# Understanding Fluctuations using rapidity window and collision energy dependences

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MK, Asakawa, Ono, Phys. Lett. B728 (2014) 386-392 Sakaida, Asakawa, MK, arXiv:1409.6866 MK, to appear soon!

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#### Beam-Energy Scan





#### Cumulants up to 4<sup>th</sup> Order in 2014





# This is a great achievement in physics!

#### Fluctuations and Elemental Charge

Asakawa, Heinz, Muller, 2000 Jeon, Koch, 2000 Ejiri, Karsch, Redlich, 2005



$$\langle \delta N_q^n \rangle_c = \langle N_q \rangle$$
$$\Longrightarrow \langle \delta N_B^n \rangle_c = \frac{1}{3^{n-1}} \langle N_B \rangle$$

$$3N_B = N_q$$



$$\langle \delta N_B^n \rangle_c = \langle N_B \rangle$$

Free Boltzmann  $\rightarrow$  Poisson  $\langle \delta N^n \rangle_c = \langle N \rangle$ 

### Shot Noise



#### Total charge in the interval

$$Q = e\langle N \rangle$$
  

$$\delta Q^{2} \rangle = e^{2} \langle \delta N^{2} \rangle = e^{2} \langle N \rangle = eQ$$
  

$$\underbrace{\langle \delta Q^{2} \rangle}{Q} = e$$
  

$$\checkmark \text{ Quantum Hall effect}$$
  
Saminadayar+, PRL**79**,2526(1997)

✓ Superconductor

X. Jehl+, Nature405,50 (2000)

#### Shot Noise



#### Total charge in the interval

$$= e\langle N \rangle$$

$$Q^{2} \rangle = e^{2} \langle \delta N^{2} \rangle = e^{2} \langle N \rangle = eQ$$

$$\frac{\langle \delta Q^{2} \rangle}{Q} = e$$

$$\frac{\langle Quantum Hall effect}{Saminadayar+, PRL79,2526(1997)}$$

Superconductor

X. Jehl+, Nature405,50 (2000)

Higher order cumulants:

$$\langle \delta Q^3 \rangle = e^3 \langle N \rangle = e^2 Q$$

3rd order ex. Beenakker+, PRL90,176802(2003) up to 5th order

Gustavsson+, Surf.Sci.Rep.64,191(2009)

#### Cumulants up to 4<sup>th</sup> Order in 2014



# This is a great achievement in physics!

However, still a lot of things to do...



## Many Things to Do

□ Message to Experimentalists:

Measure rapidity window dependences
 Determine baryon number cumulants

□ Message to Theorists:

Do not directly compare your thermal results with exp.
 Let's pursue descriptions of non-eq. non-Gaussianity.

#### □ Message to Latticians:

Do not directly compare your results with exp.Measure more cumulants more accurately



#### Experimental results on Non-equilibrium state Non-Gaussian fluctuations



#### Experimental results on Non-equilibrium state Non-Gaussian fluctuations



# Rapidity Window Dependences of Higher Order Cumulants

MK, Asakawa, Ono, Phys. Lett. B728 (2014) 386-392; MK, to appear soon!

#### Charge Fluctuation @ LHC



 $\langle \delta N_Q^2 \rangle$  is not equilibrated at freeze-out at LHC energy!

#### $\Delta\eta$ Dependence @ ALICE





Fluctuations continue to change until kinetic freezeout!!



achieved only through diffusion.

the slower diffusion

## **Conversion of Rapidities**



#### $\Delta\eta$ Dependence @ ALICE



 $\Delta \eta$  dependences of fluctuation observables encode history of the hot medium!

 $<\delta N_{\rm B}^2$  > and  $<\delta N_{\rm p}^2$  > @ LHC ?

 $\langle \delta N_Q^2 \rangle, \langle \delta N_B^2 \rangle, \langle \delta N_p^2 \rangle$ 

should have different  $\Delta \eta$  dependence.



 $<\delta N_{\rm B}^2$  > and  $<\delta N_{\rm p}^2$  > @ LHC ?

 $\langle \delta N_Q^2 \rangle, \langle \delta N_B^2 \rangle, \langle \delta N_p^2 \rangle$ 

should have different  $\Delta\eta$  dependence.



