

Non-Gaussian fluctuations in relativistic heavy ion collisions  
Masakiyo Kitazawa

一本の草も涼風宿りけり

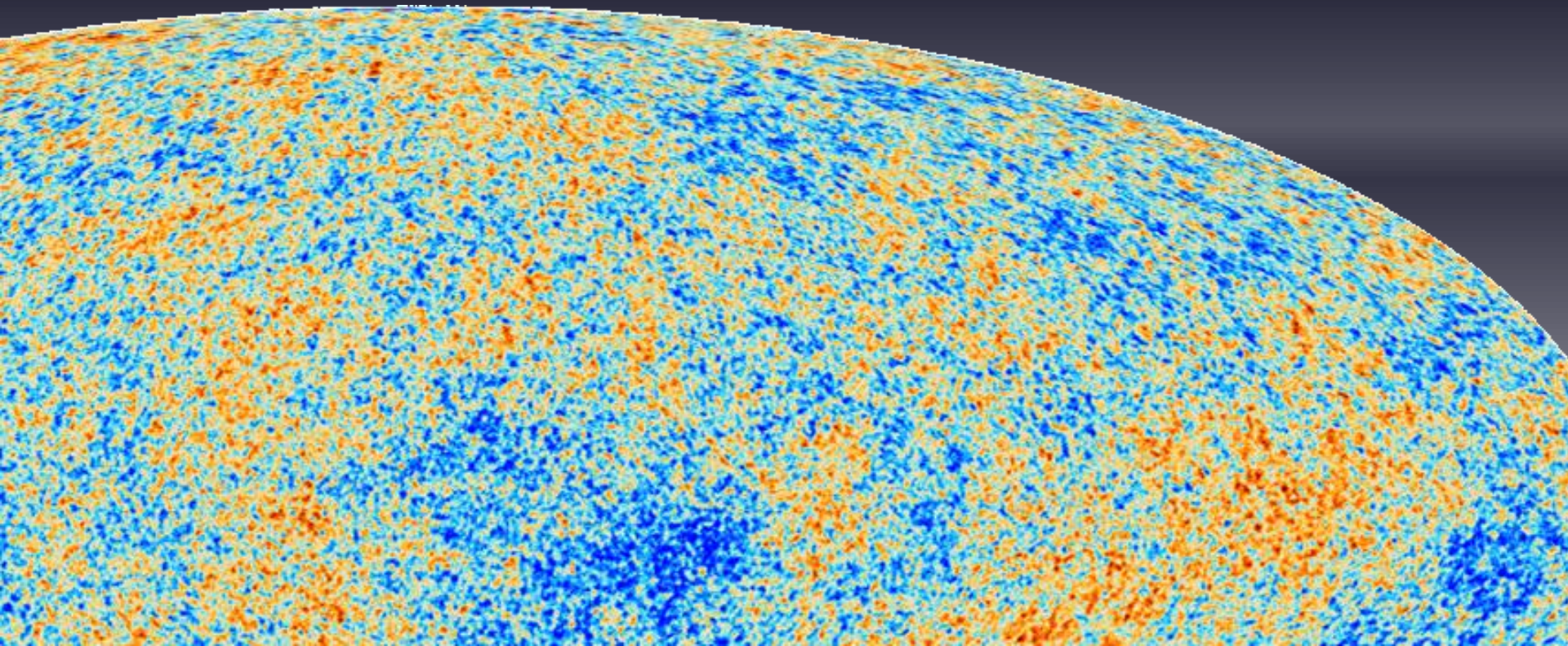
even on one blade of grass the cool wind lives

小林一茶

Issa Kobayashi

1814

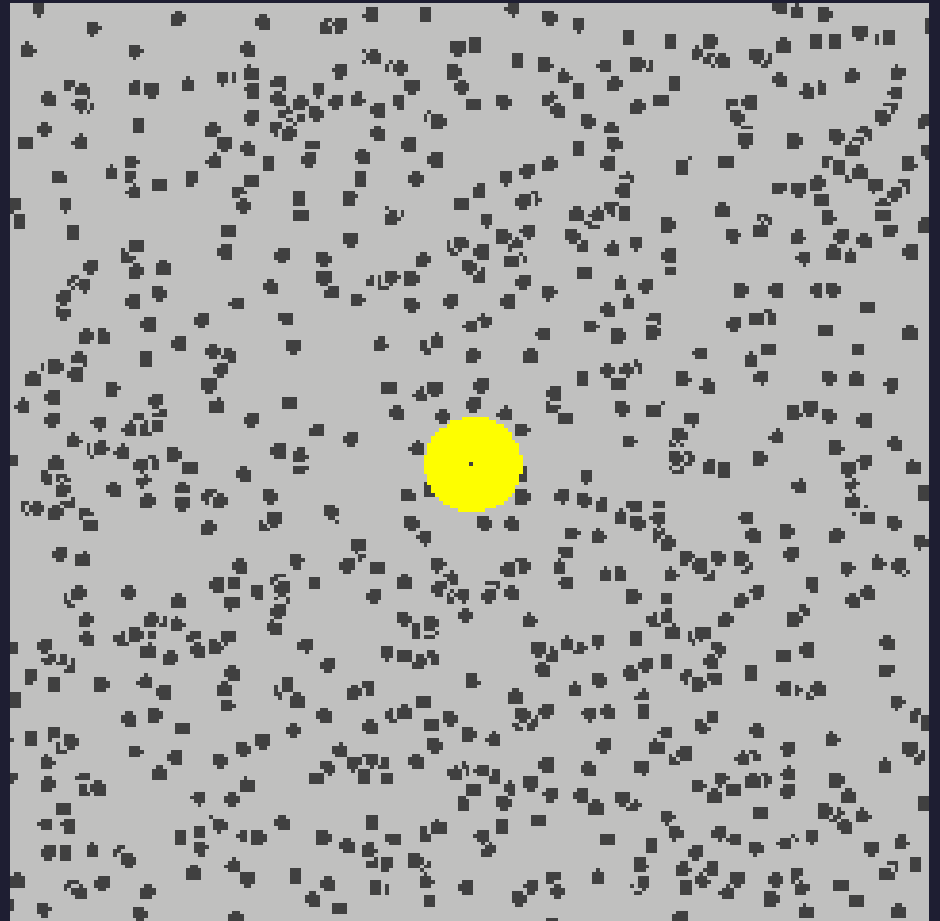
Physicists can feel **hot** early Universe  
13 800 000 000 years ago  
in tiny fluctuations of  
cosmic microwave



Physicists can feel the existence of **microscopic** atoms behind random fluctuations of Brownian pollens



A. Einstein  
1905



quarks

Feel quarks behind fluctuations  
in relativistic heavy ion collisions

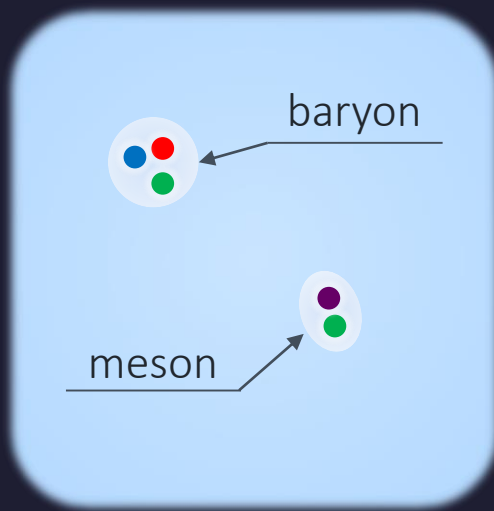
2014

# Non-Gaussian Fluctuations in Relativistic Heavy Ion Collisions

Masakiyo Kitazawa (Osaka U.)

