

Domain wall solitons and Hopf algebraic translational symmetries in noncommutative field theories

Yuya Sasai (YITP)

in collaboration with N.Sasakura (YITP)

Based on arXiv:0711.3059 (Phys.Rev.D77:045033,2008)

YITP workshop
on July 29, 2008

1.Introduction

- We believe that noncommutative field theories are important subjects for studying Planck scale physics, especially quantum gravity.
- Recently, it was pointed out that noncommutative field theories would have nontrivial symmetries, which have **Hopf algebraic structure**.

Moyal plane: Chaichian, Kulish, Nishijima, Tureanu (2004), etc.

Noncommutative gravity: Aschieri, Dimitrijevic, Meyer, Schupp, Wess (2005), etc.

SU(2) noncommutative spacetime: Freidel, Livine (2005), etc.

- In order for quantum field theories to possess Hopf algebraic symmetries, we have to include **braiding (nontrivial statistics)**.

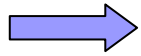
SU(2) noncommutative spacetime: Freidel, Livine (2005)

Moyal plane: Balachandran, Mangano, Pinzul, Vaidya (2006)

General case: Y.S, Sasakura (2007)

- In addition to the importance of the Hopf algebraic symmetries, the braiding can recover the unitarity and renormalization.

- We want to study more physical aspects of Hopf algebraic symmetry.



How is “symmetry breaking” of Hopf algebraic symmetry?

- We study a **domain wall soliton** in three dimensional noncommutative field theory in Lie-algebraic noncommutative space-time

$$[x^i, x^j] = 2i\kappa\epsilon^{ijk}x_k \quad .$$

- It is interesting to consider a domain wall soliton in the Lie-algebraic noncommutative space-time because

1. What is the generator of a one-parameter family of domain wall solutions which comes from a Hopf algebraic translational symmetry?

2. Is the moduli field on a domain wall massless?

Contents

1. Introduction

2. Review of noncommutative field theory in the Lie-algebraic noncommutative spacetime $[x^i, x^j] = 2i\kappa\epsilon^{ijk}x_k$

3. Derrick's theorem in the noncommutative ϕ^4 theory

4. Domain wall solitons and the moduli fields in the Lie algebraic noncommutative spacetime

5. Summary

2. Review of noncommutative field theory in the Lie-algebraic noncommutativity

$$[\hat{x}^i, \hat{x}^j] = 2i\kappa\epsilon^{ijk}\hat{x}_k, \quad (i, j, k = 0, 1, 2)$$

Imai, Sasakura (2000), Freidel, Livine (2005)

Commutation relation

$$[\hat{x}^i, \hat{x}^j] = 2i\kappa\epsilon^{ijk}\hat{x}_k,$$

$$[\hat{P}^i, \hat{x}^j] = -i\eta^{ij}\sqrt{1 + \kappa^2\hat{P}^2} + i\kappa\epsilon^{ijk}\hat{P}_k,$$

$$[\hat{P}^i, \hat{P}^j] = 0.$$

Lorentz invariance and Jacobi identity are satisfied.

These operators can be identified with Lie algebra of ISO(2,2)

$$\hat{x}_i = \kappa(\hat{J}_{-1,i} - \frac{1}{2}\epsilon_i^{jk}\hat{J}_{jk}),$$

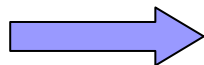
$$\hat{P}_i = \hat{P}_{\mu=i},$$

$$\left(\hat{M}_i = \kappa(\hat{J}_{-1,i} + \frac{1}{2}\epsilon_i^{jk}\hat{J}_{jk}), \right)$$

with the constraint

$$1 + \kappa^2\hat{P}^\mu\hat{P}_\mu = 0$$

$$(\mu = -1, 0, 1, 2)$$



SL(2,R) group momentum space

where ISO(2,2) Lie alg. is given by

$$[\hat{J}_{\mu\nu}, \hat{J}_{\rho\sigma}] = -i(\eta_{\mu\rho}\hat{J}_{\nu\sigma} - \eta_{\mu\sigma}\hat{J}_{\nu\rho} - \eta_{\nu\rho}\hat{J}_{\mu\sigma} + \eta_{\nu\sigma}\hat{J}_{\mu\rho}),$$

$$[\hat{J}_{\mu\nu}, \hat{P}_\rho] = -i(\eta_{\mu\rho}\hat{P}_\nu - \eta_{\nu\rho}\hat{P}_\mu),$$

