

Dynamical scalegenesis via multiple seesaws

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Based on

Phys.RevD.95 (2017)
H.Ishida, S.M, S.Okawa, Y.Omura

and also
Phys.RevD.94(2016); 1610.07137,
H.Ishida, S.M, Y.Yamaguchi

@ PPP2017
07/31 - 08/04/2017

0. Introduction

☆ Discovery of Higgs boson in 2012

--- last piece of particles predicted in SM

--- very successful SM pheno. so far



☆ It's **NOT** the end of the story!...

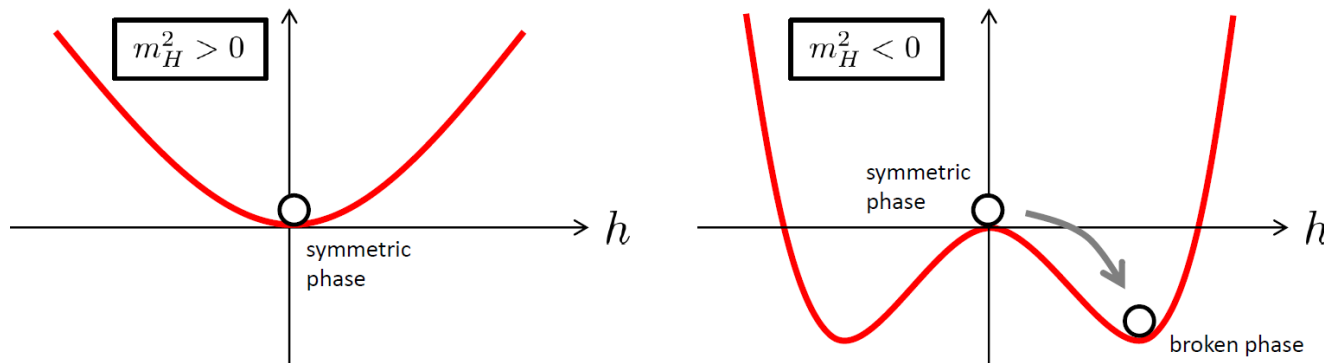
--- still lots of stuff left needed to account for

e.g. neutrino masses (mixing),
dark matter, baryon asymmetry, etc

☆ In particular,
unsatisfactory stuff on theoretical ground:

--- origin of mass is given “by hand”

Higgs potential: $V = \lambda_H (H^\dagger H)^2 + m_H^2 (H^\dagger H)$ $H = (0, \frac{1}{\sqrt{2}}(v_H + h))^T$



--- why $m_H^2 < 0$ (dynamical origin?) and $O(100 \text{ GeV})^2$?
(related to gauge hierarchy problem)

--- No answer in SM

☆ Typical ideas to access the origin of mass
(or gauge hierarchy problem)

- i) **eliminate** Higgs, replace fully by strong dynamics
e.g. Technicolor-like scenario
- ii) **protect** Higgs by some symmetry
e.g. SUSY, composite Higgs, scale symmetry

☆ Typical ideas to access the origin of mass
(or gauge hierarchy problem)

Well... it's on break

- i) **eliminate** Higgs, replace fully by strong dynamics
e.g. Technicolor-like scenario



- ii) **protect** Higgs by some symmetry

e.g. SUSY, composite Higgs, scale symmetry

--- Today, let's take **scale symmetry!**
(**classical scale-inv.**)

☆ Then, how to generate “scales”, realize EWSB?
→ dimensional transmutation

i) Coleman-Weinberg mechanism
(Radiative sym. Breaking)

But, no way only within SM
(vacuum instability due to large top loop)

$$m_{h(125)}^2 \sim \frac{1}{8\pi^2 v^2} (9m_W^4 - 12m_t^4)$$

--- Call for Beyond SM

☆ Then, how to generate “scales”, realize EWSB?
→ dimensional transmutation

i) Coleman-Weinberg mechanism
(Radiative sym. Breaking)

