

Gravitational Wave Astronomy with LIGO and Virgo

Chad Hanna - Penn State

Black Holes and Neutron Stars with Gravitational Waves - October 7, 2019

Charles E. Kaufman
Foundation

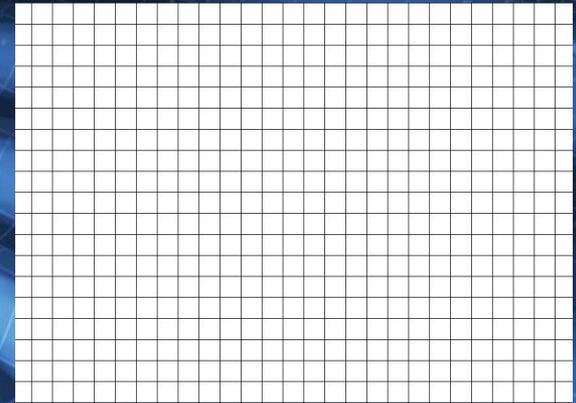


PennState Institute for Gravitation & the Cosmos
Institute for CyberScience Eberly College of Science



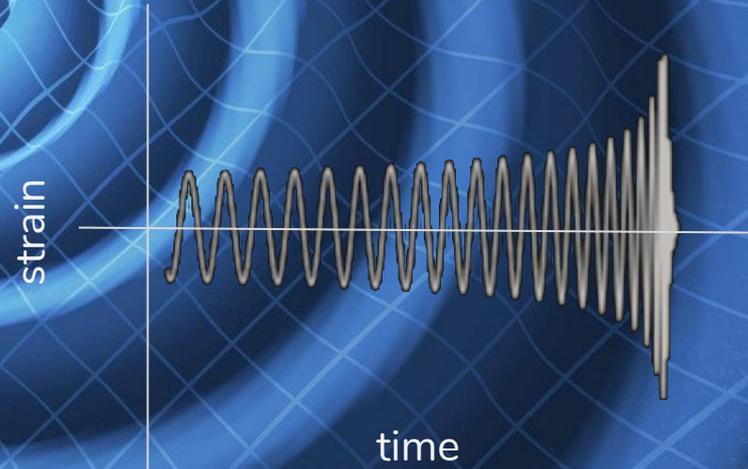
LIGO
Scientific
Collaboration

Distant observers see changes
in length: strain = $\Delta L/L$

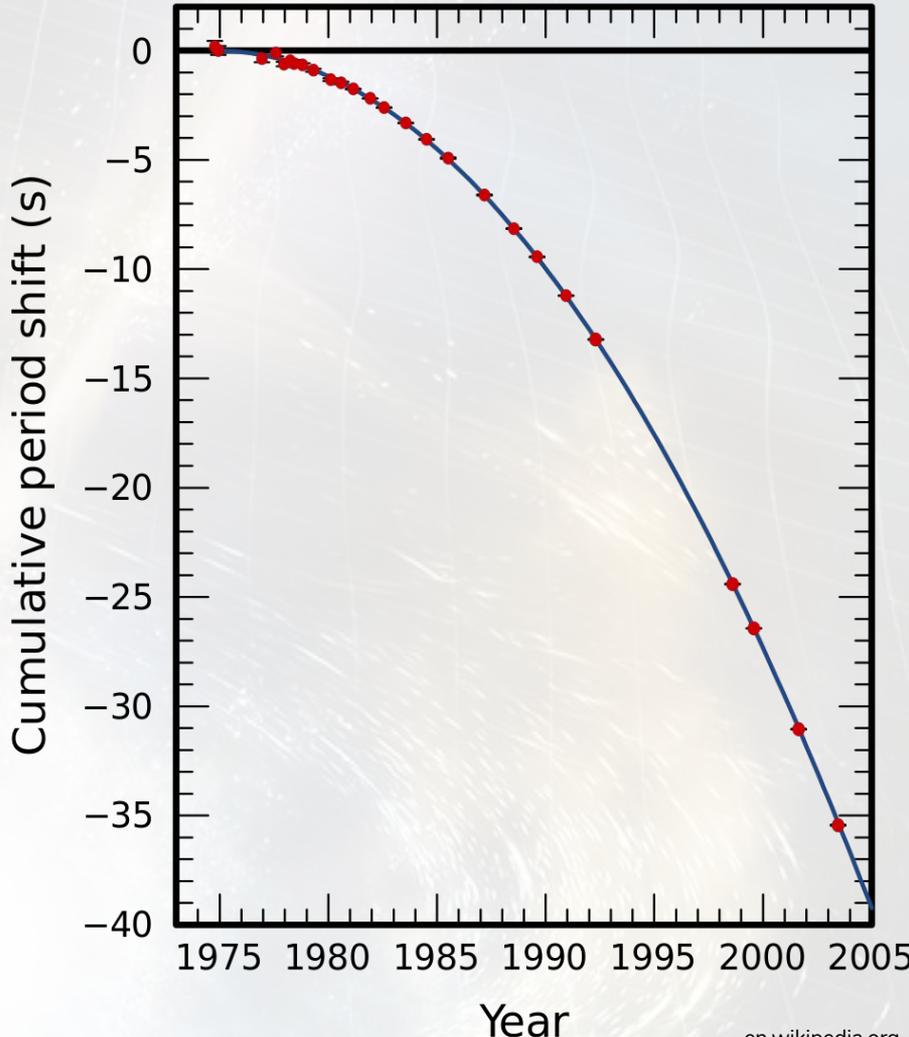


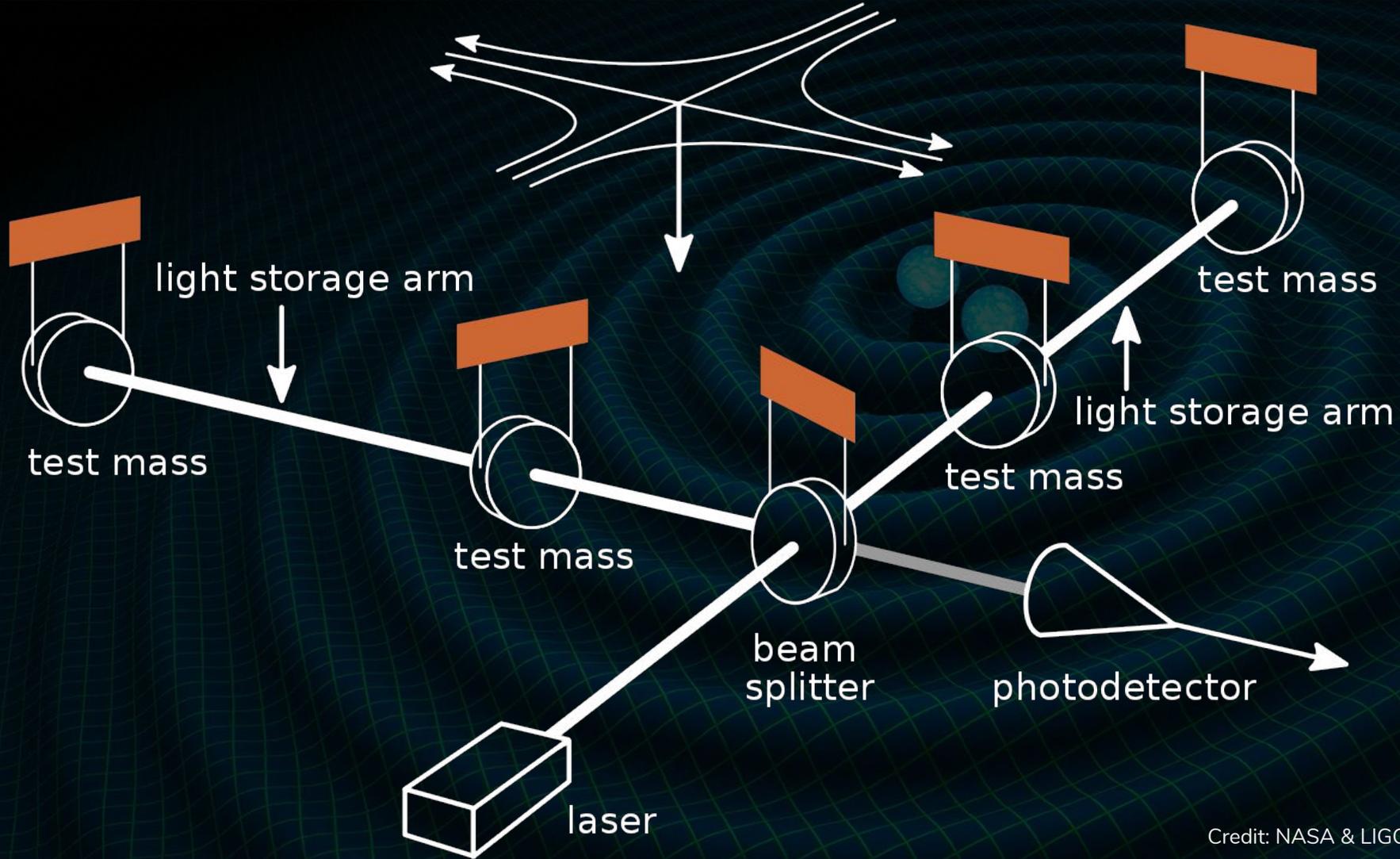
Energy loss drives the binary to
merge. Peak strain on Earth:
 10^{-21}

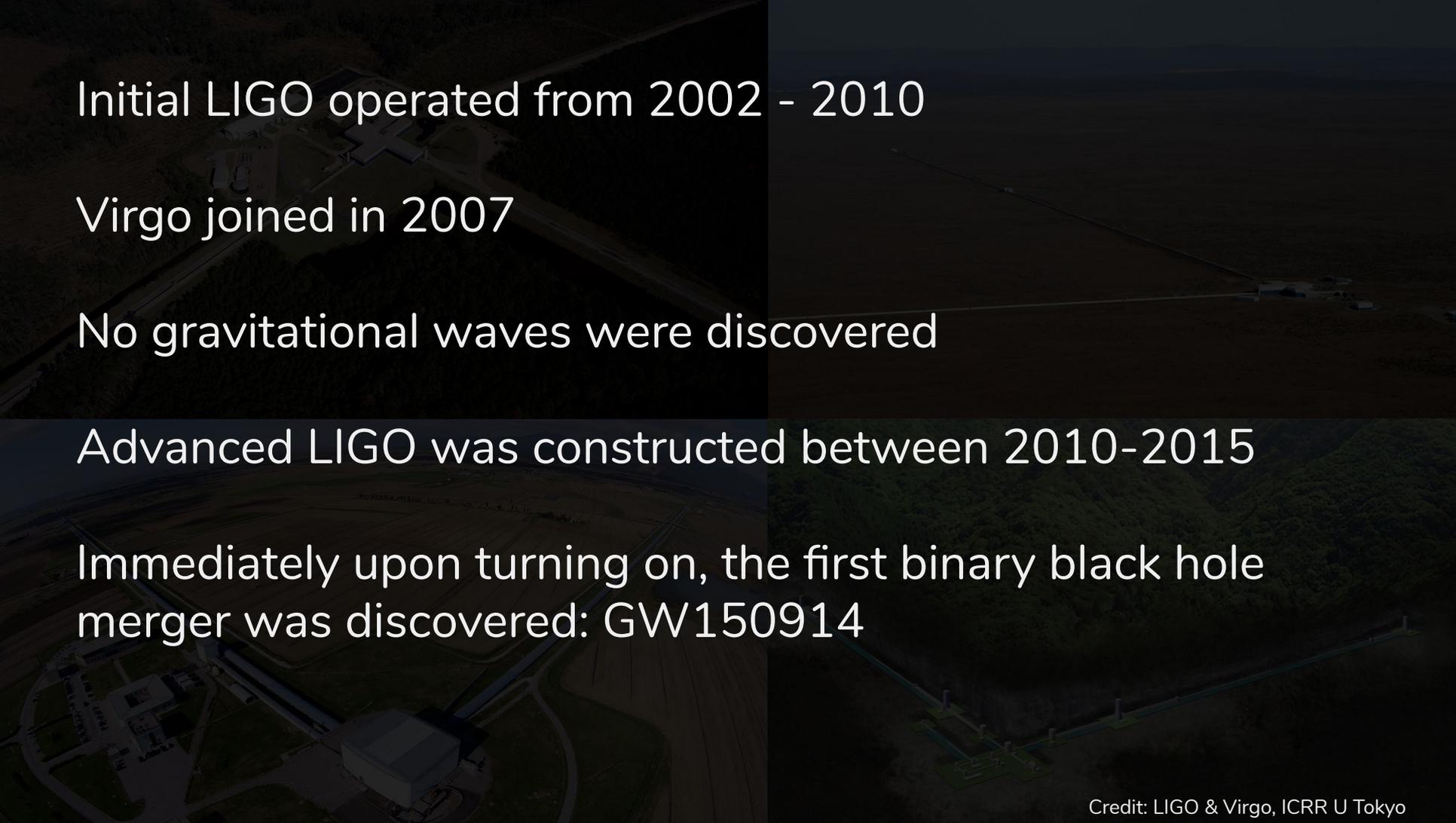
Waveform depends on
 M_1, m_2, s_1, s_2 , position, orientation, ...