# Quark-Gluon Plasma

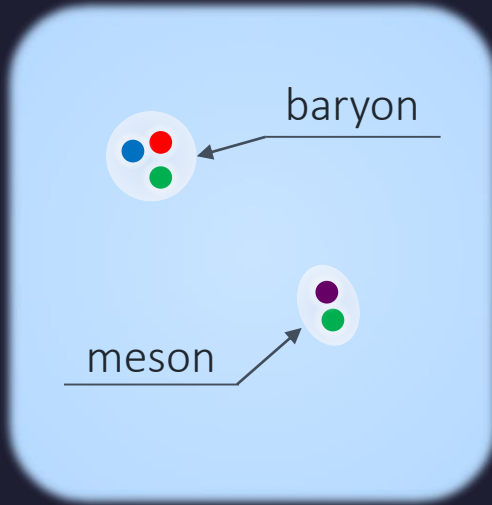
vacuum





# Quark-Gluon Plasma

vacuum



As  $T$  increases ...

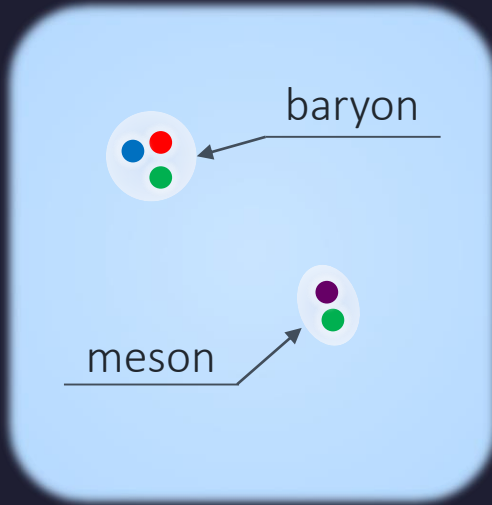


Early Universe



# Quark-Gluon Plasma (QGP)

vacuum



As  $T$  increases ...

