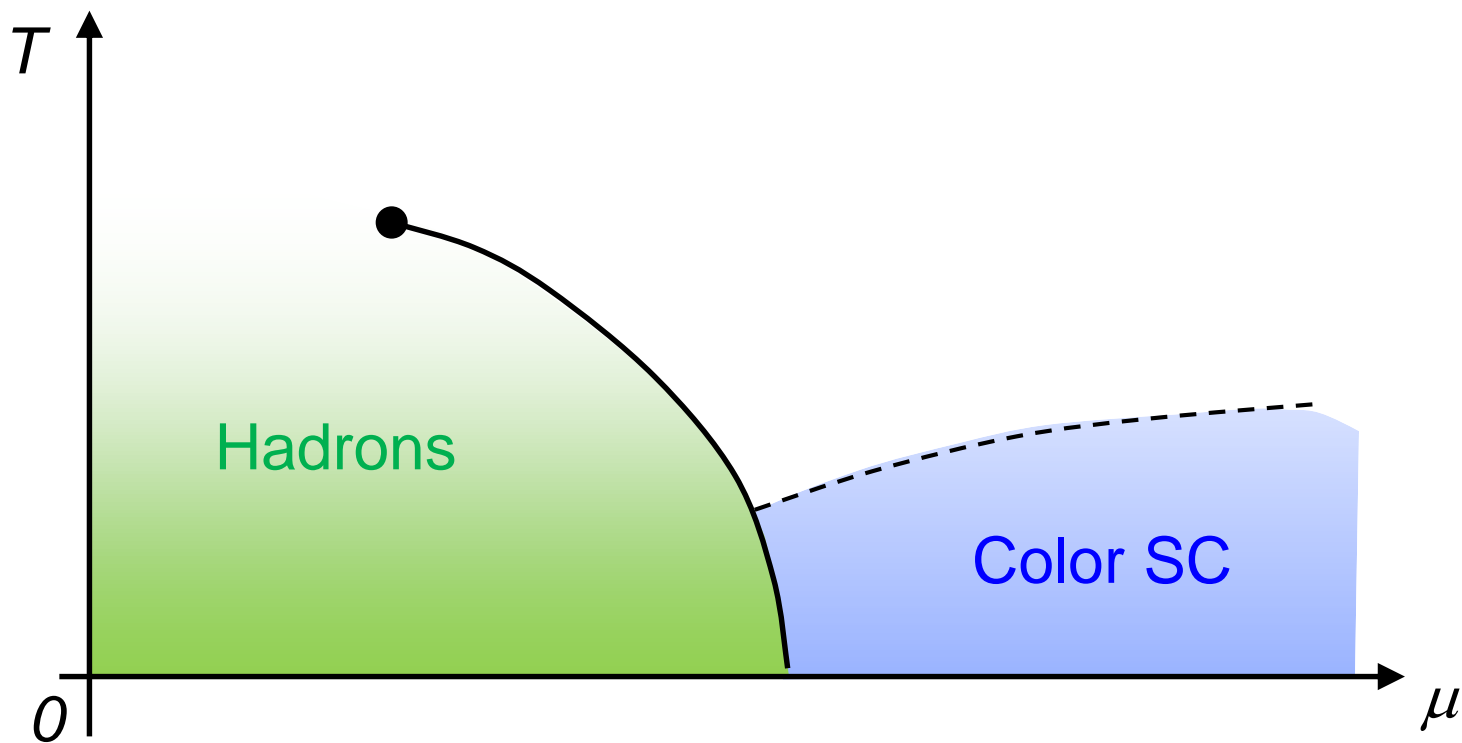


# 重イオン衝突における ゆらぎの話

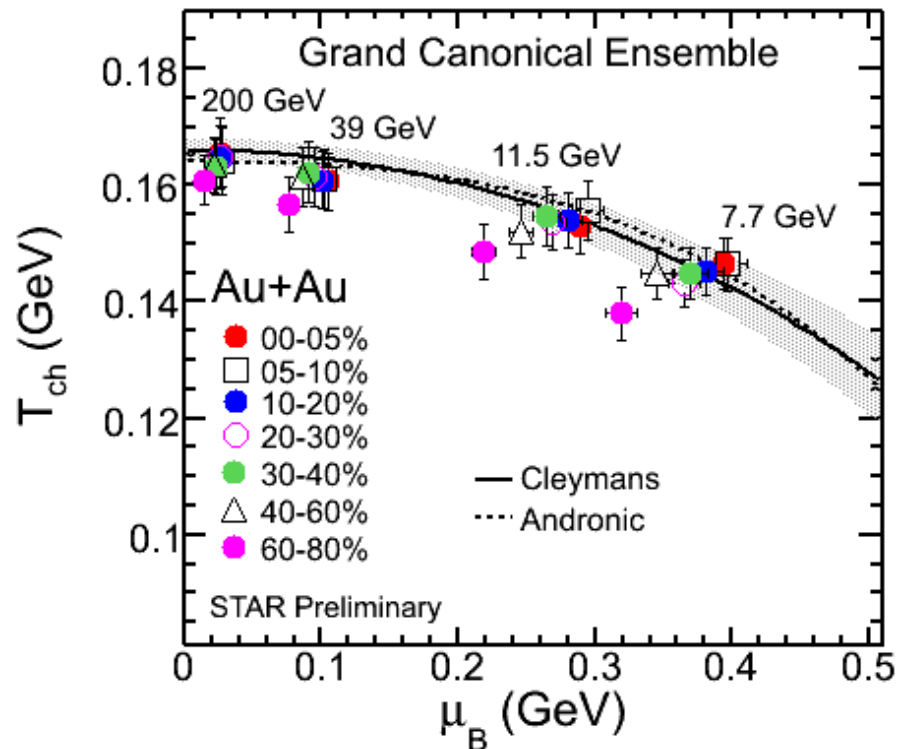
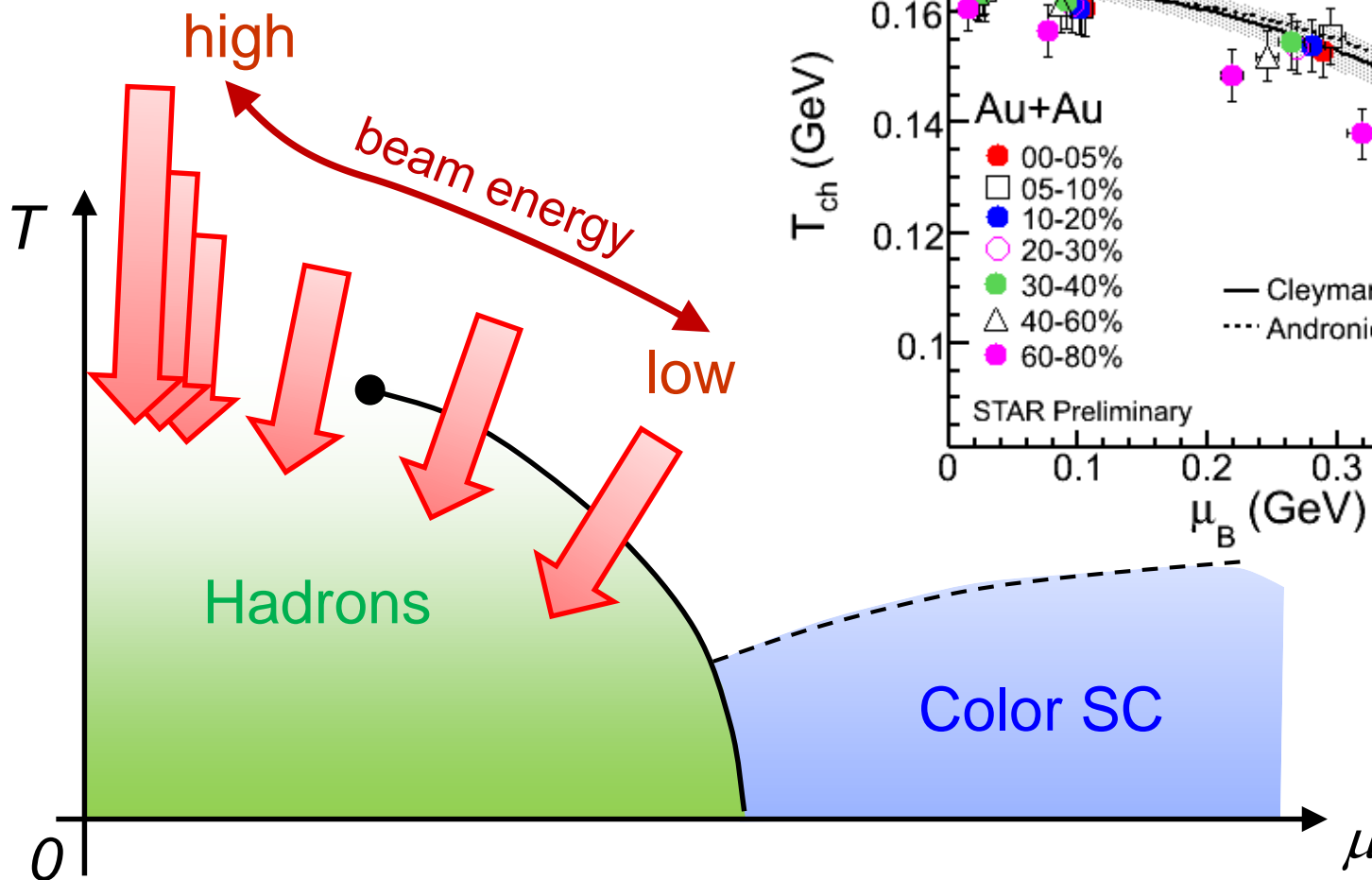
北沢 正清  
(阪大)

# Beam-Energy Scan



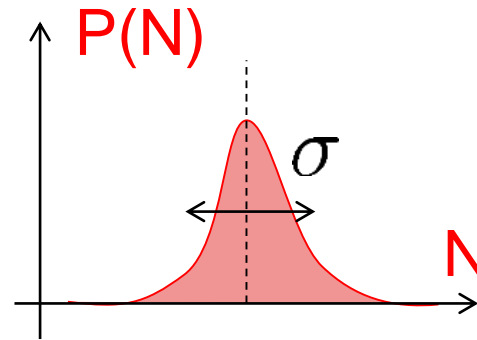
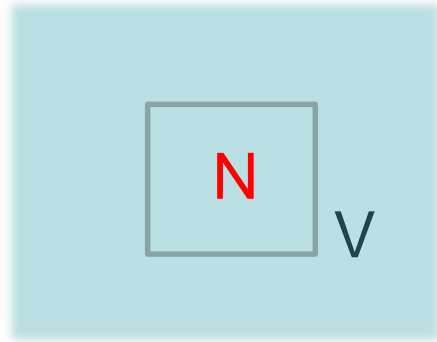
# Beam-Energy Scan

STAR 2012



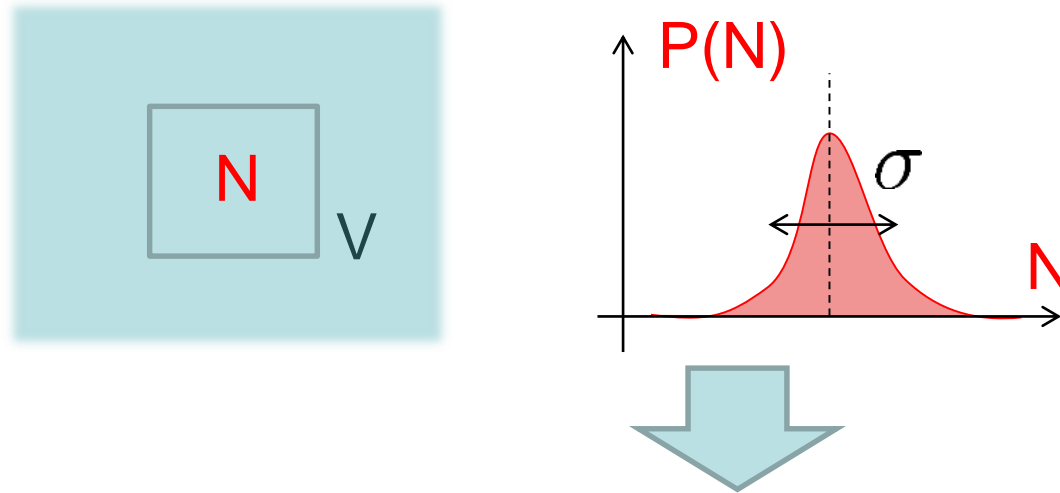
# Fluctuations

Observables in equilibrium are fluctuating.



# Fluctuations

Observables in equilibrium are fluctuating.



➤ Variance:  $\langle \delta N^2 \rangle = V \chi_2 = \sigma^2$

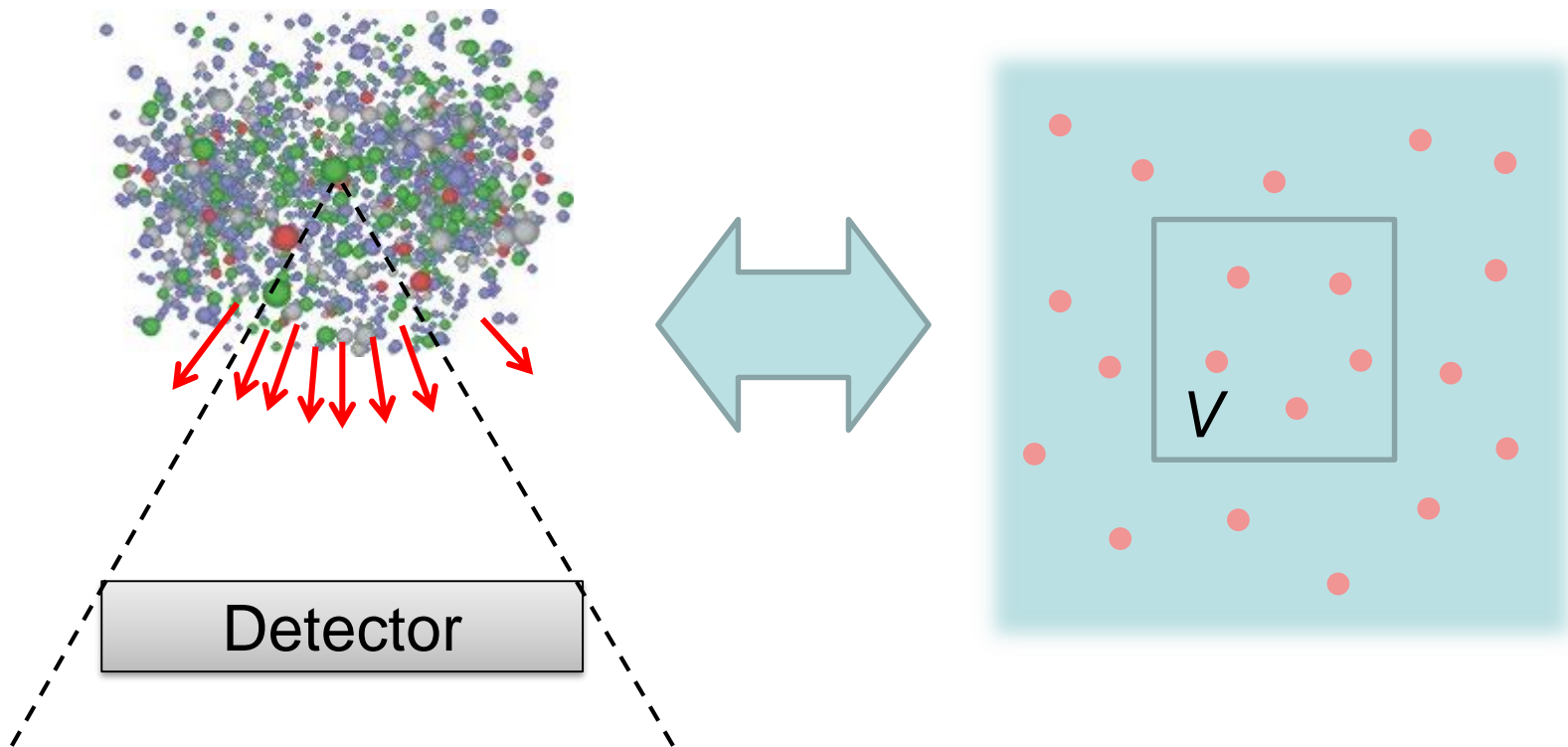
$$\delta N = N - \langle N \rangle$$

➤ Skewness:  $S = \frac{\langle \delta N^3 \rangle}{\sigma^3}$

➤ Kurtosis:  $\kappa = \frac{\langle \delta N^4 \rangle - 3\langle \delta N^2 \rangle^2}{\chi_2 \sigma^2}$

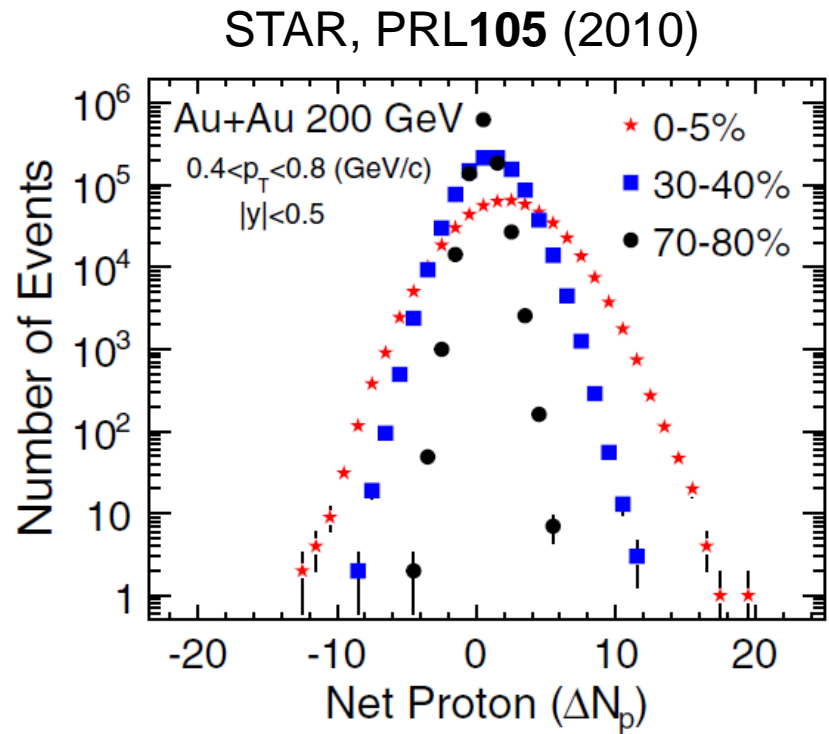
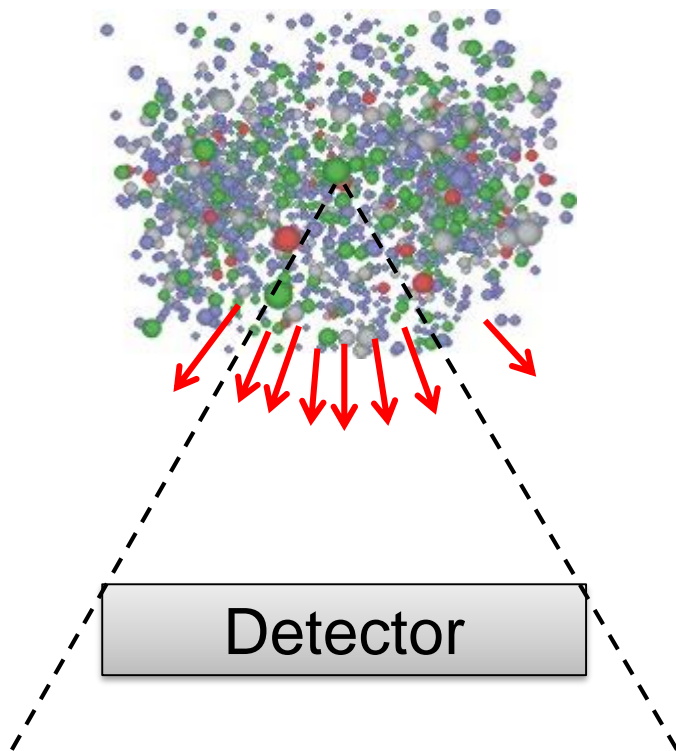
# Event-by-Event Analysis @ HIC

Fluctuations can be measured by e-by-e analysis in experiments.



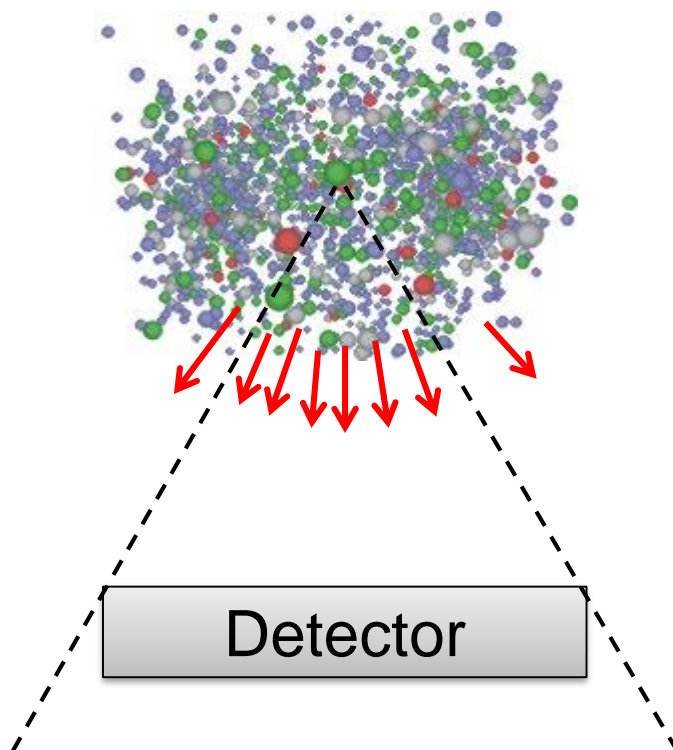
# Event-by-Event Analysis @ HIC

Fluctuations can be measured by e-by-e analysis in experiments.



