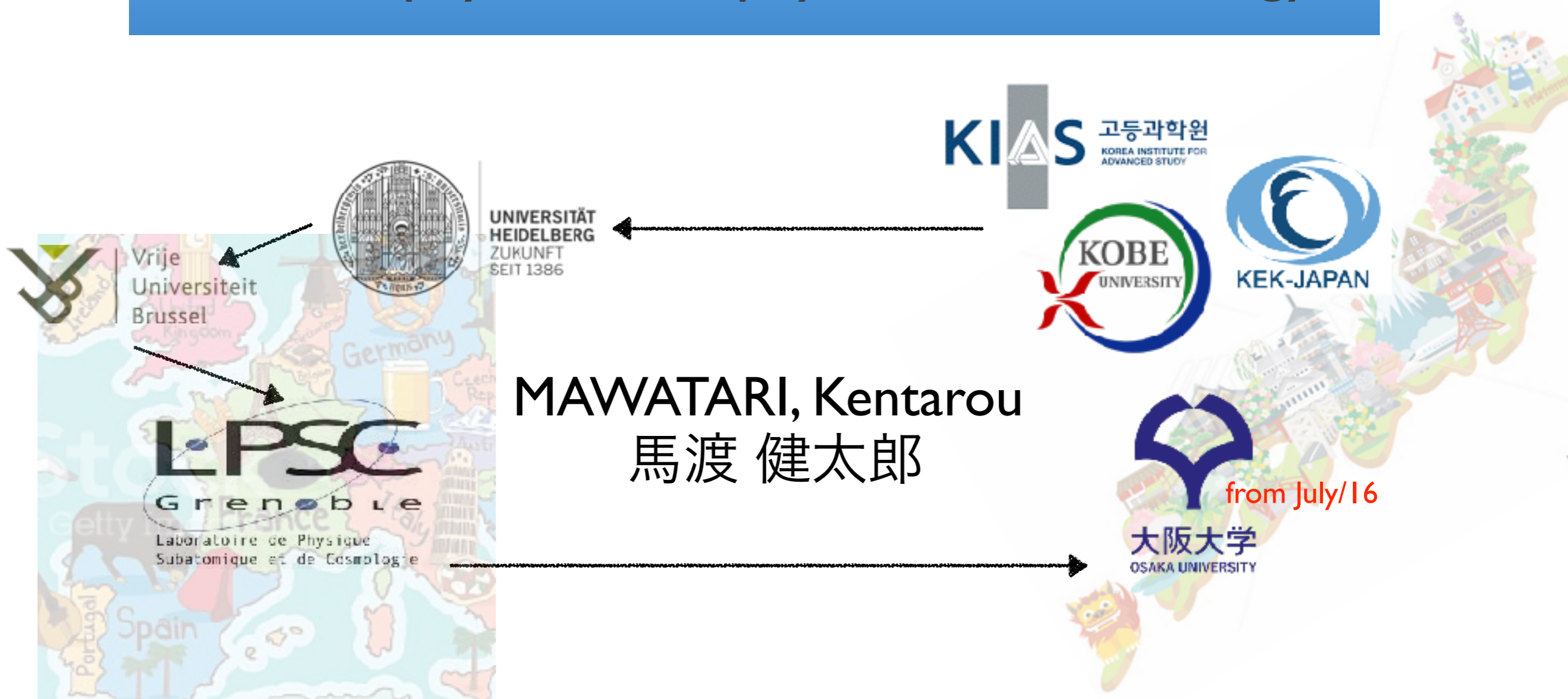
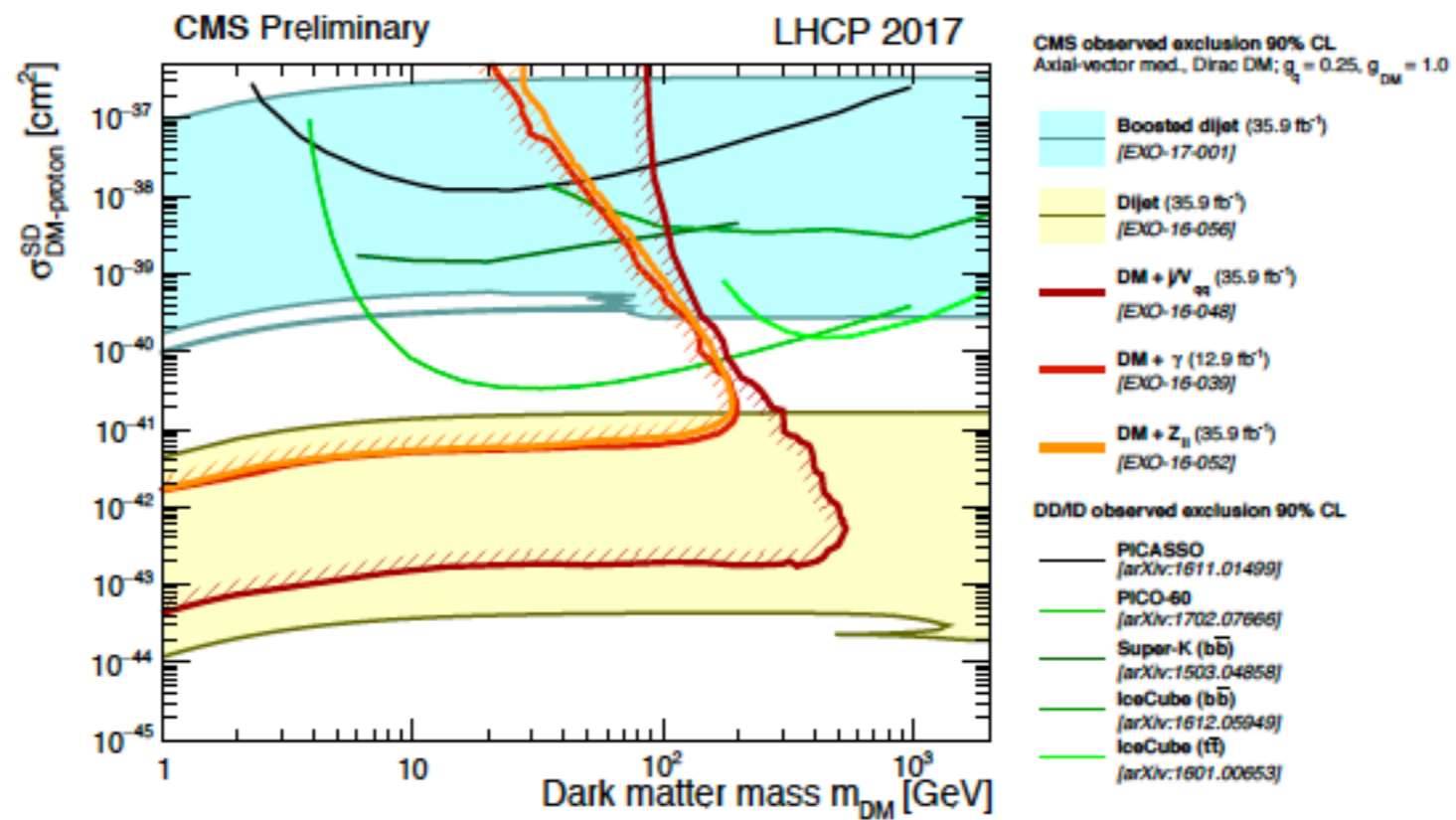
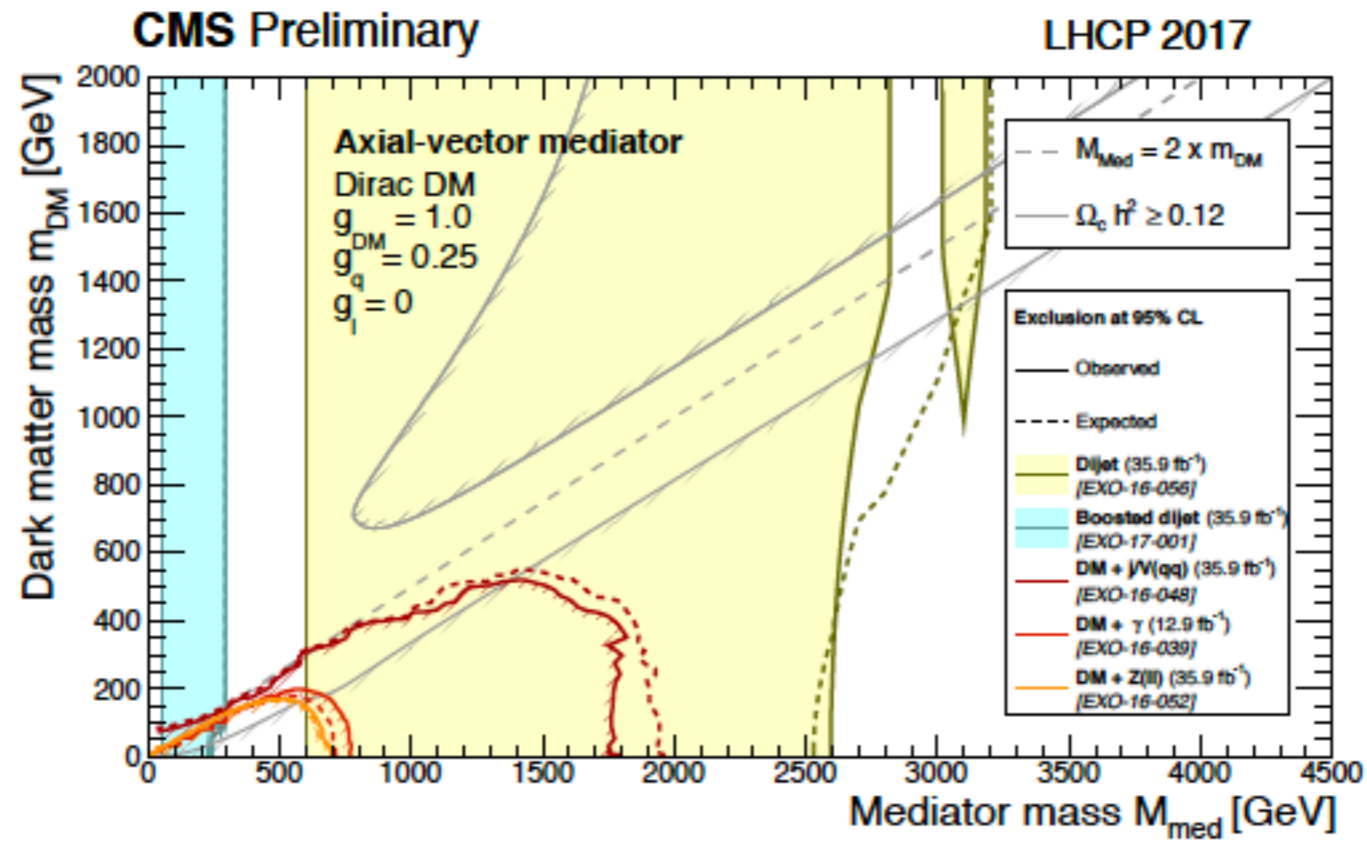
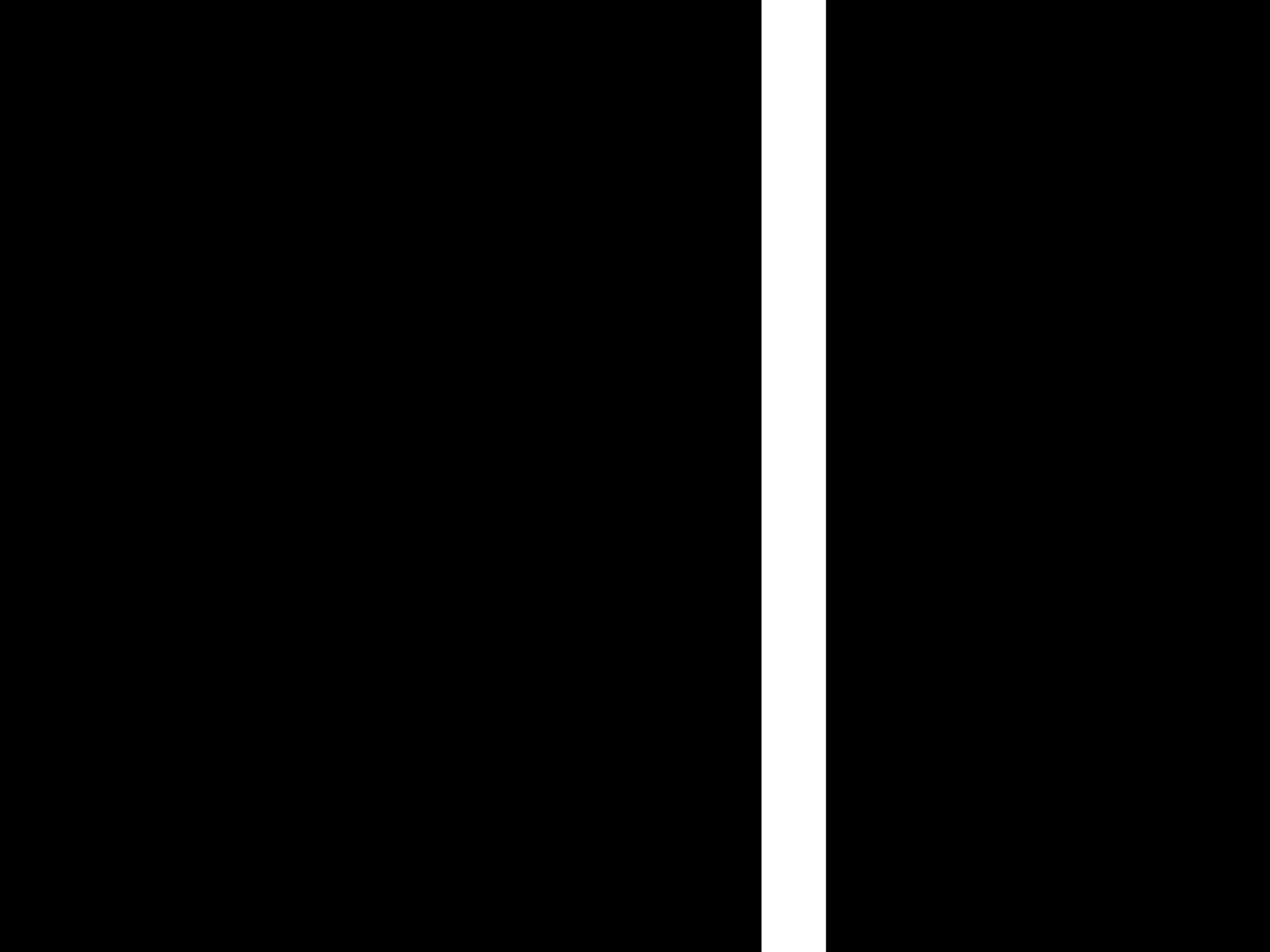


