



A model with radiative electroweak symmetry breaking and dark matter

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1. Introduction

Higgs sector in the Standard Model

- “Higgs” boson was discovered in 2012.
- Its properties are consistent with the Higgs boson in SM.
VEV, mass, spin, parity, some couplings etc.

$$V_H^{\text{SM}} = -\mu_H^2 H^\dagger H + \lambda_H (H^\dagger H)^2$$

$$H = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v_H + h \end{pmatrix} \quad v_H = 246 \text{ GeV}$$

VEV (unitary gauge) h : Higgs boson field

Higgs potential of SM

EWSB →

$$\text{const.} + \lambda_H v_H^2 h^2 + \lambda_H v_H h^3 + \frac{\lambda_H}{4} h^4$$

mass

$$m_h = \sqrt{2\lambda_H v_H} \simeq 125 \text{ GeV}$$

unmeasured

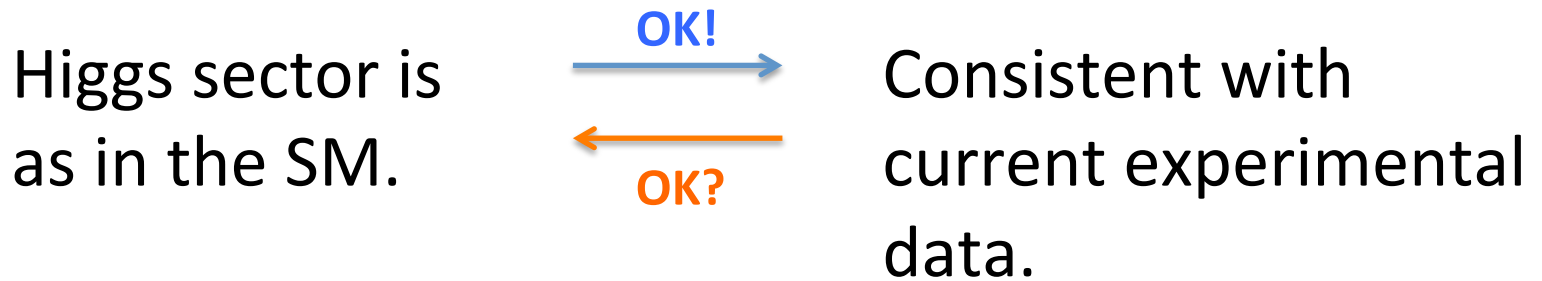
triple coupling

quartic coupling

- Measurements of Higgs self couplings are necessary to determine the Higgs potential.

Motivation

Current situation



- Is there possibility that ONLY self couplings deviate (substantially) from SM predictions?
- We are interested in information about the vicinity of the vacuum (“shape” of potential).

Motivation

- One candidate: Scale-invariant extension
 - EWSB via CW mechanism (radiative symmetry breaking)
 - Effective potential is like $\phi^4 \log(\phi)$
 - irregular at the origin (e.g. QCD, BCS theory)
 - realise Higgs VEV 246 GeV and mass 125 GeV,
 - large deviations in cubic and quartic self couplings?

We dealt with singlet-extension.

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2. Model

and

Our method of analysis

Model

$$\mathcal{L} = \mathcal{L}_{\text{SM}}|_{\mu_{\text{H}} \rightarrow 0} + \frac{1}{2}(\partial_{\mu} \vec{S})^2 - \lambda_{\text{HS}}(H^{\dagger} H)(\vec{S} \cdot \vec{S}) - \frac{\lambda_{\text{S}}}{4}(\vec{S} \cdot \vec{S})^2$$

H : Higgs doublet

\vec{S} : SM singlet (global $O(N)$ -multiplet)

$$H = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{2}G^{+} \\ v_{\text{H}} + h + iG^0 \end{pmatrix}, \quad \vec{S} = (v_{\text{S}} + s_1 \quad s_2 \quad \cdots \quad s_N)^{\text{T}}$$

K. A. Meissner et.al. [hep-ph/0612165],

R. Foot et. al. [0704.1165]

etc.