Extensions w/ Higgs portal term

$$V = \kappa |\phi|^2 (H^\dagger H)$$

(a) $\kappa = O(1)$, and >0 ; Φ loop effect included

-- assume Φ -symmetric phase: $\langle \Phi \rangle = 0$

-- requires Φ mass $> O(\text{EW})$ or large # of Φ to stabilize EW vacuum:

$$m_\phi^2 \sim \kappa v^2 \quad m_{h(125)}^2 \sim \frac{1}{8\pi^2 v^2} (9m_W^4 - 12m_t^4 + N_\phi m_\phi^4)$$

K.Endo, et al (2016); K.Hashino, et al(2016)

☆ Then, how to generate “scales”, realize EWSB?
→ dimensional transmutation

i) Coleman-Weinberg mechanism
(Radiative sym. Breaking)

Extensions w/ Higgs portal term

$$V = \kappa |\phi|^2 (H^\dagger H)$$

(b) $\kappa \ll O(1)$, and < 0 ; Φ effect decoupled, but

-- if Φ is $U(1)'$ (e.g. $U(1)_{B-L}$) charged,
then CW mechanism works for the $U(1)'$ sector

-- nonzero Φ -VEV generates Higgs mass term < 0 ,
triggers EWSB

$$V = -|\kappa| v_\phi^2 (H^\dagger H)$$

*S.Iso, et al (2009,2013); N.Okada, et al(2012,2016);
I.Oda (2013); V.V.Khoze, et al(2013); M.Hashimoto, et
al (2014); J.Guo, et al(2015)*

☆ Then, how to generate “scales”, realize EWSB?
→ dimensional transmutation

ii) Strong gauge dynamics (hidden/dark QCD)
(nonperturbative sym. Breaking)

Extensions w/ Higgs portal term

$$V = \kappa |\phi|^2 (H^\dagger H)$$

* $\kappa \ll 0(1)$, and < 0

* Φ = a scalar with strong coupling to hidden/dark QCD

--- below the dynamical scale, dark QCD generates scalar condensation:

$$\langle \phi^\dagger \phi \rangle \simeq \Lambda_{\text{dQCD}}^2$$

→ $V = -|\kappa| \Lambda_{\text{dQCD}}^2 (H^\dagger H)$

*F.Wilczek (2008); T.Hur, et al(2011);
M.Holthausen, et al (2013); J.Kubo, et al(2014)*

☆ Then, how to generate “scales”, realize EWSB?
→ dimensional transmutation

iii) Strong gauge dynamics (vectorlike-SM gauging)
= vectorlike confinement (hypercolor: HC)
--- “bosonic seesaw mechanism”

*N.Haba et al (2009,2016,2017); O.Antipin, et al (2015);
H.Ishida, S.M., Y.Yamaguchi (2016);
H.Ishida, S.M., S.Okawa, Y.Omura (2017)*

Extensions w/ Yukawa term & HC fermions ($F_{\{L/R\}}$)

$$\mathcal{L} = -y(\bar{\chi}H\psi) + \text{h.c.}$$

	$SU(N)_{\text{HC}}$	$SU(3)_c$	$SU(2)_W$	$U(1)_Y$
$\chi_{L/R}$	N	1	2	1/2
$\psi_{L/R}$	N	1	1	0

*N.Haba et al (2016,2017);
H.Ishida, S.M., Y.Yamaguchi (2016)*

☆ Then, how to generate “scales”, realize EWSB?
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 H.Ishida, S.M., S.Okawa, Y.Omura (2017)*

Extensions w/ Yukawa term & HC fermions ($F_{\{L/R\}}$)

$$\mathcal{L} = -y(\bar{\chi}H\psi) + \text{h.c.} \xrightarrow{\textcircled{\Lambda}_{\text{HC}}} -\Lambda_{\text{HC}}^2(y \cdot \Theta^\dagger H + |\Theta|^2) + \dots$$

	$SU(N)_{\text{HC}}$	$SU(3)_c$	$SU(2)_W$	$U(1)_Y$
$\chi_{L/R}$	N	1	2	1/2
$\psi_{L/R}$	N	1	1	0

$\Theta \sim \bar{\psi}\chi$ composite Higgs

Λ_{HC} :

“chiral” $SU(3)_{F_{L/R}}$ breaking scale

*N.Haba et al (2016,2017);
 H.Ishida, S.M., Y.Yamaguchi (2016)*

HC-quark condensate occurs at $\mu = \Lambda_{\text{HC}}$.