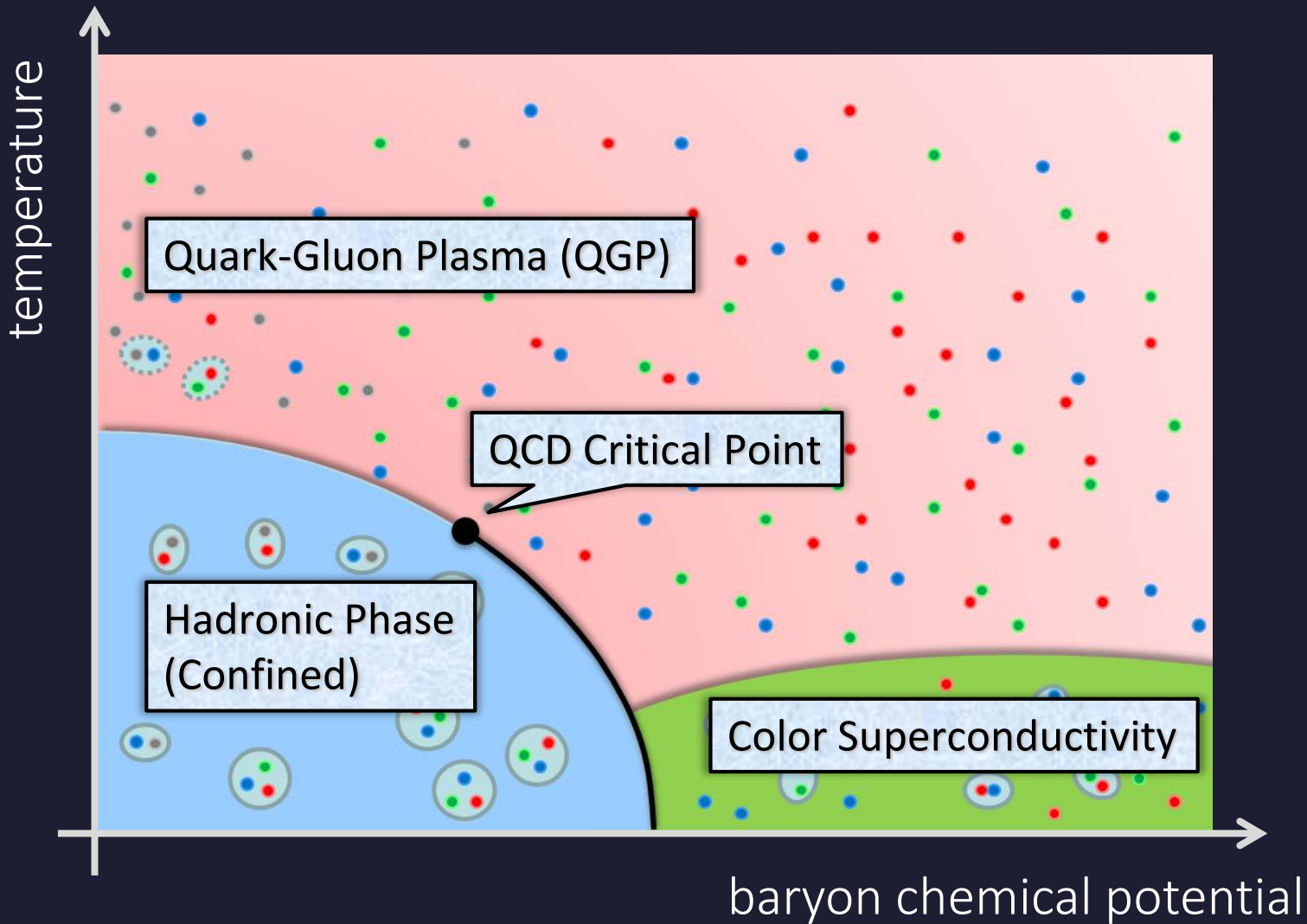


quark-gluon plasma

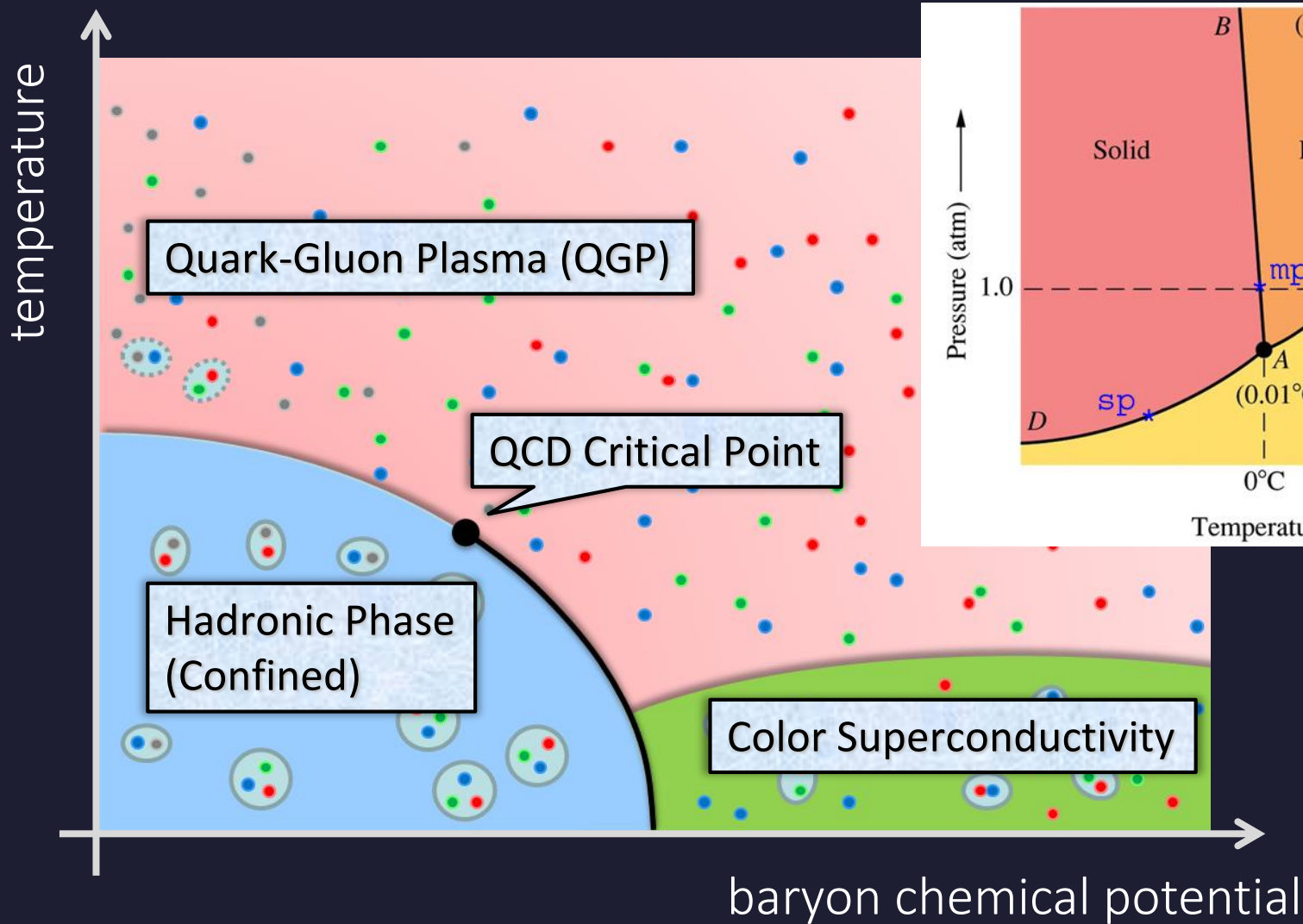
Early Universe



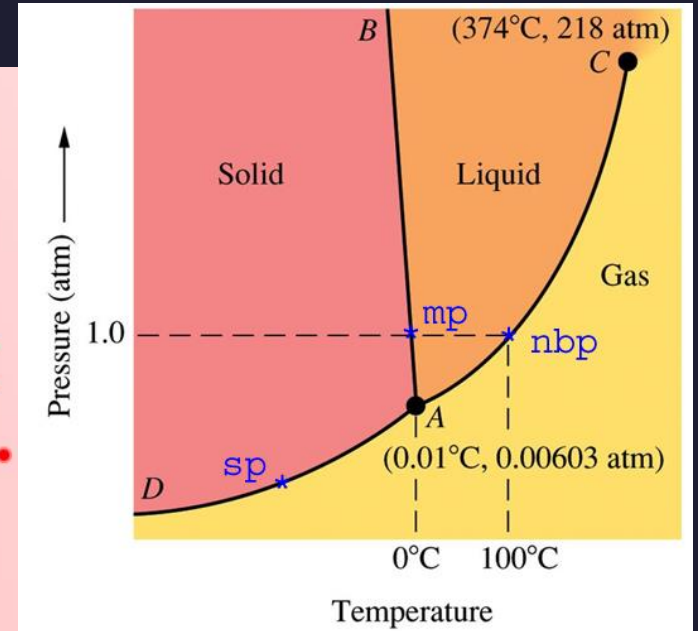
# QCD Phase Diagram



# QCD Phase Diagram



Phase diagram of water



# Relativistic Heavy Ion Collisions



LHC – Large Hadron Collider

# Relativistic Heavy Ion Collisions

For the search of  
new particles



An aerial photograph of a valley with a large red oval overlaid on it, representing the circular path of the Large Hadron Collider. The oval is marked with several small red circles at intervals. In the background, there are mountains and a city.

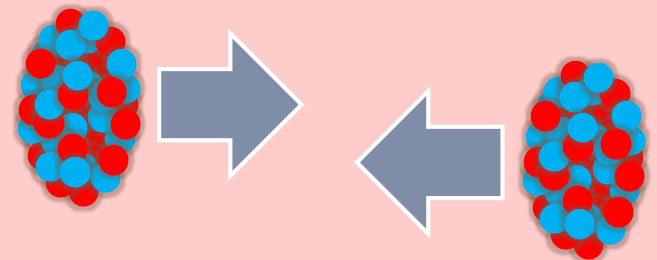
LHC – Large Hadron Collider

# Relativistic Heavy Ion Collisions

For the search of  
new particles



To create the early  
Universe



LHC – Large Hadron Collider

①

Quark-Gluon Plasma

temperature :  $T \sim 4 \times 10^{12} \text{K}$

lifetime :  $t \sim 10^{-22}$  seconds.

②

The medium then cools down with an expansion.

③

Confined particles  
arrive at the detector.



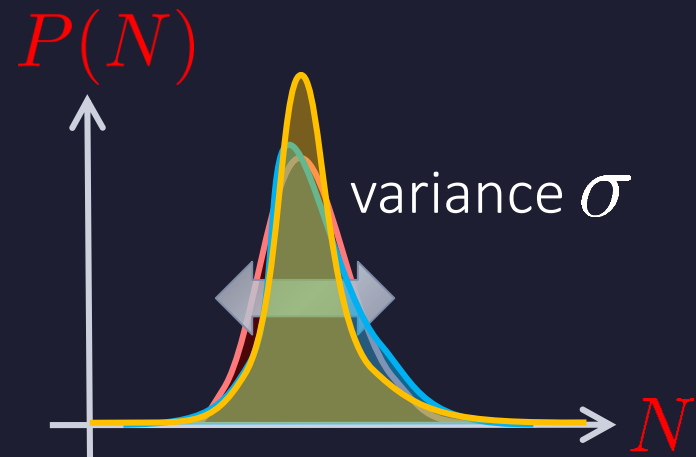
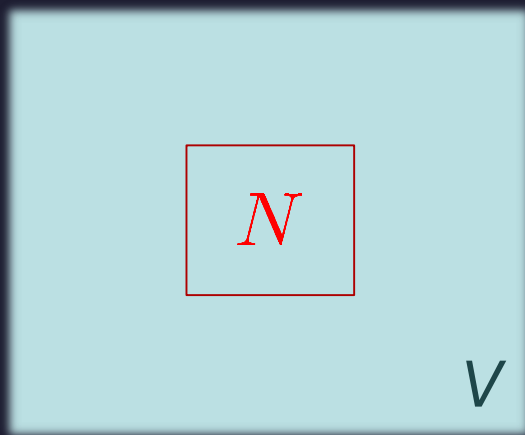
2.76 ATeV

