2016. 11. 2 Compact Stars and Gravitational Waves

### Strangeness Nuclear Physics and Neutron Stars



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# **1. Introduction**



### **Exotic flavors in NS?**

### Attractive AN interaction and hyperon mixing



### Baryon fractions in NS is sensitive to BB interactions



We need all the YN, YY interactions in free space and in nuclear matter (at high density...)

### Experimental status of YN, YY (+K<sup>bar</sup>N) interactions

Poor  $\Lambda p$ ,  $\Sigma^{\pm}p$  scattering data + <u>Various hypernuclear data</u>

Experimentally known:



- ΣN int. looks strongly repulsive. Σ
   How large repulsion?
   -> No Σ's in NS?
   Σ<sup>±</sup>-p scattering
   Λ-p scattering
- EN int. is unknown. Attractive or repulsive? -> E's in NS?
- A A int. weakly attractive (B( $\Lambda\Lambda$ ) =0.7 MeV). Need more data to confirm H dibaryon above  $\Lambda\Lambda$  threshold?
- K<sup>bar</sup>N interaction strongly attractive. How large attraction in nuclei? -> Kaon condensation in NS?

Nijmegen ESC08 model reproduces most of the existing data well. Also important to test Lattice QCD calculations.



**H** dibaryon

Ξ hypernuclei



**ΛΛ hypernuclei** 

**Experiments running or planned** 



# 2.1 S = -1 Systems $4H^{OO}$ $4H^{OO}$ $4H^{OO}$ AN interaction in neutron matter

Charge symmetry breaking in  $\Lambda$  hypernuclei

Neutron–rich  $\Lambda$  hypernuclei

Isotope effect of  $\Lambda$  binding energies

### <u>ΛN interaction in neutron stars?</u>

A has no isospin =>  $\Lambda$ N interaction should be the same in symmetric nuclear matter and in pure neutron matter



We need to study

- Charge symmetry breaking in A hypernuclei
- **Neutron-rich**  $\Lambda$  hypernuclei
- Isospin dependence of  $\Lambda$  binding energy ( $\Lambda$ nn force)

### <u>Charge Symmetry Breaking (CSB)</u> <u>in $\Lambda$ hypernuclei?</u>



### **Decay-pion measurement at Mainz**



### <u>γ spectroscopy</u> of hyeprnuclei











### <u>Charge Symmetry Breaking (CSB)</u> in A=4 hypernuclei



Existence of a large CSB effect confirmed <u>only by γ-ray data</u>

- **B**<sub> $\Lambda$ </sub> [  ${}^{4}_{\Lambda}$ H(0<sup>+</sup>) ] confirmed via  ${}^{4}_{\Lambda}$ H ->  ${}^{4}$ He +  $\pi^{-}$  decay , suggesting the emulsion  ${}^{4}_{\Lambda}$ He(0<sup>+</sup>) data also reliable
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### What is the origin of the large CSB effect?

u/d quark mass difference + EM effects => CSB in hadrons and hadron-hadron interactions

\*  $\Sigma^+\Sigma^-$  mass difference + CSB in BB forces (Nijmegen SC97e) =>  $\Delta B_{\Lambda}(0^+) \sim 70$  keV at maximum.

Nogga et al., PRL 88 (2002) 172501

Ν

Ν

π

 $\Sigma^0$ 

Ν

Λ

\* Shell model calc. using D2 =>  $\Delta B_{\Lambda}(0^{+}) \sim 200 \text{ keV}$ . A. Gal, PLB 744 (2015) 352 D2: central-only  $\Lambda \Sigma$  coupling Akaishi et al., PRL 84 (2000) 3539 (SC: tensor dominated  $\Lambda \Sigma$  coupling) N  $\Lambda \frac{\Lambda \Sigma}{coupling}$ 

\* Ab initio calc. with Bonn-Juelich EFT force (LO) =>  $\Delta(B_{\Lambda}(0^+) - \Delta B_{\Lambda}(1^+)) \sim 0.3$  MeV. = central dominated  $\Lambda \Sigma$  coupling

D. Gazda and A. Gal, PRL 116 (2016) 122501

# The CSB effect is sensitive to $\Lambda N-\Sigma N$ coupling.

CSB effects in p-shell hypernuclei will confirm the origin.

