

Toward Construction of Supergravity Superstrata States

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第6回静岡素粒子集中セミナー

stra·tum [stréitəm | strá:-]

【ラテン語「広がったもの」の意から】

一名C(總)-ta [-tə], ~s)

1 【地質】地層; 層.

2 層, 階級.

*

Based on collaboration with:

Iosif Bena (Saclay)

Jan de Boer (Amsterdam)

Stefano Giusto (Padova)

Rodolfo Russo (Queen Mary)

Nicholas Warner (USC)

1412.xxxx, 1406.4506, 1307.3115,
1209.6056, 1110.2781, 1107.2650, 1004.2521

Road to Superstratum

Black hole puzzles

- ▶ Entropy (microstate) problem

$$S_{\text{BH}} = \frac{A}{4G_N}$$

Schwarzschild: $S_{\text{BH}} = 10^{77} (M/M_{\odot})^2$

Cf. No-hair theorem: $e^S = 1$

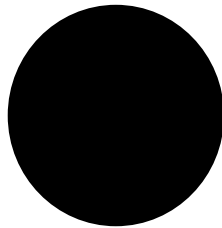
- ▶ Information paradox
- ▶ Firewall

...

Microstate counting

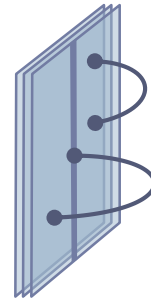
► Strominger-Vafa 1996:

Gravity
(strong coupling)



S_{BH}

Field theory
(weak coupling)



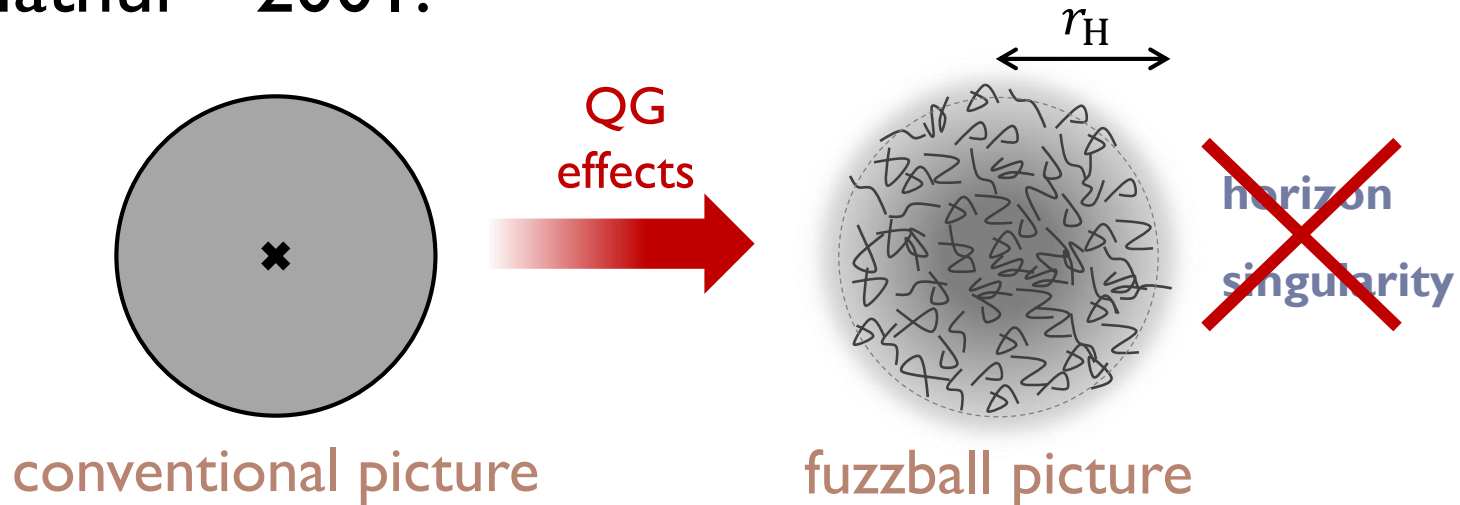
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$S_{\text{CFT}} = \log N_{\text{CFT}}$

→ Cool, but what's gravity picture of the microstates?