→ $\langle \bar{F}F \rangle = \langle \bar{\chi}\chi \rangle = \langle \bar{\psi}\psi \rangle \neq 0$ ($\langle \bar{\chi}\psi \rangle = \langle \bar{\psi}\chi \rangle = 0$ by Vafa-Witten's theorem)

$\bar{\psi}_L \chi_R$ and $\bar{\psi}_R \chi_L$ couple to a composite Higgs doublet $\Theta \sim \bar{\psi}\chi$

Then, “bosonic seesaw mechanism” can work:

$$\begin{aligned}
 & (H^\dagger, \Theta^\dagger) \begin{pmatrix} 0 & y\Lambda_{\text{HC}}^2 \\ y\Lambda_{\text{HC}}^2 & M_\Theta^2 \end{pmatrix} \begin{pmatrix} H \\ \Theta \end{pmatrix} \quad \begin{array}{l} \text{Classically scale invariance} \\ \text{HC-quark condensation} \end{array} \\
 & \approx (H_1^\dagger, H_2^\dagger) \begin{pmatrix} -\frac{y^2}{M_\Theta^2} \Lambda_{\text{HC}}^4 & 0 \\ 0 & M_\Theta^2 \end{pmatrix} \begin{pmatrix} H_1 \\ H_2 \end{pmatrix} \\
 & \text{where we have used } y \ll 1. \quad (M_\Theta \simeq \Lambda_{\text{HC}})
 \end{aligned}$$

Negative mass term is naturally generated: $m_H^2 \approx -\frac{y^2}{M_\Theta^2} \Lambda_{\text{HC}}^4$
 $\approx -(y\Lambda_{\text{HC}})^2$

☆ Constraint on scenarios of scalegenesis

~ asymptotic safety condition ~

-- scale/trace anomaly (quadratic, quartic divergence)
should be generated at quantum level

-- which may, however, be cancelled
by gravitational effects at Planck scale

*M. Shaposhnikov, et al (2010); C. Wetterich, et al (2016);
A. J. Helmboldt, et al (2016)*

--- Hence, to keep the classical scale-inv.
theory has to be safe up to Planck scale
(i.e. free from Landau poles)

☆ This talk introduces a **bosonic seesaw model**, shows that



* single **HC dynamics** explains

all masses for existing particles

-- SM particles including 125 GeV Higgs,

-- active (RH) neutrinos,

-- dark matter candidate (composite or elementary),
by EWSB and $U(1)_{\{B-L\}}$ breaking

* rich pheno. at TeV is predicted:

More interesting than other
scale-inv. models

-- HC hadrons (pions, scalar mesons, baryons,...)

-- B-L Higgs, gauge boson

-- a “new” production process for DM in thermal history,
DM detection exp. , etc.

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-- dynamical scalegeneration

2. Summary

1. Bosonic seesaw model w/ multiple seesaws

-- dynamical scalegeneration

Phys.RevD.95 (2017)

H.Ishida, S.M, S.Okawa, Y.Omura

* SM gauges x U(1)_{B-L}

x SU(3)_{HC} x CSI

		SU(3) _c	SU(2) _W	U(1) _Y	U(1) _{B-L}
quarks	q_L^α	3	2	1/6	-1/3
	u_R^α	3	1	2/3	-1/3
	d_R^α	3	1	-1/3	-1/3
leptons	l_L^α	1	2	-1/2	1
	e_R^α	1	1	-1	1
elementary EW Higgs	H	1	2	1/2	0
RHMv (3 generations)	N_R^α	1	1	0	1
elementary B-L Higgs	ϕ	1	1	0	-2

