## Probing axion dark matter with magnon

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1. Axion: axion DM, axion-electron interaction

2. Magnon as collective spin excitation of electrons

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### **Axion**

QCD axion: a Nambu-Goldstone boson of the broken Peccei-Quinn symmetry (for resolving the strong CP problem)

#### Invisible axion models

The KSVZ model J.E.Kim (1979),

M.A.Shifman, A.Vainshtein, V.I.Zakharov (1980)

The DFSZ model M.Dine, W.Fischler, M.Srednicki (1981),
 A.Zhitnitsky (1980)

(QCD) axion is a strong candidate for dark matter.

How can we detect?

#### Axion-electron interaction

An axion can interact with the electron:

$$\mathcal{L}_{\rm int} = -ig_{aee}a(x)\bar{\psi}(x)\gamma_5\psi(x)$$

$$\left( \, ilde{g}_{aee}(\partial_{\mu}a) ar{\psi} \gamma^{\mu} \gamma_5 \psi(x) \, , \quad ilde{g}_{aee} = rac{g_{aee}}{2m_e} \, \, 
ight)$$

In the non-relativistic limit,  $\left(\mu_B=rac{|e|}{2m}
ight.$  : Bohr magneton ,  $\hat{S}^i=rac{\sigma^i}{2}$  : spin ight)

$$\mathcal{H}_{\mathrm{int}} \simeq -\frac{g_{aee}\hbar}{2m_e}\hat{\boldsymbol{\sigma}}\cdot\boldsymbol{\nabla}a = -2\mu_B\hat{\boldsymbol{S}}\cdot\left(\frac{g_{aee}}{e}\boldsymbol{\nabla}a\right)$$

effective magnetic field

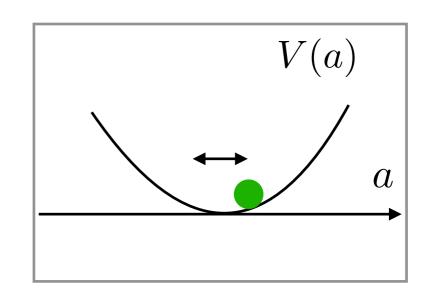


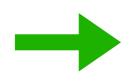
Reflecting the nature and distribution of the axion DM

### **Axion DM**

Axions can behave as DM if it oscillates around the bottom of the potential:

The equation of state parameter w=0





$$a(x) = a_0 \cos(\omega t - kx)$$

corresponding to the abundance of the axion DM

determined by the axion mass (  $\omega=m_a$  )

**※** 

In the case of the QCD axion, the axion mass around  $\,\mu eV\,$  is favored for DM.

$$m_a \sim \mu \text{eV}$$

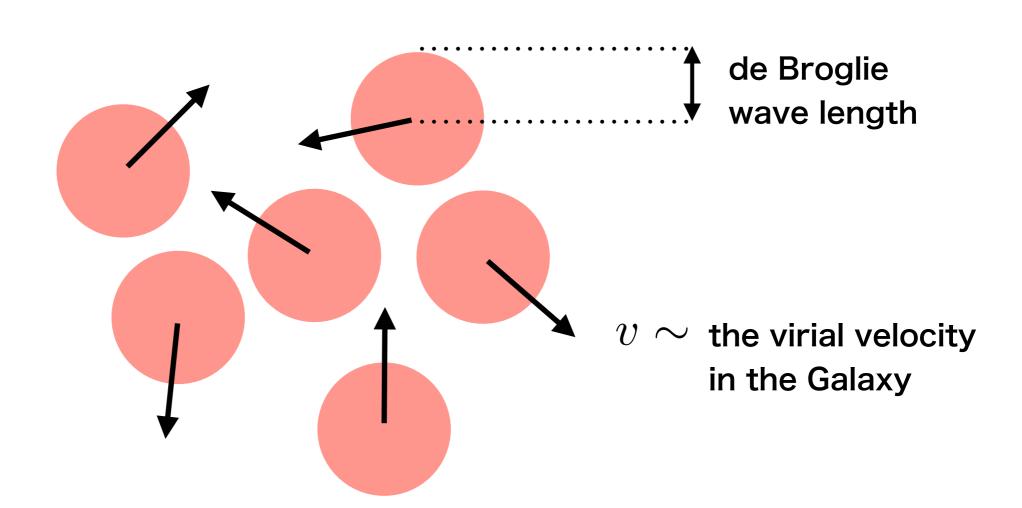
scale of cavity experiments (cm, GHz)

