

R-parity conserving $U(1)_X$ extended MSSM and its phenomenological aspects

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The model: R-parity conserving

Minimal SUSY $SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)_X$

	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_X$	R-parity
Q_i	3	2	+1/6	$x_Q = +\frac{1}{3}x_H + \frac{1}{3}x_\Psi$	-
U^c_i	3*	1	-2/3	$x_{U^c_i} = -\frac{4}{3}x_H - \frac{1}{3}x_\Psi$	-
D^c_i	3*	1	+1/3	$x_{D^c_i} = +\frac{2}{3}x_H - \frac{1}{3}x_\Psi$	-
L_i	3	2	-1/2	$x_L = -x_H - x_\Psi$	-
$N^c_{1,2}$	1	1	0	$x_{N^c_{1,2}} = +x_\Psi$	-
Ψ	1	1	0	$x_\Psi = +x_\Psi$	+
E^c_i	1	1	+1	$x_{E^c_i} = +2x_H + x_\Psi$	-
H_u	1	2	+1/2	$x_{H_u} = +x_H$	+
H_d	1	2	-1/2	$x_{H_d} = -x_H$	+

□: MSSM □: Additional parts of R-parity conserving Minimal SUSY $U(1)_X$ Model

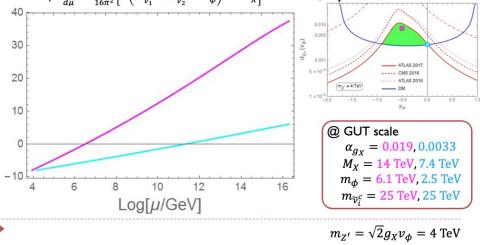
Suppose only ϕ (scalar component of Ψ) develops a VEV.
→ $U(1)_X$ symmetry is broken, while R-parity is conserved.

Radiative $U(1)_X$ symmetry breaking

Suppose $m_{\tilde{\chi}_1^0}^2 \gg m_\phi^2$ and MSSM sparticle mass²

$$\mu \frac{dm_{\tilde{\chi}_1^0}}{d\mu} = \frac{g_X^2}{16\pi^2} [2(m_{\tilde{\chi}_1^0}^2 + m_{\tilde{e}}^2 + m_{\tilde{\nu}}^2) - 8M_X^2].$$

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$$m_{Z'} = \sqrt{2}g_Xv_\phi = 4 \text{ TeV}$$

Dark Matter candidates

LSP (Lightest Super Particle) neutralino is a candidate for DM as usual in the MSSM.

New DM candidate:

$$(\chi_1) = (\cos\theta \quad -\sin\theta) \begin{pmatrix} \psi \\ \lambda_X \end{pmatrix} \leftarrow \text{fermion component of } \Psi$$

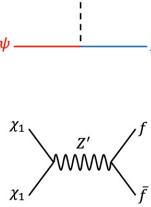
(χ_2) = ($\sin\theta \quad \cos\theta$) $\begin{pmatrix} \psi \\ \lambda_X \end{pmatrix} \leftarrow U(1)_X \text{ gaugino}$

Assuming that the lighter mass eigenstate χ_1 is the lightest neutralino.

→ χ_1 is DM candidate.

If $m_{\chi_1} = m_{DM} \sim \frac{1}{2}m_{Z'}$, the annihilation process is efficient and the DM relic abundance is reproduced.

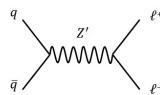
$$\Omega_{DM}h^2 = 0.120 \pm 0.01 \text{ [Planck 2018 (68% CL)]}$$



LHC Run-2 bounds on Z' boson mass

The dilepton production cross section:

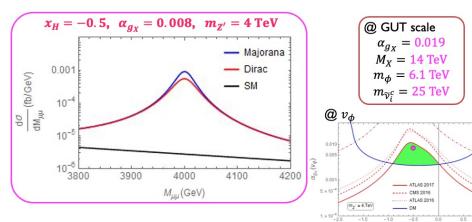
$$\frac{d\sigma(pp \rightarrow \ell^+\ell^- X)}{dM_{\ell\ell}} = \sum_{a,b} \int_{M_{CM}}^1 dx_{\ell\ell} \frac{2M_{\ell\ell}}{x_1 E_{CM}^2} f_a(x_1, M_{CM}^2) f_b(x_1, M_{CM}^2) \hat{\sigma}(\bar{q}q \rightarrow \ell^+\ell^-)$$



Dirac neutrinos at HL-LHC

1 massless Wyle + 2 massive Dirac neutrinos

3 light Majorana neutrinos + 3 heavy Majorana neutrinos



Superfields

Chiral Superfields

- $(x_H, x_\Psi) = (0, 1) \Rightarrow U(1)_{B-L}$
- $(x_H, x_\Psi) = \left(\frac{1}{2}, 0\right) \Rightarrow U(1)_Y$
- $(x_H, x_\Psi) = (-1, 1) \Rightarrow U(1)_R$

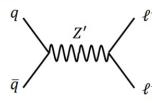
LHC physics and DM physics

For fixed X_H and $m_{Z'}$:

LHC

→ g_X Upper Bound

→ LHC Run-2

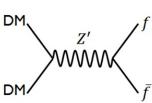


Z' portal DM

→ g_X Lower Bound

→ DM Relic Abundance

⇒ Complementarity between
DM physics and LHC physics



DM relic abundance [Non-equilibrium system]

Boltzmann Equation:

$$\frac{dY}{dx} = -\frac{x s \langle \sigma v \rangle}{H(m_{DM})} (Y^2 - Y_{EQ}^2)$$

DM relic abundance:

$$\Omega_{DM}h^2 = \frac{m_{DM}s_0 Y(\infty)}{\rho_c/h^2}$$

Observed DM relic abundance:

$$\Omega_{DM}h^2 = 0.120 \pm 0.01$$

Planck 2018 (68% CL)

→ $\langle \sigma v \rangle \sim 1 \text{ pb}$ leads to the right DM relic abundance.

yield : $Y = \frac{n}{s}$

$$x = \frac{m_{DM}}{T}$$

$$s = \frac{4\pi g_* m_{DM}^2}{45^2 M_P^2}$$

$$H(m_{DM}) = \sqrt{4\pi g_* m_{DM}^3/M_P^2}$$

$$s Y_{EQ} = \frac{g_* m_{DM}^2}{2\pi^2} K_2(x)$$

$$\langle \sigma v \rangle = \frac{s}{n} \frac{Y_{EQ}}{Y}$$

$$Y(\infty) = \frac{n}{s} = \frac{n_0}{s_0} e^{-\int_x^\infty \frac{x s \langle \sigma v \rangle}{H(m_{DM})} dx}$$

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