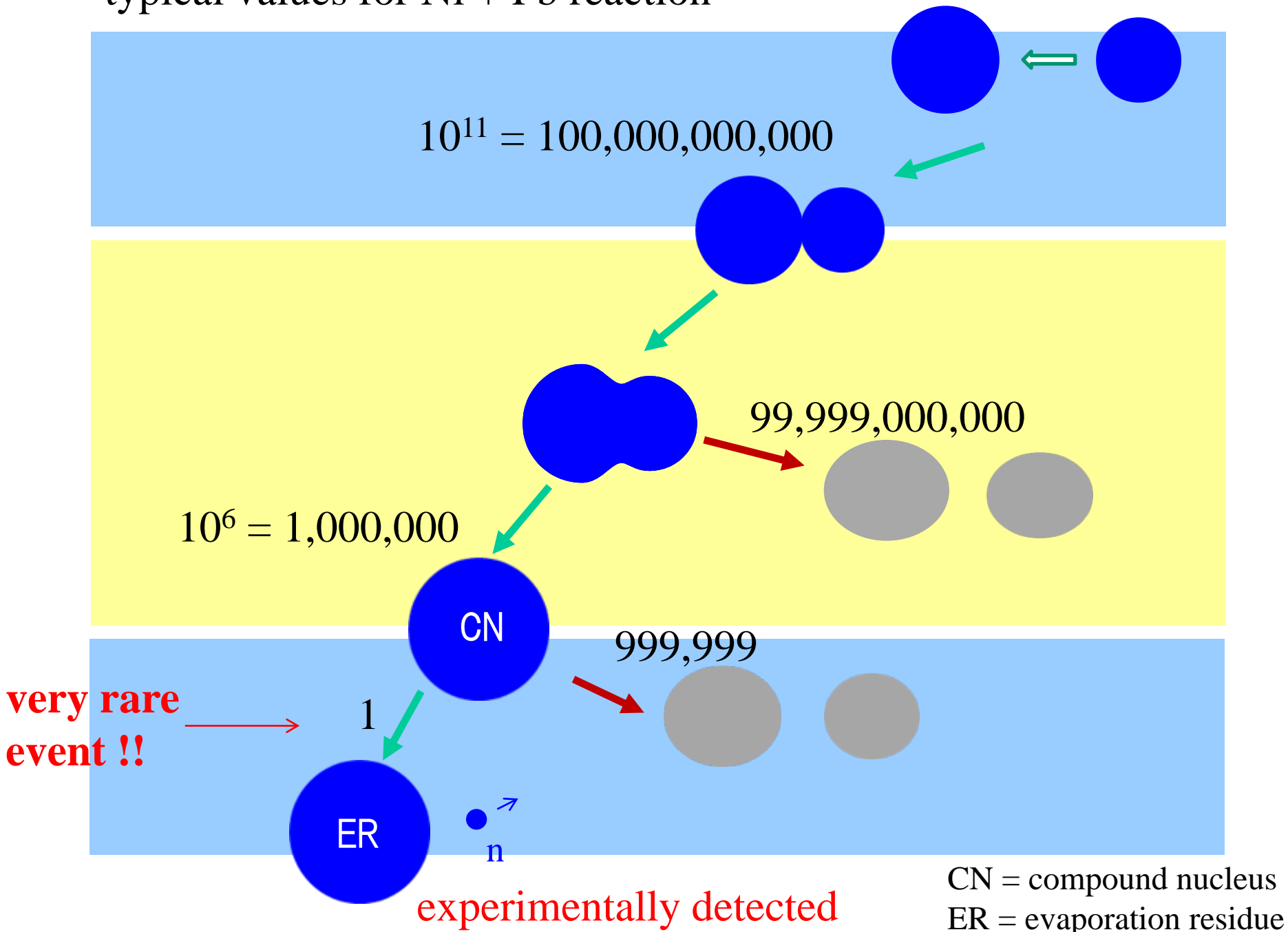


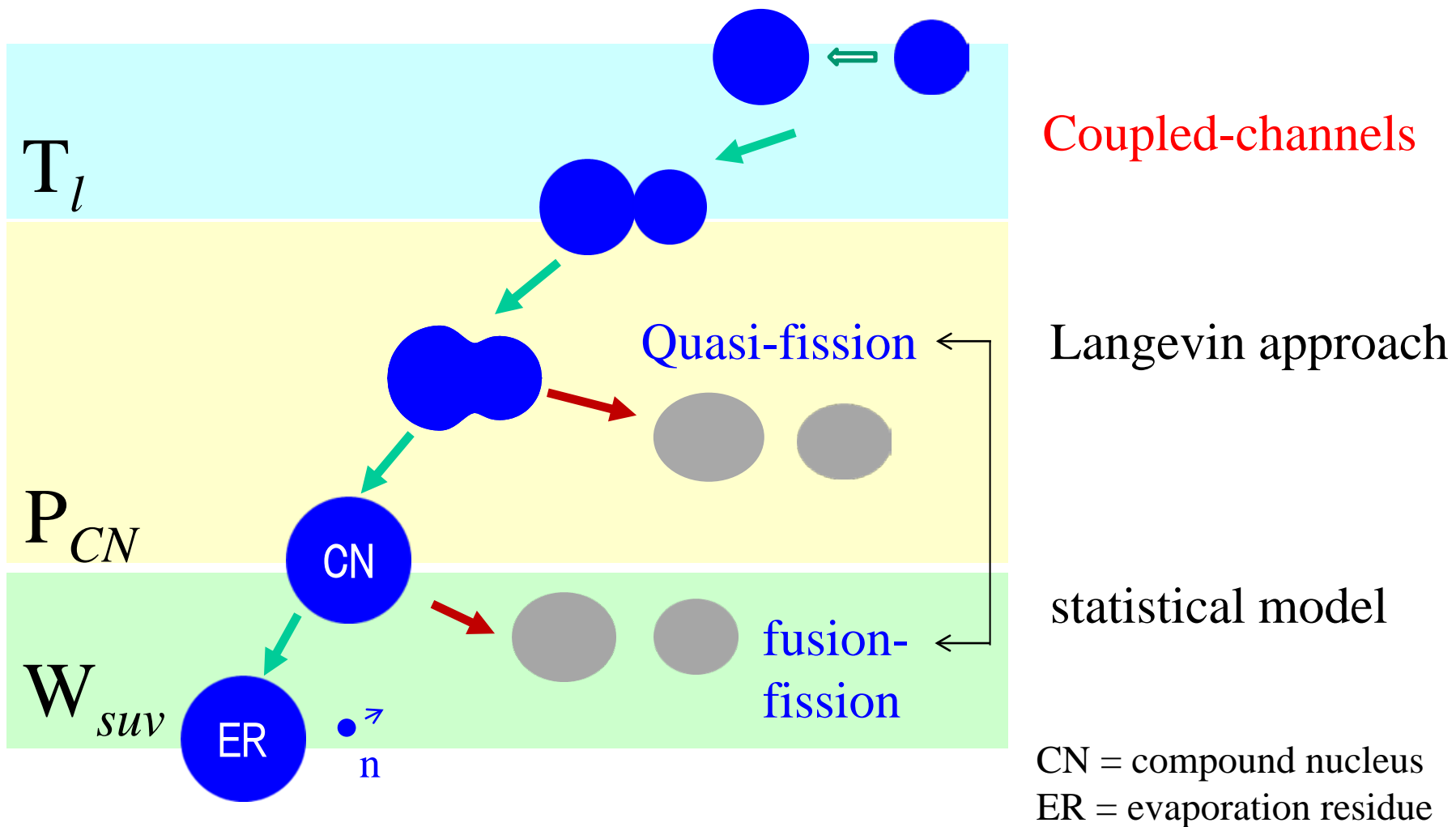
threshold:  $Z_1 \cdot Z_2 = 1600 \sim 1800$