# 観測にかかるゆらぎは、いつ形成されたのか？

ゆらぎのダイナミクス(動的振る舞い)の議論が必要

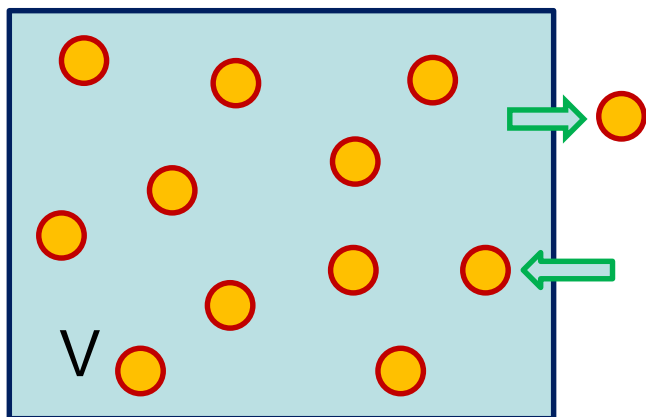




# 観測にかかるゆらぎは、いつ形成されたのか？

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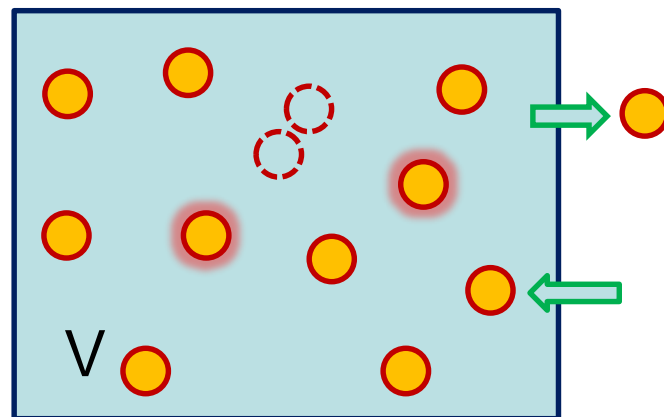
保存電荷の場合



境界を通過する電荷のみが変化に寄与

$$\tau \rightarrow \infty$$
$$\text{for } V \rightarrow \infty$$

非保存電荷の場合



体積内の任意の場所で電荷が変化できる

$$\tau \rightarrow \text{const.}$$
$$\text{for } V \rightarrow \infty$$

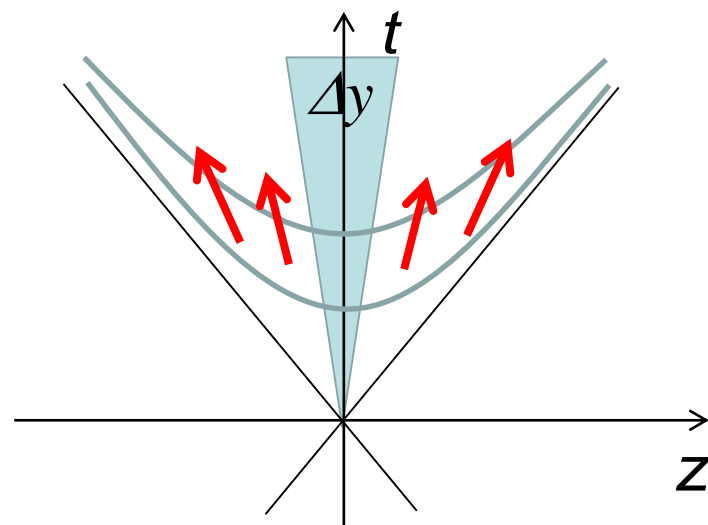
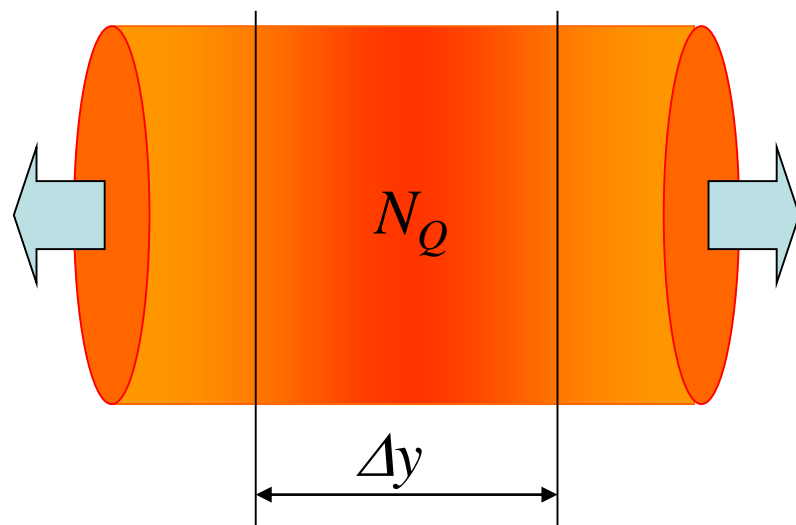
# 観測にかかるゆらぎは、いつ形成されたのか？

$\Delta\eta$ 内の保存電荷量は、初期段階のものが終状態まで生き残ることが期待できる。

Asakawa, Heinz, Muller, '00  
Jeon, Koch, '00  
Shuryak, Stephanov, '02

Note:

$$\text{STAR} \left\{ \begin{array}{l} -0.5 < \eta < 0.5 \\ 0.4 < p < 0.8 [\text{GeV}] \end{array} \right.$$

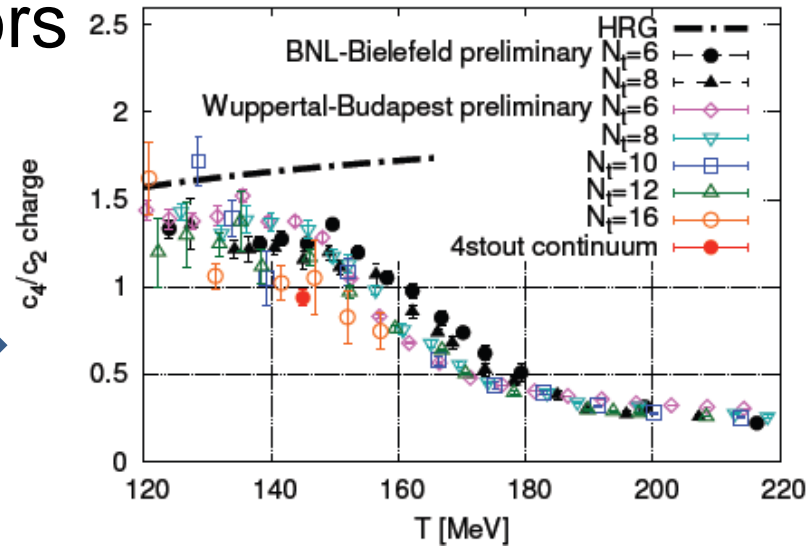


# Conserved Charges : Theoretical Advantage

## □ Definite definition for operators

- as a Noether current
- calculable on any theory

ex: on the lattice

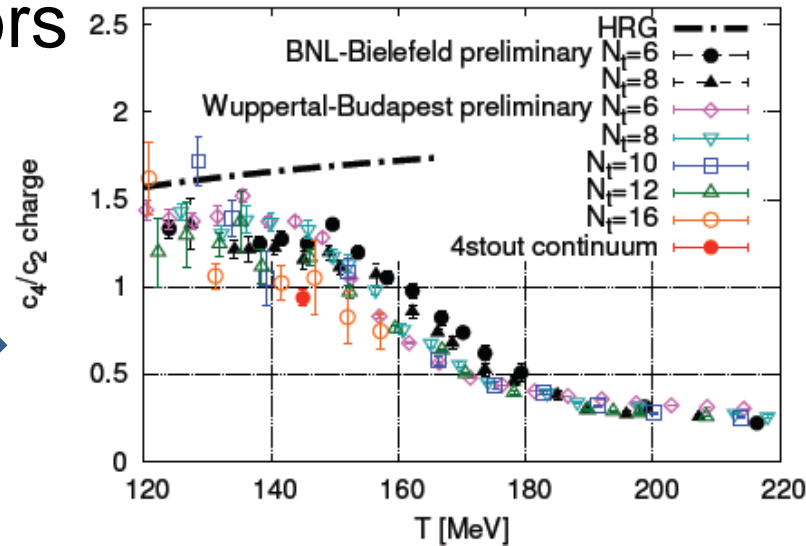


# Conserved Charges : Theoretical Advantage

## Definite definition for operators

- as a Noether current
- calculable on any theory

ex: on the lattice



## Simple thermodynamic relations

$$\langle \delta N_c^n \rangle = \frac{1}{VT^{n-1}} \frac{\partial^n \Omega}{\partial \mu_c^n}$$

- Intuitive interpretation for the behaviors of cumulants

ex:  $\langle \delta N_B^3 \rangle = \frac{1}{VT^2} \frac{\partial \langle \delta N_B^2 \rangle}{\partial \mu_B}$



# Fluctuations

Fluctuations reflect properties of matter.

Enhancement near the critical point

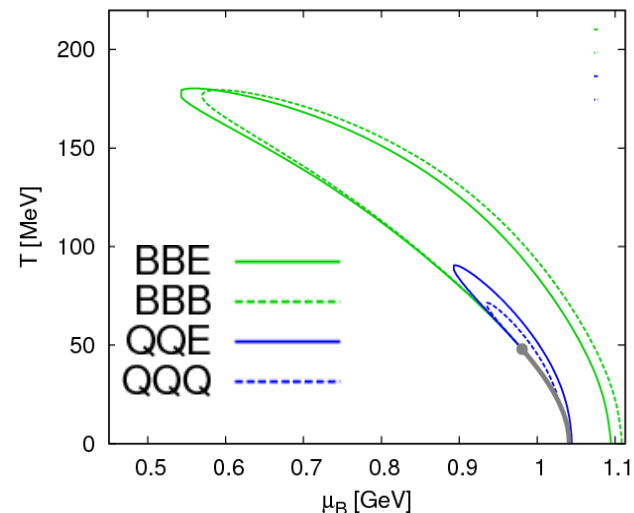
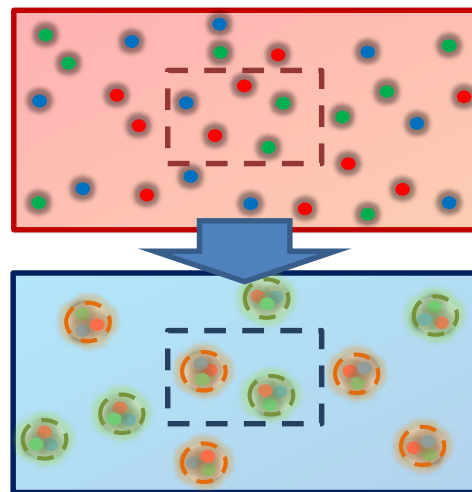
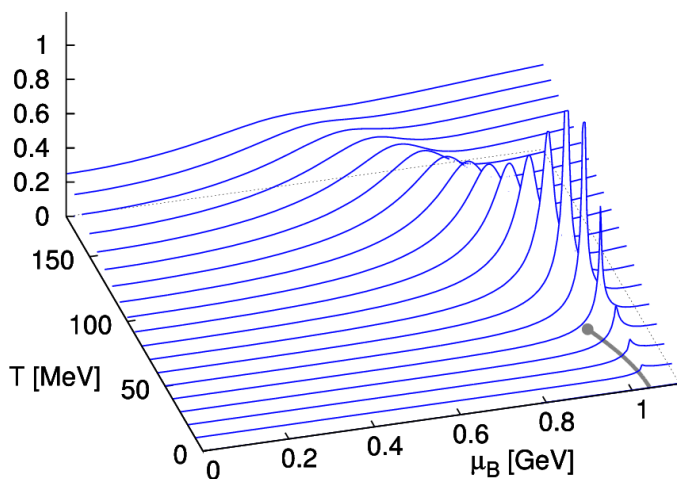
Stephanov,Rajagopal,Shuryak('98); Hatta,Stephanov('02); Stephanov('09);...

Ratios between cumulants of conserved charges

Asakawa,Heintz,Muller('00); Jeon, Koch('00); Ejiri,Karsch,Redlich('06)

Signs of higher order cumulants

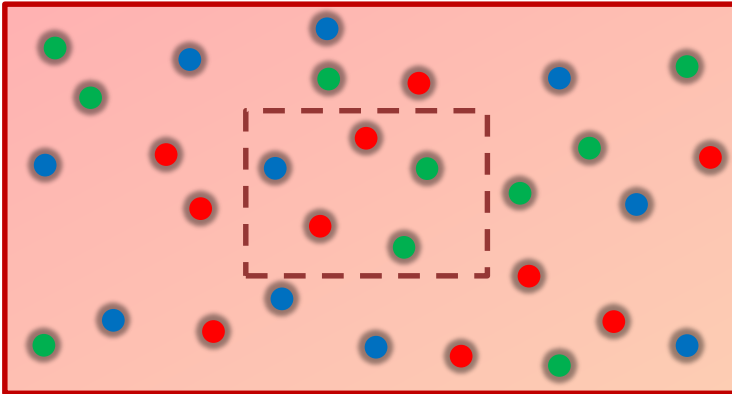
Asakawa,Ejiri,MK('09); Friman,et al.('11); Stephanov('11)



# Fluctuations

Free Boltzmann  $\rightarrow$  Poisson

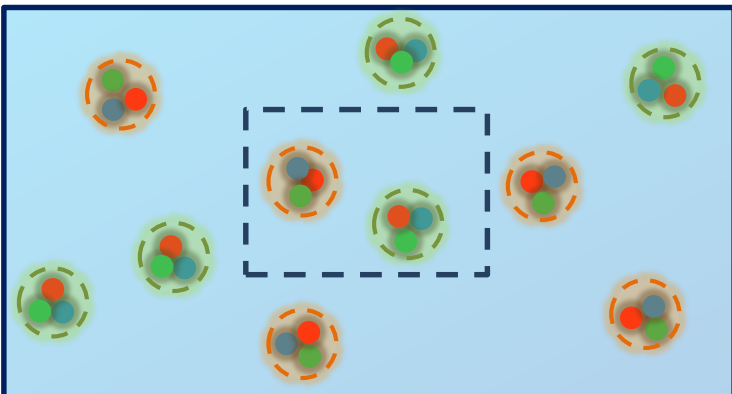
$$\langle \delta N^n \rangle_c = \langle N \rangle$$



$$\langle \delta N_q^n \rangle_c = \langle N_q \rangle$$

$$\rightarrow \langle \delta N_B^n \rangle_c = \frac{1}{3^{n-1}} \langle N_B \rangle$$

$$3N_B = N_q$$

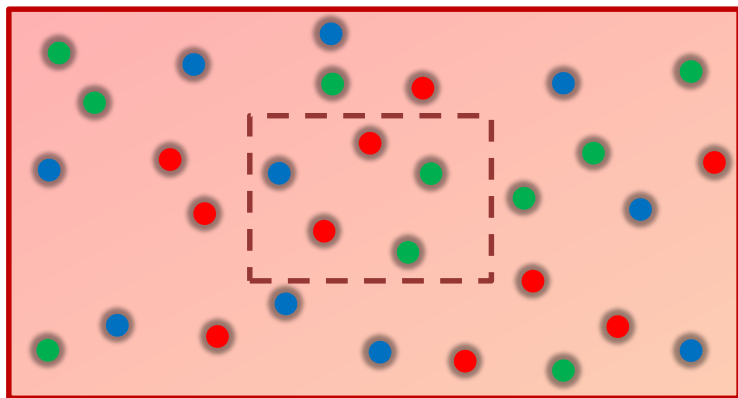


$$\langle \delta N_B^n \rangle_c = \langle N_B \rangle$$

# Fluctuations

Free Boltzmann  $\rightarrow$  Poisson

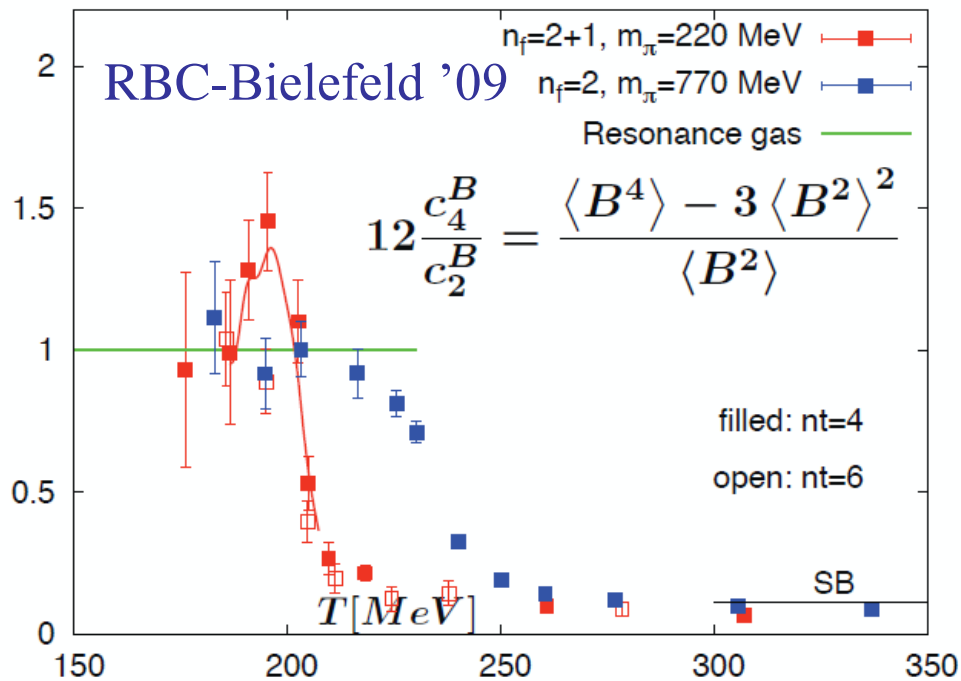
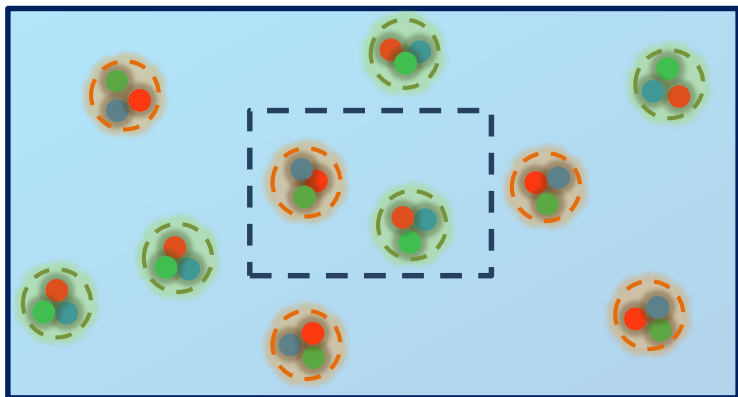
$$\langle \delta N^n \rangle_c = \langle N \rangle$$



$$\langle \delta N_q^n \rangle_c = \langle N_q \rangle$$

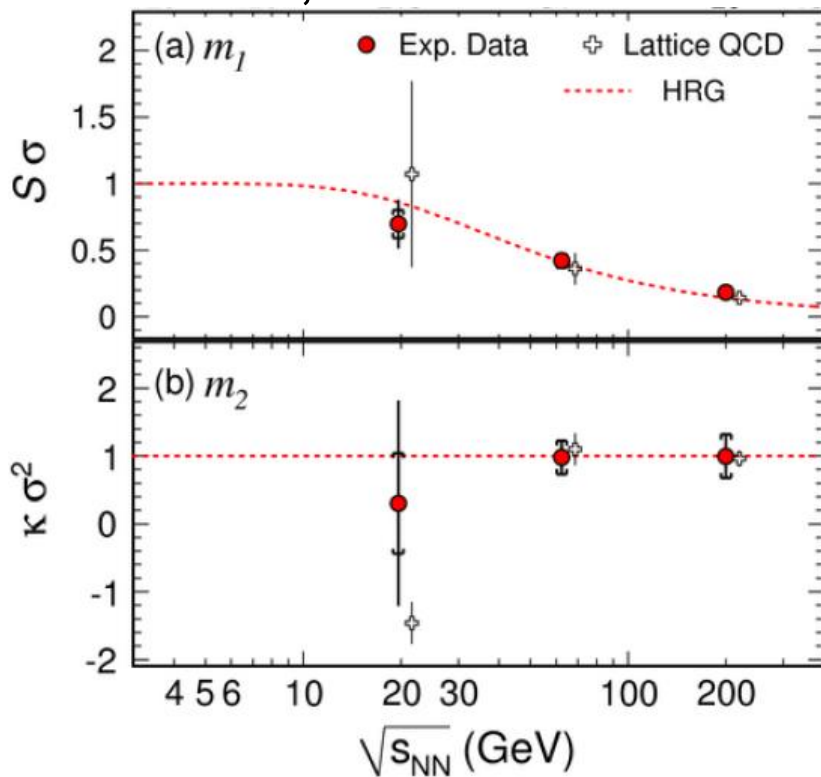
$$\Rightarrow \langle \delta N_B^n \rangle_c = \frac{1}{3^{n-1}} \langle N_B \rangle$$

$$3N_B = N_q$$



# Proton # Fluctuations @ STAR-BES

STAR, PRL2010

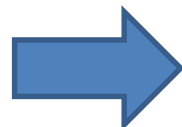
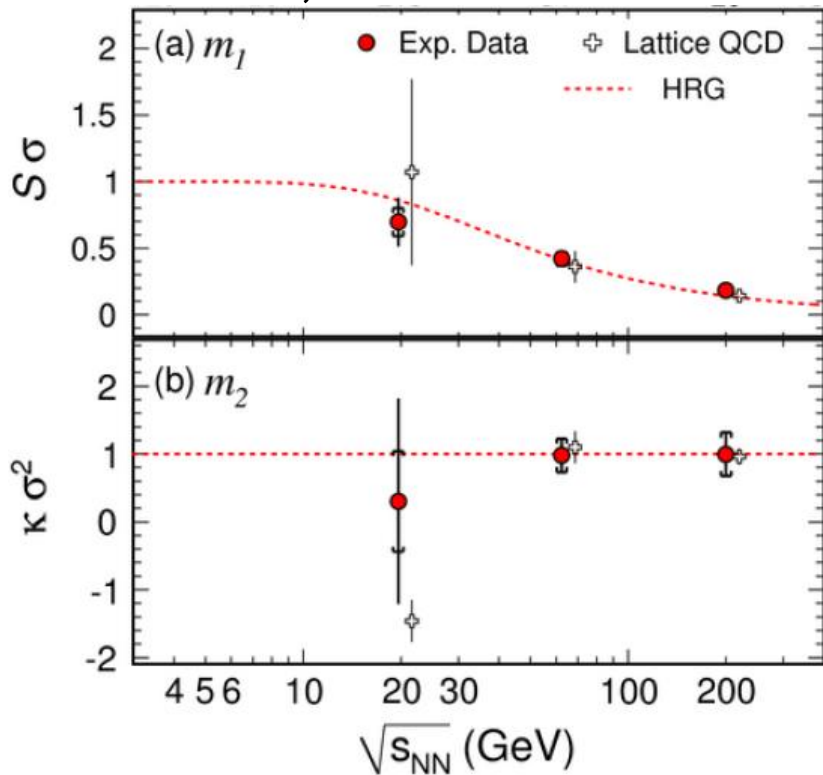


$$S\sigma = \frac{\langle (\delta N_p^{(\text{net})})^3 \rangle}{\langle (\delta N_p^{(\text{net})})^2 \rangle}, \quad \kappa\sigma^2 = \frac{\langle (\delta N_p^{(\text{net})})^4 \rangle_c}{\langle (\delta N_p^{(\text{net})})^2 \rangle}$$

