

15. May. 2019 @ Axion Cosmology in Kyoto

PBH tower in multi-phase inflation

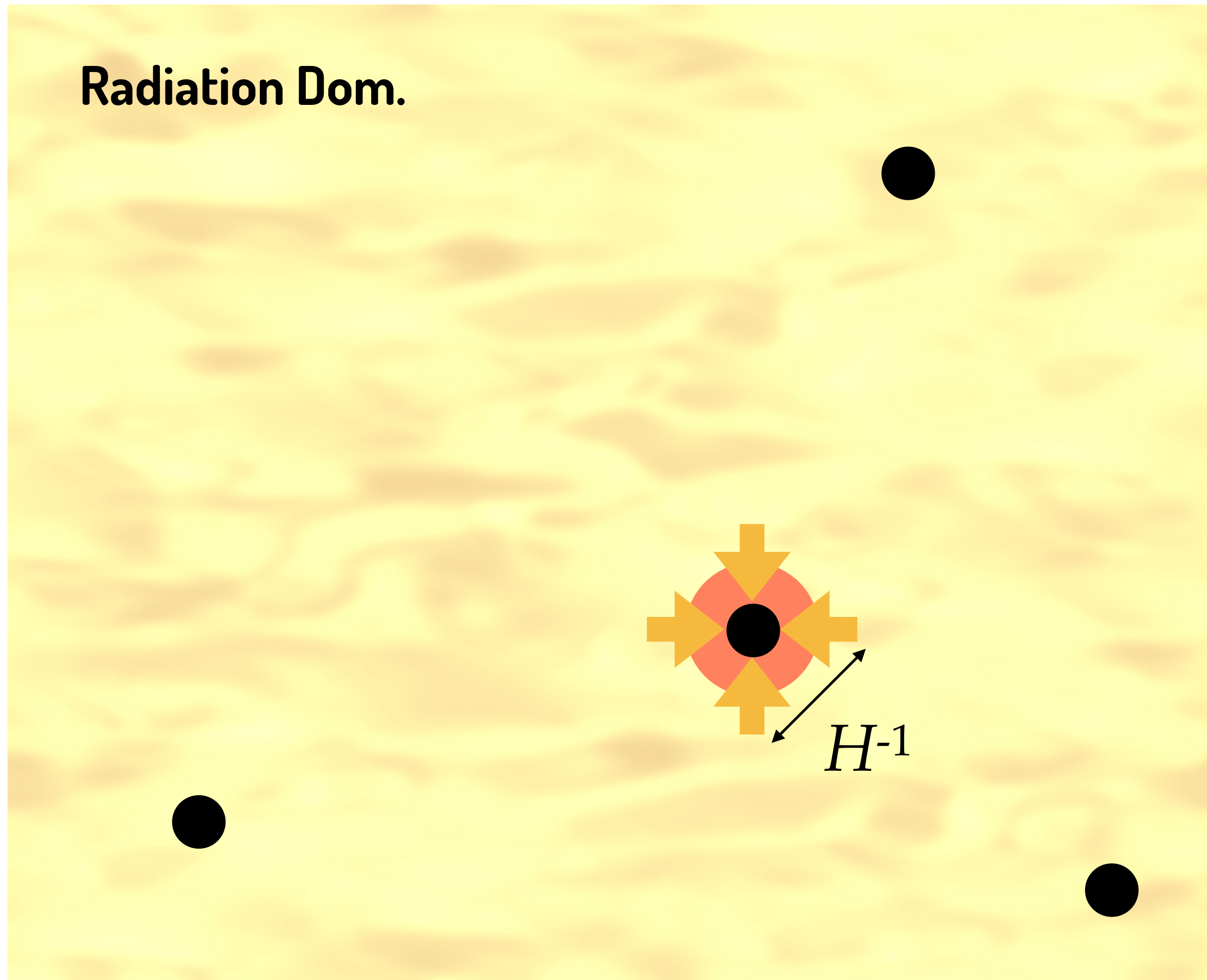


Yuichiro TADA (Nagoya U.)

w/ S. Yokoyama 1904.10298

Primordial Black Hole

Carr & Hawking 1974



- Overdensity

$$\delta > \delta_{\text{th}} \simeq 0.4$$

Musco, Miller, Rezzolla 2005, ...
Harada, Yoo, Kohri 2013

$$\left(\text{cf. } \mathcal{R}_{\text{th}} \simeq \frac{9}{4} \delta_{\text{th}} \simeq 1 \right)$$

- Rarity

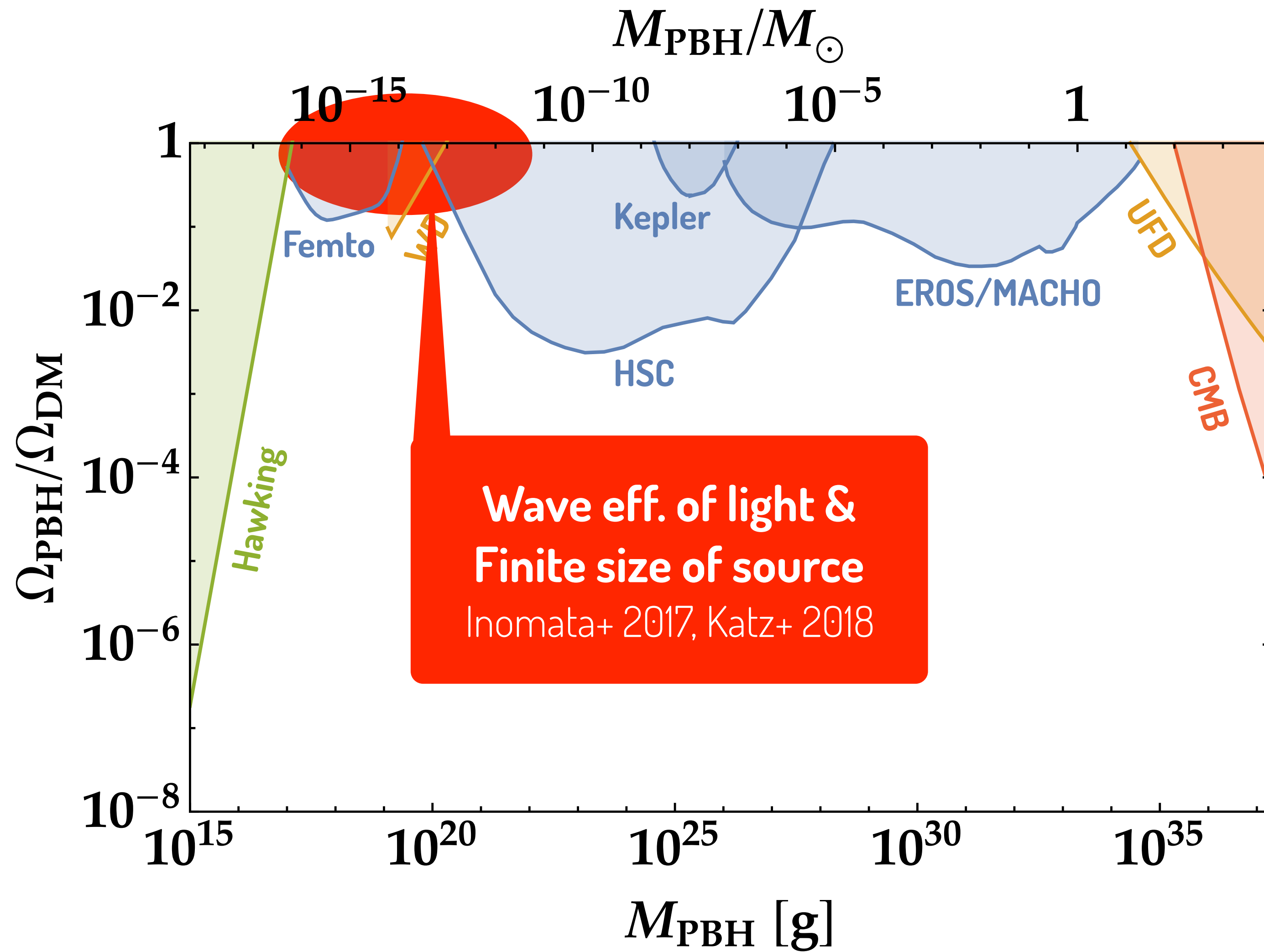
$$\frac{\Omega_{\text{PBH}}}{\Omega_{\text{DM}}} \sim \frac{\rho_{\text{PBH}}/\rho_{\text{R}}|_f}{7 \times 10^{-16}} \left(\frac{M_{\text{PBH}}}{10^{20} \text{ g}} \right)^{-1/2}$$