- We have re-analysed the model and its effective potential,
- and found the consistent vacuum by more precise perturbation.

(cf. R. Dermisck et.al. [1308.0891], C. Tamarit [1404.7673])

Particle content

$$\mathcal{L} = \mathcal{L}_{\text{SM}}|_{\mu_H \rightarrow 0} + \frac{1}{2}(\partial_\mu \vec{S})^2 - \lambda_{\text{HS}}(H^\dagger H)(\vec{S} \cdot \vec{S}) - \frac{\lambda_S}{4}(\vec{S} \cdot \vec{S})^2$$

H : Higgs doublet

\vec{S} : SM singlet (global $O(N)$ -multiplet)

$$H = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{2}G^+ \\ v_H + h + iG^0 \end{pmatrix}, \quad \vec{S} = (v_S + s_1 \quad s_2 \quad \cdots \quad s_N)^T$$

SM particles

+

N real singlet scalar bosons (s_i)

- $\text{mass}^2 \sim \lambda_{\text{HS}} v_H^2$
- singlet field VEV $\langle S \rangle = 0$ eventually
- thus singlet cannot decay

Effective potential

H : Higgs doublet

\vec{S} : SM singlet (global $O(N)$ -multiplet)

$$\langle H \rangle_J = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ \phi \end{pmatrix}, \quad \langle \vec{S} \rangle_J = (\varphi, 0, \dots, 0); \quad \phi, \varphi \in \mathbb{R}, \quad \phi, \varphi \neq 0$$

$$V_{\text{eff}}(\phi, \varphi) = V_{\text{tree}}(\phi, \varphi) + V_{1\text{-loop}}(\phi, \varphi)$$

$$V_{\text{tree}}(\phi, \varphi) = \frac{\lambda_H}{4} \phi^4 + \frac{\lambda_{HS}}{2} \phi^2 \varphi^2 + \frac{\lambda_S}{4} \varphi^4$$

$$V_{1\text{-loop}}(\phi, \varphi) = \sum_i \frac{n_i}{64\pi^2} M_i^4(\phi, \varphi) \left[\ln \frac{M_i^2(\phi, \varphi)}{\mu^2} - c_i \right] (\overline{\text{MS}} \text{ scheme})$$

$M_i^2(\phi, \varphi)$: mass-squared eigenvalues \leftarrow determined in tree level

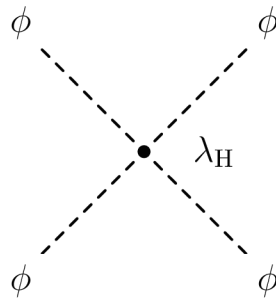
μ : renormalisation scale \rightarrow set to Higgs VEV 246GeV

c_i : scheme-dependent constants

Order counting in perturbation

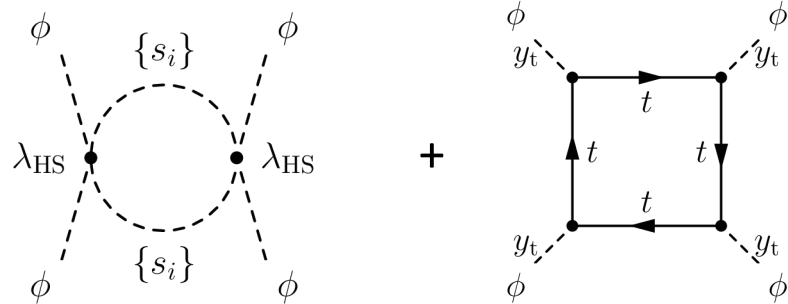
It is necessary that tree and 1-loop effects balance each other to occur correctly EWSB via CW mechanism.

tree effects



VS

1-loop effects



$$|\lambda_H| \sim \frac{N \lambda_{HS}^2}{(4\pi)^2} - \frac{N_C y_t^4}{(4\pi)^2} \ll 1 \quad (\text{massless scalar QED: } \lambda \sim \frac{e^4}{(4\pi)^2} \ll 1)$$

