# Gravitational-wave signal from binary neutron stars: a systematic analysis of the spectral properties

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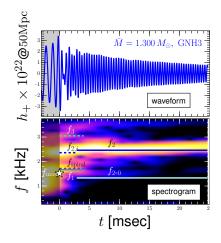
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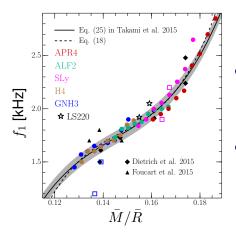
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#### **GWs from HMNSs**



- There is no doubt that GWs from HMNSs have a lot of important information, and the a series of advanced ground-based detectors such as a LIGO will be able to receive it within a few yeas.
- They are including many typical frequencies, which are named  $f_{\text{max}}$ ,  $f_1$ ,  $f_2$ ,  $f_3$ ,  $f_{2-0}$ ,  $f_{\text{spiral}}$  and so on (see the left figure).
- In our previous work, therefore, we had the systematic investigation by a large number of numerical-relativity simulations of binaries with nuclear EOSs.

#### **GWs from HMNSs**



- Indeed, the powerful correlations between the frequencies and physical quantities of the binary system were shown, *e.g.*, the figure show the universal relation between  $f_1$  and stellar compactness  $\overline{M}/\overline{R}$ .
- However, Bauswein et al. (2015) claimed "such a mass-independent, universal relation does not exist", when they considered the wide mass range  $\bar{M} = 1.2 - 1.5 M_{\odot}$ .
- In this study, we extend our previous work by using the wide mass range and additional analyses. Then the universal relation is still robust even for high- or low-mass binaries. We also realize the reason why the mismatch found in Bauswein et al. (2015) arose.

#### **GWs from HMNSs**