# Neutron-rich hypernuclei

- Strong mixing of  $\Lambda N$ - $\Sigma N$
- Scrong mixing of AN-210 B.F. Gibson et al. PRC6 (1972) 741, etc
   Coherent effect in proton/neutron-rich nuclei

Akaishi et al. PRL 84 (2000) 3539



Larger mixing in a host nucleus with larger I

=> How large mixing in n-rich hypernuclei?







=> Effect to  $\Lambda$  appearance in n star?



$${}^{\underline{6}}_{\underline{\Lambda}}$$
H is bound?

#### $FINUDA@DA\Phi NE$

M. Agnello et al., PRL 108 (2012) 042501

$$K^{-}(stop) + {}^{6}Li -> {}^{6}_{\Lambda}H + \pi^{-}$$

 ${}^{6}{}_{\Lambda}H \rightarrow {}^{6}He + \pi^{-}$ 

3 events of bound  ${}^{6}_{\Lambda}$ H reported.





#### E10@J-PARC

Sugimura et al. PLB 729 (2014) 39

$$\pi^{-}$$
+ <sup>6</sup>Li -> <sup>6</sup><sub>A</sub>H + K<sup>+</sup>

( $\pi^- p \rightarrow \Sigma^- K^+$ ,  $\Sigma^- p \rightarrow \Lambda n$  etc.)

No  ${}^{6}_{\Lambda}$ H peak observed.



### <u>Ann bound state??</u>

# HypHI@GSIC. Rappold et al.PHYSICAL REVIEW C 88, 041001(R) (2013)

Search for evidence of  ${}^{3}_{\Lambda}n$  by observing  $d + \pi^{-}$  and  $t + \pi^{-}$  final states in the reaction of  ${}^{6}\text{Li} + {}^{12}\text{C}$  at 2A GeV



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### <u>(e,e'K<sup>+</sup>) spectroscopy</u>

### for isospin dependence in $\Lambda$ binding energy

Compare  $\frac{40}{Ca} (e,e'K^+) \frac{40}{\Lambda} K$  and  $\frac{48}{Ca} (e,e'K^+) \frac{48}{\Lambda} K$ 



Figure 2-8: Latest results of  $\Lambda$  separation energies for  ${}^{40}_{\Lambda}$ K and  ${}^{48}_{\Lambda}$ K as a function of C<sub>T</sub> calculated with AFDMC. Errors are statistical errors in Monte Carlo calculation and future study will make them smaller.

S.N. Nakamura (Tohoku) et al.

AFDMC calc. by D.Lonardoni, S.Gandolfi et al.

The experiment has been approved.

# 2.2 S = -1 Systems K<sup>-</sup>pp and K<sup>bar</sup>N interaction



# K<sup>bar</sup>-N interaction and K<sup>bar</sup> nucleus

- **KN** interaction
  - Known to be strongly attractive from K<sup>-</sup>p atomic X-ray shift and low energy K<sup>-</sup> p scattering data
  - $\Lambda(1405)$  (1/2<sup>-</sup>) can be interpreted as a K<sup>-</sup>p bound state
- K<sup>-</sup>pp bound state KNN (I=1/2) The simplest Kbar nucleus
  - Theoretical prediction of B.E. and Γ depend on the KN interaction and theoretical framework.
- K in matter
  - Clarify possible existence of K<sup>-</sup> condensation in neutron stars
  - Extremely high density nuclei can be produced?

Calculated  $K^- pp$  binding energies B and widths  $\Gamma$  (in MeV).

A. Gal / Nuclear Physics A 914 (2013) 270-279

	Chiral, energy dependent			Non-chiral, static calculations			
	var. [7]	var. [8]	Fad. [9]	var. [10]	Fad [11]	Fad [12]	var. [13]
В	16	17-23	9–16	48	50-70	60-95	40-80
Г	41	40-70	34-46	61	90-110	45-80	40-85

[7] N. Barnea, A. Gal, E.Z. Liverts, Phys. Lett. B 712 (2012) 132.

[8] A. Doté, T. Hyodo, W. Weise, Nucl. Phys. A 804 (2008) 197;

A. Doté, T. Hyodo, W. Weise, Phys. Rev. C 79 (2009) 014003.

[9] Y. Ikeda, H. Kamano, T. Sato, Prog. Theor. Phys. 124 (2010) 533.