HC fermions (x4)

	$F_{L/R}$	SU(3) _{HC}	SU(3) _c	SU(2) _W	U(1) _Y	U(1) _{B-L}
SU(2) doublet	$\chi = (\chi_1, \chi_2)^T$	3	1	2	1/2 + q	q'
SU(2) singlet	ψ_1	3	1	1	q	q'
SU(2) singlet	ψ_2	3	1	1	q	-2 + q'

* Criterion for modeling HC sector:

$F_{L/R}$	$SU(3)_{\text{HC}}$	$SU(3)_c$	$SU(2)_W$	$U(1)_Y$	$U(1)_{B-L}$
$\chi = (\chi_1, \chi_2)^T$	3	1	2	$1/2 + q$	q'
ψ_1	3	1	1	q	q'
ψ_2	3	1	1	q	$-2 + q'$

Making “mirror” Higgses by HC strong dynamics

$$\Theta_1 \sim \bar{\psi}_1 \chi, \quad \Phi \sim \bar{\psi}_1 \psi_2,$$

$$H \sim (1, 2, 1/2, 0) \quad \phi \sim (1, 1, 0, -2)$$

to be seesaw partners via HC-fermion Yukawa terms:

$$y_H (\bar{\chi} H \psi_1 + \text{H.c.}) + y_\phi (\bar{\psi}_2 \phi \psi_1 + \text{H.c.})$$

* Dynamical mass generation: sequence of seesaws

- approximate “chiral”
 $SU(4) \times SU(4)$ sym. in HC sector
 (explicitly broken by gauges and
 HC-fermion Yukawa ints.)

$F_{L/R}$	$SU(3)_{HC}$	$SU(3)_c$	$SU(2)_W$	$U(1)_Y$	$U(1)_{B-L}$
$\chi = (\chi_1, \chi_2)^T$	3	1	2	$1/2 + q$	q'
ψ_1	3	1	1	q	q'
ψ_2	3	1	1	q	$-2 + q'$

up to HC-fermion axial U(1) (anomaly)

- spontaneously broken at Λ_{HC} (say, O(TeV) – O(10TeV))

$$\langle \bar{F}^i F^j \rangle \sim \Lambda_{HC}^3 \delta^{ij} \quad SU(4)_{FL} \times SU(4)_{FR} \rightarrow SU(4)_{FV}$$

- as well as CSI, simultaneously

* Dynamical mass generation: sequence of seesaws

- theory below HC scale, described by **HC hadrons**
- (ref. QCD) low-lying spectra include:

- **HC pions** x 15 (pNG bosons) {+ HC eta'}
- **HC scalar mesons** x 16
- **HC spin 1/2 baryons** x 20 (\rightarrow incl. DM), {and vector (axialvector) mesons}

$\mathcal{M} = \mathcal{S} + i\mathcal{P}$	constituent	$SU(2)_W$	$U(1)_Y$	$U(1)_{B-L}$
$(f_0^{\text{HC}}, a_0^{\text{HC}} + iP_{a_0^{\text{HC}}})_{ij}$	$\bar{\chi}_i \chi_j$	(1, 3)	0	0
$(\Theta_1 + iP_{\Theta_1})_i$	$\bar{\psi}_1 \chi_i$	2	1/2	0
$(\Theta_2 + iP_{\Theta_2})_i$	$\bar{\psi}_2 \chi_i$	2	1/2	2
$\Phi + iP_\Phi$	$\bar{\psi}_1 \psi_2$	1	0	-2
$\varphi_1 + iP_{\varphi_1}$	$\bar{\psi}_1 \psi_1$	1	0	0
$\varphi_2 + iP_{\varphi_2}$	$\bar{\psi}_2 \psi_2$	1	0	0

