

Holographic entanglement entropy of deSitter braneworld

Tetsuya Shiromizu

Math. Dept. / Kobayashi-Maskawa Institute
Nagoya University

[Contents]

1. Introduction
2. Review of braneworld
3. Holographic entanglement entropy in deSitter braneworld
4. Summary

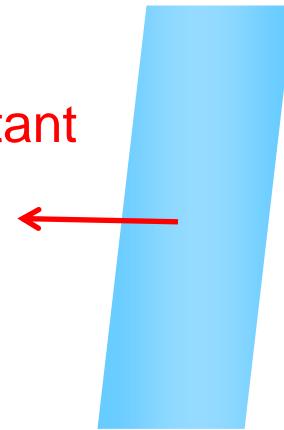
1. Introduction

Braneworld

- Higher dimensional model inspired by string theory

Simplest model: Randall-Sundrum(RS) 1999

5-dim. bulk spacetime with negative cosmological constant
(5-dim. anti-deSitter spacetime)



Gravity associated with brane is
fully taken into account

-applicable to cosmology/black hole
-similar setup to adS/CFT (\Rightarrow braneworld holography)

brane: 4-dim. spacetime
(4-dim. Minkowski spacetime)

[New lights to braneworld]

Holographic entanglement entropy Ryu & Takayanagi 2006

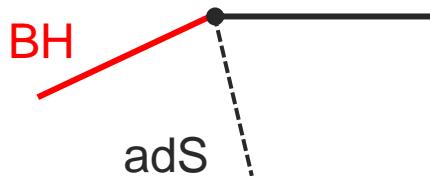
Island formula for Hawking radiation

Penington 2020, Almheiri et al 2019,2020,...

co-dimension 2 holography

Akal, Kusuki, Takayanagi & Wei 2020

Hybrid formulation of adS/CFT and braneworld



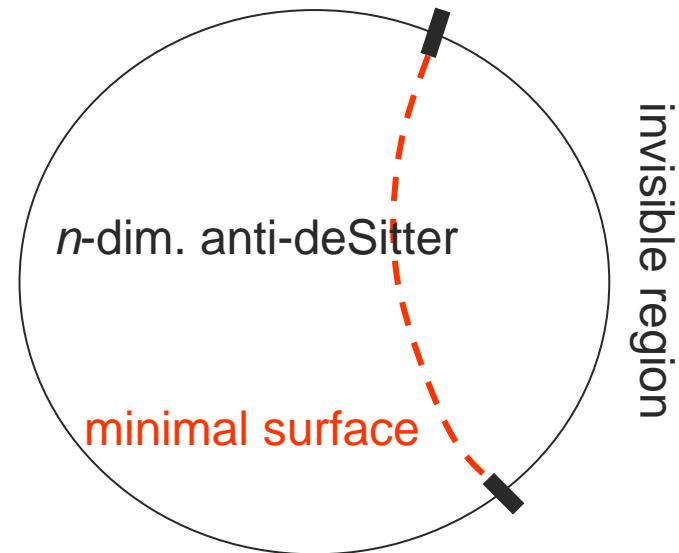
→ Revisit braneworld model with those in mind

Holographic entanglement entropy

in adS/CFT setup

Ryu-Takayanagi formula(2006)

$$S_{\text{ent}} = \frac{\text{area of minimal surface}}{4G_n}$$



Holographic entanglement entropy in braneworld?