D3BBE693



# Thermal Fluctuations

Observables in equilibrium are fluctuating!



$$\langle \delta N^2 \rangle = V \chi_2 = \sigma^2$$

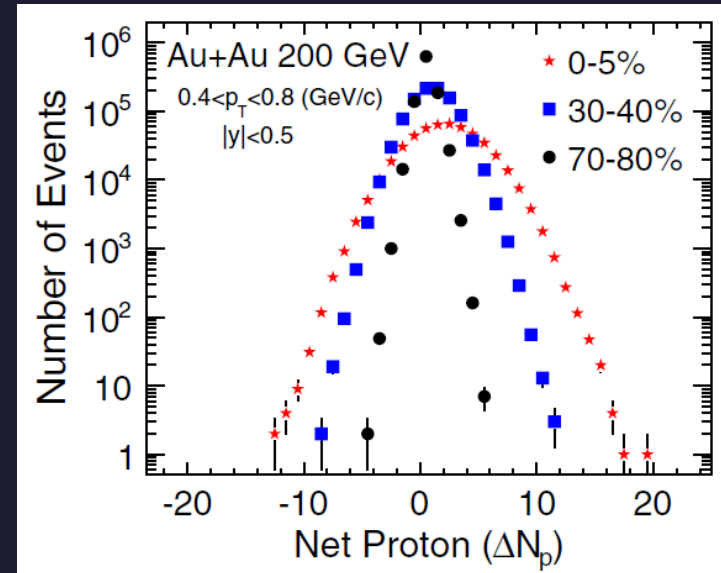
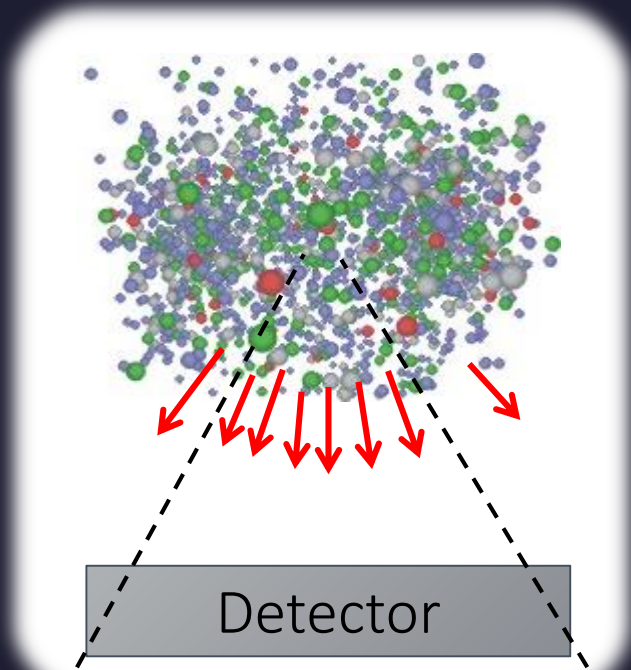
$$S = \frac{\langle \delta N^3 \rangle}{\sigma^3}$$

$$\kappa = \frac{\langle \delta N^4 \rangle - 3\langle \delta N^2 \rangle^2}{\chi_2 \sigma^2}$$