Fuzzball proposal

▶ Mathur ~2001:



- ▶ Microstates = QG/string “fuzz”
- ▶ Not describable within sugra in general (some hope for supersymmetric states)

2-charge system (1)

- ▶ Fuzzball proposal was made based on this system

- ▶ Canonical rep:

IIB on $S^1_5 \times T^4_{6789}$

	1	2	3	4	5	6	7	8	9
N_1 D1	•	•	•	•	○	~	~	~	~
N_2 D5	•	•	•	•	○	○	○	○	○

S^1 (bracketed over columns 5-6) T^4 (bracketed over columns 6-9)

- ▶ Supersymmetric (8 supercharges, 1/4 BPS)
- ▶ Large microscopic entropy:

$$S_{\text{micro}} = 2\sqrt{2}\pi\sqrt{N_1 N_2}$$

- ▶ Horizon vanishes classically

2-charge system (2)

- ▶ SUGRA microstates known:

“Microstate geometries” [Lunin-Mathur 2001]

- ▶ It is a *supertube*:

$$D1(5) + D5(56789) \rightarrow \text{KKM}(\lambda 6789, 5)$$



+



→



arbitrary curve

$$\vec{x} = \vec{f}(\lambda) \in \mathbb{R}_{1234}^4$$

- ▶ Reproduces entropy; $S \sim \sqrt{N_1 N_2}$

- ▶ Dictionary b/w CFT & sugra microstates known

RR gnd state \longleftrightarrow curve $\vec{f}(\lambda)$

2-charge system: summary

Fuzzball works for 2-charge sys,
which however is not a black hole.



Need to go to system with finite horizon
to prove / disprove fuzzball conjecture

3-charge system

- ▶ Susy BH in 5D (4 supercharges)
- ▶ Canonical rep [Strominger-Vafa 1996]

IIB on
 $S_5^1 \times T_{6789}^4$

	1	2	3	4	5	6	7	8	9
N_1 D1	•	•	•	•	○	~	~	~	~
N_2 D5	•	•	•	•	○	○	○	○	○
N_3 P	•	•	•	•	○	~	~	~	~

- ▶ Decoupling $\rightarrow AdS_3 \times S^3 \times T^4$ / D1-D5 CFT
- ▶ Macroscopic entropy: $S \sim \sqrt{N_1 N_2 N_3}$

4-charge system

- ▶ Susy BH in 4D (4 supercharges)
- ▶ Canonical rep [Maldacena-Strominger-Witten 1997]

M on T_{456789}^6

		1	2	3	4	5	6	7	8	9	A
N_1 M5		•	•	•	~	~	○	○	○	○	○
N_2 M5		•	•	•	○	○	~	~	○	○	○
N_3 M5		•	•	•	○	○	○	○	~	~	○
N_4 P		•	•	•	~	~	~	~	~	~	○

- ▶ Decoupling $\rightarrow AdS_3 \times S^2 \times T^6$ / MSW CFT
- ▶ Macroscopic entropy: $S \sim \sqrt{N_1 N_2 N_3 N_4}$

5D/4D ansatz (1)

*Want to find gravity microstates
for 3- & 4-charge systems*

(although there is no guarantee that they are describable within sugra...)

Start from 3-charge system

IIB / T_{56789}^5 D1(5), D5(56789), P(5)

↓ T_5, T_6, T_7

IIA / T_{56789}^5 D2(67), D2(89), F1(5)

↓ lift along x^A

M / T_{56789A}^6 M2(67), M2(89), M2(5A)

Nicely symmetric

5D/4D ansatz (2)

$$M / T_{56789A}^6 \quad M2(67), M2(89), M2(5A)$$

- ▶ Bena-Warner, Gauntlett-Gutowski 2004:

Classified all solutions in 5D preserving same susy

- 4D hyperkahler base \mathcal{B}_4
- Funcs & forms defined on \mathcal{B}_4

↓ Technically difficult.
Assume $U(1)$ symmetry in B_4

- 3D flat base \mathbb{R}^3
- Harmonic funcs on \mathbb{R}^3

$$H = (V, K^I, L_I, M), \quad H = h + \sum_p \frac{Q_p}{|r - r_p|}$$

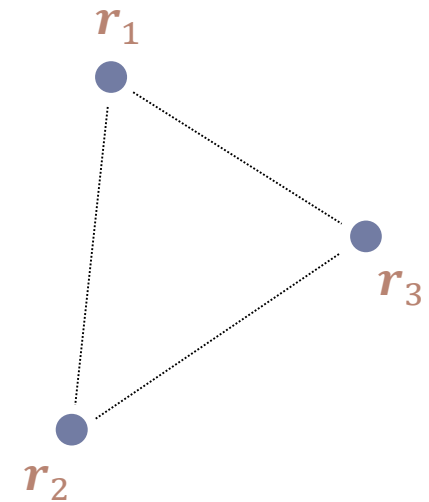
5D/4D ansatz (3)