The Hulse-Taylor Pulsar PSR B1913+16 provided conclusive evidence that double compact objects would be gravitational wave sources







Initial LIGO operated from 2002 - 2010

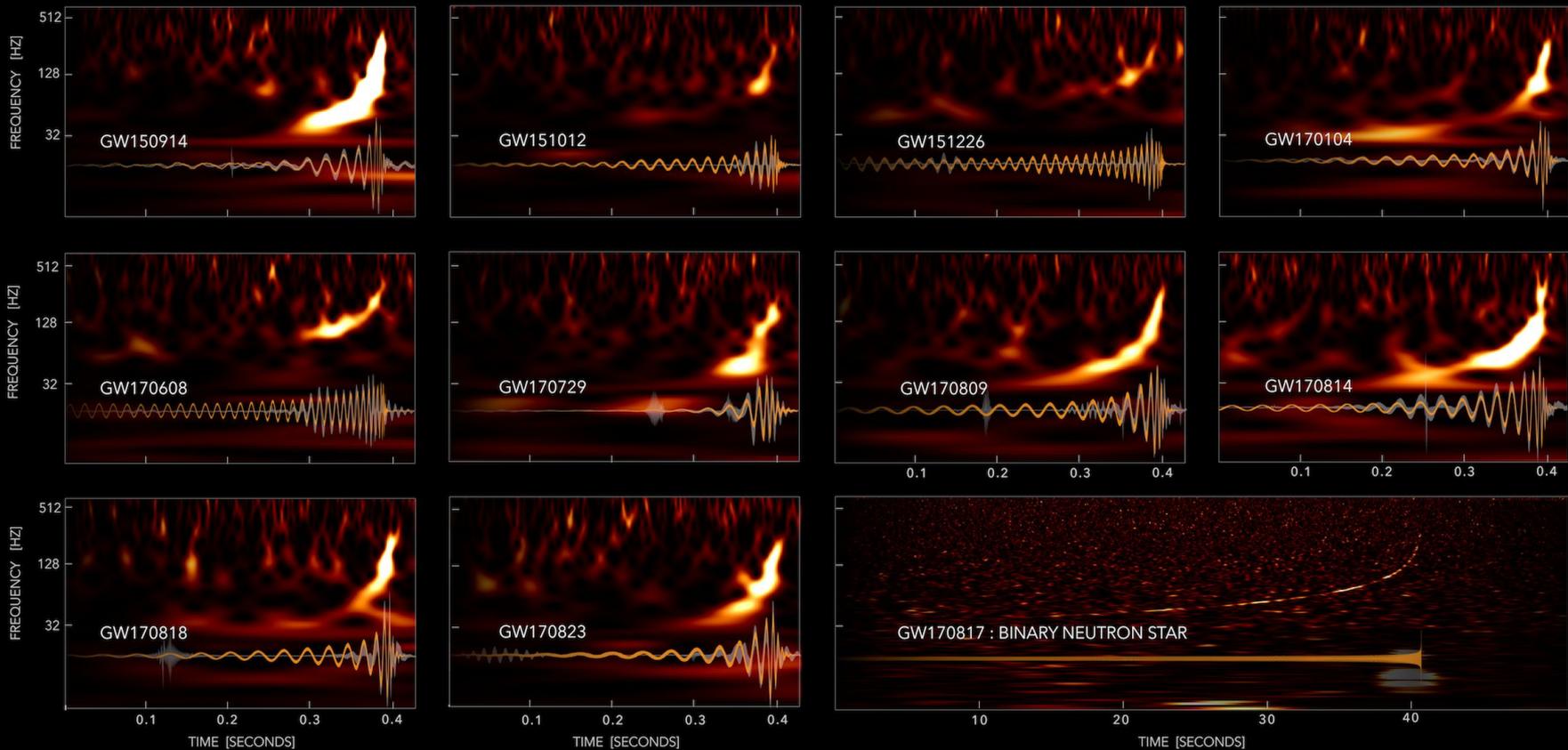
Virgo joined in 2007

No gravitational waves were discovered

Advanced LIGO was constructed between 2010-2015

Immediately upon turning on, the first binary black hole merger was discovered: GW150914

GWTC-1 - GWs discovered from September 2015 - August 2017

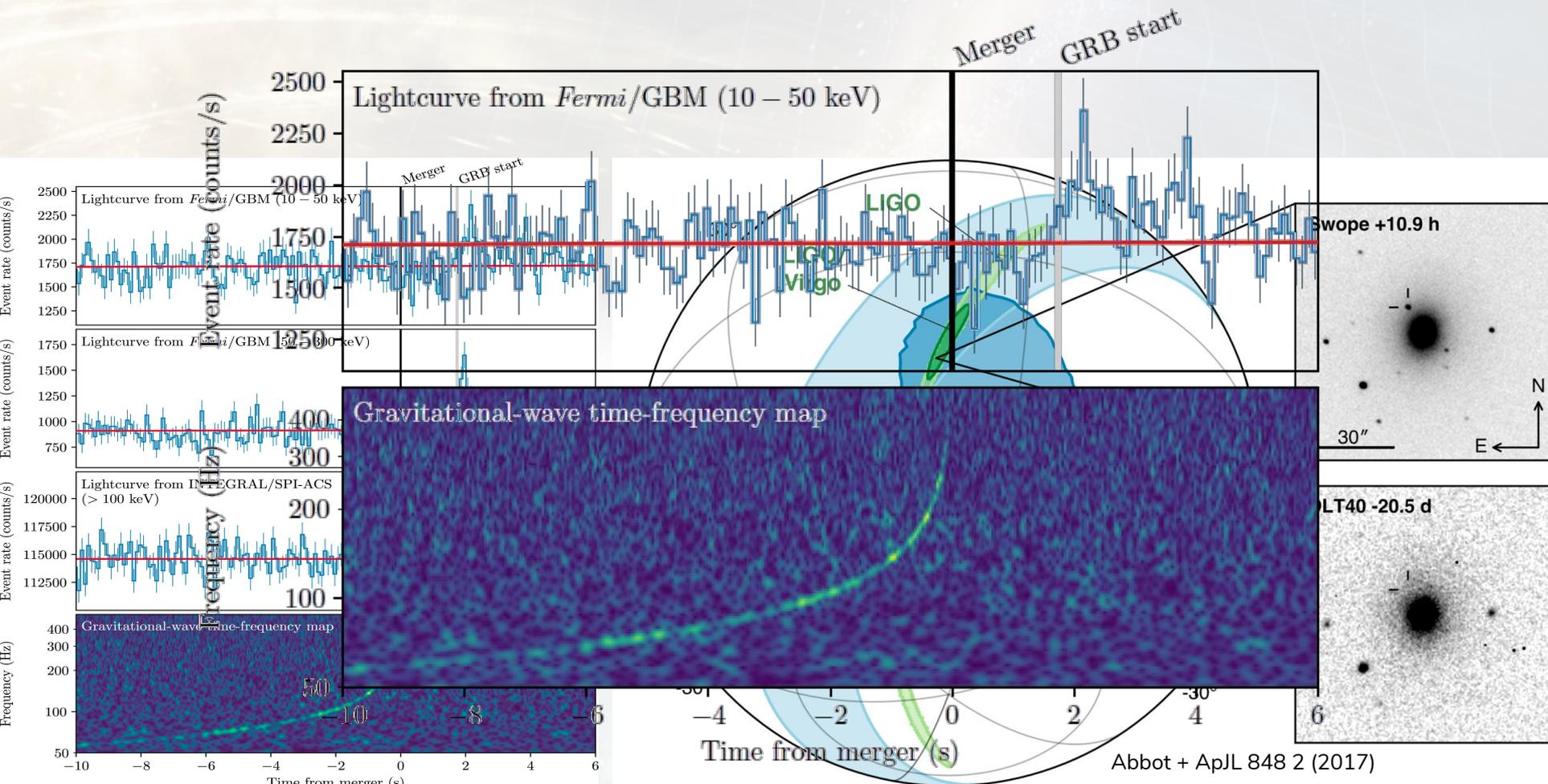


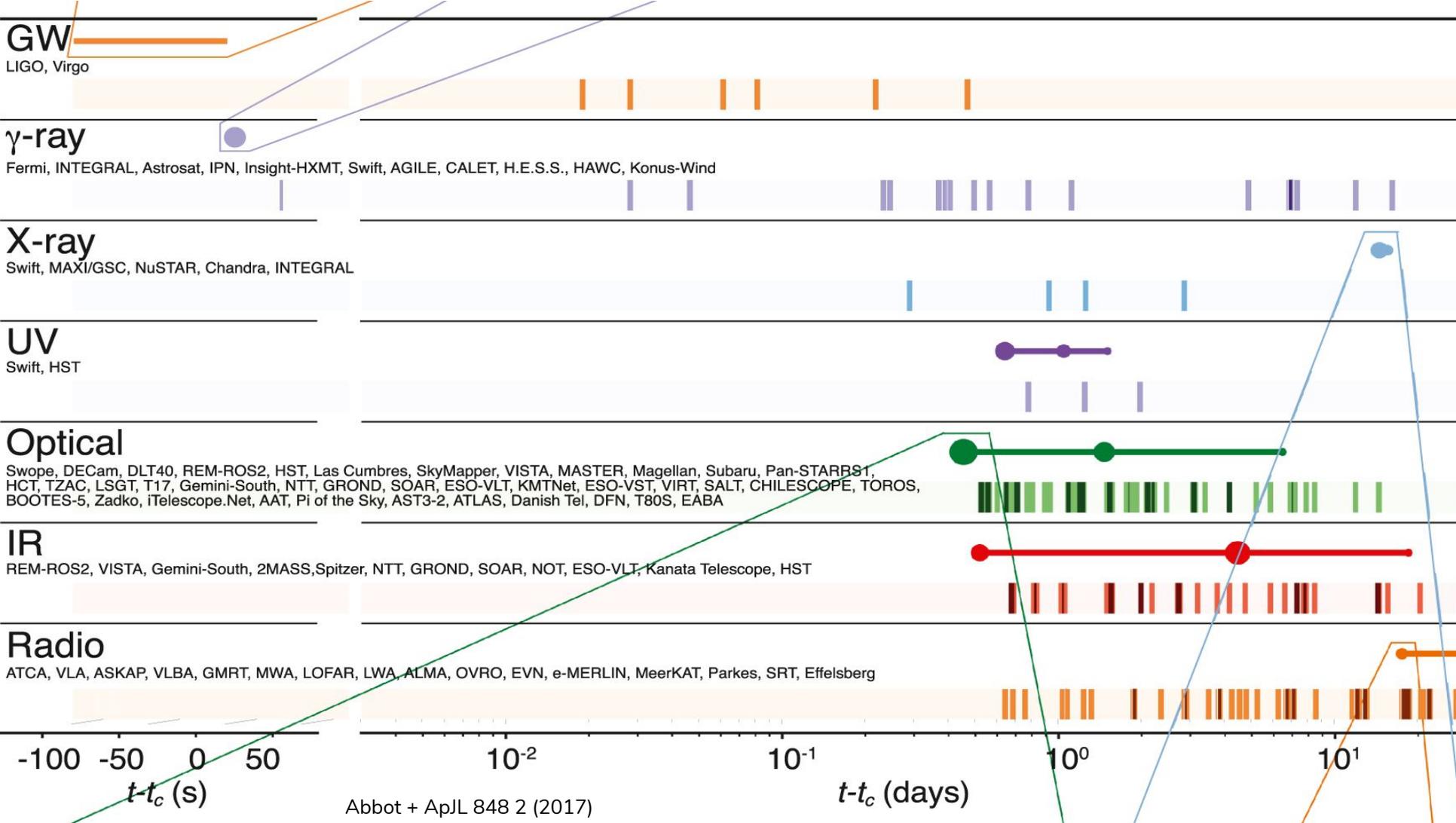
GWTC-1 - GWs discovered from September 2015 - August 2017