Gaussian

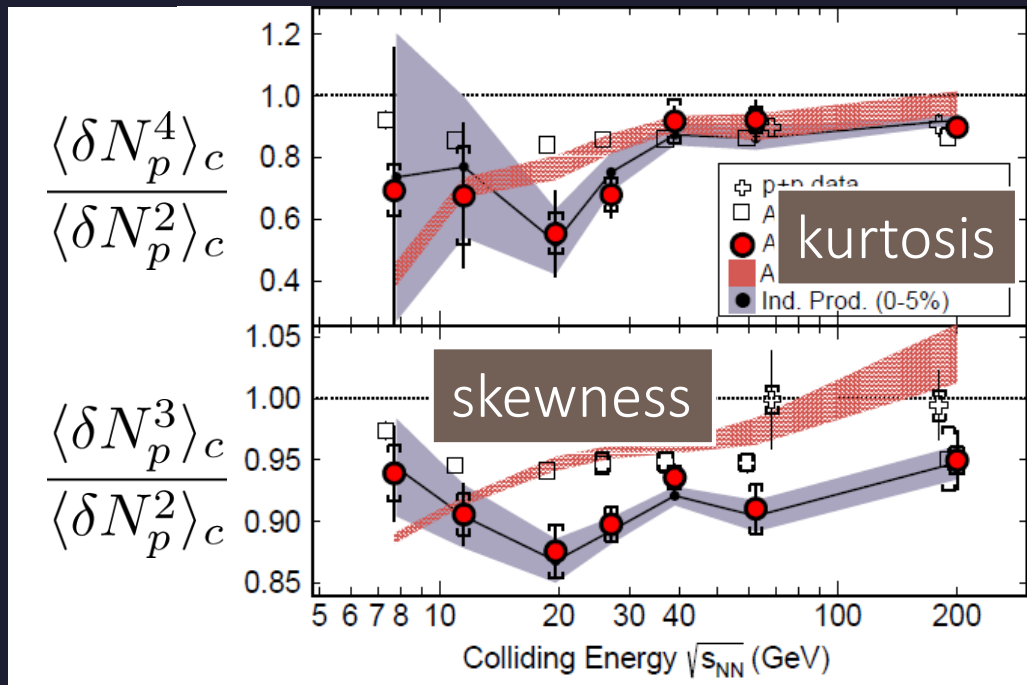
non-Gaussianity

# Event-by-Event Measurement

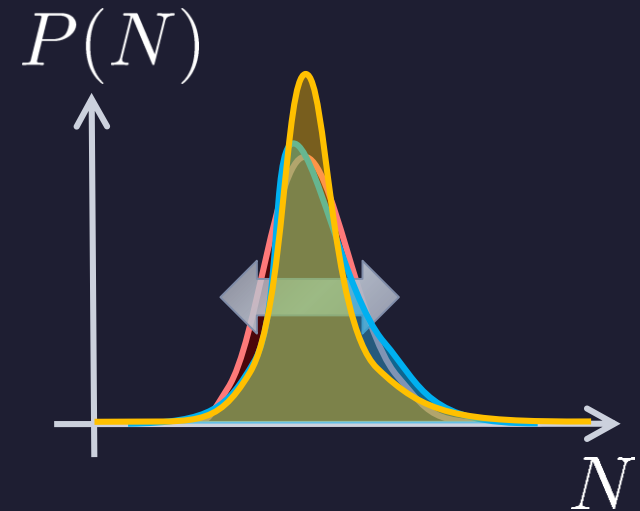


STAR Collaboration, PRL 2010

# Non-Gaussianity @ RHIC

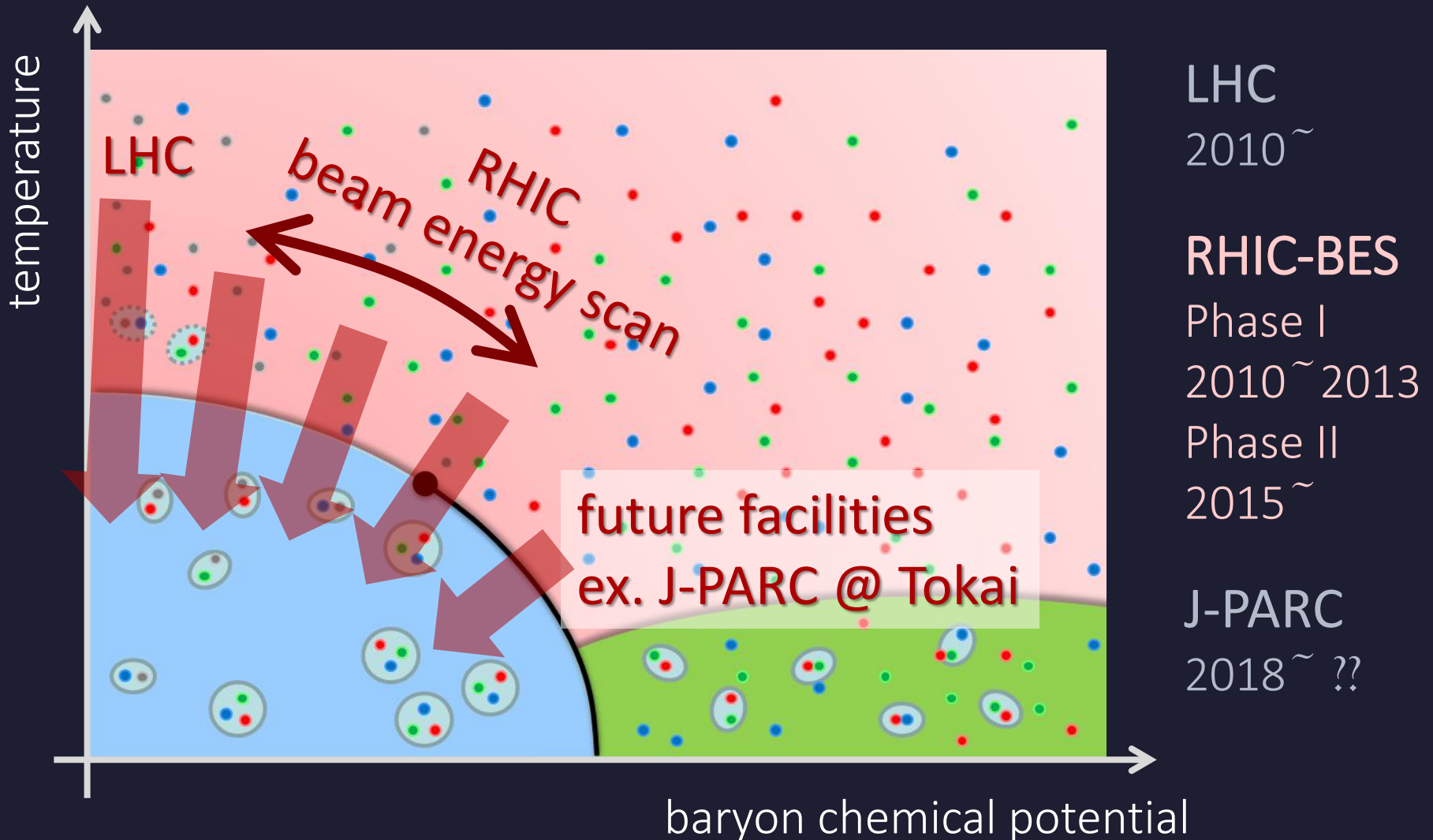


STAR Collaboration, PRL 2014



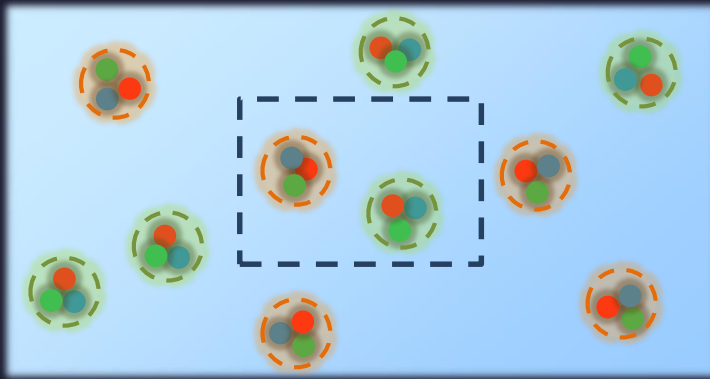
- ❑ Nonzero higher-order cumulants of conserved charges (skewness and kurtosis)
- ❑ They are not far from Poissonian values.