# Seeking dark matter at the interface of collider physics, astrophysics and cosmology



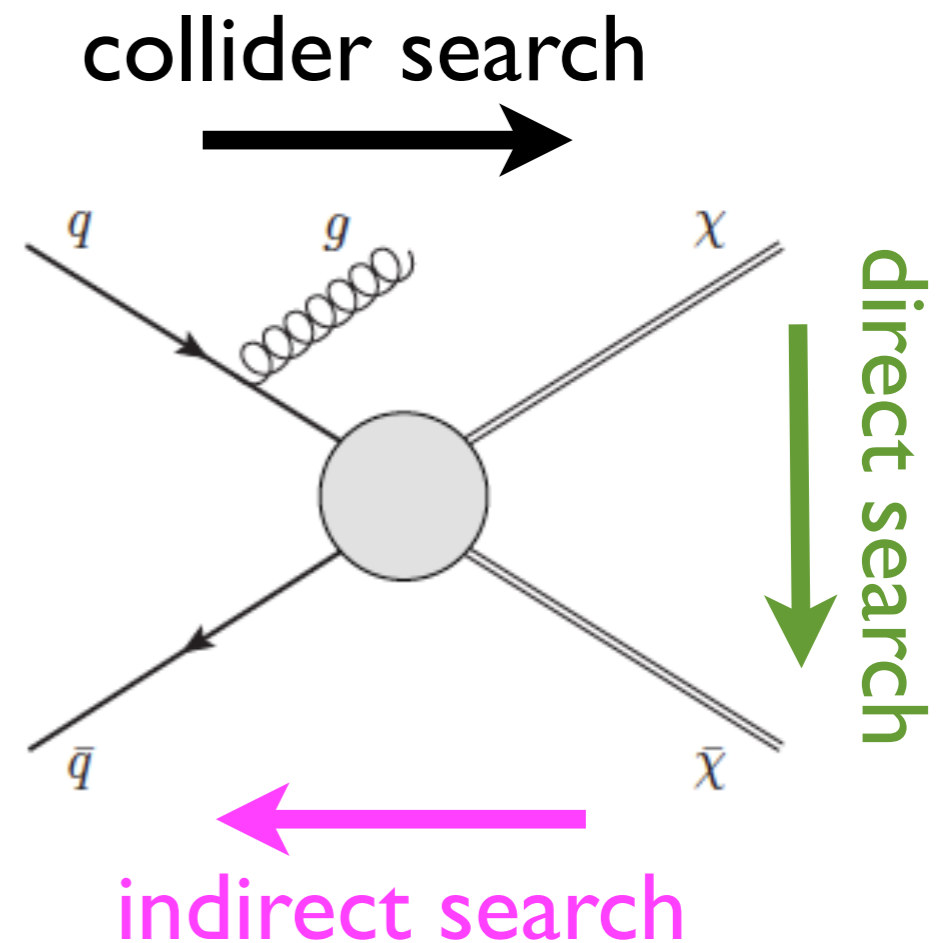
# LHC Dark Matter Working Group (LHC DMWG)



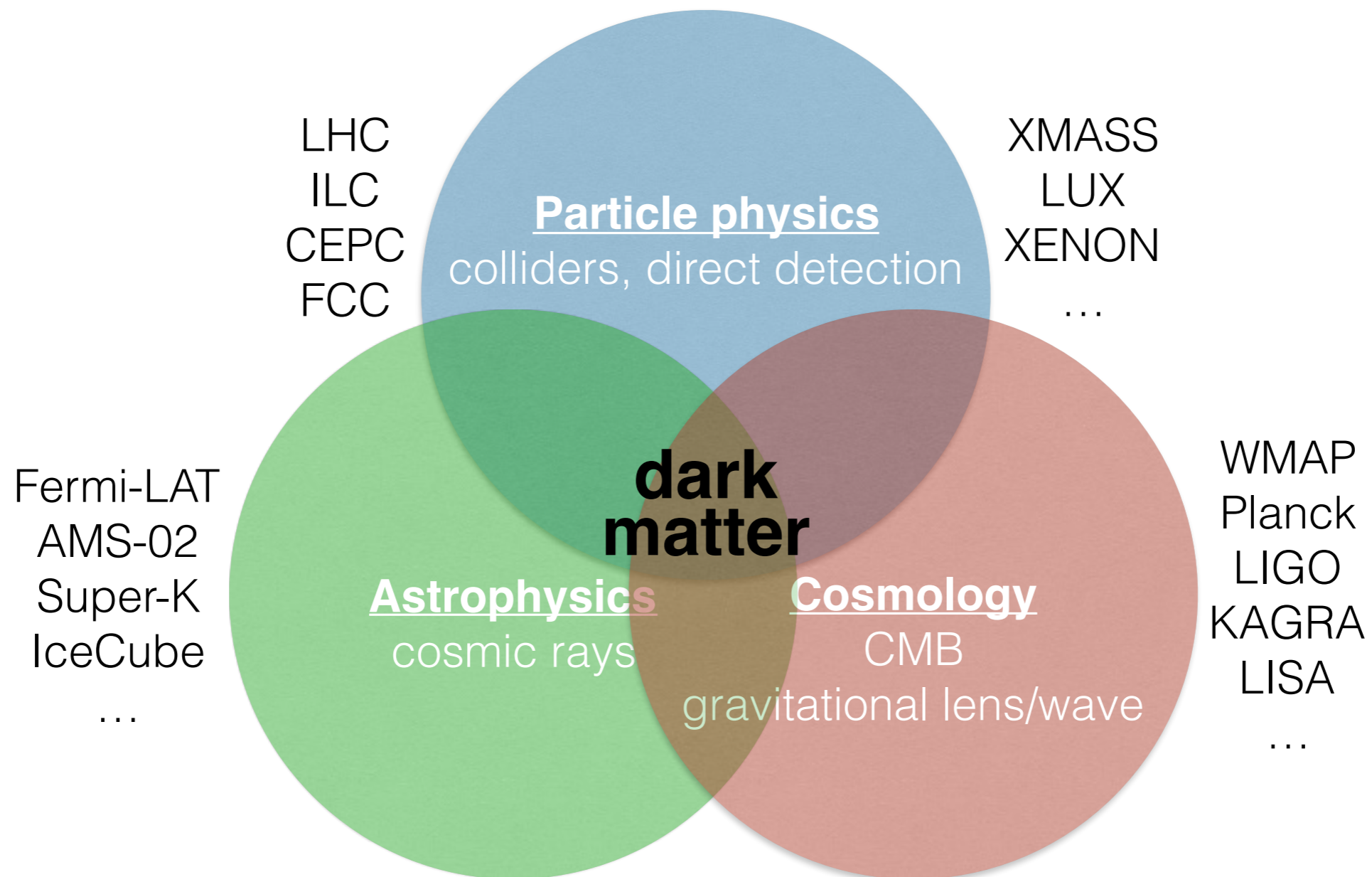


# Looking for dark matter

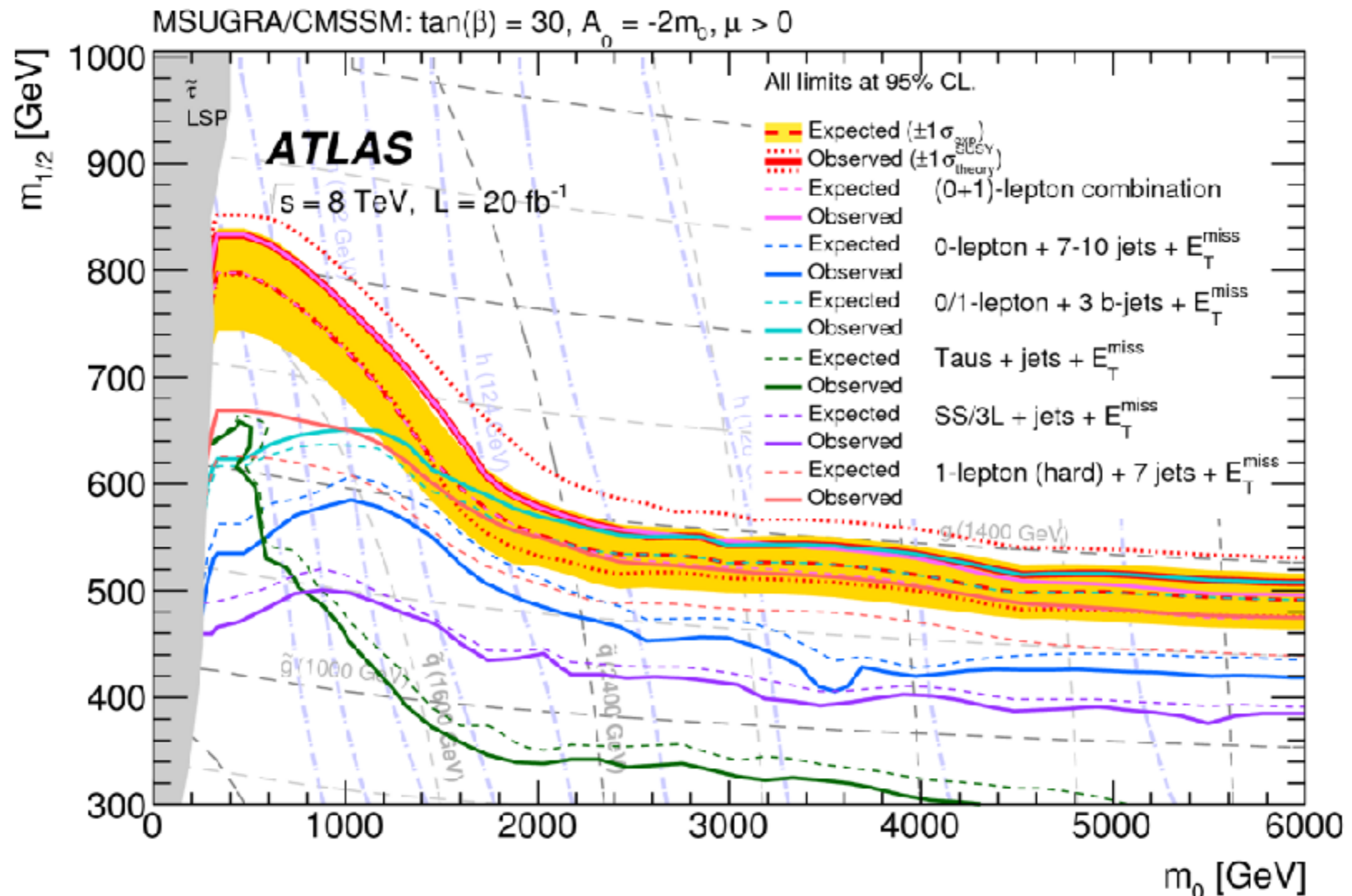
- If dark matter has non-gravitational interactions with ordinal matter, we can observe it directly and/or indirectly.
- Indeed, many types of dark matter search experiments are currently on-going all over the world.



# Dark matter Venn diagram



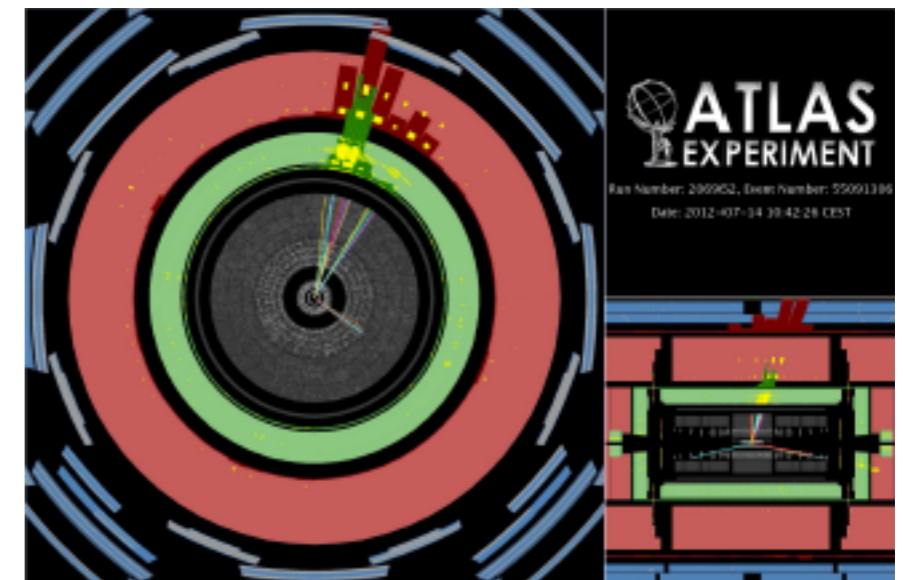
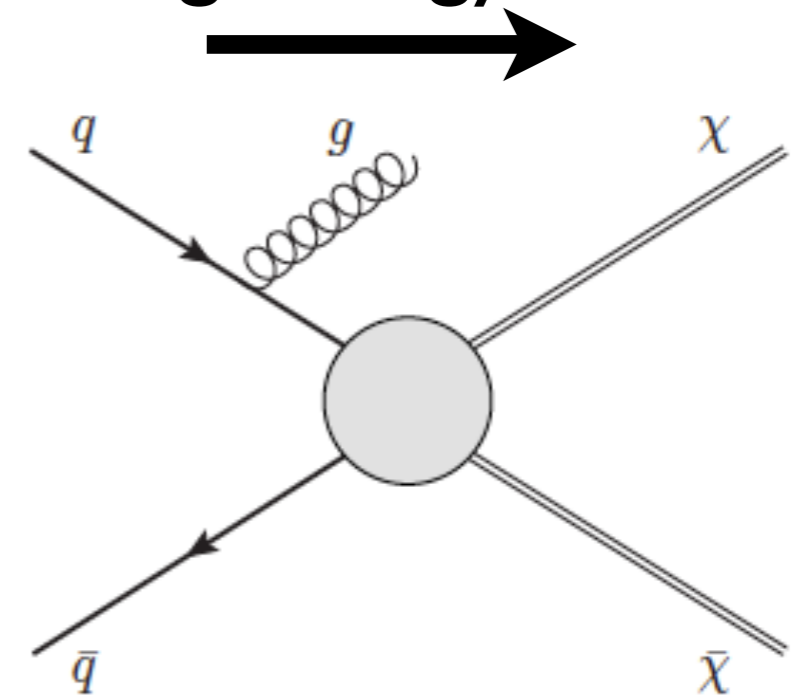
# Dark matter searches at LHC Run-I: top-down (e.g. Supersymmetry)



# Dark matter searches at LHC Run-I: bottom-up (Effective field theory)

- employed contact interaction operators in EFTs (effective field theories).
- vector  $\frac{1}{M_\star^2} \bar{\chi} \gamma^\mu \chi \bar{q} \gamma_\mu q$
- axial-vector  $\frac{1}{M_\star^2} \bar{\chi} \gamma^\mu \gamma^5 \chi \bar{q} \gamma_\mu \gamma^5 q$
- The signal is determined by the Lorentz structure, the DM mass, and the overall coupling (or the cutoff scale).
- easy interpretation to non-collider DM searches
- EFT validation;  $M_\star \cong$  (LHC accessible energy)

missing-energy search





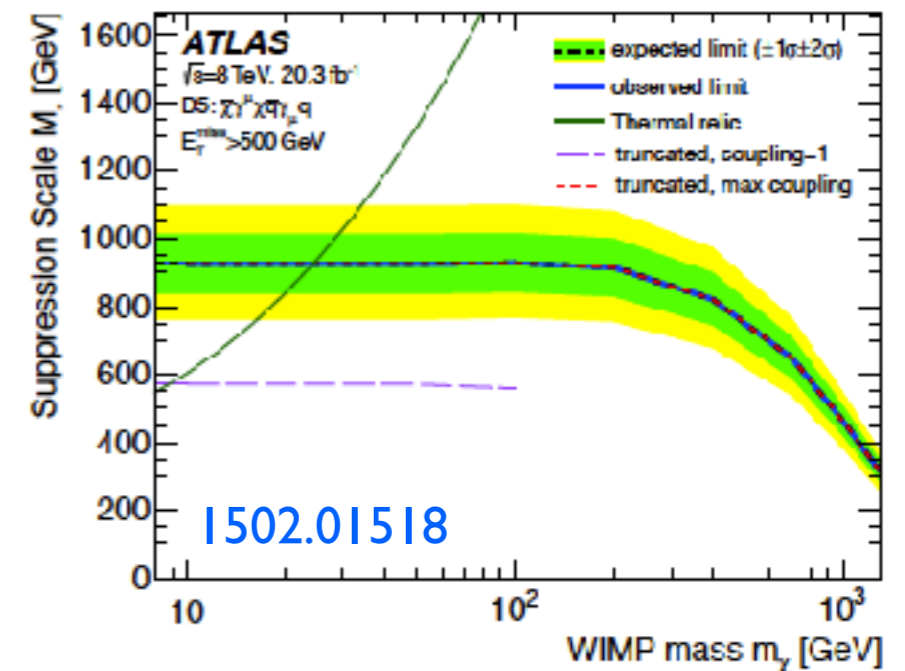
# Dark matter searches at LHC Run-I: bottom-up (Effective field theory)

- employed contact interaction operators in EFTs (effective field theories).