~ 10σ rarity

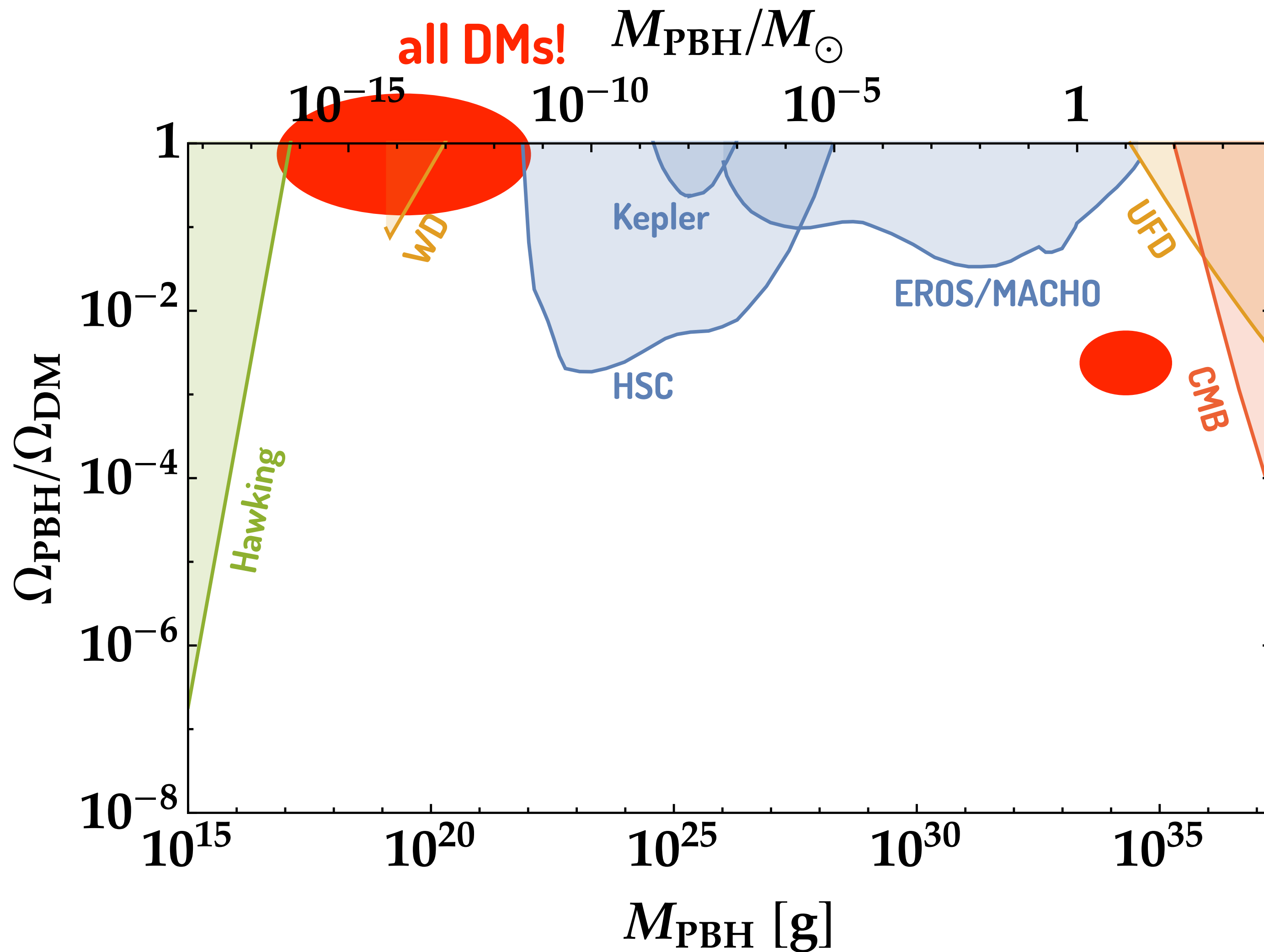
➔ $\mathcal{P}_{\mathcal{R}}(k_{\text{PBH}}) \sim \left(\frac{\mathcal{R}_{\text{th}}}{10} \right)^2 \simeq 10^{-2}$

(cf. $\mathcal{P}_{\mathcal{R}}(k_{\text{CMB}}) \simeq 2 \times 10^{-9}$)