[10] T. Yamazaki, Y. Akaishi, Phys. Lett. B 535 (2002) 70.

[11] N.V. Shevchenko, A. Gal, J. Mareš, Phys. Rev. Lett. 98 (2007) 082301;

N.V. Shevchenko, A. Gal, J. Mareš, J. Revai, Phys. Rev. C 76 (2007) 044004.

[12] Y. Ikeda, T. Sato, Phys. Rev. C 76 (2007) 035203;
 Y. Ikeda, T. Sato, Phys. Rev. C 79 (2009) 035201.

[13] S. Wycech, A.M. Green, Phys. Rev. C 79 (2009) 014001.

# **Previous positive data for K<sup>-</sup>pp**

	FINUDA	DISTO		
reaction	Stopped K <sup>-</sup> absorption on <sup>6, 7</sup> Li+ <sup>12</sup> C	p + p @ Tp=2.85GeV		
method	Invariant mass of back-to-back Ap pairs	$p+p \rightarrow X+K^+$ (missing mass) $X \rightarrow \Lambda+p$ (invariant mass)		
B.E	$115^{+6}_{-5}(stat)^{+3}_{-4}(syst)$ MeV	$105 \pm 5 \text{ MeV}$		
Width	$67^{+14}_{-11}(stat)^{+2}_{-3}(syst) MeV$	118 ± 8 MeV		
-250 -200 -150 25 25 20 25 20 15 20 15 25 20 15 25 20 15 25 20 15 25 20 15 25 20 15 25 20 15 25 20 15 25 20 15 25 20 15 25 25 20 15 25 20 15 25 20 15 25 20 15 20 15 20 15 20 15 20 15 20 15 20 15 20 15 20 15 20 15 20 15 20 15 20 15 20 15 15 20 15 20 15 15 20 15 15 20 15 20 15 15 20 20 15 15 20 15 20 15 15 20 20 15 15 20 20 15 15 20 20 20 15 15 20 20 20 15 15 20 20 20 20 20 20 20 20 20 20	M.Agnello <i>et al.</i> , PRL 94, 212303 (2005) $M(K^{-}+p+p) - 2.37 \text{ GeV/C}^{2} - 9 - 9 - 100 - 50 - 0 - 100 - 50 - 0 - 100 - 50 - 0 - 100 - 50 - 0 - 100 - 50 - 0 - 100 - 50 - 0 - 100 - 50 - 0 - 100 - 50 - 0 - 100 - 50 - 0 - 0 - 0 - 0 - 0 - 0 - 0 - 0 $	T.Yamazaki <i>et al.</i> , PRL 104, 132502 (2010) B (K <sup>-</sup> pp) [GeV] M(K <sup>-</sup> +p+p) <sub>-0.1</sub> <i>large-angle proton</i> M = 2.267 (2) M =		

# <u>K<sup>-</sup>pp search via d(π<sup>+</sup>,K<sup>+</sup>)</u>

 $d(\pi^+, K^+)X$  reaction ( $P_{\pi} = 1.7 GeV/c$ )

K<sup>-</sup>pp is produced via a  $\Lambda(1405)$  doorway.

 $\pi^{+} + n \rightarrow \Lambda(1405) + K^{+}$  $\Lambda(1405) + p \rightarrow K^{-}pp$  $(\rightarrow quasi free \Lambda^{*})$ 



Y.Akaishi, T.Yamazaki, Phys. Rev. C 76 045201 (2007)

# Coincidence with protons

 $\frac{d(\pi^+, K^+) K^- pp}{K^- pp} \rightarrow \Lambda + p$   $\downarrow p \downarrow p + \gamma$ 