- vector  $\frac{1}{M_\star^2} \bar{\chi} \gamma^\mu \chi \bar{q} \gamma_\mu q$

- axial-vector  $\frac{1}{M_\star^2} \bar{\chi} \gamma^\mu \gamma^5 \chi \bar{q} \gamma_\mu \gamma^5 q$

- The signal is determined by the Lorentz structure, the DM mass, and the overall coupling (or the cutoff scale).
- easy interpretation to non-collider DM searches
- EFT validation;  $M_\star \cong$  (LHC accessible energy)



# LPCC: LHC Physics Centre at CERN

[Welcome](#)
[About](#)
[LHC working groups](#)
[LHC publications](#)
[Events](#)
[Newsletter](#)

## LHC DM WG: WG on Dark Matter Searches at the LHC

[LHC WORKING GROUPS](#)
[Dark Matter WG](#)
[WG Meetings](#)
[WG documents](#)
[Electroweak WG](#)
[WG Documents](#)
[WG meetings](#)
[Forward Physics WG](#)
[WG TWIKI PAGE](#)
[WG documents](#)
[WG meetings](#)
[Heavy Flavour WG](#)
[WG Documents](#)
[WG Meetings](#)
[MB & UE WG](#)
[WG meetings](#)
[WG documents](#)
[Machine Learning](#)

To subscribe to the general WG mailing list, used to distribute announcements about meetings and available documents, go to

<http://simba3.web.cern.ch/simba3/SelfSubscription.aspx?groupName=lhc-dmwg>

A second mailing list is used for more technical exchanges related to the ongoing work of the WG. To subscribe, go to

<http://simba3.web.cern.ch/simba3/SelfSubscription.aspx?groupName=lhc-dmwg-contributors>

The LHC Dark Matter Working Group (LHC DM WG) brings together theorists and experimentalists to define guidelines and recommendations for the benchmark models, interpretation, and characterisation necessary for broad and systematic searches for dark matter at the LHC. As examples, the group develops and promotes well-defined signal models, specifying the assumptions behind them and describing the conditions under which they should be used. It works to improve the set of tools available to the experiments, such as higher-precision calculations of the backgrounds. It assists theorists with understanding and making use of LHC results.

The LHC DM WG develops and maintains close connections with theorists and other experimental particle DM searches (e.g. Direct and Indirect Detection experiments) in order help verify and constrain particle physics models of astrophysical excesses, to understand how collider searches and non-collider experiments complement one another, and to help build a comprehensive understanding of viable dark matter models.

# Search for dark matter produced with an energetic jet or a hadronically decaying $W$ or $Z$ boson at $\sqrt{s} = 13$ TeV

The CMS Collaboration\*

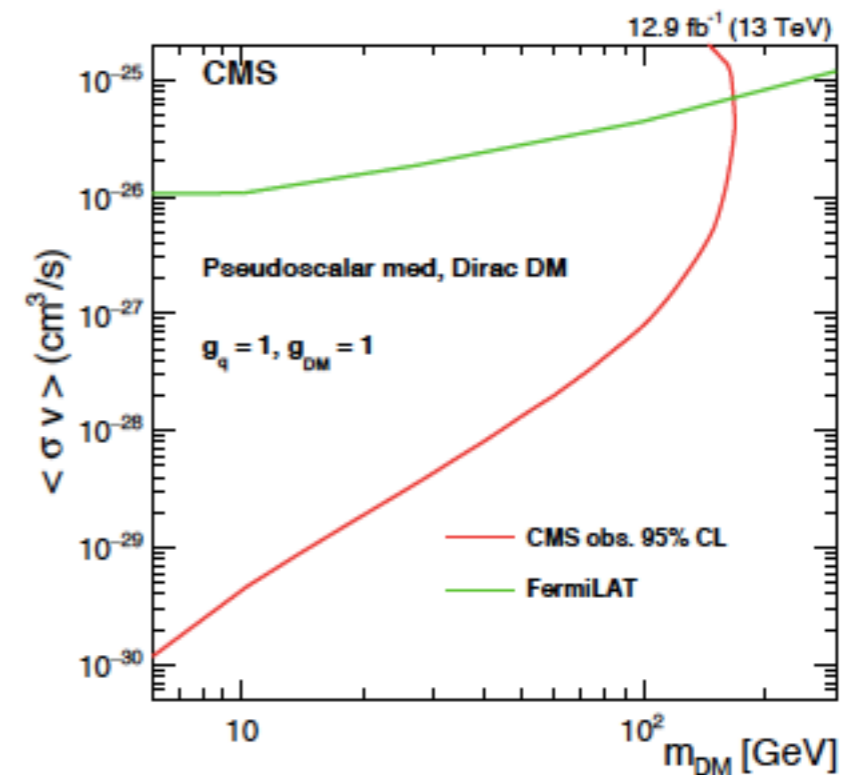
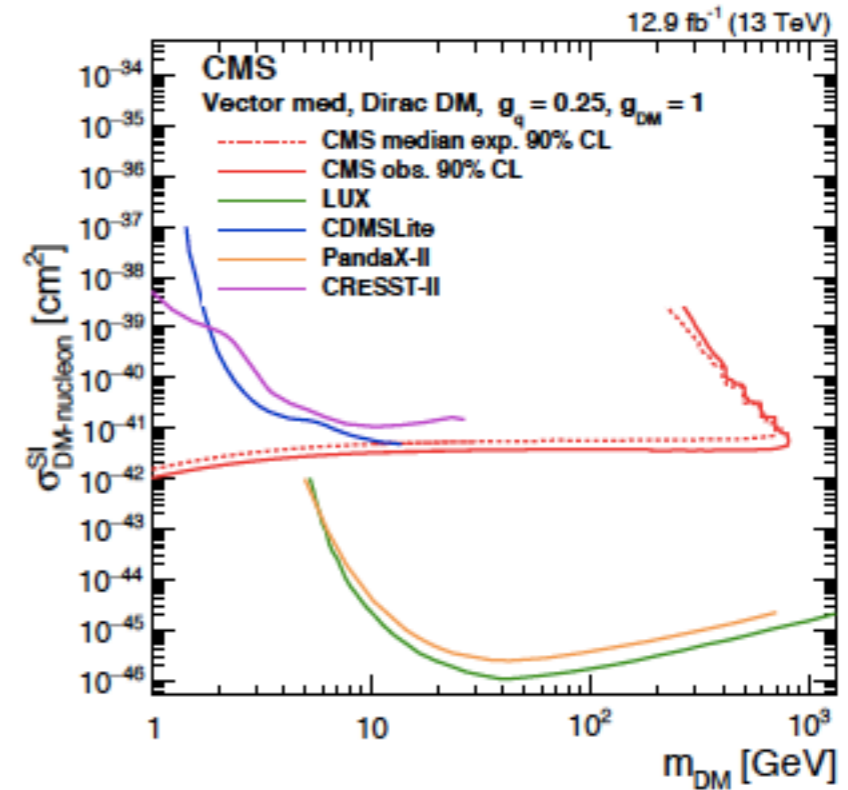
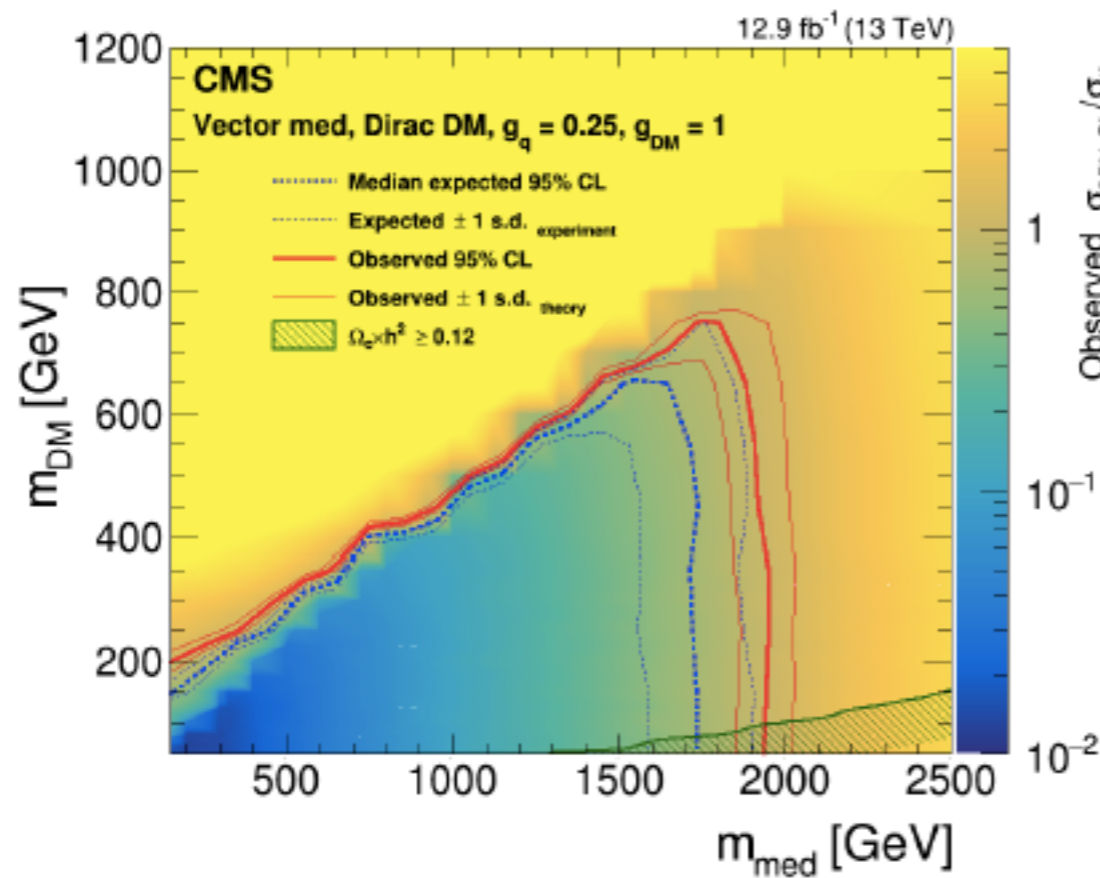
## Abstract

A search for dark matter particles is performed using events with large missing transverse momentum, at least one energetic jet, and no leptons, in proton-proton collisions at  $\sqrt{s} = 13$  TeV collected with the CMS detector at the LHC. The data sample

...

ing fraction. The results of this search provide the strongest constraints on the dark matter pair production cross section through vector and axial-vector mediators at a particle collider. When compared to the direct detection experiments, the limits obtained from this search provide stronger constraints for dark matter masses less than 5, 9, and 550 GeV, assuming vector, scalar, and axial-vector mediators, respectively. The search yields stronger constraints for dark matter masses less than 200 GeV, assuming a pseudoscalar mediator, when compared to the indirect detection results from Fermi-LAT.

# Constraints from LHC, direct, and indirect searches



# DM searches at LHC Run-II

LHC DMWG 1507.00966

- is employing simplified DM models.

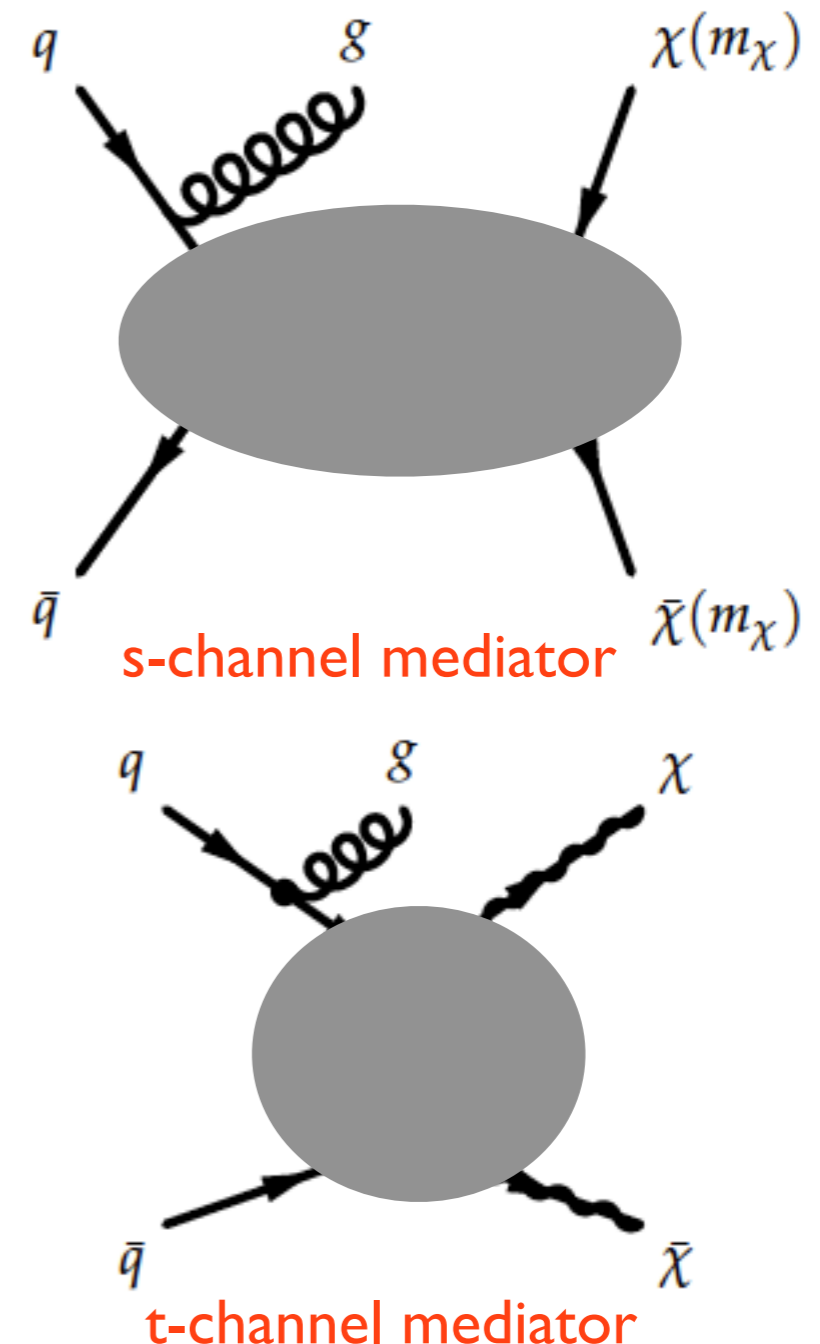
- $$\mathcal{L}_{\text{vector}} = g_q \sum_{q=u,d,s,c,b,t} Z'_\mu \bar{q} \gamma^\mu q + g_\chi Z'_\mu \bar{\chi} \gamma^\mu \chi$$

- $$\mathcal{L}_{\text{axial-vector}} = g_q \sum_{q=u,d,s,c,b,t} Z'_\mu \bar{q} \gamma^\mu \gamma^5 q + g_\chi Z'_\mu \bar{\chi} \gamma^\mu \gamma^5 \chi$$

- The signal is determined by the mediator type, the DM and mediator masses, and the two couplings.

