

SUSY Model Discrimination at an Early Stage of the LHC

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If $\text{bkg} \sim 0$,

discovery \longleftrightarrow ~ 5 events

measurement \longleftrightarrow $> O(100-1000)$ events

Measurements may require a few years.

Why can we do **before** detailed measurements?

When discovered \longleftrightarrow crosssection

\longleftrightarrow mass scale of produced particles

What discovered \longleftrightarrow mass spectrum topology

How further can we go in this direction?

We compare **mSUGRA** and (low scale) **mGMSB**

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Low scale \Leftrightarrow light gravitino mass.

Model with gravitino mass < 16 eV is free from gravitino problem.

We compare **mSUGRA** and (low scale) **mGMSB**

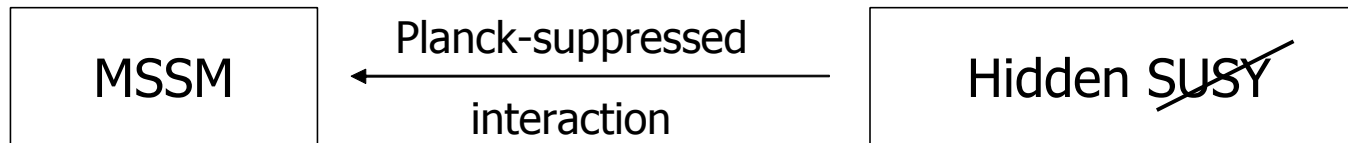
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Low scale \Leftrightarrow light gravitino mass.

Model with gravitino mass < 16 eV is free from gravitino problem.

Missing gravitinos resemble missing neutralinos.

mSUGRA



The model parameters (assuming good flavor and CP properties):

$$m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu$$

$$m_{\text{sfermion}}(m_{\text{GUT}}) = m_0$$

$$m_{\text{gaugino}}(m_{\text{GUT}}) = m_{1/2}$$

mGMSB (Gauge Mediated Supersymmetry Breaking)



The model parameters (assuming GUT relations):

$$\Lambda_s, \Lambda_g, m_{\text{mes}}, \tan \beta, \text{sign } \mu$$

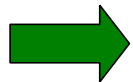
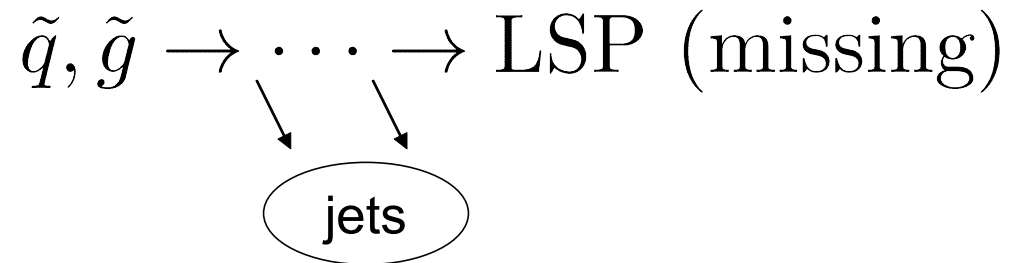
$$m_{\text{sfermion}}(m_{\text{mes}}) = \frac{\alpha_{\text{SM}}}{4\pi} \Lambda_s$$

$$m_{\text{gaugino}}(m_{\text{mes}}) = \frac{\alpha_{\text{SM}}}{4\pi} \Lambda_g$$

LHC signature

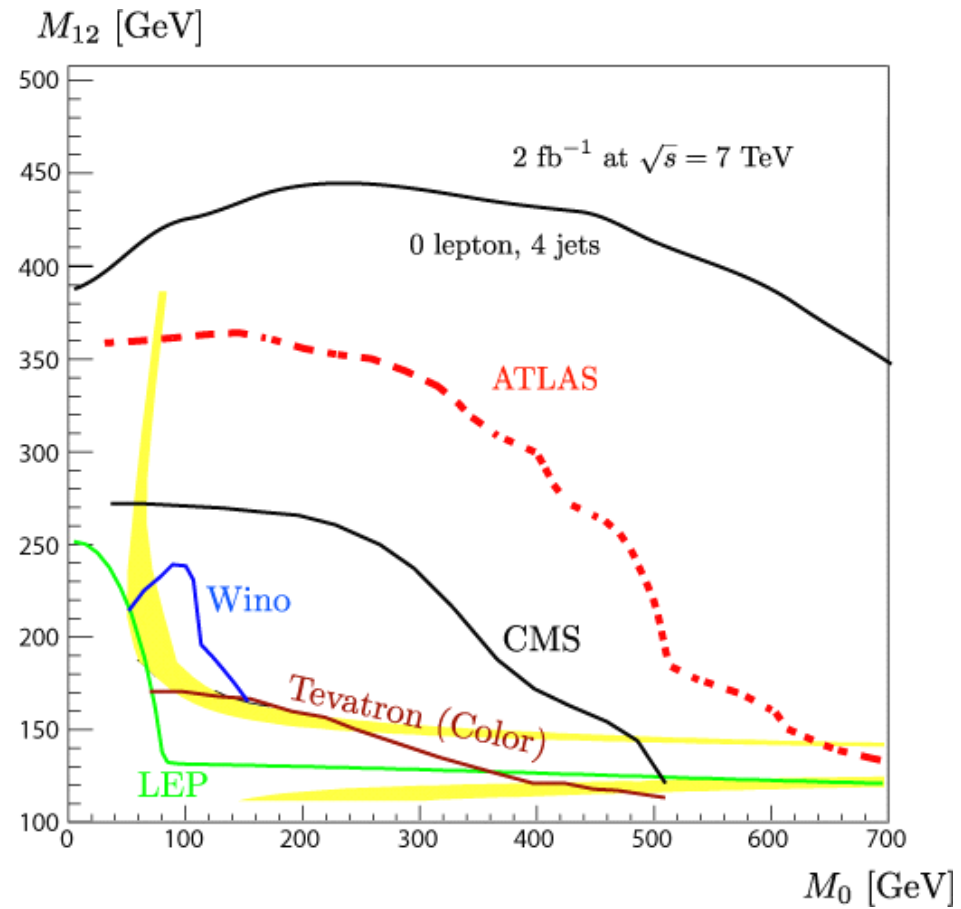
Typically, SUSY events are produced with colored particles \tilde{q} , \tilde{g}