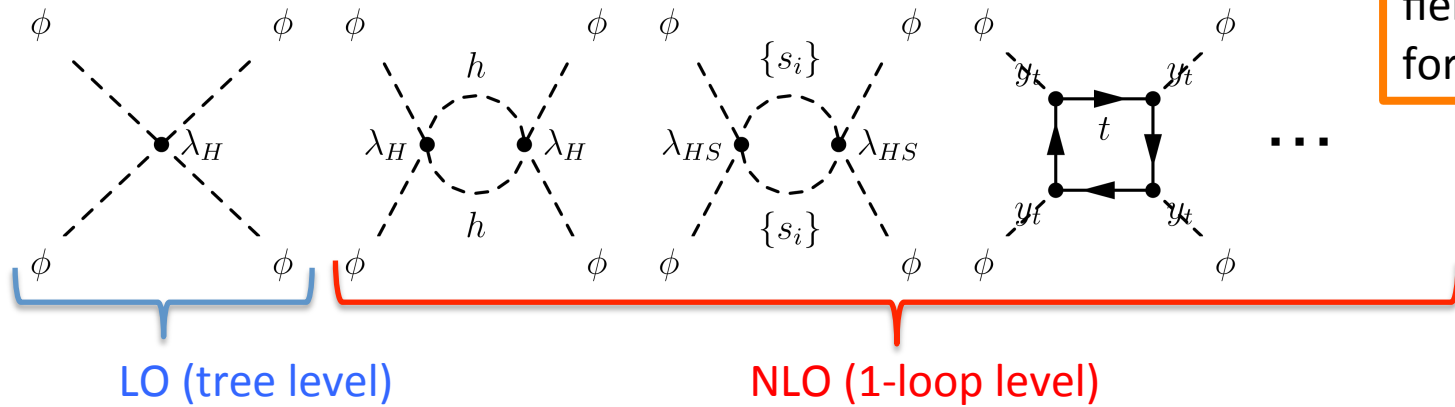
(Roughly $|\lambda_H| \ll |\lambda_{HS}|$. N_C is the colour factor.)



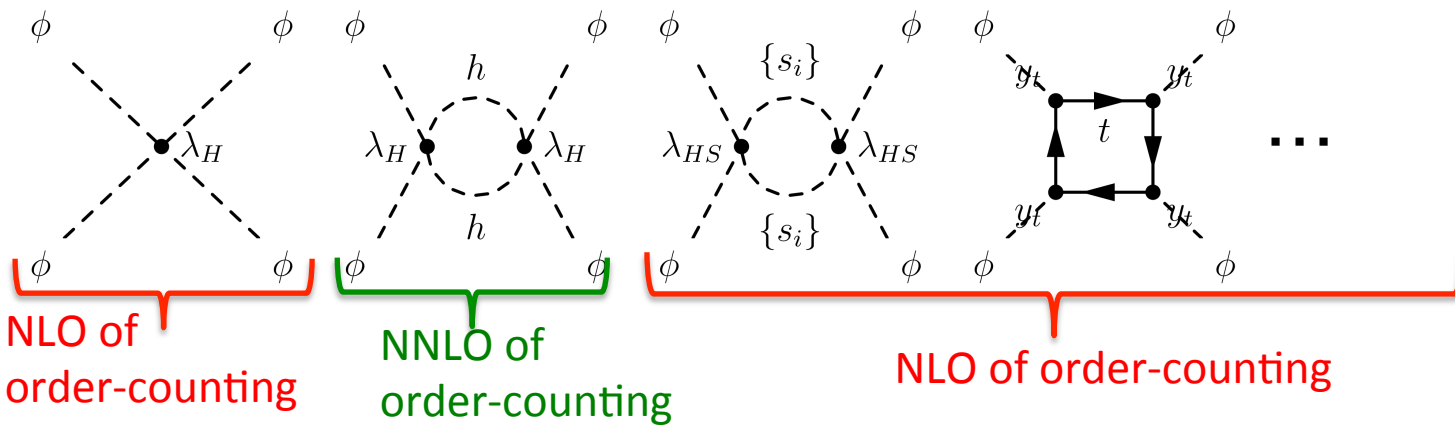
Using this relation gives us specific order-counting.