$a_s = 16.8 \text{ MeV}$   
 $a_c = 0.72 \text{ MeV}$

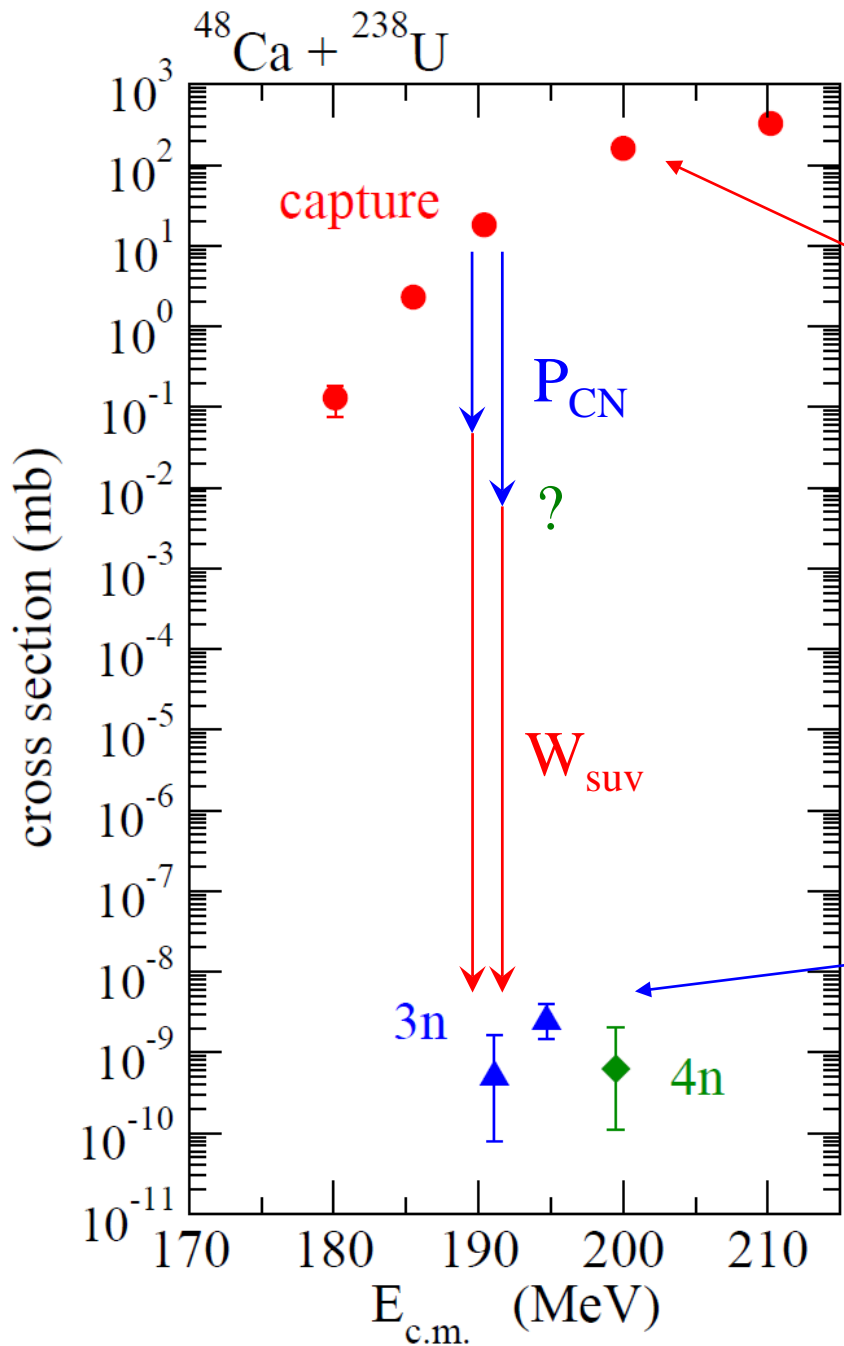


# typical values for Ni + Pb reaction





$$\sigma_{\text{ER}}(E) = \frac{\pi}{k^2} \sum_l (2l + 1) T_l(E) P_{\text{CN}}(E, l) W_{\text{suv}}(E^*, l)$$



no experimental data for  $P_{\text{CN}}$

$$\sigma_{\text{cap}}(E) = \frac{\pi}{k^2} \sum_l (2l + 1) T_l(E)$$

~~$$\sigma_{\text{CN}}(E) = \frac{\pi}{k^2} \sum_l (2l + 1) T_l(E) \times P_{\text{CN}}$$~~