Then, SUSY particles decay cascadelly to the **LSP** (Lightest SUSY Particle):

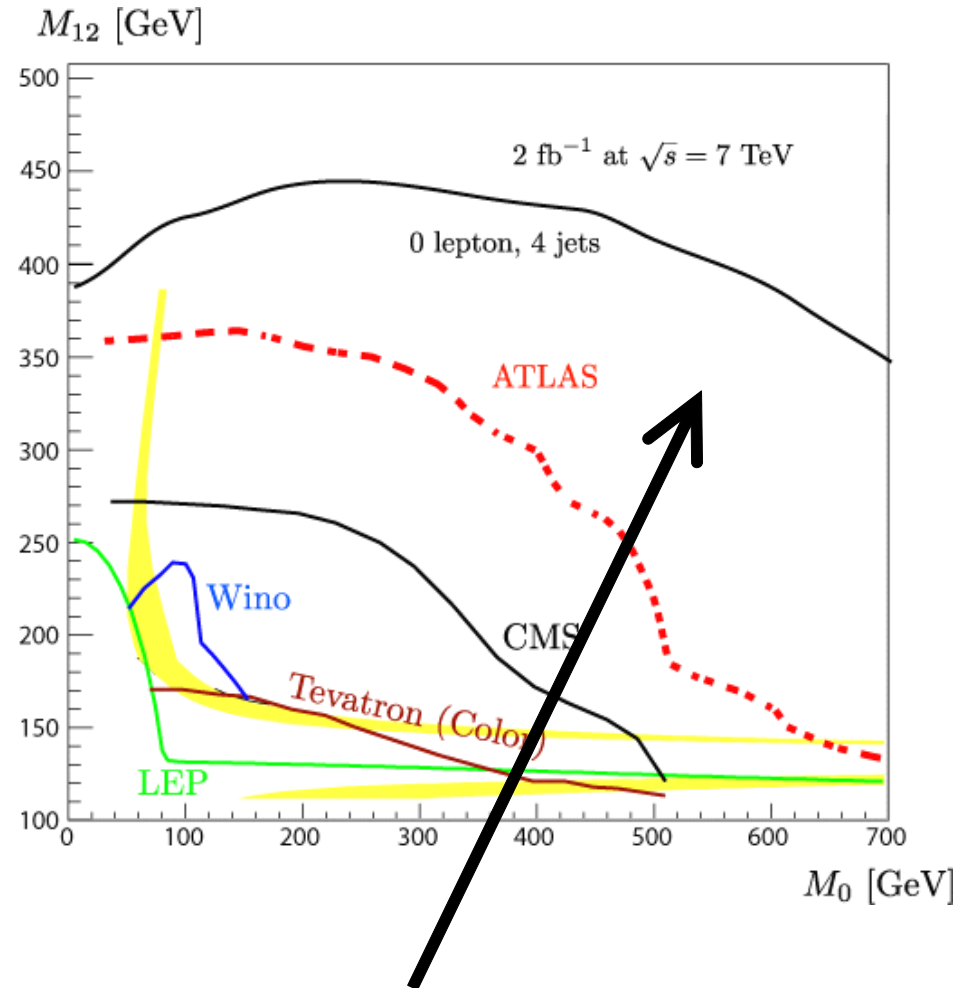


Multi-jets + missing

Example: mSUGRA discovery potential



Example: mSUGRA discovery potential



We consider this region

mSUGRA vs mGMSB

We consider how to discriminate mSUGRA events with mGMSB events.

In addition to **multi-jets + missing**,
SUSY events may contain other particles (leptons, photons, ...)

mSUGRA LSP = $\tilde{\chi}_1^0 \sim \tilde{B}$

$$\tilde{g} \xrightarrow{q} \tilde{q}^{(*)} \xrightarrow{q} \tilde{\chi}_1^+ \xrightarrow{W^+} \tilde{\chi}_1^0, \quad W \rightarrow qq', \textcircled{\ell\nu}.$$

$$\tilde{g} \xrightarrow{q} \tilde{q}^{(*)} \xrightarrow{q} \tilde{\chi}_2^0 \xrightarrow{\textcircled{\ell}} \tilde{\ell}^{(*)} \xrightarrow{\textcircled{\ell}} \tilde{\chi}_1^0.$$

mSUGRA vs mGMSB

mGMSB LSP = \tilde{G} (gravitino) (\sim massless)

$$\tilde{g} \xrightarrow{q} \tilde{q}^{(*)} \xrightarrow{q} \tilde{\chi}_1^0 \xrightarrow{\gamma/Z} \tilde{G}$$

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Coupling to gravitino is weaker than the SM couplings



mSUGRA vs mGMSB

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Coupling to gravitino is weaker than the SM couplings

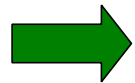
 $\tilde{q}, \tilde{g} \rightarrow \dots \rightarrow$ **NLSP** \rightarrow LSP (gravitino)
Always!

$$\text{NLSP} = \left\{ \begin{array}{l} (1) \text{ Slepton} \\ (2) \text{ Lightest neutralino } (\sim \tilde{B}) \end{array} \right.$$

mSUGRA vs mGMSB

$$\text{NLSP} = \begin{cases} (1) \text{ Slepton} \\ (2) \text{ Lightest neutralino } (\sim \tilde{B}) \end{cases}$$

$$(1) \quad \tilde{q}, \tilde{g} \xrightarrow{\text{jets}} \dots \xrightarrow{\ell} \tilde{\ell} \xrightarrow{\ell} \tilde{G}$$