Obs. const. on PBH



Obs. const. on PBH



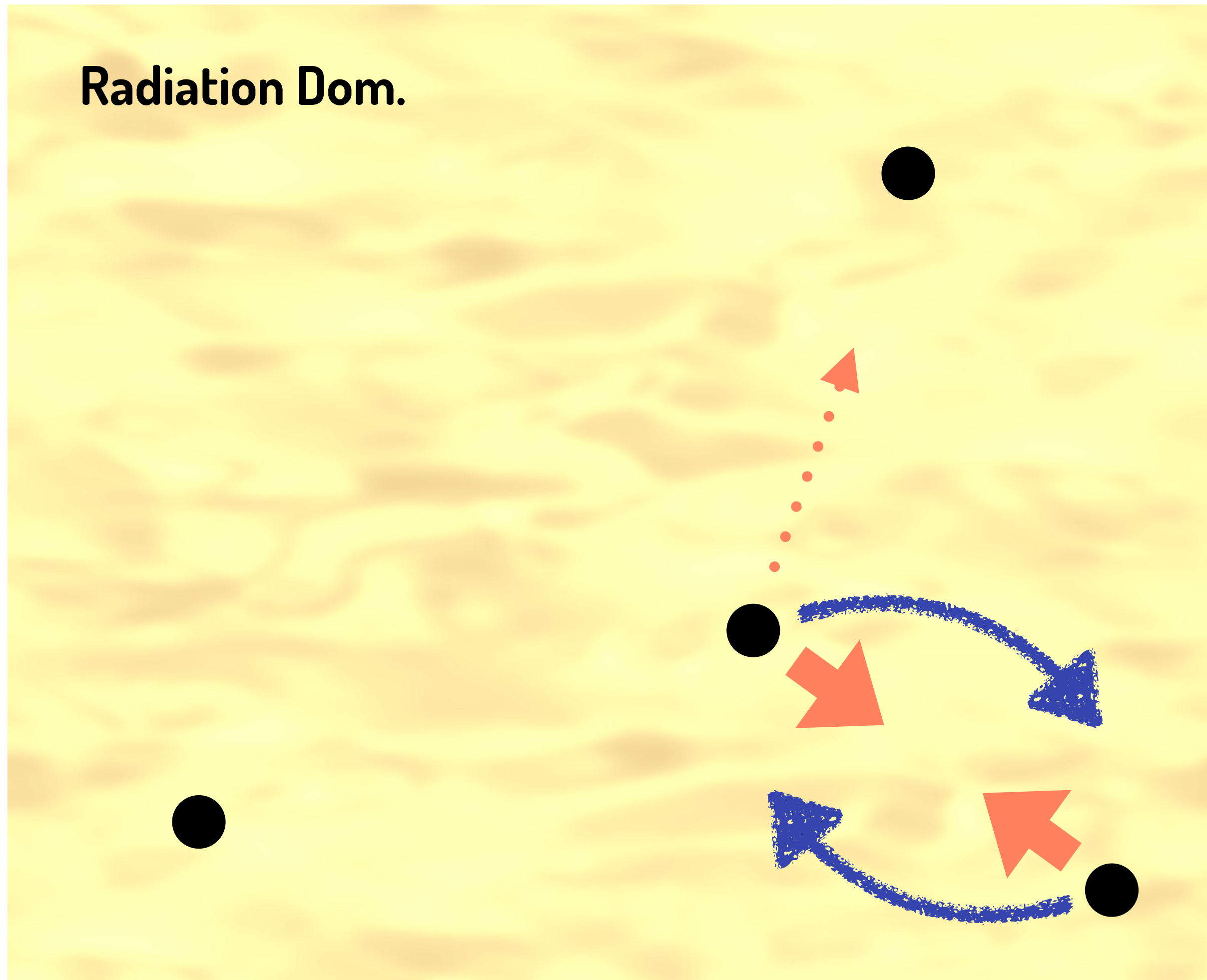
Massive than stellar BHs found

small spin

Event	m_1/M_{\odot}	m_2/M_{\odot}	M/M_{\odot}	χ_{eff}
GW150914	$35.6^{+4.8}_{-3.0}$	$30.6^{+3.0}_{-4.4}$	$28.6^{+1.6}_{-1.5}$	$-0.01^{+0.12}_{-0.13}$
GW151012	$23.3^{+14.0}_{-5.5}$	$13.6^{+4.1}_{-4.8}$	$15.2^{+2.0}_{-1.1}$	$0.04^{+0.28}_{-0.19}$
GW151226	$13.7^{+8.8}_{-3.2}$	$7.7^{+2.2}_{-2.6}$	$8.9^{+0.3}_{-0.3}$	$0.18^{+0.20}_{-0.12}$
GW170104	$31.0^{+7.2}_{-5.6}$	$20.1^{+4.9}_{-4.5}$	$21.5^{+2.1}_{-1.7}$	$-0.04^{+0.17}_{-0.20}$
GW170608	$10.9^{+5.3}_{-1.7}$	$7.6^{+1.3}_{-2.1}$	$7.9^{+0.2}_{-0.2}$	$0.03^{+0.19}_{-0.07}$
GW170729	$50.6^{+16.6}_{-10.2}$	$34.3^{+9.1}_{-10.1}$	$35.7^{+6.5}_{-4.7}$	$0.36^{+0.21}_{-0.25}$
