

Applications of chiral nuclear forces up to N³LO to nuclear matter and neutron stars

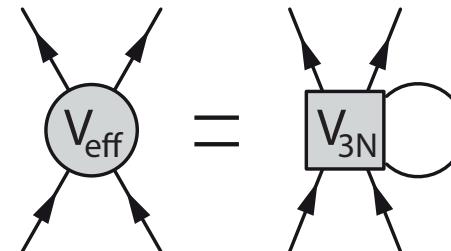
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0 Improved Normal-Ordering Method

- » include dominant 3N contributions
- » effective NN potentials – now at N³LO
- » applicable to all nuclear forces



towards *consistent* N³LO calculations

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1 Isospin-Asymmetric Matter

- » MBPT calculation of 11 proton fractions
- » N³LO NN and N²LO 3N forces
- » constrains astrophysical quantities

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2 Many-Body Convergence ?

- » N³LO NN+3N forces beyond HF
- » MBPT (third order) vs. SCGF method
- » normal-ordering at finite-temperatures

arXiv:1608.05615 (PRC in press)

3 BCS Pairing Gap

- » neutron matter, 1S_0 and $^3P_2 - ^3F_2$
- » new uncertainties
- » new (semi-)local NN potentials

(arXiv:1610.05213)