## Effective magnetic field

$$\boldsymbol{B}_a = \frac{g_{aee}}{e} \boldsymbol{\nabla} a$$

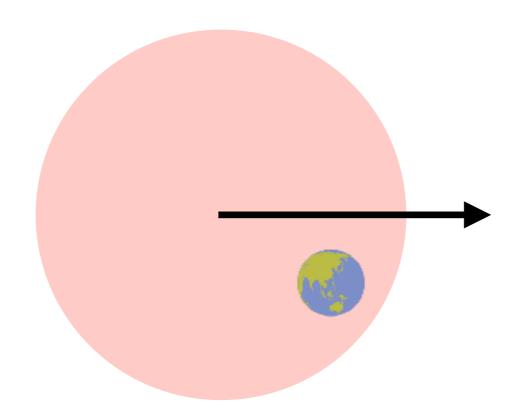
If gradient of the axion DM is non-zero, it acts as magnetic fields.

We assume that the axion DM forms clumps in the Galaxy as a stable solution of the Shrodinger-Poisson equation.



## Effective magnetic field

When a clump of axion DM is going through us, we feel "axion-wind"



Then  $\nabla a \sim m_a v a$  and the coherence time is

$$t_{\rm ob} \sim \frac{r_{\rm ob}}{v}$$
  
=  $2.3 \times 10^6 \times \left(\frac{1.0 \mu eV}{m_a}\right)^{1/2} \left(\frac{0.45 \text{ GeV/cm}^3}{\rho_{\rm ob}}\right)^{1/4} \left(\frac{300 \text{ km/s}}{v}\right) [\text{s}]$ 

## Effective magnetic field

We can estimate the amplitude of the effective magnetic field as

$$B_a \simeq 4.4 \times 10^{-8} \times g_{aee} \left(\frac{\rho_{\rm ob}}{0.45 \text{ GeV/cm}^3}\right)^{1/2} \left(\frac{v}{300 \text{ km/s}}\right) [\text{T}]$$

 $g_{aee}$  is tiny, how can we measure such small magnetic field?

- Axion-electron resonance caused by coherent oscillation of axion DM
   Use so many electrons (magnon)

1. Axion: axion DM, axion-electron interaction

2. Magnon as collective spin excitation of electrons

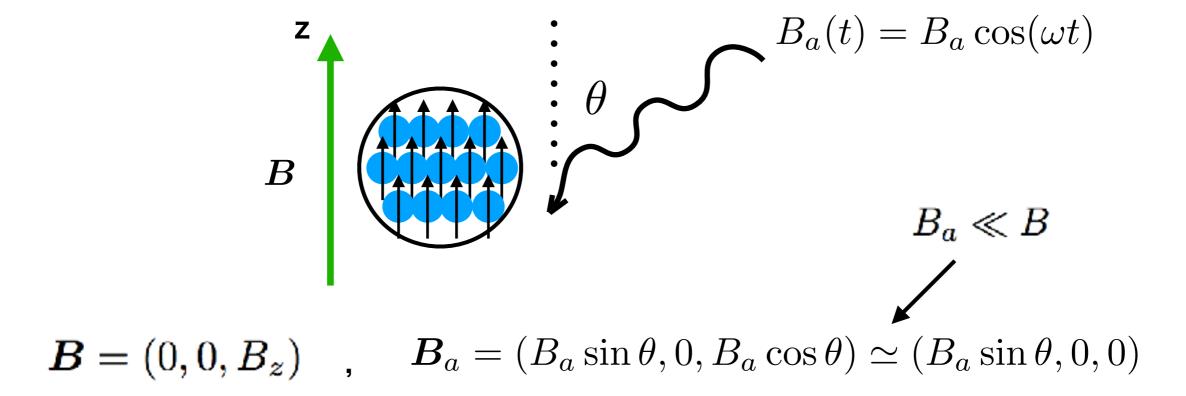
### **Magnon**

We consider a ferromagnetic sample which has N electronic spins in an external magnetic field  $\boldsymbol{B}$ .