Order counting in perturbation

[Usual perturbation]




[This prescription]



Order counting in perturbation

[For the tree potential]

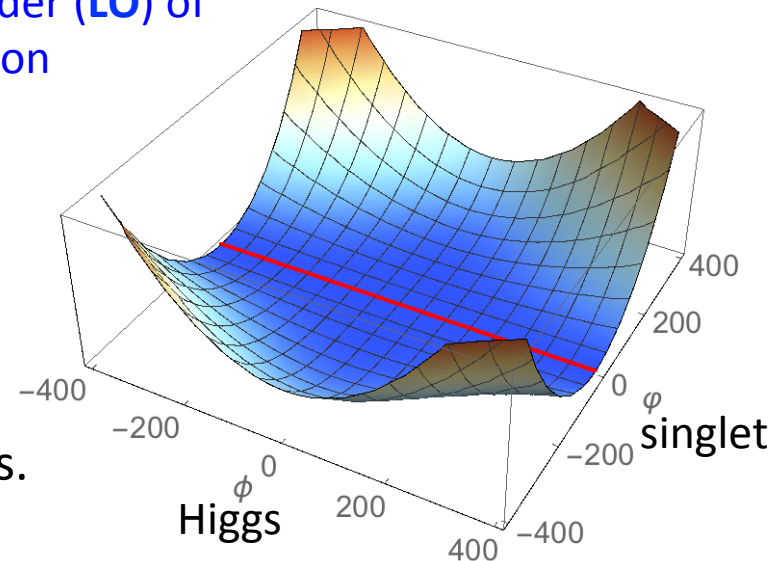
$$|\lambda_H| \ll |\lambda_{HS}|$$


$$\frac{\lambda_H}{4} \phi^4 + \frac{\lambda_{HS}}{2} \phi^2 \varphi^2 + \frac{\lambda_S}{4} \varphi^4$$

next-to-leading order (**NLO**)
of perturbation


leading order (**LO**) of
perturbation

- The **LO** potential is $\frac{\lambda_{HS}}{2} \phi^2 \varphi^2 + \frac{\lambda_S}{4} \varphi^4$
- and it has degenerate minima on the ϕ -axis.



Order counting in perturbation

[For the 1-loop corrections] $\frac{1}{64\pi^2} \text{STr} M^4(\phi) \left[\ln \left(\frac{M^2(\phi)}{\mu^2} \right) - C \right]$

$|\lambda_H| \ll |\lambda_{HS}|$ 

- $\frac{1}{64\pi^2} (\lambda_H \phi^2 + \lambda_{HS} \varphi^2)^2 \left[\ln \left(\frac{\lambda_H \phi^2 + \lambda_{HS} \varphi^2}{\mu^2} \right) - \frac{3}{2} \right]$

- $\simeq \frac{1}{64\pi^2} (\lambda_{HS} \varphi^2)^2 \left[\ln \left(\frac{\lambda_{HS} \varphi^2}{\mu^2} \right) - \frac{3}{2} \right]$

- $\frac{1}{64\pi^2} F_{\pm}^2(\phi, \varphi) \left[\ln \left(\frac{F_{\pm}(\phi, \varphi)}{\mu^2} \right) - \frac{3}{2} \right] \simeq \frac{1}{64\pi^2} F_{\pm \text{app}}^2(\phi, \varphi) \left[\ln \left(\frac{F_{\pm \text{app}}(\phi, \varphi)}{\mu^2} \right) - \frac{3}{2} \right],$

where

$$F_{\pm}(\phi, \varphi) \simeq F_{\pm \text{app}}(\phi, \varphi)$$

$$\equiv \frac{\lambda_{HS}}{2} \phi^2 + \frac{\lambda_{HS} + 3\lambda_S}{2} \varphi^2 \pm \sqrt{\left[-\frac{\lambda_{HS}}{2} \phi^2 + \frac{\lambda_{HS} - 3\lambda_S}{2} \varphi^2 \right]^2 + 4\lambda_{HS}^2 \phi^2 \varphi^2}$$