- ▶ Multi-center config of BHs & BRs in 5D

\mathbb{R}^3

$$H = h + \sum_p \frac{Q_p}{|\mathbf{r} - \mathbf{r}_p|}$$

- ▶ Positions \mathbf{r}_p satisfy “bubbling eq”
- ▶ Large family of solutions
- ▶ Reducing on $U(1)$, it becomes 4D BHs (same as Bates-Denef 2003)



$$H = (V, K^I, L_I, M)$$

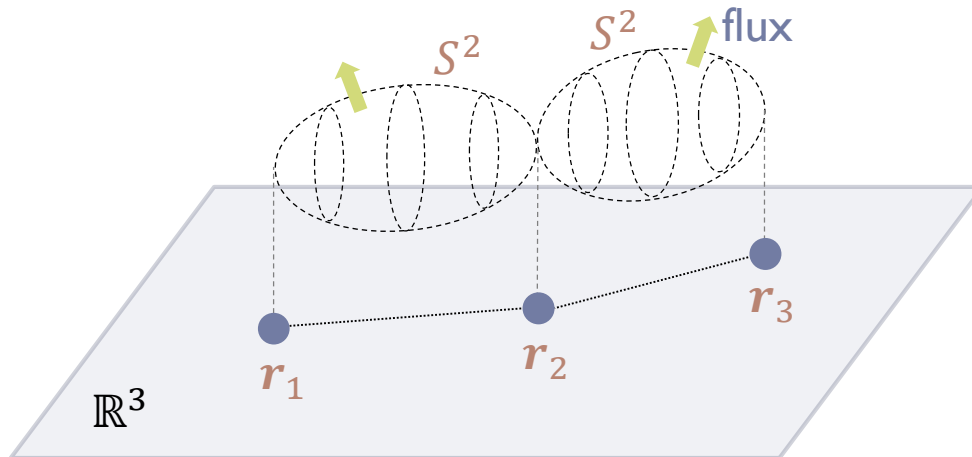
D6 D4 D2 D0

Microstates in 5D/4D ansatz (1)

- ▶ Tune charges → *Regular & horizonless solutions!*

[Bena-Warner 2006] [Berglund-Gimon-Levi 2006]

D6 = KKM with fluxes



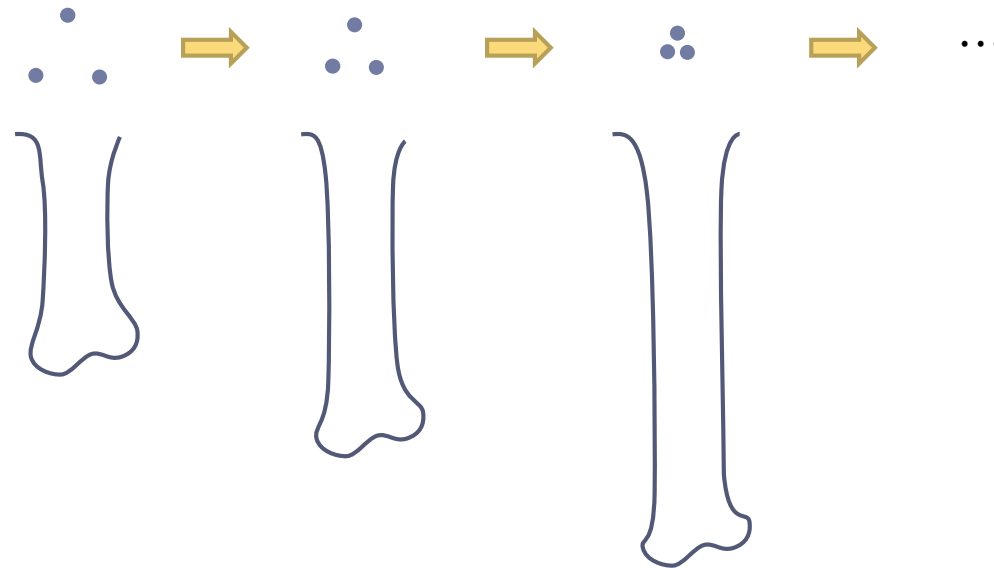
Mechanism to support horizon-sized structure!
Cf. Firewall

- Microstate geometries of 3- and 4-charge black holes 😊

Microstates in 5D/4D ansatz (2)

- ▶ Various nice properties 😊

- ▶ Scaling solutions [BW et al., 2006, 2007]



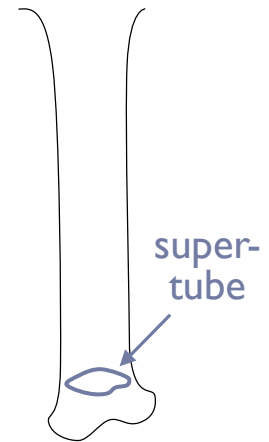
- ▶ Gap expected from CFT: $\Delta E \sim 1/N_1 N_2$

Microstates in 5D/4D ansatz (3)

The real question: are there enough?

