

# Apparent Horizons in Holographic Dual of Hydrodynamics

Shin Nakamura

(Center for Quantum Spacetime  
(CQUeST) , Sogang Univ.)

Ref.: S. Kinoshita, S. Mukohyama, S.N. and K. Oda.  
arXiv:0807.3797

# Motivation

**RHIC**: Relativistic Heavy Ion Collider  
(@ Brookhaven National Laboratory)

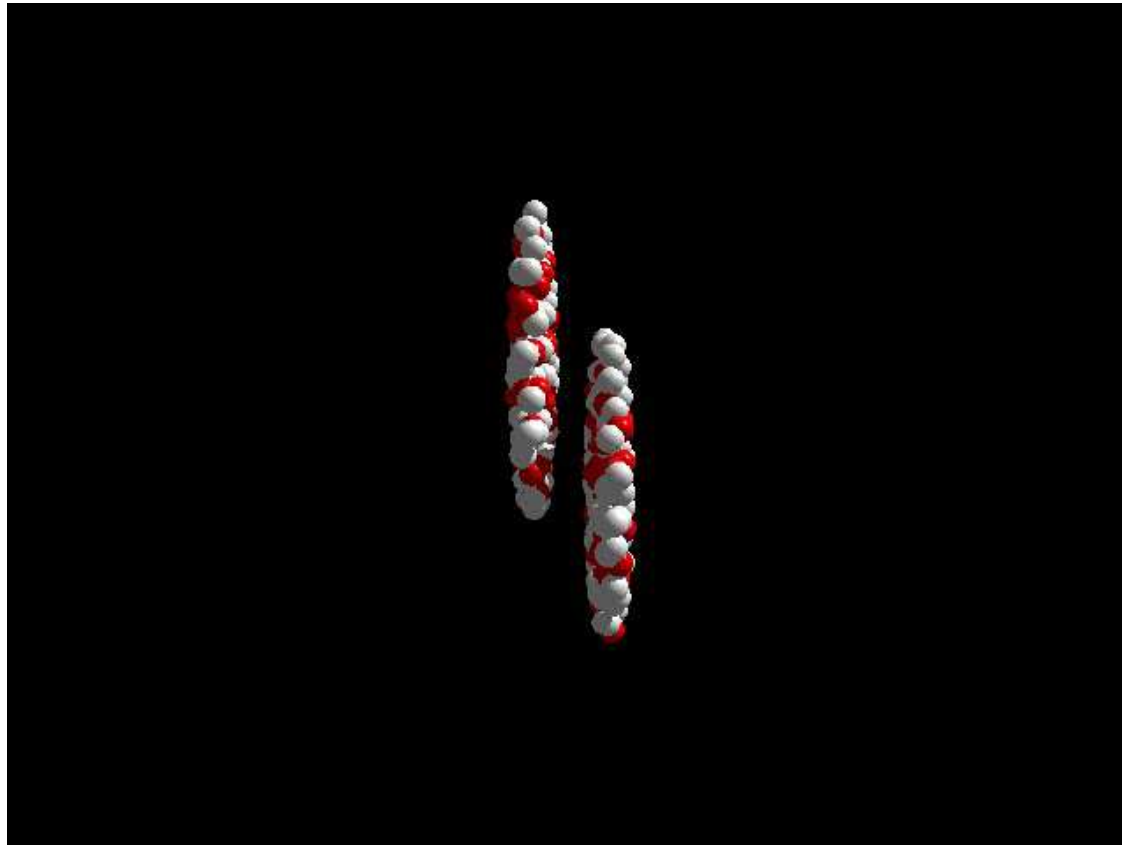
Heavy ion:  
e.g.  $^{197}\text{Au}$   
 $\sqrt{s_{NN}} \sim 200\text{GeV}$ .

Similar exp. at  
**LHC**: ALICE/CMS



[http://www.bnl.gov/RHIC/inside\\_1.htm](http://www.bnl.gov/RHIC/inside_1.htm)

# Quark-gluon plasma (QGP) is created

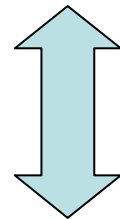


For the movie: [http://www.bnl.gov/RHIC/heavy\\_ion.htm](http://www.bnl.gov/RHIC/heavy_ion.htm)

# This is a **time-dependent** system

- Furthermore, QGP is known to be a **strongly interacting** system.