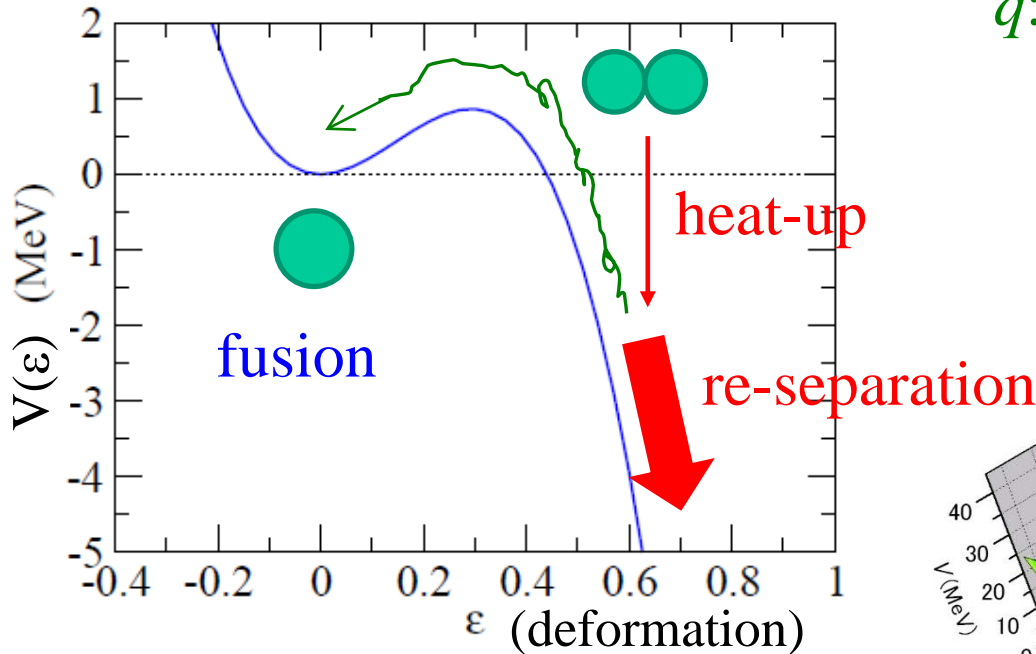
not available

$$\sigma_{\text{ER}}(E) = \frac{\pi}{k^2} \sum_l (2l + 1) T_l(E) \times P_{\text{CN}} \cdot W_{\text{suv}}$$

large uncertainties



# Langevin approach



## thermal fluctuation

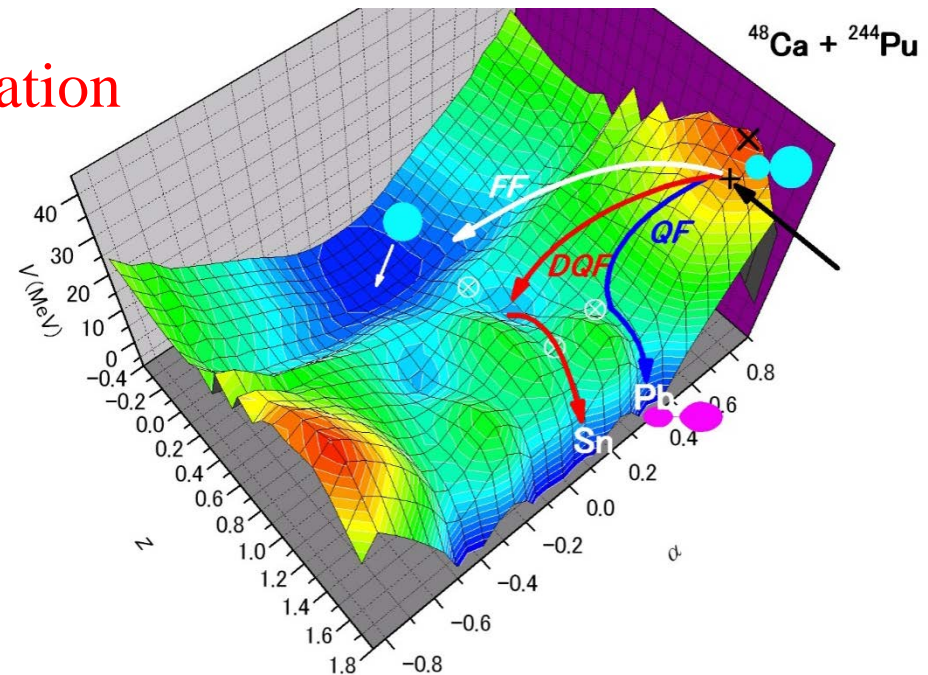
→ Langevin method  
(Brownian method)

$$m \frac{d^2 q}{dt^2} = - \frac{dV(q)}{dq} - \gamma \frac{dq}{dt} + R(t)$$

$\gamma$ : friction coefficient  
 $R(t)$ : random force

## multi-dimensional extention

- $q$ : ▪ internuclear separation,
- deformation,
- asymmetry of the two fragments



