### For the J-PARC Heavy-Ion Collisions

Masakiyo Kitazawa

## Remarks!

- No solid results in this talk. Just delusional ideas.
- I am waiting for your opinions / discussions (throughout this meeting)!
- I welcome your criticisms.

What Physics are Suitable for J-PARC HIC?

#### Characteristics of J-PARC HIC

- > High density
  - QCD phase transitions
    - Neutron stars / NS-NS merger

> High luminosity

- Various event selections
  - Rare probes
  - Event-by-event analyses

Take advantages of these properties!

 $\sqrt{s_{_{\rm NN}}} = 1 \sim 6 {
m ~GeV}$ 

What Physics are Suitable for J-PARC HIC?

#### Characteristics of J-PARC HIC





## High Density



Special thanks to C. Matsumura

Search of dense medium / QCD phase transitions

How Dense? : Freezeout

#### **Chemical Freezeout**



Randrup, Cleymans, 2006

The most dense system at CFO with  $\rho \sim 0.7 \rho_0$  is realized at

$$\int \sqrt{s_{_{\rm NN}}} \simeq 4 \,\,{
m GeV}$$
  
 $E_{
m lab.} \simeq 30 \,\,{
m GeV}$ 



## How Dense? : Highest



 $ho > 10
ho_0$  would be realized!

On the other hand, JAM can reproduce  $dv_1/dy$  and  $p_T$  spectra for J-PARC energy.

Nara-san's talk, yesterday

Why?

I.C. Arsene+, PRC 75, 034902 (2007)

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Large E-v-E Fluctuation?



Baryon stopping seems to have strong e-v-e fluctuation.

### Origin of e-v-e fluctuation 1: Stopping $\rightarrow$ Penetrate Transition



Structural Transition stopping → penetrate



### Origin of e-v-e fluctuation 2: **Softening** of EoS





### E-v-E Fluctuation of Density How to **Observe** Them?



We need observables sensitive to baryon stopping

 $p_L$ 

Ex. 
$$\langle p_L^2 
angle$$

#### E-v-E Fluctuation of Density: How to **Use** Them?



## BDS (Baryon Density Scan)

average transverse energy



#### BDS (Baryon Density Scan) average transverse energy Hadronic Phase $\langle dE_{T}/d\eta \rangle / \langle dN_{ch}/d\eta \rangle$ (GeV) PHENIX PHENIX scale error **ALICE** STAR **NA49** $E_T$ WA98 E802 + FOPI P Baryon-rich events 0.5 B $\langle B \rangle$ 0 10<sup>3</sup> 10<sup>2</sup> 10 1 $s_{NN}$

Softening of EoS by 1st transition

O Quark-Gluon O O Plasma

**Compact Stars** 

B

 $\mu_{B}$ 

#### \_\_\_\_\_

### Event-by-event Fluctuations: Compared with high energy



#### STAR Collaboration, 2015

**Thermal** fluctuations are concerned @ high  $\sqrt{s}$ .

Measurement of thermal fluc. justified only under **Bjorken picture** 

Asakawa, MK, arXiv:1512.05038

Pictorial view for low √s



detector

Other Ideas

 $\succ$  Event-by-event dv<sub>1</sub>/dy, raial flow, ...

➤ Correlation b/w baryon stopping and
 ➤ v<sub>1</sub>, strangeness, dE<sub>T</sub>/dy, ...

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Rare Probes

 $\sqrt{s_{_{\rm NN}}} = 1 \sim 6 {
m ~GeV}$ 

## Rare Probes

- Exotic hadrons
- Strangelet
- Strangeness
- Anti-protons



Stochastic Description for anti-p and s transports?



### And, of course, Other Many Important Topics



# Liquid-Gas Phase Transition of Nuclear Medium



Any possibility to perform a similar analysis in HIC...?

Summary

Characteristics of J-PARC HIC High density / High luminosity

We should be able to have more ideas to take these advantages of low-energy collisions! event selections / e-b-e analyses / rare probes

Important to accelerate J-PARC/FAIR/NICA programs!

 $\sqrt{s_{_{\rm NN}}}=1\sim 6~{\rm GeV}$