- black hole on brane [Emparan 2006](#)

- deSitter spacetime on brane

[Iwashita, Kobayashi, Shiromizu & Yoshino 2006](#)

[deSitter spacetime]

- solution to Einstein equation with **positive cosmological constant**
- maximally symmetric, Lorentz group
- exponentially expanding universe

→ cosmological horizon

analogy to Hawking-Bekenstein formula for black hole entropy

$$S = \frac{c^3 A_{\text{horizon}}}{4G\hbar}$$

→ deSitter entropy

Gibbons-Hawking 1977

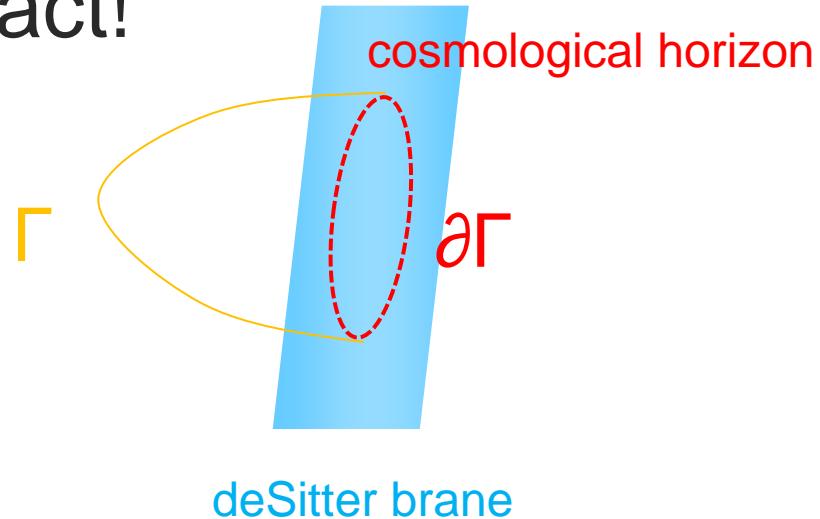
[deSitter braneworld: result]

simplest, but non-trivial, exact!

Ryu-Takayanagi formula

$$S_{\text{ent}} = \text{ext} \left[\frac{\text{Area}(\Gamma)}{4G} \right] ?$$

⇒ Ryu-Takayanagi with curvature corrections
(Hung, Myers & Smolkin 2011)



in Einstein, Gauss-Bonnet, Lovelock

$$S_{\text{ent}} = S_{\text{dS}}$$

Hawking, Maldacena & Strominger 2000,
Iwashita, Kobayashi, Shiromizu & Yoshino 2006
Kushihara, Izumi & Shiromizu 2021

2. Review of braneworld

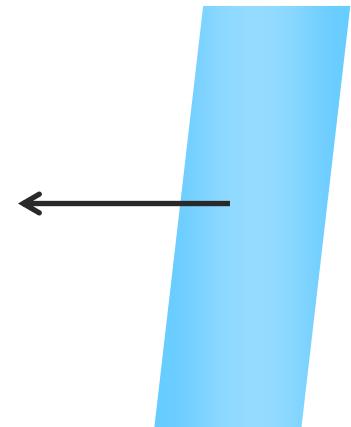
[Randall-Sundrum II]

Randall & Sundrum 1999

5-dim. bulk spacetime with negative cosmological constant

matter is confined on the brane

warped extra dimensions due to self-gravity of brane



brane: 4-dim. spacetime, tension

[4D gravity is recovered]

Randall & Sundrum 1999, Garriga & Tanaka 2000,
Giddings, Katz & Randall 2000

correction to Newton potential

$$V \simeq -\frac{GM}{r} \left(1 + O(\ell^2/r^2) \right)$$

$\ell \cdots$ adS curvature length

[Geometrical projection]

Shiromizu, Maeda & Sasaki 2000

From 5 to 4

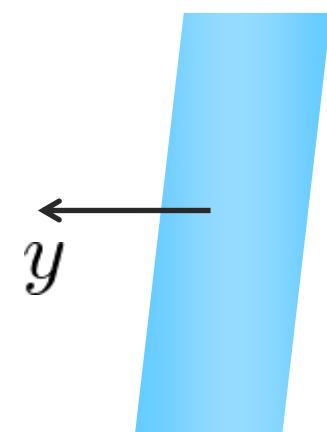
$$R_{MN} - \frac{1}{2}g_{MN}R = \kappa^2 T_{MN} - \Lambda g_{MN}$$

