

# Gravitational Self-Force under Regge-Wheeler Gauge

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# §1. Introduction

Space-based gravitational wave observatory

Super massive black hole – Compact object binary

## Black hole perturbation

Background: Black hole (mass:  $M$ )

+ Perturbation: **Point particle** ( $\mu$ )

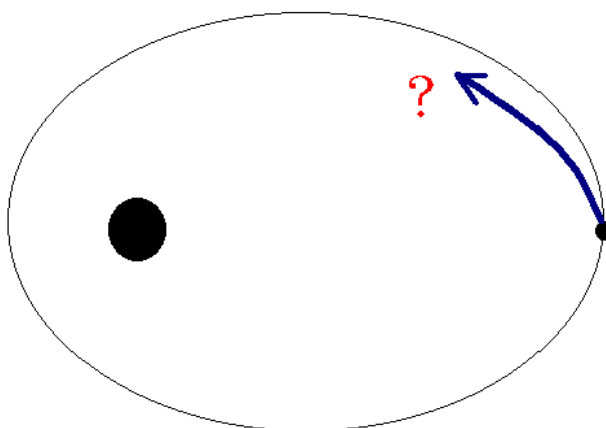
$$g_{\alpha\beta} = g_{\alpha\beta}^{(b)} + h_{\alpha\beta}^{\text{full}}$$

Lowest order in the mass ratio  $(\mu/M)^0$ :

Geodesic on the background geometry

Next order: Orbit deviates from the geodesic.

## Self-force



Point particle =====> Singular

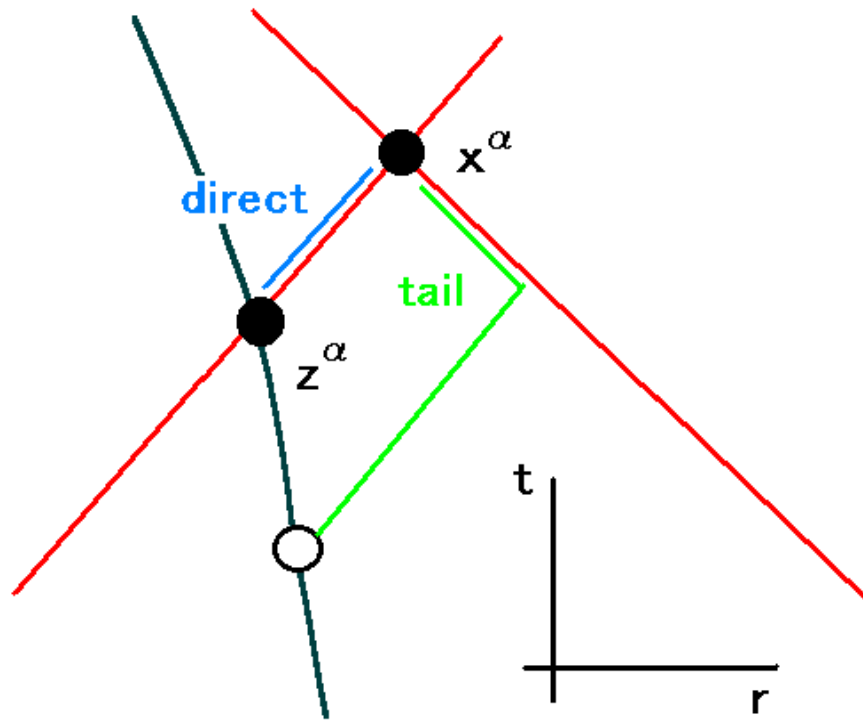
**Need Regularization!!**

\* Mino, Sasaki and Tanaka ('97)

Quinn and Wald ('97) [harmonic (H) gauge]

MiSaTaQuWa

$$h_{\alpha\beta}^{\text{full,H}} = h_{\alpha\beta}^{\text{dir,H}} + h_{\alpha\beta}^{\text{tail,H}}$$



$\{z^\alpha\}$ : Orbit of a particle

$\{x^\alpha\}$ : Regularization point

Gravitational self-force  $O(\mu/M)$ :

Deviation from geodesic

$$\mu \frac{D^2 z^\mu(\tau)}{d\tau^2} = F^\mu(z)$$

$$F^\mu = -\frac{\mu}{2} (g_{(b)}^{\mu\nu} + u^\mu u^\nu) (2h_{\nu\beta;\alpha}^{\text{tail,H}} - h_{\alpha\beta;\nu}^{\text{tail,H}}) u^\alpha u^\beta$$

$\{u^\alpha\}$ : Four velocity

$h^{\text{dir,H}}$ : Direct part

support only on the past light cone  
evaluated locally

$h^{\text{tail,H}}$ : Tail part

Support inside the past null cones

Depend on the whole history of the particle

\* Gives the physical self-force

It is difficult to calculate the tail part **directly**.