$$[\hat{P}_\mu, \hat{P}_\nu] = 0,$$

Scalar field

$$\phi(x) = \int dg \tilde{\phi}(g) e^{iP(g) \cdot x}$$

$$g = P^{-1}(g) - i\kappa P^i(g) \tilde{\sigma}_i \in SL(2, R)$$

$$\tilde{\sigma}_0 = \sigma_2, \quad \tilde{\sigma}_1 = i\sigma_3, \quad \tilde{\sigma}_2 = i\sigma_1$$

$$P^{-1}(g) = \pm \sqrt{1 + \kappa^2 P^i(g) P_i(g)}$$



From now on, we pay attention to the positive sign for simplicity.

Star product

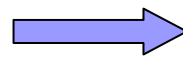
$$e^{iP(g_1) \cdot x} \star e^{iP(g_2) \cdot x} = e^{iP(g_1 g_2) \cdot x} ,$$

where

$$P^i(g_1 g_2) = P^i(g_1) \sqrt{1 + \kappa^2 P(g_2)^2} + \sqrt{1 + \kappa^2 P(g_1)^2} P^i(g_2) - \kappa \epsilon^{ijk} P_j(g_1) P_k(g_2) .$$

Thus coproduct of the translational operator is given by

$$\Delta(P^i) = P^i \otimes \sqrt{1 + \kappa^2 P^l P_l} + \sqrt{1 + \kappa^2 P^l P_l} \otimes P^i - \kappa \epsilon^{ijk} P_j \otimes P_k$$



Hopf algebraic structure!

Thus, the momentum conservation is deformed.

Action

$$S = \int d^3x \left[-\frac{1}{2}(\partial^i \phi \star \partial_i \phi)(x) - \frac{1}{2}m^2(\phi \star \phi)(x) - \frac{\lambda}{4}(\phi \star \phi \star \phi \star \phi)(x) \right]$$

- At the quantum level, **deformed momentum conservation is violated**
in the non-planar diagrams.

Imai, Sasakura (2000)

- To keep the momentum conservation law, we have to introduce **braiding**
such that

$$\psi(\tilde{\phi}_1(g_1)\tilde{\phi}_2(g_2)) = \tilde{\phi}_2(g_2)\tilde{\phi}_1(g_2^{-1}g_1g_2) \quad ,$$

which was discovered in three dimensional quantum gravity.

Freidel, Livine (2005)

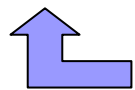
3. Derrick's theorem in the noncommutative ϕ^4 theory

Action

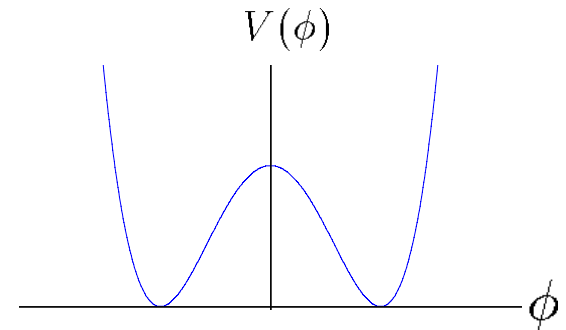
$$S = \int d^3x \left[-\frac{1}{2}(\partial^i \phi \star \partial_i \phi)(x) + \frac{1}{2}m^2(\phi \star \phi)(x) - \frac{\lambda}{4}(\phi \star \phi \star \phi \star \phi)(x) - \frac{m^4}{4\lambda} \right]$$

Equation of motion for $\phi(x)$

$$\partial^2 \phi(x) + m^2 \phi(x) - \lambda(\phi \star \phi \star \phi)(x) = 0$$

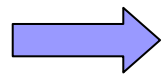


- Star product is included.
- Translational symmetry is not clear.



- Consider only one spatial direction and define

$$P = \frac{1}{\kappa} \sinh(\kappa\theta) \quad , \text{ where } -\infty < \theta < \infty.$$



$$\left\{ \begin{array}{l} \phi(x) = \int dg \tilde{\phi}(g) e^{iP(g)x} = \int \frac{d\theta}{2\pi} \tilde{\phi}(\theta) e^{\frac{i}{\kappa} \sinh(\kappa\theta)x} \\ e^{iP(g_1)x} \star e^{iP(g_2)x} = e^{iP(g_1 g_2)x} \\ \quad \longrightarrow e^{\frac{i}{\kappa} \sinh(\kappa\theta_1)x} \star e^{\frac{i}{\kappa} \sinh(\kappa\theta_2)x} = e^{\frac{i}{\kappa} \sinh(\kappa(\theta_1 + \theta_2))x} \end{array} \right.$$

usual sum!