# Search for QCD Phase Structure



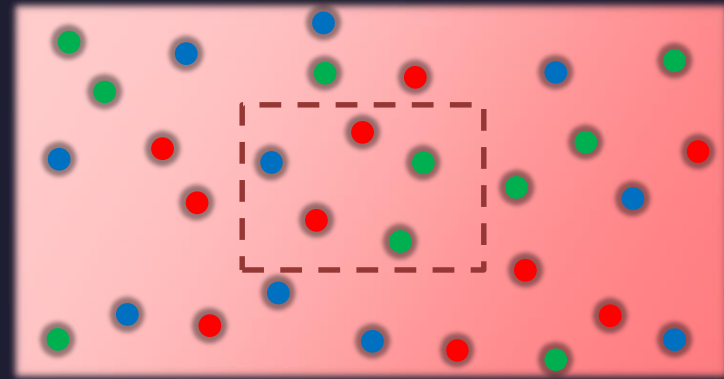
# Signal of Quark Deconfinement

Hadronic



$$|q_B| = 0, 1, \quad |q_Q| = 0, 1$$

Quark-Gluon

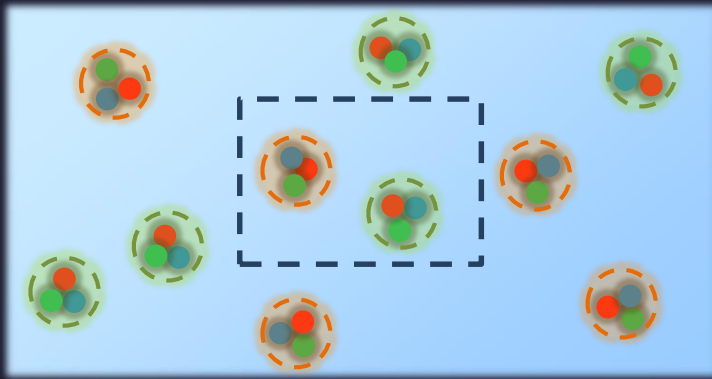


$$|q_B| = 1/3, \quad |q_Q| = 1/3, 2/3$$

Elemental charge carried by quasi-particles decreases in QGP

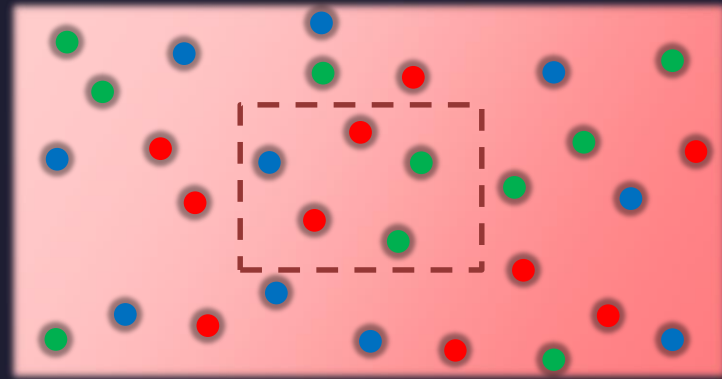
# Signal of Quark Deconfinement

Hadronic



$$|q_B| = 0, 1, \quad |q_Q| = 0, 1$$

Quark-Gluon



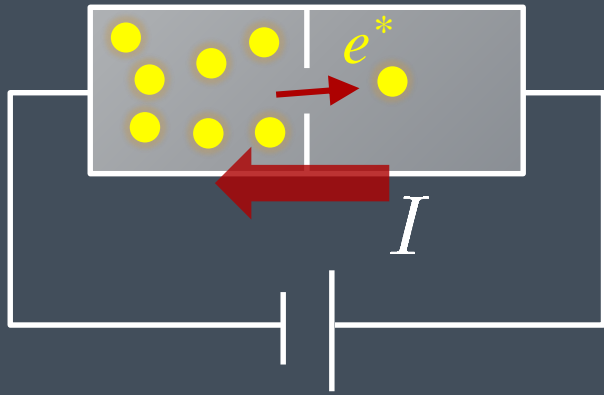
$$|q_B| = 1/3, \quad |q_Q| = 1/3, 2/3$$

Elemental charge carried by quasi-particles decreases in QGP



Corresponding thermal fluctuations decrease in QGP

# Shot Noise

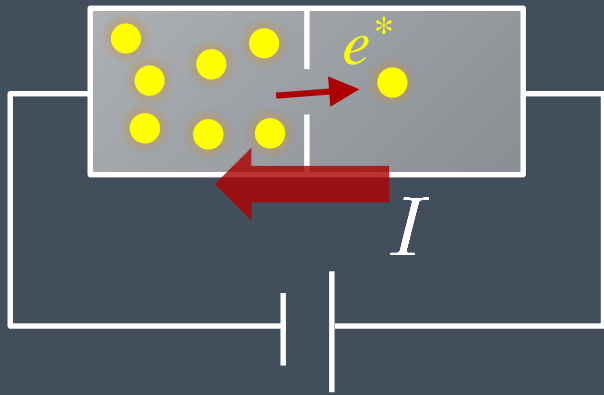


$$S_{\text{shot}} \sim \langle \delta I^2 \rangle$$

$$S_{\text{shot}} = 2e^* \langle I \rangle$$

↑  
charge of quasi-particles

# Shot Noise



$$S_{\text{shot}} \sim \langle \delta I^2 \rangle$$

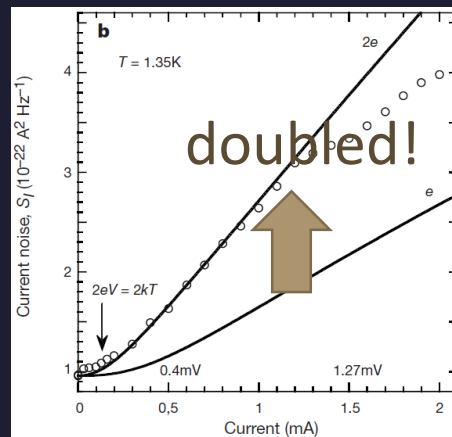
$$S_{\text{shot}} = 2e^* \langle I \rangle$$

charge of quasi-particles

Superconductors  
with Cooper Pairs

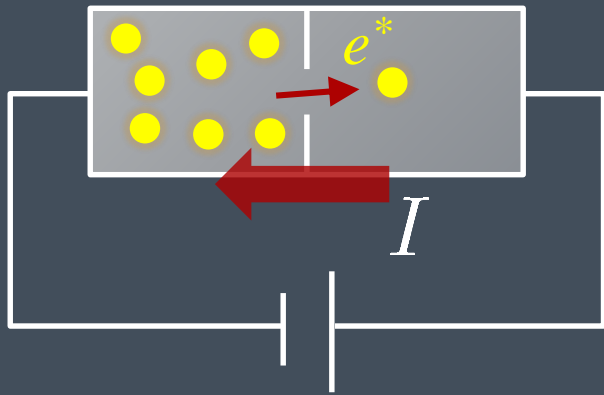
$$e^* = 2e$$

Jehl+, Nature 405,50 (2000)





# Shot Noise



$$S_{\text{shot}} \sim \langle \delta I^2 \rangle$$

$$S_{\text{shot}} = 2e^* \langle I \rangle$$

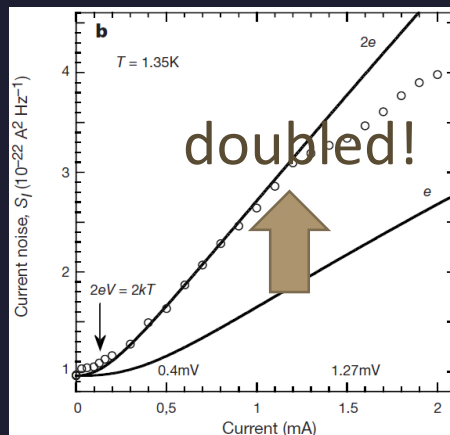


charge of quasi-particles

Superconductors  
with Cooper Pairs

$$e^* = 2e$$

Jehl+, Nature **405**,50 (2000)