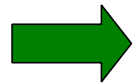


multi-jets + missing + **multi-leptons**

mSUGRA vs mGMSB

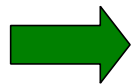
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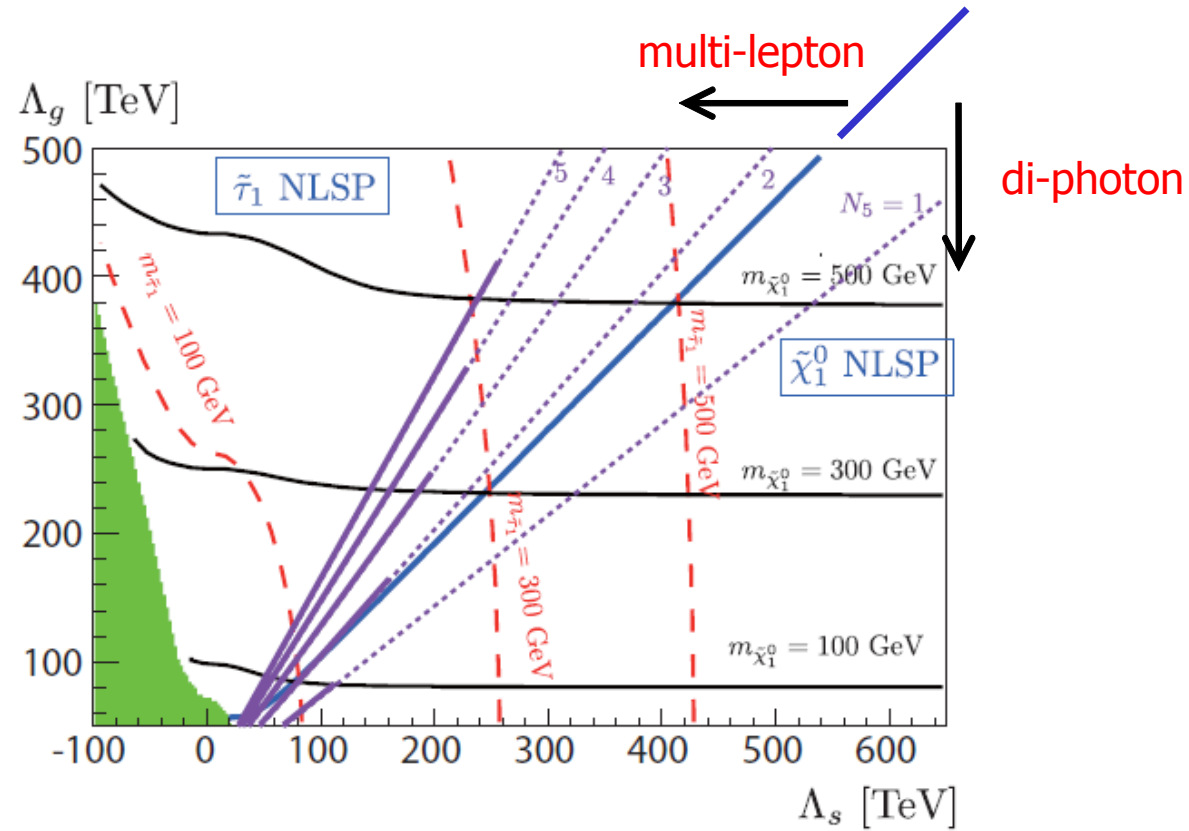
multi-jets + missing + **multi-leptons**

$$(2) \quad \tilde{q}, \tilde{g} \xrightarrow{\text{jets}} \dots \rightarrow \tilde{B} \xrightarrow{\gamma} \tilde{G}$$