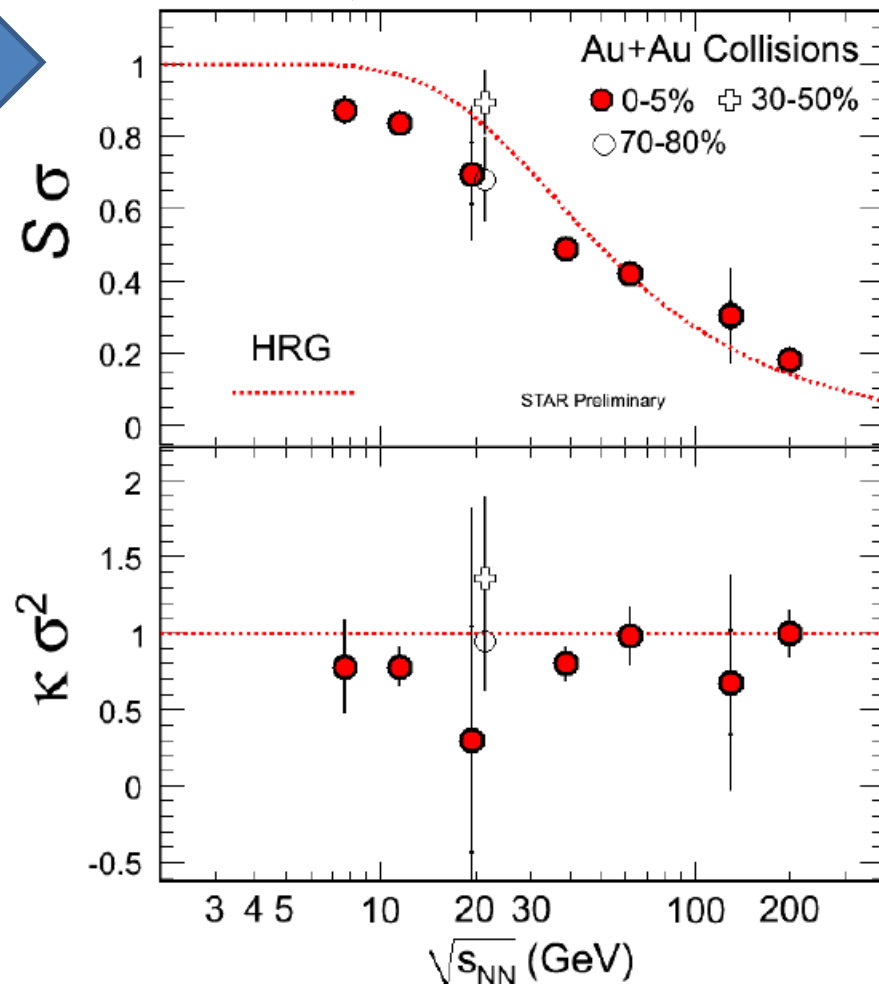


# Proton # Fluctuations @ STAR-BES

STAR, PRL2010



STAR, 2011



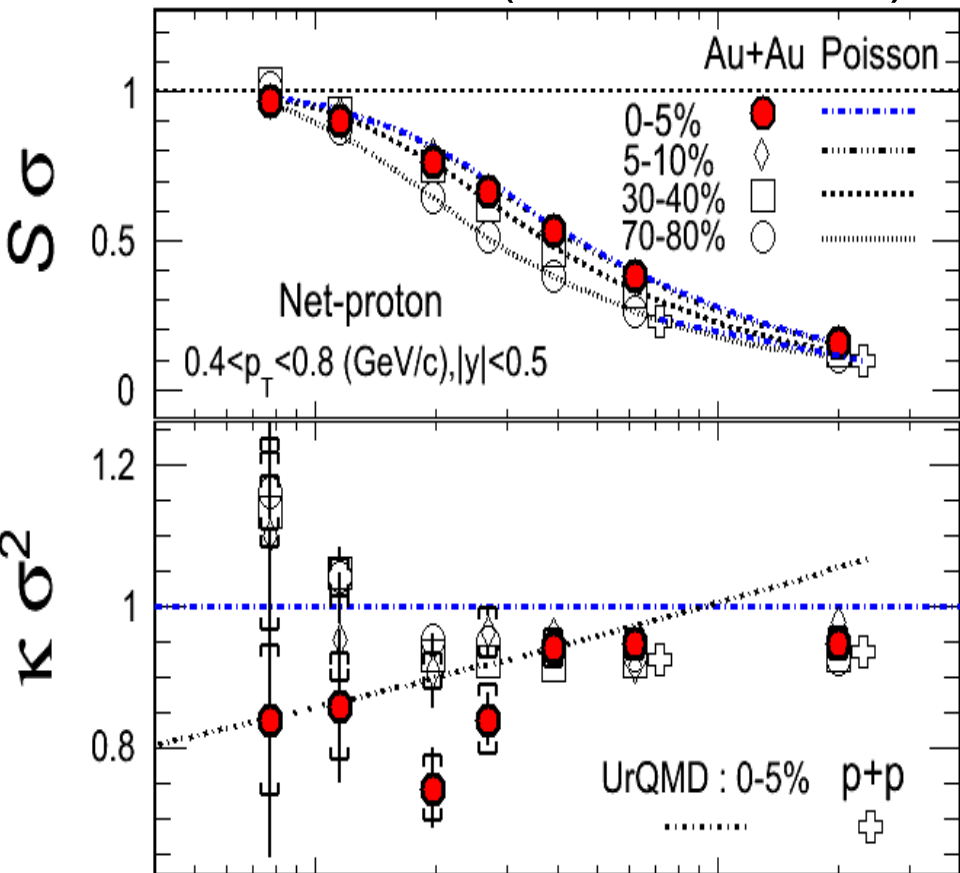
$$S\sigma = \frac{\langle(\delta N_p^{(\text{net})})^3\rangle}{\langle(\delta N_p^{(\text{net})})^2\rangle}, \quad \kappa\sigma^2 = \frac{\langle(\delta N_p^{(\text{net})})^4\rangle_c}{\langle(\delta N_p^{(\text{net})})^2\rangle}$$

high  $\mu$

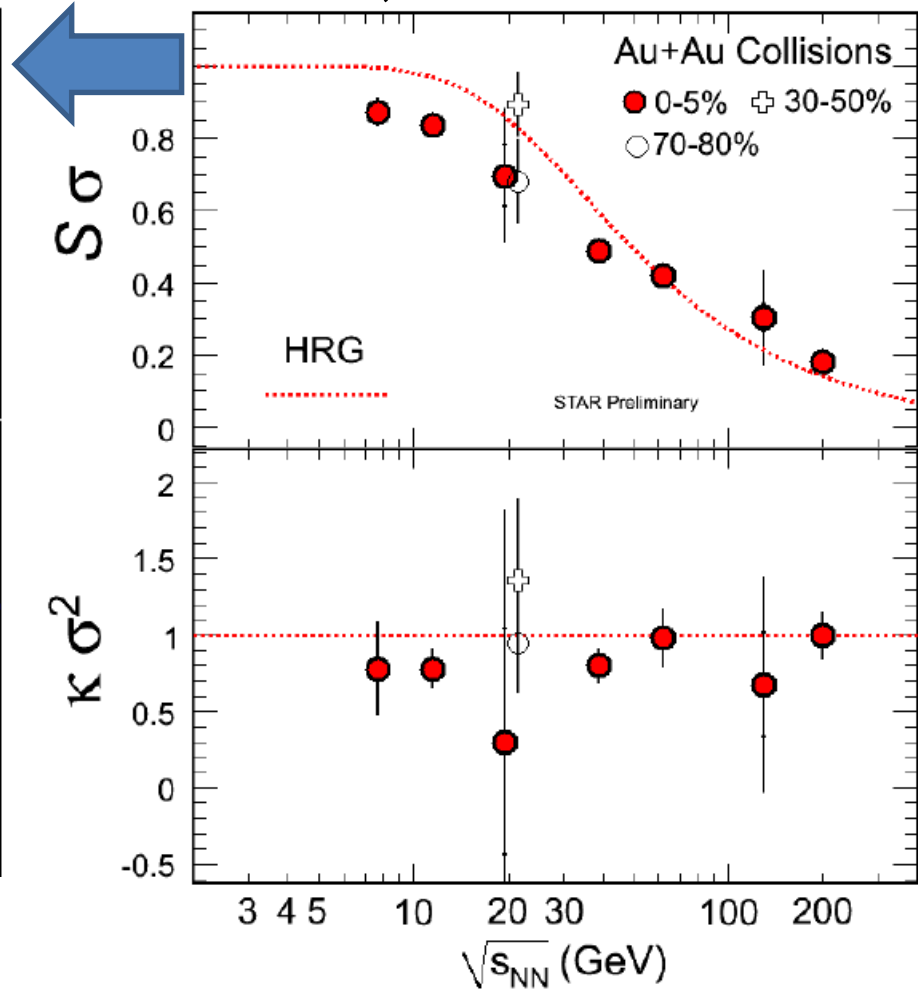
low  $\mu$

# Proton # Fluctuations @ STAR-BES

STAR, 2012 (Quark Matter)



STAR, 2011

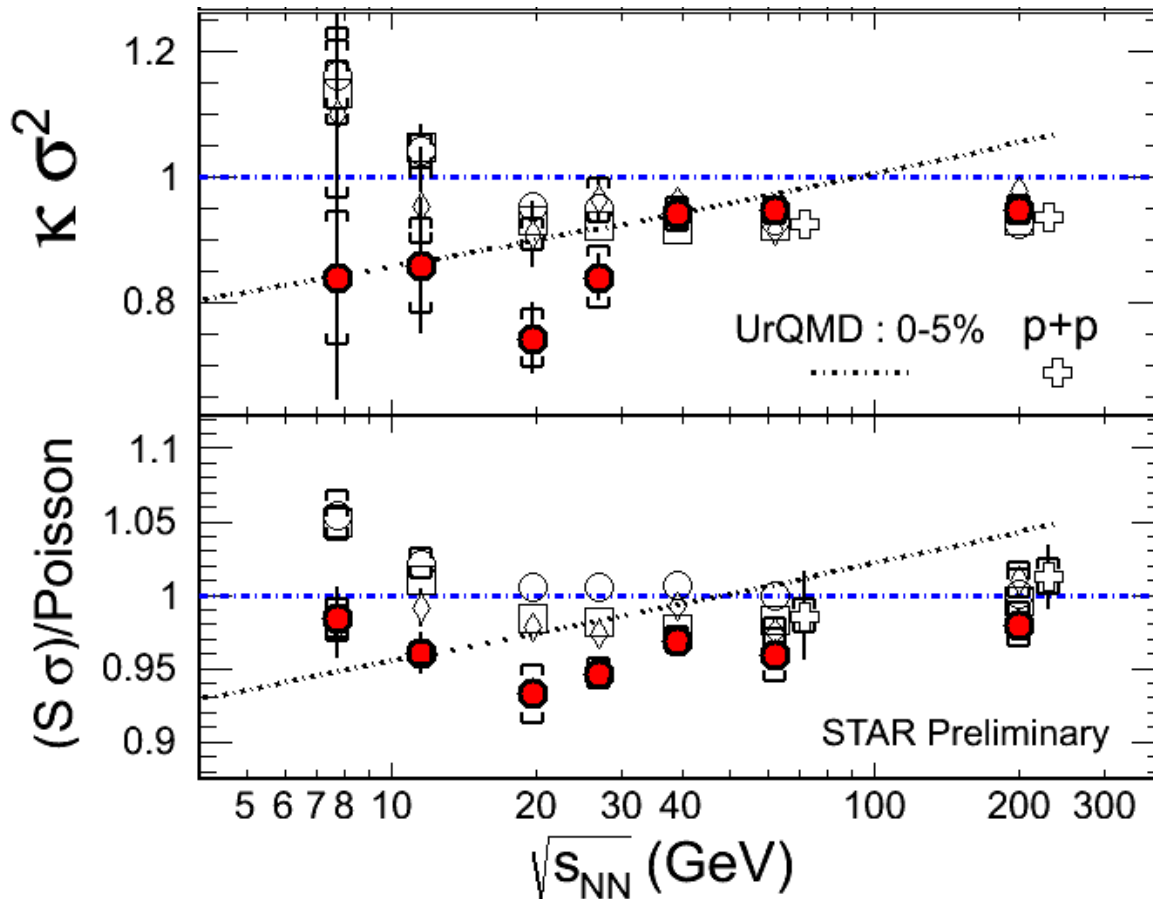


high  $\mu$

low  $\mu$

# Proton # Cumulants @ STAR-BES

STAR, QM2012



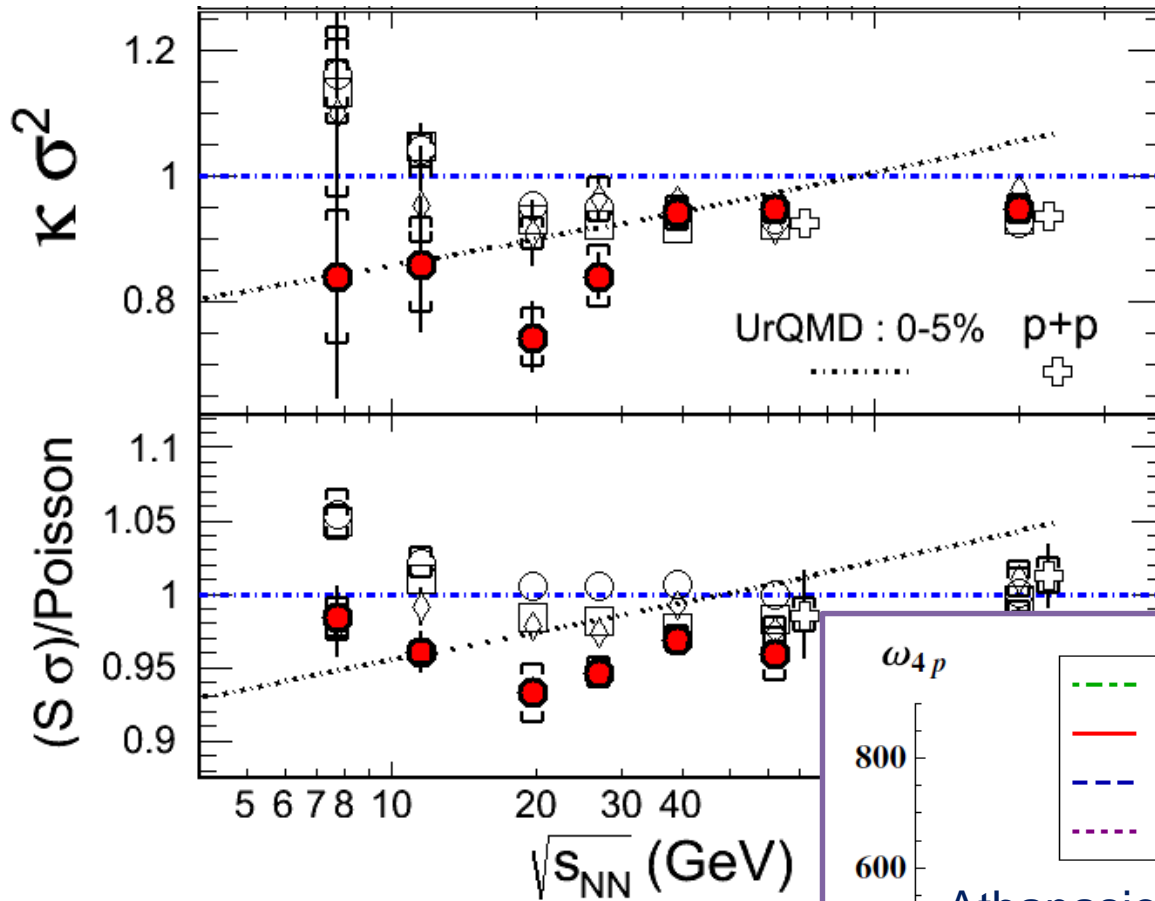
$$\frac{C_4}{C_2}$$

$$\frac{C_3}{C_1} = \frac{C_3/C_2}{\text{Poissonian}}$$

No characteristic signals on  
phase transition to QGP nor QCD CP

# Proton # Cumulants @ STAR-BES

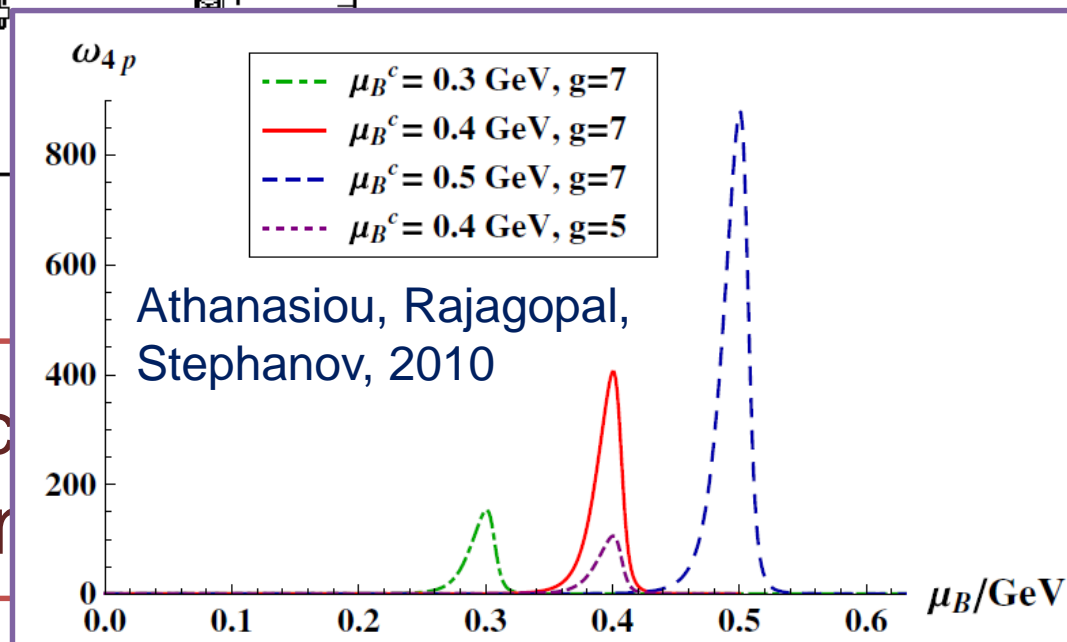
STAR, QM2012



$$\frac{C_4}{C_2}$$

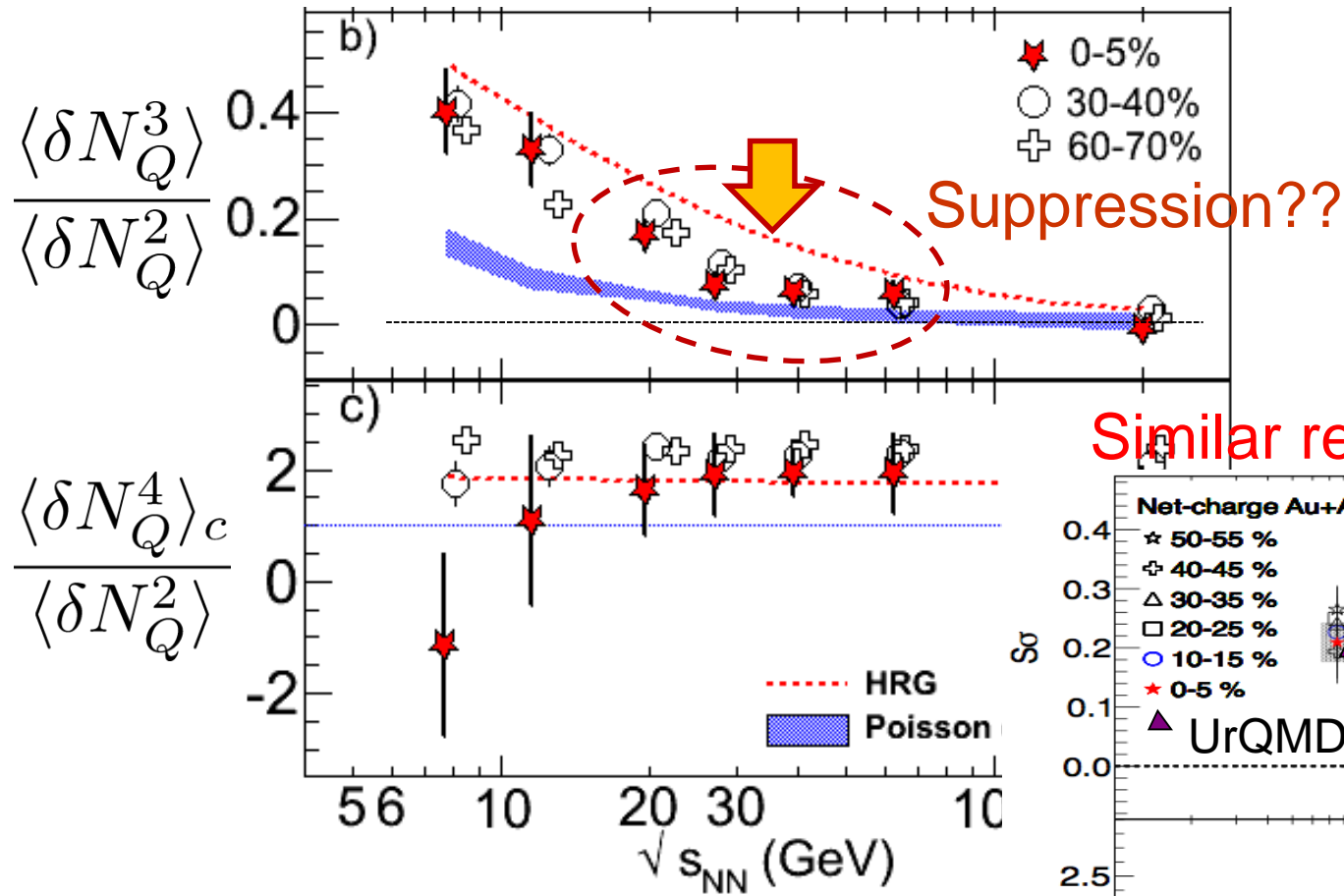
$$\frac{C_3}{C_2} = \frac{C_3}{C_2}$$

No charac  
phase transition

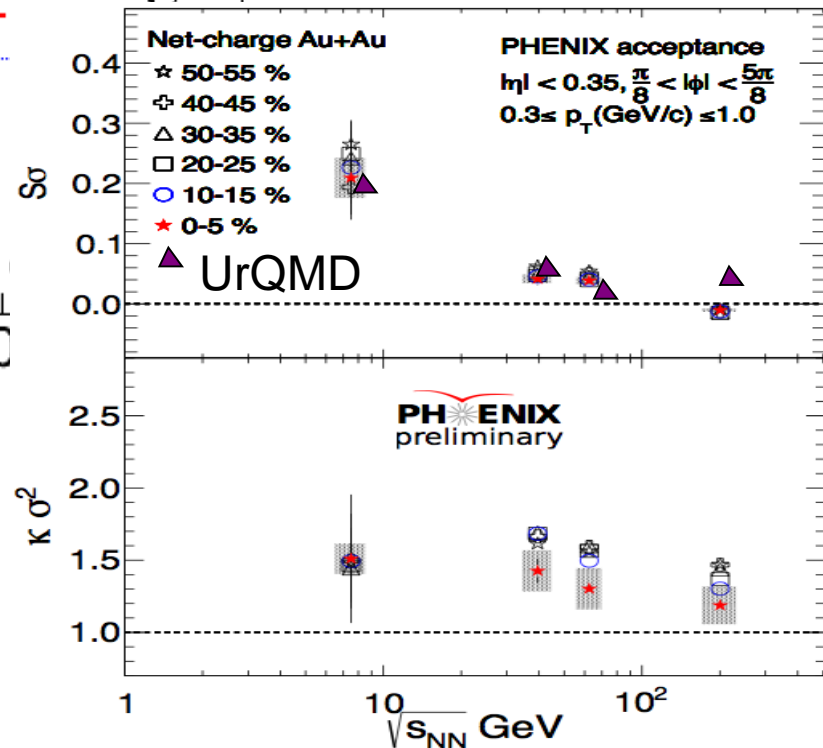


# Charge Fluctuations @ STAR-BES

STAR, QM2012

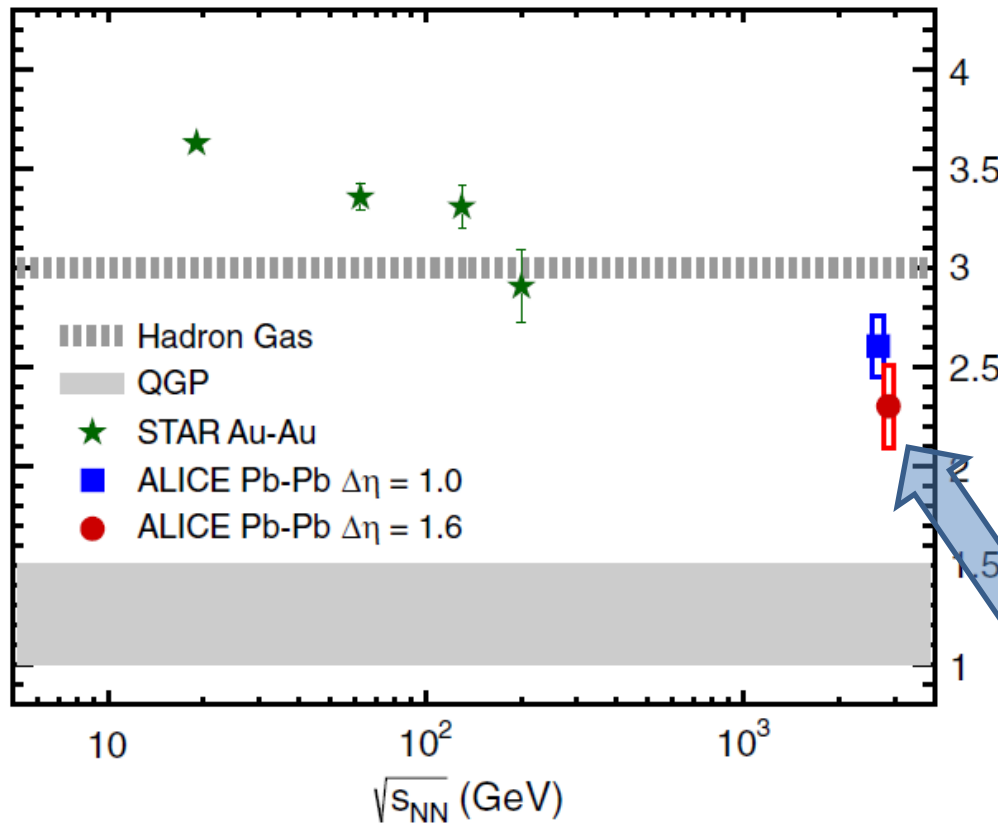


Similar result @ PHENIX



# Charge Fluctuation @ LHC

ALICE, PRL110,152301(2013)