- ✓ Abe, Wada, Bao
- ✓ Aritomo, Ohta
- ✓ Zagrebaev

# Theory: Lagenvin approach

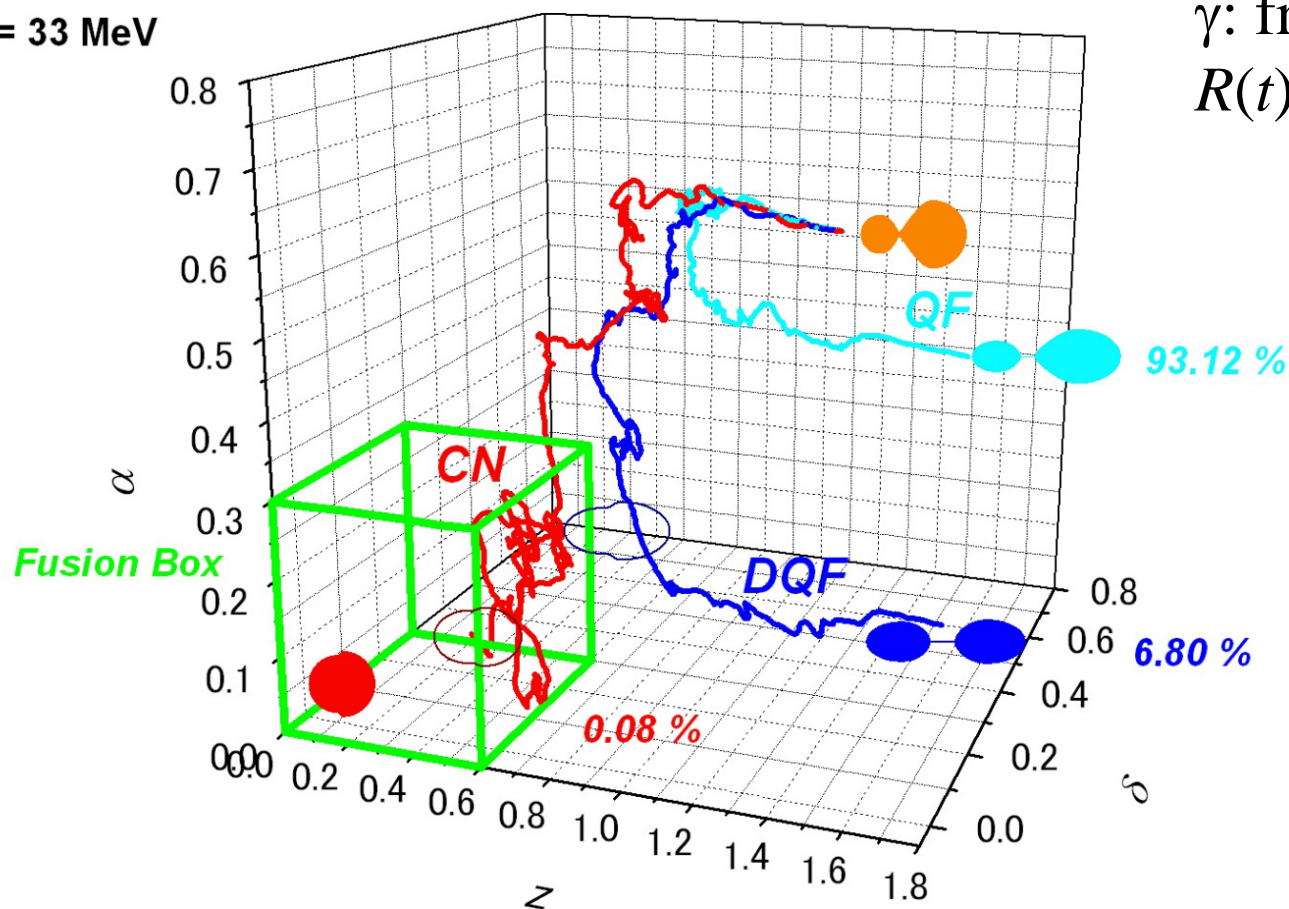
multi-dimensional extension of:

$$m \frac{d^2 q}{dt^2} = - \frac{dV(q)}{dq} - \gamma \frac{dq}{dt} + R(t)$$



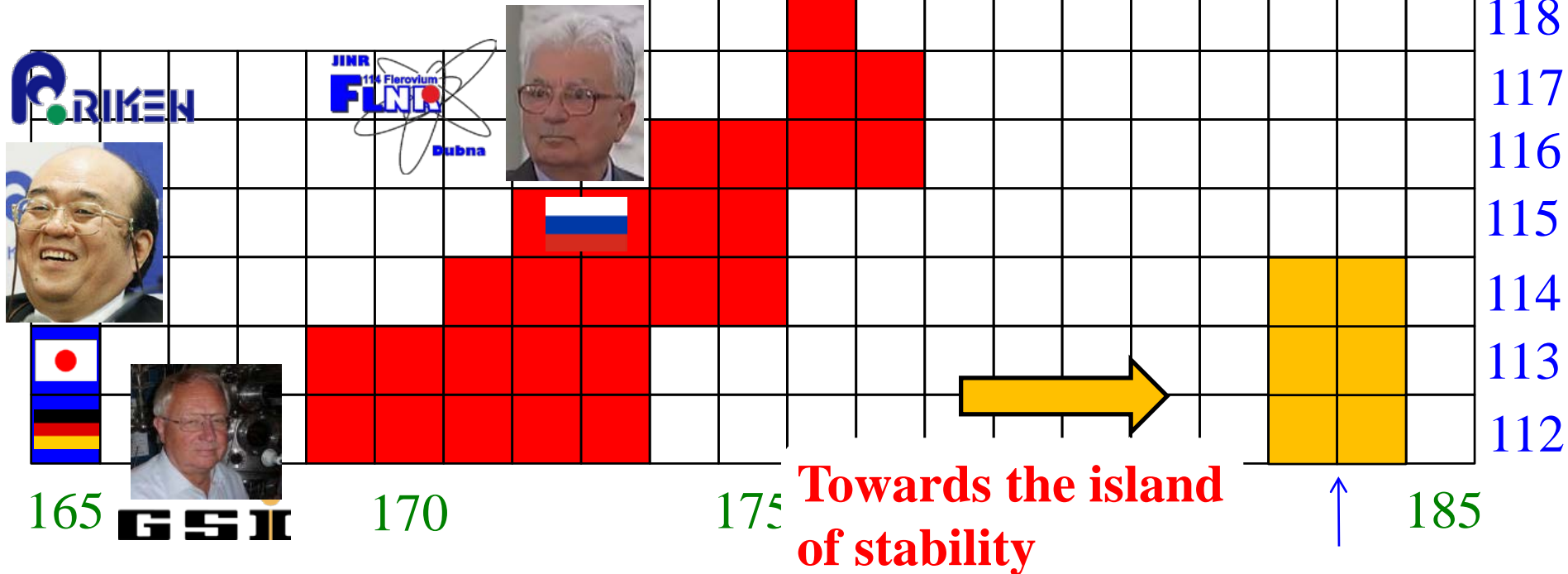
$E^* = 33 \text{ MeV}$

$\gamma$ : friction coefficient  
 $R(t)$ : random force



## Future directions

Superheavy elements synthesized so far

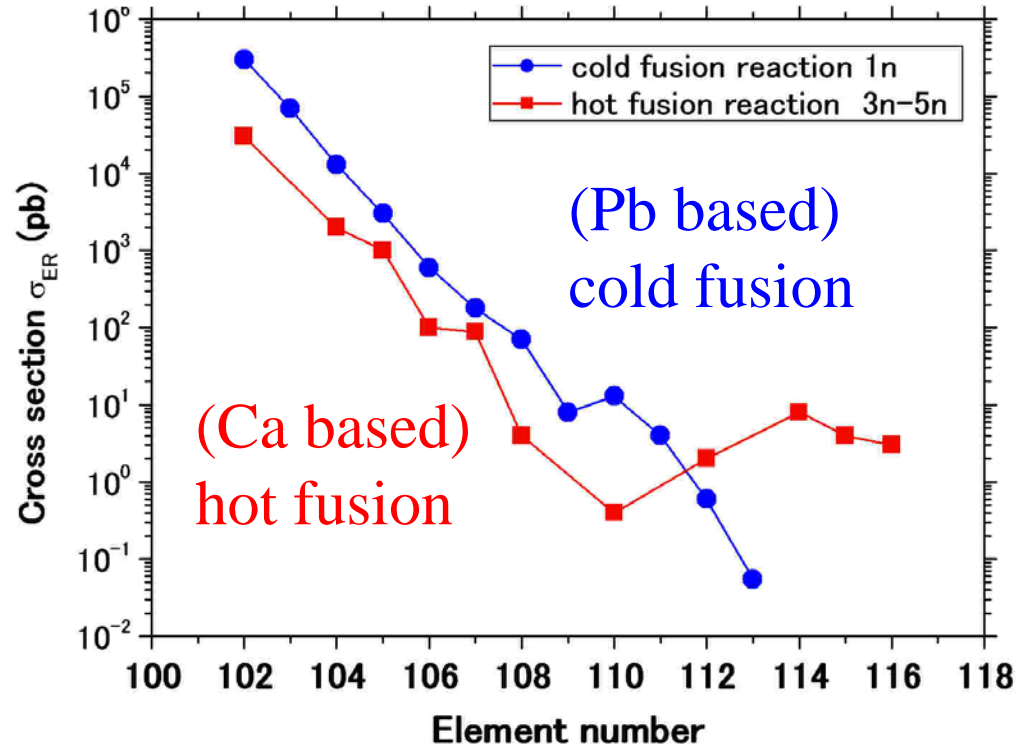
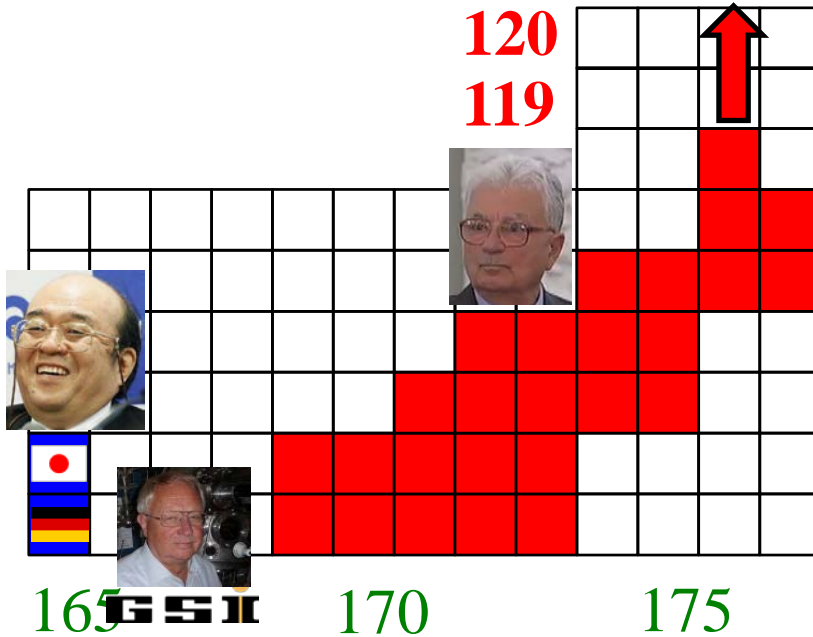


## Theoretical issues:

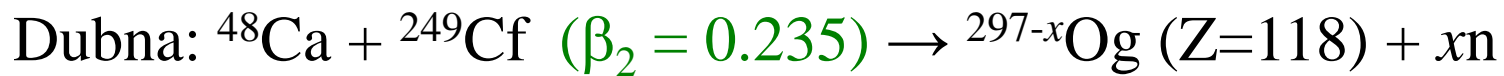
- to understand the reaction dynamics
- to make a reliable theoretical prediction for fusion cross sections