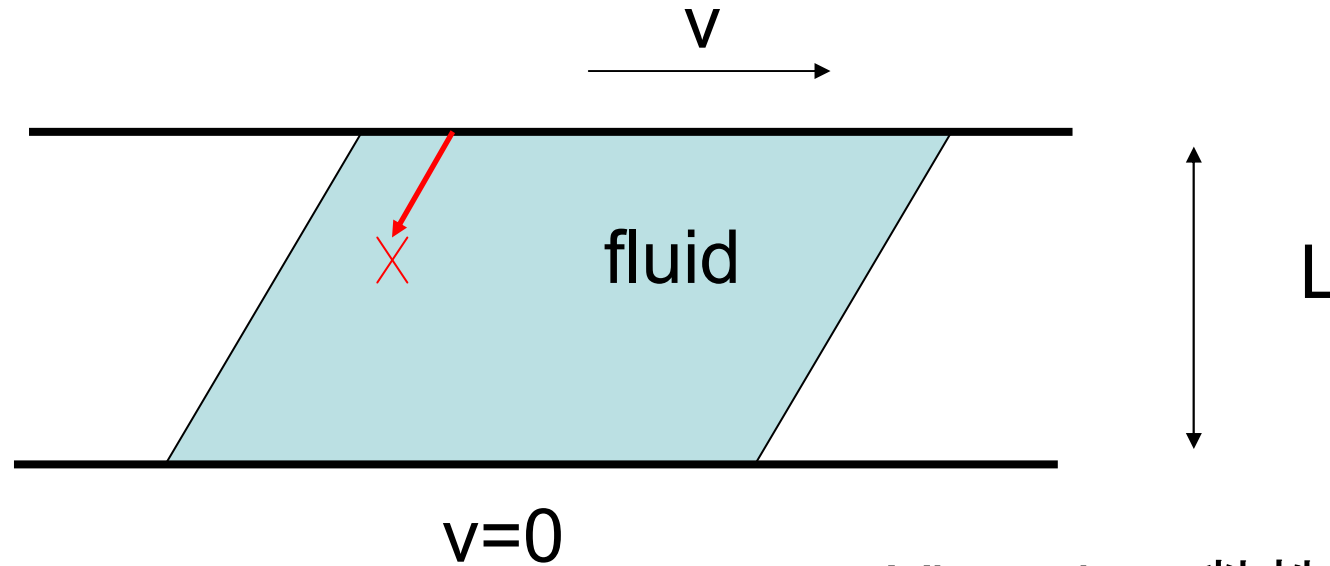
(small viscosity)



AdS/CFT

**A time-dependent 5d geometry.**

# Shear viscosity



$$\frac{F}{A} = \eta \frac{v}{L}$$

Viscosity: 粘性度

Shear: 剪断

Strong interaction  $\longleftrightarrow$  Small viscosity

# Construction of time-dependent AdS/CFT itself is a **challenge**.

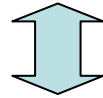
- In this talk, we consider the **Bjorken flow** of **N=4 SYM** theory.

A **standard, simplest** model of QGP

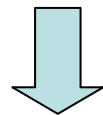


# How to obtain the geometry?

Solution to the 5d Einstein equation with  $\Lambda < 0$   
with appropriate boundary conditions  
and initial conditions.



- The 4d geometry where the YM theory lives.
- The initial stress tensor of the YM fluid.



Time-dependent solutions

# General relativity as hydrodynamics

time-dependent

5d Geometry

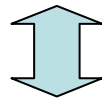


time-dependent

4d Stress tensor

GKP-Witten prescription  
(AdS/CFT dictionary)

Time evolution of (the long-range modes of)  
stress tensor



Described by hydrodynamics.

Hydrodynamics comes out of the gravity!



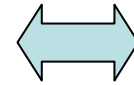
# Summary of our work

## Hydrodynamics

- hydrodynamic equation  
(energy-momentum conservation)
- equation of state  
(conformal invariance)

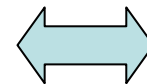
## Our model

5d Einstein's eq. at  
the vicinity of the  
boundary



- transport coefficients  
(viscosity, relaxation time...)

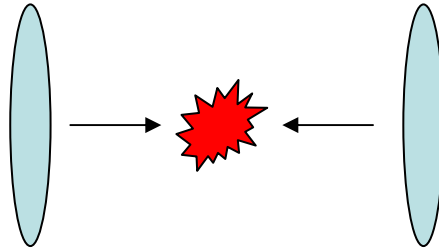
Regularity around  
the horizon



Related to thermal equilibrium

Physical parameters from Cosmic Censorship.