D-measure

$$D = 4 \frac{\langle \delta N_Q^2 \rangle}{\langle N_Q^+ + N_Q^- \rangle}$$

- $D \sim 3-4$  Hadronic
- $D \sim 1$  Quark

LHC:

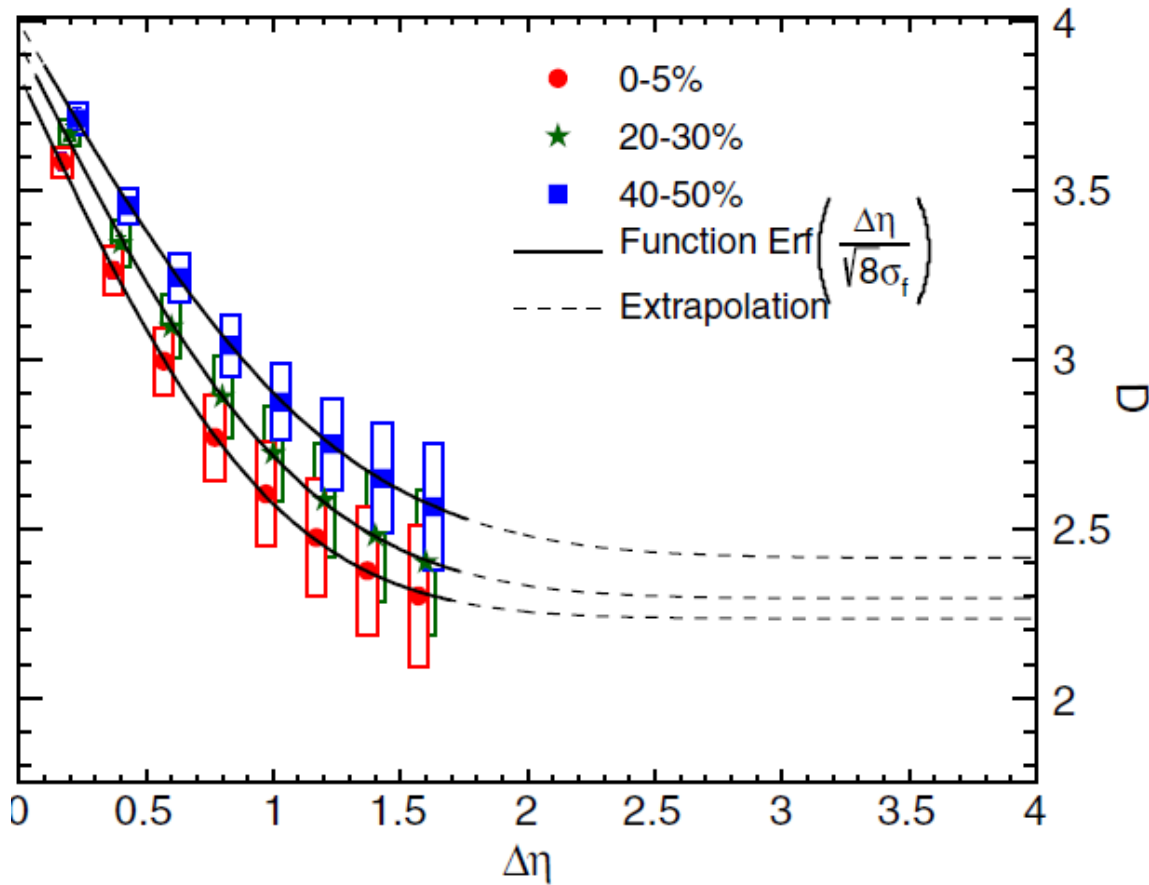
significant suppression  
from hadronic value



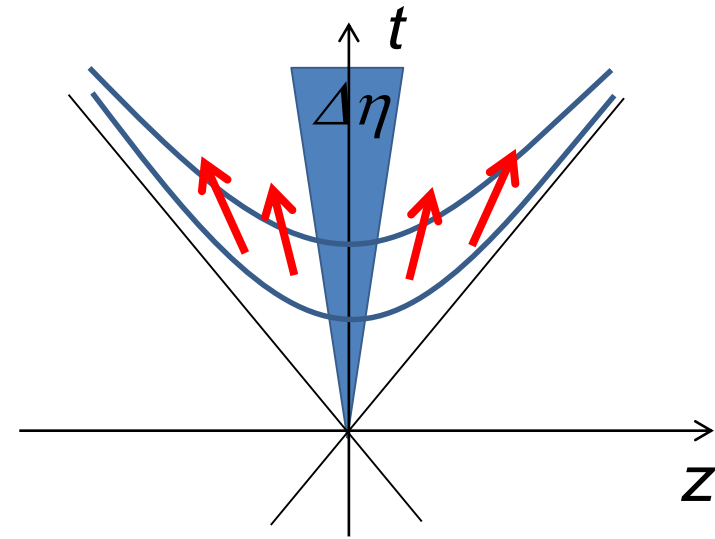
LHC終状態の電荷ゆらぎは、ハドロン化以前に生成されたものを強く反映している！

# $\Delta\eta$ Dependence @ ALICE

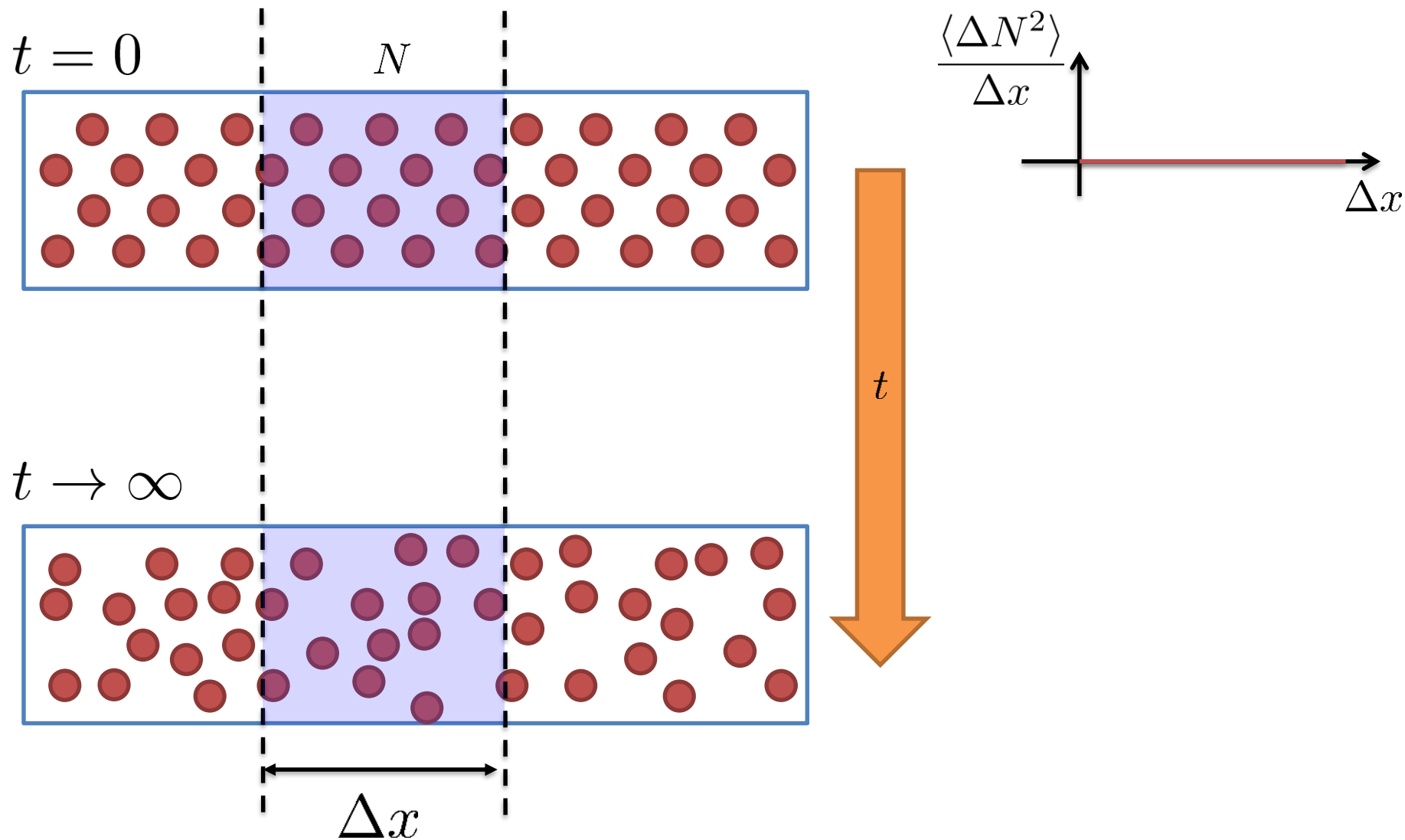
ALICE  
PRL 2013



$\Delta\eta$   
↑  
rapidity window

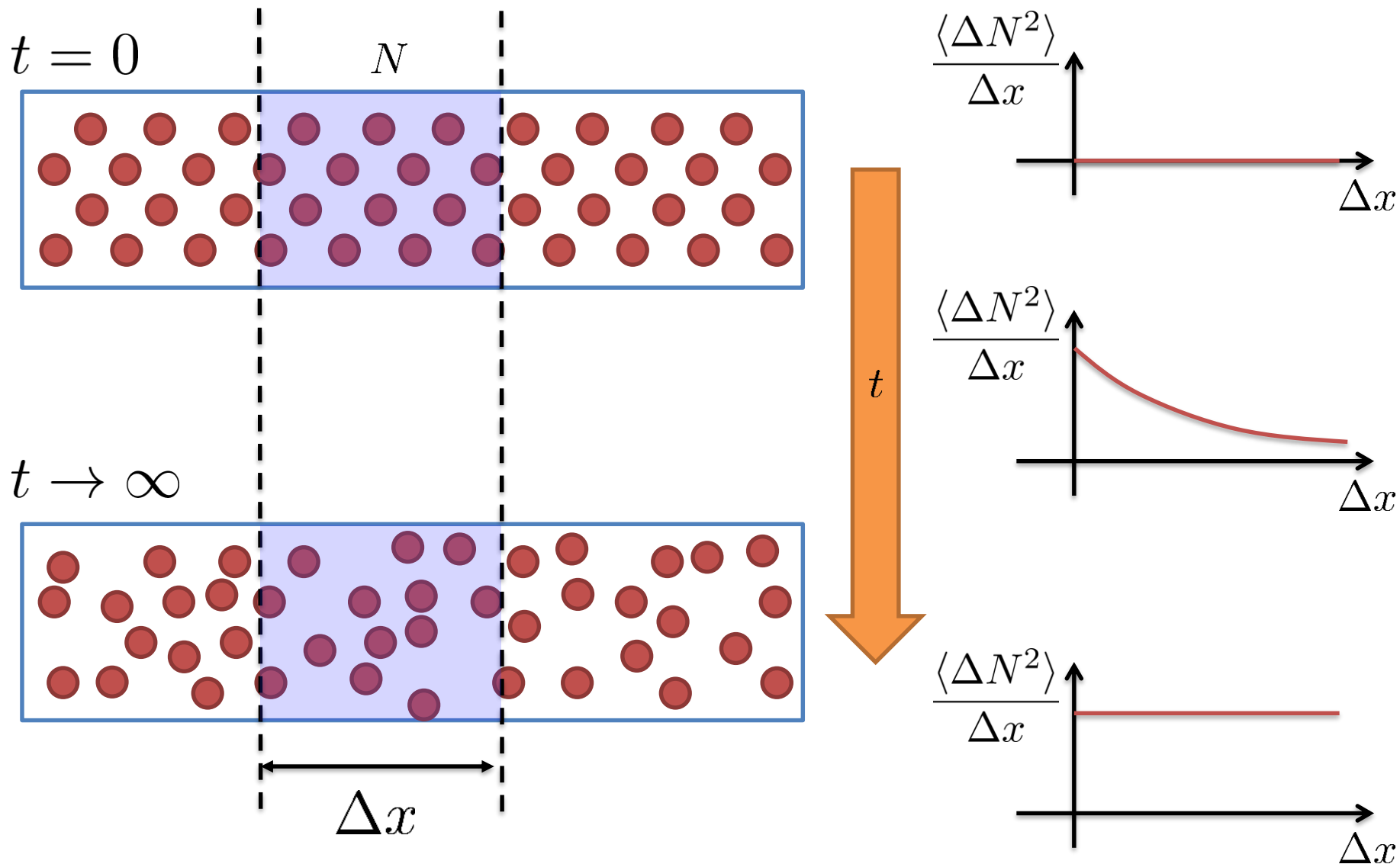


# Dissipation of a Conserved Charge



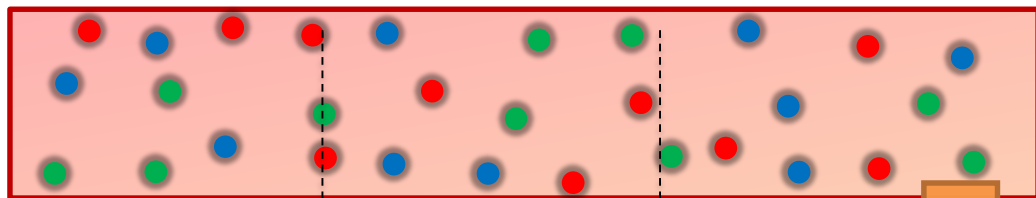


# Dissipation of a Conserved Charge

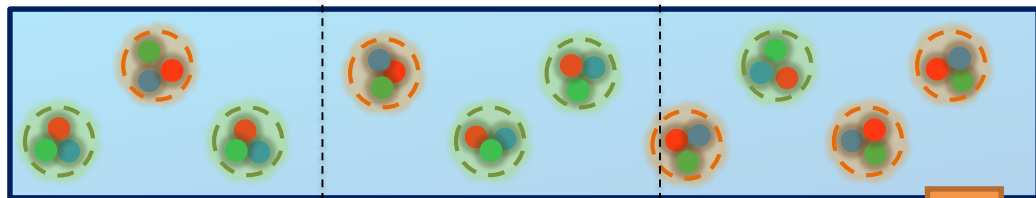


# Time Evolution in HIC

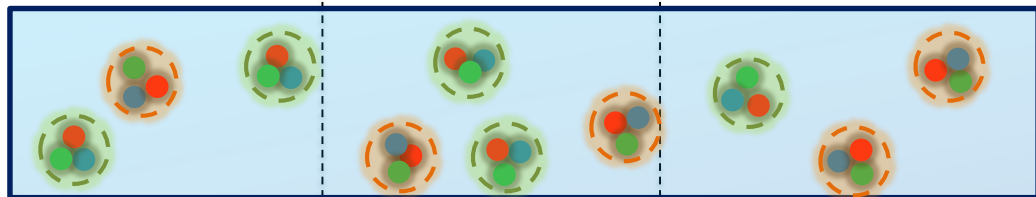
Quark-Gluon Plasma



Hadronization



Freezeout



$\Delta\eta$

$$\frac{\langle \Delta N^2 \rangle}{\Delta\eta}$$

$\chi_{\text{HAD}}$

$\chi_{\text{QGP}}$

$\Delta\eta$

$\chi_{\text{HAD}}$

$\chi_{\text{QGP}}$

$\Delta\eta$

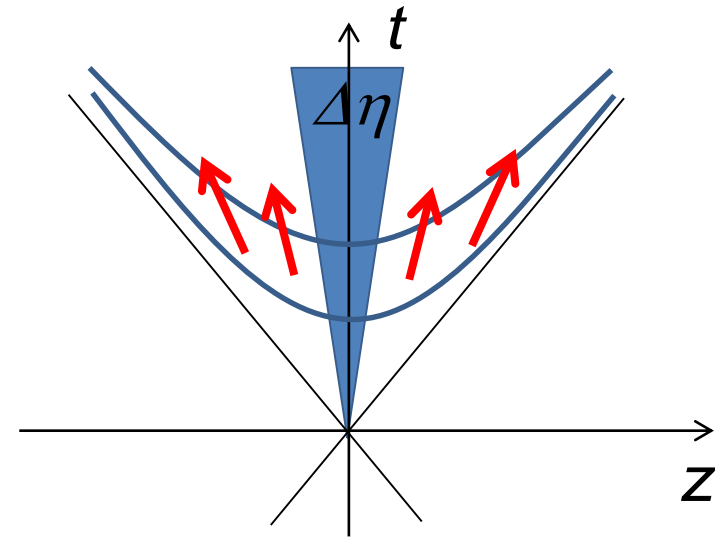
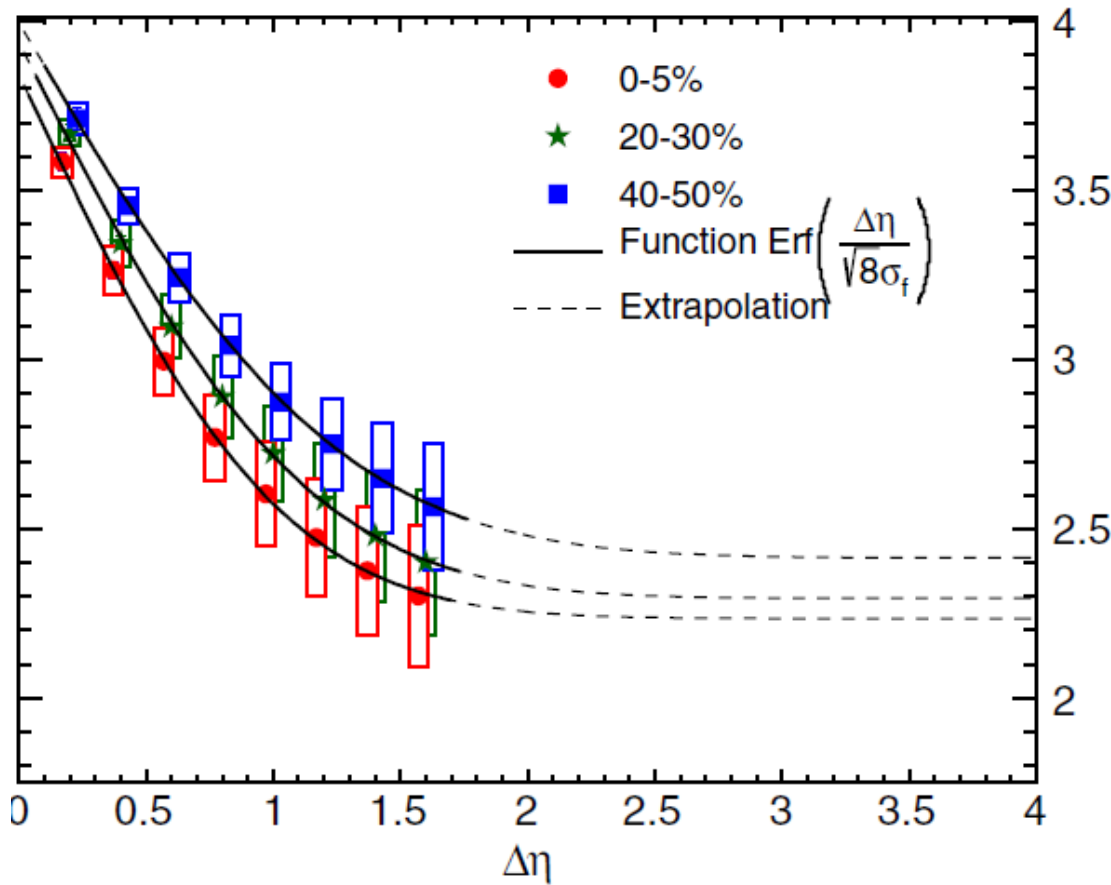
$\chi_{\text{HAD}}$

$\chi_{\text{QGP}}$

$\Delta\eta$

# $\Delta\eta$ Dependence @ ALICE

ALICE  
PRL 2013

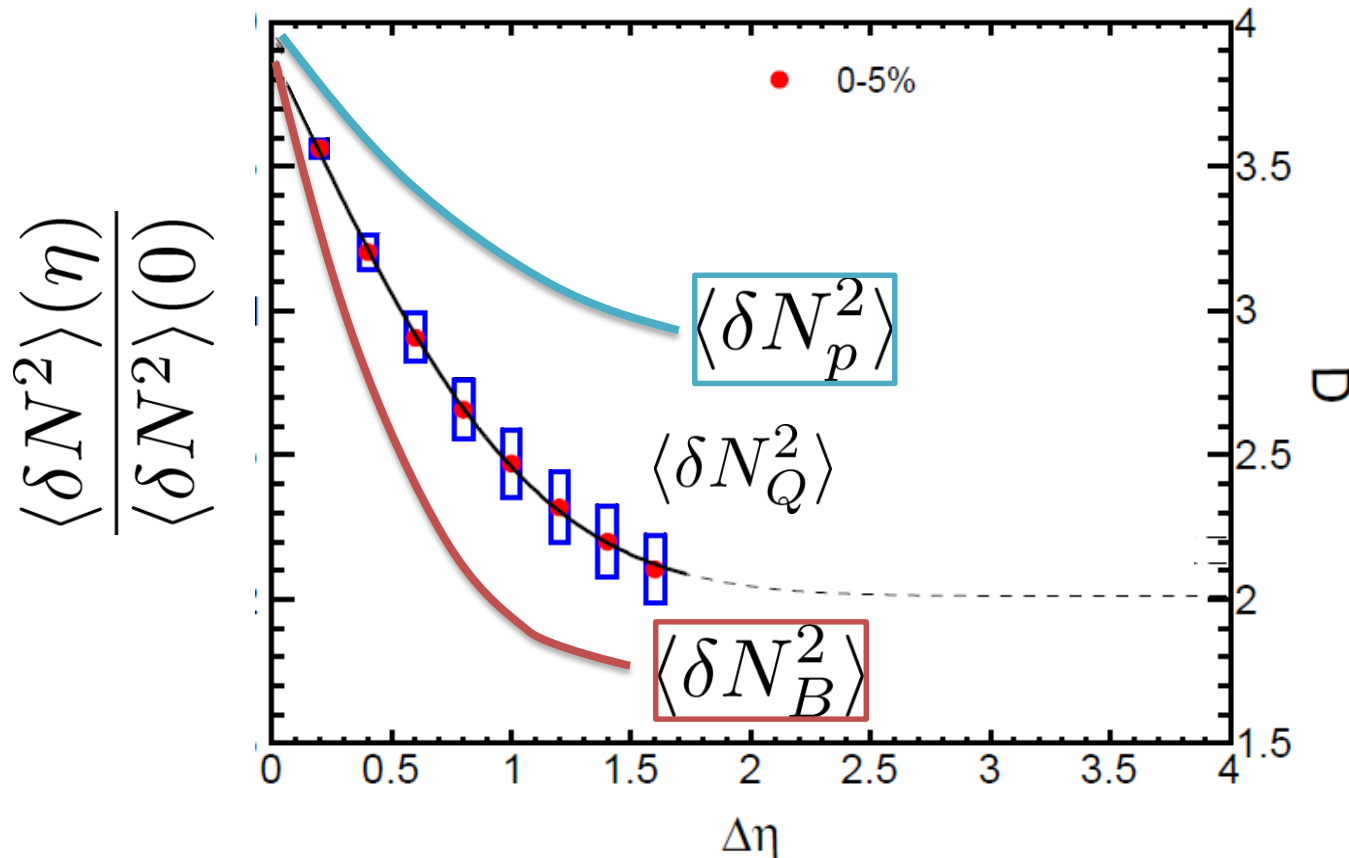


ゆらぎの $\Delta\eta$  依存性には、高温物質の  
時間発展の情報が刻まれている！

# $\langle \delta N_B^2 \rangle$ and $\langle \delta N_p^2 \rangle$ @ LHC ?

$$\langle \delta N_Q^2 \rangle, \langle \delta N_B^2 \rangle, \langle \delta N_p^2 \rangle$$

should have different  $\Delta\eta$  dependence.