GW170809	$35.2^{+8.3}_{-6.0}$	$23.8^{+5.2}_{-5.1}$	$25.0^{+2.1}_{-1.6}$	$0.07^{+0.16}_{-0.16}$
GW170814	$30.7^{+5.7}_{-3.0}$	$25.3^{+2.9}_{-4.1}$	$24.2^{+1.4}_{-1.1}$	$0.07^{+0.12}_{-0.11}$
GW170817	$1.46^{+0.12}_{-0.10}$	$1.27^{+0.09}_{-0.09}$	$1.186^{+0.001}_{-0.001}$	$0.00^{+0.02}_{-0.01}$
GW170818	$35.5^{+7.5}_{-4.7}$	$26.8^{+4.3}_{-5.2}$	$26.7^{+2.1}_{-1.7}$	$-0.09^{+0.18}_{-0.21}$
GW170823	$39.6^{+10.0}_{-6.6}$	$29.4^{+6.3}_{-7.1}$	$29.3^{+4.2}_{-3.2}$	$0.08^{+0.20}_{-0.22}$

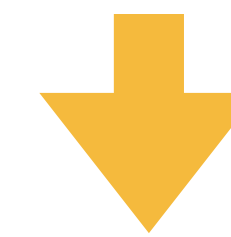
LIGO/Virgo 2018

Binary PBH



Sasaki+ 2016

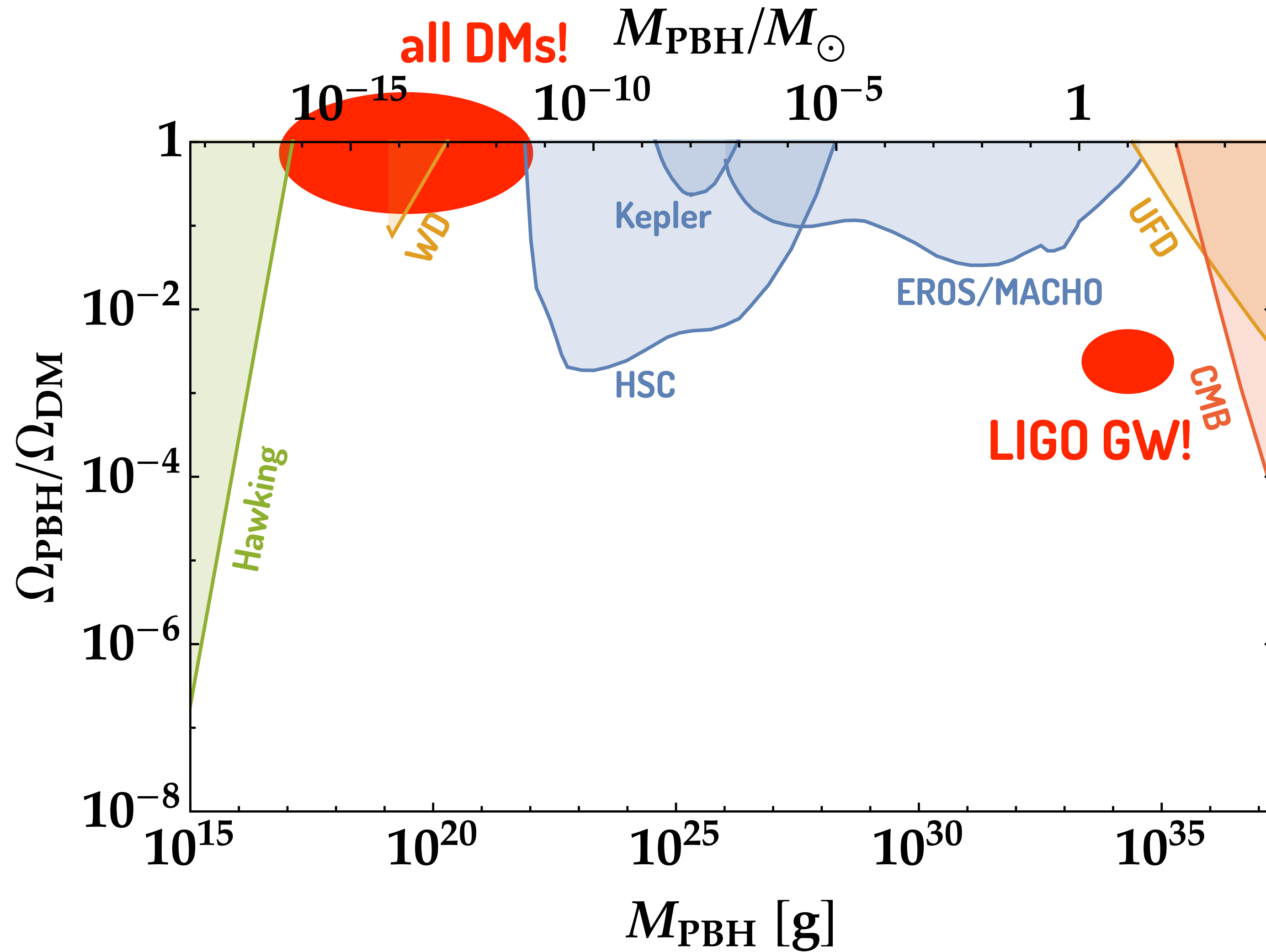
$$\text{if } \frac{\Omega_{\text{PBH}}}{\Omega_{\text{DM}}} \sim 10^{-3} - 10^{-2}$$



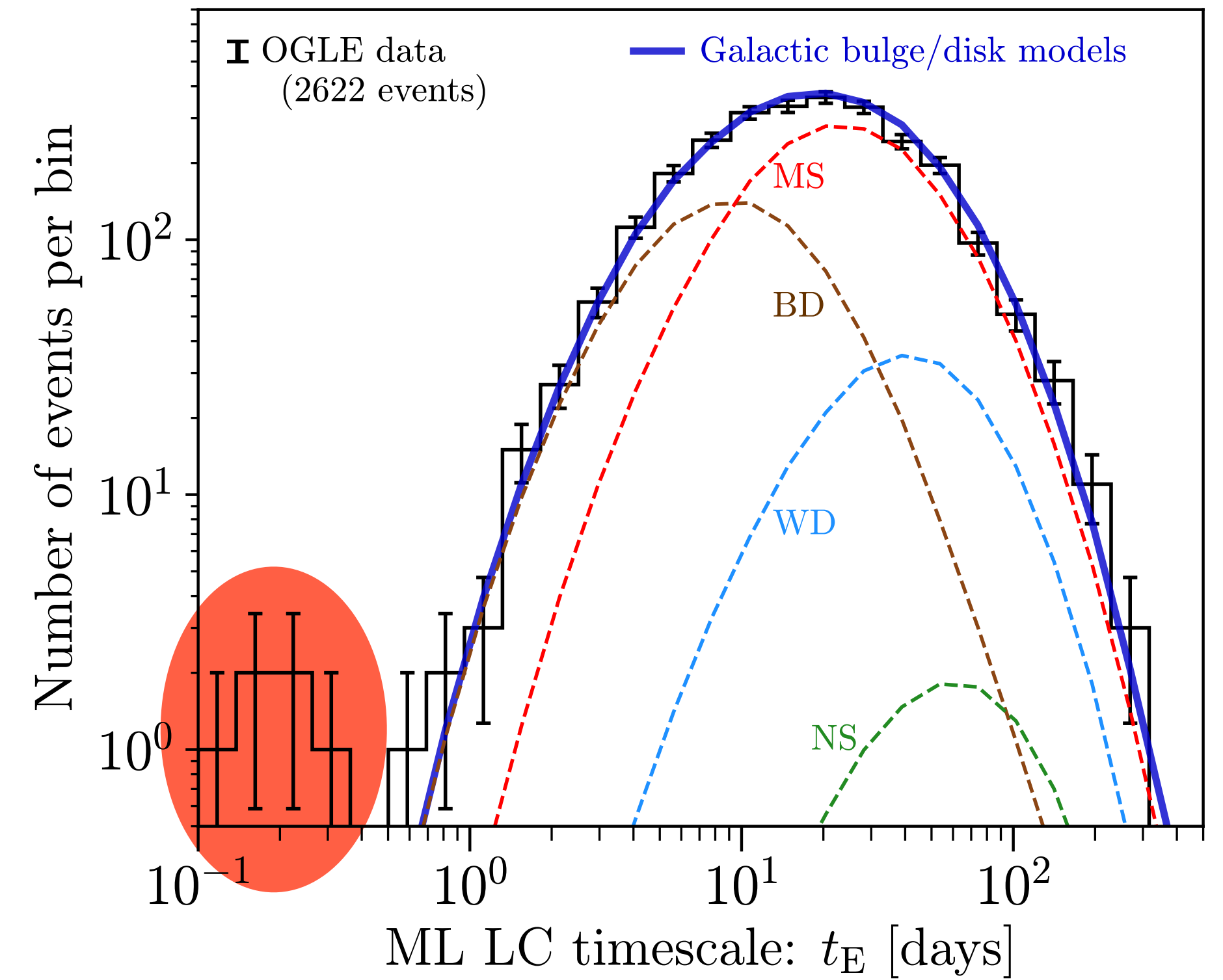
$$[\text{merger rate}] \simeq 52.9^{+55.6}_{-27.0} \text{ Gpc}^{-3} \text{ yr}^{-1}$$

