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真空崩壊確率の完全な1-LOOP ORDER での計算とその応用について ~スケール不定性を減らせるか~



Now ongoing...

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Bubble Nucleation Rate



At the Leading order,



TOY MODEL



TOY MODEL

Nι

In fact, the potential is scale-dependent.

Potential
$$V = -\bar{t}(Q)\phi + \frac{\bar{m}^2(Q)}{2}\phi^2 - \frac{\bar{A}(Q)}{2}\phi^3 + \frac{\bar{\alpha}(Q)}{8}\phi^4$$

Bounce $\partial^2 \phi = V'(\phi)$
Action $B = S_E(\phi_B) = \int d^4x \left[\frac{1}{2}(\partial\phi_B)^2 + V(\phi_B)\right]$
Here $\gamma \simeq m^4 e^{-B}$
Maybe, the best Q is a "typical" scale

. . .



スケール依存性の大きさ

$$V = -\bar{t}(Q)\phi + \frac{\bar{m}^2(Q)}{2}\phi^2 - \frac{\bar{A}(Q)}{2}\phi^3 + \frac{\bar{\alpha}(Q)}{8}\phi^4$$

Beta functions

$$\beta_t = \frac{3Am^2}{16\pi^2} \qquad \beta_{m^2} = \frac{3}{16\pi^2} (\alpha m^2 + 3A^2)$$
$$\beta_A = \frac{9\alpha A}{16\pi^2} \qquad \beta_\alpha = \frac{9\alpha^2}{16\pi^2}$$

Renormalization conditions

@Q = m $\bar{m}^2(m) = m^2, \ \bar{A}(m) = m, \ \bar{t}(m) = 0, \ \bar{\alpha}(m) = \alpha$ スケール依存性の大きさ



Much larger uncertainty in a realistic model (w/ top loop)

1-LOOP ORDERでの計算 さて、どうしましょうか。



もっと読み返してみると $\gamma = Ae^{-B}$

$$A = \frac{B^2}{4\pi^2} \left(\frac{\det' S''|_{\text{Bounce}}}{\det S''|_{\text{False}}} \right)^{-1/2}$$



SOLVE corresponding Ordinary Differential Equations

Since 1928

Many mathematical proofs, but not so many pheno. results I. M. Gelfand, A. M. Yaglom; S. Coleman; J. H. van Vleck; R. H. Cameron, W. T. Martin; R. Dashen, B. Hasslacher, A. Neveu; R. Forman; K. Kirsten, A. J. McKane;

method, proof, renormalization, zero modes, fermions, implementation, ...

Please invite me to your LAB!!



 $\gamma = Ae^{-B} \equiv m^4 e^{-B - \delta B}$











 $\alpha = 0.9$





SM+STAU SYSTEM TOP LOOP



Staus can be light $m_{\tilde{\tau}} > 103.5 \text{GeV}(\text{LEP})$

hyy coupling, co-annihilation with Bino, ...

But, the potential may become unstable toward the stau direction

$$V = +\frac{1}{\sqrt{2}}y_{\tau}X_{\tau}\tilde{\tau}_{L}\tilde{\tau}_{R}h + \frac{m_{L}^{2}}{2}\tilde{\tau}_{L}^{2} + \frac{m_{R}^{2}}{2}\tilde{\tau}_{R}^{2} + \cdots \qquad X_{\tau} = A_{\tau} - \mu \tan\beta$$
$$\tan\beta = \langle H_{\tau}^{0} \rangle / \langle H_{\tau}^{0} \rangle$$

Stable

EW vacuum is the global minimum

Meta-Stable $t_{\rm dec} \gtrsim 13.8 {\rm Gyr}$

Unstable $t_{\rm dec} \lesssim 13.8 {\rm Gyr}$

Tachyonic stau

No EW vac.

 $\mathbf{U} + \mathbf{I}$

考えるスペクトラム

For simplicity, we consider the case where only the staus are light





結果其の壱

$m_L = m_R = 600 \text{GeV}, \ X_\tau = 95 \text{TeV}, \ \tan \beta = 15$



結果其の壱

 $m_L = m_R = 600 \text{GeV}, \ X_\tau = 95 \text{TeV}, \ \tan \beta = 15$



Caveat: Overall factor is NOT determined!

結果其の弐





Caveat: The position of the green lines can be changed!

まとめ

- The bubble nucleation rate has often been estimated without calculating the pre-exponential factor.
- This estimate involves uncertainty in the renormalization scale, which, we showed, results in O(10%) uncertainty in the exponent of the bubble nucleation rate.
- To reduce the uncertainty, we explicitly calculated the preexponential factor and showed that it is greatly reduced.
- Scalars and fermions have already been implemented, but the gauge bosons are now ongoing.



"Backup called Trash"

One thing I'm concerned about is the file size...