- ▶ 4-charge sys [de Boer et al., 2008-09]
 - ▶ Quantization of D6- $\overline{D6}$ -D0 config → *much less entropy* ☹️
- ▶ 3-charge sys (+ fluctuating supertube)
 - ▶ Entropy enhancement mechanism [BW et al., 2008]
→ Much more entropy?
 - ▶ An estimate [BW et al., 2010]

$$S \sim N^{\frac{5}{4}} \ll N^{\frac{3}{2}} \quad \text{Parametrically smaller ☹️}$$



Summary

4D/5D ansatz solutions *are*
black hole microstates,
but they are *too few*.

Possibilities:

- A) SUGRA is not enough
- B) Need more general ansatz

Superstratum

Exotic branes & double bubbling

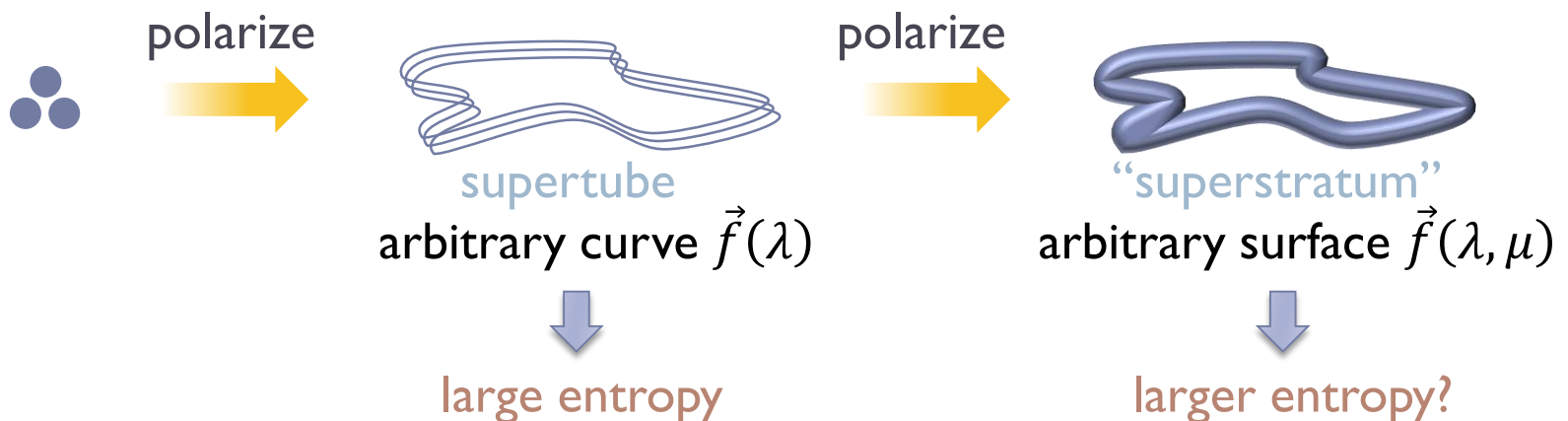
[de Boer+MS 2010]:

1. Ordinary branes can polarize into *non-geometric* (exotic) branes

e.g., MSW: $M5(6789A) + M5(4589A) \rightarrow 5^3(\lambda 4567, 89A)$

→ Microstates must involve *non-geometric* configs

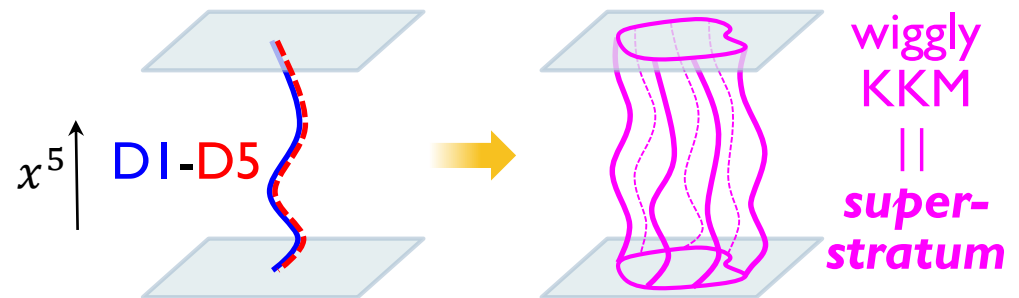
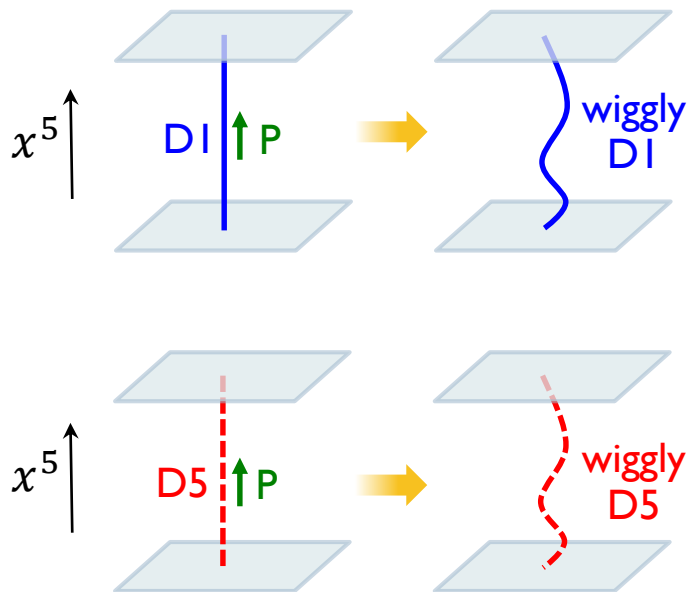
2. Double bubbling