B. P. Abbott et al. 2019

Event	m_1/M_\odot	m_2/M_\odot	M/M_\odot	χ_{eff}	M_f/M_\odot	a_f	$E_{\text{rad}}/(M_\odot c^2)$	z	$\Delta\Omega/\text{deg}^2$
GW150914	$35.6^{+4.8}_{-3.0}$	$30.6^{+3.0}_{-4.4}$	$28.6^{+1.6}_{-1.5}$	$-0.01^{+0.12}_{-0.13}$	$63.1^{+3.3}_{-3.0}$	$0.69^{+0.05}_{-0.04}$	$3.1^{+0.4}_{-0.4}$	$0.09^{+0.03}_{-0.03}$	180
GW151012	$23.3^{+14.0}_{-5.5}$	$13.6^{+4.1}_{-4.8}$	$15.2^{+2.0}_{-1.1}$	$0.04^{+0.28}_{-0.19}$	$35.7^{+9.9}_{-3.8}$	$0.67^{+0.13}_{-0.11}$	$1.5^{+0.5}_{-0.5}$	$0.21^{+0.09}_{-0.09}$	1555
GW151226	$13.7^{+8.8}_{-3.2}$	$7.7^{+2.2}_{-2.6}$	$8.9^{+0.3}_{-0.3}$	$0.18^{+0.20}_{-0.12}$	$20.5^{+6.4}_{-1.5}$	$0.74^{+0.07}_{-0.05}$	$1.0^{+0.1}_{-0.2}$	$0.09^{+0.04}_{-0.04}$	1033
GW170104	$31.0^{+7.2}_{-5.6}$	$20.1^{+4.9}_{-4.5}$	$21.5^{+2.1}_{-1.7}$	$-0.04^{+0.17}_{-0.20}$	$49.1^{+5.2}_{-3.9}$	$0.66^{+0.08}_{-0.10}$	$2.2^{+0.5}_{-0.5}$	$0.19^{+0.07}_{-0.08}$	924
GW170608	$10.9^{+5.3}_{-1.7}$	$7.6^{+1.3}_{-2.1}$	$7.9^{+0.2}_{-0.2}$	$0.03^{+0.19}_{-0.07}$	$17.8^{+3.2}_{-0.7}$	$0.69^{+0.04}_{-0.04}$	$0.9^{+0.05}_{-0.1}$	$0.07^{+0.02}_{-0.02}$	396
GW170729	$50.6^{+16.6}_{-10.2}$	$34.3^{+9.1}_{-10.1}$	$35.7^{+6.5}_{-4.7}$	$0.36^{+0.21}_{-0.25}$	$80.3^{+14.6}_{-10.2}$	$0.81^{+0.07}_{-0.13}$	$4.8^{+1.7}_{-1.7}$	$0.48^{+0.19}_{-0.20}$	1033
GW170809	$35.2^{+8.3}_{-6.0}$	$23.8^{+5.2}_{-5.1}$	$25.0^{+2.1}_{-1.6}$	$0.07^{+0.16}_{-0.16}$	$56.4^{+5.2}_{-3.7}$	$0.70^{+0.08}_{-0.09}$	$2.7^{+0.6}_{-0.6}$	$0.20^{+0.05}_{-0.07}$	340
GW170814	$30.7^{+5.7}_{-3.0}$	$25.3^{+2.9}_{-4.1}$	$24.2^{+1.4}_{-1.1}$	$0.07^{+0.12}_{-0.11}$	$53.4^{+3.2}_{-2.4}$	$0.72^{+0.07}_{-0.05}$	$2.7^{+0.4}_{-0.3}$	$0.12^{+0.03}_{-0.04}$	87
GW170817	$1.46^{+0.12}_{-0.10}$	$1.27^{+0.09}_{-0.09}$	$1.186^{+0.001}_{-0.001}$	$0.00^{+0.02}_{-0.01}$	≤ 2.8	≤ 0.89	≥ 0.04	$0.01^{+0.00}_{-0.00}$	16
GW170818	$35.5^{+7.5}_{-4.7}$	$26.8^{+4.3}_{-5.2}$	$26.7^{+2.1}_{-1.7}$	$-0.09^{+0.18}_{-0.21}$	$59.8^{+4.8}_{-3.8}$	$0.67^{+0.07}_{-0.08}$	$2.7^{+0.5}_{-0.5}$	$0.20^{+0.07}_{-0.07}$	39
GW170823	$39.6^{+10.0}_{-6.6}$	$29.4^{+6.3}_{-7.1}$	$29.3^{+4.2}_{-3.2}$	$0.08^{+0.20}_{-0.22}$	$65.6^{+9.4}_{-6.6}$	$0.71^{+0.08}_{-0.10}$	$3.3^{+0.9}_{-0.8}$	$0.34^{+0.13}_{-0.14}$	1651

GW170817: First multi-messenger source of GWs





Masses in the Stellar Graveyard

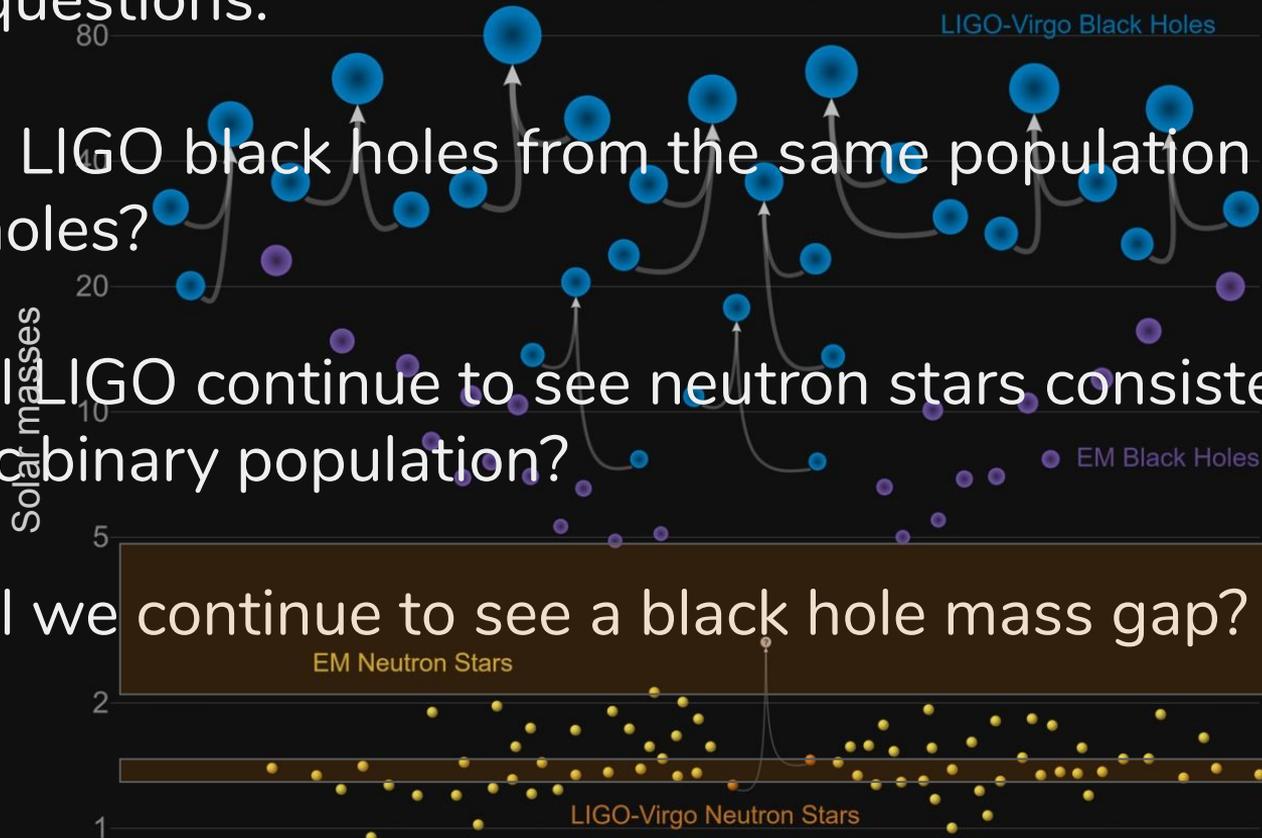
in Solar Masses

Open questions:

Are LIGO black holes from the same population as “EM” black holes?

Will LIGO continue to see neutron stars consistent with galactic binary population?

Will we continue to see a black hole mass gap?

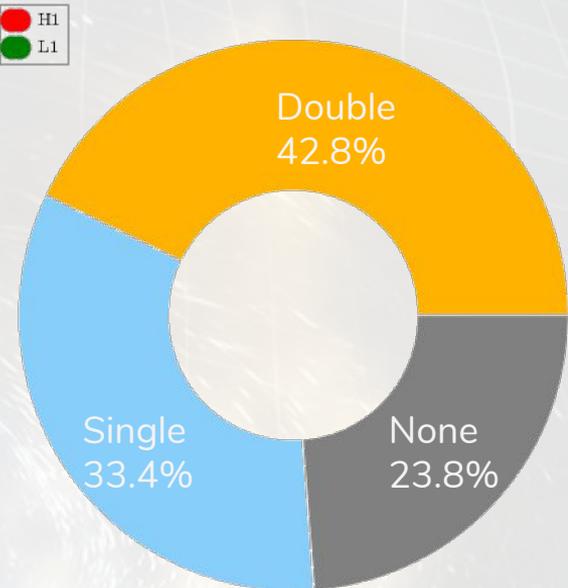
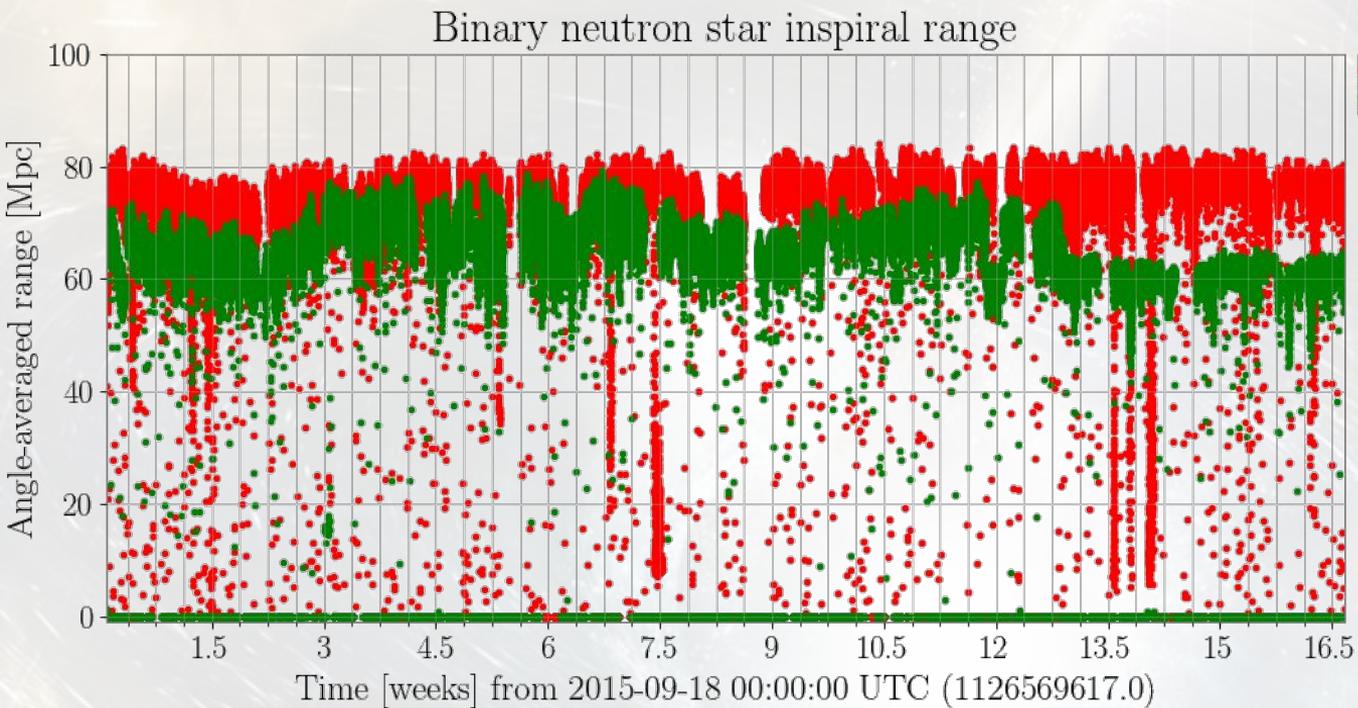