$$\langle (\delta N_p^{(\text{net})})^2 \rangle = \frac{1}{4} \langle (\delta N_B^{(\text{net})})^2 \rangle + \frac{1}{4} \langle N_B^{(\text{tot})} \rangle$$

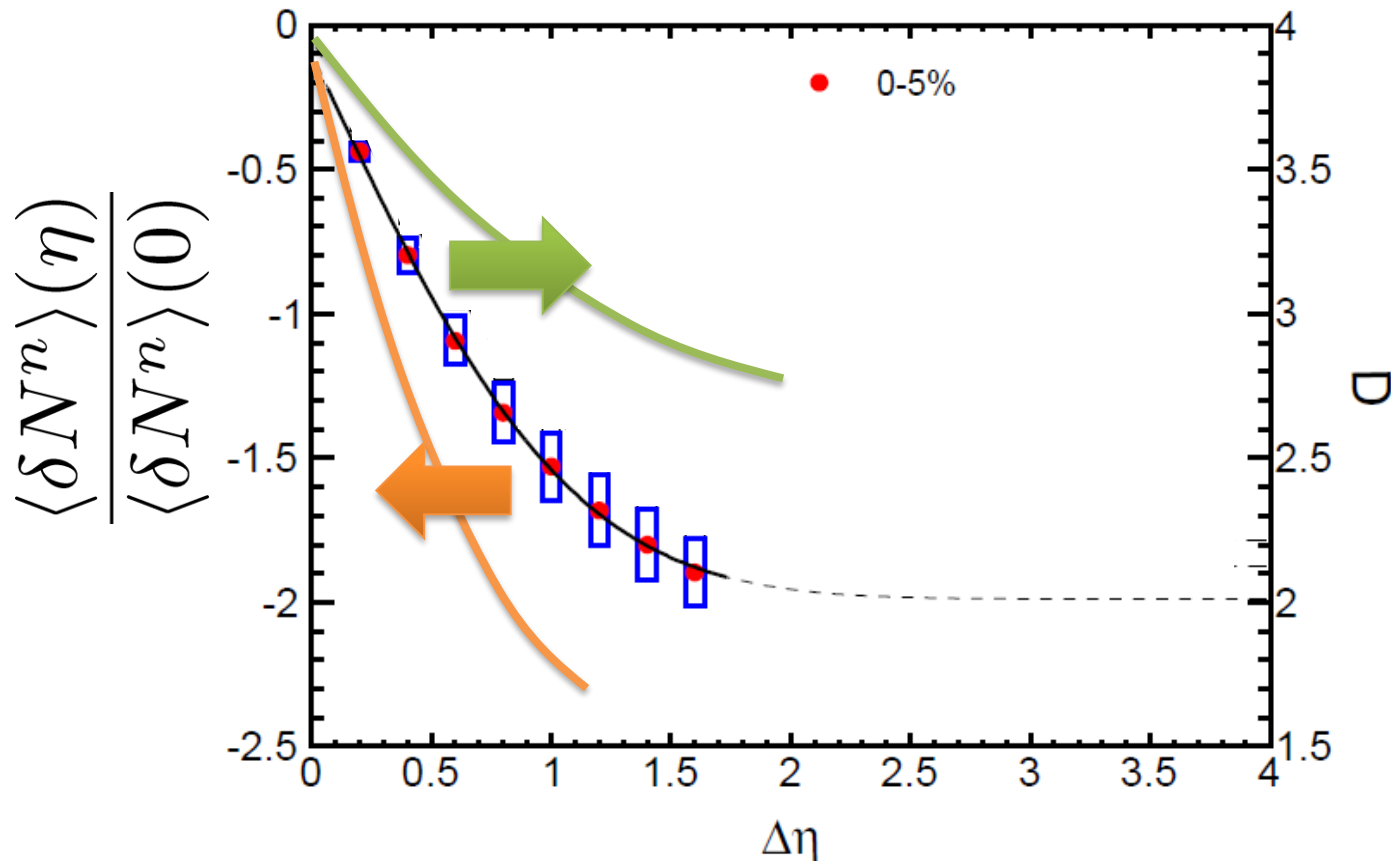
# $\langle \delta N_Q^4 \rangle @ \text{LHC ?}$

How does  $\langle \delta N_Q^4 \rangle_c$  behave as a function of  $\Delta\eta$ ?

Left  
(suppression)

or

Right  
(hadronic)

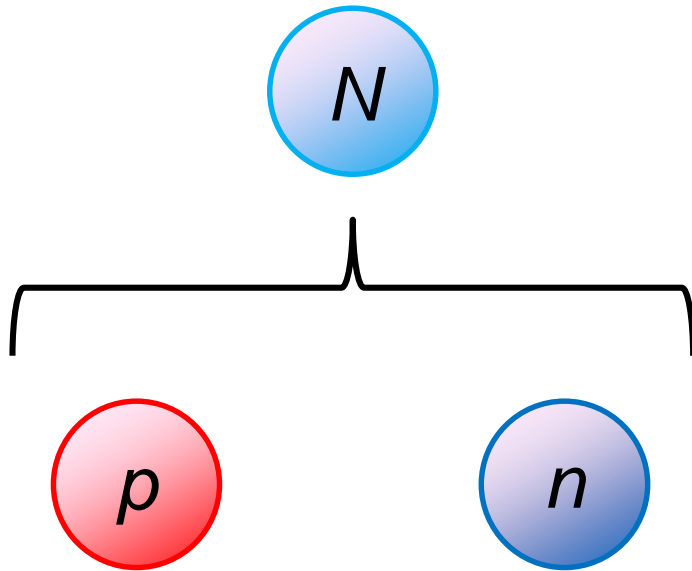


# Baryon vs Proton Number Fluctuations

MK, Asakawa, PRC85,021901C(2012); PRC86, 024904(2012)

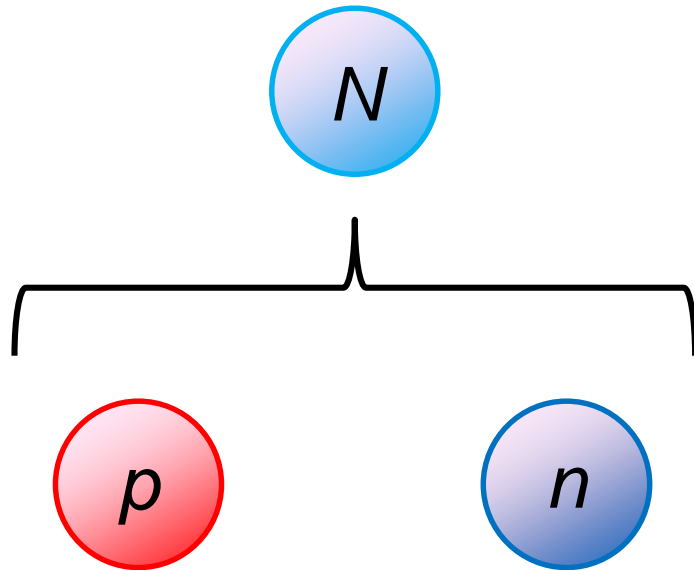
- $\frac{\langle \delta N_B^n \rangle_c}{\langle \delta N_B^m \rangle_c} \neq \frac{\langle \delta N_p^n \rangle_c}{\langle \delta N_p^m \rangle_c}$
- $\langle \delta N_B^n \rangle_c$  are experimentally observable

# Nucleon Isospin as Two Sides of a Coin

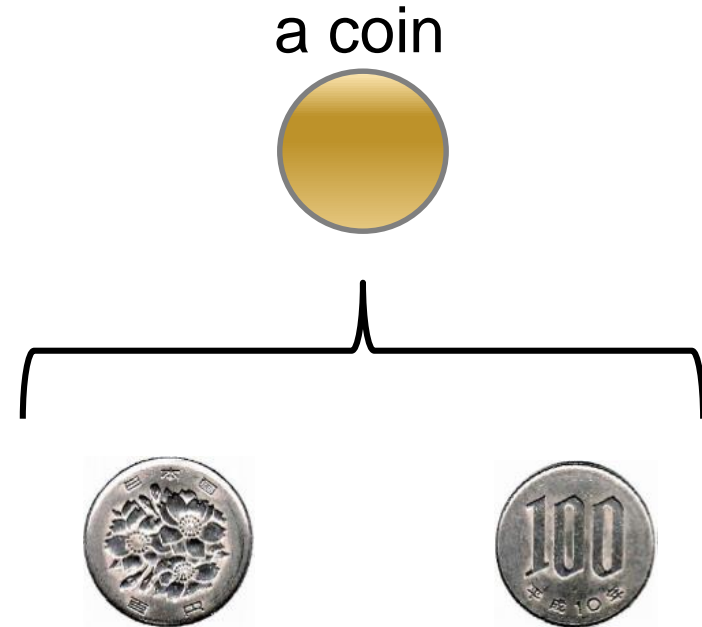


Nucleons have  
two isospin states.

# Nucleon Isospin as Two Sides of a Coin



Nucleons have two isospin states.



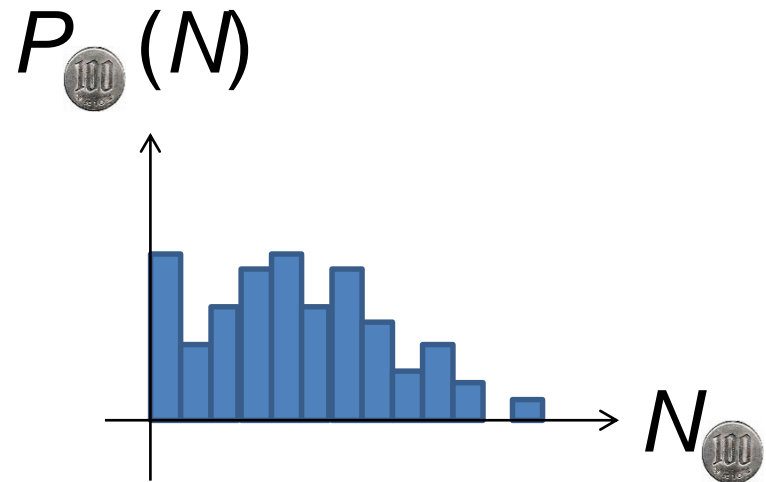
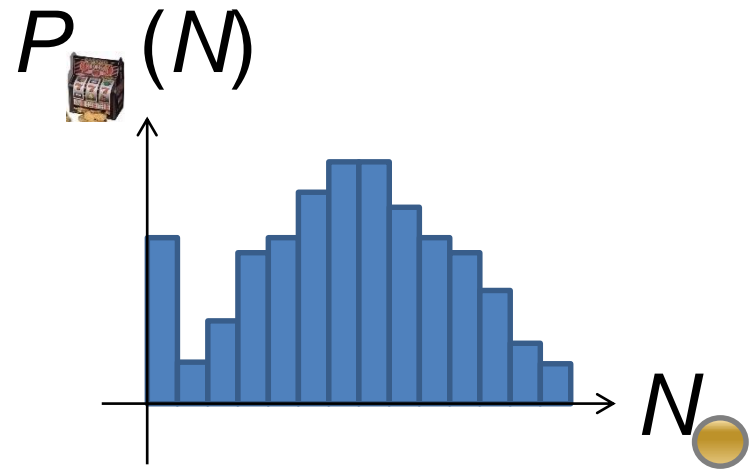
Coins have two sides.



# Slot Machine Analogy



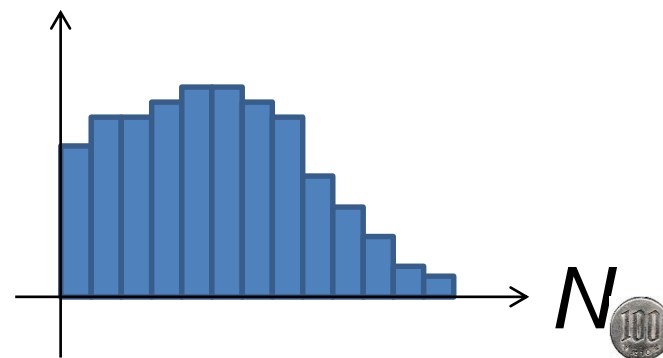
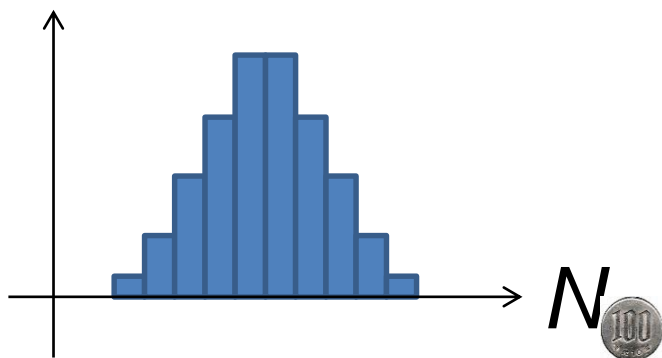
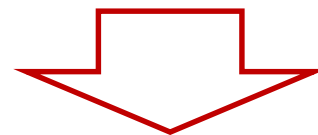
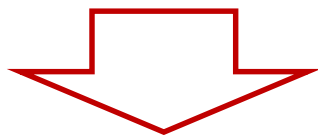
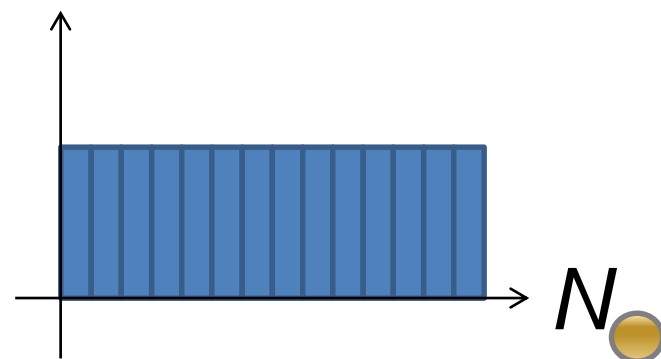
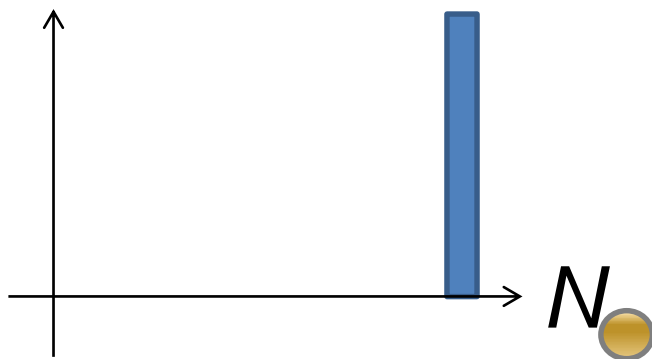
$$\text{Gold Coin} = 100 \text{ Silver Coin} + \text{Small Silver Coin}$$



# Extreme Examples

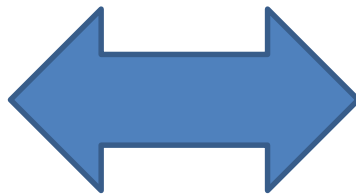
Fixed # of coins

Constant probabilities



# Reconstructing Total Coin Number

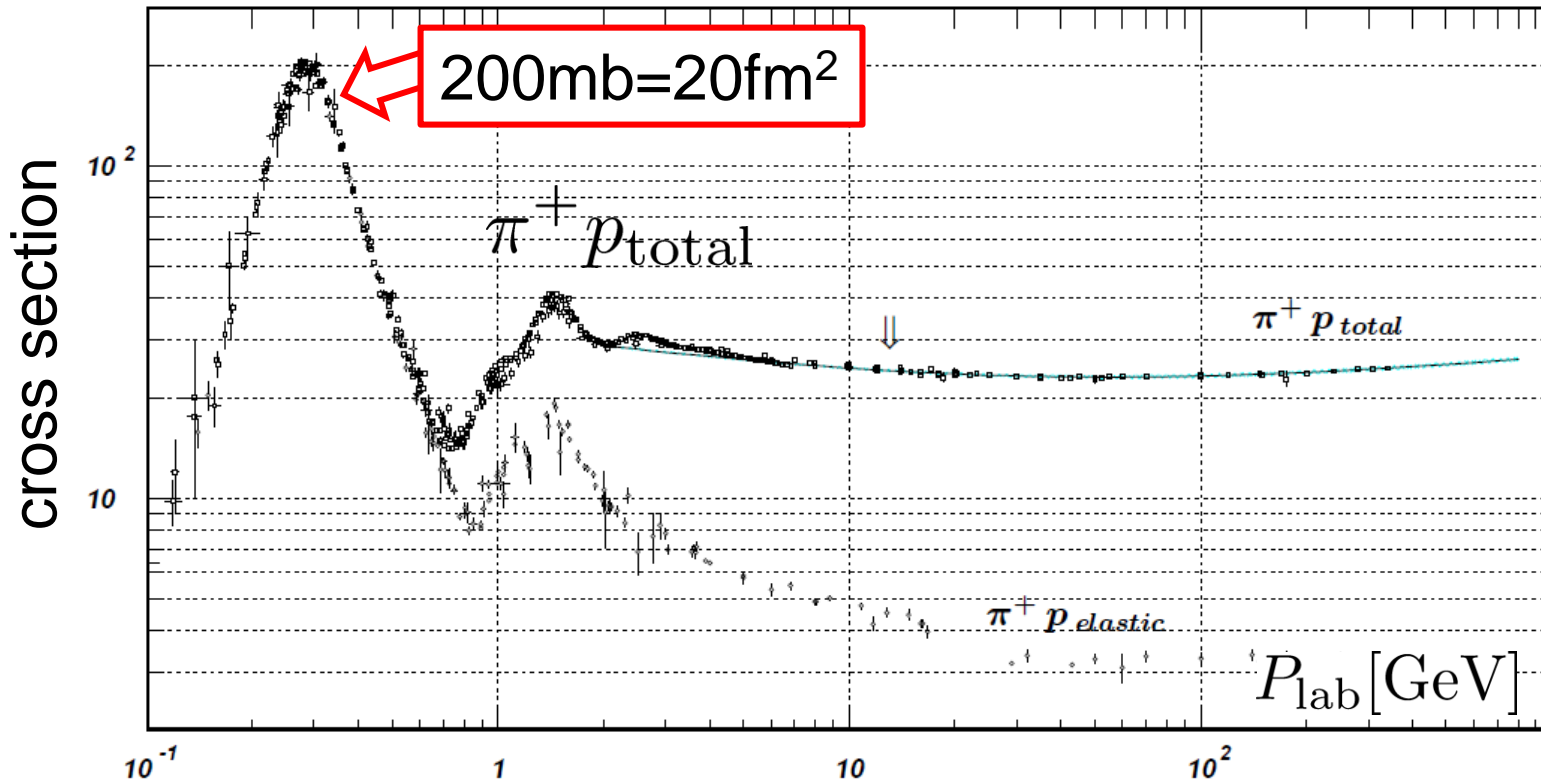
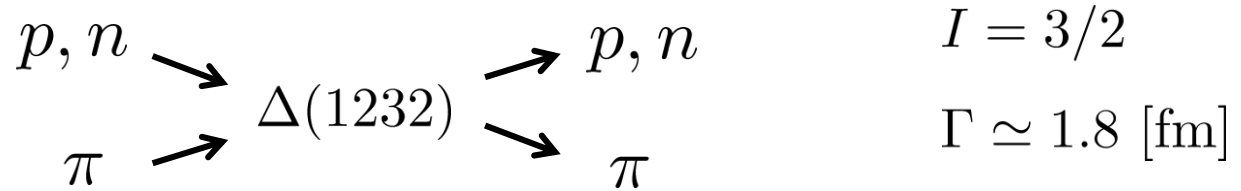
$$P_{\text{100}}(N_{\text{100}}) = \sum_{\text{slot}} P_{\text{slot}}(N_{\text{slot}}) B_{1/2}(N_{\text{100}}; N_{\text{slot}})$$



$$B_p(k; N) = p^k (1 - p)^{N-k} {}_k C_N \quad \text{:binomial distr. func.}$$

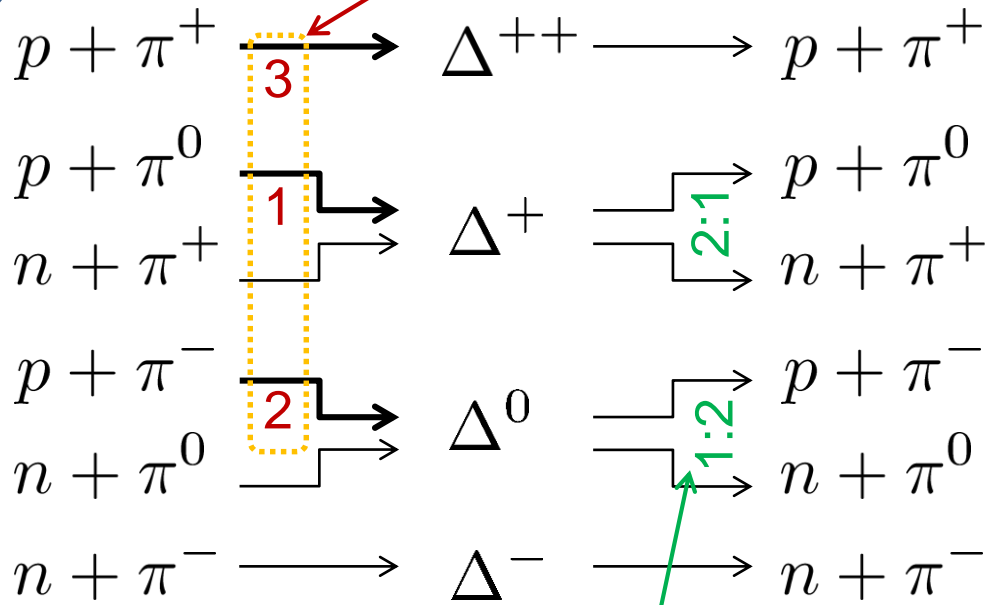
# Nucleon Isospin in Hadronic Medium

- Isospin of baryons can vary after chemical freezeout via charge exchange reactions mediated by  $\Delta(1232)$ :