Order counting in perturbation

$$\begin{aligned} V_{\text{LO}} &= \frac{\lambda_{\text{HS}}}{2} \phi^2 \varphi^2 + \frac{\lambda_{\text{S}}}{4} \varphi^4 \\ V_{\text{NLO}} &= \frac{\lambda_{\text{H}}}{4} \phi^4 + \frac{F_{+\text{app}}^2(\phi, \varphi)}{64\pi^2} \left[\ln \left(\frac{F_{+\text{app}}(\phi, \varphi)}{\mu^2} \right) - \frac{3}{2} \right] \\ &\quad + \frac{3}{64\pi^2} (\lambda_{\text{HS}} \varphi^2)^2 \left[\ln \left(\frac{\lambda_{\text{HS}} \varphi^2}{\mu^2} \right) - \frac{3}{2} \right] \\ &\quad + \frac{N-1}{64\pi^2} (\lambda_{\text{HS}} \phi^2 + \lambda_{\text{S}} \varphi^2)^2 \left[\ln \left(\frac{\lambda_{\text{HS}} \phi^2 + \lambda_{\text{S}} \varphi^2}{\mu^2} \right) - \frac{3}{2} \right] \\ &\quad + (\text{vector boson parts}) + (\text{fermion parts}) \end{aligned}$$

@ LO

- Higgs field has degenerate minima on the ϕ -axis.
- Singlet field does not have VEV ($O(N)$ -symmetry is not broken).

@ NLO

- Higgs VEV determines and Higgs boson gets its mass.
- Higgs becomes lighter than singlet(s).
(cf. E. Gildener and S. Weinberg [*PRD*13, 3333 (1976)])

RG-improved potential

We used the two kinds of RG-improved potential.

[M. Bando et. al., *PLB* 301, 83 (1993)]

$$V_{\text{eff}}^{(\text{LL})}(\phi, \varphi) = \frac{\lambda_{\text{H}}(t)}{4}\phi^4(t) + \frac{\lambda_{\text{HS}}(t)}{2}\phi^2(t)\varphi^2(t) + \frac{\lambda_{\text{S}}(t)}{4}\varphi^4(t)$$

$$V_{\text{eff}}^{(\text{imp-NLO})}(\phi, \varphi) = \frac{\lambda_{\text{H}}(t)}{4}\phi^4(t) + \frac{\lambda_{\text{HS}}(t)}{2}\phi^2(t)\varphi^2(t) + \frac{\lambda_{\text{S}}(t)}{4}\varphi^4(t) \\ + \sum_i \frac{n_i}{64\pi^2} M_i^4(\phi(t), \varphi(t)) \left[\ln \frac{M_i^2(\phi(t), \varphi(t))}{\mu^2(t)} - c_i \right]$$

(t : running parameter)

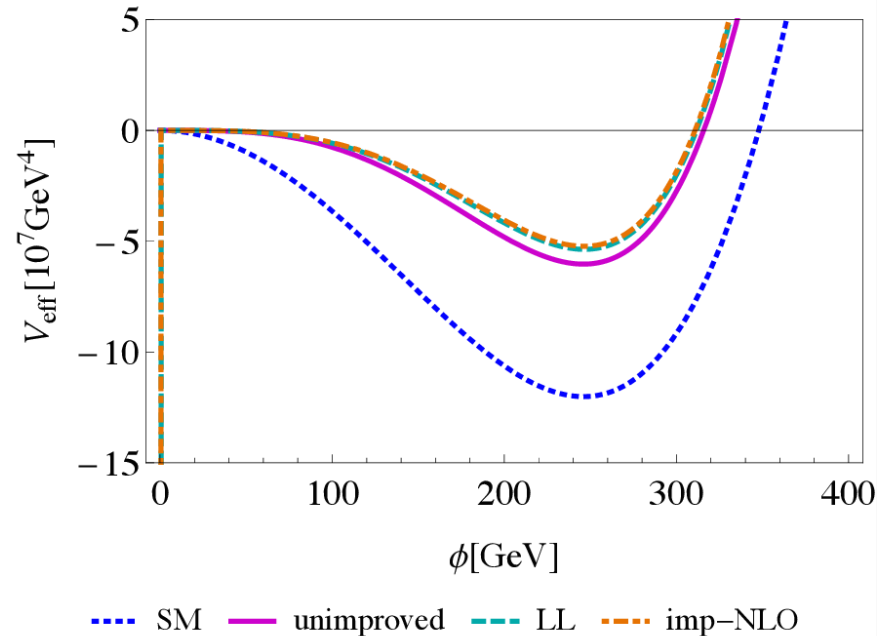
Using both, we can compare the “shape” around the vacuum and can confirm the validity of perturbation there.