Ozel et al 2012

Timeline

O1

Sept '15 -- Jan '16



Timeline

O1

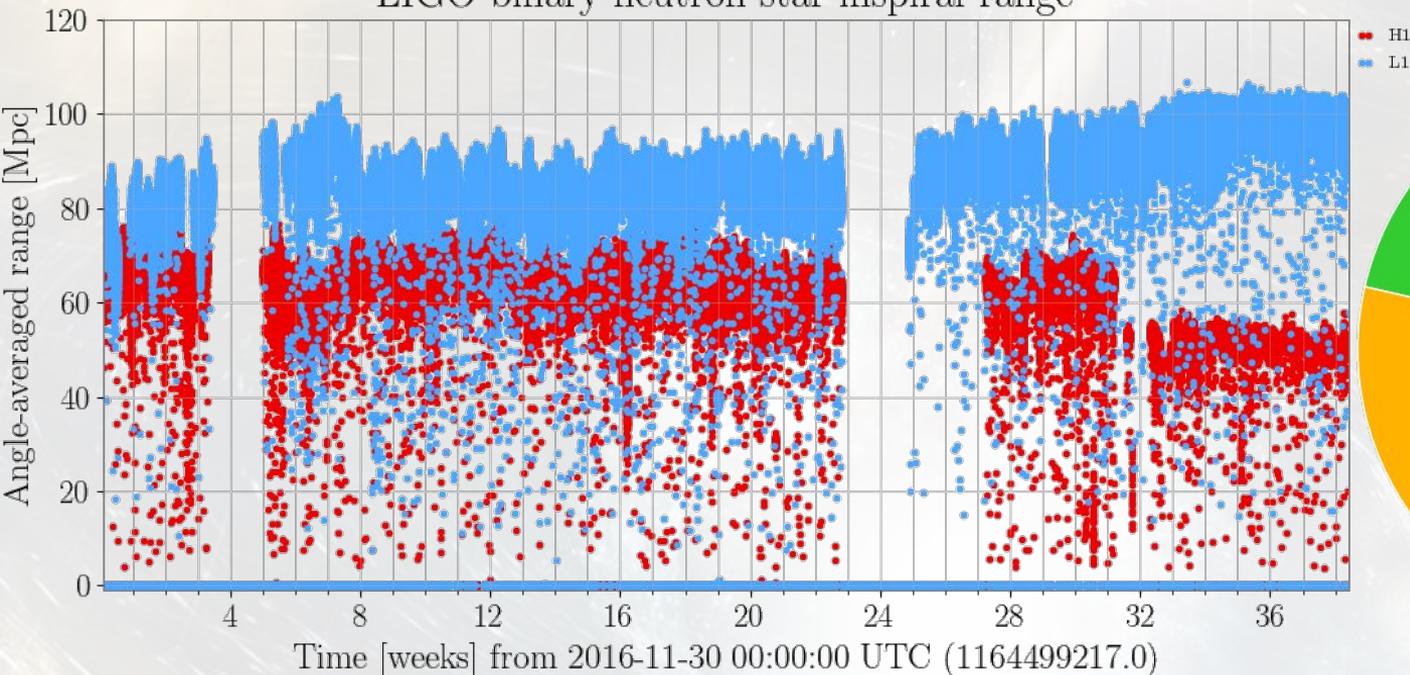
Commissioning -
No data

O2

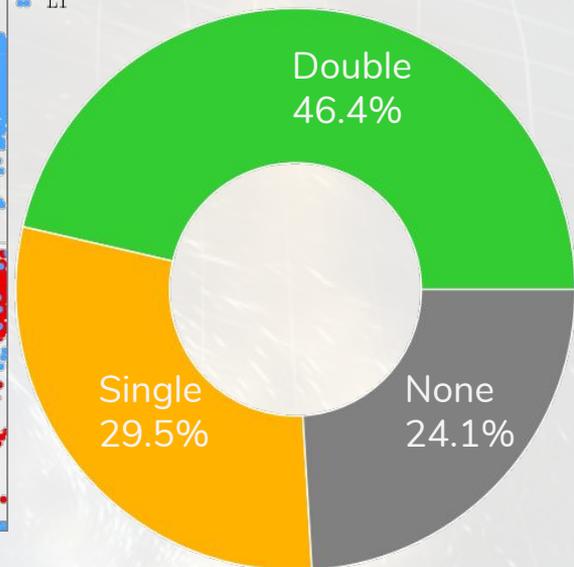
Sept '15 -- Jan '16

Nov '16 -- Aug '17

LIGO binary neutron star inspiral range



•• H1
•• L1



Timeline

O1

Commissioning -
No data

O2

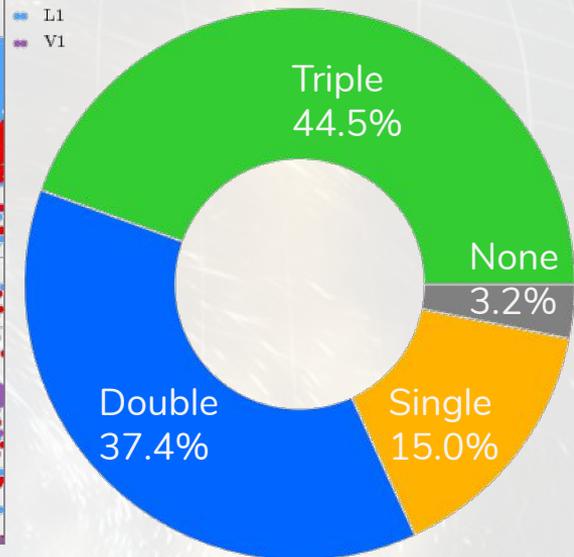
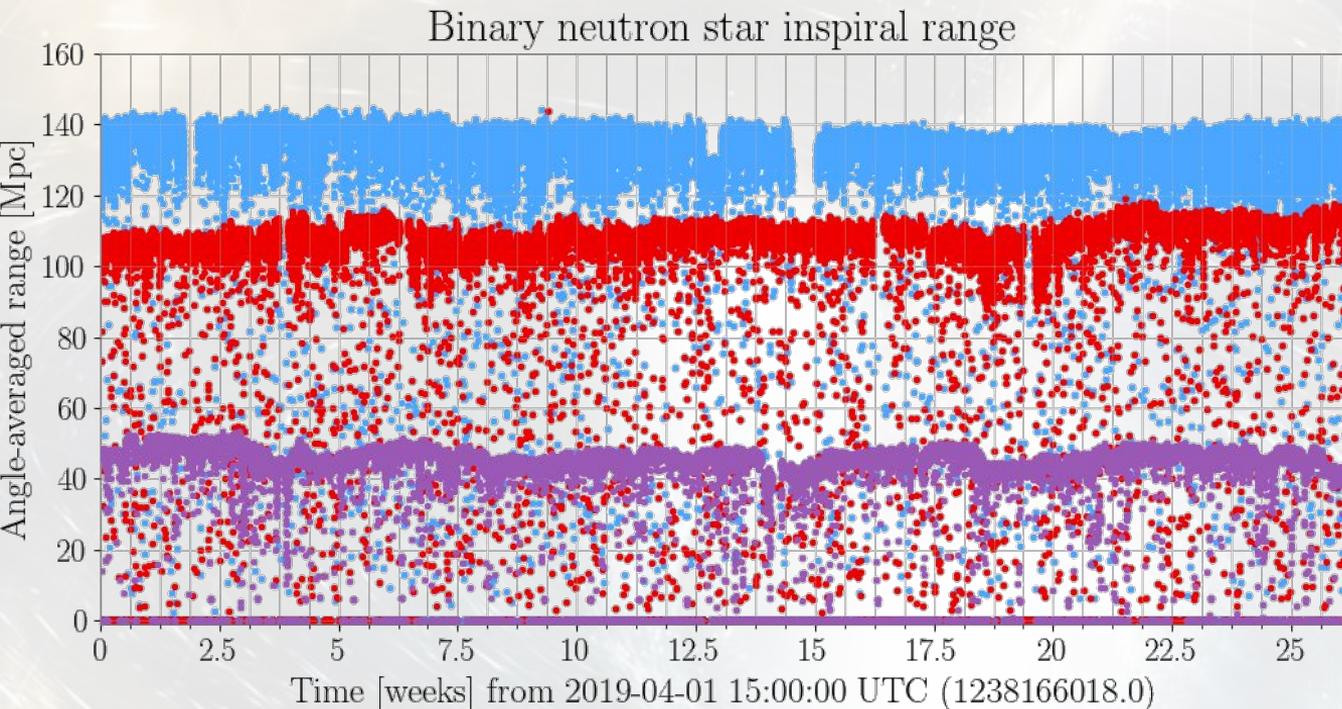
Commissioning -
No data