#### Slide by Y. Ichikawa





## <u>K-pp search via d( $\pi^+$ ,K+)</u>

### J-PARC E27 Nagae et al.

1.69 GeV/c pion beam, n ( $\pi^+$ ,K<sup>+</sup>)  $\Lambda$ (1405),  $\Lambda$ (1405) + p -> K<sup>-</sup>pp



Y. Ichikawa et al., PTEP 2015 (2015) 021D01

# Another K<sup>-</sup>pp search exp.

#### J-PARC E15 Iwasaki et al.



# <u>The latest result: <sup>3</sup>He(K<sup>-</sup>,Λp)n</u>



Slide by T.Yamaga

# 3. S = -2 Systems Ξ hypernuclei



#### Slide by Nakazawa

### <u>ΛΛ hypernuclei from emulsion (KEK E373)</u>

#### Nagara event



# "Kiso event"

### found by overall scanning method (E373)

K. Nakazawa et al. PTEP 2015, 033D02



 $B_{\Xi^-} = 4.38 \pm 0.25 \text{ MeV} - 1.11 \pm 0.25 \text{ MeV}$  $^{10}_{\Lambda}$ Be production: in the ground state in the highest excited state >>: 3D atomic state of the  $\Xi^{-14}$ N system (0.17 MeV) First evidence for a deeply bound  $\Xi$  state =  $\Xi$ -N attractive  $\rightarrow \Xi^-$  can exist in neutron stars

## J-PARC E07: S=-2 systems by emulsion



 $\Lambda \rightarrow N\pi$ ,  $\Lambda N \rightarrow NN$ 

The experiment just started this spring.

### **E-atomic X-rays via "Hyperball-X" (Ge array)**



### <u>Spectroscopy of E-hypernuclei via (K<sup>-</sup>,K<sup>+</sup>) reaction</u>

**J-PARC E05** 

Nagae et al.

- Discovery of Ξ-hypernuclear states as a peak(s)
- Measurement of Ξ-nucleus potential depth and width
- Coupling between Ξ-nucleus and ΛΛ-nucleus



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J-PARC E05 Nagae et al.

Slide by T.Nagae

### Pilot run (2015) using the existing SKS spectrometer



 $\Delta E \sim 5.4$  MeV FWHM: much better than the previous BNL exp (~14 MeV)

# 4. Hyperon puzzle and future prospects for ΛNN interaction

J-PARC Hadron Hall Extension Plan

### Hyperon puzzle

Hyperons (at least  $\Lambda$ 's) must appear at  $\rho = 2 \sim 3 \rho_0$  2 reliable samples EOS with hyperons (or kaons) too soft to support  $\langle \rho \rangle$  of  $\sim 2.0M_{\odot}$  NS

heavy ( >1.5  $M_{sun}$ ) NS's in "standard" frameworks

=> Unknown repulsion should exist at high  $\rho$ 



<u>Various models to support  $2 M_{sun} NS$  with hyperons</u>

- Introduce strong repulsion in 3-body force (NNN, YNN, YYN, YYY)
- Quark Meson Coupling model
- Density dependence in coupling const., hadron mass, etc.
- Relativistic framework
- Lattice QCD
- Phase transition to quark matter = quark star or "crossover"

YN strong repulsion at high density really exists ? Any experimental evidence for that?

### **3-body nuclear force**



3-body force from chiral effective field theory

G. Hagen et al., PRL 109 (2012) 032502



### Density dependence of $\Lambda N$ interaction affects $\Lambda$ hypernuclear data?

We have almost no information on BB forces in high density ( $\rho > \rho_0$ ) matter Ab-initio calc. of nuclear binding energies => NNN repulsion necessary Similar YNN (YYN, YYY) repulsive forces? How large are they?









# <u>Summary</u>

YN, YY interactions in free space and in nuclear (neutron) matter should be studied to understand the inner core of NS.

S=-1

- A larger charge symmetry breaking effect in ΛN int. is confirmed for A=4 hypernuclei via γ-ray measurement.
- Exciting data for n-rich hypernuclei ( ${}^{6}_{\Lambda}H$ ,  $\Lambda$ nn) should be confirmed.
- New data for K<sup>-</sup>pp appeared, further investigation necessary.

### S=-2

- A Ξ-nuclear bound system was discovered in emulsion, suggesting Ξ to appear in NS. Ξ-atomic X-rays will be also measured.
- <sup>12</sup><sub>Ξ</sub>Be hypernuclear state(s) were also observed in (K-,K+) reaction, indicating a rather deep Ξ-nuclear potential.

### Future

A's binding energies in wide mass numbers will be precisely measured at Jlab and J-PARC to approach the  $\Lambda$ NN force.