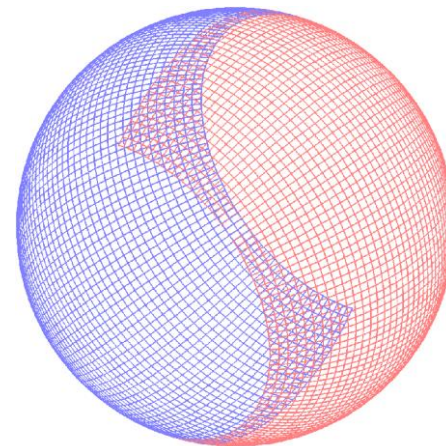
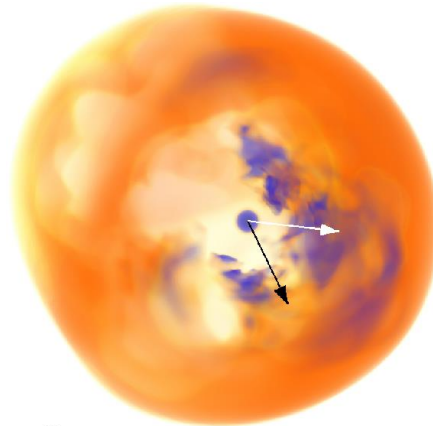
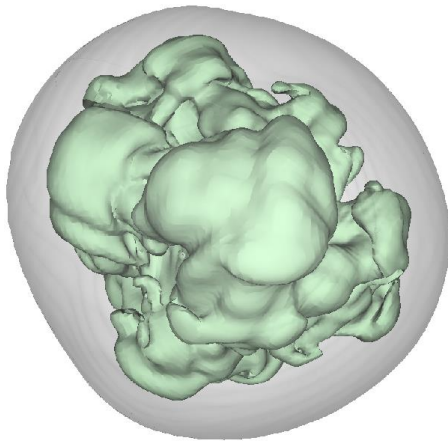
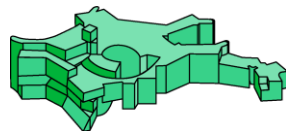


# Neutron star kicks by gravitational tug-boat mechanism

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für Astrophysik



Astrophysical Big Bang  
Laboratory



1 Nov 2016

Neutron star kicks

Kick mechanism??