## Regularization

$$h_{\alpha\beta}^{\text{tail,H}} = h_{\alpha\beta}^{\text{full,H}} - h_{\alpha\beta}^{\text{dir,H}}$$

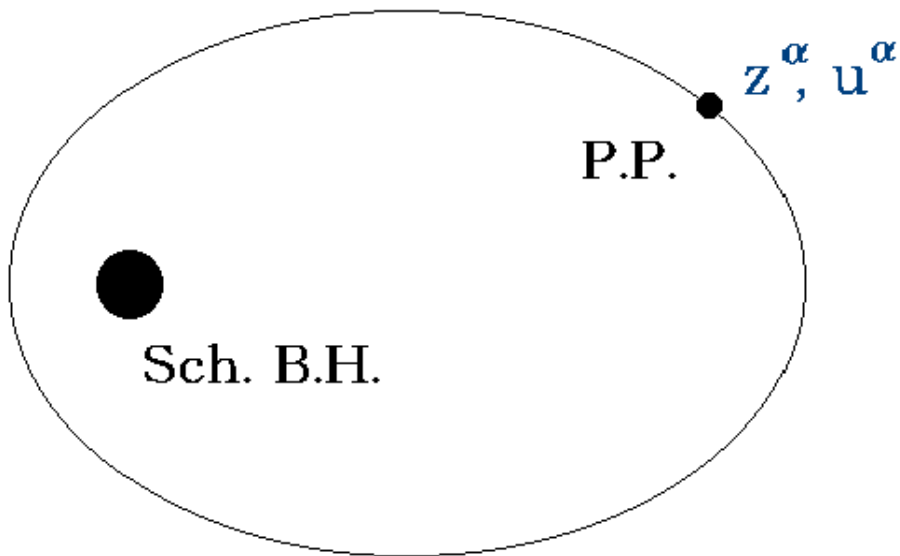
<Two Problems>

- 1) Subtraction problem
- 2) Gauge problem

## Background: Schwarzschild black hole

$$g_{\mu\nu}^{(b)} dx^\mu dx^\nu = -f(r)dt^2 + f(r)^{-1}dr^2 + r^2(d\theta^2 + \sin^2\theta d\phi^2)$$

$$f(r) = 1 - \frac{2M}{r}$$



Black hole: mass  $M$

Point particle: mass  $\mu$

## <Subtraction problem>

$$h_{\alpha\beta}^{\text{tail,H}} = h_{\alpha\beta}^{\text{full,H}} - h_{\alpha\beta}^{\text{dir,H}}$$

Full metric perturbation:

Form of spherical harmonic series

in the Regge-Wheeler/Teukolsky formalism

Fourier expansion: In order to treat analytically

Summation of (Fourier-)harmonic series

→ In general, numerical calculation

Direct part:

Given only **locally**

around the location of the particle

Difficult transformation

to the (Fourier-)harmonic series form

**Two different form → Same form**

**\* We use spherical harmonics expansion!**

## Mode sum (decomposition) regularization

Barack and Ori ('00)

Barack, Mino, Nakano, Ori and Sasaki ('02)

Full force:  $F_{\text{full}}^{\mu} = F^{\mu}[h^{\text{full}}]$

$$\begin{aligned} F_{\text{full}}^{\mu}(x) &= \sum_{\ell m \omega} F_{\text{full}}^{\mu}(\ell m \omega; x) \\ &= \sum_{\ell} F_{\text{full}}^{\mu}(\ell; x). \end{aligned}$$

Direct part:  $F_{\text{dir}}^{\mu} = F^{\mu}[h^{\text{dir}}]$

$$\begin{aligned} F_{\text{dir}}^{\mu}(x) &= \sum_{\ell m} F_{\text{dir}}^{\mu}(\ell m; x) \\ &= \sum_{\ell} F_{\text{dir}}^{\mu}(\ell; x). \end{aligned}$$

\* Derived by Legendre expansion directly.