- Richer phenomenology

- Interpretations to non-collider DM searches are complicated.



# s-channel simplified DM models

LHC DMWG [1507.00966, 1603.04156]

- Simplified DM models (s-channel):

- spin-1 mediator

$$\mathcal{L}_{\text{vector}} = -g_{\text{DM}} Z'_\mu \bar{\chi} \gamma^\mu \chi - g_q \sum_{q=u,d,s,c,b,t} Z'_\mu \bar{q} \gamma^\mu q,$$

$$\mathcal{L}_{\text{axial-vector}} = -g_{\text{DM}} Z'_\mu \bar{\chi} \gamma^\mu \gamma_5 \chi - g_q \sum_{q=u,d,s,c,b,t} Z'_\mu \bar{q} \gamma^\mu \gamma_5 q.$$

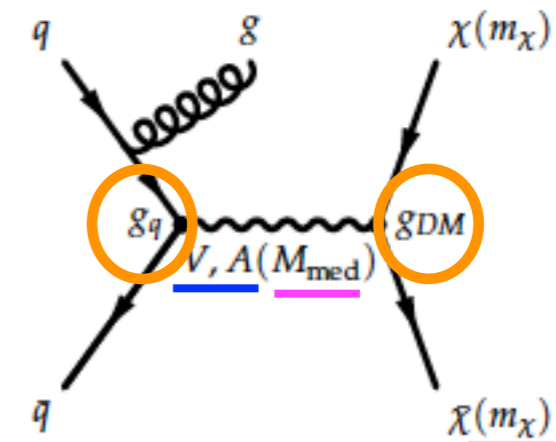
- spin-0 mediator

$$\mathcal{L}_{\text{scalar}} = -g_{\text{DM}} \phi \bar{\chi} \chi - g_q \frac{\phi}{\sqrt{2}} \sum_{q=u,d,s,c,b,t} y_q \bar{q} q,$$

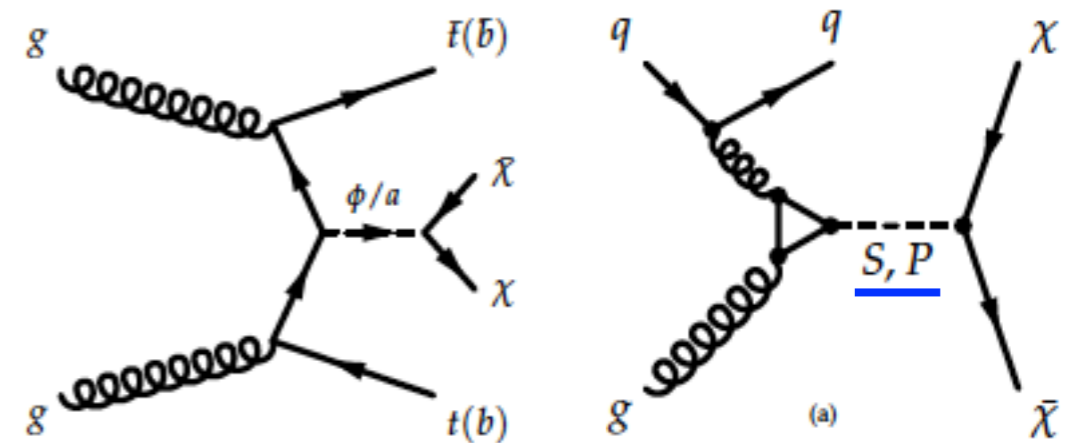
$$\mathcal{L}_{\text{pseudo-scalar}} = -ig_{\text{DM}} \phi \bar{\chi} \gamma_5 \chi - ig_q \frac{\phi}{\sqrt{2}} \sum_{q=u,d,s,c,b,t} y_q \bar{q} \gamma_5 q,$$

- The signal is determined by

- the mediator type (V, A, S, P)
  - the DM and mediator masses
  - the two couplings



spin-1 mediator



spin-0 mediator

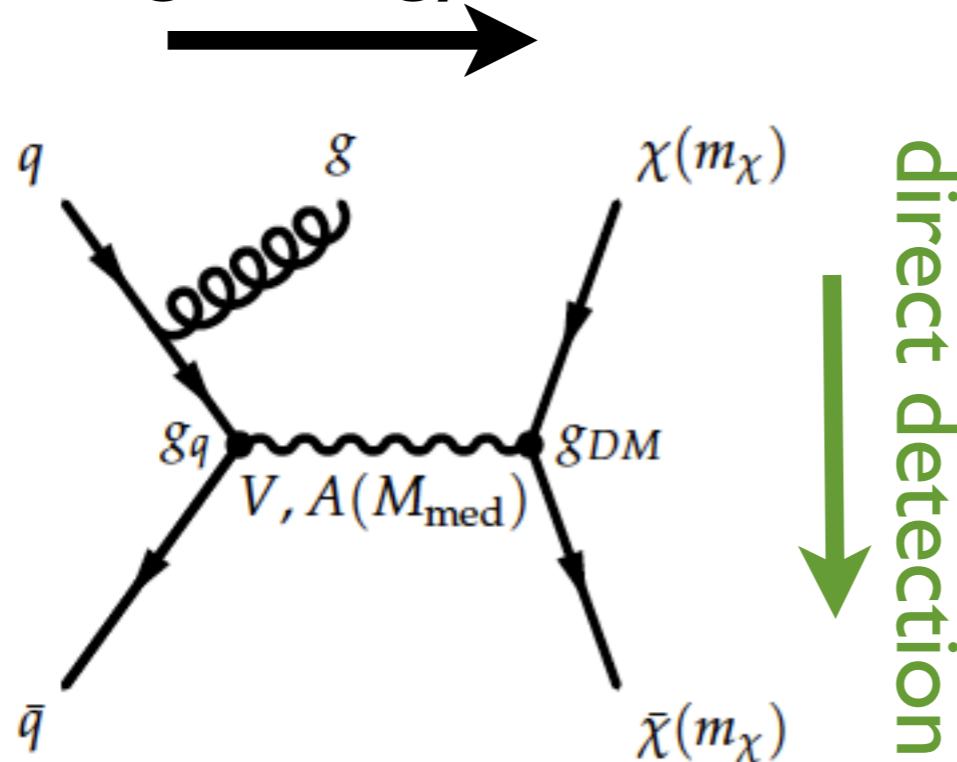
👉 Top-philic DM models

# Signatures of simplified DM models

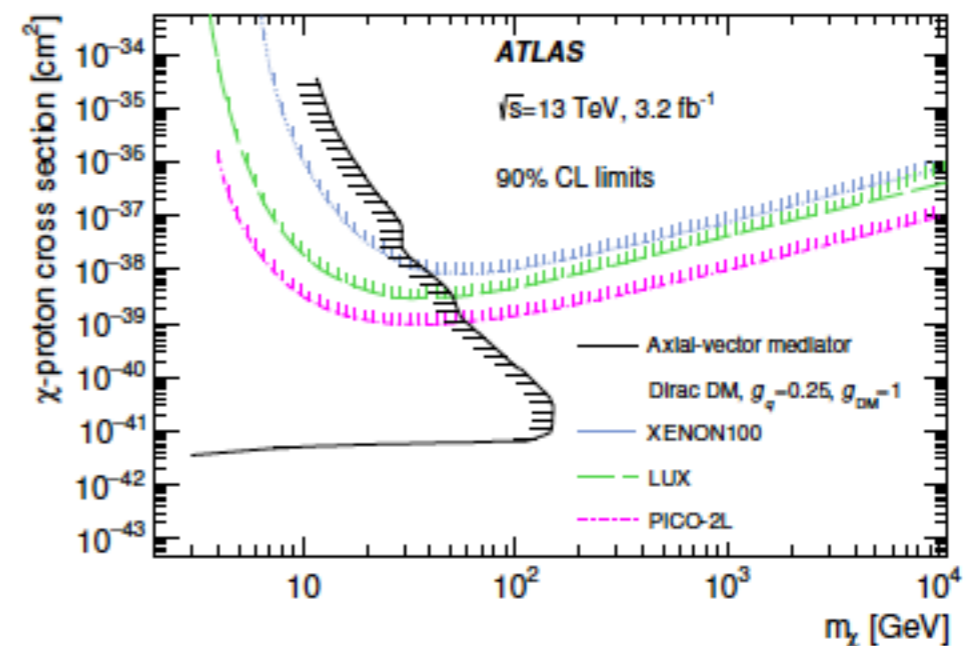
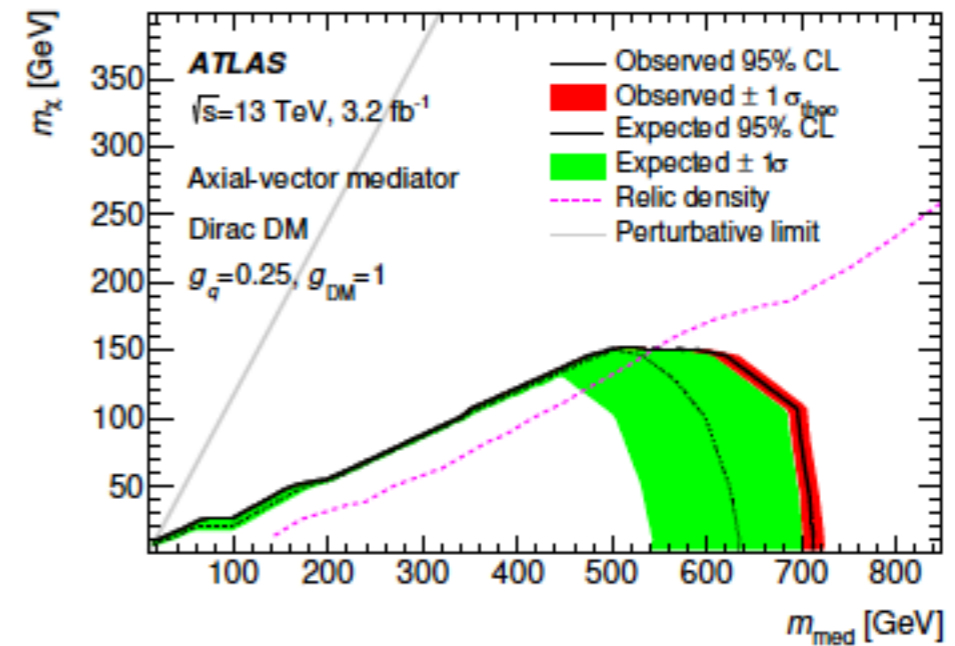
LHC DMWG [1507.00966, 1603.04156]

[ATLAS 1604.01306]

missing-energy search



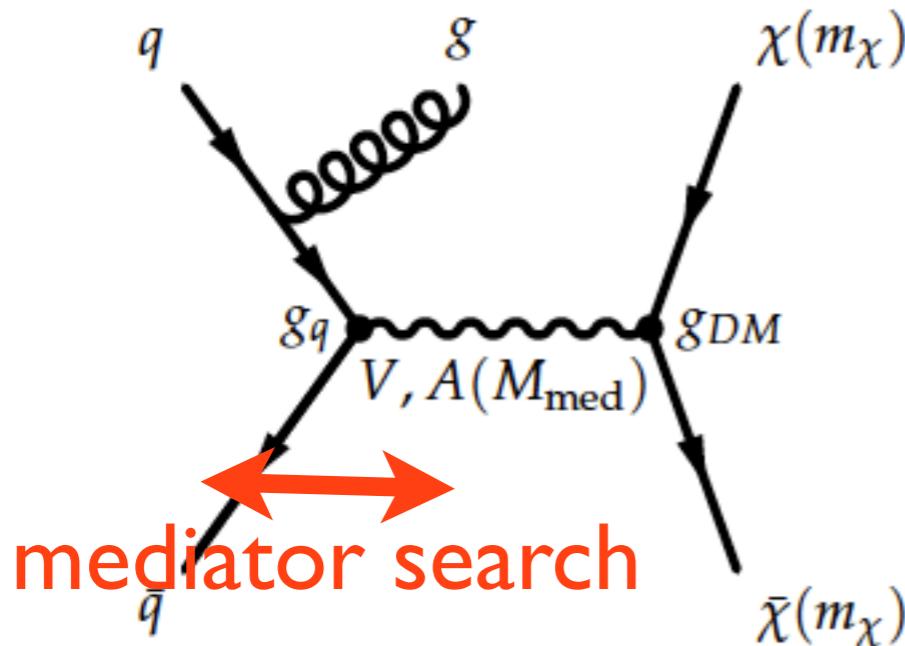
relic density  
indirect detection



# Signatures of simplified DM models

LHC DMWG [1507.00966, 1603.04156]

missing-energy search



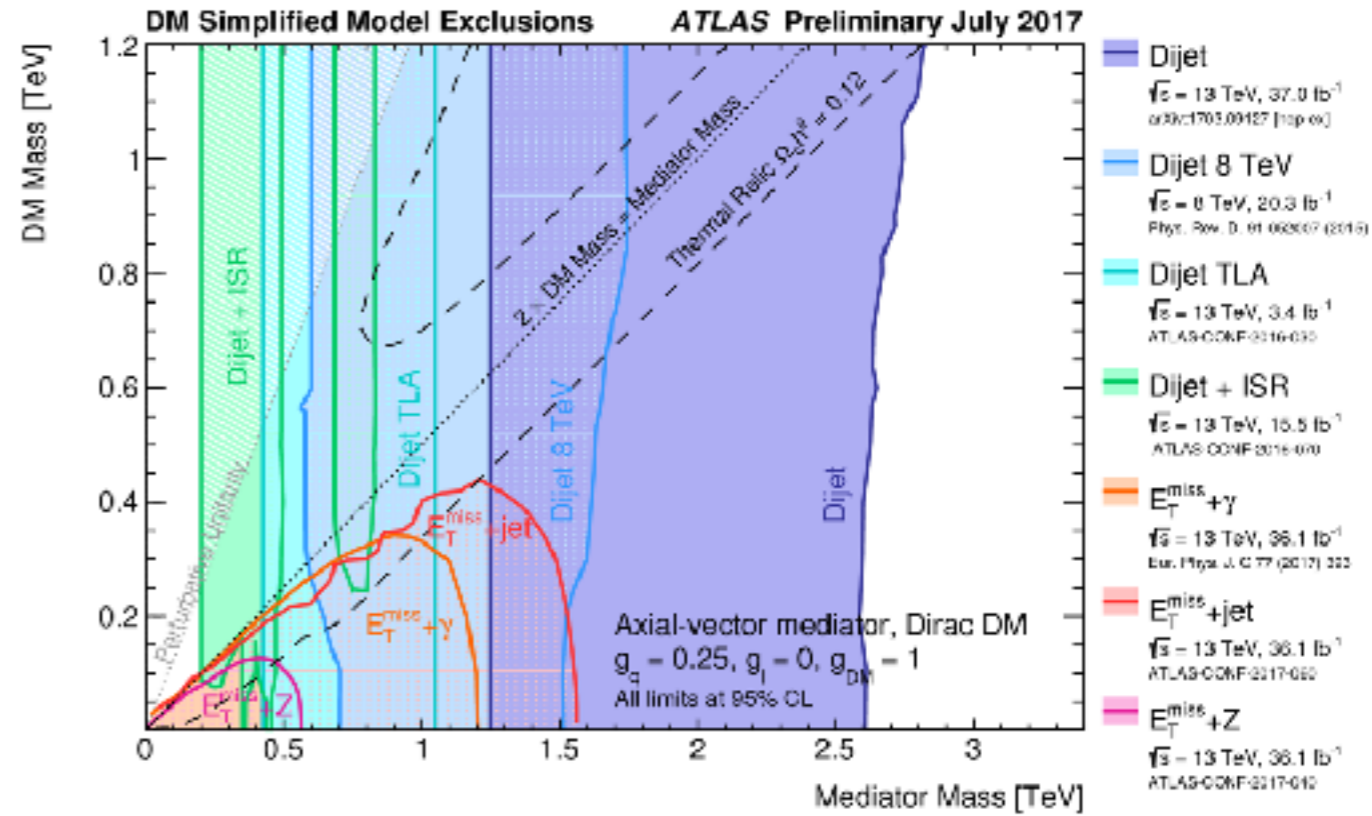
mediator search



relic density  
indirect detection



direct detection





# A comprehensive approach to DM studies: simplified top-philic models

[1605.09242, JHEP]