- Next, we define

$$h(x) = \int \frac{d\theta}{2\pi} \tilde{\phi}(\theta) e^{i\theta x}.$$

- Equation of motion becomes

$$\frac{1}{\kappa^2} \sin^2(\kappa \partial) h(x) + m^2 h(x) - \lambda h^3(x) = 0.$$

- no star product
- translation invariant

- To analyze this, we consider an action for $h(x)$

$$S_h = \int dx \left[-\frac{1}{2\kappa^2} \sin(\kappa \partial) h(x) \sin(\kappa \partial) h(x) + \frac{1}{2} m^2 h^2(x) - \frac{\lambda}{4} h^4(x) - \frac{m^4}{4\lambda} \right].$$

- Expanding with κ , the energy is

$$E_h = -S_h = \int dx \left[\frac{1}{2} \left(\sum_{n=1}^{\infty} \kappa^{2n-2} C_n \partial^n h(x) \partial^n h(x) \right) + V(h(x)) \right],$$

where $C_n = 2^{n-1} / (n!(2n-1)!!) > 0$

$$V(h(x)) = -\frac{1}{2} m^2 h^2(x) + \frac{\lambda}{4} h^4(x) + \frac{m^4}{4\lambda} \geq 0.$$

- Rescaling $x^i \rightarrow x'^i = \mu x^i$ ($0 < \mu < \infty$) and defining $h^{(\mu)}(x) = h(\mu x)$, the energy for $h^{(\mu)}(x)$ becomes

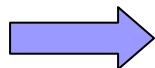
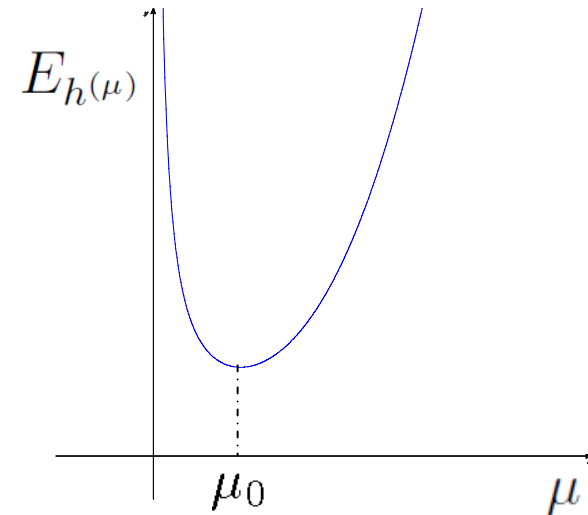
$$\begin{aligned}
 E_{h^{(\mu)}} &= \int dx \left[\frac{1}{2} \left(\sum_{n=1}^{\infty} \kappa^{2n-2} C_n (\partial^n h^{(\mu)}(x))^2 \right) + V(h^{(\mu)}(x)) \right] \\
 &= \int dx' \frac{1}{\mu} \left[\frac{1}{2} \left(\sum_{n=1}^{\infty} \kappa^{2n-2} C_n (\partial'^n h(x'))^2 \right) + V(h(x')) \right] \\
 &= \frac{1}{\mu} E_0 + \sum_{n=1}^{\infty} \mu^{2n-1} E_{2n},
 \end{aligned}$$

where

$$E_0 = \int dx V(h(x)),$$

$$E_{2n} = \frac{C_n}{2} \int dx (\partial^n h(x))^2.$$

- Since all the E_0 and E_{2n} are non-negative, $E_{h^{(\mu)}}$ takes a minimal value at a positive finite μ .



Domain wall solitons may exist.