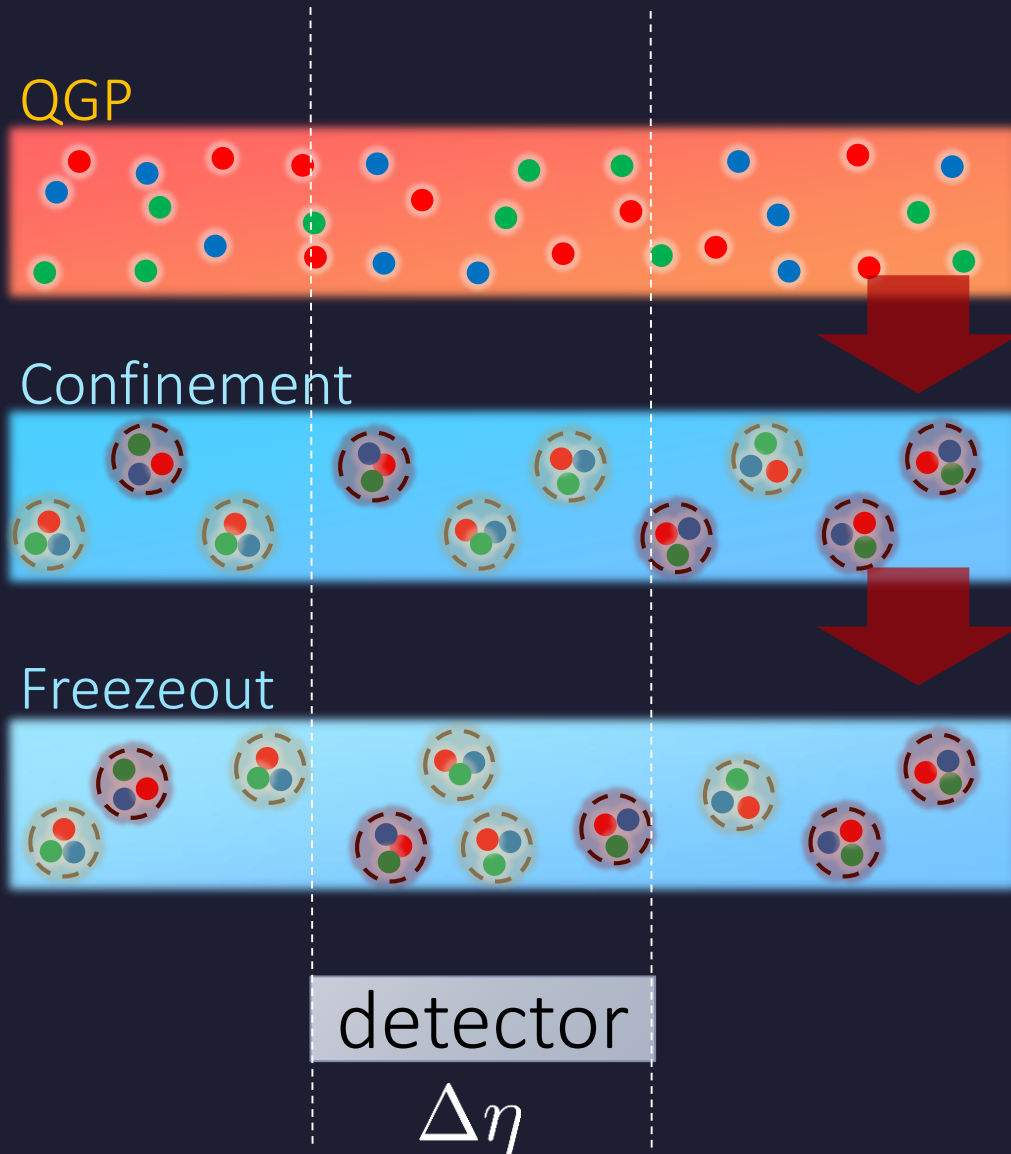


Fractional Quantum  
Hall Systems

$$e^* = \frac{q}{p} e$$

Saminadayar+, PRL**79**,2526 (1997)

# Diffusion of Fluctuations



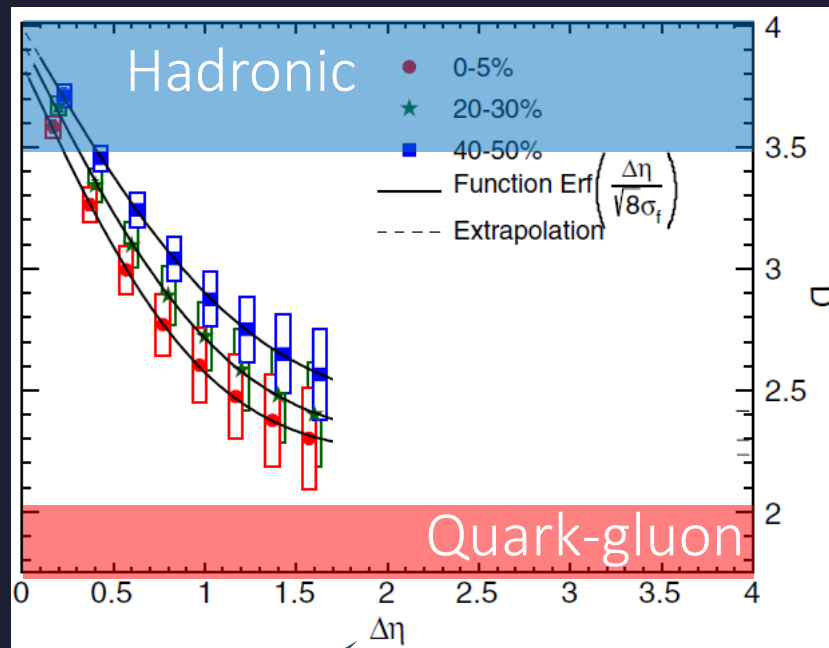
Experiments can vary spatial volume to measure fluctuations



The larger  $\Delta\eta$ , the earlier fluctuations

# Electric Charge Fluctuations @ LHC

ALICE Collaboration,  
PRL **110**, 152301 2013



$$\sim \frac{\langle \delta N_Q^2 \rangle}{V}$$

$$\sim V$$

宿りけり 一本の草も涼風  
一茶

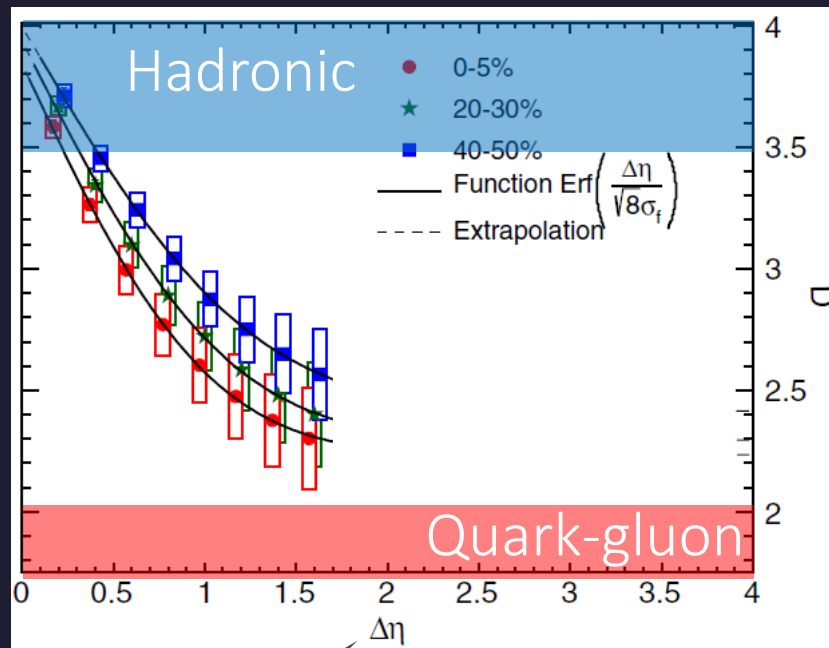
Fluctuation is more QGP-like as  $V$  becomes larger.  
The  $\Delta\eta$  dependence encodes history of the medium!