Baryon # cumulants are experimentally observable! MK, Asakawa, 2011;2012

 $<\delta N_{0}^{4} > @ LHC ?$ 



#### Hydrodynamic Fluctuations

Landau, Lifshitz, Statistical Mechaniqs II Kapusta, Muller, Stephanov, 2012

#### Stochastic diffusion equation



#### How to Introduce Non-Gaussianity?

Stochastic diffusion equation

$$\partial_{\tau} n = D \partial_{\eta}^2 n + \partial_{\eta} \xi(\eta, \tau)$$

Choices to introduce non-Gaussianity in equil.:

- $\square$  *n* dependence of diffusion constant *D*(*n*)
- colored noise
- □ discretization of *n*

#### How to Introduce Non-Gaussianity?

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Choices to introduce non-Gaussianity in equil.:

*n* dependence of diffusion constant *D*(*n*)
 colored noise
 discretization of *n* our choice

**REMARK**:

Fluctuations measured in HIC are almost Poissonian.

#### A Brownian Particle's Model

#### Hadronization (specific initial condition)



(1) Describe time evolution of Brownian particles exactly (2) Obtain cumulants of particle # in  $\Delta\eta$ 

#### **Baryons in Hadronic Phase**



#### **Diffusion Master Equation**

MK, Asakawa, Ono, 2014



### **Diffusion Master Equation**

MK, Asakawa, Ono, 2014



#### Solve the DME **exactly**, and take $a \rightarrow 0$ limit

No approx., ex. van Kampen's system size expansion

#### Net Charge Number

Prepare 2 species of (non-interacting) particles



Time evolution of Q up to Gauusianity is consistent with the stochastic diffusion equation

# **Time Evolution in Hadronic Phase**

#### Hadronization (initial condition)



Boost invariance / infinitely long system
 Local equilibration / local correlation



# **Time Evolution in Hadronic Phase**

#### Hadronization (initial condition)







#### Freezeout

**Fime evolution via DME** 





### **Total Charge Number**

In recombination model,



 $\square$   $N_B^{(\text{tot})}$  can fluctuate, while  $N_B^{(\text{net})}$  does not.

#### $\Delta\eta$ Dependence: 4<sup>th</sup> order



Charcteristic  $\Delta \eta$  dependences!





Relaxation from a specific initial condition to Skellam value

- We assume perfect equilibration at T<sub>c</sub>
- Cumulants near T<sub>c</sub> can be studied by Lattice.

□ No pair creation/annihilation below T<sub>chem</sub>

- For baryons, this would be justified below T<sub>c</sub>
- But, may not for electric charges.

□ No finite volume effect

# Effect of Global Charge Conservation (Finite Volume Effect)

Sakaida, Asakawa, MK, arXiv:1409.6866

#### **Global Charge Conservation**

Conserved charges in the total system do no fluctuate!



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Conserved charges in the total system do no fluctuate!



Jeon, Koch, PRL2000; Bleicher, Jeon, Koch (2000)

#### **Diffusion in Finite Volume**

Solve the diffusion master equation in finite volume



## **Diffusion in Finite Volume**

Solve the diffusion master equation in finite volume



#### **Physical Interpretation**



 $d(\tau)$  : Averaged Diffusion Distance  $D(\tau)$  : Diffusion Coefficient  $\eta_{tot}$  : Total Length of Matter



Effects of the GCC appear only near the boundaries.

## **Comparison with ALICE Result**



 $d(\tau)$ 

 $\eta_{\rm tot}$ 

T





- No GCC effect in ALICE experiments!
- Same conclusion for higher order cumulants

### Very Low Energy Collisions

Large contribution of global charge conservationViolation of Bjorken scaling



Careful treatment is required to interpret fluctuations at low beam energies! Many information should be encoded in  $\Delta\eta$  dep.

#### Summary

Measurement of non-Gaussianity at STAR
 Non-equilibrium behavior at ALICE

Theoretical description to treat both "non"s is needed!

Plenty of information in  $\Delta\eta$  dependences of various cumulants  $\langle N_Q^2 \rangle_c, \ \langle N_Q^3 \rangle_c, \ \langle N_Q^4 \rangle_c, \ \langle N_B^2 \rangle_c, \ \langle N_B^3 \rangle_c, \ \langle N_B^4 \rangle_c$ 

and those of non-conserved charges, mixed cumulants...

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## $\Delta\eta$ Dependence

Shuryak, Stephanov, 2001

□ Initial condition:  $\langle \delta n(\eta_1, 0) \delta n(\eta_2, 0) \rangle = \sigma_2 \delta(\eta_1 - \eta_2)$ 

Translational invariance



#### **Open Questions & Future Work**

- Why the primordial fluctuations are observed only at LHC, and not RHIC ?
- Extract more information on each stage of fireballs using fluctuations

- Model refinement
  - Including the effects of nonzero correlation length / relaxation time global charge conservation