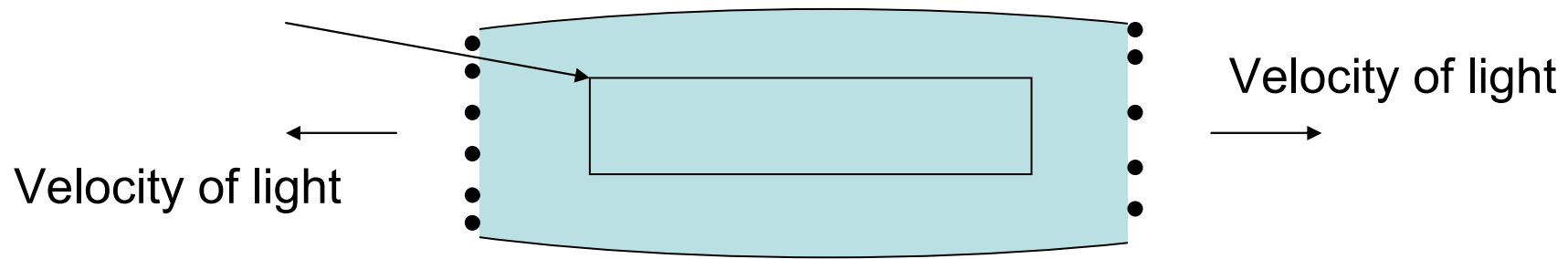
# Our system: Bjorken flow



Relativistically accelerated heavy nuclei

↑  
A standard model  
of QGP fluid

## Central Rapidity Region (CRR)



After collision

(Bjorken 1983)

- (Almost) **one-dimensional expansion**.
- We have **boost symmetry** in the CRR.

→ Time dependence of the physical quantities are written by the **proper time**.