# Hot fusion for $Z = 119$ and $120$

## Towards $Z=119$ and $120$ isotopes

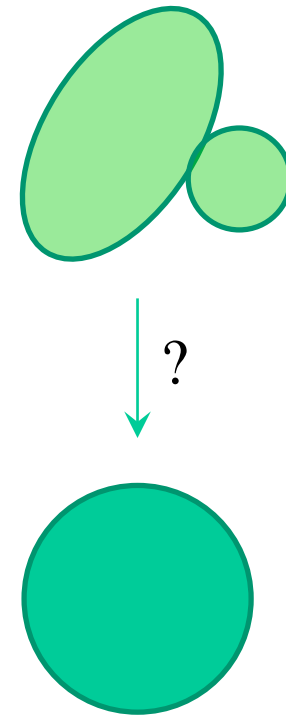
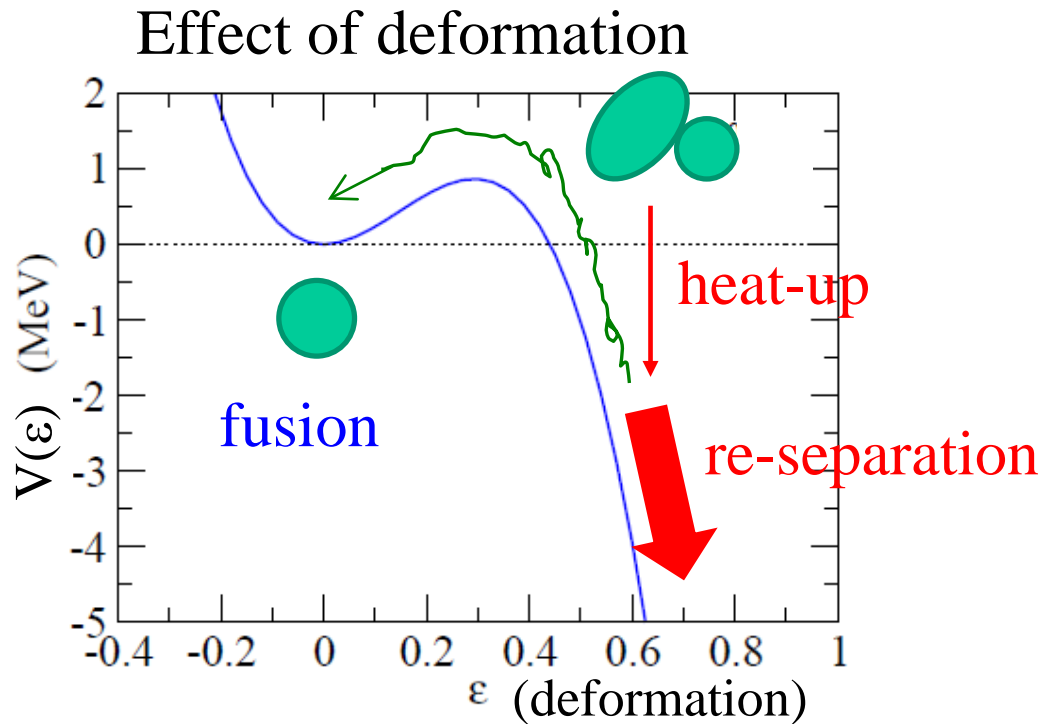


hot fusion:  $^{48}\text{Ca} + \text{actinide targets}$



role of deformation?

# Hot fusion: $^{48}\text{Ca}$ + deformed actinide target



## Open problems

- how is the shape evolved to a compound nucleus?
- Deformation: a quantum effect
  - how does the deformation disappear during heat-up?

Quantum friction

## Quantum friction

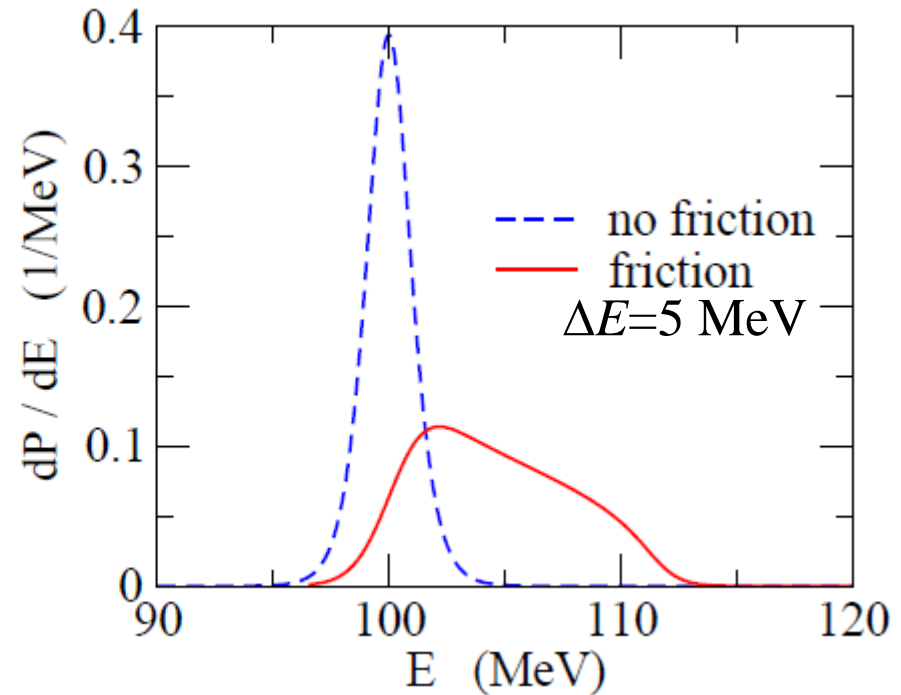
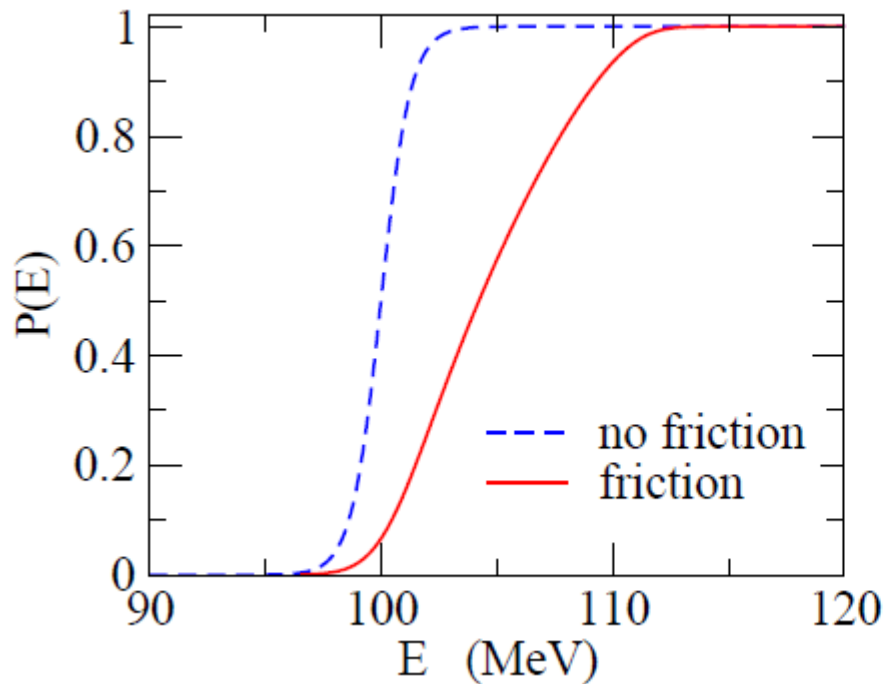
classical eq. of motion  $\dot{p} = -V'(x) - \gamma p$

a quantization: Kanai model E. Kanai, PTP 3 (1948) 440

$$H = \frac{p^2}{2m} + V(x) \rightarrow \frac{\pi^2}{2m} e^{-\gamma t} + e^{\gamma t} V(x) \quad (\pi = e^{\gamma t} p)$$