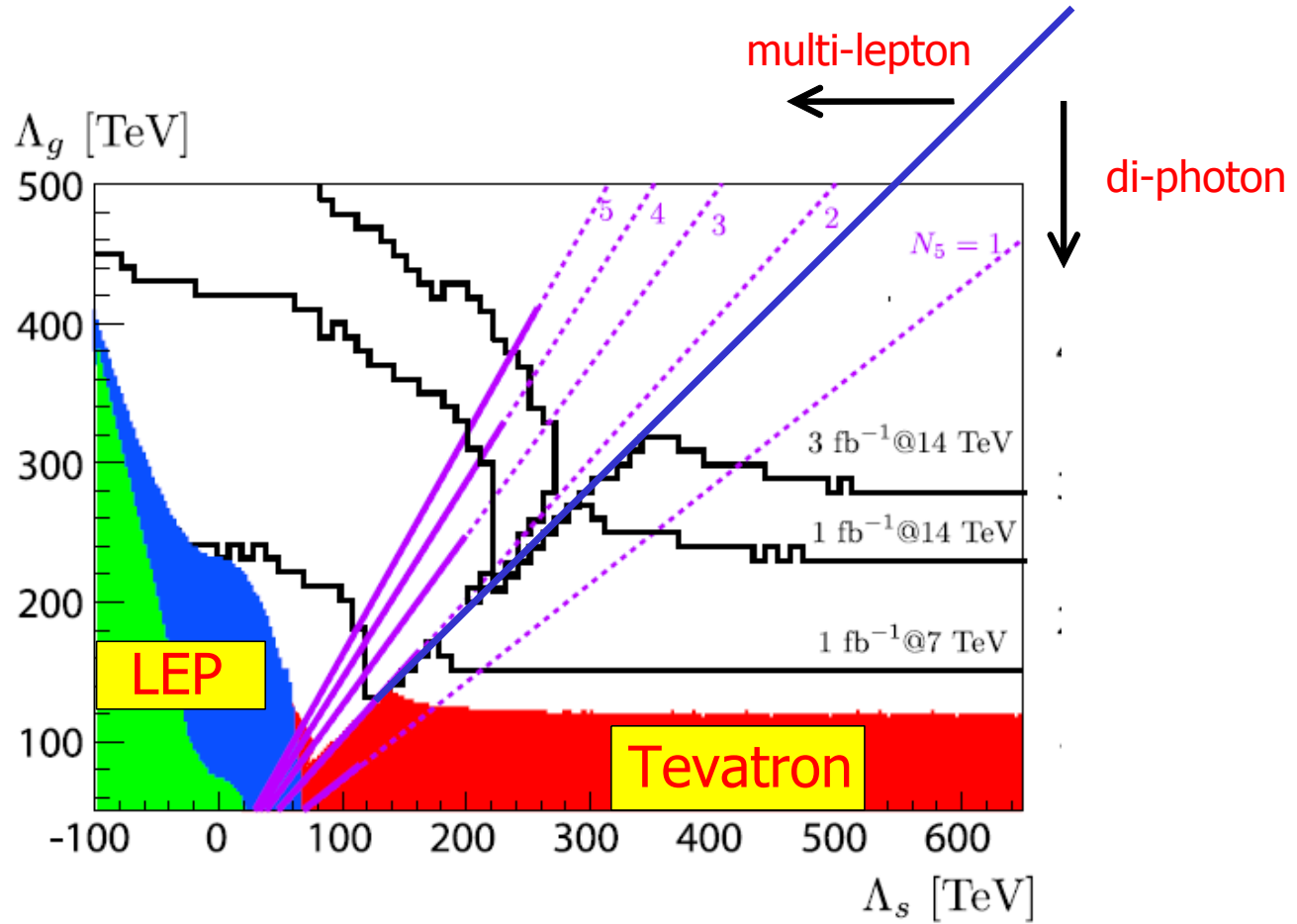


multi-jets + missing + **di-photons**

mGMSB discovery potential



mGMSB discovery potential



mSUGRA vs mGMSB

Naive estimation:

	2 photons	Same-sign 2 leptons	0 lepton 4 jets
mSUGRA	×	○	⊙
mGMSB $\tilde{\tau}$ -NLSP	×	⊙	△
mGMSB $\tilde{\chi}_1^0$ -NLSP	⊙	△	△

We now try to discriminate the models according to this table.

Significance variable

We use the **significance** Z
to express "goodness" of each mode.

$$Z = Z(N_s, N_b)$$

Statistical uncertainty
SM background contribution] is automatically incorporated.

$$Z > 5 \Leftrightarrow 5\sigma \text{ discovery}$$

The error of Z depends on N_s and N_b .
(Typically, $\delta Z \simeq 1$)

Discrimination of mGMSB with neutralino-NLSP

We scan the model points in the
(sfermion mass)-(gaugino mass) plane.

For simplicity, A-term is set to 0, $\mu > 0$.

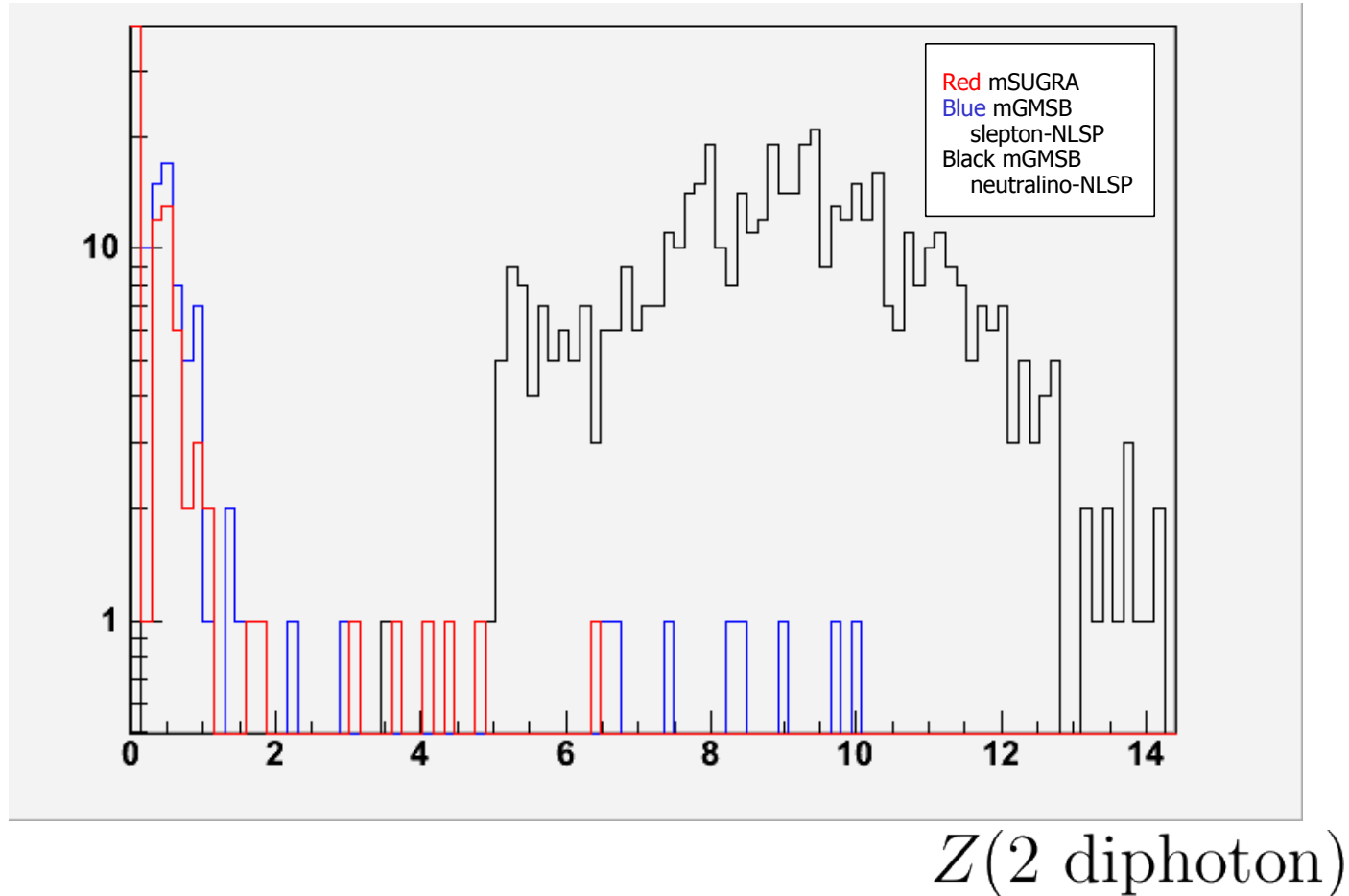
tan beta = 10,20,...,50 (mSUGRA)
10,40 (mGMSB)

Cuts

	$E_{\text{miss}}^{\text{T}}$	M_{eff}	$E_{\text{miss}}^{\text{T}}/M_{\text{eff}}$	4/2 jets	other
SS2 ℓ	100	1000		(100,50)	$l_2 > 20$
2 γ	200			(90,60)	
0 ℓ +4 jets	200	1000	0.2/0.3	(100,50)	$l < 10$
0 ℓ +2 jets	300	1200	0.3	(300,200)	$l < 10$