# Local rest frame(LRF)

$$t = \tau \cosh y, x^1 = \tau \sinh y$$

$$ds^2 = -d\tau^2 + \tau^2 dy^2 + dx_{\perp}^2$$

Rapidity

$y = \text{const.}$

Proper-time

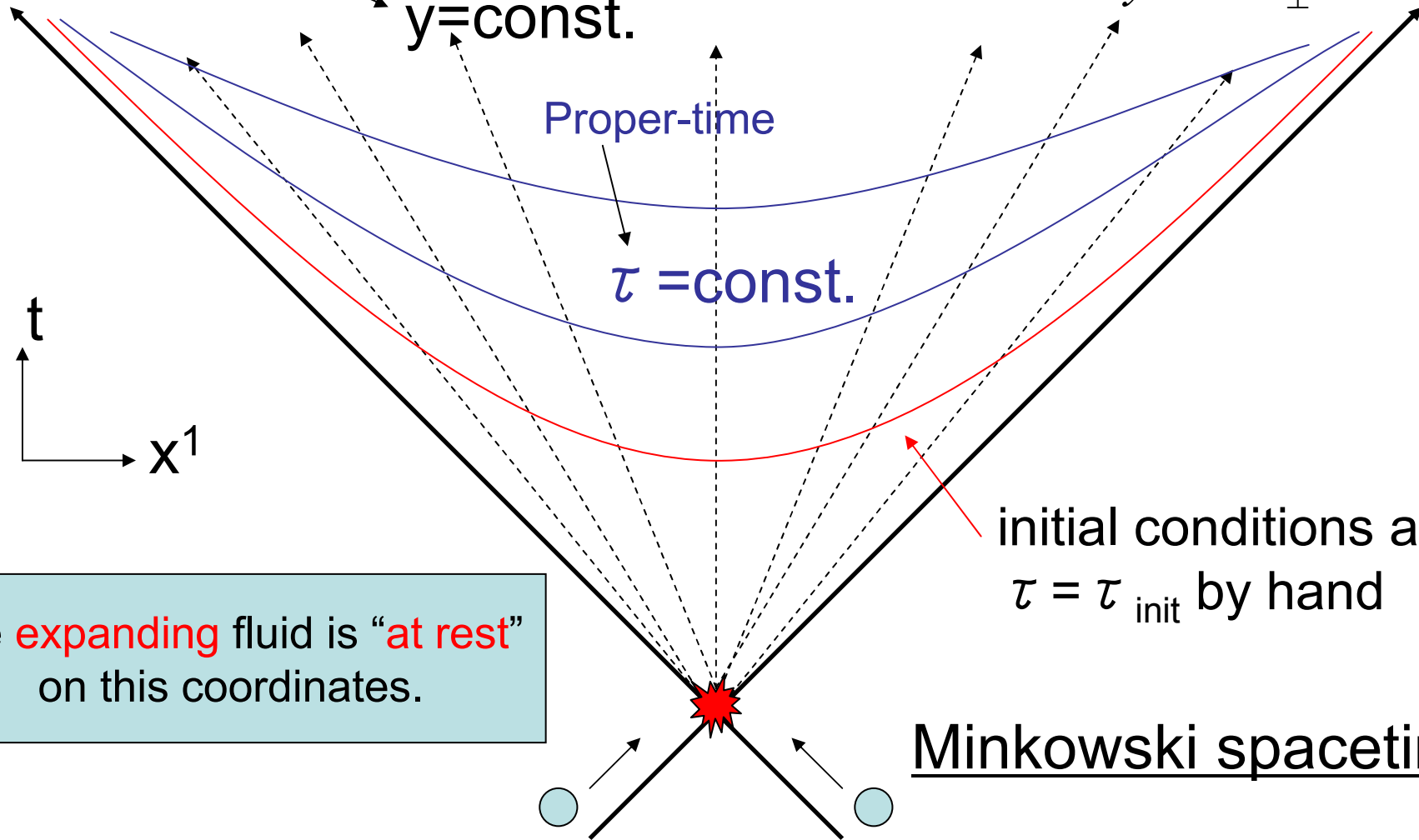
$\tau = \text{const.}$

$t$   
 $x^1$

initial conditions at  
 $\tau = \tau_{\text{init}}$  by hand

The **expanding** fluid is “**at rest**”  
on this coordinates.

Minkowski spacetime



# Stress tensor on LRF

The stress tensor becomes **diagonal**:

$$T_{\mu\nu} = \begin{pmatrix} \varepsilon & 0 & 0 & 0 \\ 0 & \tau^2 \left( p - \frac{4}{3} \frac{\eta}{\tau} \right) & 0 & 0 \\ 0 & 0 & p + \frac{2}{3} \frac{\eta}{\tau} & 0 \\ 0 & 0 & 0 & p + \frac{2}{3} \frac{\eta}{\tau} \end{pmatrix}$$