52 pages, 23 figs, 8 tables

Cosmology	relic		$m_X > m_t$		
	indirect		$m_X < m_t$	Planck, FermiLAT	
Astrophysics			$m_X > m_Y$		Arina
	direct		$m_X > 1 \text{ GeV}$	LUX, CDMSLite	Heisig
Colliders	$\cancel{E}_T$		$m_Y > 2m_X$	$+t\bar{t}$	Conte, Fuks, Guo
			$m_Y > 2m_X$	$+j, +Z, +h$	Martini, Vryonidou
	no $\cancel{E}_T$		$m_Y > 2m_t$	$4t$	Hespel
			$m_Y > 2m_t$	$t\bar{t}$	Pellen
			$m_Y < 2m_X, 2m_t$	$jj, \gamma\gamma$	

Kraemer

Maltoni

Arina

Backovic

Heisig

Conte, Fuks, Guo

Martini, Vryonidou

Mawatari

Hespel

Pellen

$$\mathcal{L}_{t,X}^{Y_0} = -\left(g_t \frac{yt}{\sqrt{2}} \bar{t}t + g_X \bar{X}X\right)Y_0$$

**X=DM**

**Y=mediator**

# BSM phenomenology workflow

- take a BSM model (symmetry, particle contents,...), i.e. Lagrangian

- derive the Feynman rules **Model provider**

- draw Feynman diagrams for our interesting processes

- compute the amplitude (squared)

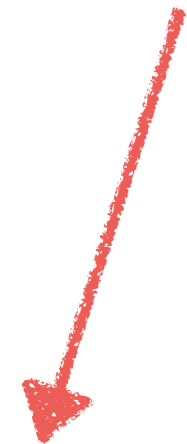
**Matrix-element  
Generator**

- generate parton-level events

- parton-shower/hadronisation **Shower Monte-Carlo programs**

- detector effect **Detector simulation tool**

- analysis **Recasting tool**



**DM physics tool**

DM annihilation

DM-N cross section

# BSM phenomenology workflow

- take a BSM model (symmetry, particle contents,...), i.e. Lagrangian

- derive the Feynman rules `FeynRules`

- draw Feynman diagrams for our interesting processes

- compute the amplitude (squared)

- generate parton-level events

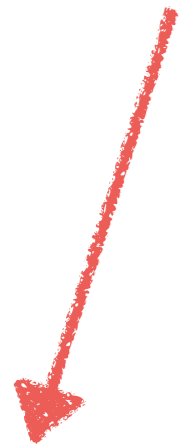
`MadGraph`

- parton-shower/hadronisation `Herwig, Pythia, Sherpa`

- detector simulation `Fastjet, Delphes`

- analysis `GAMBIT, Checkmate, FastLim, MadAnalysis`

`micrOMEGAs`  
`MadDM`



# MC tools for Run-2 DM searches

LHC DM WG [1507.00966]

Benchmark models for ATLAS and CMS Run-2 DM searches		
vector/axial vector mediator, <i>s</i> -channel (Sec. 2.1)		
Signature	State of the art calculation and tools	Implementation
jet + $\cancel{E}_T$	NLO+PS (powheg, SVN r3059)	[Forl; Foro]
	NLO+PS ( <u>DMsimp UFO</u> + MADGRAPH5_AMC@NLO v2.3.0)	[New]
	NLO (MCFM v7.0)	Upon request
W/Z/ $\gamma$ + $\cancel{E}_T$	LO+PS (UFO + MadGraph5_aMC@NLO v2.2.3)	[Fora]
	NLO+PS ( <u>DMsimp UFO</u> + MADGRAPH5_AMC@NLO v2.3.0)	[New]
scalar/pseudoscalar mediator, <i>s</i> -channel (Sec. 2.2)		
Signature	State of the art calculation and tools	Implementation
jet + $\cancel{E}_T$	LO+PS, top loop (powheg, r3059)	[Forn; Form]
	LO+PS, top loop ( <u>DMsimp UFO</u> + MADGRAPH5_AMC@NLO v2.3.0)	[New]
	LO, top loop (MCFM v7.0)	Upon request
W/Z/ $\gamma$ + $\cancel{E}_T$	LO+PS (UFO + MadGraph5_aMC@NLO v2.2.3)	
$t\bar{t}, b\bar{b}$ + $\cancel{E}_T$	LO+PS (UFO + MadGraph5_aMC@NLO v2.2.3)	[Ford]
	NLO+PS ( <u>DMsimp UFO</u> + MADGRAPH5_AMC@NLO v2.3.0)	[New]

Backovic, Kramer, Maltoni, Martini, KM, Pellen [1508.05327, EPJC]

## Simplified dark matter models

### Authors

- s-channel (spin-0 and spin-1)
  - Antony Martini (Université catholique de Louvain) & Kentarou Mawatari (LPSC Grenoble)
    - Emails: kentaro.mawatari@lpsc.in2p3.fr
- s-channel (spin-0 and spin-1 electroweak)
  - Jian Wang (Johnnas Gutenberg University of Mainz) & Cen Zhang (Brookhaven National Laboratory)
    - Emails: cenzhang@bnl.gov
- s-channel (spin-2)
  - Goutam Das (Saha Inst.), Celine Degrande (CERN) & Kentarou Mawatari (LPSC Grenoble)
    - Emails: goutam.cas@saha.ac.in, celine.degrande@cern.ch, kentaro.mawatari@lpsc.in2p3.fr

### Description of the model

This is simplified dark matter models for NLO. Our lagrangian consists of different types of DM:

- $X_r$  (real scalar DM)
- $X_c$  (complex scalar DM)
- $X_d$  (Dirac spinor DM)
- $X_m$  (Majorana spinor DM) [to be done.]
- $X_v$  (vector DM)
- ...

and different types of mediators:

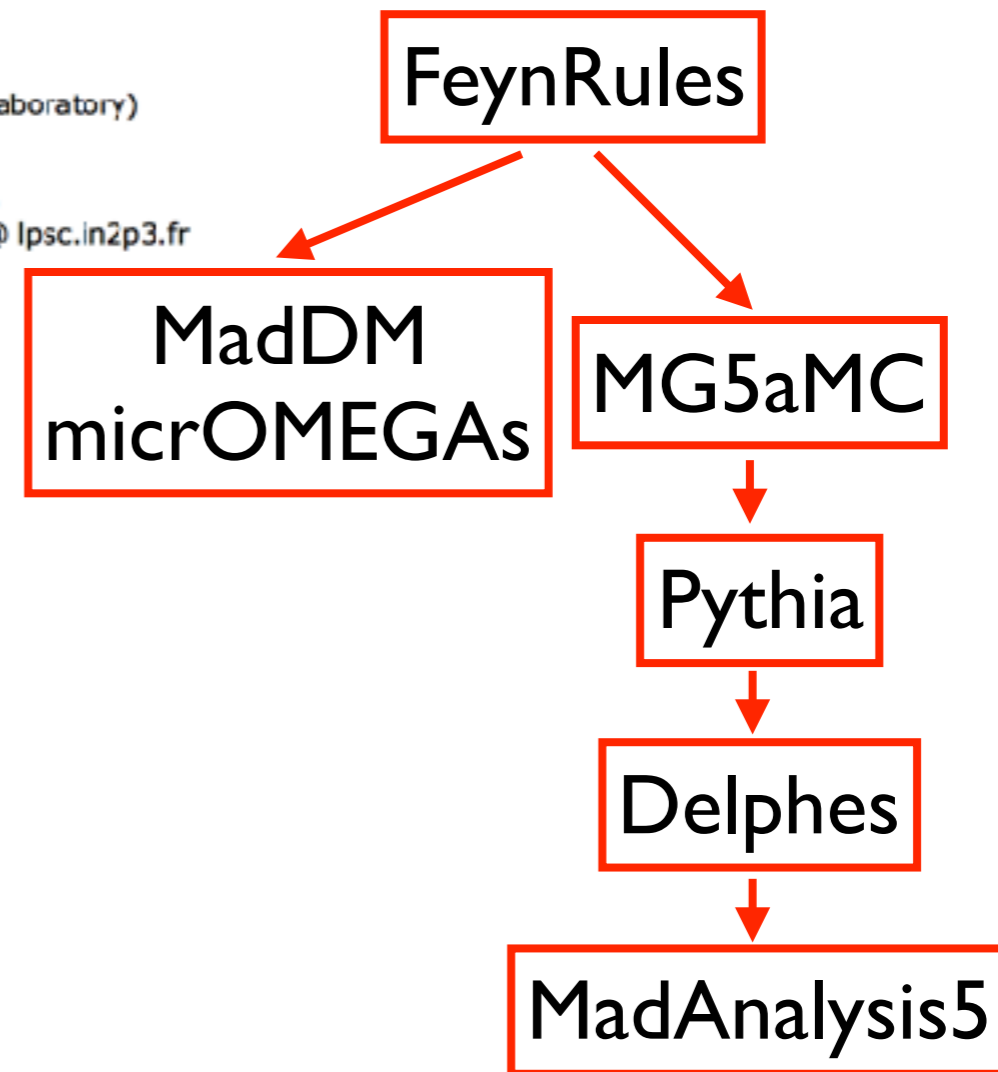
- s-channel
  - $Y_0$  (spin-0)
  - $Y_1$  (spin-1)
  - $Y_2$  (spin-2)
  - ...
- t-channel [to be done.]

See more details in

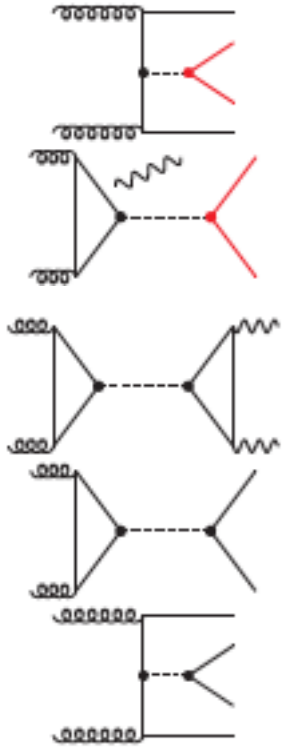
- [1508.00564](#) : O. Mattelaer, E. Vryonidou, "Dark matter production through loop-induced processes at the LHC: the s-channel mediator case" (EPJC75(2015)436).
- [1508.05327](#) : M. Backovic, M. Kramer, F. Maltoni, A. Martini, K. Mawatari, M. Pellen, "Higher-order QCD predictions for dark matter production at the LHC in simplified models with s-channel mediators" (EPJC75(2015)482).
- [1509.05785](#) : M. Neubert, J. Wang, C. Zhang, "Higher-order QCD predictions for dark matter production in mono-Z searches at the LHC" (JHEP1602(2016)082).
- [1605.09359](#) : G. Das, C. Degrande, V. Hirschi, F. Maltoni, H. Shao, "NLO predictions for the production of a spin-two particle at the LHC"
- [1701.07008](#) : S. Kraml, U. Laa, K. Mawatari, K. Yamashita, "Simplified dark matter models with a spin-2 mediator at the LHC".

$$\mathcal{L}_{t,X}^{Y_0} = -\left(g_t \frac{y_t}{\sqrt{2}} \bar{t}t + g_X \bar{X}X\right)Y_0$$