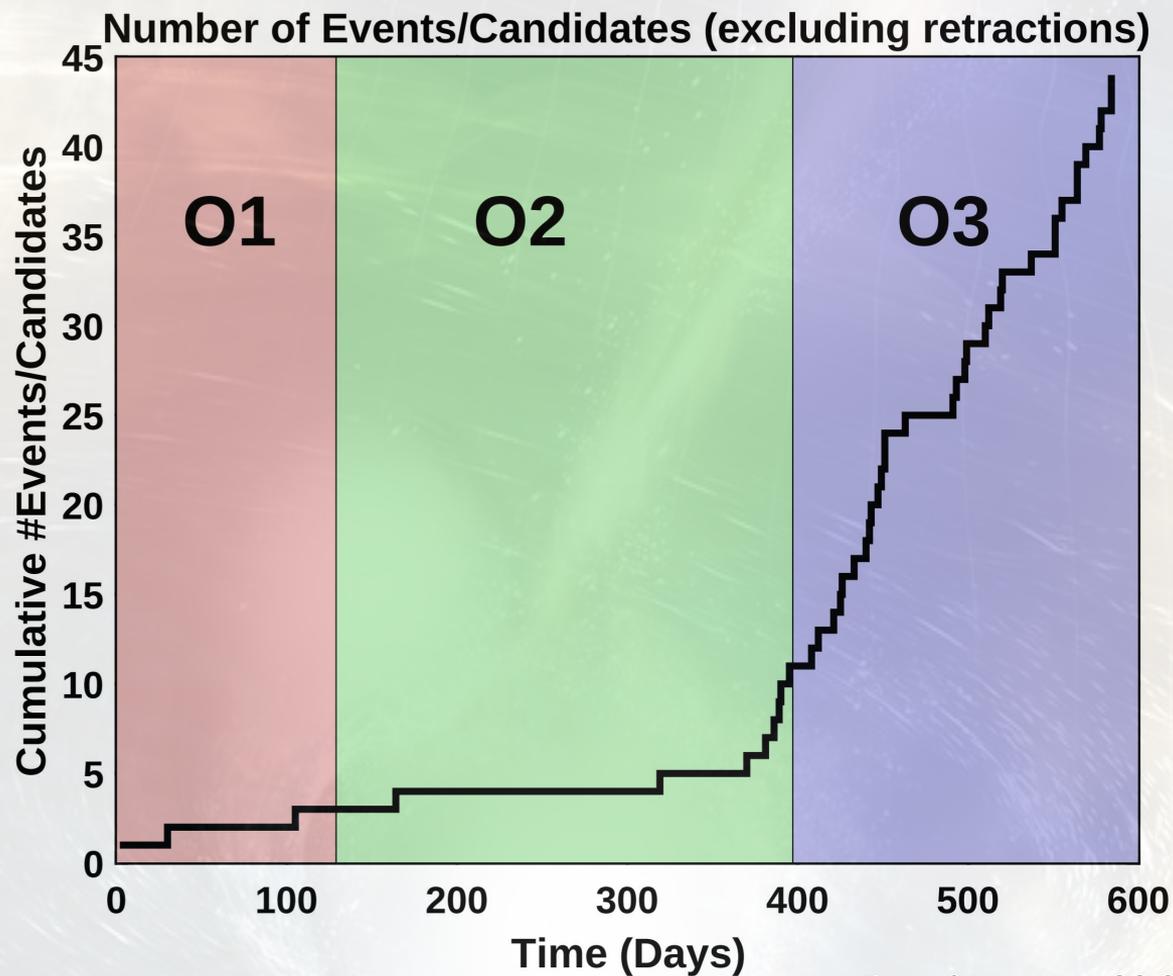
O3a

Sept '15 -- Jan '16

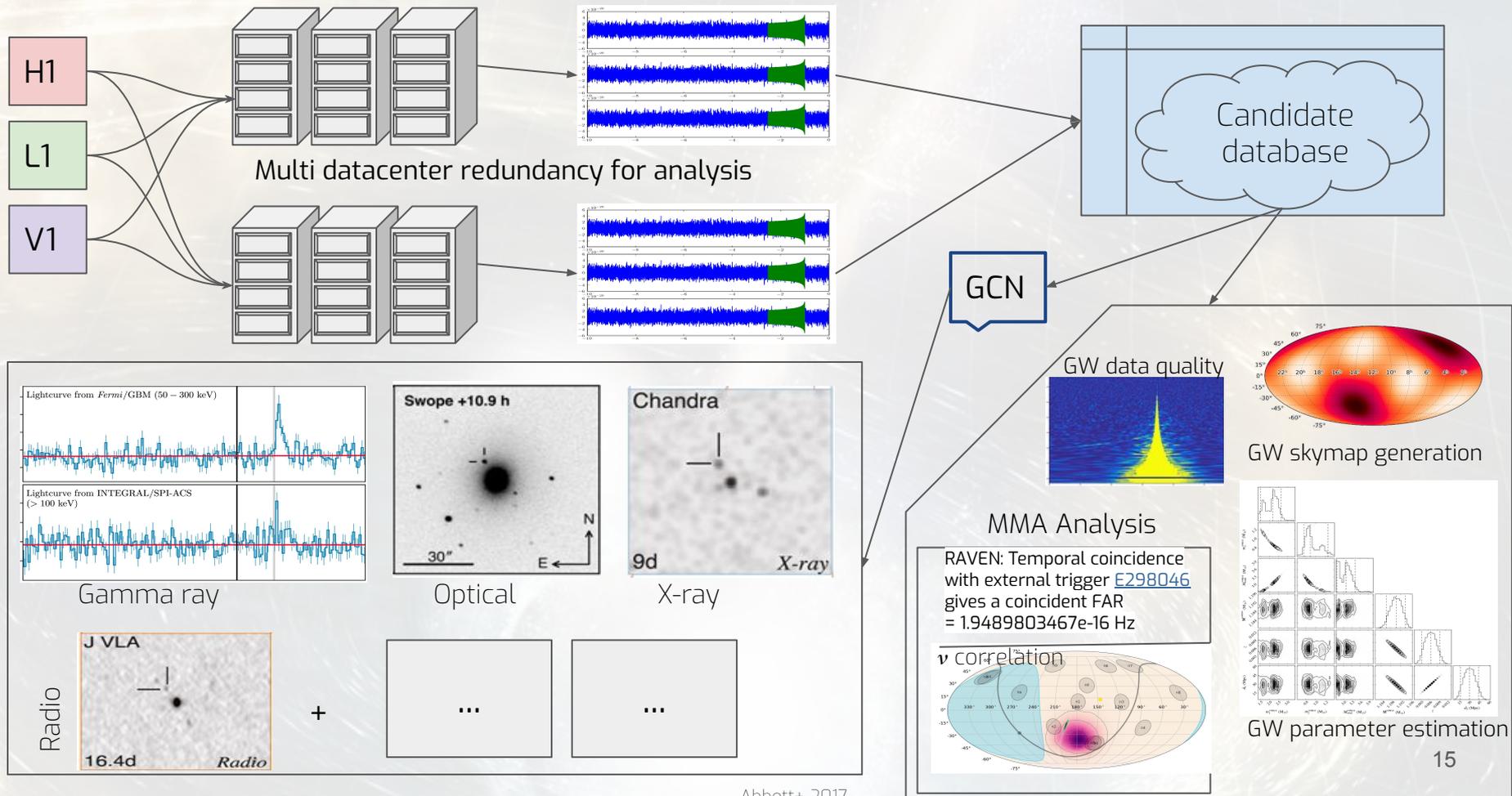
Nov '16 -- Aug '17

Apr '19 -- Sep '19





Real-time GW data processing (this is not a complete diagram)

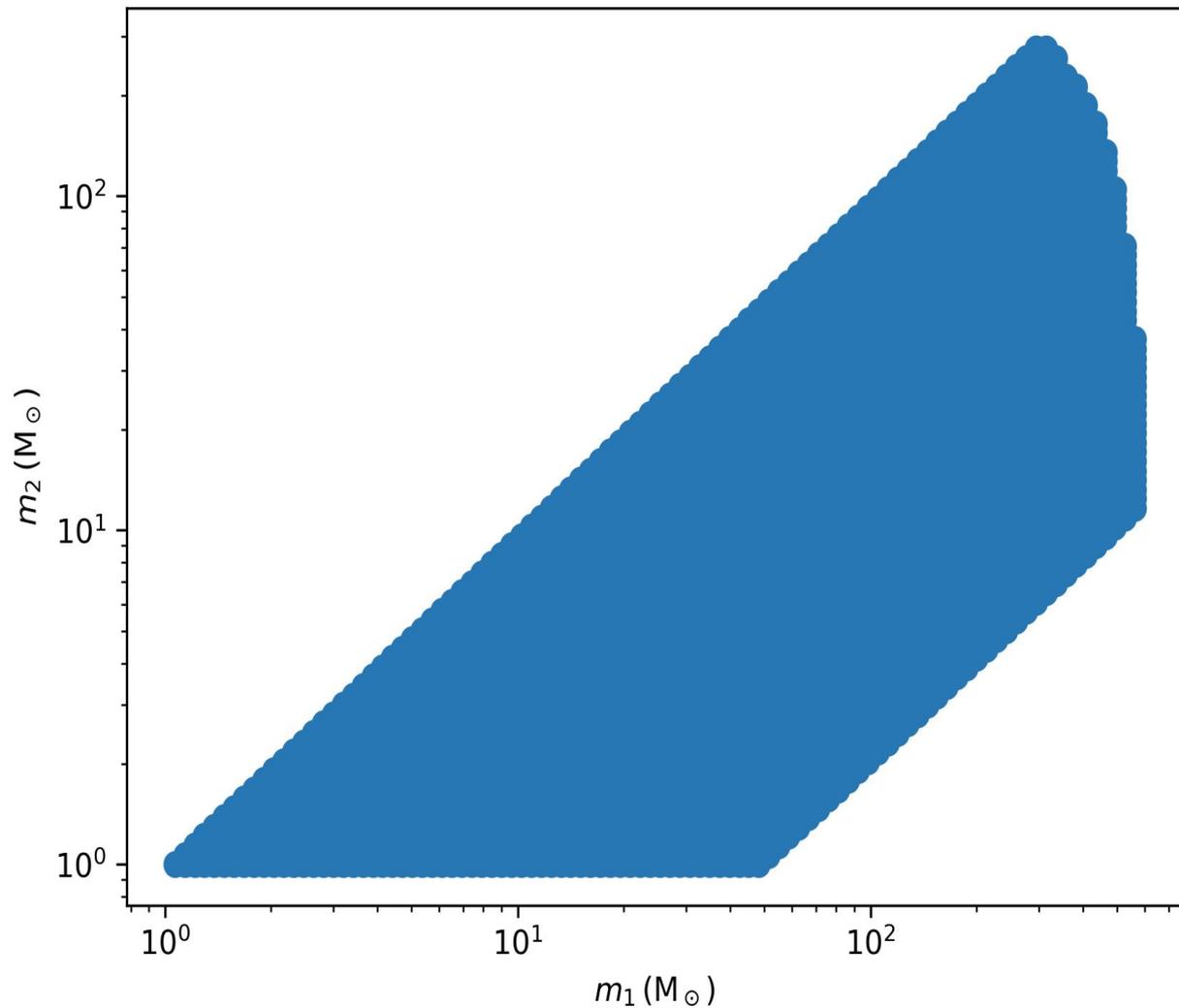


Real-time analysis covers:

- Component: 1 - 600 M_{\odot}
- Total < 600 M_{\odot}
- NS spin < 0.05
- BH spin < 1
- $Q < 50$

Millions of templates are searched in parallel

Candidates are uploaded to GW database within 20 seconds of data acquisition.



Open public GW alerts:

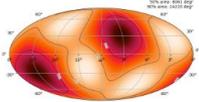
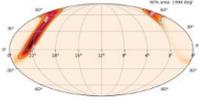
<https://gracedb.ligo.org/superevents/public/O3/>