energy density

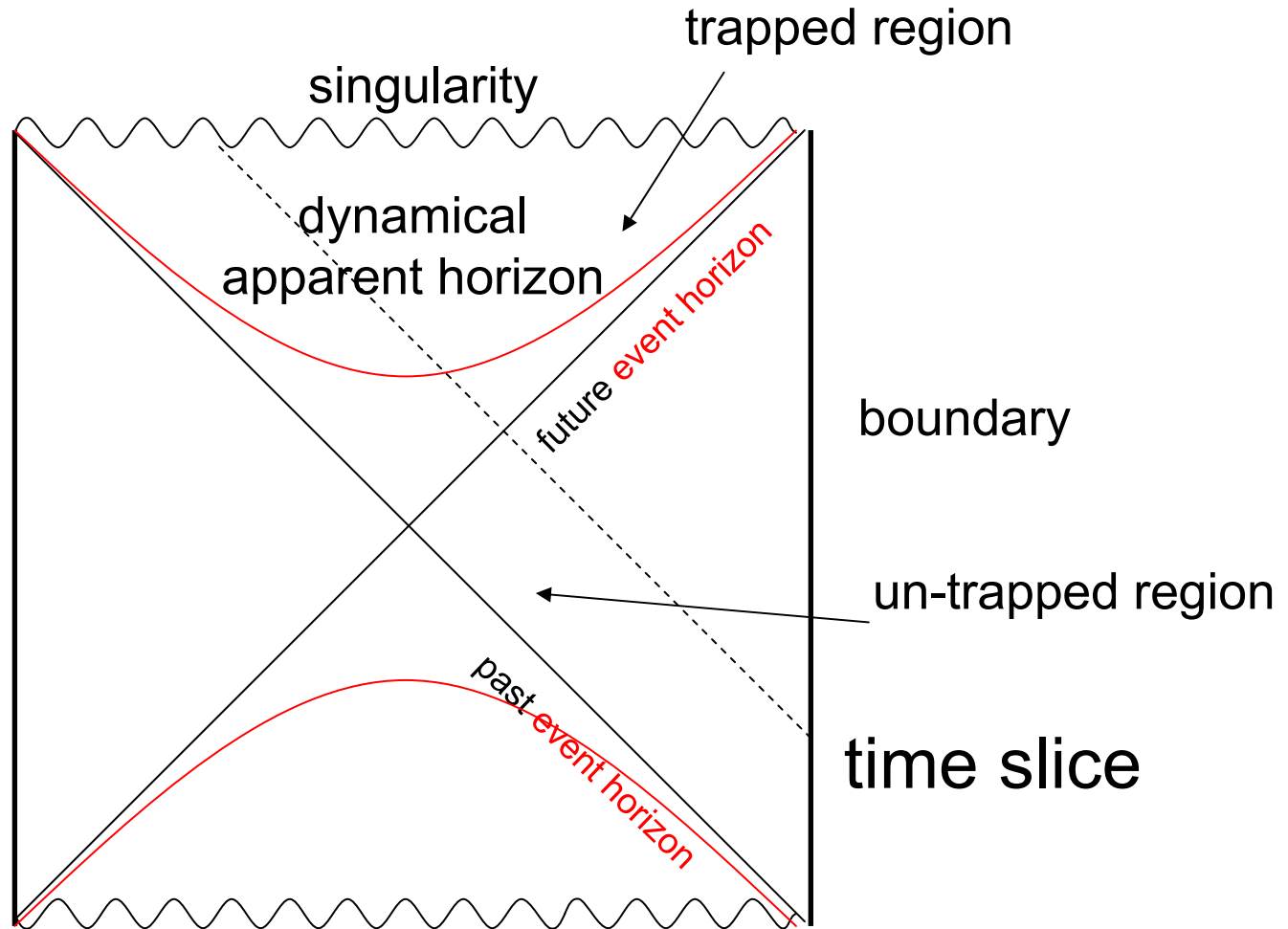
pressure

shear viscosity

If the stress tensor is **diagonal**

→ you are **on the flow**.

# Penrose diagram



Cf. Janik-Peschanski (2005)  
on the Fefferman-Graham coordinates.

**Eddington-Finkelstein**  
coordinates

# Our dual geometry

- Parametrization (Eddington-Finkelstein type):

$$ds^2 = -r^2 a d\tau^2 + 2d\tau dr + r^2 \tau^2 e^{2b-2c} \left(1 + (r\tau)^{-1}\right)^2 dy^2 + r^2 e^c d\vec{x}_\perp^2$$

- $r$  is the radial (5th) direction, the boundary is at  $r = \infty$ .
- the boundary condition (LRF):  
 $a \rightarrow 1, b \rightarrow 0, c \rightarrow 0, \text{ at } r = \infty.$

The Einstein's equation



Differential equations for  $a, b, c$ .

→ Difficult to get the exact solution.

# Approximation:

Hydrodynamics:  $\tau^{-2/3}$  expansion works for Bjorken flow.

Expansion around equilibrium.

Gravity dual:  $\tau^{-2/3}$  expansion, with  $r$  fixed?

We found that

$\tau^{-2/3}$  expansion with  $u=r \tau^{1/3}$  fixed works well.

$$ds^2 = -r^2 a d\tau^2 + 2d\tau dr + r^2 \tau^2 e^{2b-2c} \left(1 + (r\tau)^{-1}\right)^2 dy^2 + r^2 e^c d\vec{x}_\perp^2$$

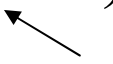
$$a = a_0(u) + a_1(u)\tau^{-2/3} + a_2(u)\tau^{-4/3} + \dots$$

Similar for  $b$  and  $c$ .

Solve the equation order by order.

# The zeroth-order solution (at the leading order)

$$ds^2 = -r^2 \left( 1 - \frac{w^4}{(r\tau^{1/3})^4} \right) d\tau^2 + 2d\tau dr + r^2 (\tau^2 dy^2 + d\vec{x}_\perp^2)$$


 This is u

$$w^4 = \varepsilon_0 \left( \frac{3N_c^2}{8\pi^2} \right)$$

- This reproduces the **correct** zeroth-order stress tensor of the Bjorken flow.

$$T_{\tau\tau} = \varepsilon = \varepsilon_0 \left( \frac{1}{\tau^{4/3}} + \dots \right)$$

- We have an **apparent horizon**.

$$e^F \theta_+ \theta_- = -\frac{9}{2} (1 - u^{-4} w^4) \quad \text{trapped region if } u < w.$$

Normalized product of expansions

The location of the **apparent horizon**:  $u = w + O(\tau^{-2/3})$



# The (event) horizon is necessary

$$\left(R_{\mu\nu\rho\lambda}\right)^2 = 8\left(5 + \frac{9w^8}{u^8}\right) + O(\tau^{-2/3})$$

We have a **physical singularity** at the **origin**.