# $\Delta(1232)$

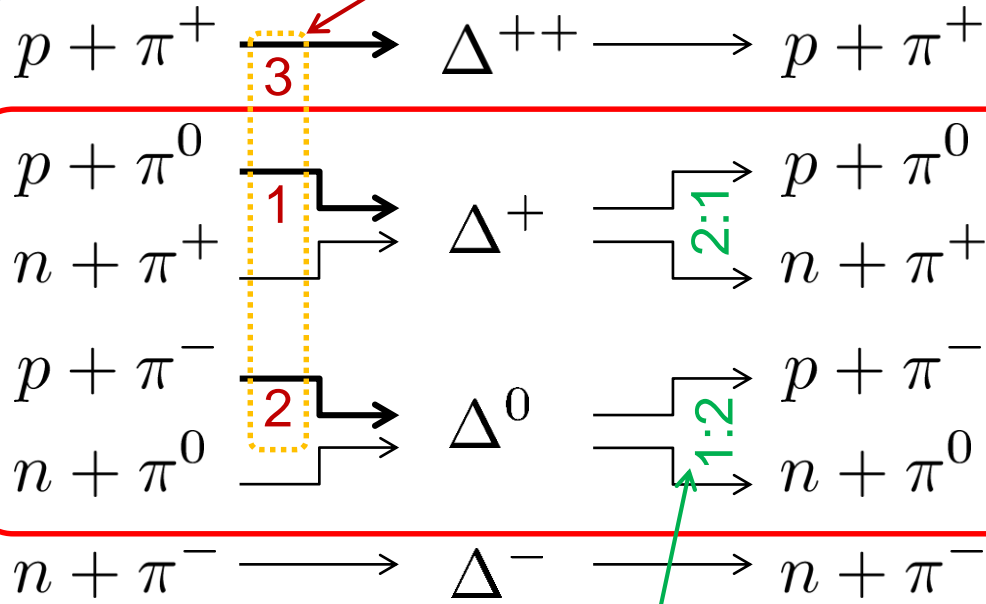
cross sections of  $p$



decay rates of  $\Delta$

# $\Delta(1232)$

cross sections of  $p$

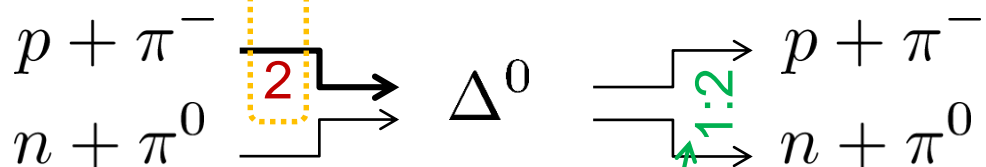
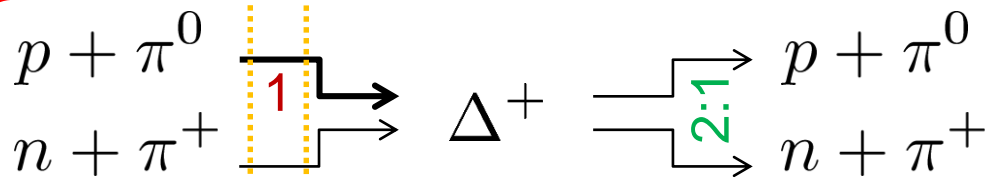


decay rates of  $\Delta$

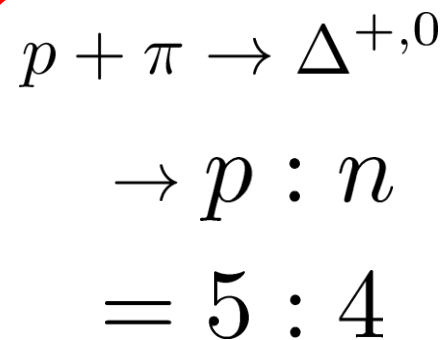
$$\begin{aligned} p + \pi &\rightarrow \Delta^{+,0} \\ &\rightarrow p : n \\ &= 5 : 4 \end{aligned}$$

# $\Delta(1232)$

cross sections of  $p$



decay rates of  $\Delta$

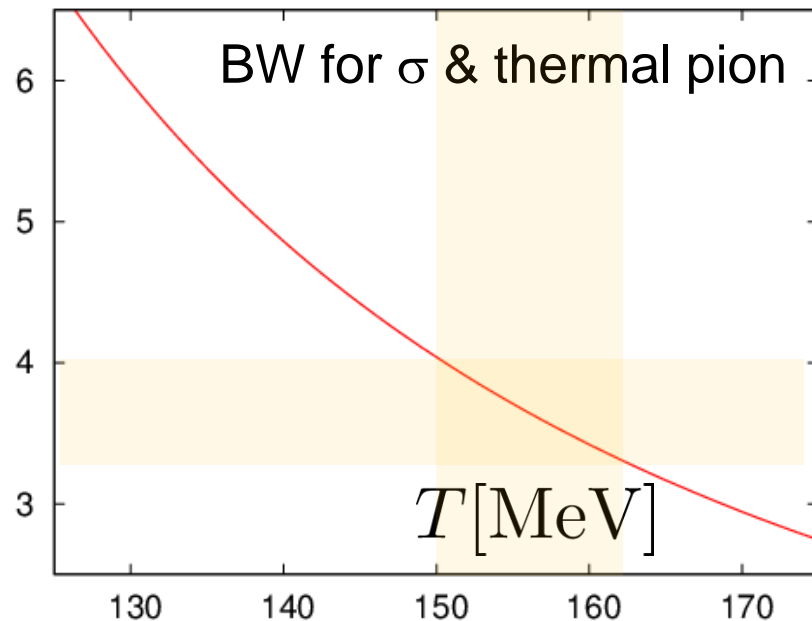


Lifetime to create  $\Delta^+$  or  $\Delta^0$

$$\tau^{-1} = \int \frac{d^3 k_\pi}{(2\pi)^3} \sigma(E_{\text{cm}}) v_\pi n(E_\pi)$$

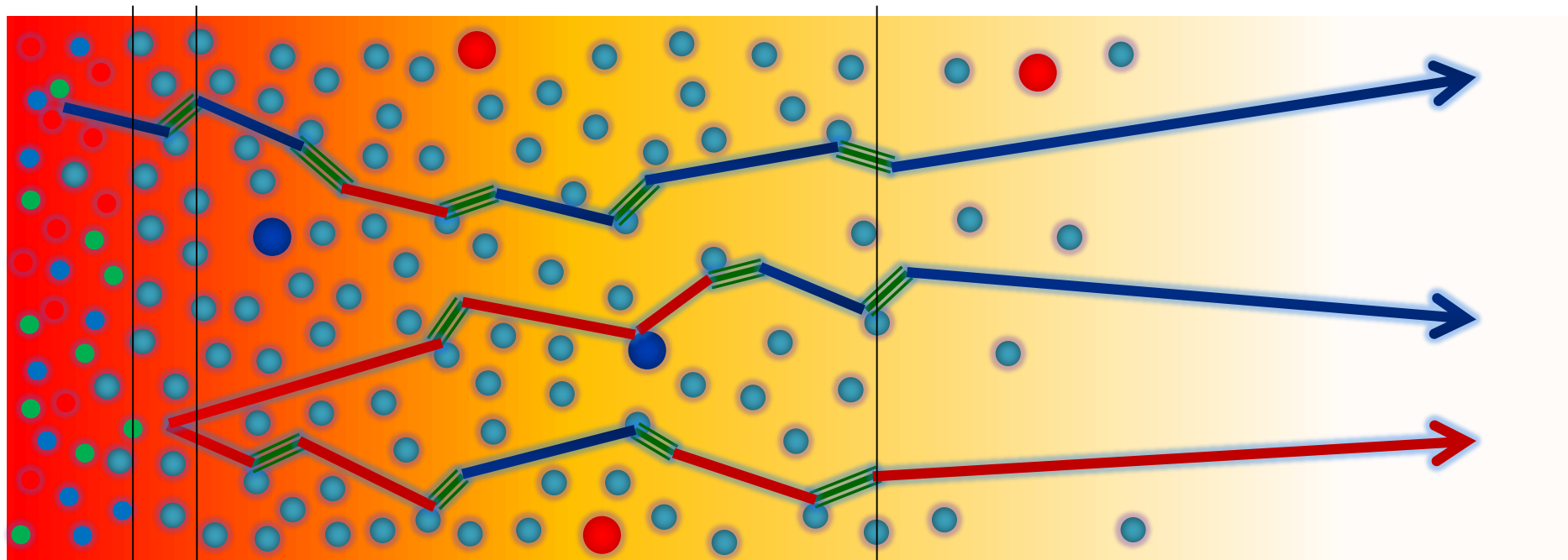
$\tau$  [fm]

(freezeout time)  $\simeq 20$  [fm]



# Nucleons in Hadronic Phase

time →



hadronize  
chem. f.o.

10~20fm

kinetic f.o.

- $p, \bar{p}$
- $n, \bar{n}$
- ≡≡≡  $\Delta(1232)$
- mesons
- baryons

$$m_\pi \simeq T \ll m_N - \mu_N$$

$$n_N \ll 1$$

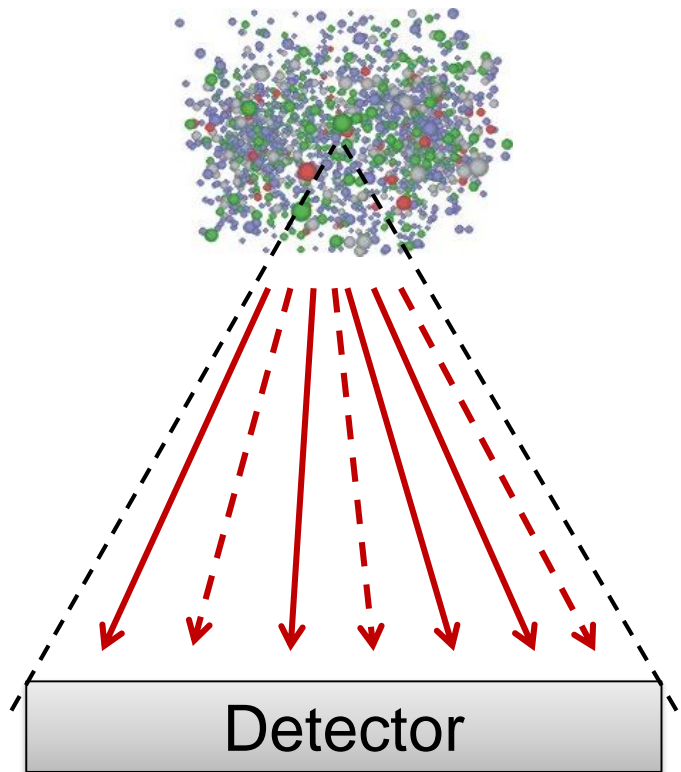
- rare NN collisions
- no quantum corr.

$$n_N \ll n_\pi$$

- many pions



# Probability Distribution $\mathcal{P}(N_p, N_n, N_{\bar{p}}, N_{\bar{n}})$



$\square$   $\left\{ \begin{array}{l} \longrightarrow N_N \text{ nucleons} \\ \dashrightarrow N_{\bar{N}} \text{ anti-nucleons} \end{array} \right.$

$\longrightarrow F(N_N, N_{\bar{N}})$

$\square$   $N_N \left\{ \begin{array}{l} N_p \text{ protons} \\ N_n \text{ neutrons} \end{array} \right.$

$\longrightarrow B(N_p; N_N)$

binomial distribution func.

$$\mathcal{P}(N_p, N_n, N_{\bar{p}}, N_{\bar{n}})$$

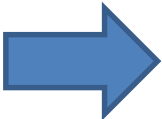
$$= F(N_N, N_{\bar{N}}) B(N_p; N_N) B(N_{\bar{p}}; N_{\bar{N}})$$

➤ for any phase space in the final state.

# Difference btw Baryon and Proton Numbers

(1)  $N_B^{(\text{net})} = N_B - N_{\bar{B}}$  deviates from the equilibrium value.

(2) Boltzmann (Poisson) distribution for  $N_B, N_{\bar{B}}$ .



$$\begin{aligned}
 2\langle(\delta N_p^{(\text{net})})^2\rangle &= \frac{1}{2}\langle(\delta N_B^{(\text{net})})^2\rangle + \frac{1}{2}\langle(\delta N_B^{(\text{net})})^2\rangle_{\text{free}} \\
 2\langle(\delta N_p^{(\text{net})})^3\rangle &= \frac{1}{4}\langle(\delta N_B^{(\text{net})})^3\rangle + \frac{3}{4}\langle(\delta N_B^{(\text{net})})^3\rangle_{\text{free}} \\
 2\langle(\delta N_p^{(\text{net})})^4\rangle_c &= \frac{1}{8}\langle(\delta N_B^{(\text{net})})^4\rangle_c + \dots
 \end{aligned}$$

genuine info.
noise

For free gas

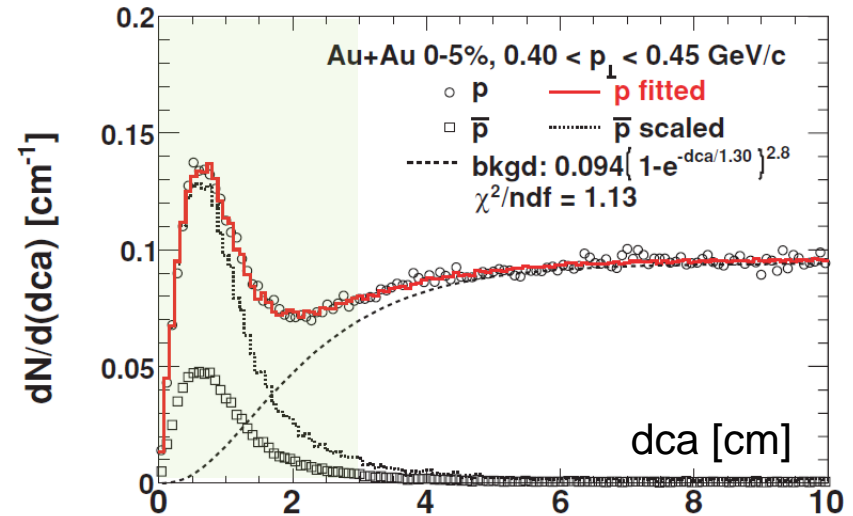
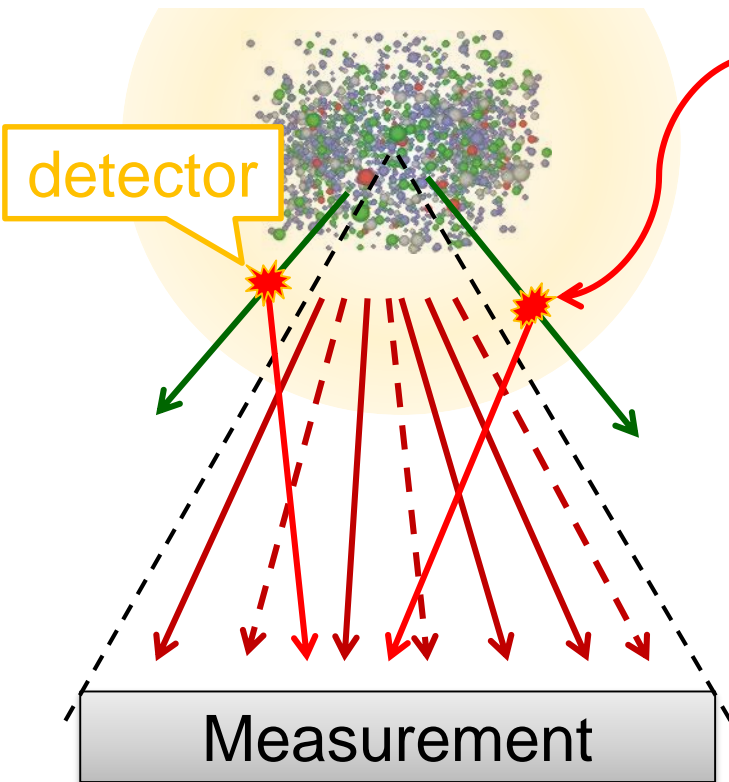
$$2\langle(\delta N_p^{(\text{net})})^n\rangle_c = \langle(\delta N_N^{(\text{net})})^n\rangle_c$$

# Secondary Protons

MK+, in preparation

Secondary (knockout) protons

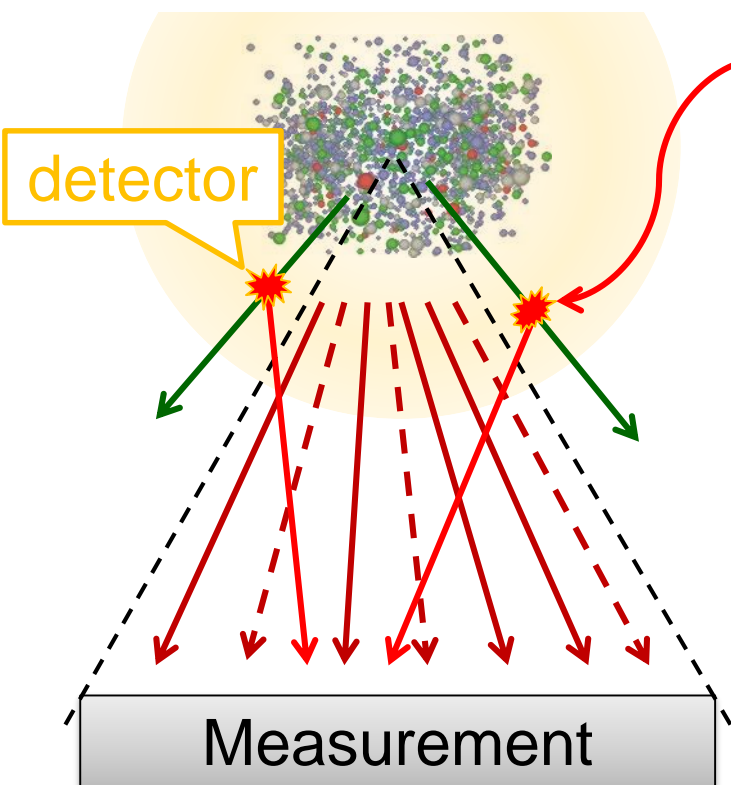
20% of observed protons @ STAR



STAR, PRC79,034909(2009)

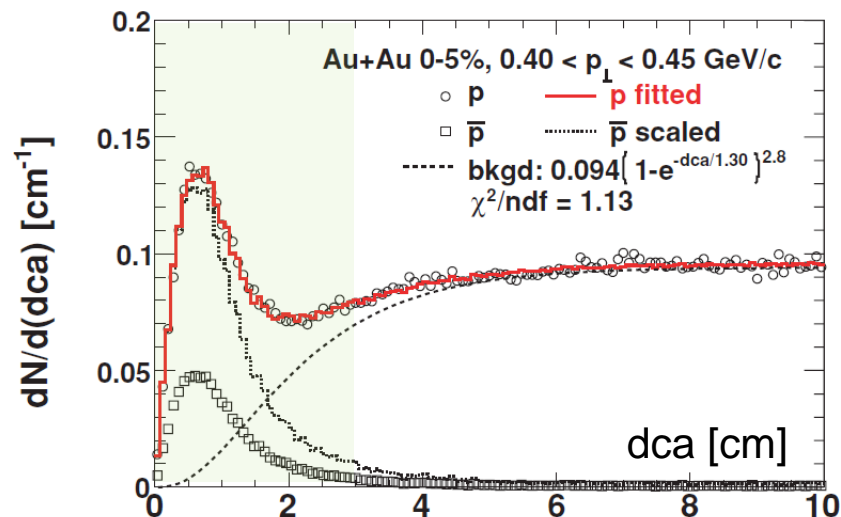
# Secondary Protons

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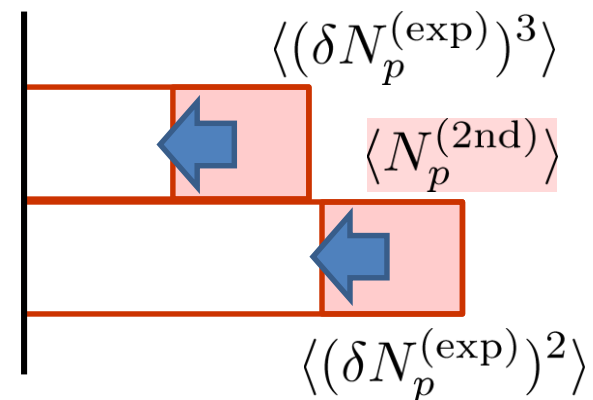
20% of observed protons @ STAR



STAR, PRC79,034909(2009)

Their contribution can be eliminated!

$$\langle (\delta N_p^{(\text{QGP})})^n \rangle_c = \langle (\delta N_p^{(\text{exp})})^n \rangle_c - \langle N_p^{(2\text{nd})} \rangle$$

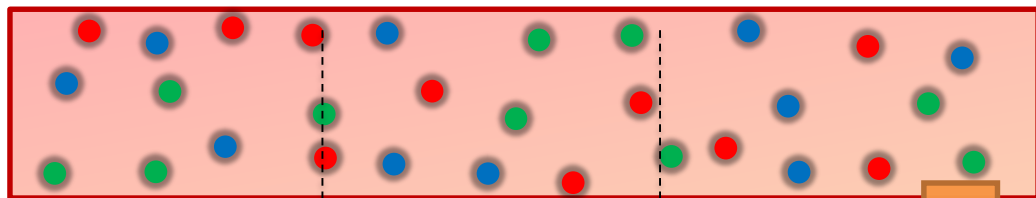


# Time Evolution of Higher Order Cumulants

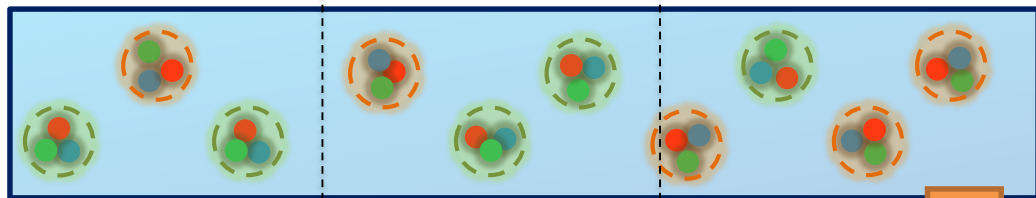
MK, Asakawa, Ono, arXiv:1307.xxxx

# Time Evolution in HIC

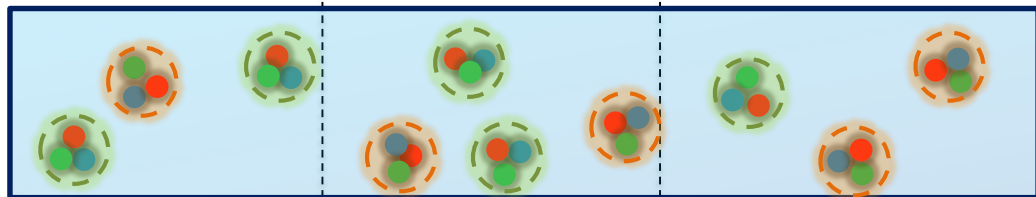
Quark-Gluon Plasma



Hadronization

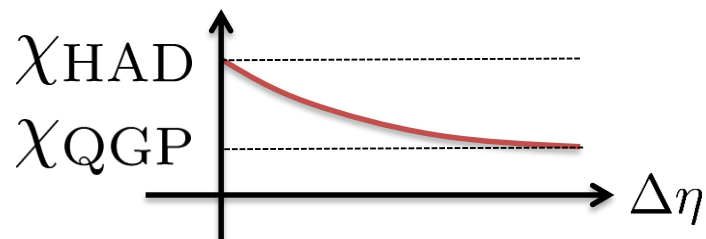
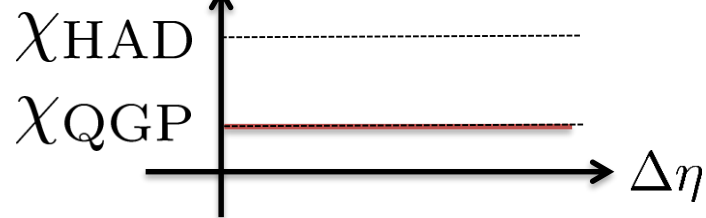