Discrimination of mGMSB with neutralino-NLSP

#model points

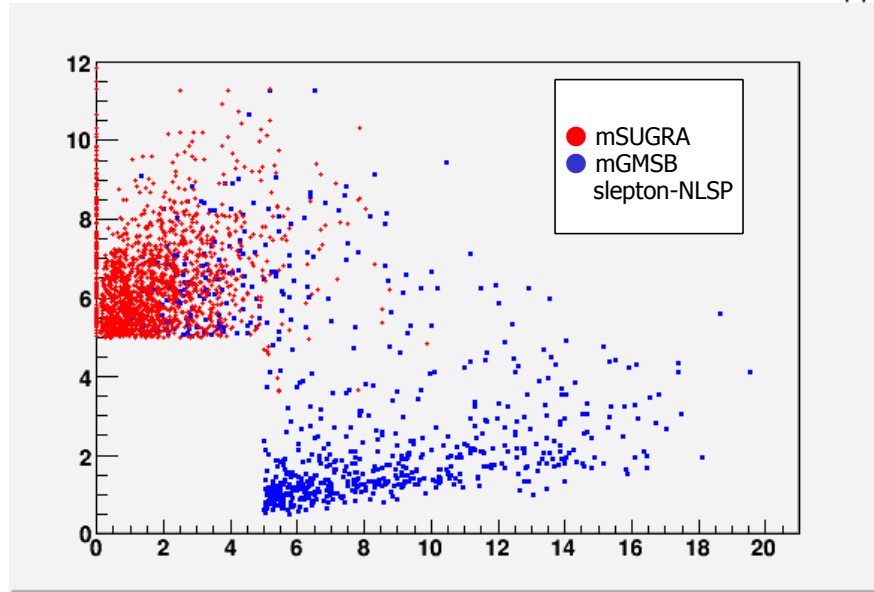


The neutralino-NLSP case can easily be discriminated.

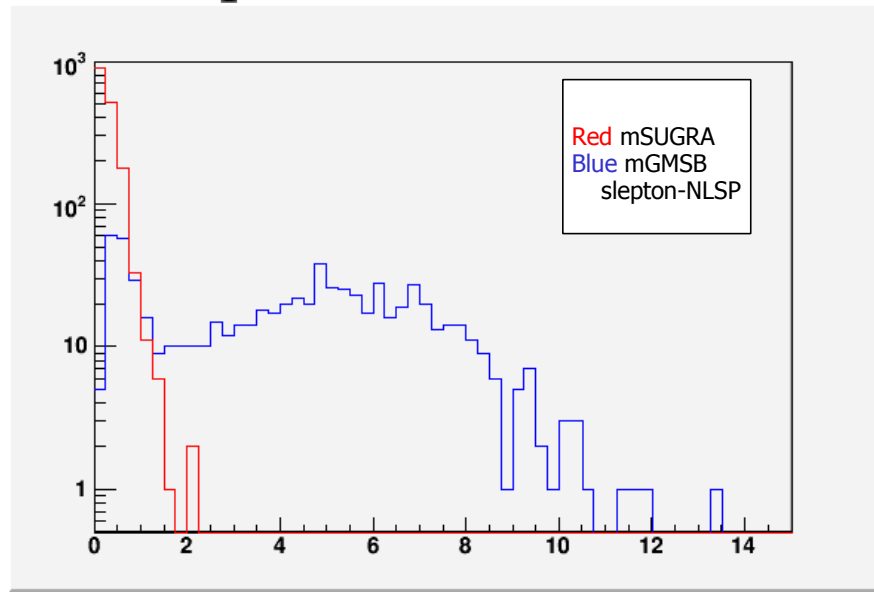
mSUGRA vs mGMSB with slepton-NLSP

$Z(0\ell+4 \text{ jets})$

#model points



$Z(\text{Same-sign } 2\ell)$



$Z(\text{Same-sign } 2\ell)/Z(0\ell+4 \text{ jets})$

Most mGMSB model points are discriminated from mSUGRA points
But, mGMSB with nearly degenerated slepton-bino contaminates.