HC baryon	constituent	$SU(2)_W$	I_3	Y	$B-L$
p_{HC}^{2+3q}	$\chi_1 \chi_1 \chi_2$	2	1/2	$3/2 + 3q$	$3q'$
n_{HC}^{1+3q}	$\chi_1 \chi_2 \chi_2$	2	-1/2	$3/2 + 3q$	$3q'$
$\Lambda_{(1)}^{1+3q}$	$\chi_1 \chi_2 \psi_1$	1	0	$1 + 3q$	$3q'$
$\Sigma_{(1)}^{2+3q}$	$\chi_1 \chi_1 \psi_1$	3	1	$1 + 3q$	$3q'$
$\Sigma_{(1)}^{1+3q}$	$\chi_1 \chi_2 \psi_1$	3	0	$1 + 3q$	$3q'$
$\Sigma_{(1)}^{3q}$	$\chi_2 \chi_2 \psi_1$	3	-1	$1 + 3q$	$3q'$
$\Xi_{(11)}^{1+3q}$	$\chi_1 \psi_1 \psi_1$	2	1/2	$1/2 + 3q$	$3q'$
$\Xi_{(11)}^{3q}$	$\chi_2 \psi_1 \psi_1$	2	-1/2	$1/2 + 3q$	$3q'$
$\Lambda_{(2)}^{1+3q}$	$\chi_1 \chi_2 \psi_2$	1	0	$1 + 3q$	$-2 + 3q'$
$\Omega_{(12)}^{3q}$	$\psi_1 \psi_1 \psi_2$	1	0	$3q$	$-2 + 3q'$
$\Sigma_{(2)}^{2+3q}$	$\chi_1 \chi_1 \psi_2$	3	1	$1 + 3q$	$-2 + 3q'$
$\Sigma_{(2)}^{1+3q}$	$\chi_1 \chi_2 \psi_2$	3	0	$1 + 3q$	$-2 + 3q'$
$\Sigma_{(2)}^{3q}$	$\chi_2 \chi_2 \psi_2$	3	-1	$1 + 3q$	$-2 + 3q'$
$\Xi_{(12)}^{1+3q}$	$\chi_1 \psi_1 \psi_2$	2	1/2	$1/2 + 3q$	$-2 + 3q'$
$\Xi_{(12)}^{3q}$	$\chi_2 \psi_1 \psi_2$	2	-1/2	$1/2 + 3q$	$-2 + 3q'$
$\Omega_{(22)}^{3q}$	$\psi_1 \psi_2 \psi_2$	1	0	$3q$	$-4 + 3q'$
$\Xi_{(22)}^{1+3q}$	$\chi_1 \psi_2 \psi_2$	2	1/2	$1/2 + 3q$	$-4 + 3q'$
$\Xi_{(22)}^{3q}$	$\chi_2 \psi_2 \psi_2$	2	-1/2	$1/2 + 3q$	$-4 + 3q'$

* Dynamical mass generation: sequence of seesaws

Among HC hadrons, focus on “mirror” composite Higgses

$$\Theta_1 \sim \bar{\psi}_1 \chi, \quad \Phi \sim \bar{\psi}_1 \psi_2,$$

via HC-fermion Yukawa terms:

$$y_H(\bar{\chi} H \psi_1 + \text{H.c.}) + y_\phi(\bar{\psi}_2 \phi \psi_1 + \text{H.c.})$$

trigger two seesaws:
$$\begin{pmatrix} 0 & y_{H/\phi} \Lambda_{\text{HC}}^2 \\ y_{H/\phi} \Lambda_{\text{HC}}^2 & \Lambda_{\text{HC}}^2 \end{pmatrix} \quad y_{H/\phi} \ll 1$$

$m_H^2 \simeq -y_H^2 \Lambda_{\text{HC}}^2$ and $m_\phi^2 \simeq -y_\phi^2 \Lambda_{\text{HC}}^2$ realize EWSB & U(1)B-L breaking w/
 $-\lambda_H(H^\dagger H)^2 - \lambda_\phi(|\phi|^2)^2$