Implication for D1-D5-P

[Bena+de Boer+
MS+Warner | 107.2650]

$D1(5)$ \longrightarrow $D5(\lambda 6789)$ \longrightarrow $5^2_3(\mu 6789, \lambda 5)$
 $D5(56789)$ \longrightarrow $D1(\lambda)$ \longrightarrow $1^6_3(\mu, \lambda 56789)$
 $P(5)$ \longrightarrow $KKM(\lambda 6789, 5)$ \longrightarrow $KKM(\mu 6789, \lambda)$



\rightarrow Need fluctuation along x^5

Recall we T-dualized along 5 to get to the frame of 4D/5D ansatz...

Implication for D1-D5-P



- ▶ Expect *geometric superstrata* sol'ns in 6D sugra
 - parametrized by a function of *two* variables
- ▶ *Generic* solutions must be *non-geometric* superstrata
- ▶ But we have more intuition for geometric ones



Study susy solutions in 6D sugra!

Susy solutions in 6D sugra

[Bena+Giusto+MS+Warner 1110.2781]

6D theory

- ▶ 6D $\mathcal{N} = 2$ sugra with a vector multiplet
- ▶ Bosonic fields
 - ▶ Metric $g_{\mu\nu}$
 - ▶ Dilaton ϕ
 - ▶ 2-form B_2 , field strength $G_3 = dB_2$
- ▶ IIB on T_{6789}^4 :
 - D1(5) \rightarrow I-brane coupled to B_2
 - D5(56789) \rightarrow I-brane coupled to \tilde{B}_2

Susy sol'n (1): Base

6D spacetime: (u, v, x^m) u : isometry, $v \sim x^5$
 x^m : 4D base

► 4D base $\mathcal{B}^4(v)$: almost hyper-Kähler

$$ds_4^2 = h_{mn}(x, v) dx^m dx^n, \quad m, n = 1, 2, 3, 4$$

$\beta(x, v)$: 1-form (\leftrightarrow KKM)

$J^{(A)}(x, v)$, $A = 1, 2, 3$: almost HK 2-forms

$$J^{(A)m}_n J^{(B)n}_p = \epsilon^{ABC} J^{(C)m}_p - \delta^{AB} \delta^m_p$$

$$d_4 J^{(A)} = \partial_v (\beta \wedge J^{(A)}), \quad D \equiv d_4 - \beta \wedge \partial_v$$

Susy sol'n (2): Fields

► Fields on \mathcal{B}^4

Z_1 : scalar \leftrightarrow D1(v)

Z_2 : scalar \leftrightarrow D5($v6789$)

Θ_1 : 2-form \leftrightarrow D1(λ)

Θ_2 : 2-form \leftrightarrow D5($\lambda6789$)

ω : 1-form \leftrightarrow J

\mathcal{F} : scalar \leftrightarrow P(v)

► 6D fields

$$ds_6^2 = \frac{2}{\sqrt{Z_1 Z_2}} (dv + \beta) \left(du + \omega + \frac{1}{2} \mathcal{F} (dv + \beta) \right) - \sqrt{Z_1 Z_2} ds_4^2$$

$$G_3 = d \left[-\frac{1}{2} Z_1^{-1} (du + \omega) \wedge (dv + \beta) \right] + \frac{1}{2} *_4 (DZ_2 + \dot{\beta} Z_2) + (dv + \beta) \wedge \Theta_1$$

$$e^{\sqrt{2}\phi} = \sqrt{Z_1/Z_2}$$

Susy sol'n (3): Linear structure

► First layer (Z, Θ)

$$D * _4 (DZ_I + \dot{\beta} Z_I) + 2D\beta \wedge \Theta_J = 0 \quad \{I, J\} = \{1, 2\}$$

$$D\Theta_J - \dot{\beta} \wedge \Theta_J - \partial_v \left[\frac{1}{2} * _4 (DZ_I + \dot{\beta} Z_I) \right] = 0 \quad \cdot \equiv \partial_v$$