LIGO/Virgo 2018

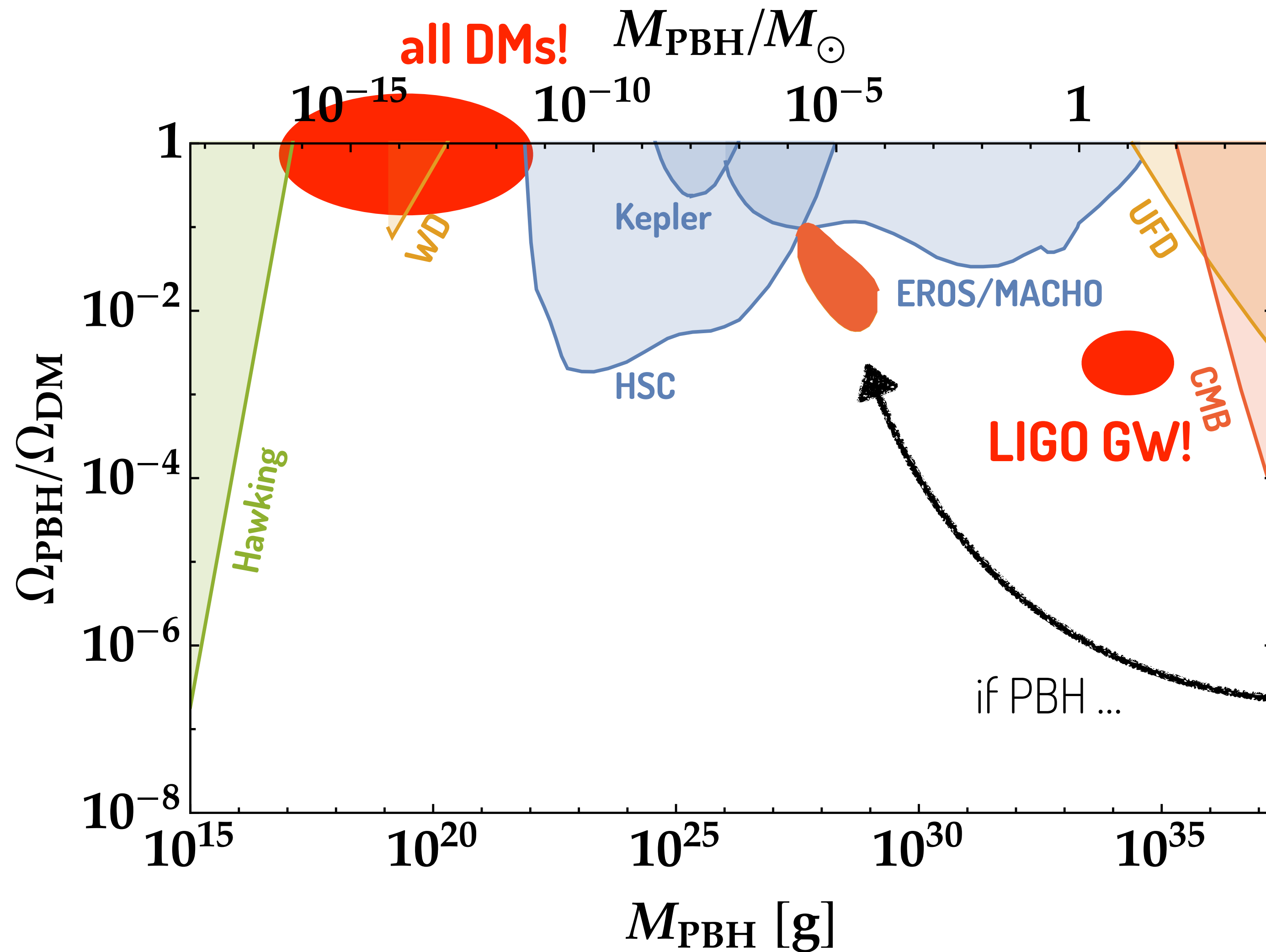
Obs. const. on PBH



5-yr Optical Gravitational Lensing Experiment (OGLE)
Niikura+ 2019

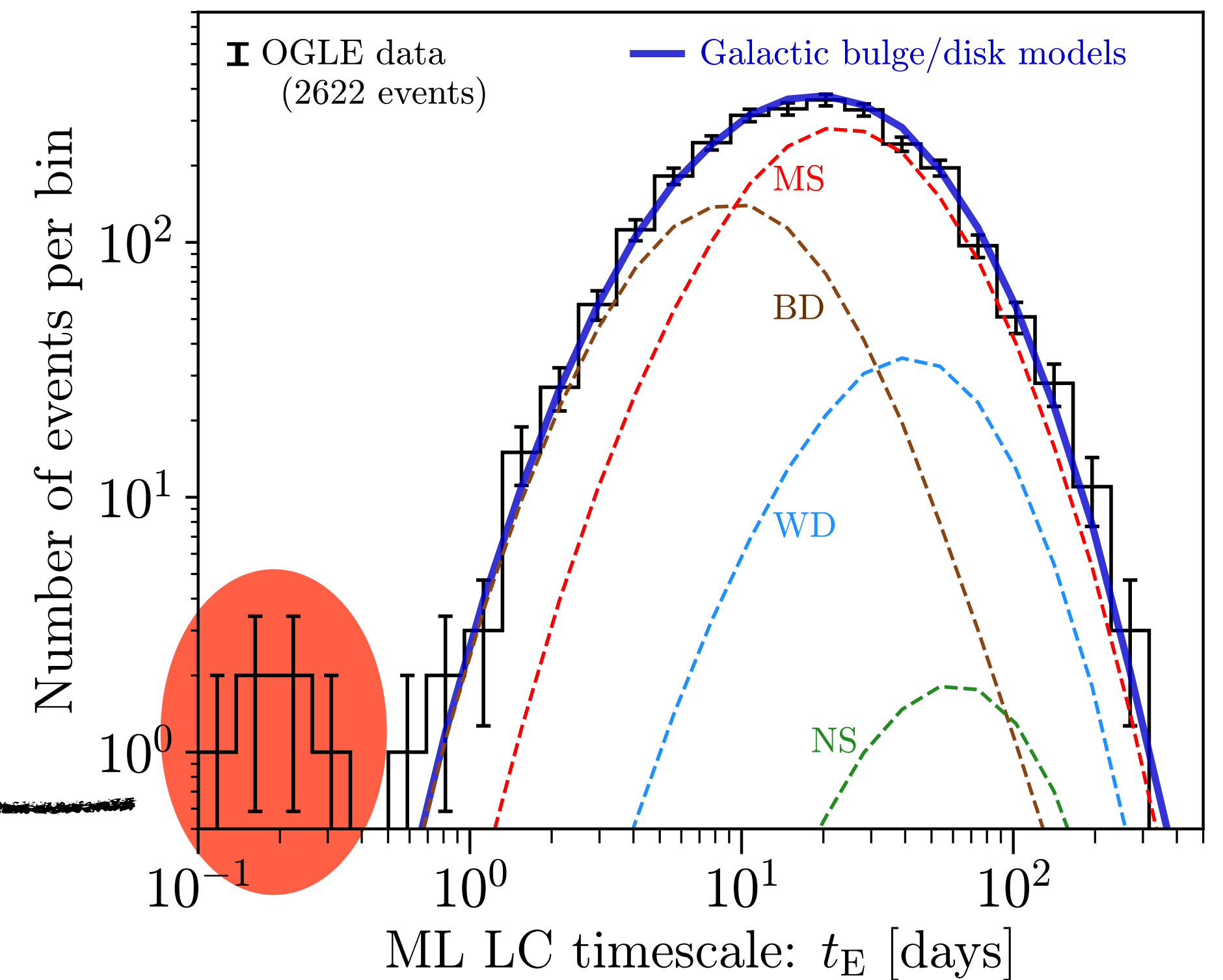


Obs. const. on PBH

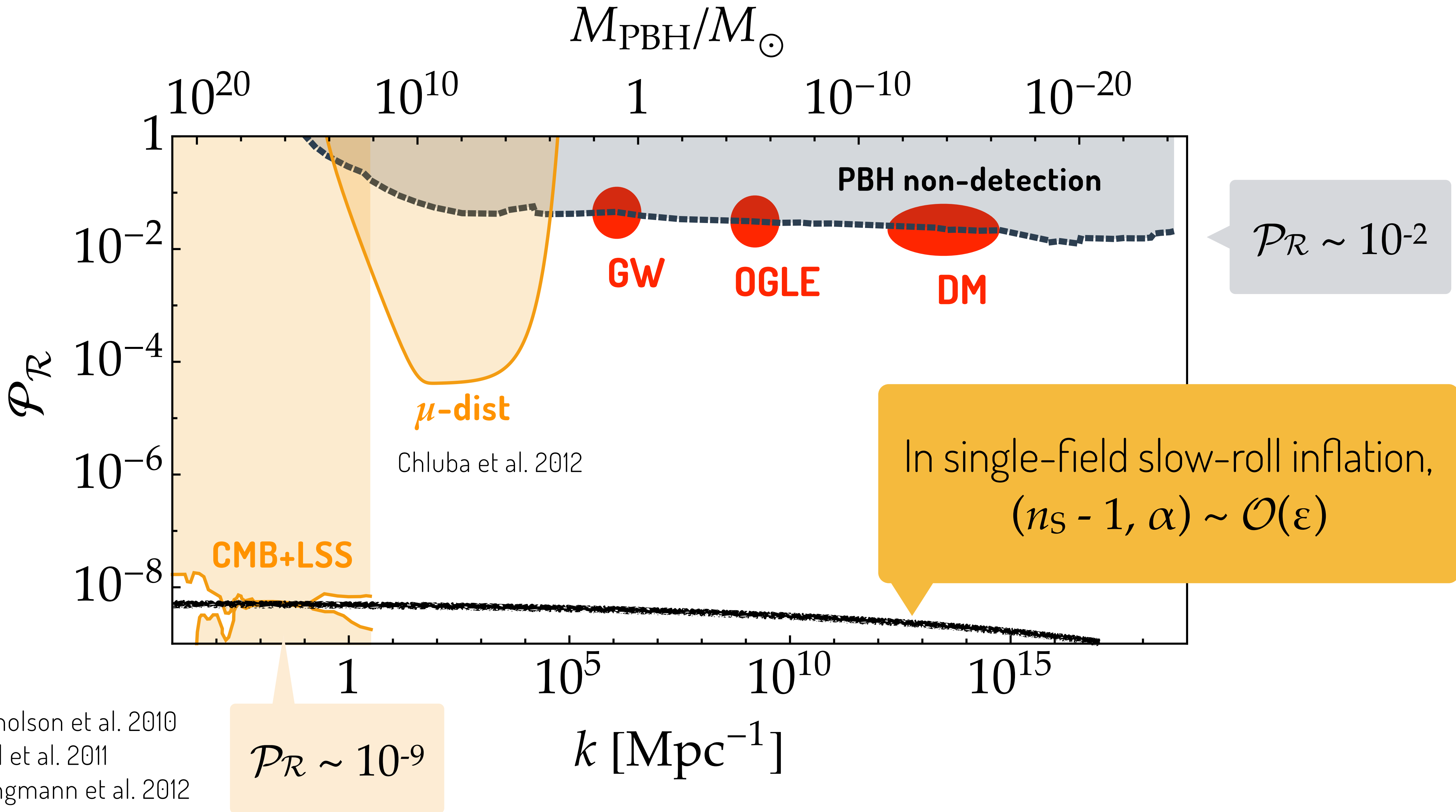


5-yr Optical Gravitational Lensing Experiment (OGLE)