$$T_{MN} = (-\sigma q_{MN} + \tau_{MN})\delta(y)$$

$${}^{(4)}R_{\mu\nu} - \frac{1}{2}q_{\mu\nu}{}^{(4)}R = -\Lambda_4 q_{\mu\nu} + 8\pi G \tau_{\mu\nu} + \kappa^4 \pi_{\mu\nu} - E_{\mu\nu}$$

$$\left\{ \begin{array}{l} \Lambda_4 = \frac{\Lambda}{2} + \frac{\kappa^4 \sigma^2}{12} \\ 8\pi G = \frac{\kappa^4 \sigma}{6} \\ \pi_{\mu\nu} = -\frac{1}{4}\tau_{\mu\alpha}\tau_{\nu}^{\alpha} + \frac{1}{12}\tau\tau_{\mu\nu} + \frac{1}{8}q_{\mu\nu}\tau_{\alpha\beta}\tau^{\alpha\beta} - \frac{1}{24}q_{\mu\nu}\tau^2 \\ E_{\mu\nu} = C_{\mu y \nu y} \end{array} \right.$$

a part of 5-dimensional Weyl tensor



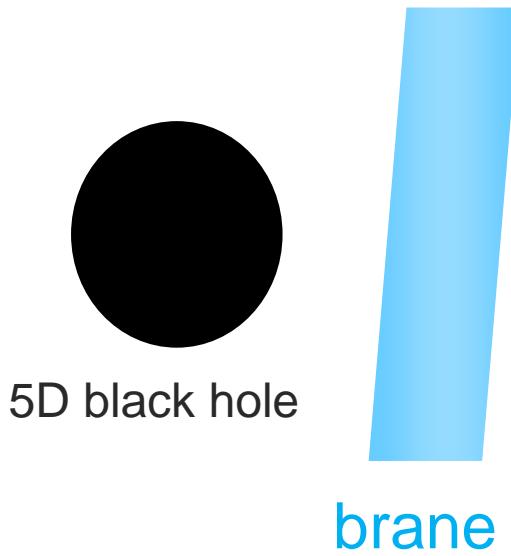
$(M_4, q_{\mu\nu})$

$$[E_{\mu\nu}$$

$${}^{(4)}R_{\mu\nu} - \frac{1}{2}q_{\mu\nu}{}^{(4)}R = -\Lambda_4 q_{\mu\nu} + 8\pi G\tau_{\mu\nu} + \kappa^4\pi_{\mu\nu} - E_{\mu\nu}$$

]

Binetury, Deffayet, Ellwanger & Langlois 2000,
Ida 2000, Mukohyama 2000,
Mukohyama, Shiromizu & Maeda 2000, ...



$E_{\mu\nu}$ ~ holographic 4D radiation
proportional to black hole mass

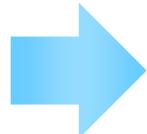
[Braneworld holography]

Gubser 2001, Giddings, Katz & Randall 2000,
Hawking, Hertog & Reall 2000, Shiromizu & Ida 2001,
Koyama & Soda 2001, Soda & Kanno 2002,...

$$\begin{aligned} {}^{(4)}R &= -8\pi G\tau - \frac{\kappa^2}{4} \left(\tau_{\mu\nu}\tau^{\mu\nu} - \frac{1}{3}\tau^2 \right) \\ &\simeq -8\pi G\tau - \frac{\ell^2}{4} \left({}^{(4)}R_{\mu\nu}{}^{(4)}R^{\mu\nu} - \frac{1}{3}{}^{(4)}R^2 \right) \end{aligned}$$



trace anomaly of CFT



4D gravity+matter+holographic CFT

[Black hole on brane]

Exact solution?

-not discovered in 5D.

only numerical studies

Shiromizu & Shibata 2000, Wiseman 2002&2003,
Kudoh, Tanaka & Nakamura 2003, Kudoh 2004, Tanahashi & Tanaka 2008,
Yoshino 2009, Figueras & Wiseman 2011, Wang & Choptuik 2016

-discovered in 4D.