It is well described by the Heisenberg model:

$$\mathcal{H} = -2\mu_B \sum_{i} \hat{m{S}}_i \cdot (m{B} + m{B}_a) - \sum_{i,j} J_{ij} \hat{m{S}}_i \cdot \hat{m{S}}_j$$

i=1...N specify the sites of electrons.



### Holstein-Primakoff transformation

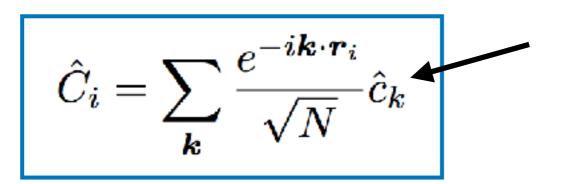
Spin operators can be rewritten in terms of bosonic operators by using the Holstein-Primakoff transformation:

$$\begin{cases} \hat{S}_{(i)}^z = \frac{1}{2} - \hat{C}_i^{\dagger} \hat{C}_i \;, \\ \hat{S}_{(i)}^+ = \sqrt{1 - \hat{C}_i^{\dagger} \hat{C}_i} \; \hat{C}_i \;, & \text{where} \quad [\hat{C}_i, \hat{C}_j^{\dagger}] = \delta_{ij} \\ \hat{S}_{(i)}^- = \hat{C}_i^{\dagger} \sqrt{1 - \hat{C}_i^{\dagger} \hat{C}_i} \;, & \end{cases}$$

A spin flip corresponds to creation of a boson.

$$B_z \qquad = \text{ no boson }, \qquad B_z \qquad = \qquad \text{a boson}$$

Furthermore, we can move on to Fourier space:



A bosonic operator of "spin wave"

II Magnon

## **Magnon**

$$\mathcal{H} = -2\mu_B \sum_{i} \hat{m{S}}_i \cdot (m{B} + m{B}_a) - \sum_{i,j} J_{ij} \hat{m{S}}_i \cdot \hat{m{S}}_j$$

Holstein-Primakoff transformation

$$\mathcal{H} \simeq 2\mu_B B_z \hat{c}_{k=0}^{\dagger} \hat{c}_{k=0} + 2\mu_B \frac{B_a \sin \theta}{4} \sqrt{N} \left( \hat{c}_{k=0}^{\dagger} e^{-i\omega_a t} + \hat{c}_{k=0} e^{i\omega_a t} \right)$$
$$+ \sum_{i=1..N} \mathcal{H}(\hat{c}_{k=i})$$

The coupling constant is effectively increased by  $\sqrt{N}$  . Typically,  $\sqrt{N}\sim \sqrt{10^{20}}\sim 10^{10}$  .

The axion DM excites the uniform mode of the magnon!

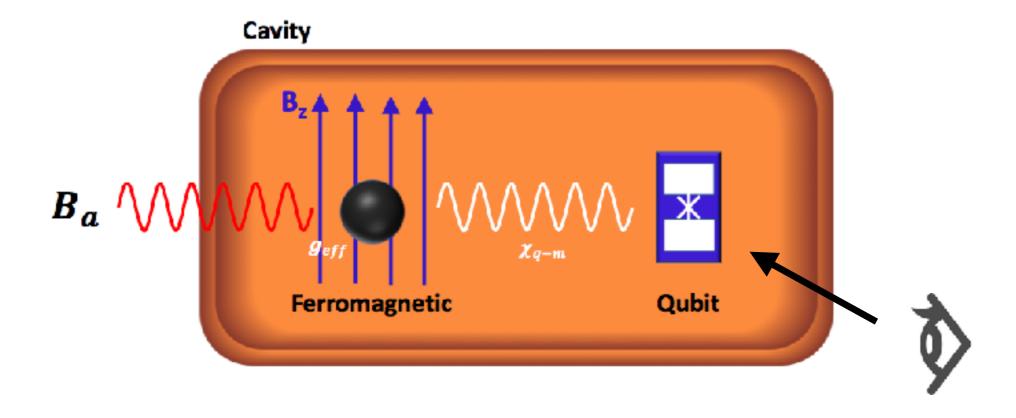
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# **Experiment**

We measured the quantum state of a magnon with qubit

(qubit: A two-state system)



We can see the quantum state of the magnon by observing qubit spectrum!