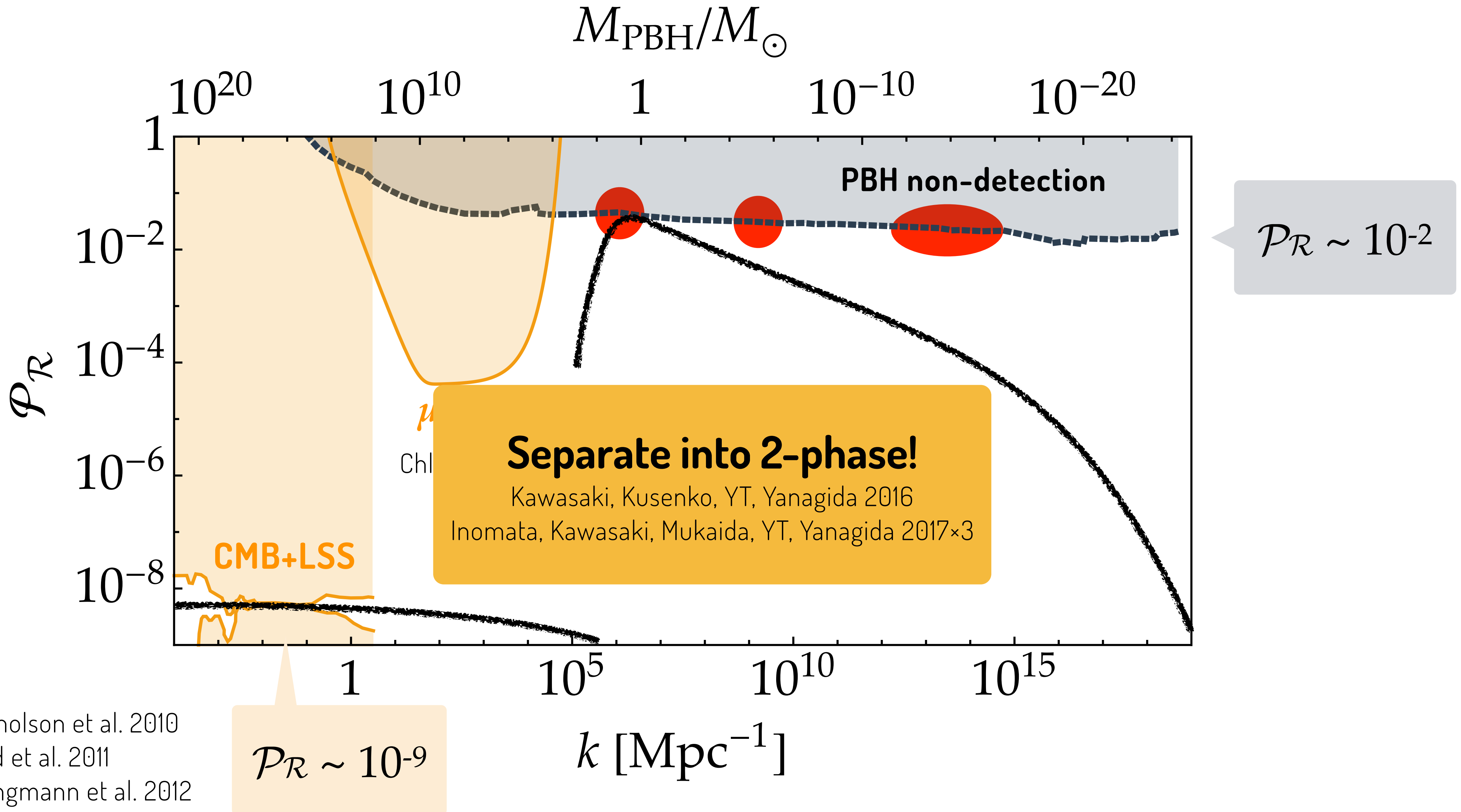
Niikura+ 2019



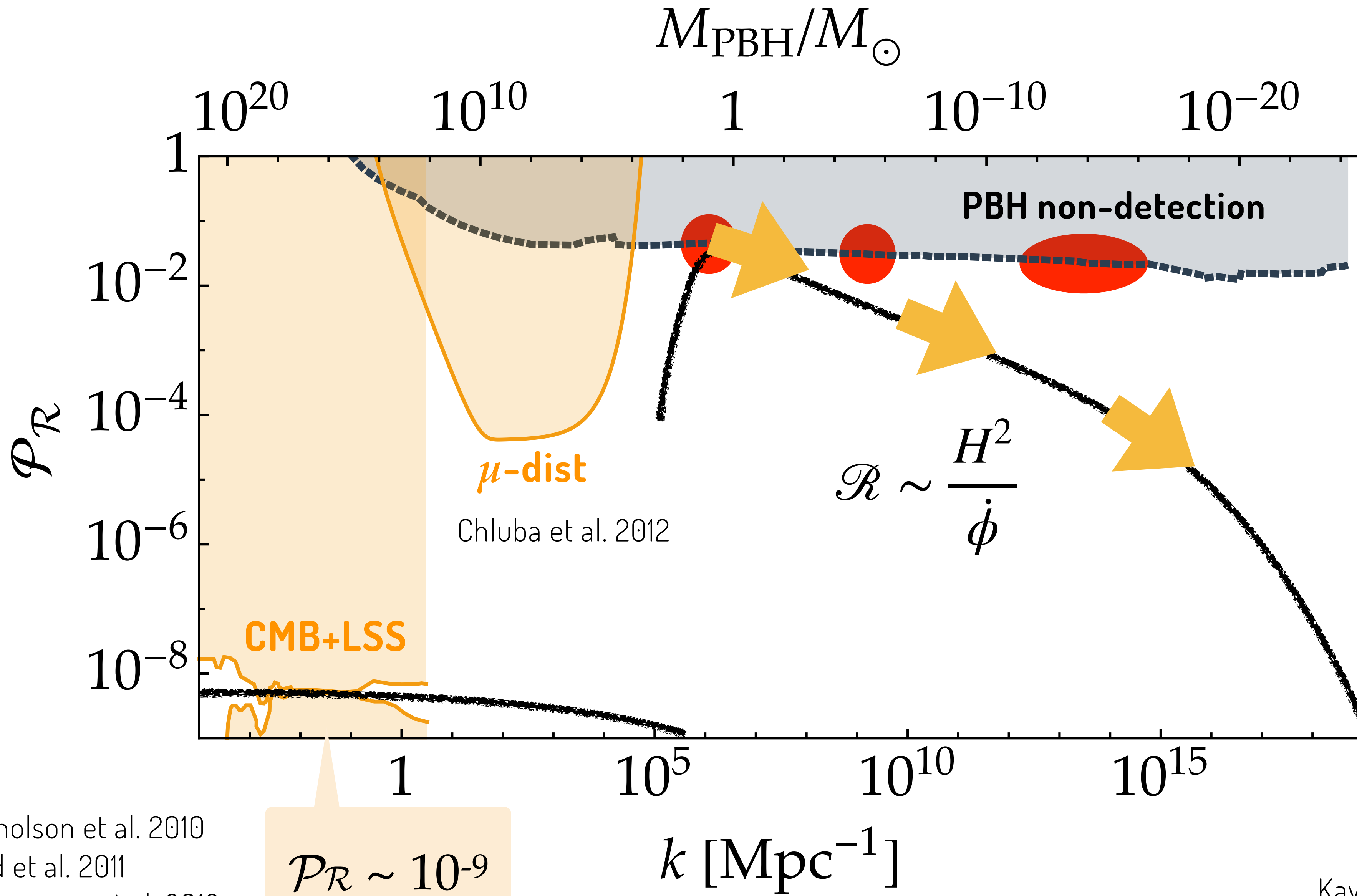
Inflationary sector ...



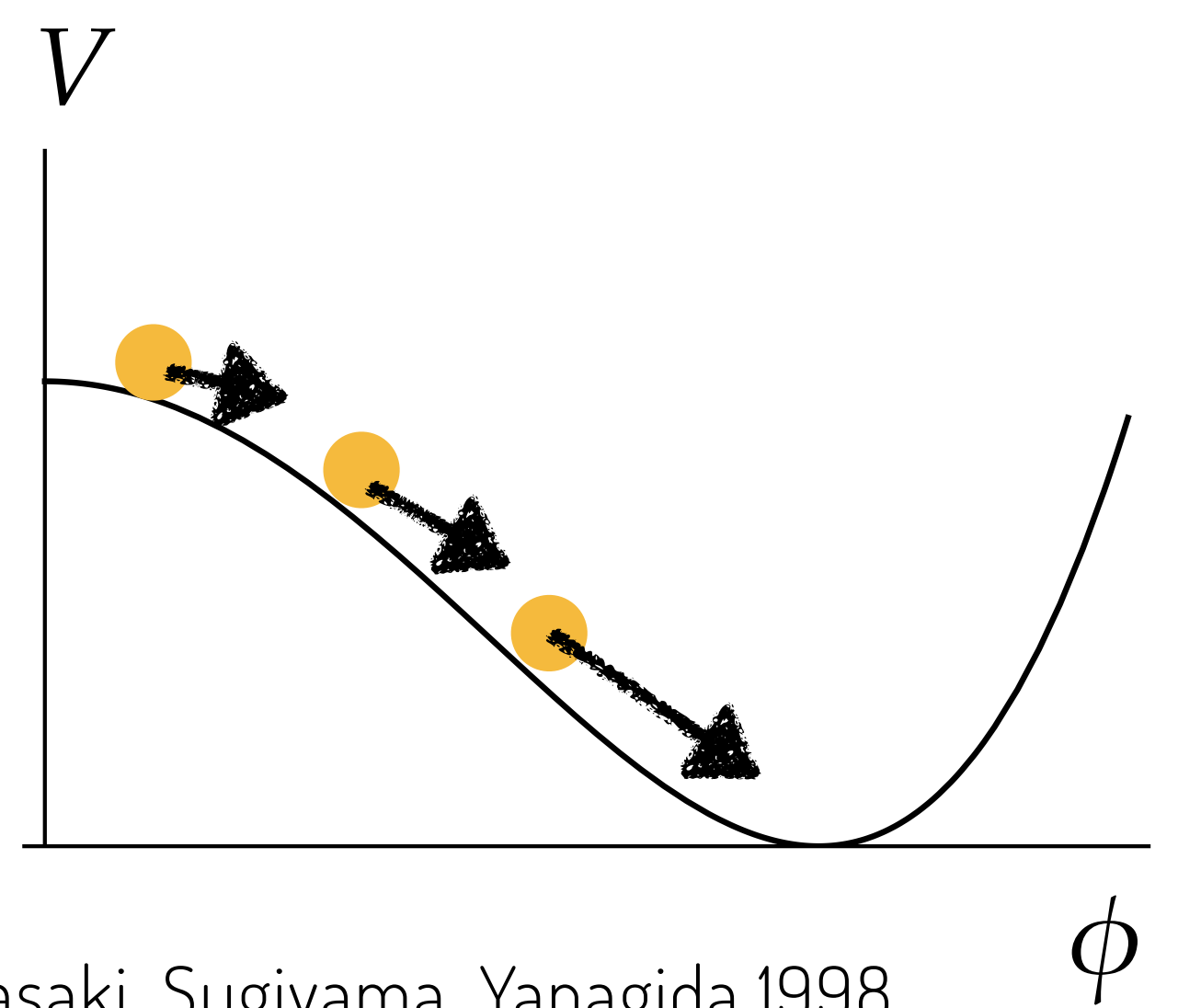
Nicholson et al. 2010
 Bird et al. 2011
 Bringmann et al. 2012