3. Couplings of the Higgs and singlets

Results: $N = 1$ case

$N = 1$	
μ	$v_H = 246 \text{ GeV}$
$y_t(v_H)$	0.919
$g(v_H)$	0.644
$g'(v_H)$	0.359
$\lambda_H(v_H)$	-0.059
$\lambda_{HS}(v_H)$	4.5
$\lambda_S(v_H)$	0.10
$\langle H \rangle$	246 GeV
m_h	126 GeV
$\langle S \rangle$	0
m_s	527 GeV
Landau pole	4.1 TeV

K.E. and Y. Sumino



Almost the same as the previous work (R. Dermisck et.al. [1308.0891]).

→ Our analysis is valid around the VEV $\langle \phi \rangle = 246 \text{ GeV}$.

Results: $N = 4$ and 12 cases

$N = 4$		$N = 12$	
μ	$v_H = 246 \text{ GeV}$	μ	$v_H = 246 \text{ GeV}$
$y_t(v_H)$	0.919	$y_t(v_H)$	0.919
$g(v_H)$	0.644	$g(v_H)$	0.644
$g'(v_H)$	0.359	$g'(v_H)$	0.359
$\lambda_H(v_H)$	-0.061	$\lambda_H(v_H)$	-0.063
$\lambda_{HS}(v_H)$	2.3	$\lambda_{HS}(v_H)$	1.4
$\lambda_S(v_H)$	0.10	$\lambda_S(v_H)$	0.10
$\langle H \rangle$	246 GeV	$\langle H \rangle$	246 GeV
m_h	126 GeV	m_h	126 GeV
$\langle S \rangle$	0	$\langle S \rangle$	0
m_s	378 GeV	m_s	293 GeV
Landau pole	19 TeV	Landau pole	37 TeV

K.E. and Y. Sumino

As N increases,

- portal coupling λ_{HS} decreases
- and Landau pole goes far away,
- singlet mass decreases ($m_s^2 = \lambda_{HS} v_H^2$).

Results:

Couplings of the Higgs and Singlets

	$N = 1$	$N = 4$	$N = 12$
$\lambda_{hhh}/\lambda_{hhh}^{(\text{SM})}$	1.8	1.7	1.6
$\lambda_{hhhh}/\lambda_{hhhh}^{(\text{SM})}$	4.3	3.2	2.8
λ_{hss}	20	10	5.9
λ_{hhss}	13	5.7	3.2
λ_{ssss}	8.3	1.9	0.9

K.E. and Y. Sumino

$$\begin{aligned}
 V_{\text{eff}} = & \text{const.} + \frac{1}{2}m_h^2 h^2 + \frac{1}{2}m_s^2 \vec{s} \cdot \vec{s} + \frac{\lambda_{hhh}}{3!} v_H h^3 + \frac{\lambda_{hhhh}}{4!} h^4 \\
 & + \frac{\lambda_{hss}}{2} v_H h \vec{s} \cdot \vec{s} + \frac{\lambda_{hhss}}{4} h^2 \vec{s} \cdot \vec{s} + \frac{\lambda_{ssss}}{4!} (\vec{s} \cdot \vec{s})^2 + \dots,
 \end{aligned}$$