Emparan, Horowitz & Myers 2000

[deSitter on brane]

Nihei(1999), Kaloper(1999), Garriga & Sasaki(2000),...

$$ds^2 = dr^2 + (\ell H)^2 \sinh^2(r/\ell) [-dt^2 + H^{-2} \cosh^2(Ht) d\Omega_{n-2}^2]$$

(n-1)-dimensional deSitter spacetime in complete chart

$$= dr^2 + (\ell H)^2 \sinh^2(r/\ell) [-(1 - H^2 \rho^2) dT^2 + (1 - H^2 \rho^2)^{-1} d\rho^2 + \rho^2 d\Omega_{n-3}^2]$$

(n-1)-dimensional deSitter spacetime in static chart

$$\Lambda = -\frac{(n-1)(n-2)}{\ell^2} \quad \ell \dots \text{adS curvature length}$$

$$r = r_0 \quad \text{brane location}$$

$$H^{-1} = \ell \sinh(r_0/\ell) \quad H: \text{expansion rate}$$

$$\frac{n-2}{\ell} \cosh(r_0/\ell) = 4\pi G_n \sigma \quad \sigma: \text{brane tension}$$

[cosmological horizon]

deSitter spacetime

$$ds^2 = -(1 - H^2 \rho^2) dT^2 + (1 - H^2 \rho^2)^{-1} d\rho^2 + \rho^2 d\Omega_2^2$$

$$\rho = H^{-1} \quad \text{cosmological horizon}$$

cf) Schwarzschild black hole

$$ds^2 = -\left(1 - \frac{2Gm}{r}\right) dt^2 + \left(1 - \frac{2Gm}{r}\right)^{-1} dr^2 + r^2 d\Omega_2^2$$

$$r = 2Gm \quad \text{event horizon } (m: \text{mass})$$

$$S_{\text{Bekenstein-Hawking}} = \frac{A_{\text{horizon}}}{4G} \sim \left(\frac{\text{horizon size}}{l_{\text{pl}}}\right)^2$$

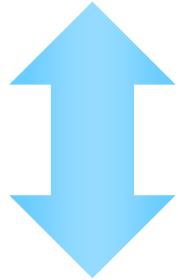
4. Holographic entanglement entropy in deSitter braneworld

Holographic entanglement entropy v.s. deSitter entropy

Braneworld holography

Holographic entanglement entropy in braneworld?