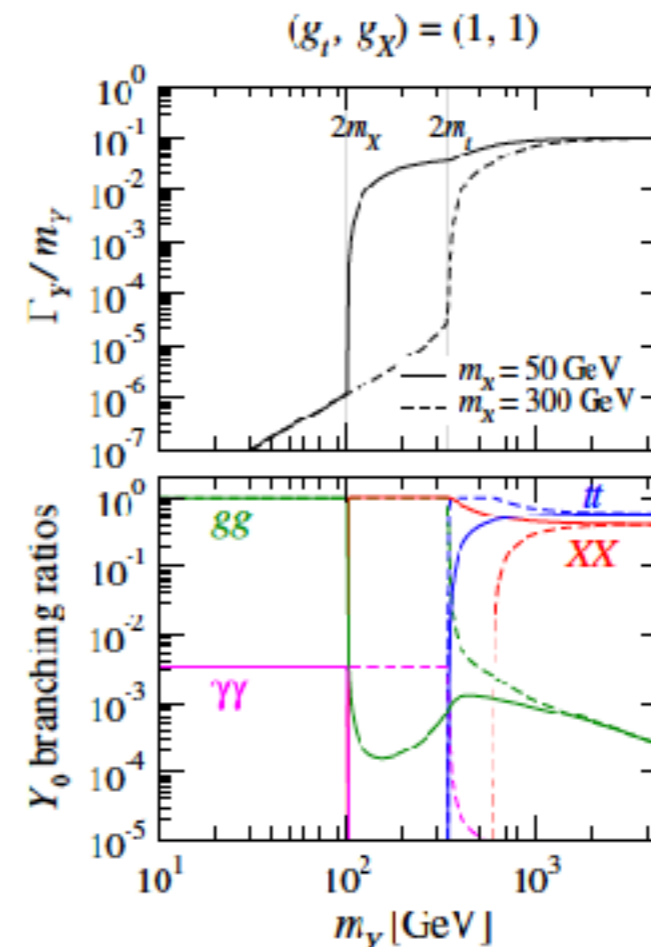
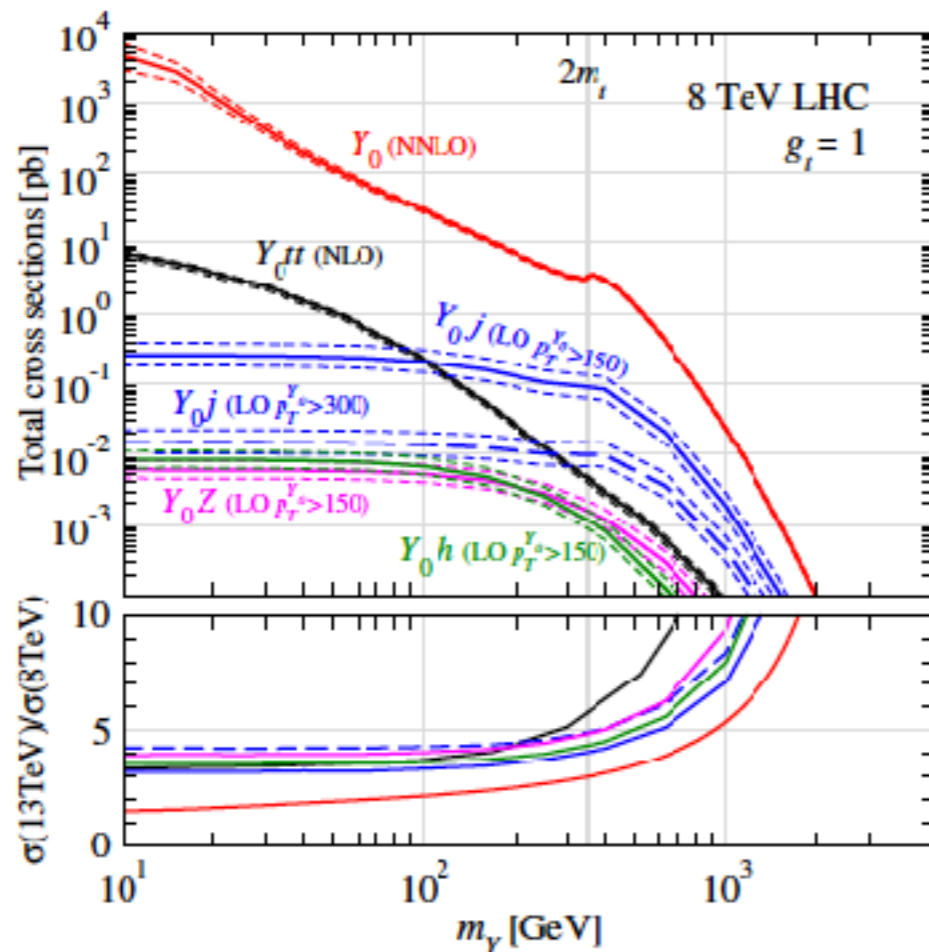
$\{g_t, g_X, m_X, m_Y\}$



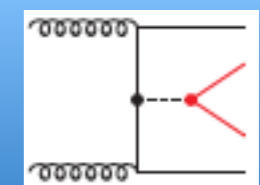
# LHC(8TeV) constraints for top-philic DM



Final state	Imposed constraint	Reference	Comments
$\cancel{E}_T + t\bar{t}$	MADANALYSIS5 PAD (new)	CMS [1504.03198]	Semileptonic top-antitop decay
$\cancel{E}_T + j$	MADANALYSIS5 PAD (new)	CMS [1408.3583]	
$\cancel{E}_T + Z$	$\sigma(\cancel{E}_T > 150 \text{ GeV}) < 0.85 \text{ fb}$	CMS [1511.09375]	Leptonic Z-boson decay
$\cancel{E}_T + h$	$\sigma(\cancel{E}_T > 150 \text{ GeV}) < 3.6 \text{ fb}$	ATLAS [1510.06218]	$h \rightarrow b\bar{b}$ decay
$jj$	$\sigma(m_Y = 500 \text{ GeV}) < 10 \text{ pb}$	CMS [1604.08907]	Only when $m_Y > 500 \text{ GeV}$
$\gamma\gamma$	$\sigma(m_Y = 150 \text{ GeV}) < 30 \text{ fb}$	CMS [1506.02301]	Only when $m_Y > 150 \text{ GeV}$
$t\bar{t}$	$\sigma(m_Y = 400 \text{ GeV}) < 3 \text{ pb}$	ATLAS [1505.07018]	Only when $m_Y > 400 \text{ GeV}$
$t\bar{t}t\bar{t}$	$\sigma < 32 \text{ fb}$	CMS [1409.7339]	Upper limit on the SM cross section



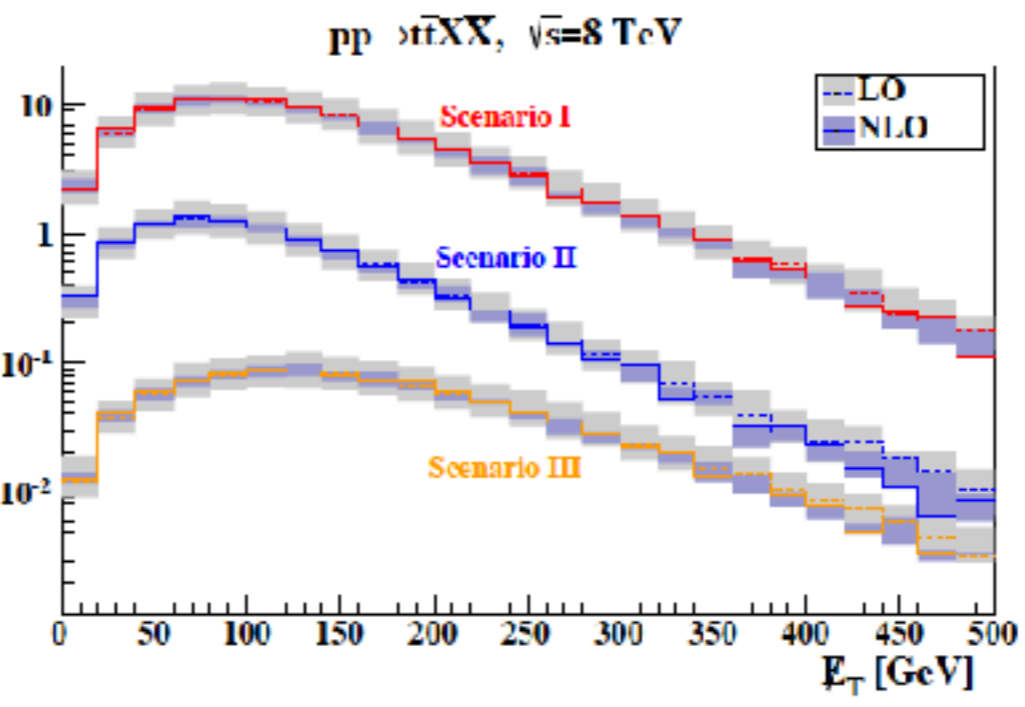
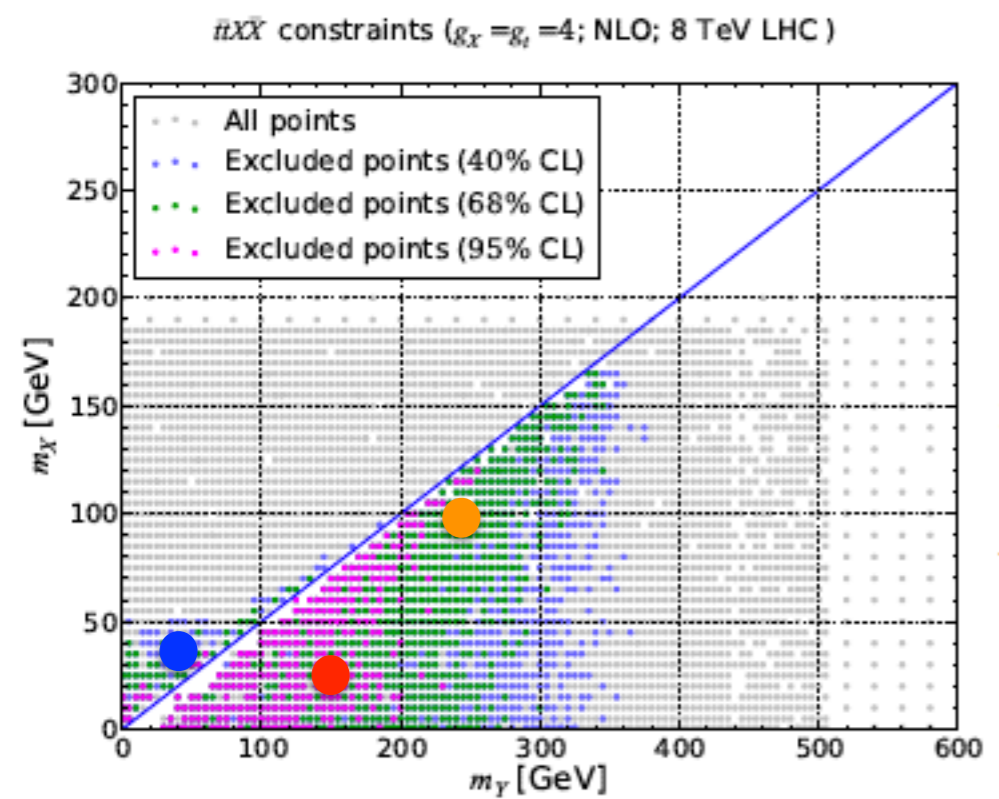
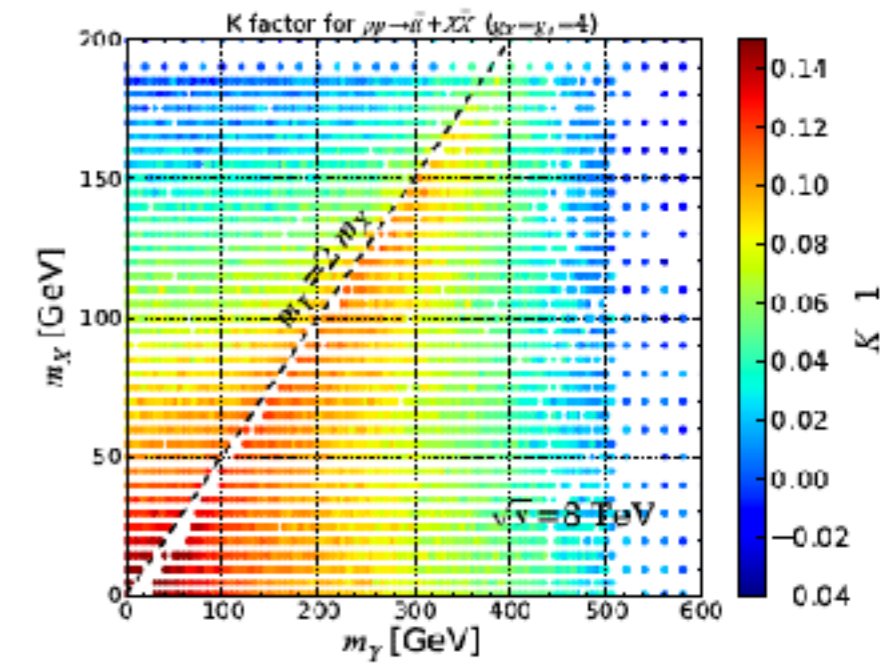
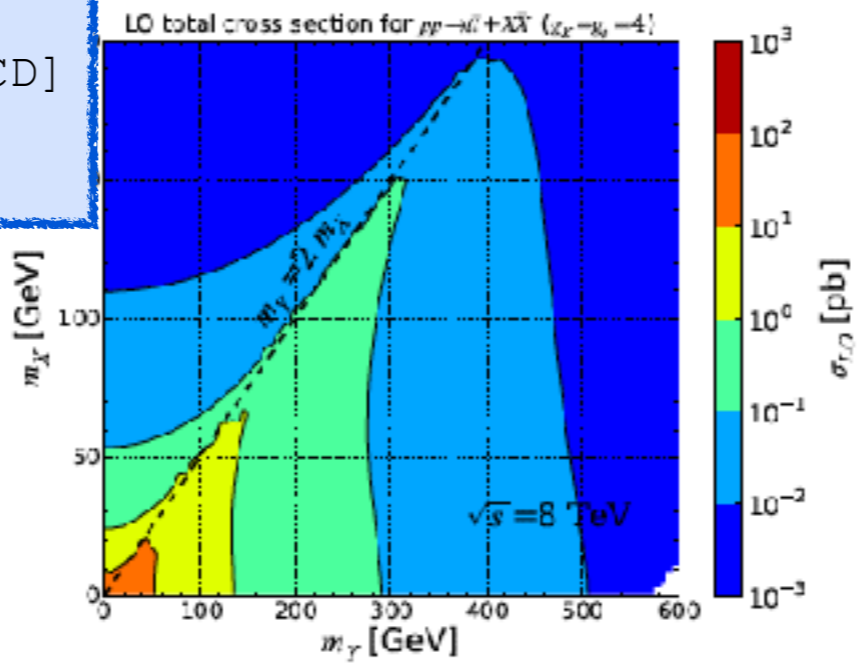
# MET + a top-quark pair (NLO)



```

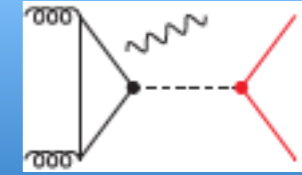
./bin/mg5_aMC
>import model DMsimp_s_spin0
>generate p p > t t~ xd xd~ [QCD]
>output
>launch
    
```

$(m_Y, m_X)$	$\sigma_{LO}$ [pb]	$\sigma_{NLO}$ [pb]
(150, 25) GeV	$0.658^{+34.9\%}_{-24.0\%}$	$0.773^{+6.1\%}_{-10.1\%}$
(40, 30) GeV	$0.776^{+34.2\%}_{-24.1\%}$	$0.926^{+5.7\%}_{-10.4\%}$
(240, 100) GeV	$0.187^{+37.1\%}_{-24.4\%}$	$0.216^{+6.7\%}_{-11.4\%}$



**NLO predictions reduce the theoretical uncertainty.**

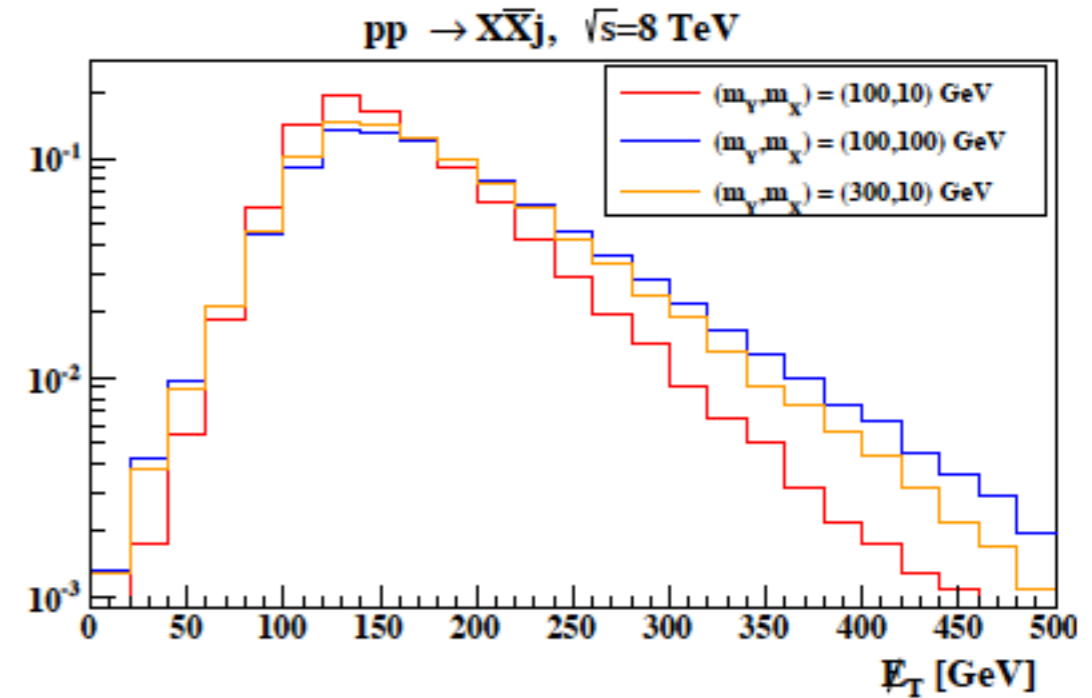
# Mono-jet (LO loop-induced)