-- physical Higgs masses

$$m_{h_1} \simeq \sqrt{2\lambda_H} v_{\text{EW}} \simeq 125 \text{ GeV}$$

$$m_{\phi_1} \simeq 2\sqrt{2\lambda_\phi} v_{\phi_1} \\ = \mathcal{O}(10) \text{ TeV}$$

$$v_\phi = \mathcal{O}(\Lambda_{\text{HC}}) \\ \text{and } \lambda_\phi = \mathcal{O}(1)$$

* Dynamical mass generation: sequence of seesaws

Furthermore, via RH Majorana ν -Yukawa terms:

$$y_{IN}^{\alpha\beta} \overline{l_L^\alpha} \tilde{H} N_R^\beta + y_N^{\alpha\alpha} (\phi \overline{N_R^{c\alpha}} N_R^\alpha + \text{H.c.})$$

trigger neutrino seesaw (Type I):

$$\begin{pmatrix} 0 & y_{IN} v_{EW} \\ y_{IN}^T v_{EW} & m_{N_R} \end{pmatrix}$$

-- active and heavy RHM neutrino masses

$$m_\nu \simeq y_{IN}^2 v_{EW}^2 / m_{N_R} = \mathcal{O}(0.1 \text{eV}) \quad \text{for } y_{IN} = \mathcal{O}(10^{-5}).$$

$$m_{N_R}^{\alpha\alpha} = \mathcal{O}(y_N^{\alpha\alpha} \Lambda_{\text{HC}}) = \mathcal{O}(5-10 \text{TeV}) \quad \text{with } y_N^{\alpha\alpha} = \mathcal{O}(1)$$
$$v_\phi = \mathcal{O}(\Lambda_{\text{HC}})$$

DM candidates

	HC baryon	constituent	$SU(2)_W$	I_3	Y	$B - L$
(I)	$\Lambda_{(1)}^{1+3q}$	$\chi_1 \chi_2 \psi_1$	1	0	$1 + 3q$	$3q'$
	$\Lambda_{(2)}^{1+3q}$	$\chi_1 \chi_2 \psi_2$	1	0	$1 + 3q$	$-2 + 3q'$
(II)	$\Omega_{(12)}^{3q}$	$\psi_1 \psi_1 \psi_2$	1	0	$3q$	$-2 + 3q'$
	$\Omega_{(22)}^{3q}$	$\psi_1 \psi_2 \psi_2$	1	0	$3q$	$-4 + 3q'$

DM candidate = SM-gauge singlet, lightest HC baryons (TeV- O(10) TeV)

(I) $q = -1/3$ and (II) $q = 0$,

* This charge choice is consistent w/ asymptotic safety

Other baryons decay to DMs + weak bosons(*) + HC pions (*)

$$p_{\text{HC}}^+ \rightarrow n_{\text{HC}}^0 + W^{+(*)} \quad n_{\text{HC}}^0 \rightarrow \mathcal{P}_{\Theta_2}^0 + \Lambda_{(2)}^0$$

DM production in early universe

* B-L singlet DM,
for simplicity

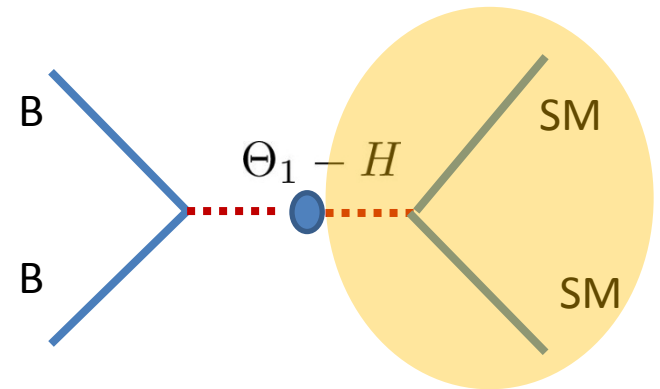
- lightest HC baryons ($B = \Lambda_{(1)}^0 (q' = 0)$ or $\Lambda_{(2)}^0 (q' = 2/3)$ for $q = -1/3$ case I)
- generated @ $T = \text{HC baron mass scale } O(1) - O(10) \text{ TeV}$:
DM=nonrelativistic; might have never been thermalized
(* HC pions can be as heavy as HC baryons;
HC pions might have never been in thermal EQ as well.)