► Second layer (\mathcal{F}, ω)

$$* _4 D * _4 L = \dot{Z}_1 \dot{Z}_2 + \ddot{Z}_1 Z_2 + Z_1 \ddot{Z}_2 + \frac{1}{2} \partial_v (Z_1 Z_2) h^{mn} \dot{h}_{mn}$$

$$+ \frac{1}{2} Z_1 Z_2 \left(h^{mn} \ddot{h}_{mn} - \frac{1}{2} h^{mn} \dot{h}_{np} h^{pq} \dot{h}_{qm} \right) - 2 \dot{\beta}_m L^m - 2 * _4 (\Theta_1 \wedge \Theta_2 - \hat{\psi} \wedge D\omega)$$

$$(1 + * _4) D\omega = 2(Z_1 \Theta_1 + Z_2 \Theta_2) - \mathcal{F} D\beta - 4Z_1 Z_2 \hat{\psi}$$

$$L \equiv \dot{\omega} + \frac{1}{2} \mathcal{F} \dot{\beta} - \frac{1}{2} D\mathcal{F} \quad \hat{\psi} = \frac{1}{16} \epsilon^{ABC} J^{(A)mn} j_{mn}^{(B)} J^{(C)}$$

— Linear if solved in the right order

6D susy sol'ns: Summary

- ▶ 6D eqs have nice linear structure
 - can be solved in principle
- ▶ Difficult in practice
 - Need some physical intuition / organizing principle to proceed

A CFT view on superstrata

[Bena+MS+Warner | 404.4506]

Questions

- ▶ What sector of CFT states are expected to be visible in sugra?
- ▶ What is the structure of solutions?

D1-D5 CFT (1)

▶ D1-D5 system

	1	2	3	4	5	6	7	8	9
D1	•	•	•	•	○	~	~	~	~
D5	•	•	•	•	○	○	○	○	○

S^1 (above columns 5-6) T^4 (above columns 6-9)

$SO(4)_{1234} \cong SU(2)_L \times SU(2)_R$ $SO(4)_{6789} \cong SU(2)_1 \times SU(2)_2$

▶ $d = 2$, (large) $\mathcal{N} = (4,4)$ SCFT

▶ Sigma model with target space $(T^4)^N / S_N$, $N \equiv N_1 N_2$

D1-D5 CFT (2)

- ▶ Matter content: 4 bosons, 4 fermions

	$SU(2)_L \times SU(2)_R$	$SU(2)_1 \times SU(2)_2$
$X_{(r)}^{\dot{A}\dot{A}}(z, \bar{z})$	(1, 1)	(2, 2)
$\psi_{(r)}^{\alpha\dot{A}}(z)$	(2, 1)	(1, 2)
$\tilde{\psi}_{(r)}^{\dot{\alpha}\dot{A}}(\bar{z})$	(1, 2)	(1, 2)

$r = 1, \dots, N$:
copy index

- ▶ RR gnd states \sim chiral primaries = twisting of N copies



Maximally spinning:



Visible sector

Conjecture:

R-symmetry $SO(4)_{1234} = SU(2)_L \times SU(2)_R$
is visible from 6D sugra. Carriers: $\psi, \tilde{\psi}$

In particular, sector generated by $SU(2)$ currents:

$$J_{(r)}^{\alpha\beta}(z) = \frac{1}{2} \epsilon_{\dot{A}\dot{B}} \psi_{(r)}^{\alpha\dot{A}}(z) \psi_{(r)}^{\beta\dot{B}}(z), \quad \tilde{J}_{(r)}^{\dot{\alpha}\dot{\beta}}(\bar{z}) = \dots$$

namely, Sugawara CFT

$$[SU(2)_L \times SU(2)_R]^N / S_N, \quad c = N$$

must be visible (the rest has $c = 5N$)



Expected entropy from strata

By considering multiple superstrata,
expect in the bulk to see entropy for $c = N$:

$$S_{\text{strata}} = 2\pi \sqrt{\frac{NN_3}{6}}$$

Instead of the full entropy for a $c = 6N$ system:

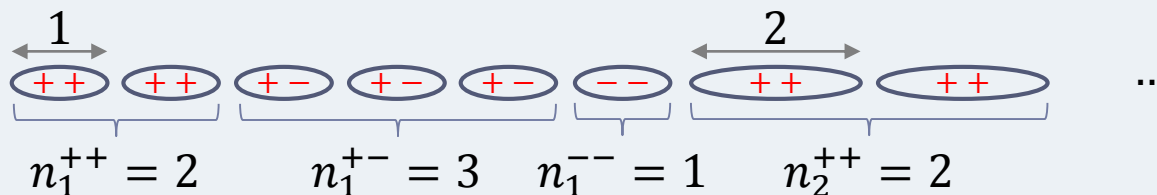
$$S_{\text{full}} = 2\pi \sqrt{NN_3}$$

Correct scaling must be reproducible!