Freezeout



$\Delta\eta$

$$\frac{\langle \Delta N^2 \rangle}{\Delta\eta}$$



# Hydrodynamic Fluctuations

Landau, Lifshitz, Statistical Mechanics II  
Kapusta, Muller, Stephanov, 2012

Diffusion equation

$$\partial_\tau n = D \partial_\eta^2 n$$



**Stochastic** diffusion equation

$$\partial_\tau n = D \partial_\eta^2 n + \partial_\eta \xi(\eta, \tau)$$

# $\Delta\eta$ Dependence

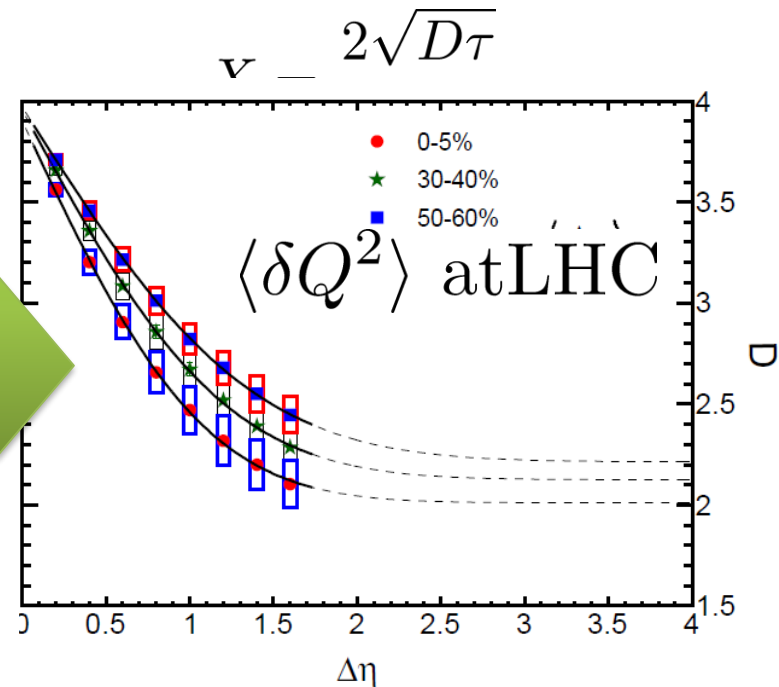
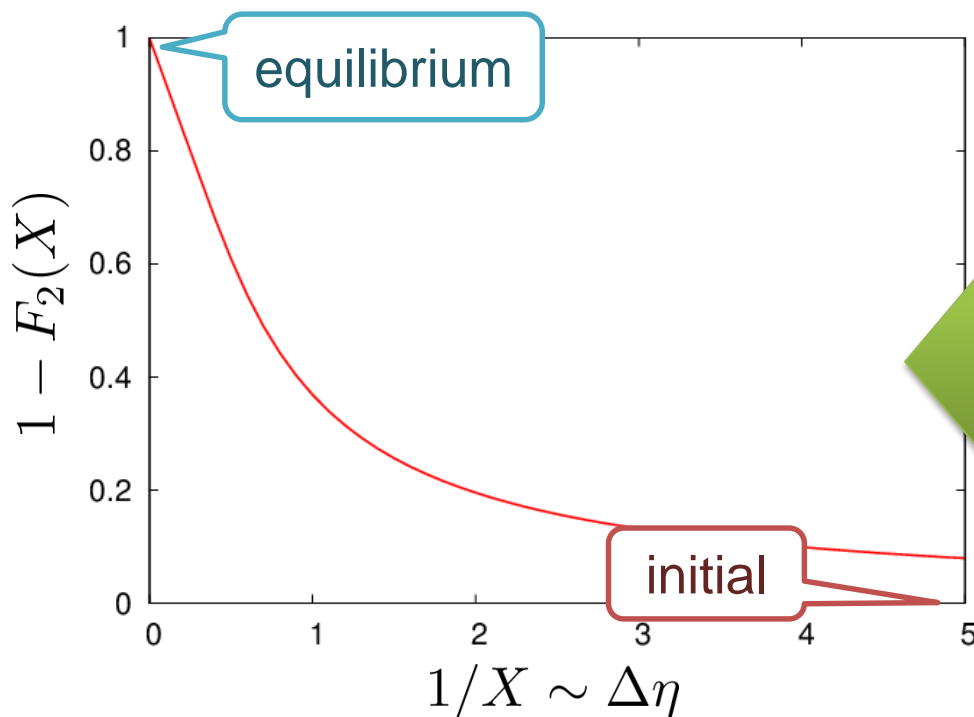
Shuryak, Stephanov, 2001

- Initial condition:  $\langle \delta n(\eta_1, 0) \delta n(\eta_2, 0) \rangle = \sigma_2 \delta(\eta_1 - \eta_2)$
- Translational invariance



$$\langle \delta Q(\tau)^2 \rangle = \underbrace{\sigma_2 F_2(X)}_{\text{initial}} + \underbrace{\chi_2 (1 - F_2(X))}_{\text{equilibrium}}$$

$$Q(\tau) = \int_0^{\Delta\eta} d\eta n(\eta, \tau)$$





# Thee “NON”s

重イオン衝突での高次ゆらぎの観測・解析は、  
物理学として相当に特殊な問題である。

## □ Non-Gaussian

通常、高次ゆらぎは観測困難。  
適度に小さい系

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観測されたゆらぎの値は、  
自由ガスとたかだか2倍のずれ

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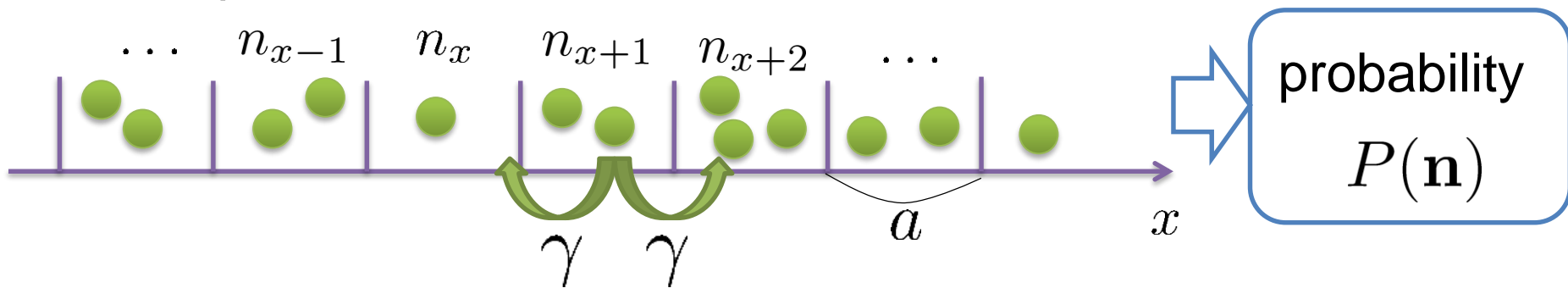
観測されたゆらぎの値は、  
自由ガスとたかだか2倍のずれ

## □ Non-equilibrium

平衡に至る非定常過程を  
記述する必要性。

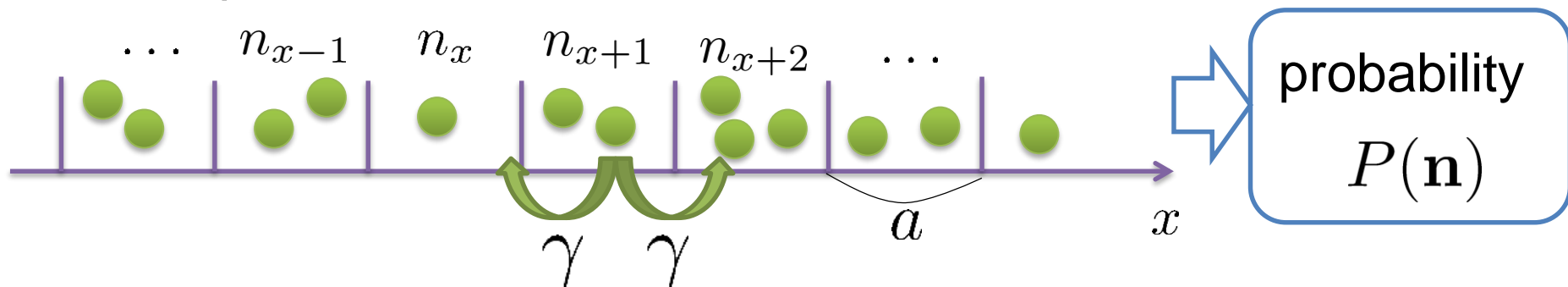
# Diffusion Master Equation

Divide spatial coordinate into discrete cells



# Diffusion Master Equation

Divide spatial coordinate into discrete cells



Master Equation for  $P(n)$

$$\frac{\partial}{\partial t} P(\mathbf{n}) = \gamma \sum_x [(n_x + 1) \{ P(\mathbf{n} + \hat{x} - \widehat{x+1}) + P(\mathbf{n} + \hat{x} - \widehat{x-1}) \} - 2n_x P(\mathbf{n})]$$


x-hat: lattice-QCD notation


Solve the DME **exactly**, and take  $a \rightarrow 0$  limit

No approx., ex. van Kampen's system size expansion

# Solution of DME

1st  $\langle \tilde{n}_k \rangle(t) = e^{-\omega_k t} \langle \tilde{n}_k \rangle_0$        $\omega_k \simeq \gamma a^2 k^2$

 initial


 Deterministic part  $\leftrightarrow$  diffusion equation at long wave length ( $1/a \ll k$ )


$$\partial_t \langle n_x(t) \rangle = \gamma a^2 \partial_x^2 \langle n_x(t) \rangle$$

 Appropriate continuum limit with  $\gamma a^2 = D$

# Solution of DME


1st  $\langle \tilde{n}_k \rangle(t) = e^{-\omega_k t} \langle \tilde{n}_k \rangle_0$        $\omega_k \simeq \gamma a^2 k^2$

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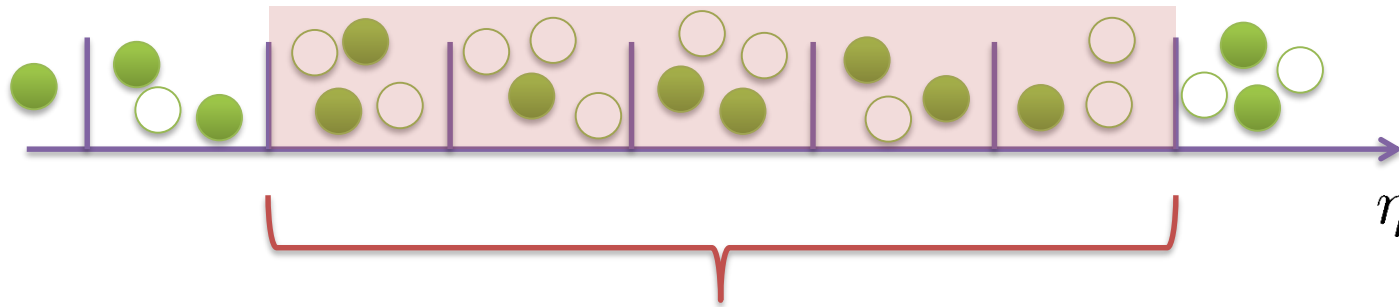
 Appropriate continuum limit with  $\gamma a^2 = D$

2nd  $\langle \delta \tilde{n}_{k_1} \delta \tilde{n}_{k_2} \rangle(t) = \langle \tilde{n}_{k_1+k_2} \rangle_0 (e^{-\omega_{k_1+k_2} t} - e^{-(\omega_{k_1} + \omega_{k_2}) t})$   
 $+ \langle \delta \tilde{n}_{k_1} \delta \tilde{n}_{k_2} \rangle_0 e^{-(\omega_{k_1} + \omega_{k_2}) t}$

 Consistent with stochastic diffusion eq. (for sufficiently smooth initial condition)

# Net Charge Number

Prepare 2 species of (non-interacting) particles



$$\bar{Q}(\tau) = \int_0^{\Delta\eta} d\eta (n_1(\eta, \tau) - n_2(\eta, \tau))$$

Let us investigate

$$\langle \bar{Q}^2 \rangle_c \quad \langle \bar{Q}^4 \rangle_c \quad \text{at freezeout time } t$$



# Initial Condition at Hadronization

- Boost invariance / infinitely long system
- Local equilibration / local correlation
- Initial fluctuations

$$\langle \bar{Q}^2 \rangle_c \quad \langle \bar{Q}^4 \rangle_c \quad \langle \bar{Q}^2 Q_{(\text{tot})} \rangle_c \quad \langle Q_{(\text{tot})}^2 \rangle_c$$

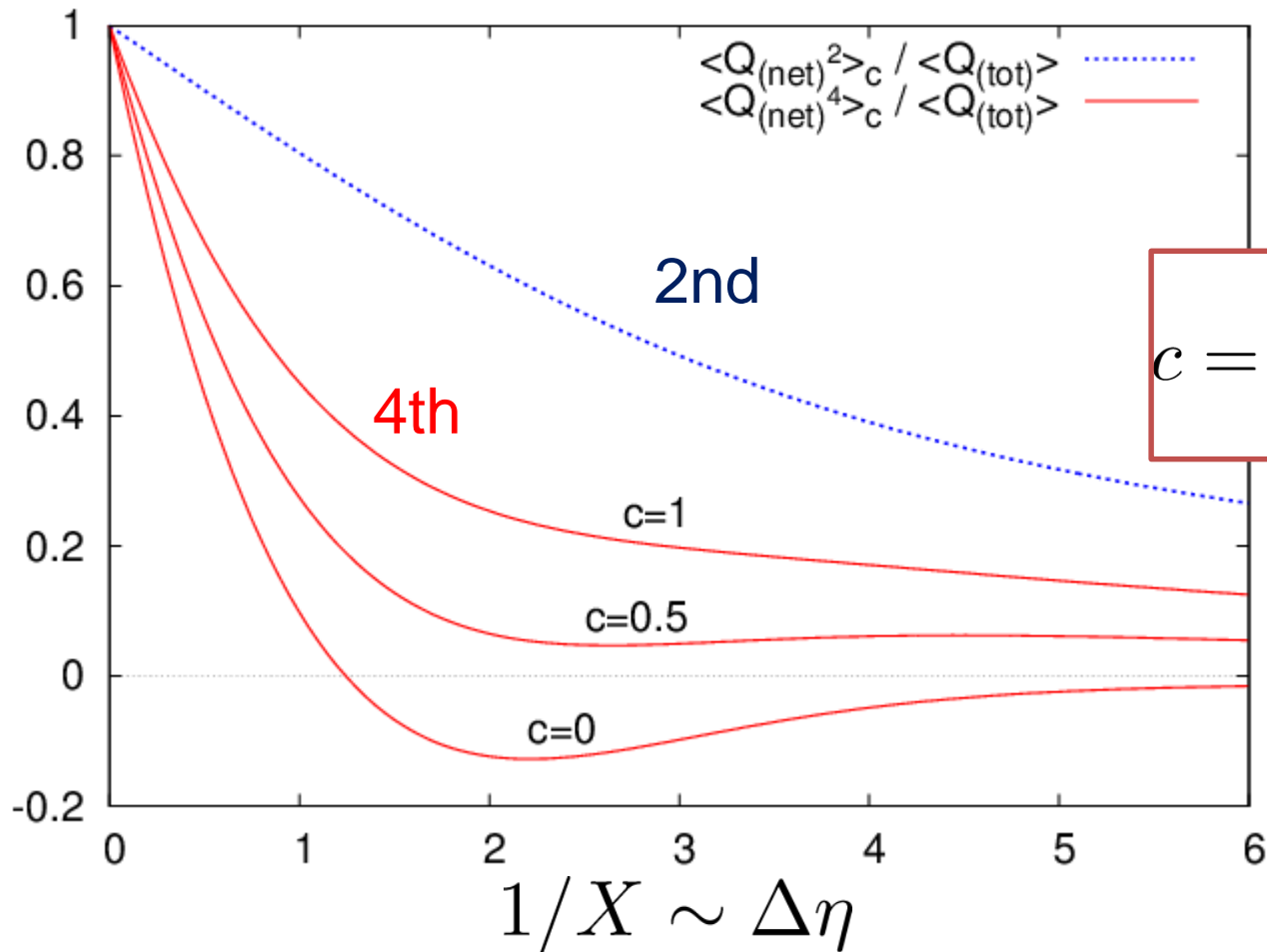
↑  
suppression owing to  
local charge conservation

↑  
strongly dependent on  
hadronization mechanism

# $\Delta\eta$ Dependence at Freezeout

Initial fluctuations:

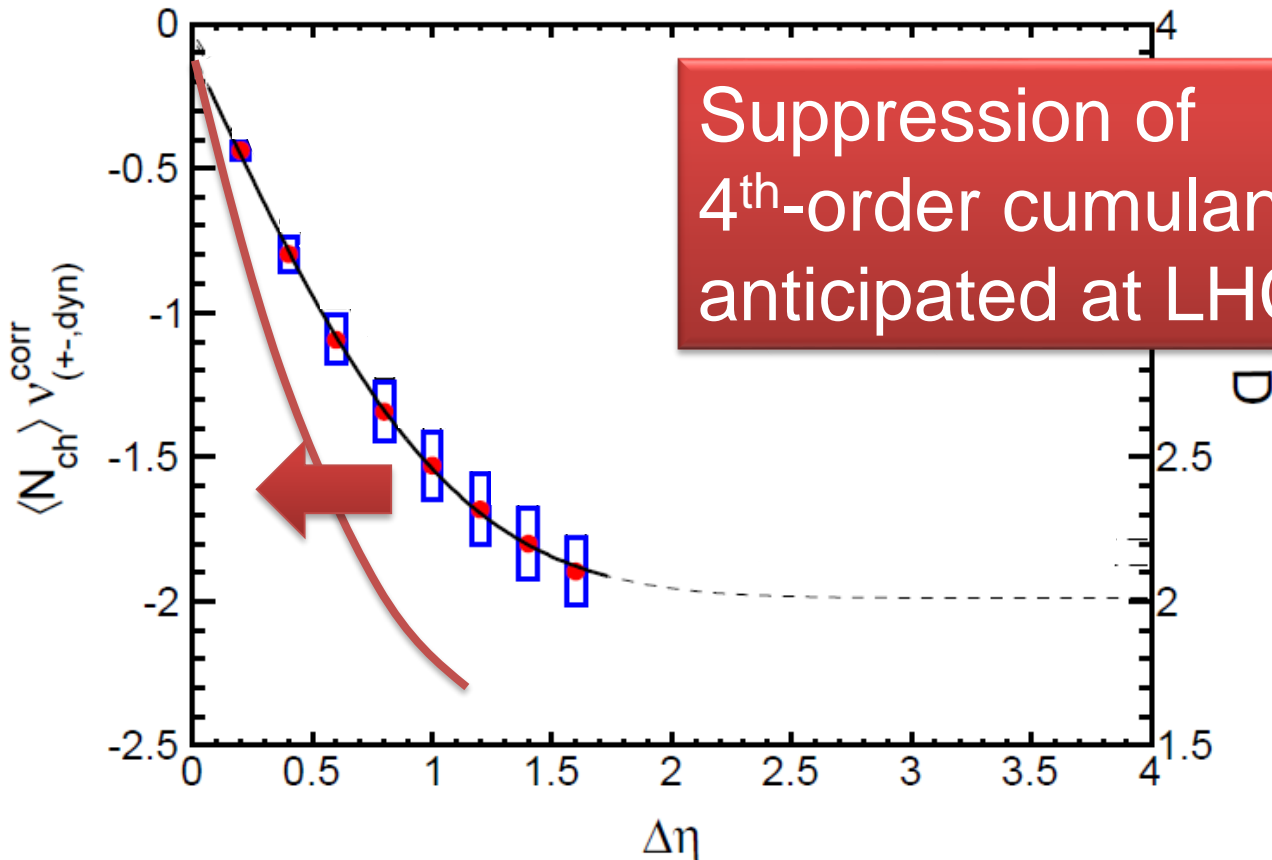
$$\langle \bar{Q}^2 \rangle_c = \langle \bar{Q}^4 \rangle_c = \langle \bar{Q}^2 Q_{(\text{tot})} \rangle_c = 0$$



# $\langle \delta N_Q^4 \rangle @ \text{LHC}$

Assumptions

- boost invariant system
- small fluctuations of CC at hadronization
- short correlation in hadronic stage

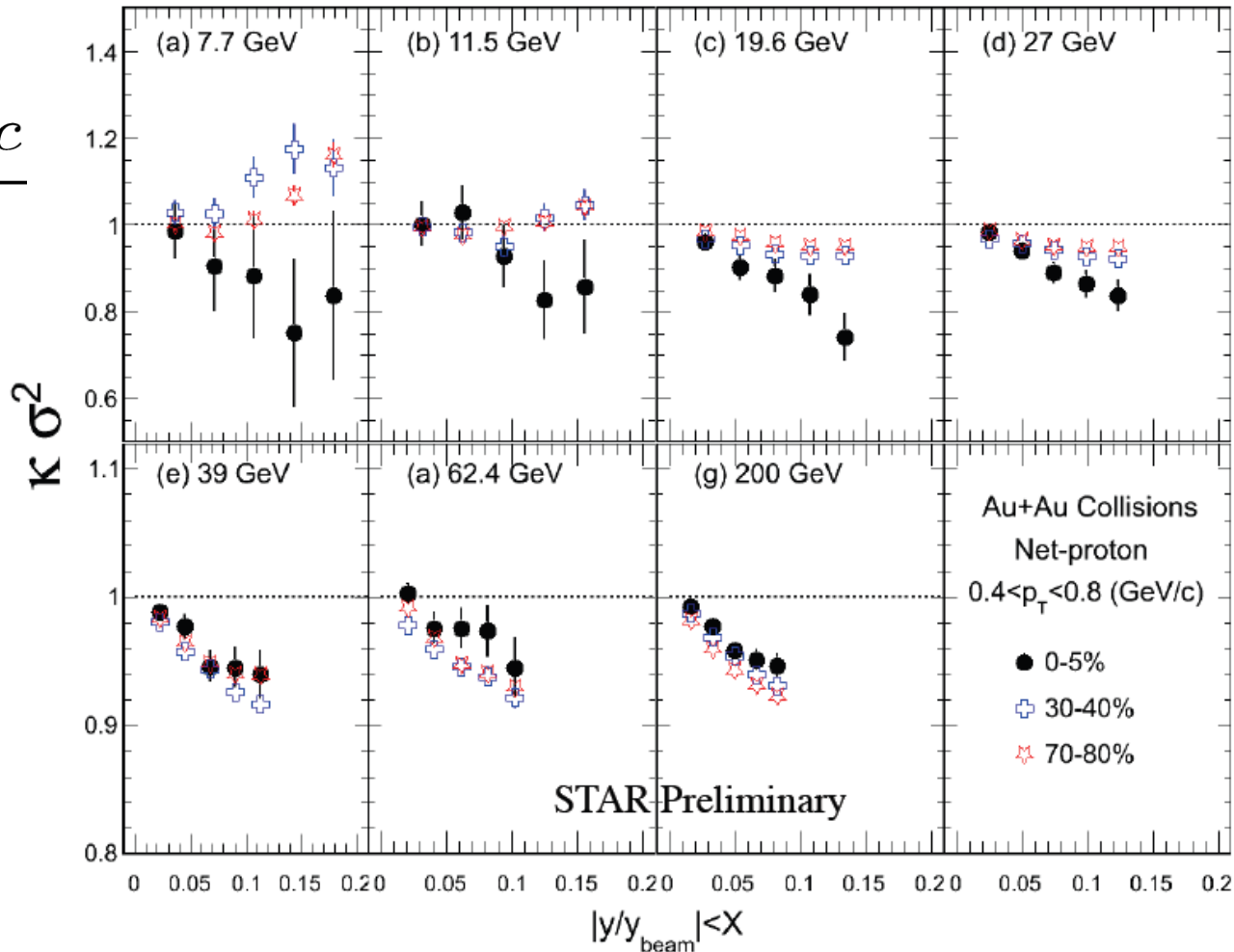


Suppression of  
4<sup>th</sup>-order cumulant is  
anticipated at LHC energy!

