# Evolution of Critical Fluctuations / Non-binomial Efficiency correction

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- 1. Diffusion
- 2. Evolution of Critical Fluctuations: 2<sup>nd</sup> order
- 3. Evolution of Critical Fluctuations: 3<sup>rd</sup> order
- 4. Non-binomial Efficiency Correction
  - Previous methods
  - > New general method

General Review: Asakawa, MK, PPNP (2016)



Non-zero non-Gaussian cumulants have been established!

Have we measured critical fluctuations?



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MK, Asakawa, Ono, PLB 728, 386 (2014) Sakaida, Asakawa, MK, PRL90, 064911 (2014) MK, NPA942, 65 (2015)



### (Non-Interacting) Brownian Particle Model



## (Non-Interacting) Brownian Particle Model



### 4<sup>th</sup> Order Cumulant

#### MK+ (2014) MK (2015)



## 4<sup>th</sup> Order Cumulant

#### MK+ (2014) MK (2015)



**□** Cumulant at small  $\Delta \eta$  is modified toward a Poisson value. **□** Non-monotonic behavior can appear.

### **Time Evolution of Fluctuations**





Is non-monotonic Δη dependence already observed?
 Different initial conditions give rise to different characteristic Δη dependence. → Study initial condition

Finite volume effects: Sakaida+, PRC90 (2015)

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Sakaida, Asakawa, Fujii, MK, PRC95,064905(2017)

## Effect of Dynamical Evolution

Hac



Growth of the critical fluctuation is limited by the critical slowing down.

#### Decay by diffusion

Fluctuations developed at CP are modified by the time evolution in later stage before observation.

μ

## **Dynamical Evolution of Critical Fluctuations**

#### $\ensuremath{\square}$ Evolution of spatially uniform " $\sigma$ " mode



See also, Kapusta, Torres-Rincon (2012); Herold, Nahrgang, ... (2015)

### Soft Mode of QCD-CP = Conserved Mode

Fujii 2003; Fujii, Ohtani, 2004; Son, Stephanov, 2004



# Evolution of baryon number density **Stochastic Diffusion Equation**

$$\partial_t n = D(t) \partial_x^2 n + \partial_x \xi$$

$$\langle \xi(x_1, t_1)\xi(x_2, t_2) \rangle = \chi_2(t) \delta^{(2)}(1-2)$$

 $D(t), \ \chi_2(t)$  :parameters characterizing criticality

We study the 2<sup>nd</sup> order cumulant as well as correlation function.



## Parametrizing $D(\tau)$ and $\chi(\tau)$

#### Critical behavior

- 3D Ising (r,H)
- model H

Berdnikov, Rajagopal (2000) Stephanov (2011); Mukherjee+(2015)

#### □Temperature dep.











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Murata, MK, in preparation

## Analysis of 3<sup>rd</sup>-order Cumulant Murata, MK in preparation SDE: Higher order cumulants vanish in equi. Include a non-linear effect into SDE $\lambda_3 = \frac{\chi_3}{\chi_2^3}$ $\partial_t n = D(t)\partial_x^2 \frac{\delta\Omega[n]}{\delta n(x)} + \partial_x \xi$ $\Omega[n] = \int dx (\lambda_2 n(x)^2 + \lambda_3 n(x)^3)$ $\begin{aligned} \lambda_3 &= 0\\ \mathsf{SDE} \end{aligned}$ See, Nahrgang, QM2017 Analytic solution at the leading order in $\lambda_3$ for $\langle N^3 \rangle_c, \ \langle \delta n(x_1) n(x_2) n(x_3) \rangle$

#### Time Evolution: Near CP

# Murata, MK in preparation



#### Time Evolution: Near CP

# Murata, MK in preparation



### Time Evolution: Near CP

# Murata, MK in preparation



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Nonaka, Esumi, MK, to appear soon

## Efficiency / Efficiency Correction

Experimental detectors have miscounting & misidentification...





## **Efficiency Correction**

### Binomial model

Independence of efficiency loss for individual particles

 $\mathbf{P} \mathcal{R}(n;N)$  :Binomial distribution func.

Cumulants of n can be represented by those of N

Bialas, Peschanski (1986); MK, Asakawa (2012); Bzdak, Koch (2012)

#### □ Unfolding

- Construct true distribution func.
- Numerically demanding



### **Binomial Model**

An efficient algorithm for multi-variable system Nonaka, MK, Esumi, PRC2017









By truncating the Taylor exp. at *m*th order, "true" moments up to *m*th order are obtained.



## Comments

 $\square \ \langle n^m \rangle_R = \sum_n n^m \mathcal{R}(n;N)$  can be obtained from R(n;N)



- □ The truncation has to be well justified.
- □ Some distributions are automatically truncated.
  - Correct efficiency correction is possible.
  - binomial, hyper-geometric, beta-binomial, ..., binomial with fluctuating probability He, Luo, last Friday
- Compared to unfolding method,
  - numerically cheaper and would be more stable
  - origin of error is more apparent

Extension to multi-variable case is straightforward.

### Proton v.s. Baryon Number Cumulants

MK, Asakawa, 2012; 2012



□ The difference would be large.

**\square** Reconstruction of  $\langle N_B^n \rangle_c$  is possible using the binomial model.

- □ The use of binomial model is justified by "isospin randomization."
- And the loss due to momentum cut...

## Summary

□ Understanding dynamical aspects of fluctuations is important!
 □ Plenty information in ∆y dependence of cumulants:

- □ Higher order cumulants can behave non-monotonically.
- $\Box \rightarrow$  can be used for constraining parameters.
- Non-monotonicity in 2<sup>nd</sup> order cumulant is an experimental signal for the existence of the QCD-CP.
- A general algorithm for the efficiency correction:
  Correct reconstruction for non-binomial response.
  Smaller numerical cost than unfolding methods.
- Reonstructing baryon # cumulants is important!
- Let's continue the search for the QCD-CP!

### **Baryons in Hadronic Phase**



time



#### 4<sup>th</sup> order : w/ Critical Fluctuation



## **Translating Languages**

#### Brownian particle model



#### From Bzdak's talk



condition

$$\langle n^m \bar{n}^{\bar{m}} \rangle_{\rm fc} = \kappa_{m\bar{m}} \Delta y F_{m+\bar{m}} (\Delta y/d)$$

$$\begin{split} c_{m\bar{m}}^{0} &= \frac{1}{2} \frac{\partial^{2}}{\partial \Delta y^{2}} \langle n^{m} \bar{n}^{\bar{m}} \rangle_{\rm fc} \Big|_{\Delta y \to 0} = \kappa_{m\bar{m}} \frac{\partial}{\partial \Delta y} F_{m+\bar{m}} (\Delta y/d) \Big|_{\Delta y \to 0} \\ &= \frac{\kappa_{m\bar{m}}}{d} \frac{1}{\sqrt{(m+\bar{m})(2\pi)^{m+\bar{m}-1}}} \\ &\kappa_{m\bar{m}}: \ {\rm F \ cumulants \ at \ initial} \\ d: \ {\rm diffusion \ distance} \end{split}$$