Nicholson et al. 2010
Bird et al. 2011
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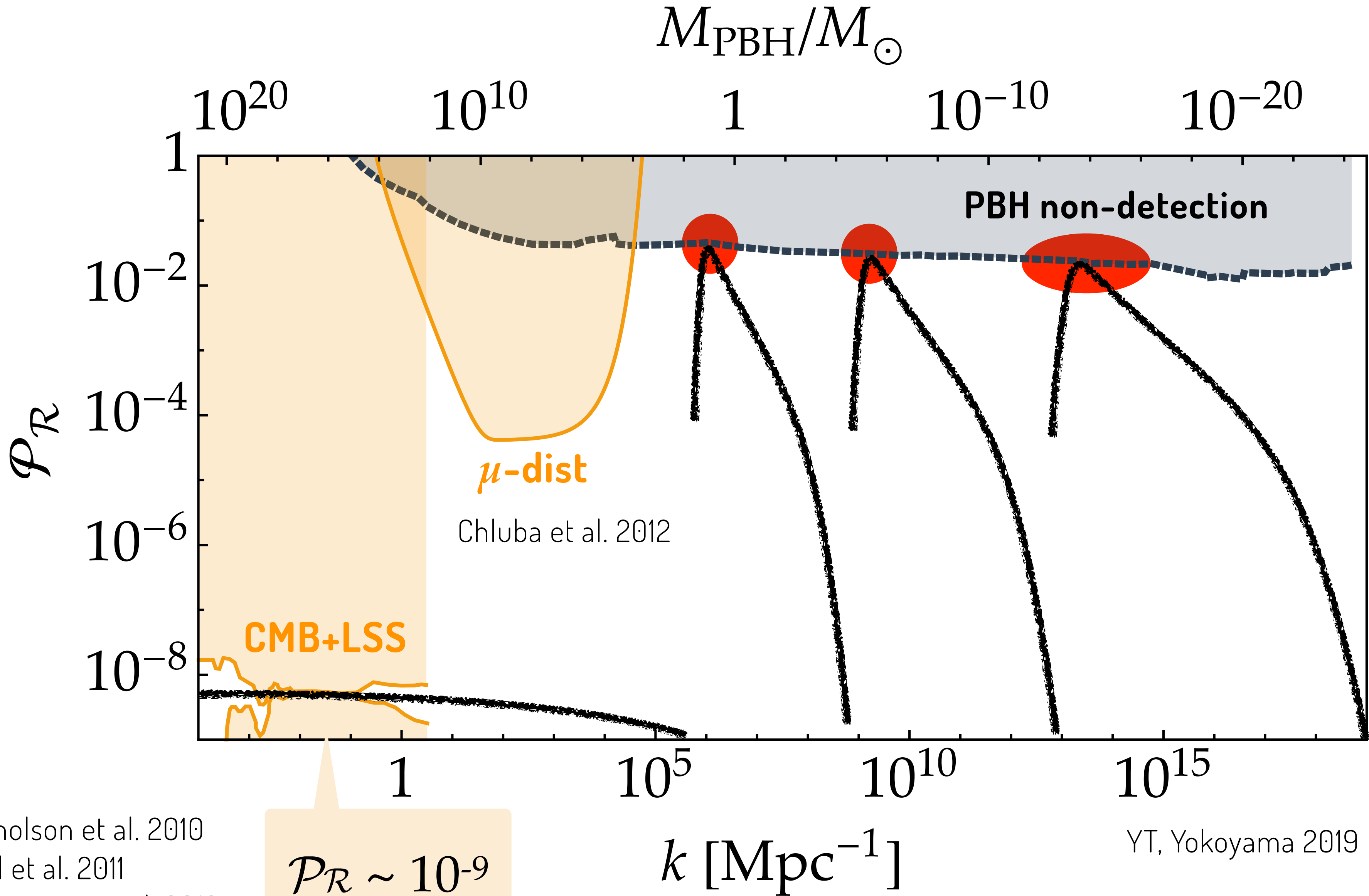


$\mathcal{P}_{\mathcal{R}} \sim 10^{-2}$



Kawasaki, Sugiyama, Yanagida 1998

Nicholson et al. 2010
 Bird et al. 2011
 Bringmann et al. 2012



- 4-hilltop

$$V_{\text{hill}} = \sum_{i=1}^4 V_{\text{hill},i}$$

$$\simeq \sum_i v_i^4 \left(1 - \frac{1}{2} \kappa_i \frac{\phi_i^2}{M_{\text{Pl}}^2} \right)$$

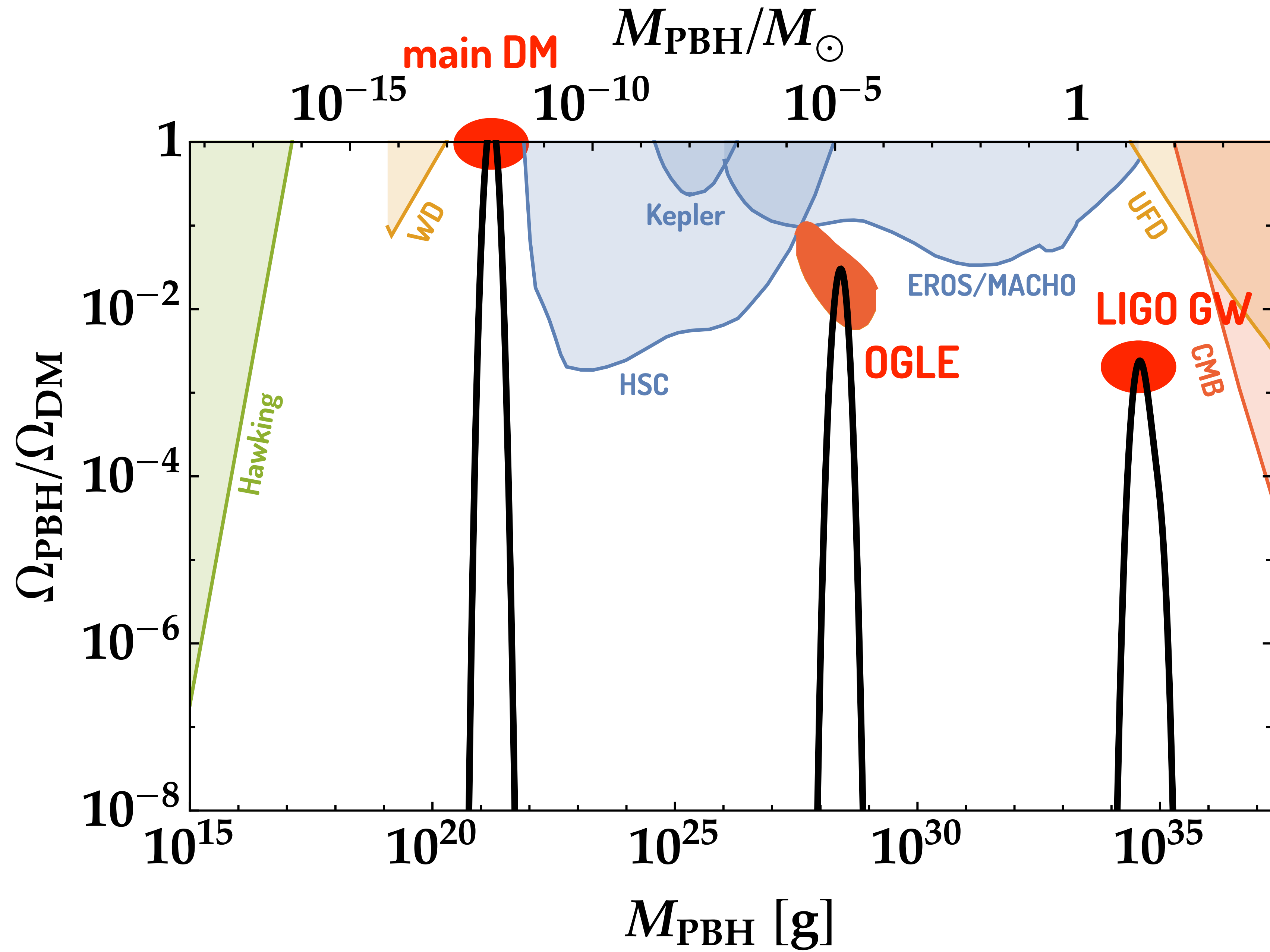
+

- Stabilization

$$V_{\text{stab}} = \sum_{i \neq j} \frac{c_{ij}}{2} V_{\text{hill},i} \frac{\phi_j^2}{M_{\text{Pl}}^2}$$

- during phase- i :
stabilize ϕ_{i+1}
- after $V_{\text{hill},i}$ decays:
start phase- $(i+1)$