# Diffusion of non-Gaussianity

MK,Asakawa,Ono,PLB728,386(2014)

# Electric Charge Fluctuations @ LHC

ALICE Collaboration,  
PRL **110**, 152301 2013



$$\sim \frac{\langle \delta N_Q^2 \rangle}{V}$$

$$\sim V$$

- Experimental results only for 2<sup>nd</sup> order fluctuation
- No results on  $\Delta\eta$  dependence of higher-order cumulants

# Stochastic Formalism

- Fluctuating hydrodynamics  
(stochastic hydrodynamics)

Landau, Lifshitz,  
Statistical Mechanics II



- Counterpart for diffusive processes

Stochastic diffusion equation

$$\partial_{\tau} n = D \partial_x^2 n + \partial_x \xi(\eta, \tau)$$

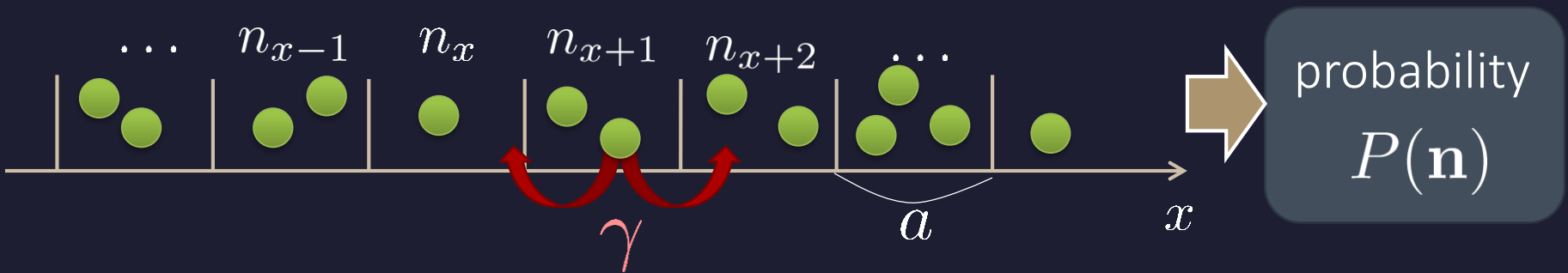
Random force determined by FDR

This formalism cannot describe non-Gaussianity!

# Diffusion Master Equation

MK,Asakawa,Ono,PLB728,386(2014)

Divide spatial coordinate into discrete cells

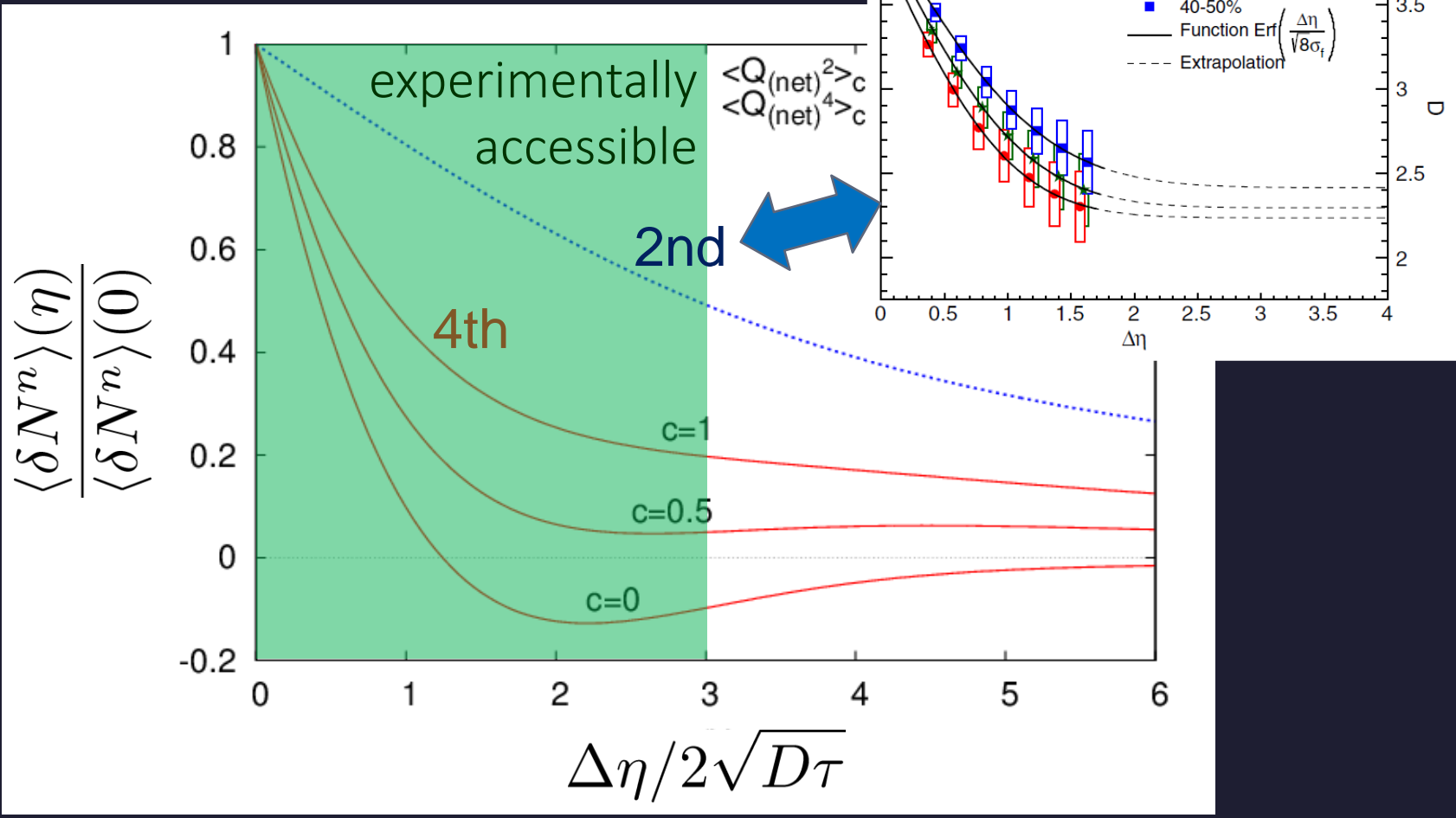


## Master Equation

$$\frac{\partial}{\partial t} P(\mathbf{n}) = \gamma \sum_x [(n_x + 1) \{P(\mathbf{n} + \mathbf{e}_x - \mathbf{e}_{x+1}) + P(\mathbf{n} + \mathbf{e}_x - \mathbf{e}_{x-1})\} - 2n_x P(\mathbf{n})]$$

Solve the DME **exactly**, and take  $a \rightarrow 0$  limit

# Our Predictions



Volume dep. of non-Gaussianity encodes more information!



# Summary

- ❑ Fluctuations are invaluable tools in physics, as well as in our daily life.
- ❑ Fluctuations acquires much attention in relativistic heavy-ion collisions. In particular, their non-Gaussianity is one of the latest topics in this realm.

1998

Rolf Landauer

The noise is the signal

A poet said

一本の草も涼風宿りけり

even on one blade of grass the cool wind lives

小林一茶

Issa Kobayashi

1814

A physicist said

1998

Rolf Landauer

The noise is the signal