LIGO/Virgo O3 Public Alerts

Detection candidates: 33

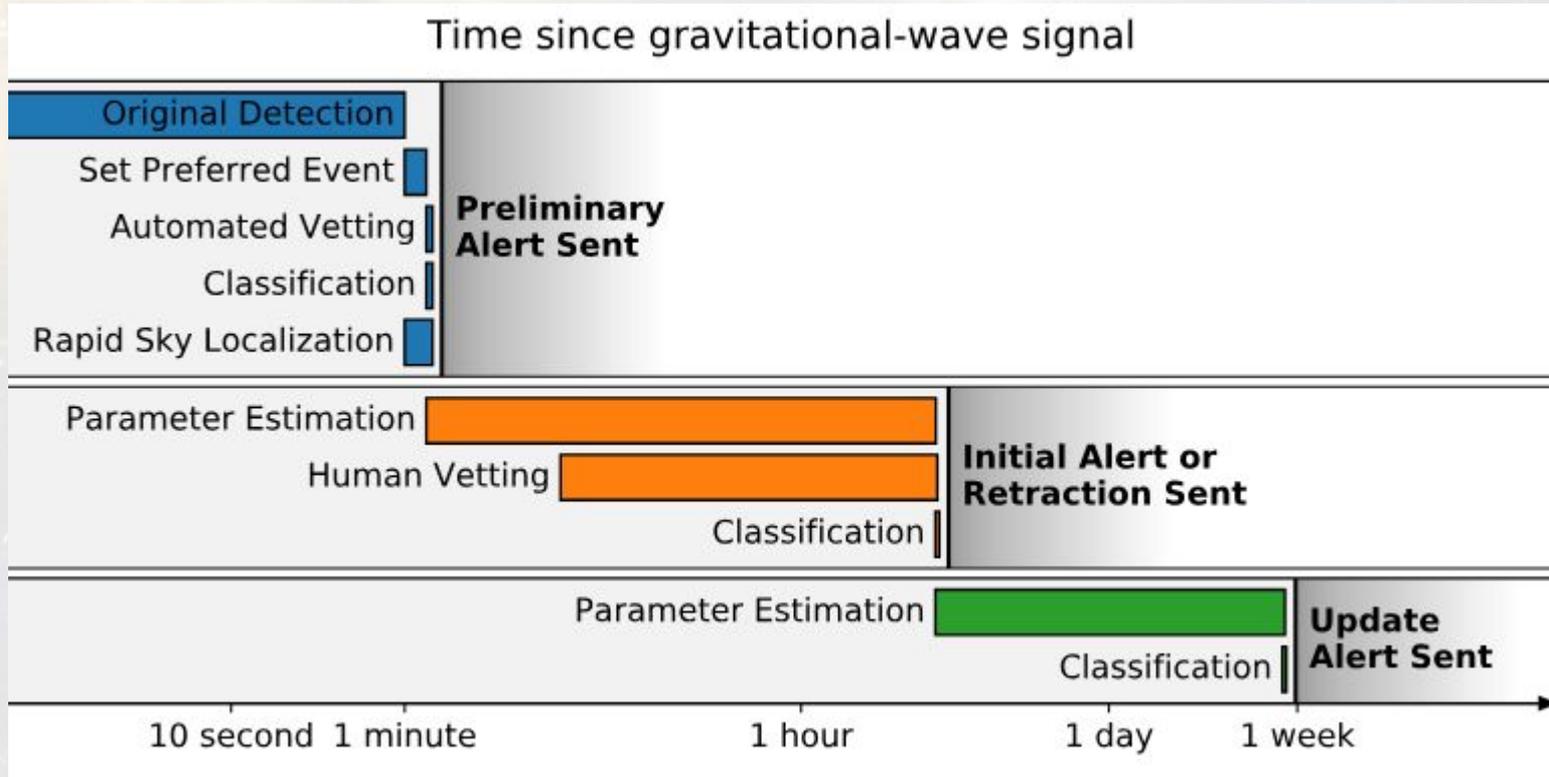
SORT: EVENT ID (A-Z) ▾



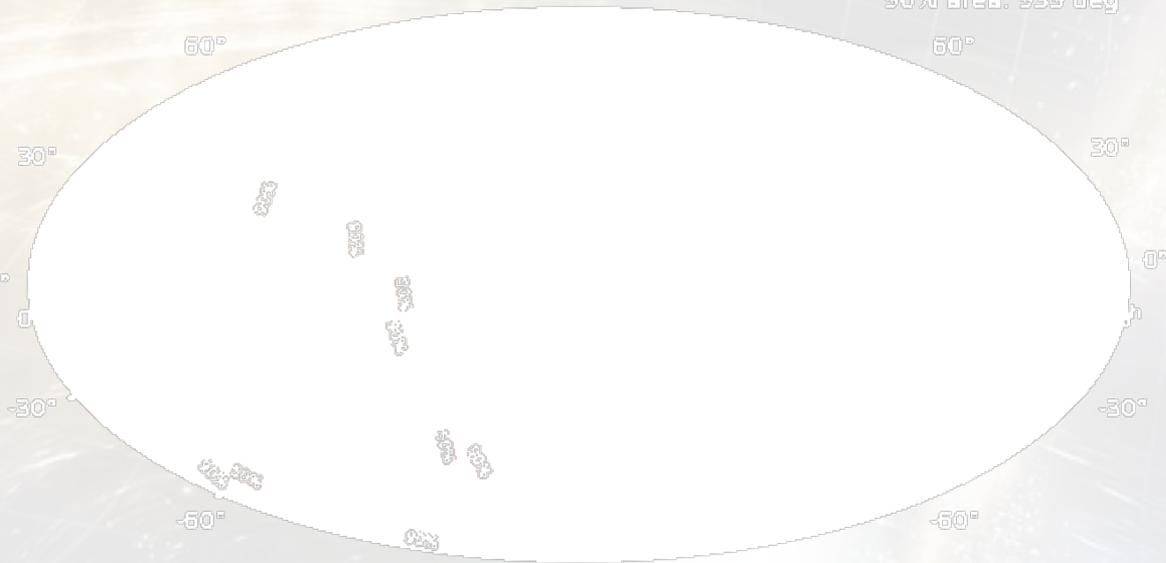
Event ID	Possible Source (Probability)	UTC	GCN	Location	FAR
S190930t	NSBH (74%), Terrestrial (26%)	Sept. 30, 2019 14:34:07 UTC	GCN Circulars Notices VOE		1 per 2.0536 years
S190930s	MassGap (95%), Terrestrial (5%)	Sept. 30, 2019 13:35:41 UTC	GCN Circulars Notices VOE		1 per 10.534 years

Open public GW alerts:

<https://emfollow.docs.ligo.org/userguide>



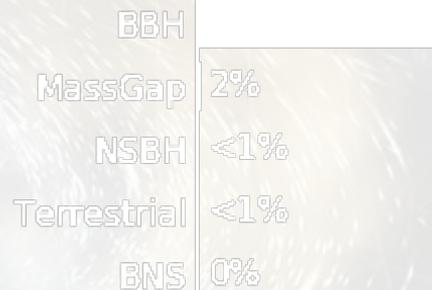
event ID: G333132
50% area: 152 deg²
90% area: 939 deg²

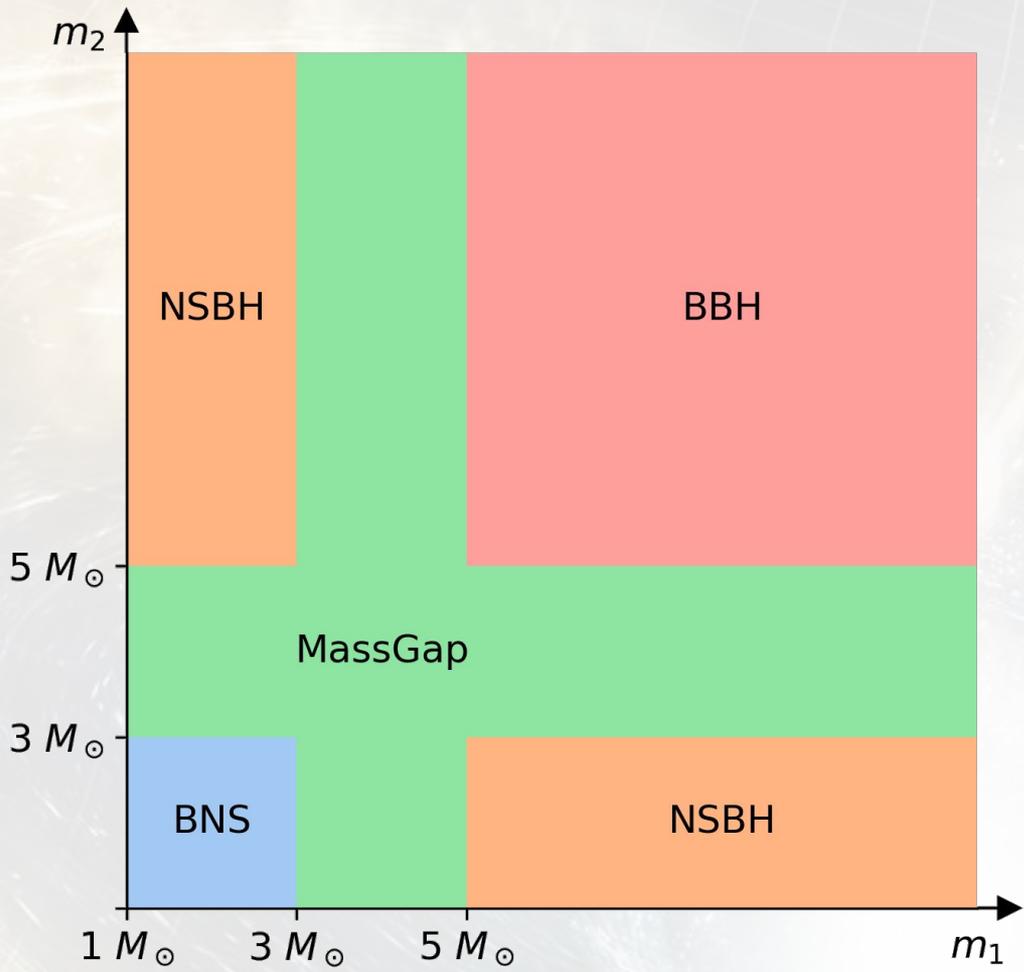


event ID: G333132
distance: 2950±1033 Mpc



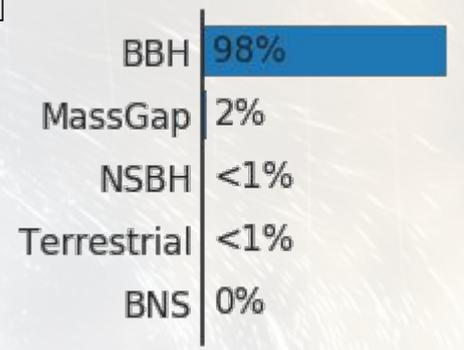
PROB_NS: 0.00 [range is 0.0-1.0]
PROB_REMNANT: 0.00 [range is 0.0-1.0]
FAR (yr⁻¹) 1 per 13.354 years





We provide “p-astro” classification based on the mass definitions to the left.

PROB_NS: 0.00 [range is 0.0-1.0]
 PROB_REMNANT: 0.00 [range is 0.0-1.0]



Number classified as BBH/Mass Gap: 23

Number classified as BNS: 4 (1 definite **, 1 reclassified as NSBH if real)

Number classified as NSBH: 4 (1 definite **)

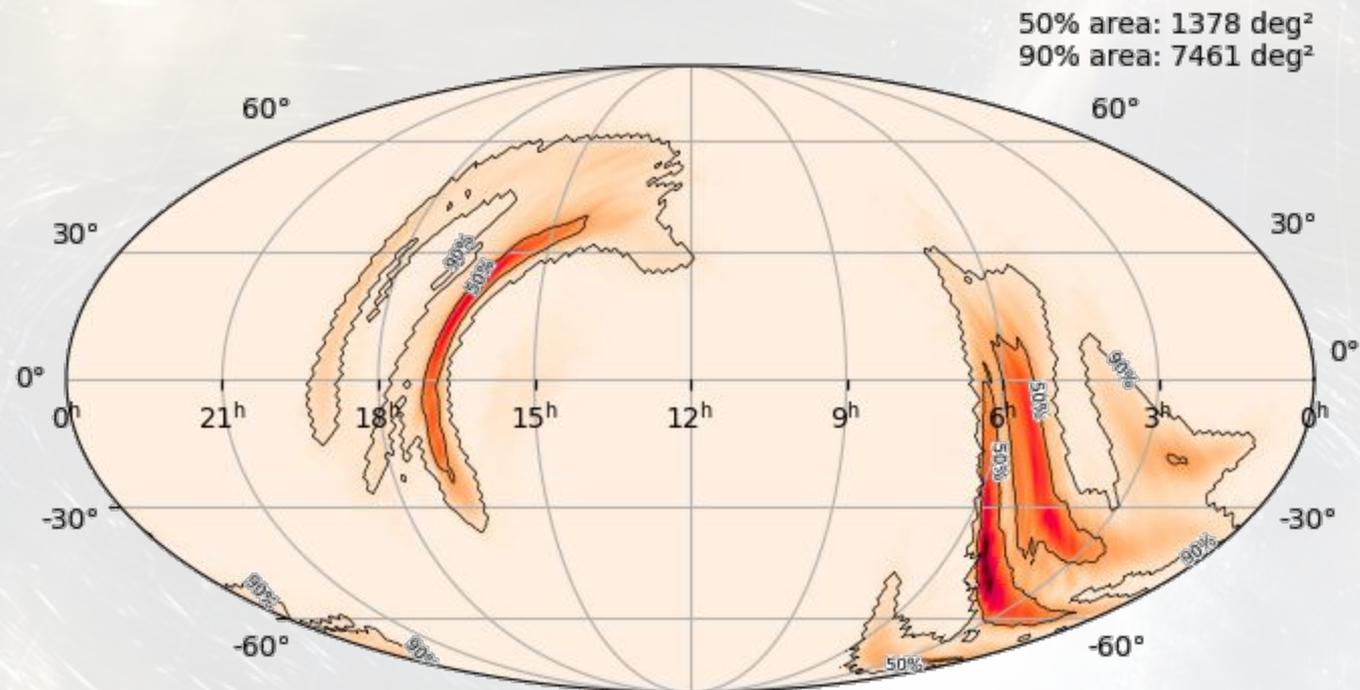
Number classified as terrestrial: 2

There have been >1000 GCNs associated with follow-up

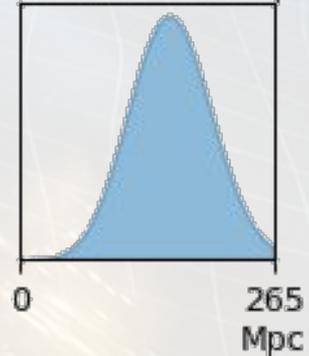
**based on mass alone - we do not know that there was an NS

BNS : S190425z (Definite GW)

Originally discovered as L only (V contributed very low SNR)



distance: 156 ± 41 Mpc



BNS	>99%
Terrestrial	<1%
NSBH	0%
MassGap	0%
BBH	0%

BNS -> NSBH : S190426c (Marginal)