Nicholson et al. 2010
Bird et al. 2011
Bringmann et al. 2012



Testability

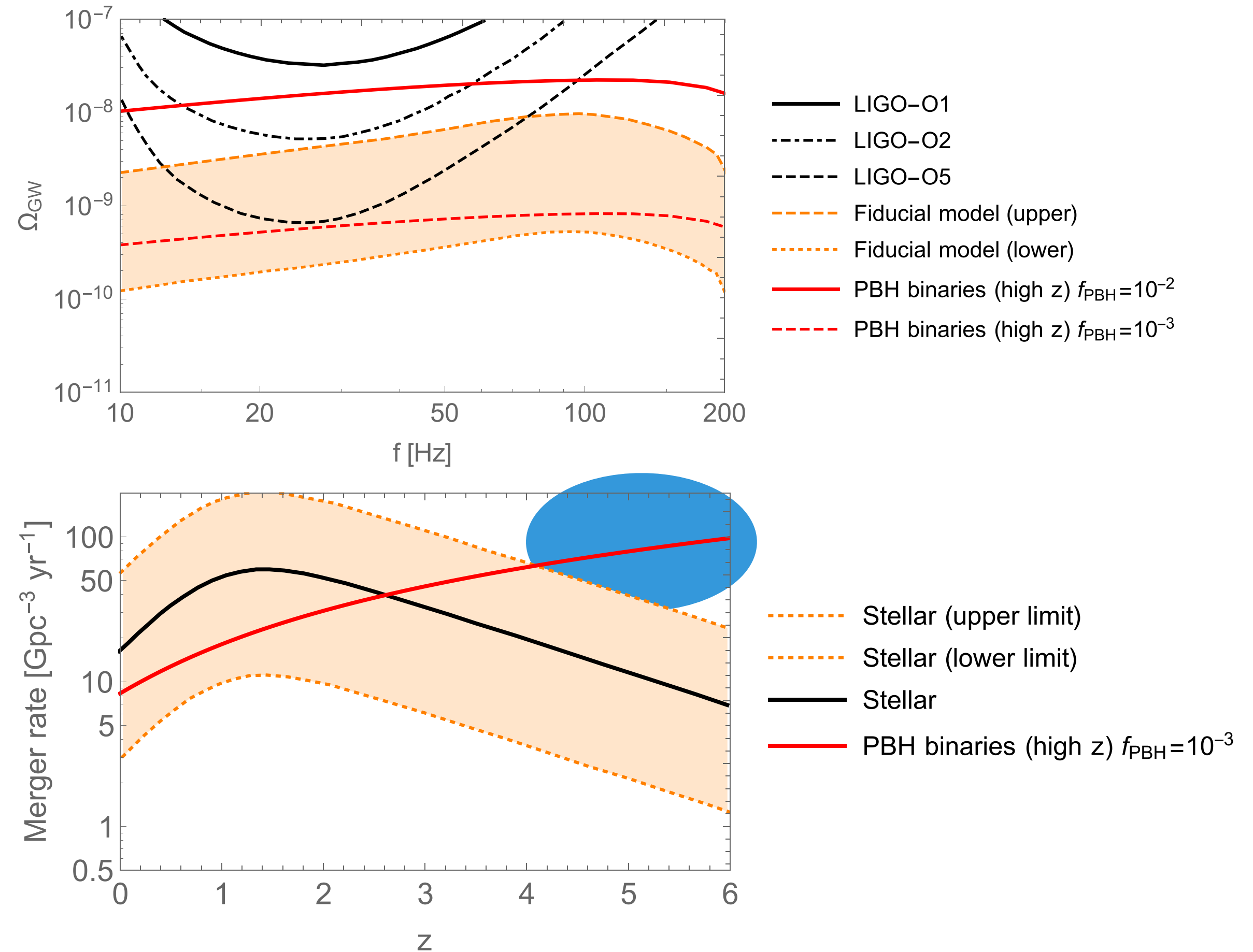
- LIGO/Virgo PBH

PBH tends to be spinless
Chiba & Yokoyama 2017

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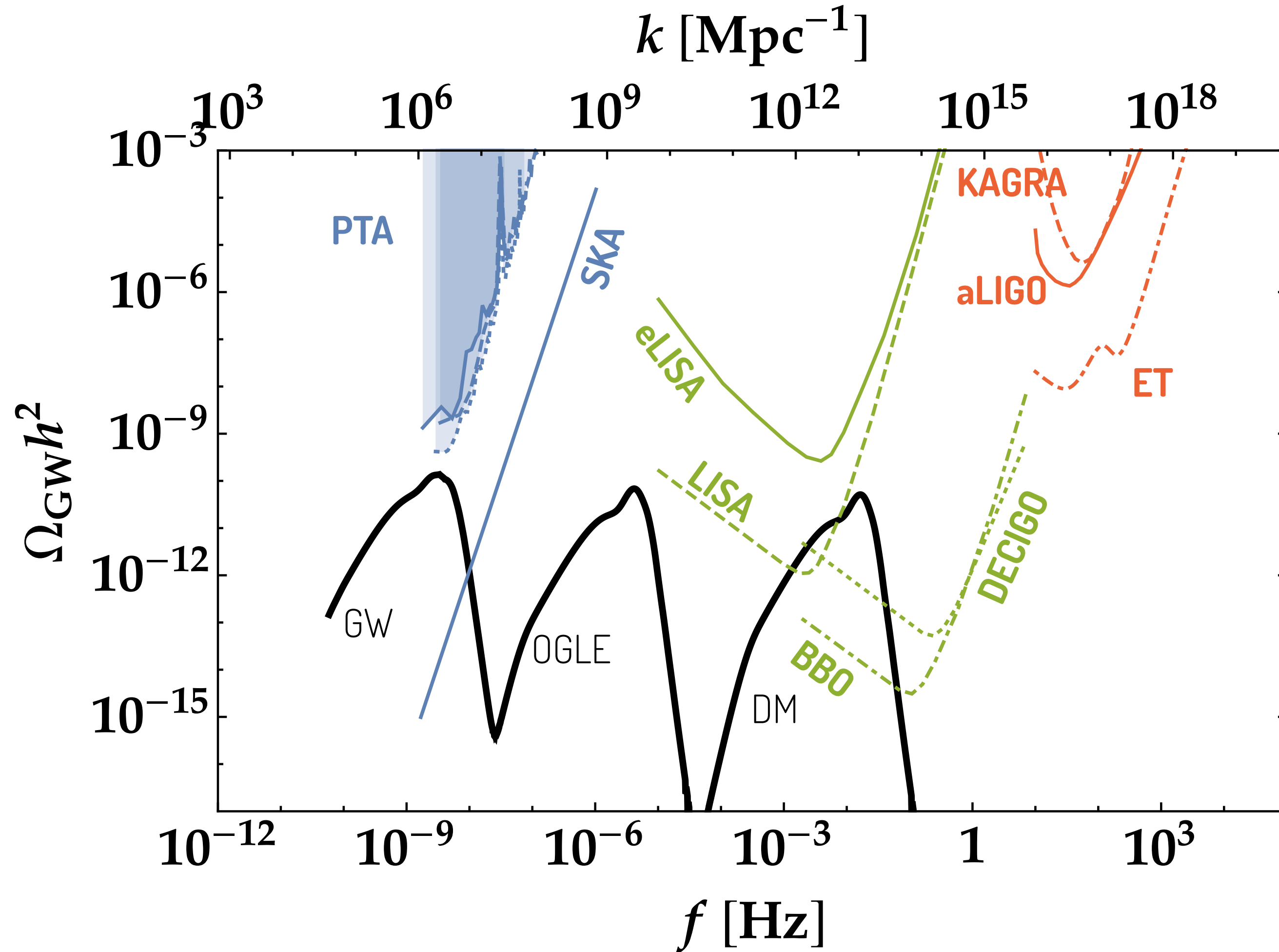
LIGO/Virgo 2018

Sasaki+ 2018



Testability

large scalar ptb. \rightarrow secondary tensor ptb. (stochastic GW): $\Omega_{\text{GW}} h^2 \sim 10^{-9} \left(\frac{\mathcal{P}_{\mathcal{R}}}{10^{-2}} \right)^2$



Conclusions

- 3 interesting mass region for PBH
 - $10 M_{\odot}$: LIGO/Virgo GW
 - $10^{-5} M_{\odot}$: OGLE lensing
 - $10^{-12} M_{\odot}$: main component of DMs
- multi-phase inflation can realize them simultaneously
 - cf. string swampland conjecture Ooguri & Vafa+ 2018
 - “dS vacua will be unstable in UV-complete theories”*
- testable by GW