Ryu-Takayanagi (+Jacobson-Myers) formula



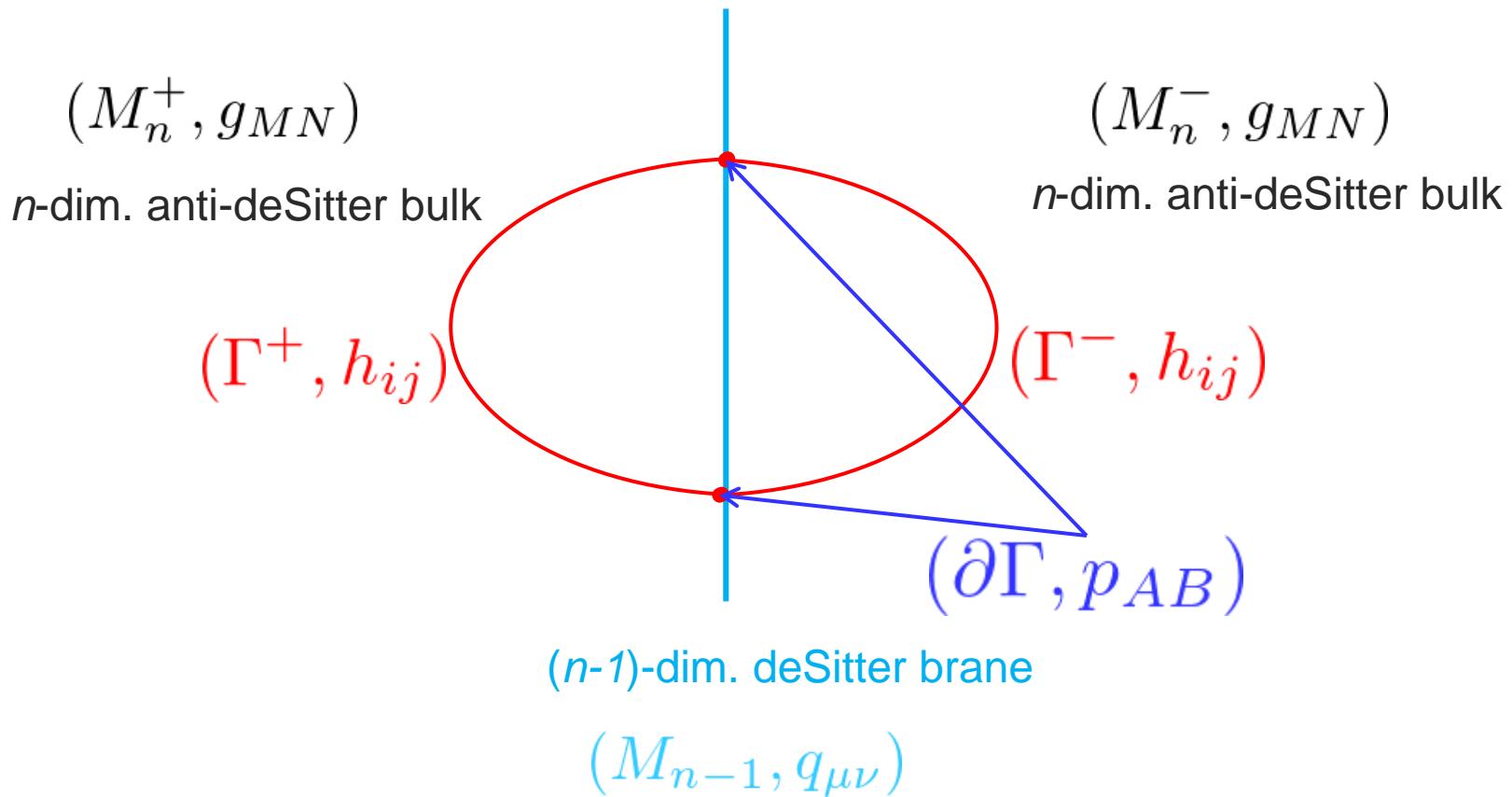
$$S_{\text{ent}} = \frac{A_{n-2}}{4G_n} \quad \text{area of minimal surface}$$

deSitter entropy in braneworld

Euclidean path integral I_E in full spacetime

$$S_{\text{dS}} = \beta \langle E \rangle - I_E = -I_E$$

[Common setup]



[in Einstein]

Hawking, Maldacena & Strominger (2001),
Iwashita, Kobayashi, Shiromizu & Yoshino (2006)

$$S_{\text{dS}} = -I_E = \frac{1}{16\pi G_n} \int_{M_n^+ \cup M_n^-} d^n x \sqrt{g} (R - 2\Lambda) + \int_{M_{n-1}} d^{n-1} x \sqrt{q} \left(-\sigma + \frac{[K]^-}{8\pi G_n} \right)$$

$$= \frac{(n-2)\Omega_{n-1}\ell^{n-2}}{4\pi G_n} \int_0^{r_0/\ell} dx \sinh^{n-3} x$$

$$S_{\text{ent}} = \frac{A_{\min}}{4G_n} = \frac{\Omega_{n-3}\ell^{n-2}}{2G_n} \int_0^{r_0/\ell} dx \sinh^{n-3} x$$

$$\Omega_{n-3} = \Omega_{n-1} \frac{n-2}{2\pi}$$



$$S_{\text{dS}} = S_{\text{ent}}$$

[in Gauss-Bonnet]

higher curvature terms

Maeda & Torii (2004), ...