Initially Classified as a BNS but very, very marginal

distance: 377±100 Mpc

```
////////////////////////////////////  
TITLE:    GCN CIRCULAR  
NUMBER:   24411  
SUBJECT:  LIGO/Virgo S190426c: Update on Source Classification  
DATE:     19/05/06 16:21:43 GMT  
FROM:     Deep Chatterjee at University of Wisconsin, Milwaukee <deep@uwm.edu>
```

LIGO/Virgo S190426c: Update on Source Classification

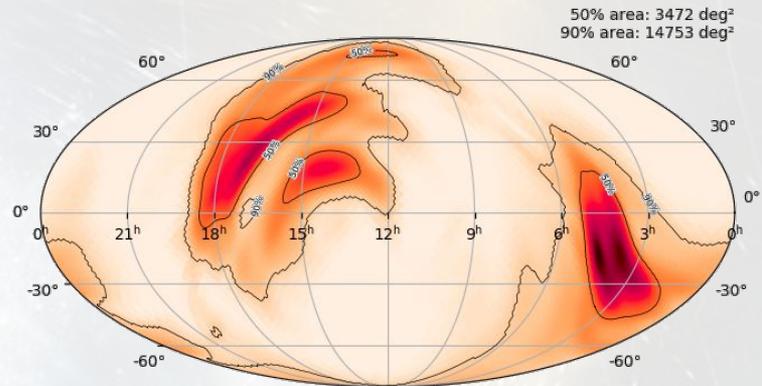
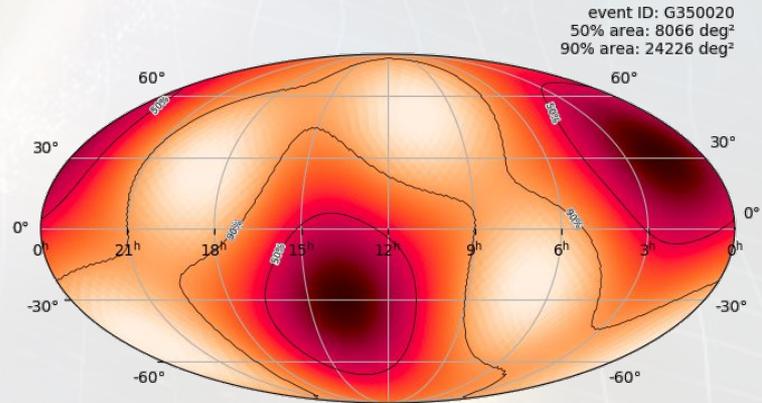
The LIGO Scientific Collaboration and Virgo Collaboration report:

Based on posterior support from preliminary parameter estimation [1,2], under the assumption that the candidate S190426c is astrophysical in origin, the relative probabilities amongst the signal categories **NSBH** : MassGap : BNS : BBH are revised to be approximately 12 : 5 : 3 : 0.

Other marginal BNS : S190901ap, S190910h

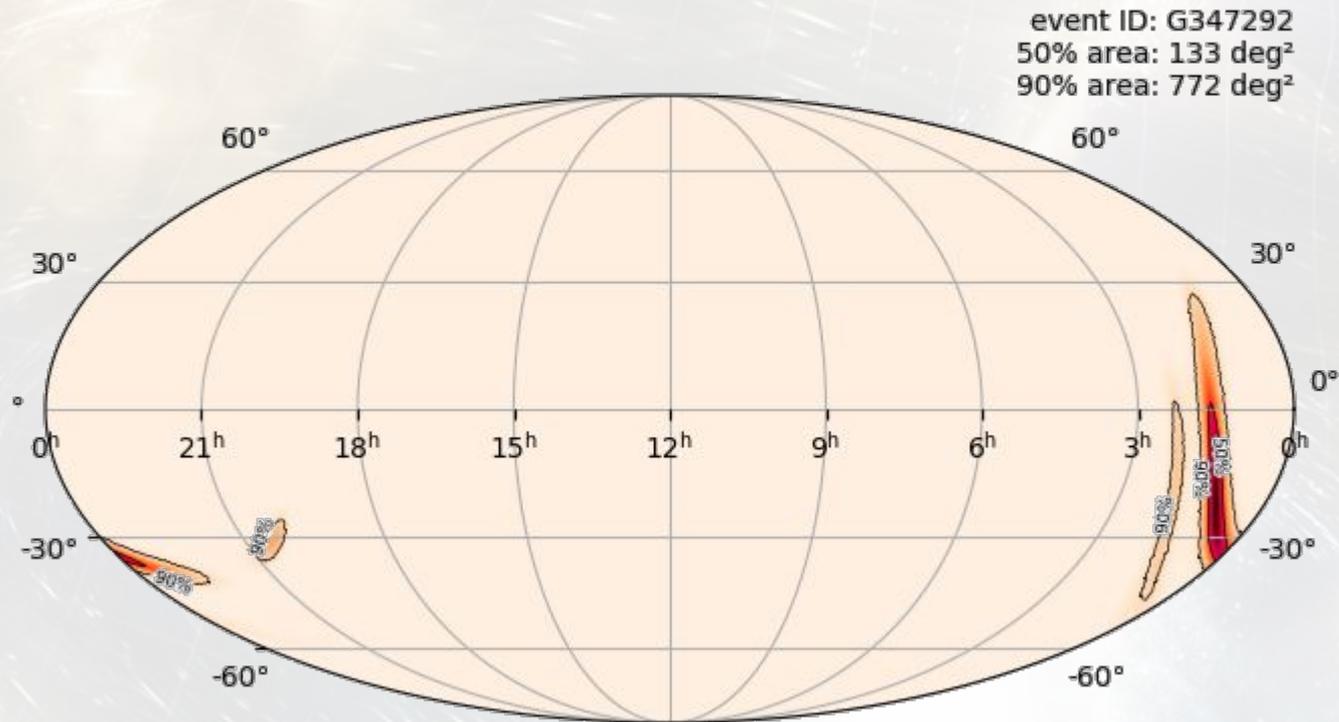
S190910h: BNS (61%), Terrestrial (39%)
L1

S190901ap: BNS (86%), Terrestrial (14%)
L1, V1

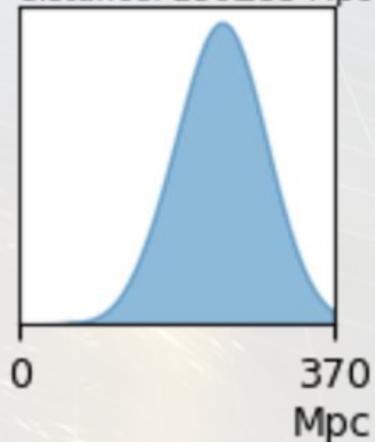


S190814bv (Definite GW)

Originally discovered as L,V only



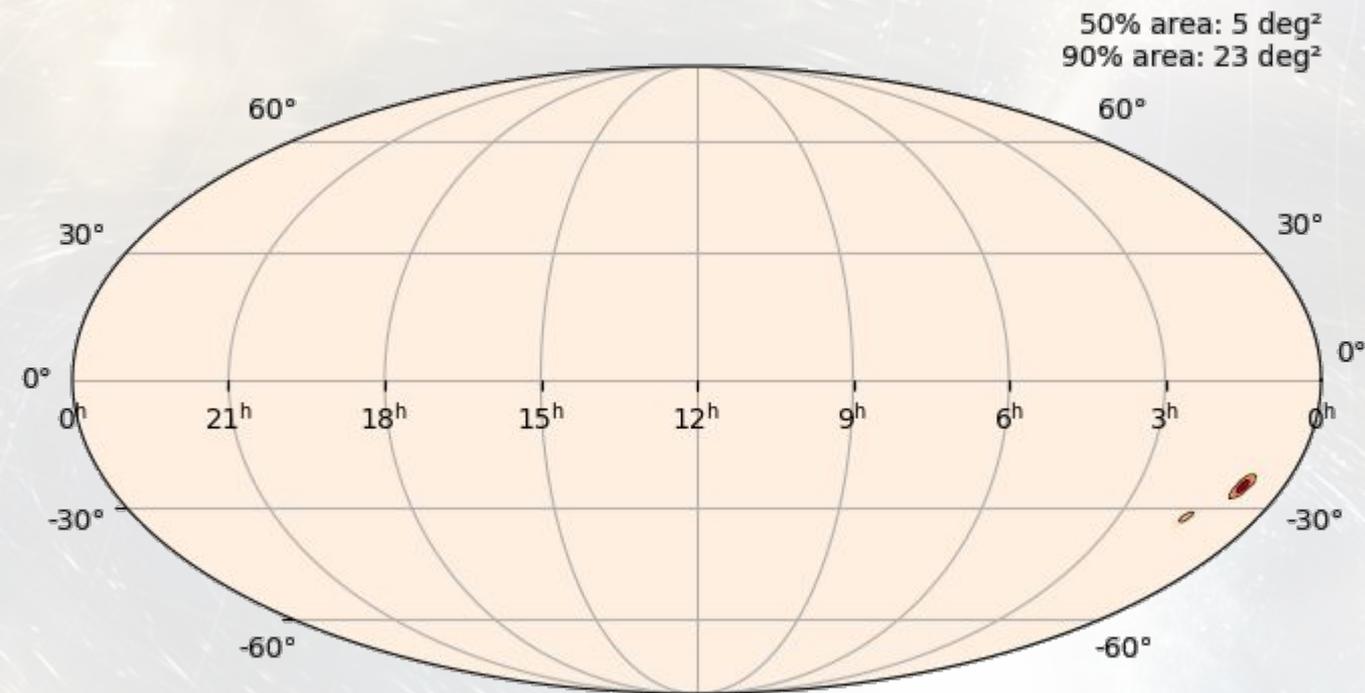
event ID: G347292
distance: 236±53 Mpc



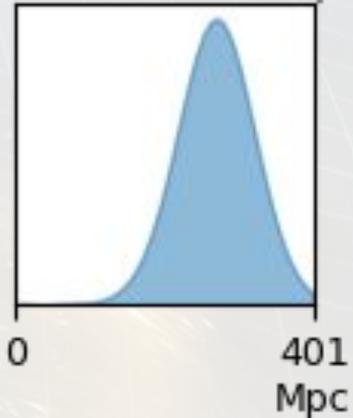
MassGap	100%
Terrestrial	<1%
NSBH	0%
BNS	0%
BBH	0%

S190814bv (Definite GW)