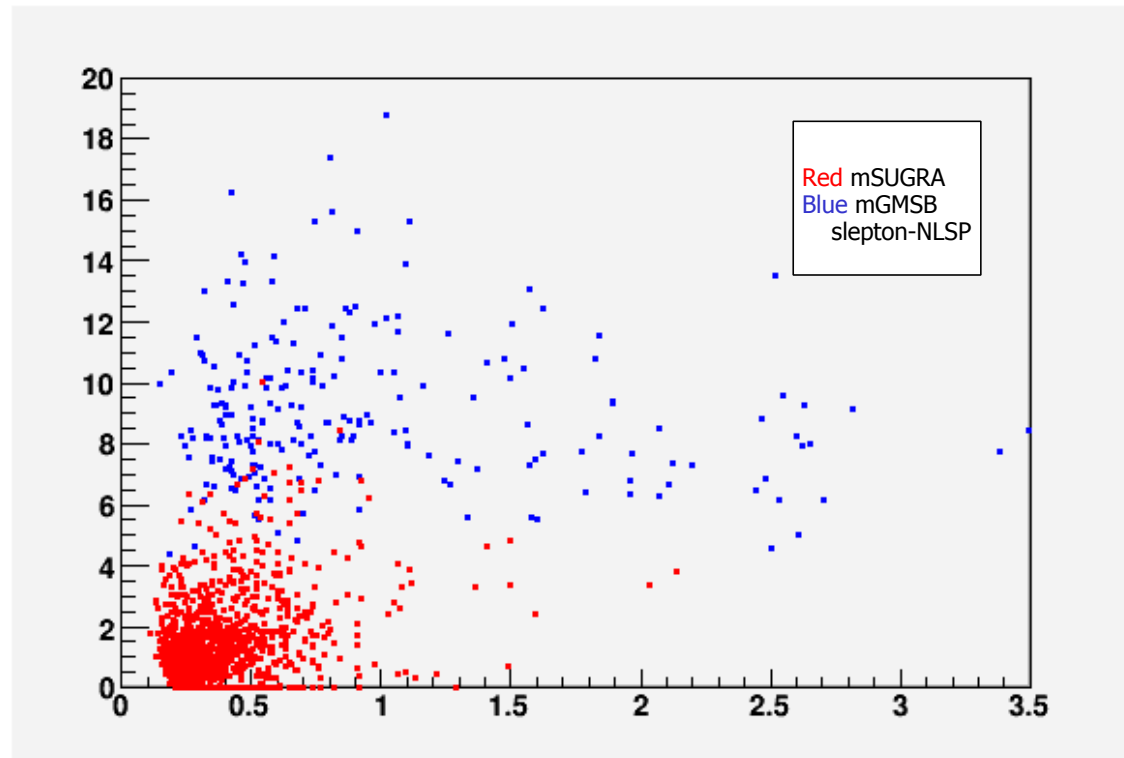
mSUGRA vs mGMSB with slepton-NLSP

mGMSB with nearly degenerated slepton-bino contaminates.

In this case, **more taus** are expected.

mSUGRA vs mGMSB with slepton-NLSP

$Z(2\tau)$

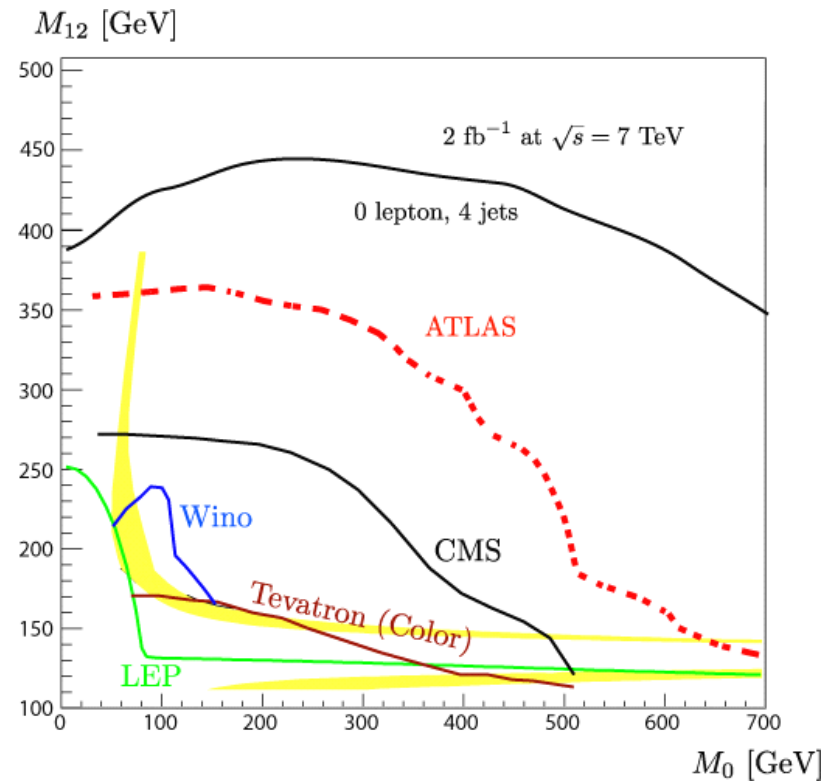


$Z(\text{Same-sign } 2\ell)/Z(0\ell+4\text{jets})$

かなり満足

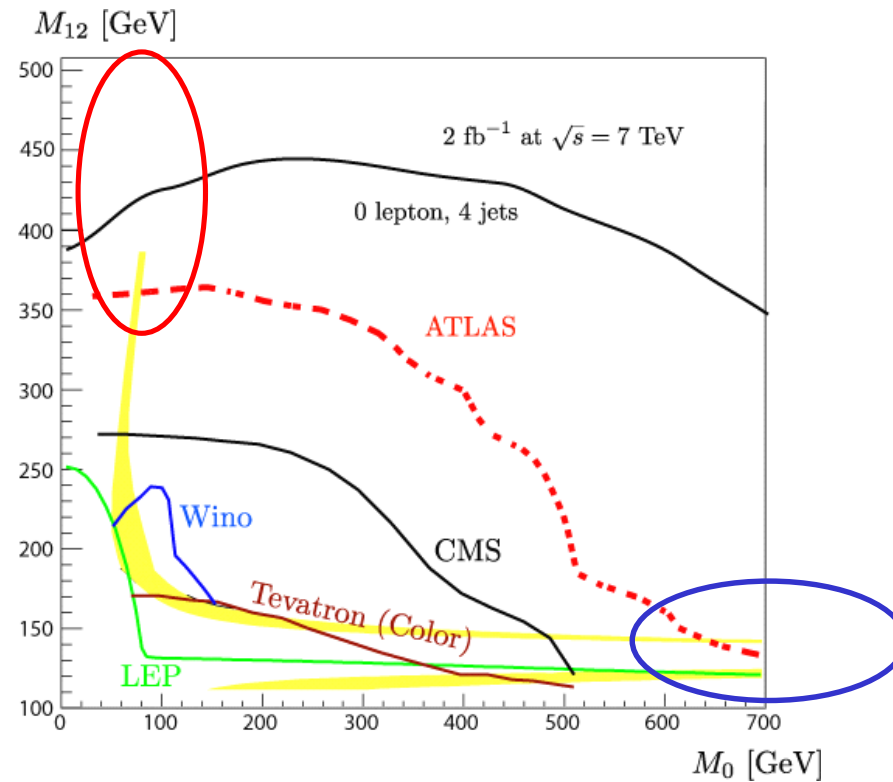
Discremination among mSUGRA model points

For neutralino DM, the focus point and coannihilation point are interesting.