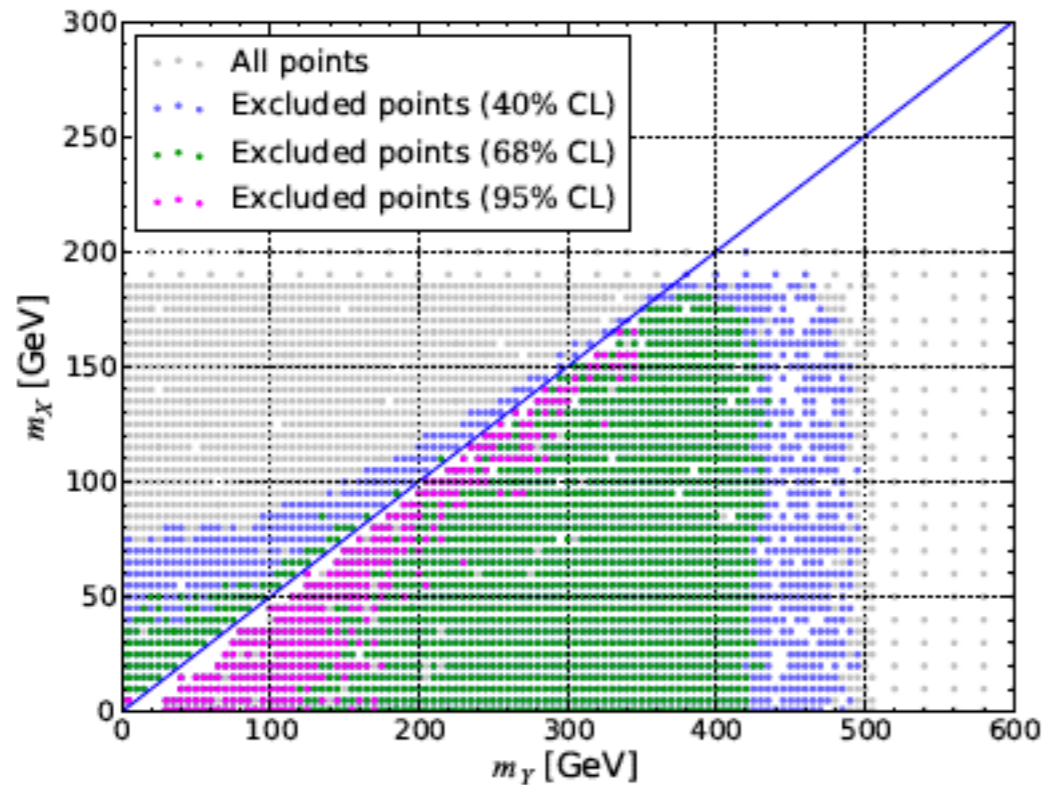
```

./bin/mg5_aMC
>import model DMsimp_s_spin0
>generate p p > j xd xd~ [QCD]
>output
>launch
    
```

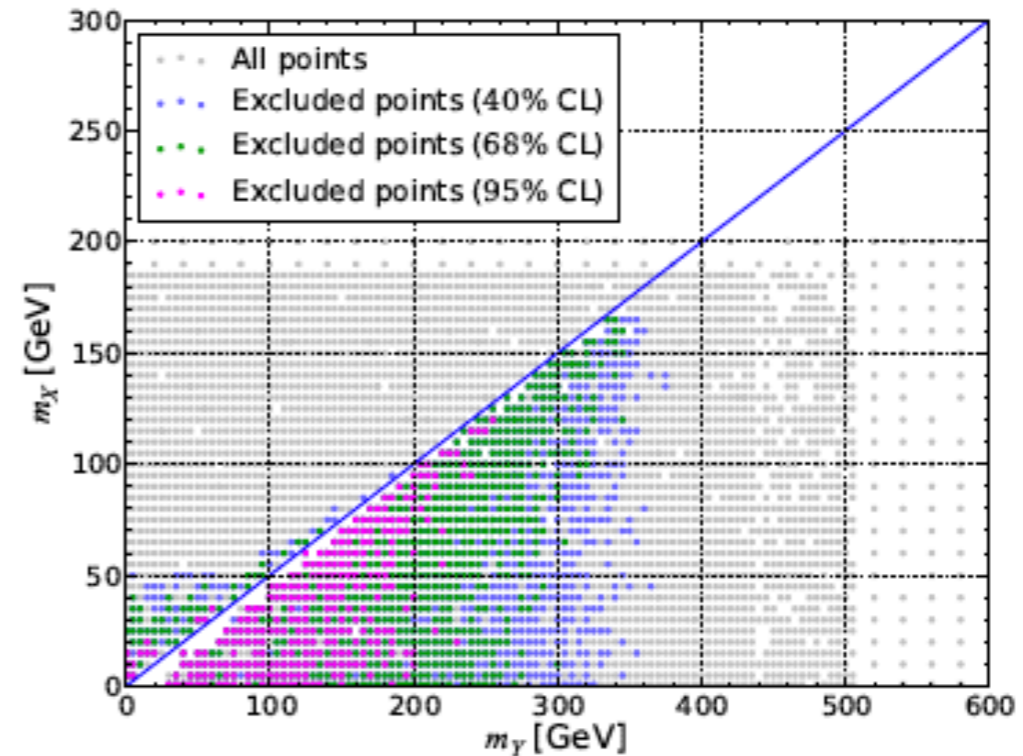
**Complementarity between the different searches.**



Monojet constraints ( $g_X=g_t=4$ ; 8 TeV LHC)



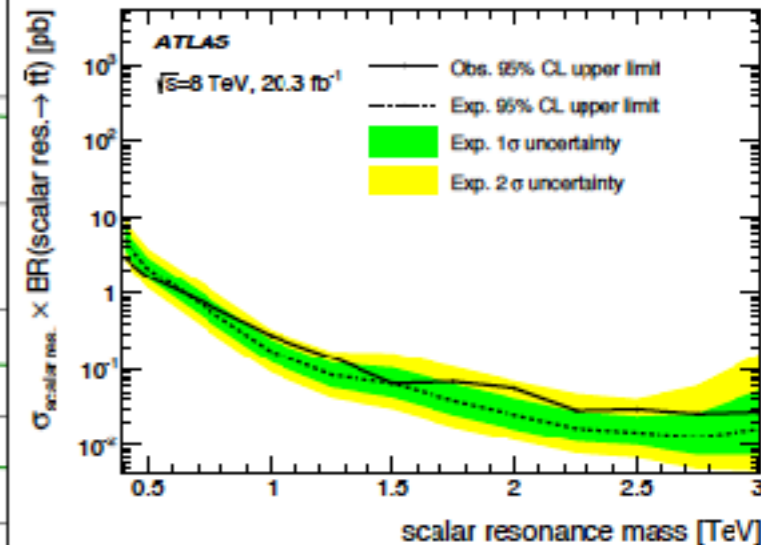
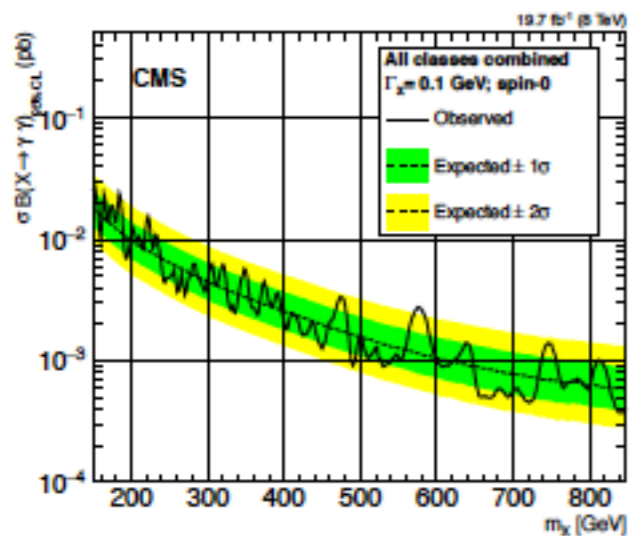
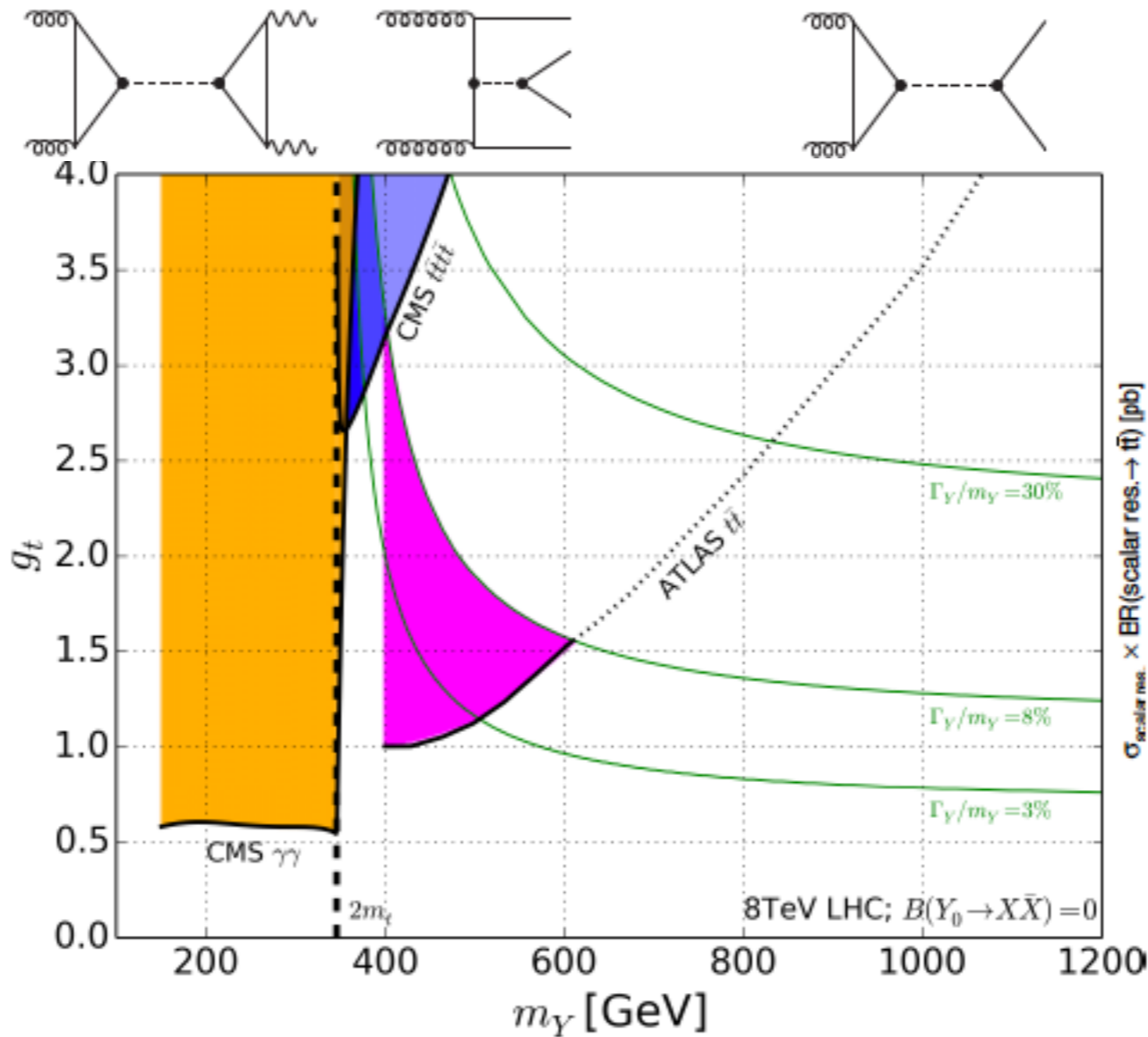
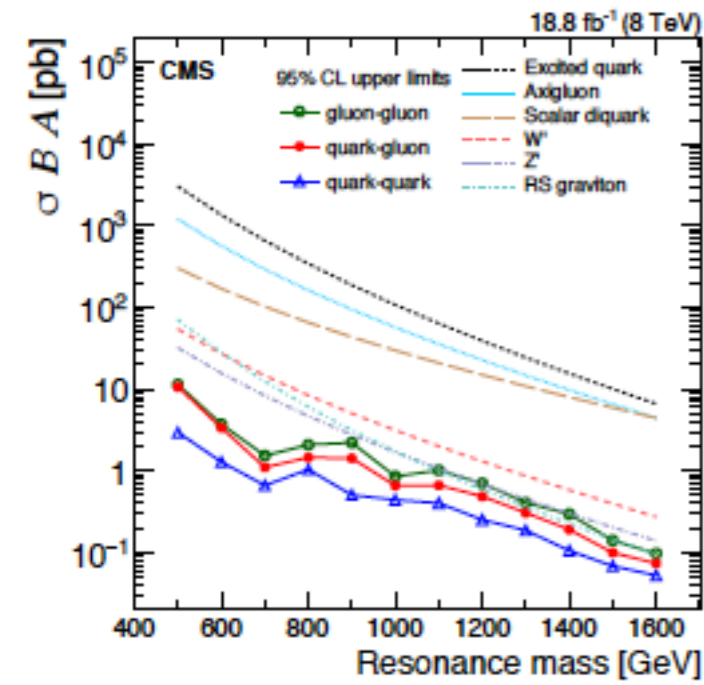
$\bar{t}tXX$  constraints ( $g_X=g_t=4$ ; NLO; 8 TeV LHC)





# Resonance search constraints

Extension of the search range is desirable.