H had “low noise” data that was reanalyzed



distance: 267 ± 52 Mpc



MassGap 100%

Terrestrial <1%

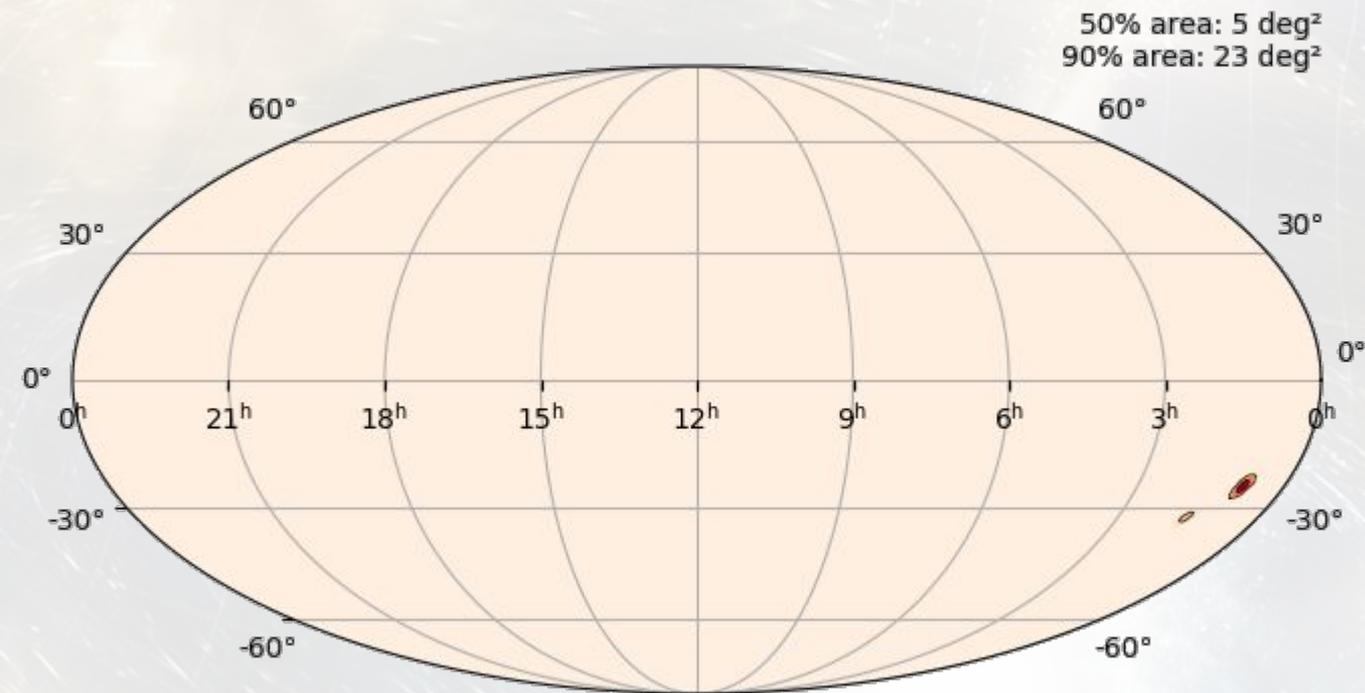
NSBH 0%

BNS 0%

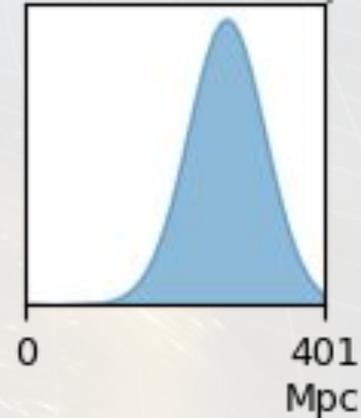
BBH 0%

S190814bv (Definite GW)

After follow-up, was reclassified as NSBH



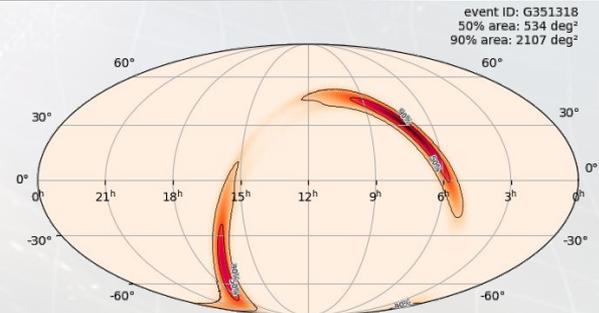
distance: 267 ± 52 Mpc



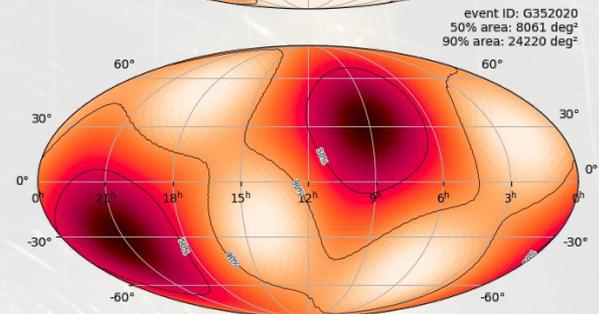
NSBH	>99%
MassGap	<1%
Terrestrial	0%
BNS	0%
BBH	0%

Other marginal NSBH : S190923y, S190930t, S190910d

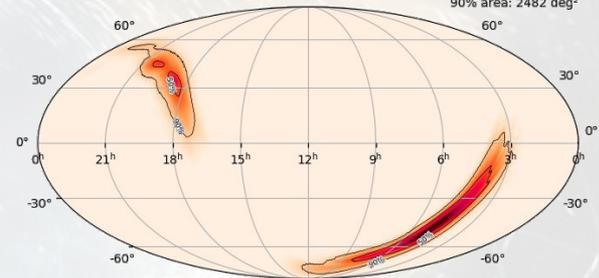
S190923y NSBH (68%), Terrestrial (32%)



S190930t NSBH (74%), Terrestrial (26%)



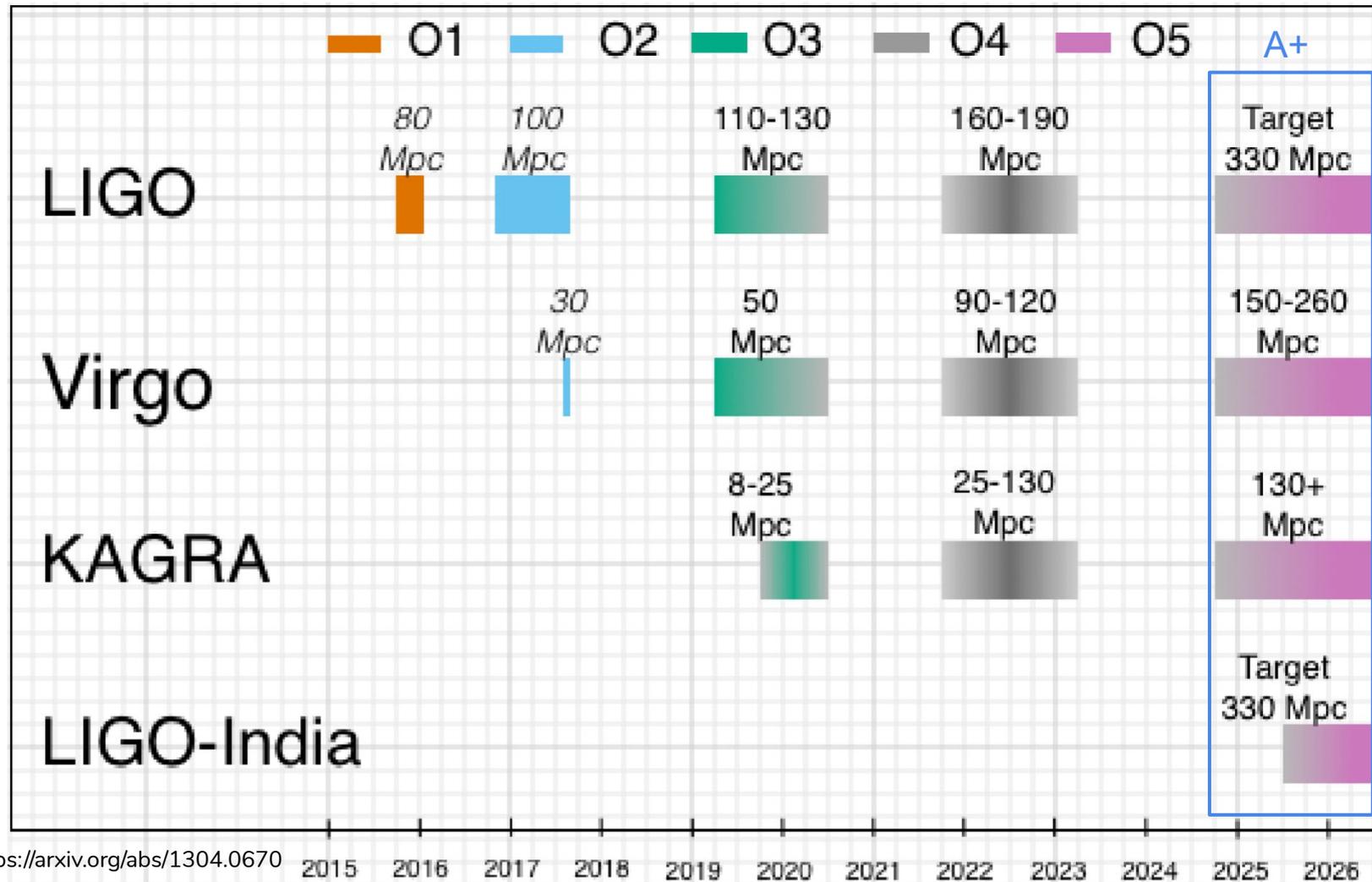
S190910d NSBH (98%), Terrestrial (2%)

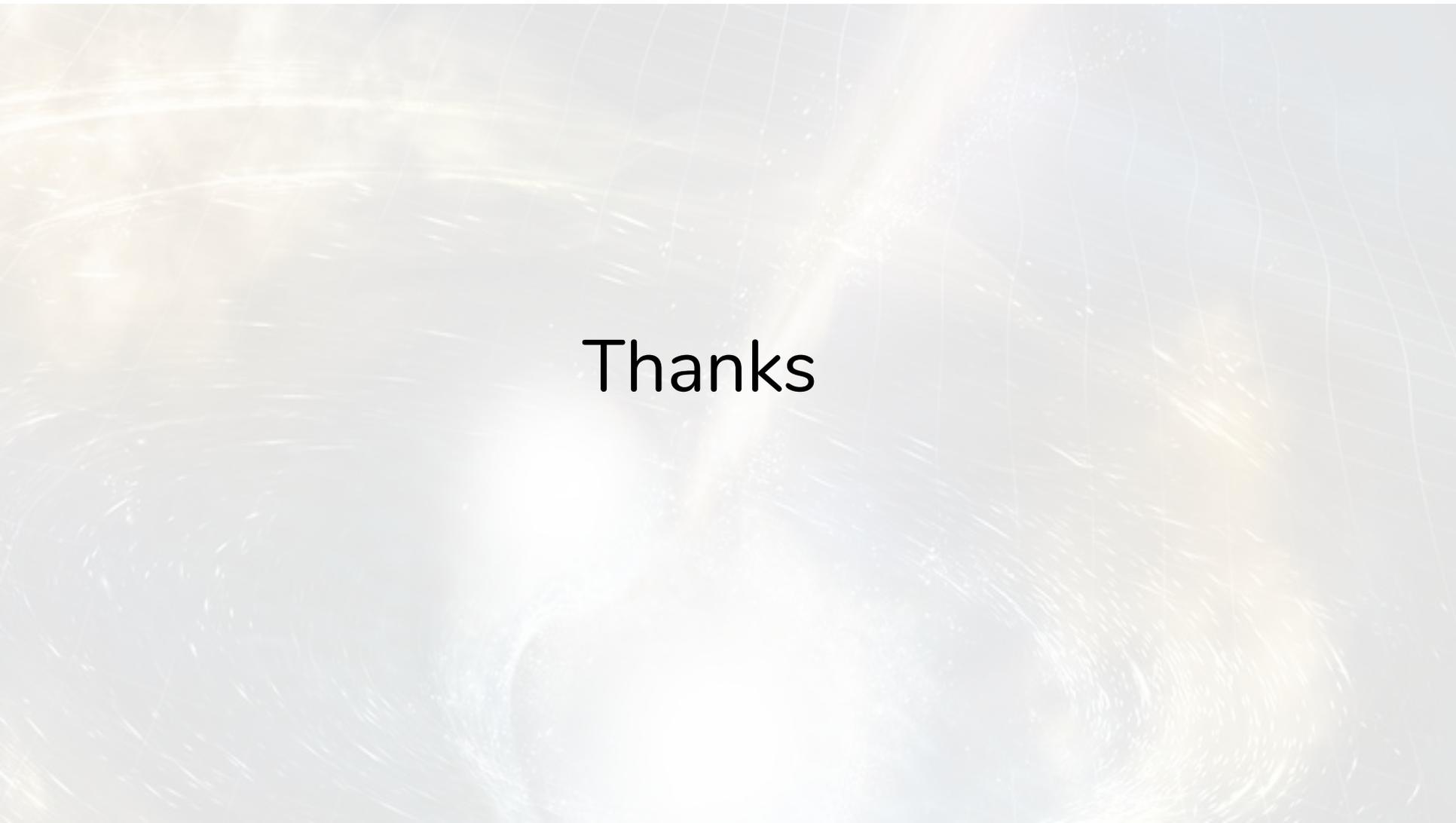


Going forward

O3b will begin on November 1, 2019 and run until April 30, 2020.

We are on track for > 50 confident GW detections by the end of the run.



The background features a complex, abstract design. It consists of numerous thin, light-colored lines that form a grid-like pattern, overlaid with dynamic, swirling light trails in shades of yellow and white. These trails create a sense of motion and depth, resembling a digital or particle-based environment. The overall color palette is soft and ethereal, with a mix of pale blues, greys, and warm yellows.

Thanks