Discremination among mSUGRA model points

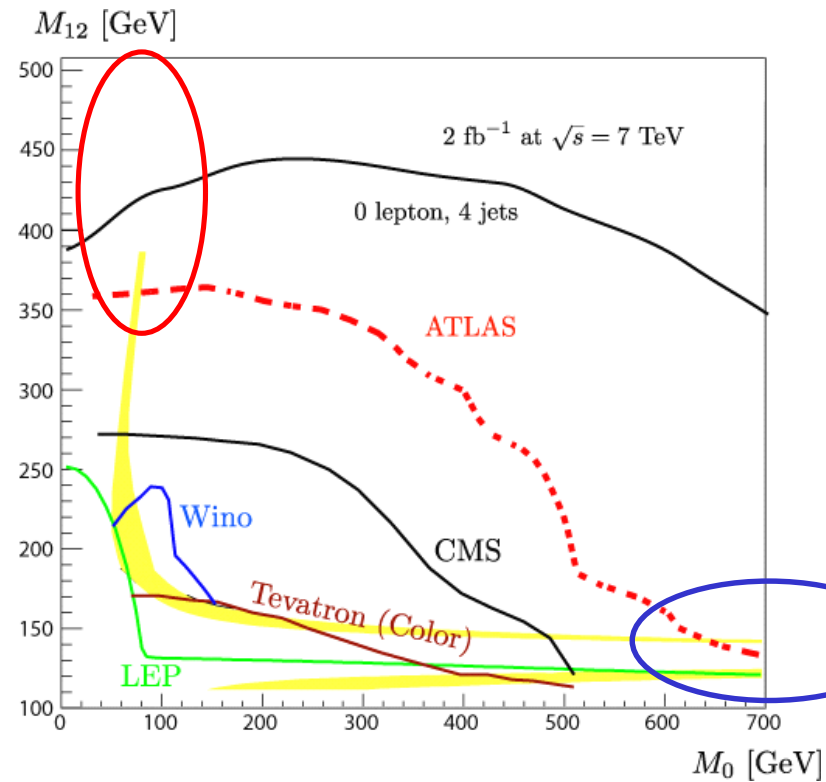
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Discremination among mSUGRA model points

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$\tilde{q}\tilde{q}, \tilde{q}\tilde{g}$
dominated