- dominant source for DM production
= bosonic seesaw portal process

$$\frac{1}{\Lambda_{\text{HC}}} \bar{B} B \Theta_1^\dagger \Theta_1 \rightarrow \frac{y_H^2}{\Lambda_{\text{HC}}} \bar{B} B H^\dagger H$$

$\Theta_1 - H$ mixing $y_H \ll 1$

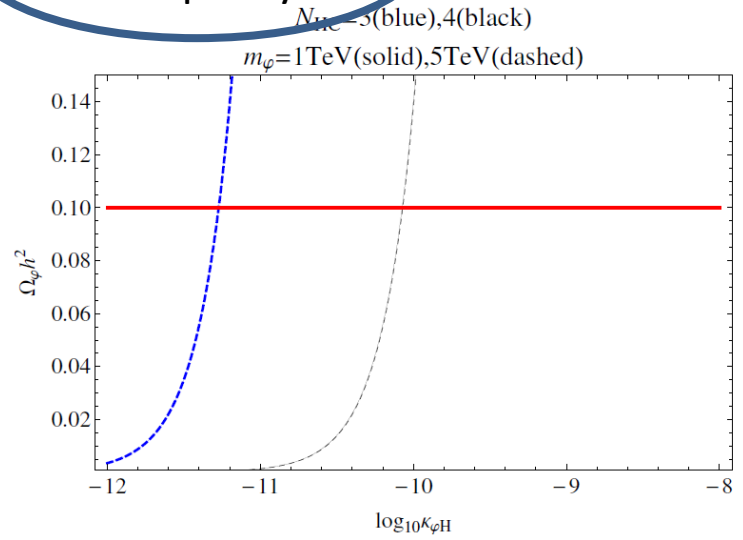
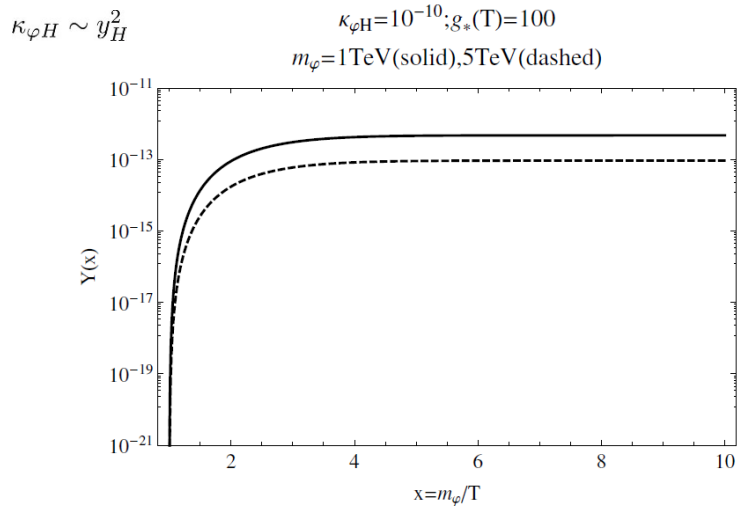
H. Ishida, S.M, Y. Yamaguchi (2016)



- non-thermal production like “freeze-in” scenario (FIMP, E-WIMP)

