

# Supersymmetry breaking in moduli-mixing racetrack model

YITP, Kyoto University    Hiroyuki Abe

E-mail: [abe@yukawa.kyoto-u.ac.jp](mailto:abe@yukawa.kyoto-u.ac.jp)

In this talk we show some structures of moduli stabilization and SUSY breaking in so-called racetrack model with double gaugino condensations where gauge couplings are given by more than one modulus field<sup>1</sup> (two moduli in practice). The 4D effective supergravity (within type IIB  $O3/O7$  framework for concreteness) is given by the Kähler potential and superpotential

$$K = -n_T \ln(T + \bar{T}) - n_S \ln(S + \bar{S}), \quad W = Ae^{-af_a} - Be^{-bf_b}, \quad f_{a,b} = m_{a,b}S + w_{a,b}T, \quad (1)$$

where  $S$  and  $T$  are the dilaton and Kähler (size) modulus respectively, and  $m_{a,b}$ ,  $w_{a,b}$  are respectively the magnetic flux and winding number of the  $D$ -brane, where the gaugino condensation occurs. We assume that the existence of three-form flux (in ten-dimensions) stabilize the complex structure (shape) moduli  $\langle U \rangle \sim 1$  around the Planck scale. Note that, due to the flux, there is a significantly warped region in the Calabi-Yau (CY) space.

If the same three-form flux induces a SUSY mass like  $W \sim SU$ , the dilaton is also stabilized  $\langle S \rangle \sim 1$  at the same scale. In this case we replace  $S$  in Eq. (1) by its VEV,  $\langle S \rangle$ . Then the effective superpotential becomes  $W = A'e^{-aw_a T} - B'e^{-bw_b T}$ , which is in the same form as the racetrack model with single modulus, but the coefficients are exponentially suppressed  $A' = Ae^{-am_a \langle S \rangle}$ ,  $B' = Be^{-bm_b \langle S \rangle}$  where  $a = 8\pi^2/N_a$  and  $b = 8\pi^2/N_b$  for  $SU(N_{a,b})$  gaugino condensation. The minimum of the scalar potential corresponds to a SUSY  $AdS_4$  local minimum with negative vacuum energy and  $a\langle \text{Re} T \rangle = aT_{SUSY} \sim \ln(M_{Pl}/m_{3/2})$ , where  $m_{3/2} \approx 10^{-14}M_{Pl}$  is the gravitino mass. To be phenomenologically viable, we uplift the vacuum energy by introducing anti  $D3$ -branes at the top of warped region in the CY space. Then the SUSY is broken due to the slight shift  $\delta T = (T_{SUSY} - \langle T \rangle) \ll T_{SUSY}$  caused by an additional potential energy of  $\overline{D3}$ s. In this case, the ratio between the VEV of auxiliary component in the chiral compensator  $C$ ,  $F^C \sim m_{3/2}$ , and one in  $T$ ,  $F^T$ , is given by ( $w_a = 0$  for simplicity)  $\alpha = \frac{F^C}{\ln(M_{Pl}/m_{3/2})} \frac{T+\bar{T}}{F^T} \simeq (1 + m_b \langle S \rangle / w_b T_{SUSY})^{-1}$ . This corresponds to the ratio between the so-called anomaly mediation and the modulus mediation for the visible sector SUSY breaking. We find that  $\alpha$  varies in a wide range with various magnetic flux  $m_{a,b}$ , compared to  $\alpha \sim 1$  without the magnetic flux.

On the other hand, if the three-form flux does not contain the SUSY mass term  $W \sim SU$ , the dilaton remains as a light modulus as well as  $T$  in Eq. (1). A careful analysis of (1) shows that a SUSY  $AdS_4$  stationary point of the scalar potential corresponds to a saddle point, and we have a SUSY breaking  $AdS_4$  local minimum close to the SUSY point in this case. By uplifting this local minimum, we can obtain Minkowski minimum but with modulus-dominated SUSY breaking  $\alpha \ll 1$ . This is because SUSY is broken before uplifting.

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<sup>1</sup>This talk is based on works in collaboration with T. Higaki and T. Kobayashi