(a quantal Hamiltonian which reproduces the classical eq. of motion)

time-dep. wave packet approach

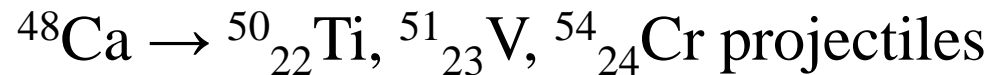
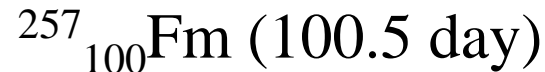
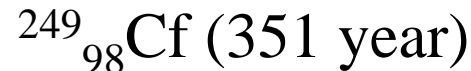
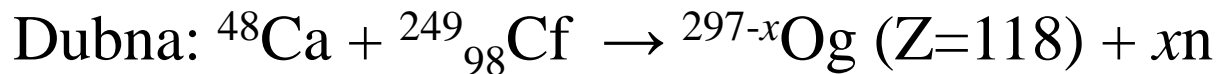


## ➤ Towards Z=119 and 120 nuclei

### Another issue



the targets: not available with sufficient amounts

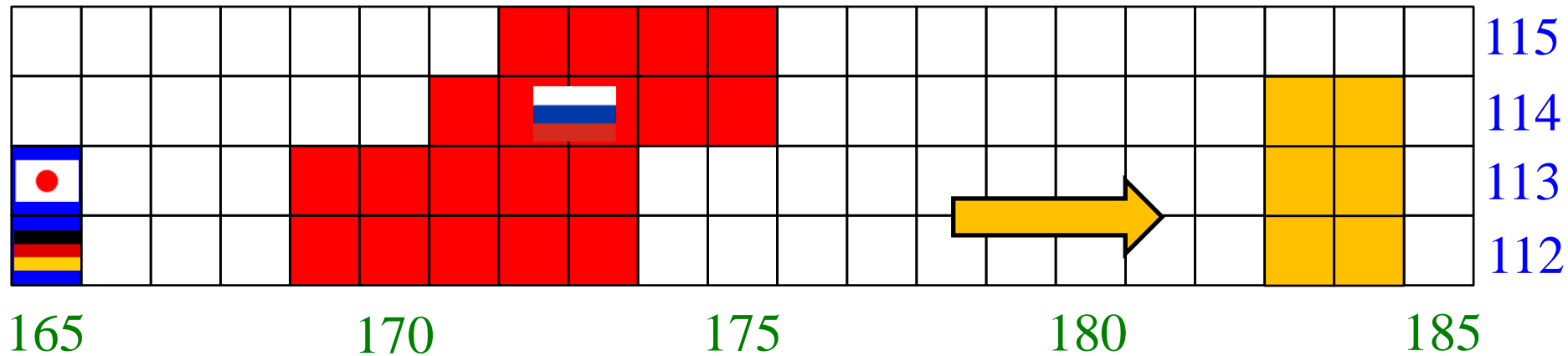


cf.  ${}^{46}_{21}\text{Sc}_{25}$  : relatively small neutron number

how much will fusion cross sections be reduced?

nobody still knows

# Towards the island of stability



neutron-rich beams: indispensable

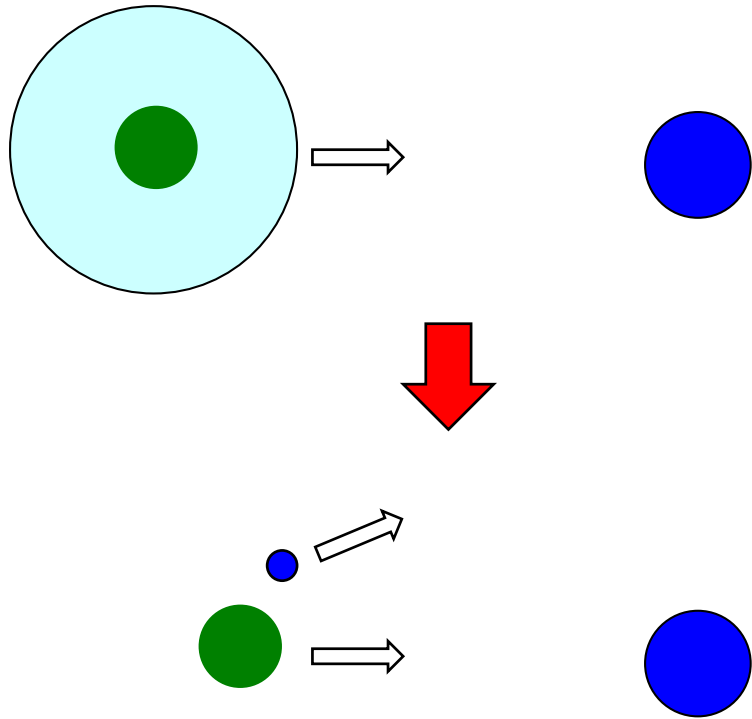
- how to deal with low beam intensity?
- reaction dynamics of neutron-rich beams?
  - ✓ capture: role of breakup and (multi-neutron) transfer?
  - ✓ diffusion: neutron emission during a shape evolution?
  - ✓ survival: validity of the statistical model?