# Qubit spectrum

$$\mathcal{H}_{\text{tot}} = \hbar \omega_{m} \hat{c}^{\dagger} \hat{c} + \frac{\hbar \omega_{q}}{2} \hat{\sigma}_{z} + \hbar g_{q-m} (\hat{c}^{\dagger} \hat{\sigma}_{-} + \hat{\sigma}_{+} \hat{c})$$

$$+ g_{eff} \left( \hat{c}^{\dagger} e^{-i\omega_{a}t} + \hat{c} e^{i\omega_{a}t} \right)$$

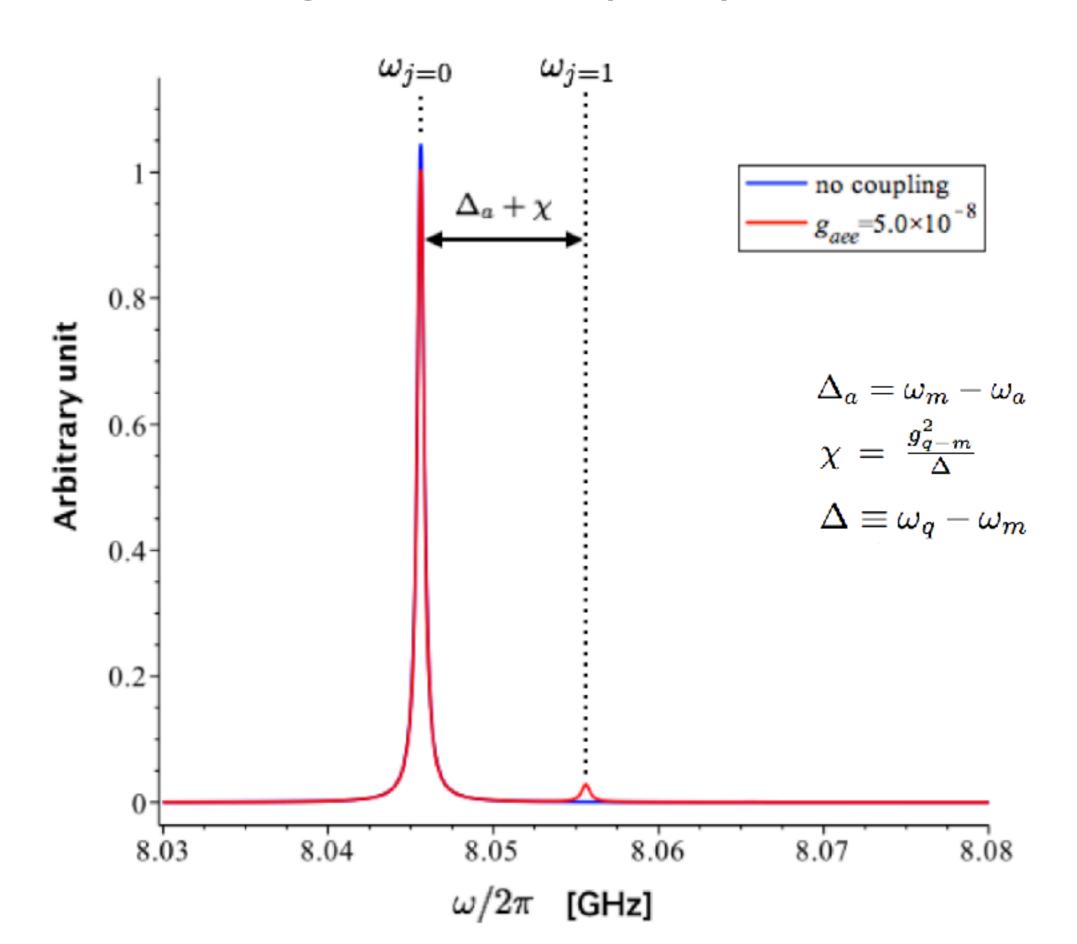
$$+ \mathcal{H}_{\text{noise}}$$

 $\begin{pmatrix} \omega_m : \text{magnon frequency} \\ \omega_q : \text{qubit frequency} \\ g_{q-m} : \text{magnon-qubit coupling constant} \\ g_{eff} : \text{magnon-axion coupling constant} \\ \end{pmatrix}$ 