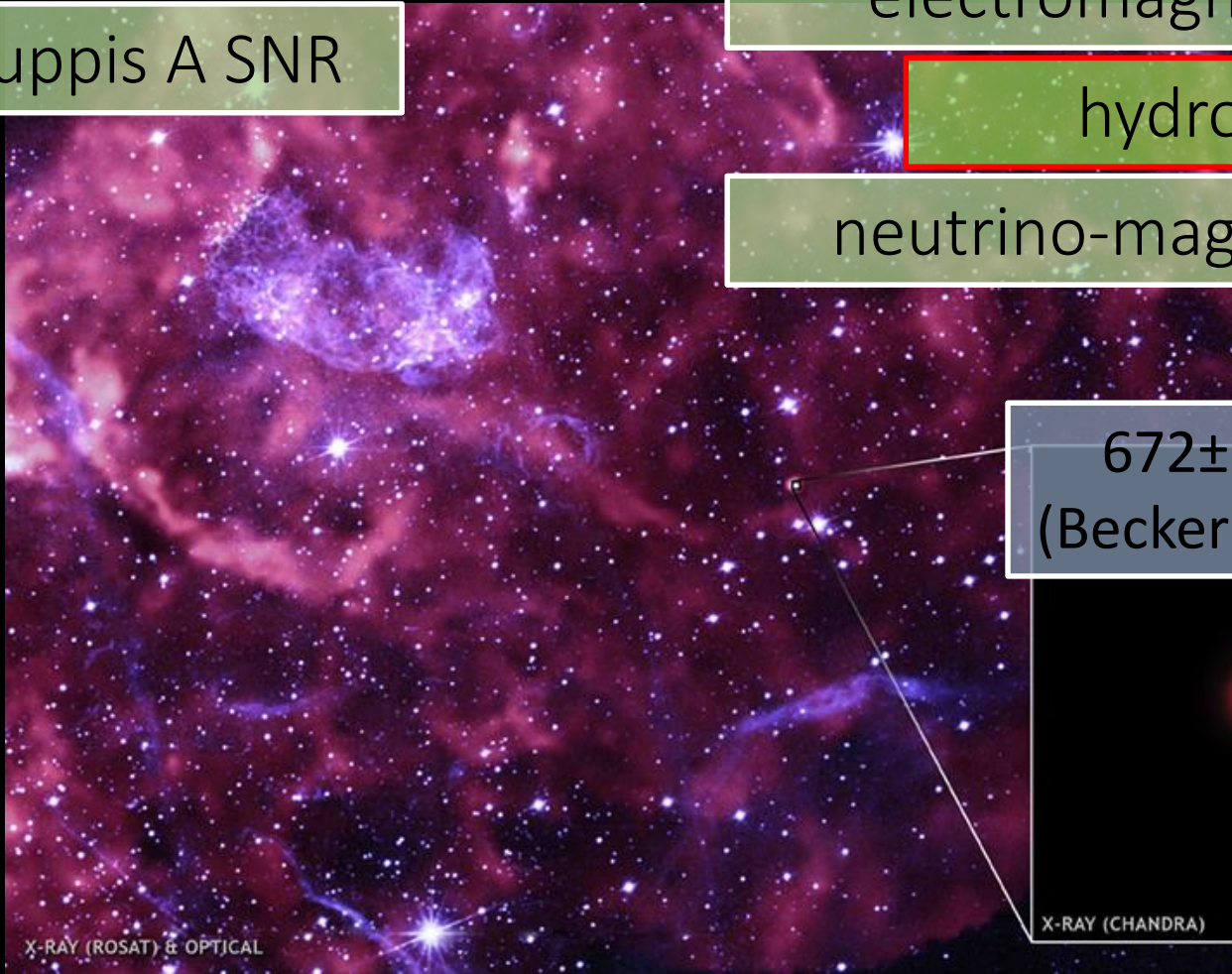
Puppis A SNR

electromagnetic

hydrodynamic

neutrino-magnetic

$672 \pm 115$  km/s  
(Becker et al. 2012)



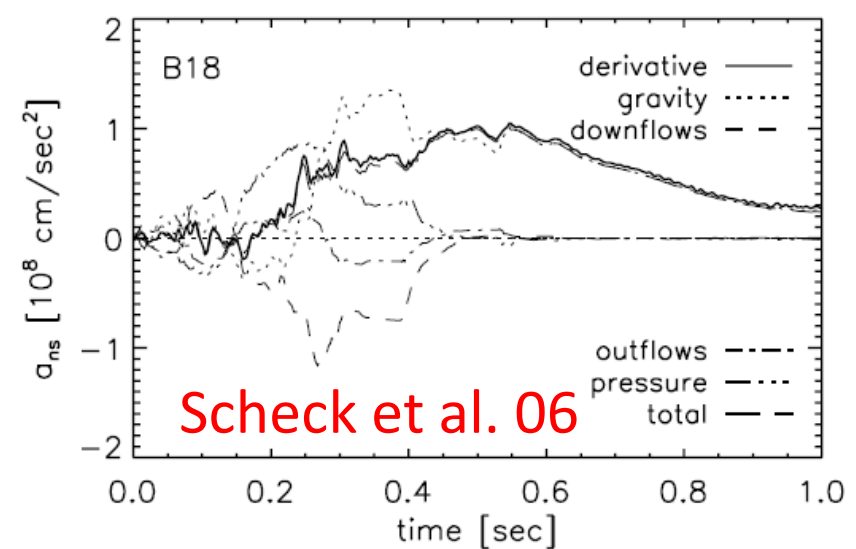
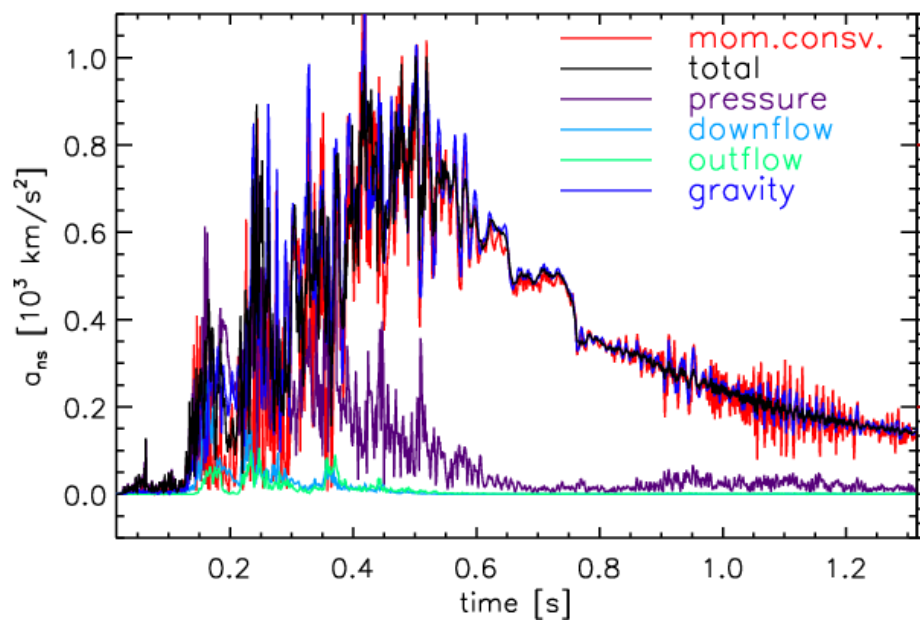
average pulsars velocity: 200-500 km/s