# $\Delta\eta$ Dependence at STAR

STAR, QM2012

$$\frac{\langle \delta N_p^4 \rangle_c}{\langle \delta N_p^2 \rangle}$$



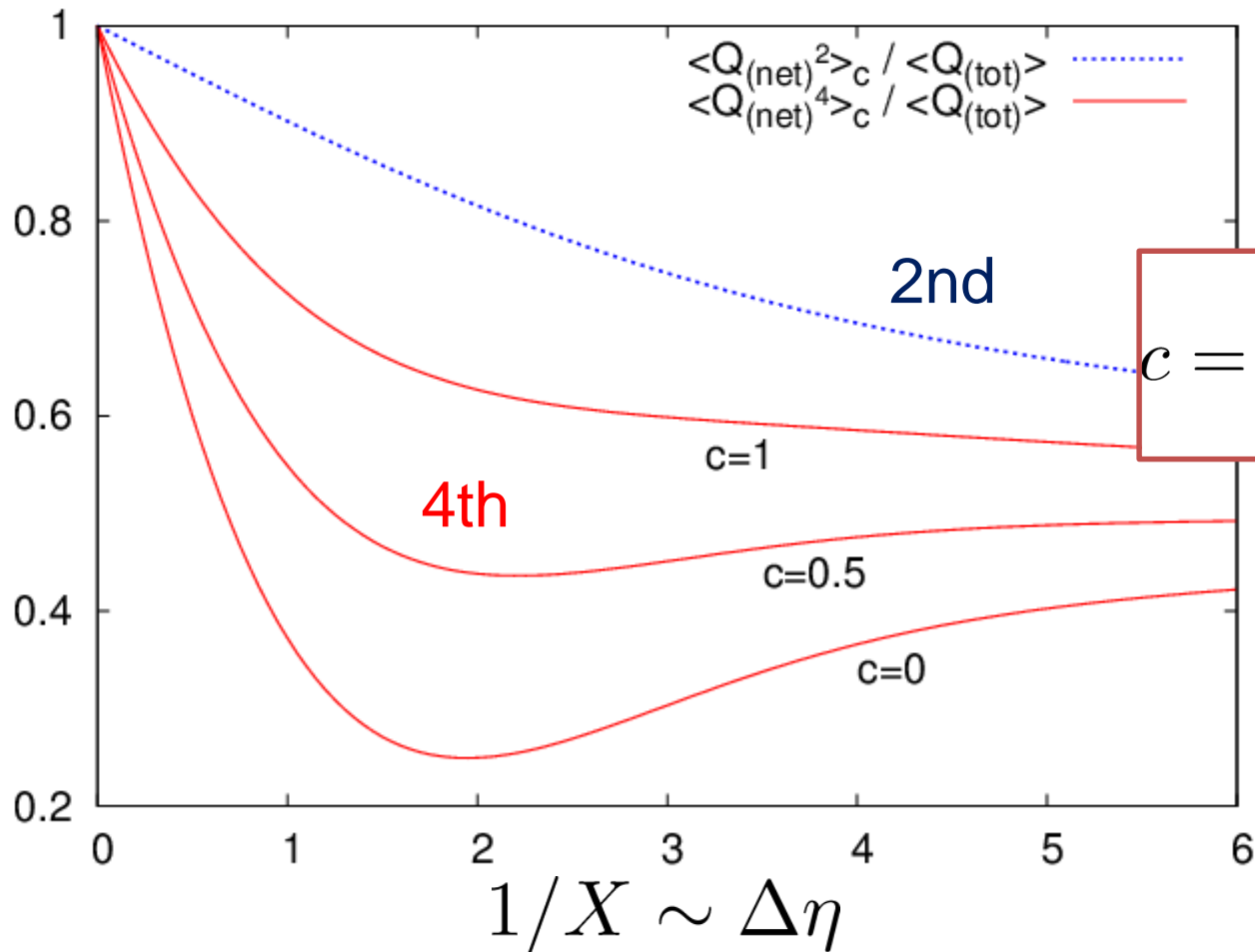
$$\frac{\langle \delta N_p^4 \rangle_c}{\langle \delta N_p^2 \rangle}$$

decreases as  $\Delta\eta$  becomes larger at RHIC.

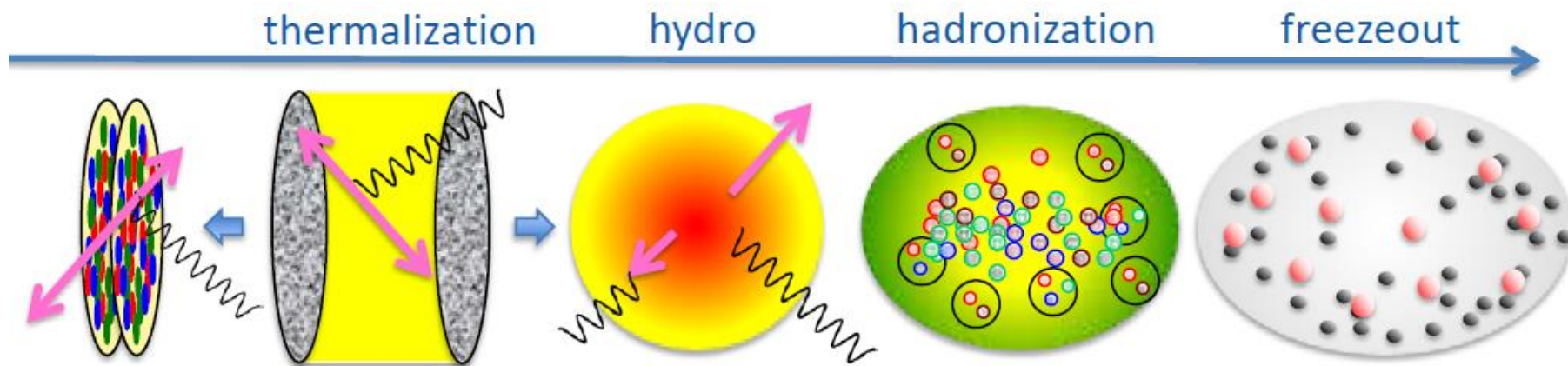
# $\Delta\eta$ Dependence at Freezeout

Initial fluctuations:

$$\langle \bar{Q}^2 \rangle_c = \langle \bar{Q}^4 \rangle_c = \langle \bar{Q}^2 Q_{(\text{tot})} \rangle_c = 0.5 \langle Q_{(\text{tot})} \rangle$$



# 高温物質の時間発展



熱平衡化  
段階での  
ゆらぎ

QGP相内での  
時間発展

拡散による  
HRGへの接近

検出器

初期エネルギー密度 → 体積ゆらぎ

# まとめ

保存量初期ゆらぎの痕跡は終状態に残されており  
LHC、RHICで観測にかかり始めている

ただし、RHIC-BES領域では即座にQCD相構造が  
直ちに見えるほど明瞭ではないので、

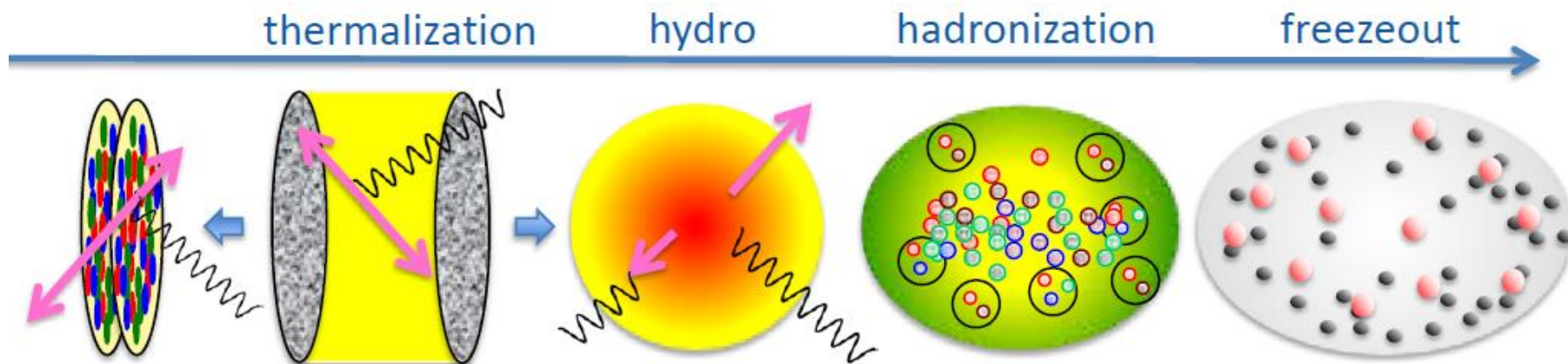
## 実験的課題

- バリオン数ゆらぎの構成
- $Q_4/Q_2$ @ALICE
- efficiency, acceptance, ...
- エネルギーゆらぎ
- BES from RHIC to LHC

## 理論的課題

- RHIC D-puzzleの解決
- RHIC-BESの結果の解釈
- ゆらぎの動的時間発展
- 2点相関関数等との比較
- 新しい観測量の提案

# Evolution of Fluctuations



Fluctuation  
in initial state



Time evolution  
in the QGP



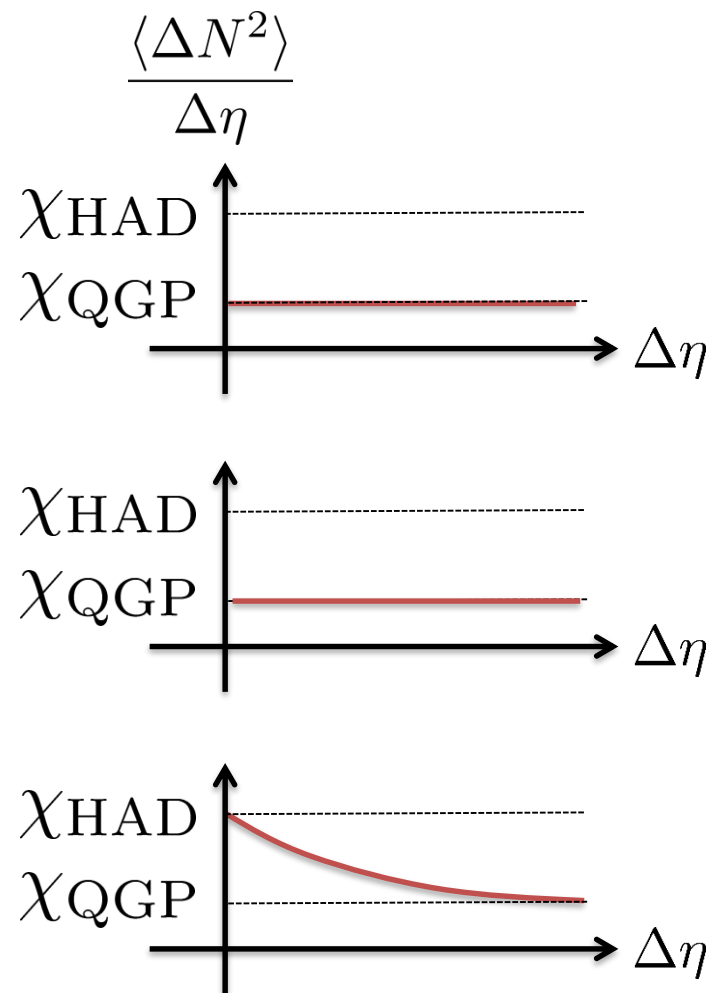
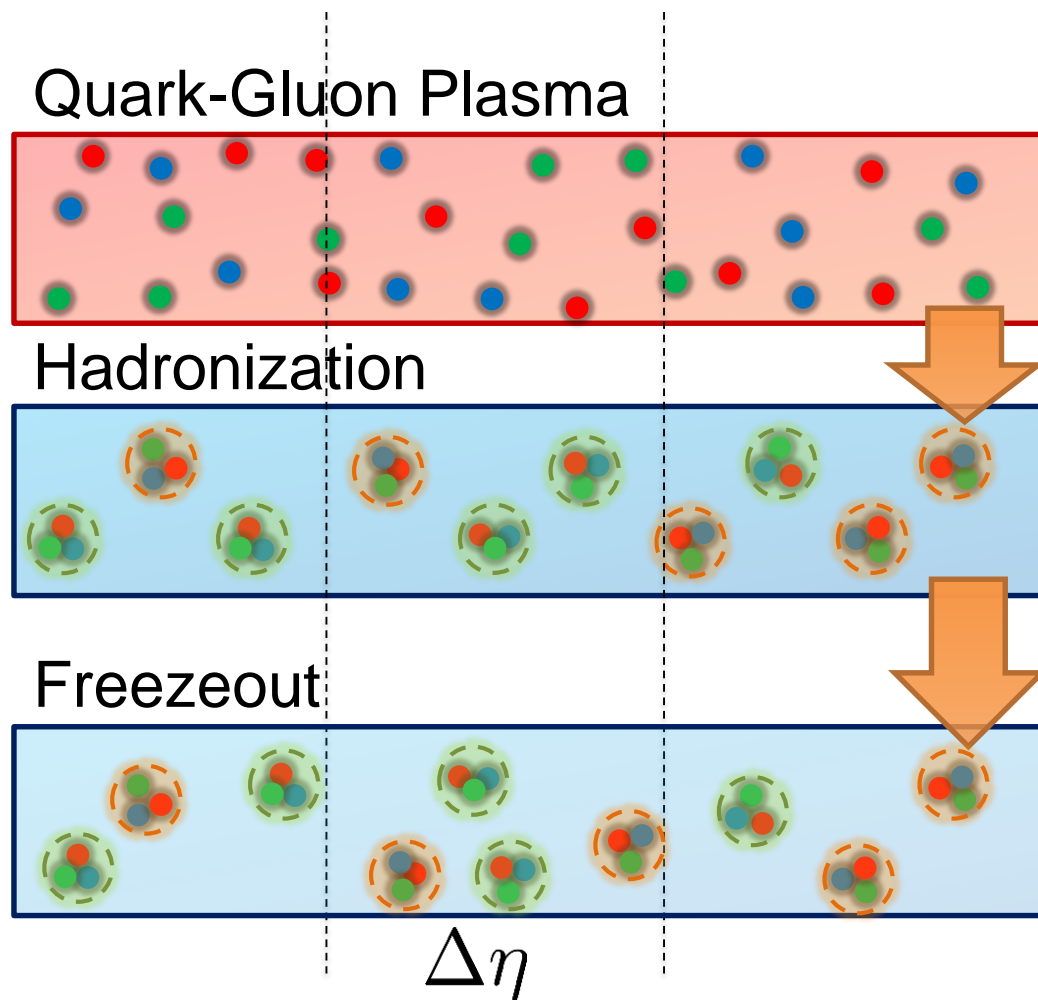
approach to HRG  
by diffusion

volume fluctuation

experimental effects  
particle missID, etc.

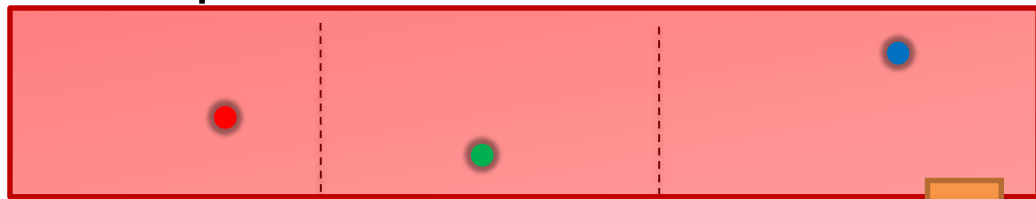


# Time Evolution in HIC

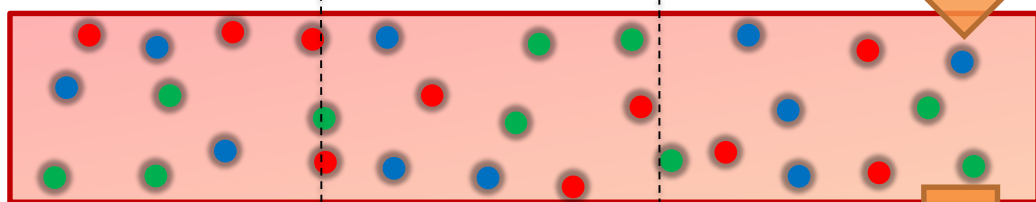


# Time Evolution in HIC

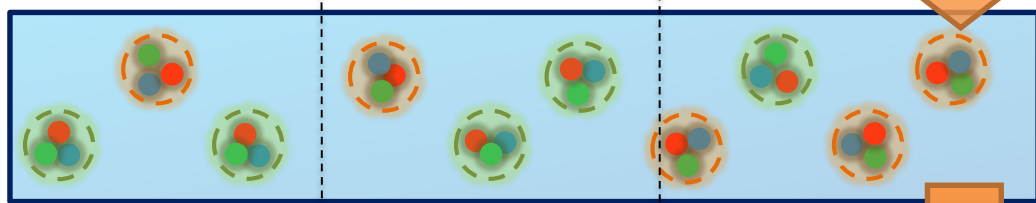
Pre-Equilibrium



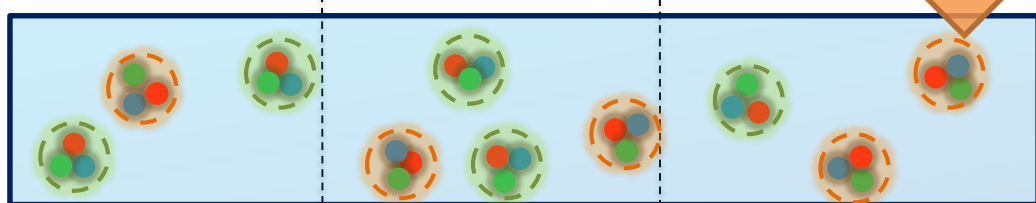
Quark-Gluon Plasma



Hadronization

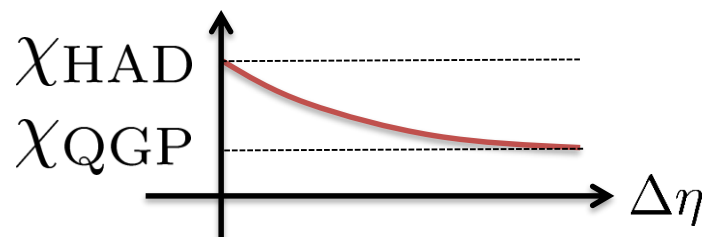
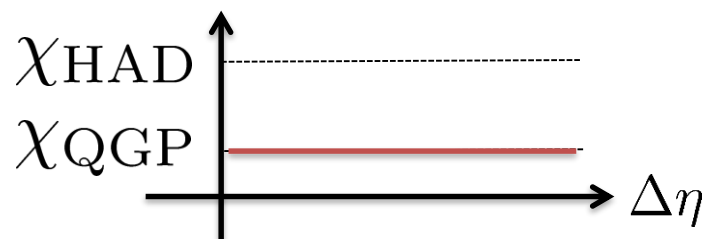
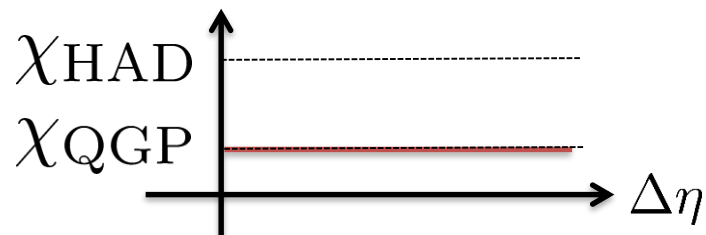
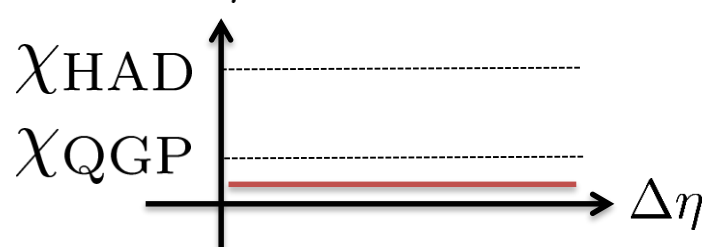


Freezeout



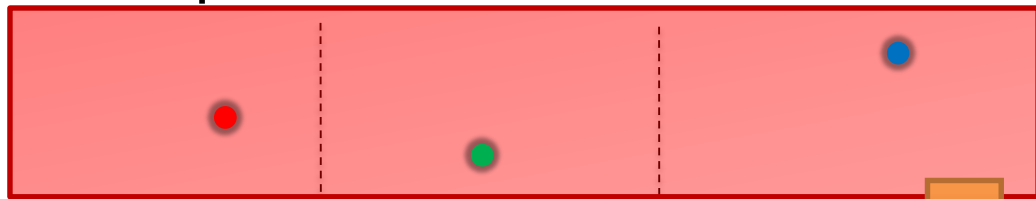
$\Delta\eta$

$$\frac{\langle \Delta N^2 \rangle}{\Delta\eta}$$

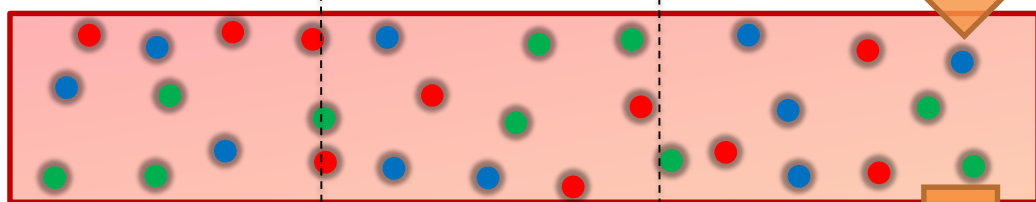


# Time Evolution in HIC

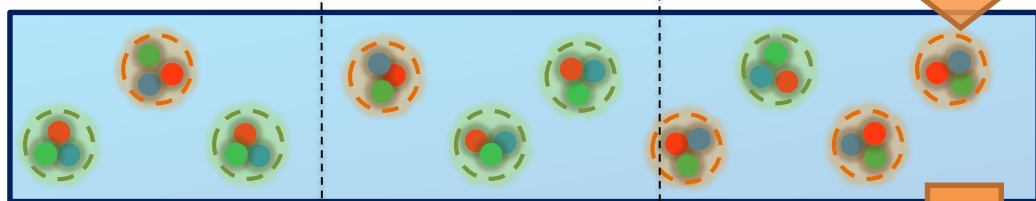
Pre-Equilibrium



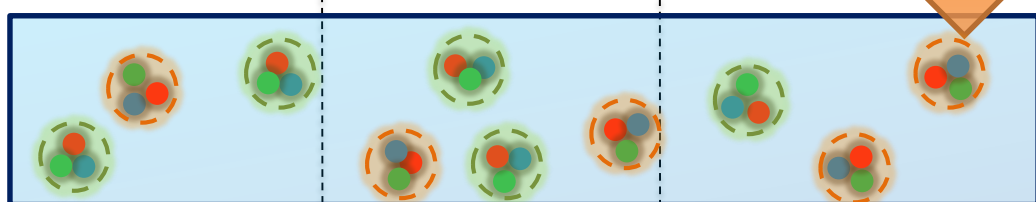
Quark-Gluon Plasma



Hadronization



Freezeout



$\Delta\eta$

$$\frac{\langle \Delta N^2 \rangle}{\Delta\eta}$$