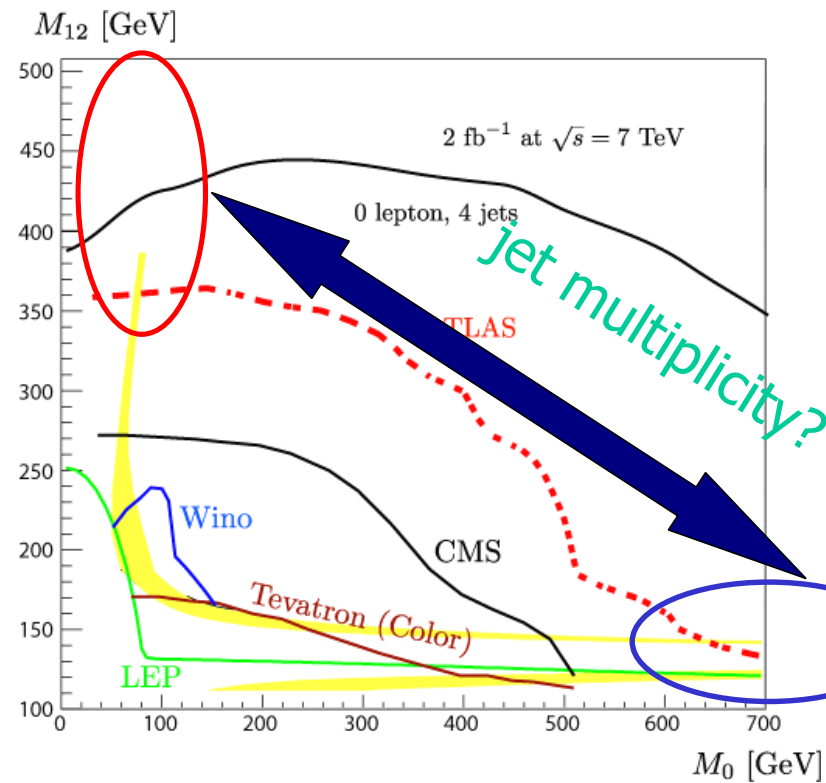


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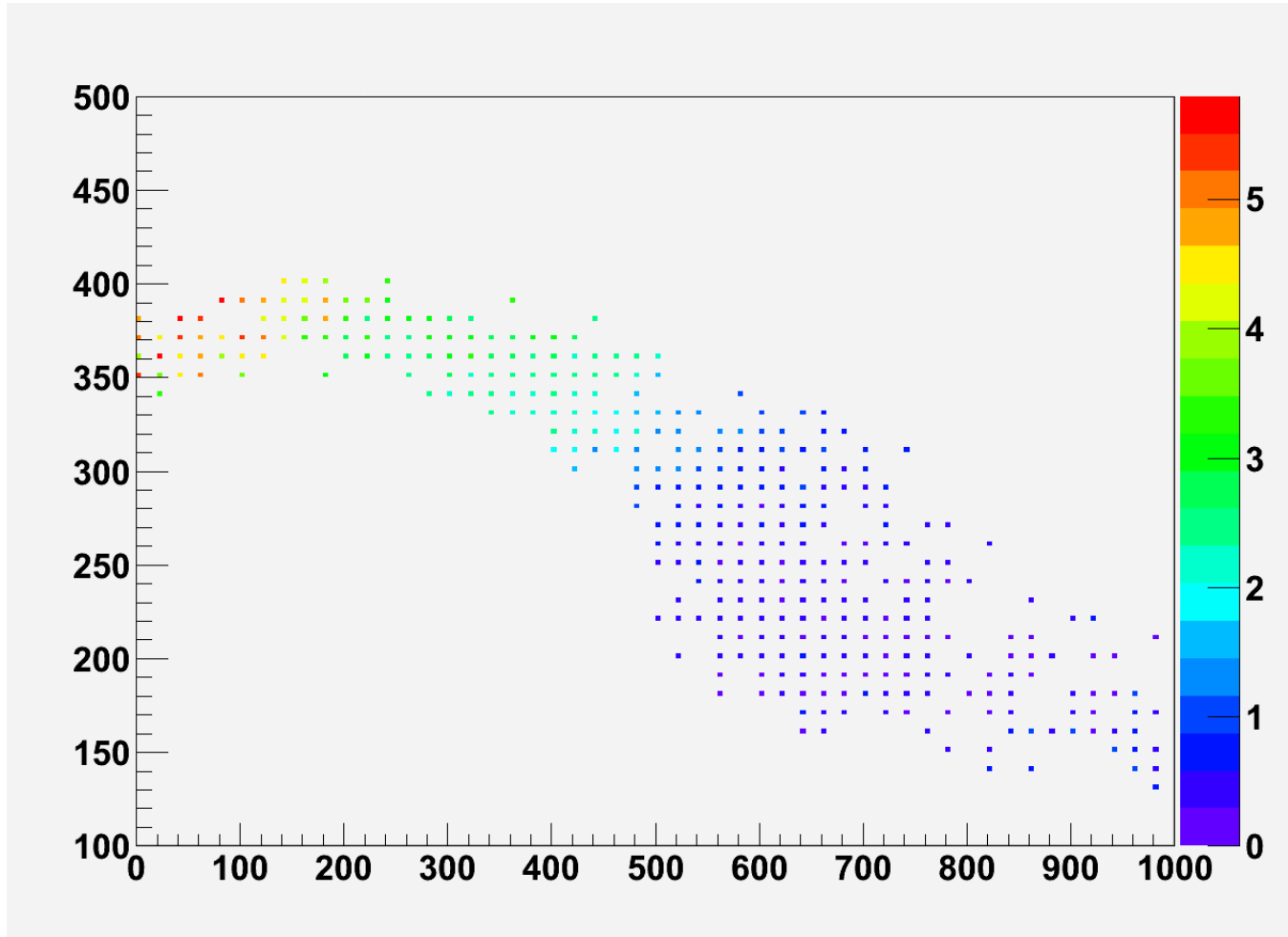
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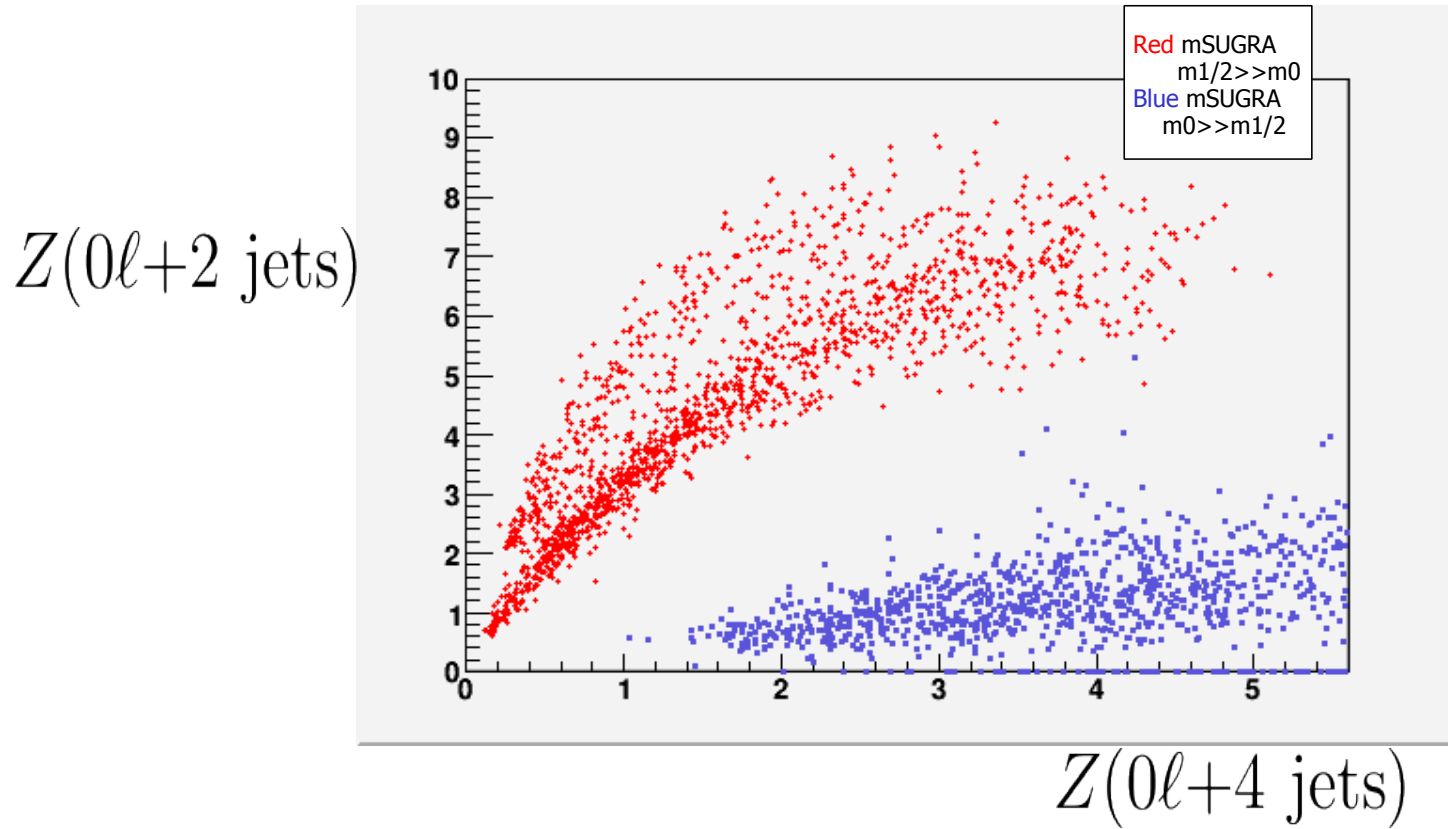
$\tilde{g}\tilde{g}$ dominated

Discremination among mSUGRA model points

$$Z(0\ell+2 \text{ jets})/Z(0\ell+4 \text{ jets})$$



Discremination among mSUGRA model points



Summary

Early LHC (2fb-1 @ 7 TeV) でのmSUGRAとmGMSBの discriminationを考えた.

0 lepton+4 jets/SS 2 lepton(tau)/2 photon mode
でかなり区別が出来る.

2j mode/4j mode (jet multiplicity)を使って
mSUGRAの(m0-m1/2)-planeの位置の情報もある程度
得られそう.