$$S = \frac{1}{16\pi G_n} \int_{M_n^+ \cup M_n^-} d^n x \sqrt{-g} \left(R - 2\Lambda + \frac{\beta \ell^2}{4} \mathcal{L}_{\text{GB}} \right) + \int_{M_{n-1}} d^{n-1} x \sqrt{-q} \left(-\sigma + \frac{1}{16\pi G_n} [Q]^- \right)$$

$$\left. \begin{array}{l} \mathcal{L}_{\text{GB}} = R^2 - 4R_{AB}R^{AB} + R_{ABCD}R^{ABCD} \\ Q = 2K + \beta \ell^2 (J - 2^{(n-1)}G_{\mu\nu}K^{\mu\nu}) \end{array} \right\}$$

Hung, Myers & Smolkin (2011)

$$S_{\text{JM}} = \frac{1}{4G_n} \int_{\Gamma^+ \cup \Gamma^-} d^{n-2} x \sqrt{h} \left(1 + \frac{\beta \ell^2}{2} {}^{(n-2)}R \right) + \frac{1}{2G_n} \int_{\partial\Gamma} d^{n-3} x \sqrt{p} \frac{\beta \ell^2}{2} [{}^{(n-3)}k]^-$$

variation 

$$\left. \begin{array}{l} {}^{(n-2)}k - \beta \ell^2 {}^{(n-2)}G_{ij} {}^{(n-2)}k^{ij} = 0 \quad (\text{in bulk}) \\ \beta \left[{}^{(n-3)}k^{AB} - {}^{(n-3)}kp^{AB} \right]^- {}^{(n-2)}k_{AB} = 0 \quad (\text{on brane}) \end{array} \right\}$$

[HEE and deSitter entropy]

Kushihara, Izumi & Shiromizu 2021

$$S_{\text{dS}} = S_{\text{ent}}$$

$$\begin{aligned} &= \frac{(n-2)\ell^{n-2}}{4\pi G_n} \Omega_{n-1} \left(\left[1 - \beta \frac{(n-2)(n-3)}{2} \right] \int_0^{r_0/\ell} dx \sinh^{n-3} x \right. \\ &\quad \left. + \beta(n-3) \sinh^{n-4}(r_0/\ell) \cosh(r_0/\ell) \right) \end{aligned}$$

$$S_{\text{dS}} = S_{\text{ent}}$$

]

Lovelock setup

$$S = \frac{1}{16\pi G_n} \int_{M_n^+ \cup M_n^-} d^n x \sqrt{-g} \left(-2\Lambda + \sum_m c_m \mathcal{L}_m \right) + \int_{M_{n-1}} d^{n-1} x \sqrt{-q} \left(-\sigma + \sum_m c_m \frac{[Q_m]^-}{16\pi G_n} \right)$$

$$\begin{cases} \mathcal{L}_m = \frac{1}{2^m} g_{M_1 N_1 \dots M_m N_m}^{K_1 L_1 \dots K_m L_m} R_{K_1 L_1}{}^{M_1 N_1} \dots R_{K_m L_m}{}^{M_m N_m} \\ Q_m = \frac{4m}{2^m} \int_0^1 ds q_{\mu_1 \nu_1 \dots \mu_{m-1} \nu_{m-1} \mu_m}^{\alpha_1 \beta_1 \dots \alpha_{m-1} \beta_{m-1} \alpha_m} \left({}^{(n-1)}R_{\alpha_1 \beta_1}{}^{\mu_1 \nu_1} - 2s^2 K_{\alpha_1}^{\mu_1} K_{\beta_1}^{\nu_1} \right) \\ \quad \dots \left({}^{(n-1)}R_{\alpha_{m-1} \beta_{m-1}}{}^{\mu_{m-1} \nu_{m-1}} - 2s^2 K_{\alpha_{m-1}}^{\mu_{m-1}} K_{\beta_{m-1}}^{\nu_{m-1}} \right) K_{\alpha_m}^{\mu_m} \\ = \frac{4m}{2^m} q_{\mu_1 \nu_1 \dots \mu_{m-1} \nu_{m-1} \mu_m}^{\alpha_1 \beta_1 \dots \alpha_{m-1} \beta_{m-1} \alpha_m} \sum_{k=0}^{m-1} \binom{m-1}{k} \frac{(-2)^k}{2k+1} K_{\alpha_1}^{\mu_1} K_{\beta_1}^{\nu_1} \dots K_{\alpha_k}^{\mu_k} K_{\beta_k}^{\nu_k} K_{\alpha_m}^{\mu_m} \\ \quad {}^{(n-1)}R_{\alpha_{k+1} \beta_{k+1}}{}^{\mu_{k+1} \nu_{k+1}} \dots {}^{(n-1)}R_{\alpha_{m-1} \beta_{m-1}}{}^{\mu_{m-1} \nu_{m-1}} \end{cases}$$