4. Domain wall solitons and the moduli fields in Lie algebraic noncommutative spacetime

4.1 The moduli space of the domain wall and the generator

- Equation of motion for $h(x)$ is given by

$$\frac{1}{\kappa^2} \sin^2(\kappa \partial) h(x) + m^2 h(x) - \lambda h^3(x) = 0.$$

This equation is invariant under the usual translation $x \rightarrow x + a$

- We can obtain the perturbative solutions as follows.

$$h(x) = h_0(x) + \kappa^2 h_2(x) + \kappa^4 h_4(x) + \dots$$

$$h_0(x) = \tanh(x + a),$$

$$h_2(x) = \frac{2(x + a)}{3 \cosh^2(x + a)} - \frac{4 \tanh(x + a)}{3 \cosh^2(x + a)}$$

$$h_4(x) = \frac{134(x + a)}{45 \cosh^2(x + a)} - \frac{8(x + a)}{3 \cosh^4(x + a)} - \frac{40 \tanh(x + a)}{9 \cosh^2(x + a)} - \frac{4(x + a)^2 \tanh(x + a)}{9 \cosh^2(x + a)} + \frac{52 \tanh(x + a)}{9 \cosh^4(x + a)}.$$

where

$$m^2 = 2, \lambda = 2.$$

- Solutions of $\phi(x)$ are in principle given from $h(x + a)$.

$$h(x) = \int \frac{d\theta}{2\pi} \tilde{\phi}(\theta) e^{i\theta x}$$

$$\downarrow$$

$$\phi(x) = \int \frac{d\theta}{2\pi} \tilde{\phi}(\theta) e^{\frac{i}{\kappa} \sinh(\kappa\theta)x}$$

$$h(x + a) = \int \frac{d\theta}{2\pi} \tilde{\phi}(\theta) e^{i\theta a} e^{i\theta x}$$

$$\downarrow$$

$$T_a \phi(x) = \int \frac{d\theta}{2\pi} \tilde{\phi}(\theta) e^{i\theta a} e^{\frac{i}{\kappa} \sinh(\kappa\theta)x}$$

$$= e^{ia\hat{\theta}} \phi(x)$$

where $\hat{\theta} = \frac{1}{\kappa} \sinh^{-1}(\kappa \hat{P})$

Thus, $\hat{\theta}$ is the generator of a one-parameter family of domain wall solutions.

In the usual case ($\kappa \rightarrow 0$), the generator of a one-parameter family of domain wall solutions is given by $\hat{\theta} = \hat{P}$.

4.2 The moduli field from the Hopf algebraic translational symmetry

- Let us assume $\phi_{sol}^a(x^1)$ is a general solution of the one dimensional equation of motion for $\phi(x)$. We expand $\phi_{sol}^a(x^1)$ with respect to a as

$$\phi_{sol}^a(x^1) = \phi_{sol}(x^1) + \underline{a} g(x^1) + \dots,$$

where $g(x^1)$ should satisfy the following equation,

$$\partial^2 g(x^1) + m^2 g(x^1) - 3\lambda \phi_{sol}(x^1) \star \phi_{sol}(x^1) \star g(x^1) = 0.$$

- In order to obtain an equation for a moduli field, we replace a to a moduli field $a(x^0, x^2)$

$$\phi(x) = \phi_{sol}(x^1) + \underline{a(x^0, x^2)} \star g(x^1) + \dots.$$

This should satisfy the three dimensional equation of motion for $\phi(x)$.

- Inserting $\phi(x) = \phi_{sol}(x^1) + a(x^0, x^2) \star g(x^1)$ into the equation of motion, and taking the first order of $a(x^0, x^2)$, we obtain

$$\partial^2(a(x^0, x^2) \star g(x^1)) + m^2(a(x^0, x^2) \star g(x^1)) - 3\lambda\phi_{sol}(x^1) \star \phi_{sol}(x^1) \star a(x^0, x^2) \star g(x^1) = 0,$$

where we have used the braiding property

$$\phi_1(x) \star \phi_2(x) = \phi_2(x) \star \phi_1(x).$$

- Using the equation for $g(x^1)$, we obtain

$$(g(x^1) - \partial^2 g(x^1)) \star \partial^2 a(x^0, x^2) = 0.$$

- Since $g(x^1) - \partial^2 g(x^1) \neq 0$, we obtain

$$\partial^2 a(x_0, x_2) = 0.$$

Thus we find that the moduli field, which propagates on the domain wall, is massless.

5. Summary

- We studied the **domain wall solution** and its **moduli** in the Lie-algebraic noncommutative space-time.
- We found that the **generator of a one-parameter family of the domain wall solutions** is given by $\hat{\theta} = \frac{1}{\kappa} \sinh^{-1}(\kappa \hat{P})$
- We checked the **moduli field** on the domain wall is **massless**.

Question

- **A scalar field $\phi(x)$ is not a c-number in braided quantum field theory.**
Can we interpret the classical solutions with the braid statistics physically?