Comparison with GW's method

E. Gildener and S. Weinberg [*PRD13*, 3333 (1976)]

$N = 1$	λ_{hhh}	λ_{hhhh}	λ_{hss}	λ_{hhss}	λ_{ssss}
SM prediction	0.78	0.78	none	none	none
GW's framework	1.3	2.9	$2\lambda_{HS} = 9.6$	$2\lambda_{HS} = 9.6$	$6\lambda_S = 0.6$
our analysis (LL)	1.4	3.4	10.2	13.0	6.5

$N = 4$	λ_{hhh}	λ_{hhhh}	λ_{hss}	λ_{hhss}	λ_{ssss}
SM prediction	0.78	0.78	none	none	none
GW's framework	1.3	2.9	$2\lambda_{HS} = 4.8$	$2\lambda_{HS} = 4.8$	$6\lambda_S = 0.6$
our analysis (LL)	1.3	2.5	5.0	5.7	1.9

$N = 12$	λ_{hhh}	λ_{hhhh}	λ_{hss}	λ_{hhss}	λ_{ssss}
SM prediction	0.78	0.78	none	none	none
GW's framework	1.3	2.9	$2\lambda_{HS} = 2.8$	$2\lambda_{HS} = 2.8$	$6\lambda_S = 0.6$
our analysis (LL)	1.3	2.2	3.0	3.2	0.92

4. Veltman's condition for the Higgs mass

Is Veltman's condition favoured?

[Veltman's condition]

Coefficient of quadratic divergence

in quantum corrections for Higgs mass **vanishes**.

→ If it is, there is no fine-tuning.

Cutoff-regularised quantum correction:

$$V_1(\phi) = \frac{1}{64\pi^2} \text{STr} \left[\Lambda^4 \left(\ln \Lambda^2 - \frac{1}{2} \right) + 2M^2(\phi)\Lambda^2 + M^4(\phi) \left(\ln \frac{M^2(\phi)}{\Lambda^2} - \frac{1}{2} \right) \right] + \text{c.t.}$$

[O. Antipin et. al, 1310.0957v3] etc.

Coefficient of quadratic divergence for Higgs mass:

$$\frac{1}{2} \frac{\partial^2}{\partial h^2} \text{STr} M^2(h) \Big|_{\mu=\mu_0} = \frac{1}{v_H^2} (6m_W^2 + 3m_Z^2 - 12m_t^2 + Nm_s^2)$$

Is Veltman's condition favoured?

SM case

$$\frac{1}{v_H^2} [6m_W^2 + 3m_Z^2 - 12m_t^2 + m_h^2] \simeq -4$$

Our model

$$\frac{1}{v_H^2} [6m_W^2 + 3m_Z^2 - 12m_t^2 + Nm_s^2] \simeq \begin{cases} 0.3-0.8 & \text{for } N = 1 \\ 4.9-5.1 & \text{for } 4 \\ 12-13 & \text{for } 12 \end{cases}$$

based on K.E. and Y. Sumino

$N = 1$ case seems to be relatively and approximately favoured, so that Fine-tuning seems to be relaxed.