[HEE formula in Lovelock]

$$S_{\text{JM}} = \frac{1}{4G_n} \int_{\Gamma^+ \cup \Gamma^-} d^{n-2}x \sqrt{h} \sum_m c_m \tilde{\mathcal{L}}_m + \frac{1}{4G_n} \int_{\partial\Gamma} d^{n-3}x \sqrt{p} \sum_m c_m [\tilde{Q}_m]^-$$

Hung, Myers & Smolkin (2011)

$$\left\{ \begin{array}{l} \tilde{\mathcal{L}}_m = \frac{m}{2^{m-1}} h^{i_1 j_1 \dots i_{m-1} j_{m-1}}_{k_1 l_1 \dots k_{m-1} l_{m-1}} {}^{(n-2)} R_{i_1 j_1}{}^{k_1 l_1} \dots {}^{(n-2)} R_{i_{m-1} j_{m-1}}{}^{k_{m-1} l_{m-1}} \\ \tilde{Q}_m = \frac{4m(m-1)}{2^{m-1}} \int_0^1 ds p^{A_1 B_1 \dots A_{m-2} B_{m-2} A_{m-1}}_{C_1 D_1 \dots C_{m-2} D_{m-2} C_{m-1}} \left({}^{(n-3)} R_{A_1 B_1}{}^{C_1 D_1} - 2s^{2(n-3)} k^{C_1 (n-3)}_{A_1} k^{D_1}_{B_1} \right) \\ \quad \dots \left({}^{(n-2)} R_{A_{m-2} B_{m-2}}{}^{C_{m-2} D_{m-2}} - 2s^{2(n-3)} k^{C_{m-2} (n-3)}_{A_{m-2}} k^{D_{m-2}}_{B_{m-2}} \right) {}^{(n-3)} k^{C_{m-1}}_{A_{m-1}} \end{array} \right.$$

[HEE and deSitter entropy]

Kushihara, Izumi & Shiromizu 2021

$$S_{\text{dS}} = S_{\text{ent}}$$

$$\begin{aligned} &= \frac{\Omega_{n-1}}{4\pi G_n} \sum_m m c_m \ell^{n-2m} \frac{(n-2)!}{(n-2m)!} \left((-1)^m (n-1) \int_0^{r_0/\ell} dx \sinh^{n-1} x \right. \\ &\quad \left. + (2m-1) \cosh(r_0/\ell) \sinh^{n-2m}(r_0/\ell) \int_0^1 ds \left(1 - s^2 \cosh^2(r_0/\ell)\right)^{m-1} \right) \end{aligned}$$

$$S_{\text{dS}} = S_{\text{ent}}$$

5. Summary and issues

[Summary]

In deSitter braneworld

**Holographic entanglement entropy
= deSitter entropy**

[Issues

- General formulation for holographic entanglement entropy in braneworld?
- Extension to co-dimension 2 braneworld?
- Revisit braneworld black hole?
- ...



Hints for quantum gravity