Model	$M_{\text{ns}}$ [ $M_{\odot}$ ]	$t_{\text{exp}}$ [ms]	$E_{\text{exp}}$ [B]	$v_{\text{ns}}$ [ $\text{km s}^{-1}$ ]	$a_{\text{ns}}$ [ $\text{km/s}^2$ ]	$v_{\text{ns},\nu}$ [ $\text{km s}^{-1}$ ]	$\alpha_{\text{kv}}$ [ $^{\circ}$ ]	$v_{\text{ns}}^{\text{long}}$ [ $\text{km s}^{-1}$ ]	$a_{\text{ns}}^{\text{long}}$ [ $\text{km/s}^2$ ]	$J_{\text{ns},46}$ [ $10^{46} \text{ g cm}^2/\text{s}$ ]	$\alpha_{\text{sk}}$ [ $^{\circ}$ ]	$T_{\text{spin}}$ [ms]
W15-1	1.37	246	1.12	331	167	2	151	524	44	1.51	117	652
W15-2	1.37	248	1.13	405	133	1	126	575	49	1.56	58	632
W15-3	1.36	250	1.11	267	102	1	160	–	–	1.13	105	864
W15-4	1.38	272	0.94	262	111	4	162	–	–	1.27	43	785
W15-5-lr	1.41	289	0.83	373	165	2	129	–	–	1.63	28	625
W15-6	1.39	272	0.90	437	222	2	136	704	71	0.97	127	1028
W15-7	1.37	258	1.07	215	85	1	81	–	–	0.45	48	2189
W15-8	1.41	289	0.72	336	168	3	160	–	–	4.33	104	235
L15-1	1.58	422	1.13	161	69	5	135	227	16	1.89	148	604
L15-2	1.51	382	1.74	78	14	1	150	95	4	1.04	62	1041
L15-3	1.62	478	0.84	31	27	1	51	–	–	1.55	123	750
L15-4-lr	1.64	502	0.75	199	123	4	120	–	–	1.39	93	846
L15-5	1.66	516	0.62	267	209	3	147	542	106	1.72	65	695
N20-1-lr	1.40	311	1.93	157	42	7	118	–	–	5.30	122	190
N20-2	1.28	276	3.12	101	12	4	159	–	–	7.26	43	127
N20-3	1.38	299	1.98	125	15	5	138	–	–	4.42	54	225
N20-4	1.45	334	1.35	98	18	1	98	125	9	2.04	45	512
B15-1	1.24	164	1.25	92	16	1	97	102	1	1.03	155	866
B15-2	1.24	162	1.25	143	37	1	140	–	–	0.12	162	7753
B15-3	1.26	175	1.04	85	19	1	24	99	3	0.44	148	2050

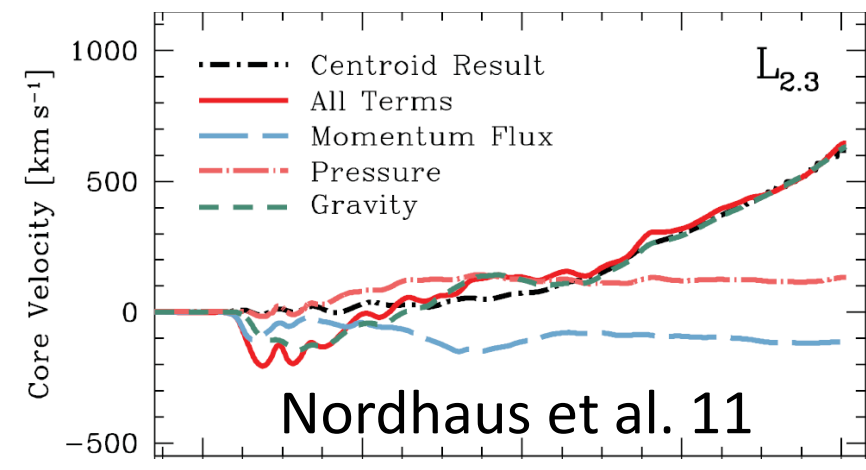
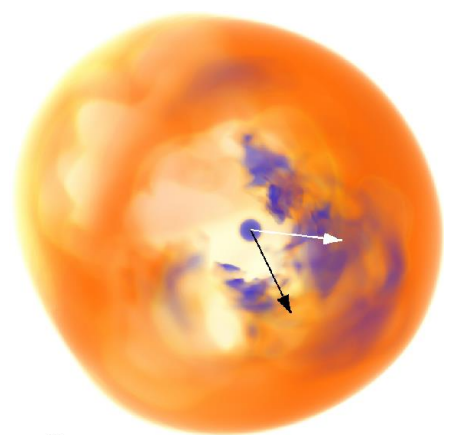
$v_{\text{ns}}$  from 31-437 km/s

700 km/s later

$$\vec{P}_{ns} \approx - \oint_{r=r_0} P d\vec{S} - \oint_{r=r_0} \rho \vec{v} v_r dS + \int_{r>r_0} \frac{GM_{ns} \vec{r}}{r^3} dm$$



Scheck et al. 06



Nordhaus et al. 11

gravitational drag term is dominant