5. Singlets as dark matter?

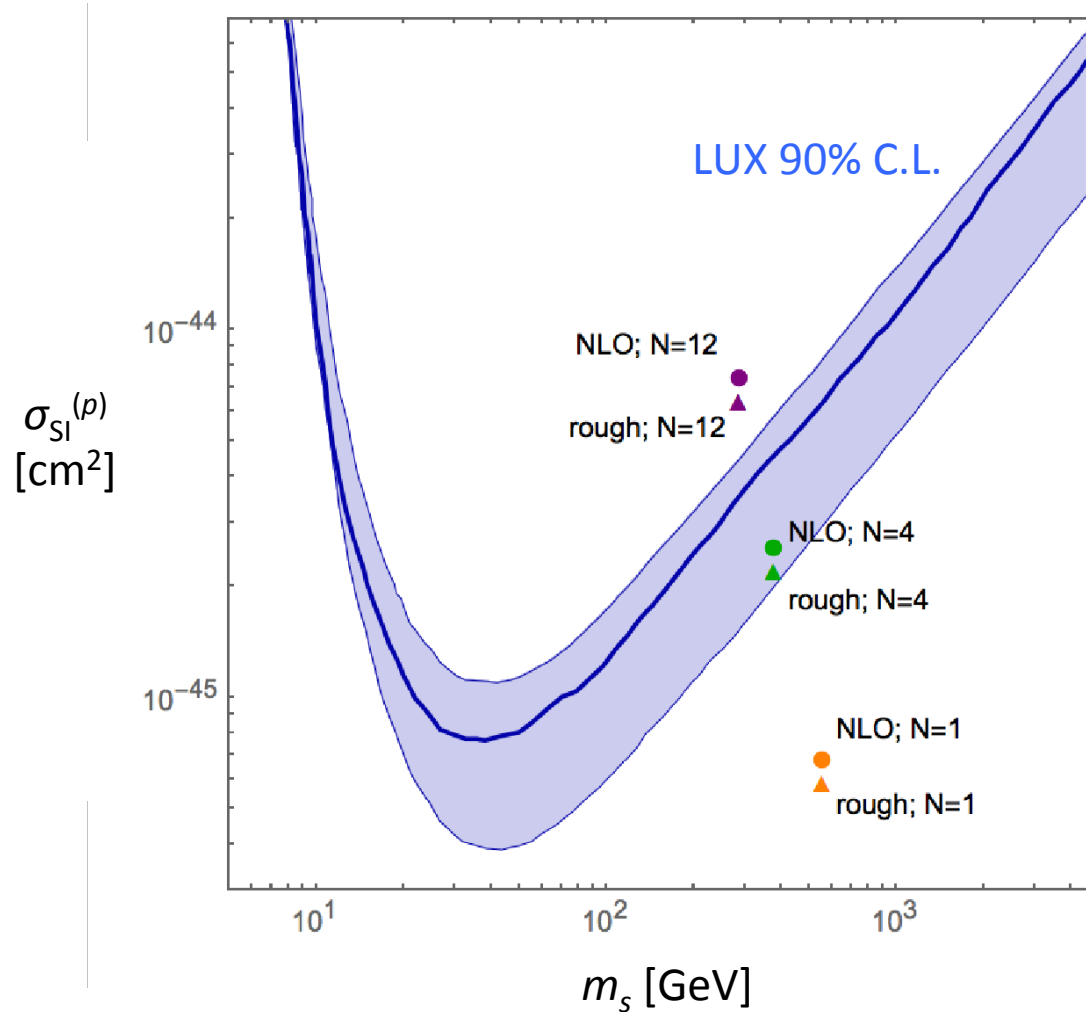
Singlets as dark matter?

- Singlets can be the dark matter candidates.
- Their thermal relic abundance are very small.
- Spin-independent cross sections with proton are predicted. Precision has been improved recently.

We used the framework in *JHEP* 1506 (2015) 097
(Hisano, Ishiwata, Nagata).

N	1	4	12
$\Omega_{s_i}/\Omega_{\text{DM}}$	2.01×10^{-4}	4.54×10^{-4}	8.07×10^{-4}
$\tilde{\sigma}_{\text{SI}}^{(p)} [10^{-46} \text{ cm}^2]$	6.77	25.6	74.5

Singlets as dark matter?



6. Summary

Summary

- Can only Higgs self couplings deviate (significantly) from SM predictions? → classically scale-inv. model
- We found perturbatively valid vacuum taking into account consistent order counting.
- Higgs self couplings: $\lambda_{3h}/\lambda_{3h}^{\text{SM}} \sim 1.7$, $\lambda_{4h}/\lambda_{4h}^{\text{SM}} \sim 3.0 - 4.0$
- Singlet scalars get mass around 300-500GeV and are only pair-produced.
- $N = 1$ case
 - Singlet can be dark matter; although relic abundance is small, it is detectable.
 - may approximately satisfy Veltman's condition.

Thank you for your attention.