DM production in early universe

* B-L singlet DM, for simplicity

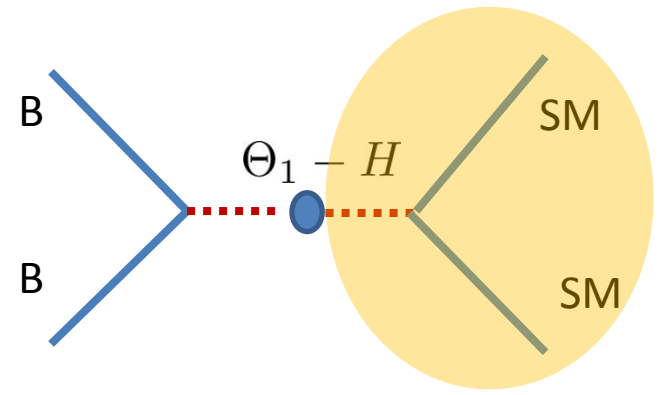


-- dominant source for DM production
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H. Ishida, S.M, Y. Yamaguchi (2016)

$$\frac{1}{\Lambda_{\text{HC}}} \bar{B} B \Theta_1^\dagger \Theta_1 \rightarrow \frac{y_H^2}{\Lambda_{\text{HC}}} \bar{B} B H^\dagger H$$

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-- non-thermal production like “freeze-in” scenario (FIMP, E-WIMP)

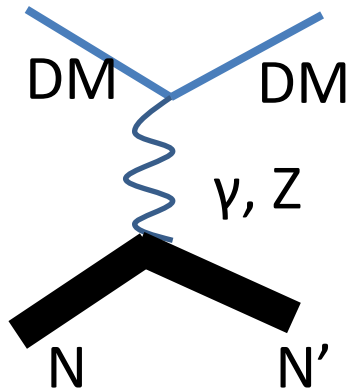
@ DM detection experiments

(I) $q = -1/3$ and (II) $q = 0$,

Ref: $q' = 0$

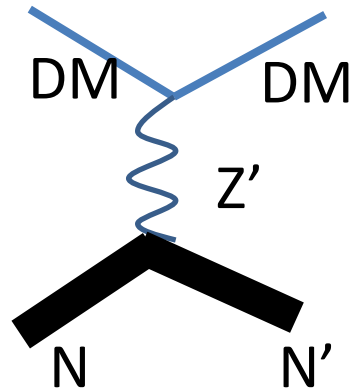
Ref: $q' = 1$

$$\frac{g \cdot g_Y}{2m_{\Lambda^0}} \bar{\Lambda}^0 \sigma_{\mu\nu} B^{\mu\nu} \Lambda^c$$



mag.moment int.
w/ g-factor: $g \sim 1$

$$g_{B-L} \bar{\Omega}^0 \gamma_\mu Z'^\mu \Omega^0$$



B-L gauge int.:
 $g_{\{B-L\}^2/m_{Z'}^2}$
 $= 1/v_\phi^2$

preference to be clarified
in future exps.

$F_{L/R}$	$SU(3)_{HC}$	$SU(3)_c$	$SU(2)_W$	$U(1)_Y$	$U(1)_{B-L}$
$\chi = (\chi_1, \chi_2)^T$	3	1	2	$1/2 + q$	q'
ψ_1	3	1	1	q	q'
ψ_2	3	1	1	q	$-2 + q'$

(I)

(II)

HC baryon	constituent	$SU(2)_W$	I_3	Y	$B-L$
$\Lambda_{(1)}^{1+3q}$	$\chi_1 \chi_2 \psi_1$	1	0	$1 + 3q$	$3q'$
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Current bounds on SI cross sections
(LUX2016&XENON1T)

(I) $m_{DM} > \mathcal{O}(10)$ TeV

(II) $m_{DM} > \mathcal{O}(10)$ TeV for $v_\phi > 5$ TeV

3. Summary



- * Bosonic seesaw models:
 - interesting scenario giving dynamical explanation for origin of mass, solving gauge hierarchy problem

 - can provide rich pheno. for TeV scale physics
HC composite spectra: incl natural DM candidate

- * We discussed a dynamical scalegenesis:
 - **all particle masses generated by a single HC**

- * Future direction
 - HC hadron collider signatures at LHC
 - effect on flavor physics
 - GW detectability of HC phase transition, etc.
 - under consideration