$jj$	$\sigma(m_Y = 500 \text{ GeV}) < 10 \text{ pb}$	CMS [1604.08907]	Only when $m_Y > 500 \text{ GeV}$
$\gamma\gamma$	$\sigma(m_Y = 150 \text{ GeV}) < 30 \text{ fb}$	CMS [1506.02301]	Only when $m_Y > 150 \text{ GeV}$
$t\bar{t}$	$\sigma(m_Y = 400 \text{ GeV}) < 3 \text{ pb}$	ATLAS [1505.07018]	Only when $m_Y > 400 \text{ GeV}$
$t\bar{t}\bar{t}$	$\sigma < 32 \text{ fb}$	CMS [1409.7339]	Upper limit on the SM cross section

# Relic vs. Direct detection vs. LHC

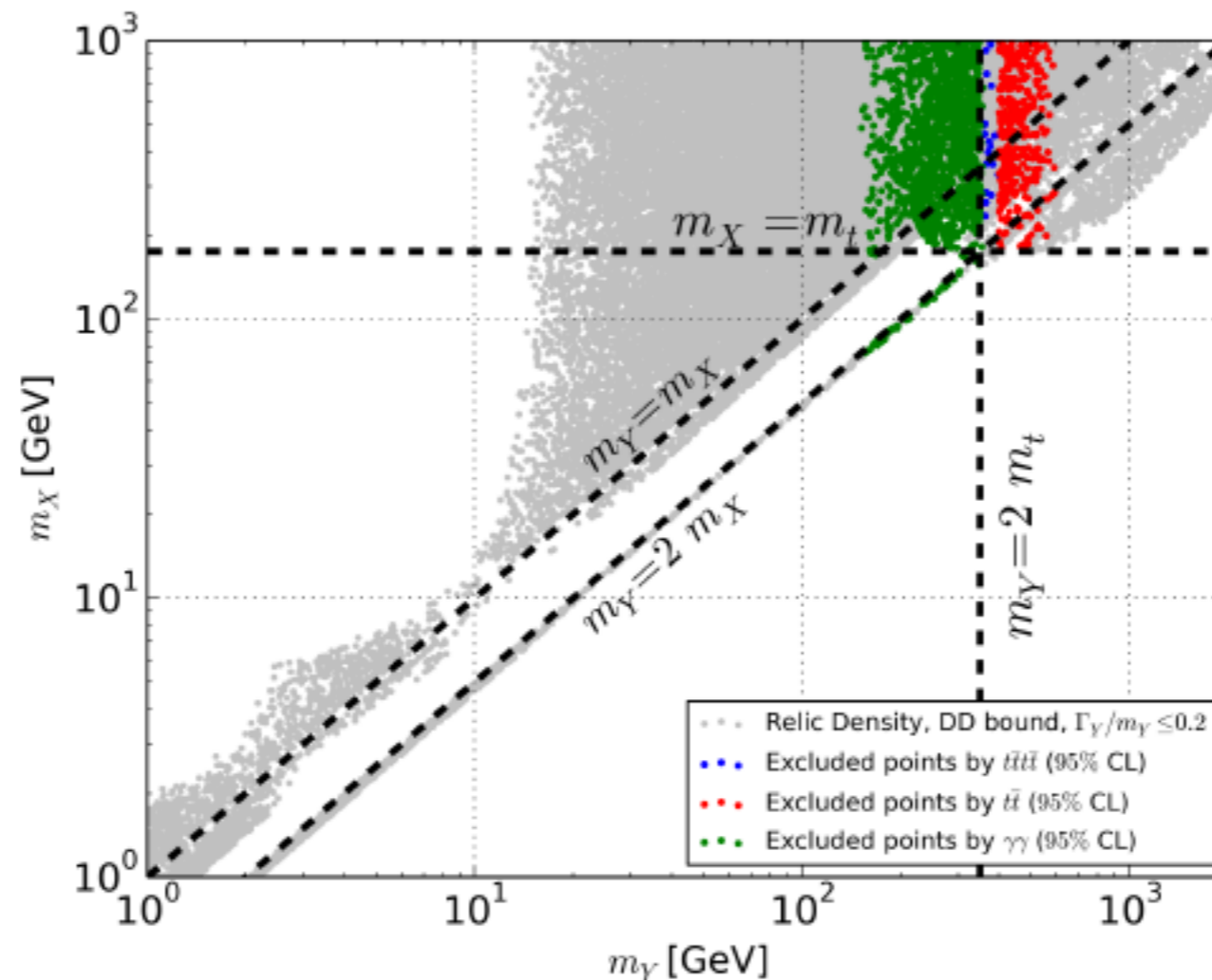
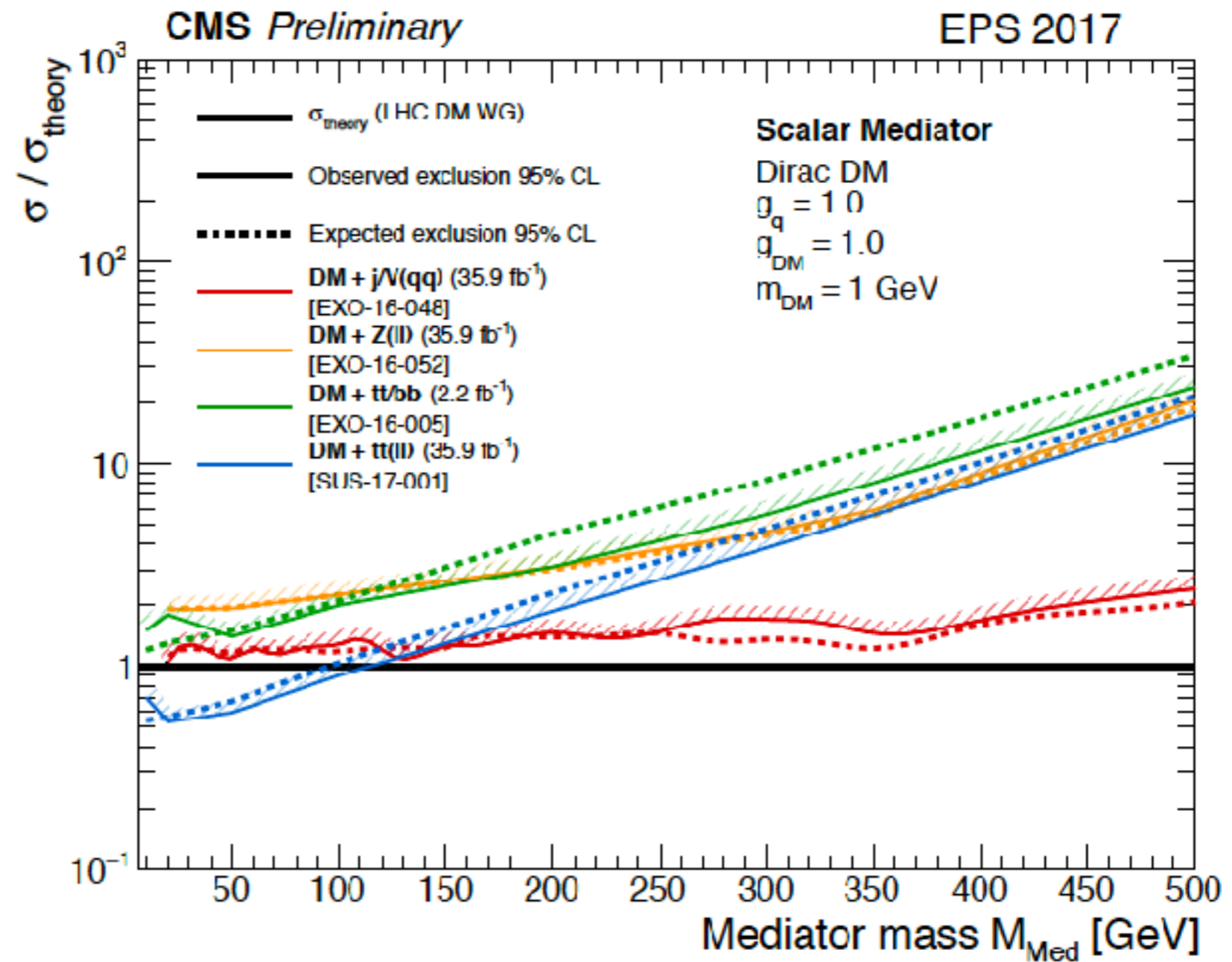
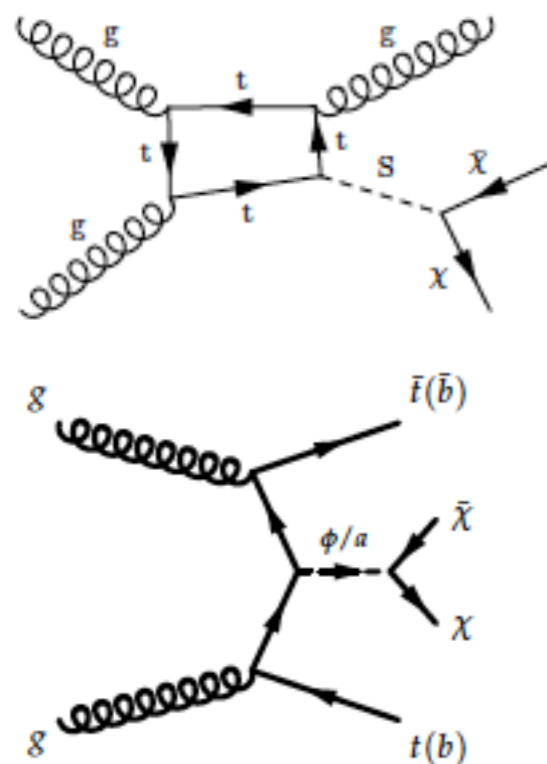


Figure 14. Results of our four-dimensional parameter scan projected onto the  $(m_Y, m_X)$  plane once constraints set from the LHC results are imposed. The points excluded by the diphoton, the  $t\bar{t}$  and the four-top considered searches all satisfy the relic density, narrow width and direct detection constraints.

# 13TeV MET constraints on top-philic DM



# Towards the next generation of simplified Dark Matter models

---

- 2 **Scalar singlet model with mixing**
  - 2.1 Scenario A
  - 2.2 Scenario B
  - 2.3 Scenarios C and D
  - 2.4 VBM production
  - 2.5 Relic density
  
- 3 **Connecting an LHC discovery of a mediator particle with DM signals**
  - 3.1 Simplified DM model
  - 3.2 Monojet signatures
  - 3.3 LHC constraints
  - 3.4 Direct detection
  - 3.5 Indirect detection
  - 3.6 Benchmark scenarios
  
- 4 **Other simplified models of interest**
  - 4.1 A few representative  $t$ -channel simplified models
  - 4.2 Spin-2 models
  - 4.3 Pseudo-Dirac DM
  
- 5 **What can we learn about simplified DM models from SUSY?**
  - 5.1 The DM mechanisms in SUSY
  - 5.2 Collider signatures
  - 5.3 Interplay of Collider and Direct Detection Searches in SUSY
  - 5.4 Lessons from SUSY for simplified DM models

arXiv:1607.06680

# Spin-2 mediated simplified DM models

Gravity-mediated DM:

Lee, Park, Sanz [1306.4107, 1401.5301]

Rueter, Rizzo, Hewett [1706.07540]

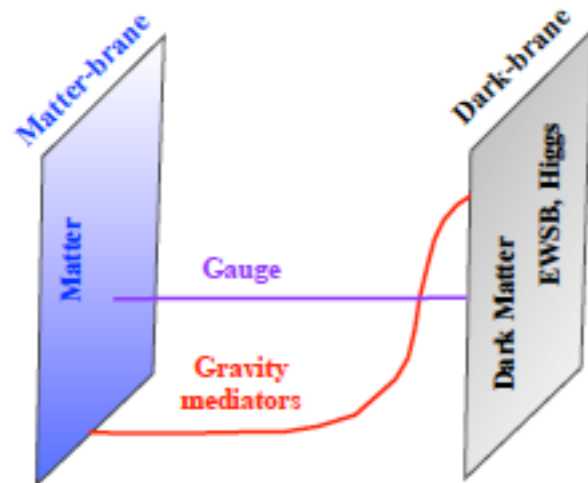
Kraml, Laa, KM, Yamashita [1701.07008, EPJC]

$$\mathcal{L}_X^{Y_2} = -\frac{1}{\Lambda} g_X^T T_{\mu\nu}^X Y_2^{\mu\nu}$$

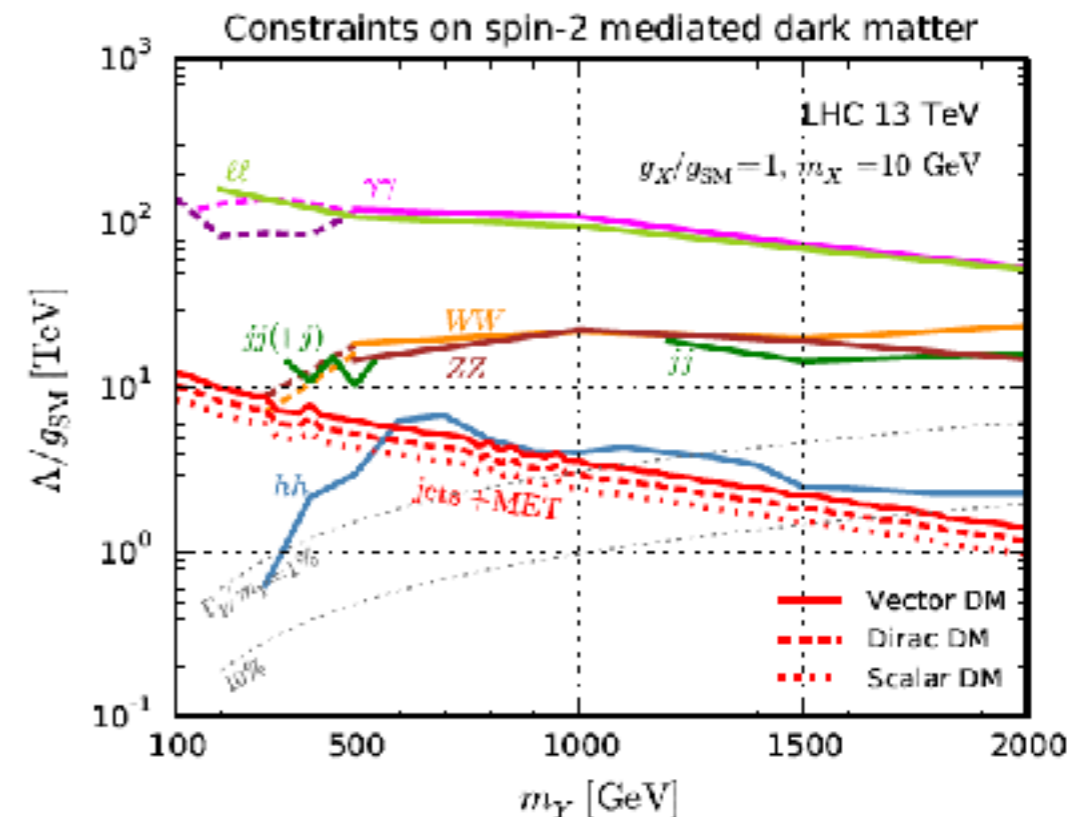
$$\mathcal{L}_{\text{SM}}^{Y_2} = -\frac{1}{\Lambda} \sum_i g_i^T T_{\mu\nu}^i Y_2^{\mu\nu}$$

$$g_{\text{SM}} \equiv g_H^T = g_q^T = g_\ell^T = g_g^T = g_W^T = g_B^T$$

$$\{m_X, m_Y, g_X/\Lambda, g_{\text{SM}}/\Lambda\}$$



- Which analyses can constrain the model at the LHC?
- Which analyses can provide the best constraint?
- Can we find some interesting parameter space?



# Summary and outlook

- 暗黒物質探索(collider + non-collider)のための系統的シミュレーションフレームワークが発展している。
  - LHC Dark Matter Working Group
  - EFT  $\leftrightarrow$  Simplified models  $\leftrightarrow$  UV-completed models (e.g. SUSY)

- 高次補正の重要性

- 信頼できる生成断面積、理論の不定性の縮小

- 一つのDMモデルを多角的に攻める。

- Recasting/reinterpretation の重要性