(To get numerical factor right, need to look into compact T^4 .)

Evidence: 2-chg states (RR gnd states)

CFT: $n_k^{\alpha\dot{\alpha}}$: number of component strings with length k , spin $\alpha, \dot{\alpha}$



Note: $J_0^- \textcircled{++} = \textcircled{-+}$ etc.

Sugra: LM geometry



Profile $\vec{x} = \vec{f}(w) \in \mathbb{R}^4$

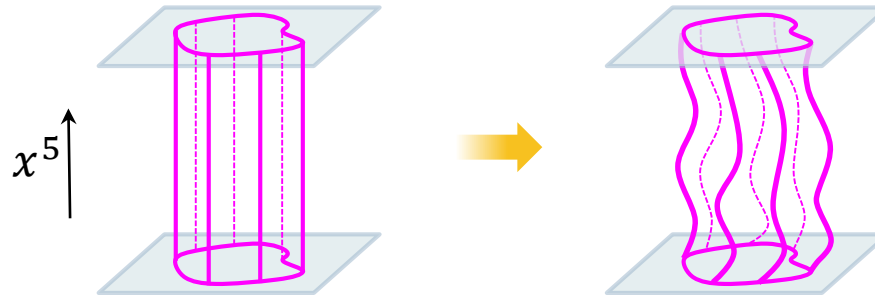
$$f^{\alpha\dot{\alpha}}(w) \sim \sum_{k \neq 0} a_k^{\alpha\dot{\alpha}} e^{2\pi i k w}$$

Dictionary: $n_k^{\alpha\dot{\alpha}} \leftrightarrow |a_k^{\alpha\dot{\alpha}}|^2$ [Lunin-Mathur]
[Kanitscheider-Skenderis-Taylor]

$SU(2)_L \times SU(2)_R$ current sector is precisely visible

3-charge states

- ▶ Add $P(\nu)$ \rightarrow fluctuation along ν
 - ▶ Expect: $f^{\alpha\dot{\alpha}}(w) \rightarrow f^{\alpha\dot{\alpha}}(w, \nu)$
 - Depend on two variables

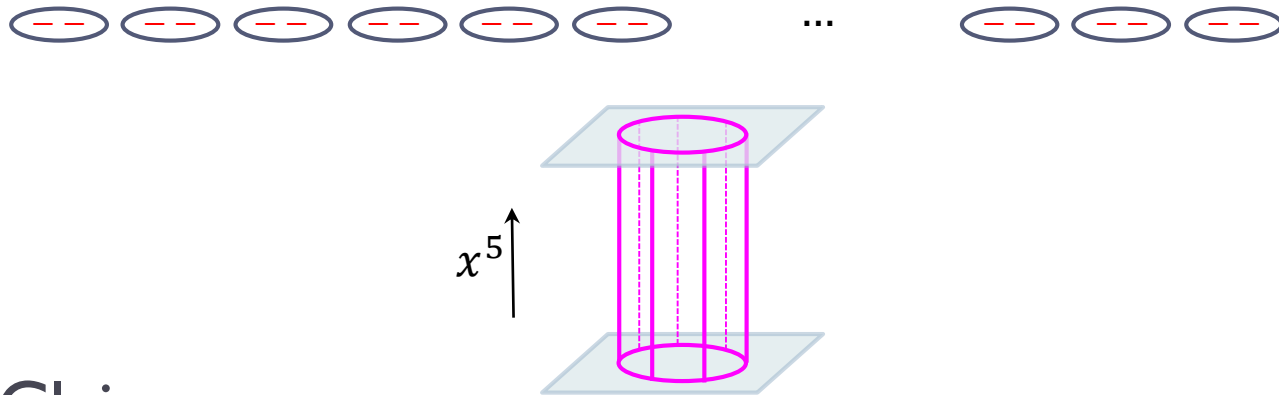


- ▶ CFT: (any, gnd)

v -dep fluct. for 3-charge states (1)

Consider:

Circular superstratum = Maximally spinning RR gnd state
= pure $AdS_3 \times S^3$



Claim:

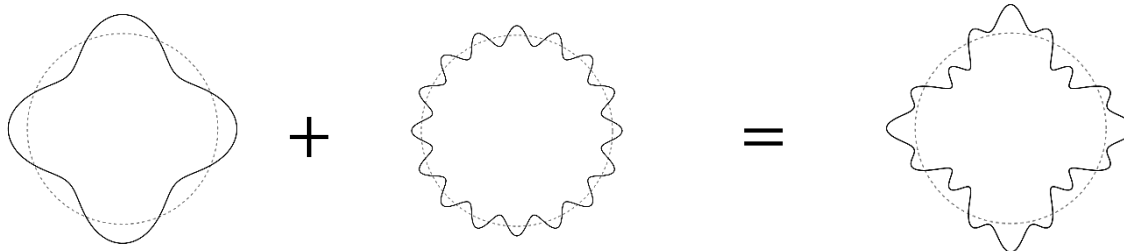
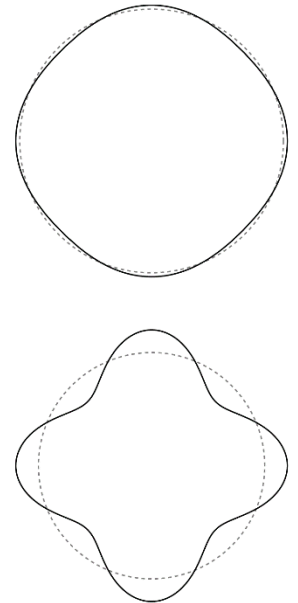
Fluctuations around this state are parametrized by functions of two variables.

v -dep fluct. for 3-charge states (2)

- ▶ Fluctuation around $AdS_3 \times S^3$ is in some $SU(2)_L \times SU(2)_R$ rep.
- ▶ Other RR ground state: $(\ell, \ell; \tilde{\ell}, \tilde{\ell}), |\ell - \tilde{\ell}| \leq 2$
 - One quantum number
 - One variable (LM profile function $f(w)$)
- ▶ Act by J (but not \tilde{J}) modes → $(\ell, m; \tilde{\ell}, \tilde{\ell})$
 - Two quantum numbers
 - Two variables (superstratum, $f(w, v)$)

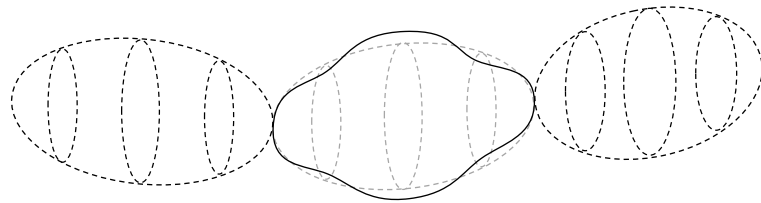
Comments

- ▶ At linear level, can realize bulk action of J on linear fluctuation around $AdS_3 \times S^3$
- ▶ Can use linear structure of 6D eqs to nonlinearly complete it (work in progress).
- ▶ Superposing multiple modes in bulk
↔ considering non-chiral primaries in CFT

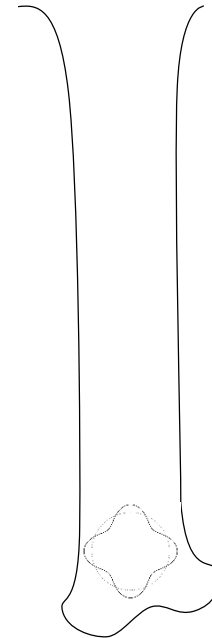


Multiple superstrata

- ▶ More generally, one has multiple S^3 's
- ▶ Can fluctuate each S^3 — multi-superstratum



- ▶ Can use $AdS_3 \times S^3$ as local model
- ▶ Large redshift in scaling geometries
 - entropy enhancement
 - $S \sim \sqrt{N_1 N_2 N_3}$?



Summary

- ▶ $SU(2)$ current algebra sector:
expected to be visible in sugra
- ▶ Sugra states: superstrata depending on 2 vars.
- ▶ Have to solve 6D system
by nonlinearly completing linear fluctuations
(work in progress)
- ▶ Multiple superstrata, scaling solution
→ entropy enhancement?

Conclusions

Conclusions

- ▶ Superstrata: conjectural microstate (non)geometries
- ▶ They live in 6D sugra (or generalization thereof)
- ▶ $SU(2)$ current algebra of D1-D5 CFT describes their fluctuations
- ▶ A LOT more stuff to do, more fun to enjoy!
 - ▶ Construct
 - ▶ Count — Stay tuned! 😊
 - ▶ Hit them to death

Thanks!