However, this is **hidden** by the **apparent horizon** at  **$u=w$**  hence the **event horizon** (outside it).

→ **Not** a **naked singularity**.

# The first-order solution (at the order of $\tau^{-2/3}$ )

$$ds^2 = -r^2 a d\tau^2 + 2d\tau dr + r^2 \tau^2 e^{2b-2c} \left(1 + (r\tau)^{-1}\right)^2 dy^2 + r^2 e^c d\vec{x}_\perp^2$$

$$a_1 = -\frac{2(1 + \xi_1)u^4 + \xi_1 w^4 - 3\eta_0 u w^4}{3u^5}$$

$$b_1 = -\frac{1 + \xi_1}{u}$$

gauge degree of freedom

$$c_1 = \frac{1}{3w} \left\{ \arctan(uw^{-1}) - \frac{\pi}{2} + \frac{1}{2} \log(u-w) - \frac{1}{2} \log(u+w) \right\} \\ - \frac{1}{2} \eta_0 \log(1 - w^4 u^{-4}) - \frac{2\xi_1}{3u}$$

The dual geometry is regular at the horizon, only when

$$(R_{\mu\nu\rho\lambda})^2 = \frac{4(9\eta_0^2 w^2 - 1)}{3(u-w)} \left( \frac{1}{u-w} - 2 \right) \tau^{-2/3} + \text{regular} + O(\tau^{-4/3})$$

$$\eta_0 = \frac{1}{3w} .$$

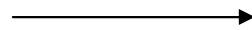
# What is $\eta_0$ ?

Read the **stress tensor** from the geometry



$\eta_0$  is found to be proportional to the **shear viscosity  $\eta$** .

$$\eta_0 = \frac{1}{3w}$$



$$\frac{\eta}{s} = \frac{1}{4\pi}$$

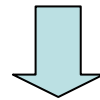
The famous ratio by Kovtun-Son-Starinets (2004)

# Higher-order results:

- From the regularity of the 2nd-order geometry, “relaxation time” is uniquely determined.

consistent with Heller-Janik, Baier et. al., and Bhattacharyya et. al.

Furthermore, we can show by using induction that the same mechanism holds to arbitrary higher order.



No un-removable singularity like

Benincasa-Buchel-heller-Janik, arXiv:0712.2025.

Our model is totally well-defined.

# Non-staticity of the local geometry

## Projected Weyl tensor

$$C_{x^1x^2}^{x^1x^2} = \frac{w^4}{u^4} - \frac{4w^4}{3u^5} \tau^{-2/3} + \dots$$

$$C_{x^1y}^{x^1y} = \frac{w^4}{u^4} - \left( \frac{4w^4}{3u^5} + \frac{3\eta_0 w^4}{u^4} \right) \tau^{-2/3} + \dots$$

An-isotropy **evolves** in time.

The dual geometry is **not static**, if we include **dissipation**.

# Area of the apparent horizon

entropy creation due to dissipation

$$A_{ap} = w^3 - \frac{3w^3 \eta_0}{2} \tau^{-2/3} + \dots$$

**Consistent** with the time evolution of the **entropy density** to the first order.

# What we have done

- Gravity dual of the Bjorken flow on an Eddington-Finkelstein type coordinates.
- Presence of the apparent horizon and the event horizon: a dynamical black hole.
- Transport coefficients from the absence of naked singularity.
- Our late-time expansion is consistent and the geometry can be regular for all orders.

# Advertisement

- Our model provides **the first consistent gravity dual** of the **Bjorken flow**.  
(cf. Heller-Loganayagam-Spalinski-Surowka-Vazquez, arXiv:0805.3774)
- Our model is a **concrete well-defined example** of **time-dependent AdS/CFT** based on a well-controlled approximation.



# I hope that

I **hope** that the gravity dual provides a computational method **beyond** what we already know.

I hope that:

Einstein + Penrose <sup>?</sup> > Kubo + Landau

Challenging problems:

Description of phenomena that appear **only**  
**in time-dependent systems**, such as

- anomalous viscosity
- turbulence
- .....

## For the details

- Please take a look at our paper  
arXiv:0807.3797.
- Please feel free to make a contact with me.