structure of exotic nuclei

more studies are required



# Fusion of halo nuclei



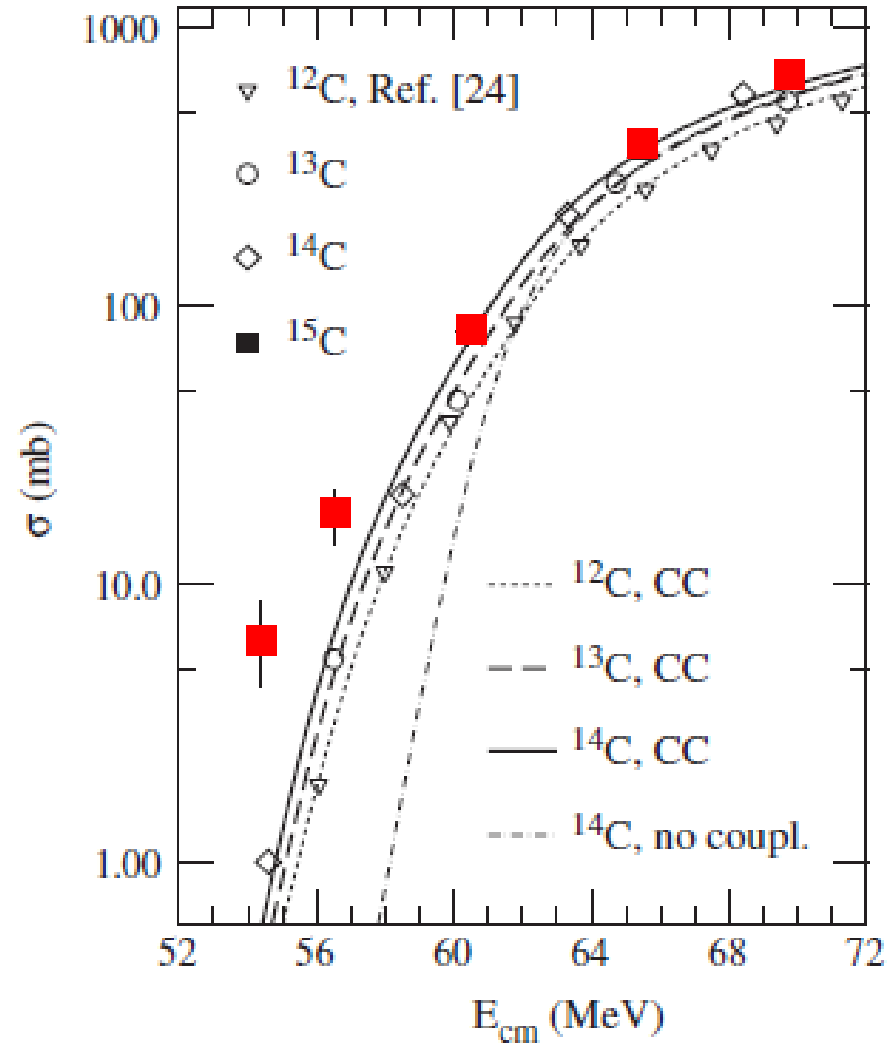
1. Lowering of potential barrier due to a halo structure

→ enhancement

2. effect of breakup

3. effect of transfer: equally important

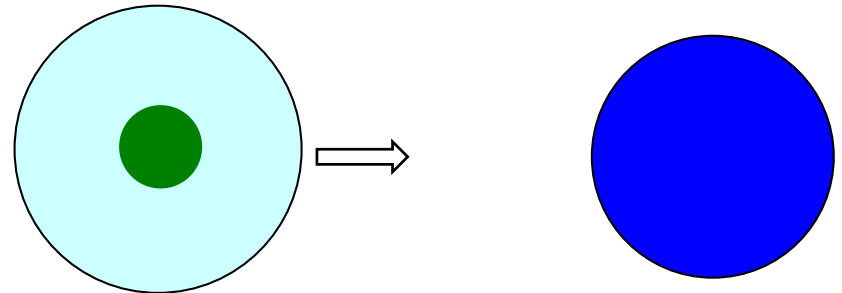
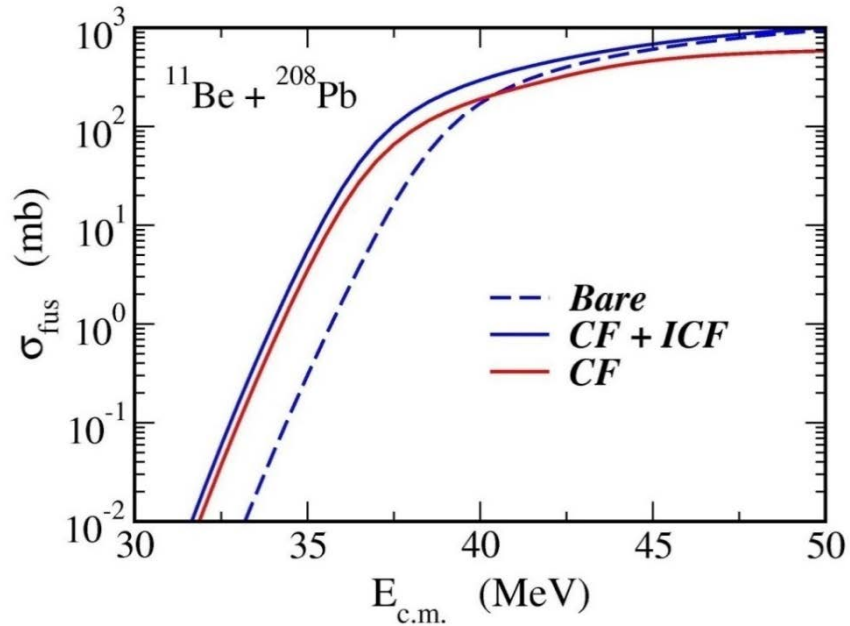
$^{12,13,14,15}\text{C} + ^{232}\text{Th}$



M. Alcorta et al.,  
PRL106('11)172701

## Future directions - 2

- Towards the island of stability  
neutron-rich beams: indispensable



simultaneous treatment  
of **breakup** and **transfer**

→ an important future problem

K. Hagino, A. Vitturi, C.H. Dasso,  
and S.M. Lenzi, Phys. Rev. C61 ('00) 037602

# Summary

## Heavy-ion fusion reactions around the Coulomb barrier

- ✓ Strong interplay between nuclear structure and reaction
- ✓ Quantum tunneling with various intrinsic degrees of freedom
- ✓ coupled-channels approach

## Remaining challenges

- ✓ microscopic understanding of heavy-ion fusion reactions

## Future perspectives: superheavy elements

- ✓ how to reduce theoretical uncertainties?
- ✓ Towards heavier SHE ( $Z = 119, 120$ )
- ✓ Towards the island of stability

investigations of physics of SHE with neutron-rich nuclei as a keyword