This system is approximately solvable and we can obtain the qubit spectrum:

$$S(\omega) = \operatorname{Re}\left[\frac{1}{\sqrt{2\pi}} \int_0^\infty \mathrm{dt} < \hat{\sigma}_-(t)\hat{\sigma}_+(0) > \mathrm{e}^{\mathrm{i}\omega t}\right]$$

## Magnon state and qubit spectrum



## **Upper limit**

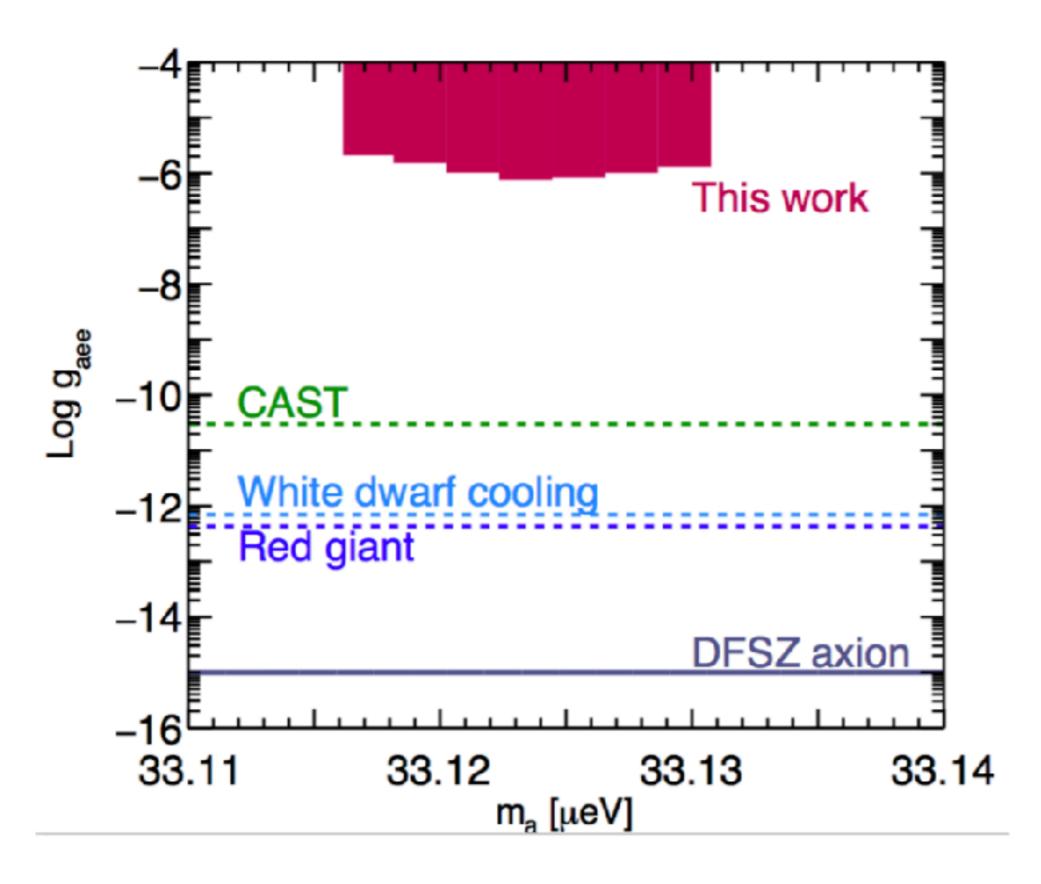
We observed the qubit spectrum and found no evidence of the axion DM.



$$B_a < 4.1 \times 10^{-14} [T]$$
 or  $g_{aee} < 1.3 \times 10^{-6}$ 

at 
$$m_a=33~\mu {\rm eV}$$

## **Upper limit**



# <u>Summary</u>

Axion is a strong candidate for DM

Axion DM can induce resonant spin precession of electrons

Axion-magnon coupling get effective factor  $\sqrt{N} \sim \sqrt{10^{20}} \sim 10^{10}$ 

We measured the quantum state of a magnon and gave an upper limit  $g_{aee} < 1.3 \times 10^{-6} \quad {\rm at} \quad m_a = 33 \quad \mu {\rm eV}$