Each  $\ell$  mode is finite at the particle location.

→ Possible take the coincidence limit  $x \rightarrow z_0$ .

$$\begin{aligned} F^{\mu}(\ell; z_0) &= F^{\mu}[h^{\text{tail}}](\ell; z_0) \\ &= F_{\text{full}}^{\mu}(\ell; z_0) - F_{\text{dir}}^{\mu}(\ell; z_0). \end{aligned}$$

Regularized self-force:

$$F^{\mu}(z_0) = \sum_{\ell} F^{\mu}(\ell; z_0).$$

<Standard form>

**Direct part of the self-force under H gauge**

**Regularization parameter:** ( $L = \ell + 1/2$ )

$$F_{\text{dir,H}}^{\mu(\pm)}|_{\ell} = \pm A^{\mu}L + B^{\mu} + \frac{C^{\mu}}{L} + D_{\ell}^{\mu}$$

$A^{\mu}$ -term: Quadratic divergence

$B^{\mu}$ -term: Linear divergence

$A^{\mu}$  and  $B^{\mu}$  are **independent of  $\ell$**

$\pm$  denotes that the limit to  $r_0$  is taken from the greater or smaller value of  $r$

$C^{\mu}$ -term: Logarithmic divergence

$$C^{\mu} = 0.$$

$D_{\ell}^{\mu}$ -term: Remaining finite contribution

**[Ambiguity]**

$$\begin{aligned} D_{\ell}^{\mu} &= \frac{d^{\mu}}{(2\ell - 1)(2\ell + 3)} \\ &\quad + \frac{e^{\mu}}{(2\ell - 1)(2\ell + 3)(2\ell - 3)(2\ell + 5)} + \dots \\ &= \frac{d^{\mu}}{4(L^2 - 1)} + \frac{e^{\mu}}{16(L^2 - 1)(L^2 - 4)} + \dots \end{aligned}$$

The summation of  $D_{\ell}^{\mu}$  over  $\ell$  (from  $\ell = 0$  to  $\infty$ ) **vanishes**.



<The difference of time/frequency domain>

“Analytic calculation”

$h_{\alpha\beta}^{\text{full,H}}$ : in the **frequency** domain

Slow motion approximation

considered as the post-Newtonian (PN) Approx.

\* Arbitrarily higher order calculation

$h_{\alpha\beta}^{\text{dir,H}}$ : in the **time** domain

evaluated locally

To use mode sum regularization

→ Need the integration w.r.t. the frequency.

Numerical calculation for general orbits.

**New regularization method** : **W. Hikida's Talk**

Decomposition of the Green function

Subtraction analytically for general orbits

Convergency of the slow motion Approx.

**Useful for numerical regularization!**

## <Gauge problem>

Direct part:  $h_{\alpha\beta}^{\text{dir,H}}$

Hadamard prescription in the harmonic gauge

Full metric perturbation:  $h_{\alpha\beta}^{\text{full,H}}$

in the Regge-Wheeler/Teukolsky formalism

**NOT** the one in the harmonic gauge

**Regge-Wheeler**/radiation gauge

\* Appropriate gauge transformation is needed.

Harmonic gauge approach:

**Sago, Nakano and Sasaki ('03)**

From RW gauge to harmonic gauge

for full metric perturbation

In practice, it is difficult to calculate.

Need double integrals.

**From harmonic to RW?**

Regularization:

$$h_{\alpha\beta}^{\text{tail,H}} = h_{\alpha\beta}^{\text{full,H}} - h_{\alpha\beta}^{\text{dir,H}}$$

Slight modification, **but very important!**

by **Detweiler & Whiting ('03)**

$$h_{\alpha\beta}^{\text{R,H}} = h_{\alpha\beta}^{\text{full,H}} - h_{\alpha\beta}^{\text{S,H}}$$

$h_{\mu\nu}^{\text{S,H}}$  : S part

Inhomogeneous solution of  
the linearized Einstein equation under H gauge

$$\begin{aligned} \bar{h}_{\mu\nu;\alpha}^{\text{H}} + 2R_{\mu\alpha\nu\beta}\bar{h}^{\text{H}\alpha\beta} &= -16\pi T_{\mu\nu} \\ \bar{h}_{\alpha\beta} &= h_{\alpha\beta} - \frac{1}{2}g_{\alpha\beta}h_{\mu}{}^{\mu} \end{aligned}$$

$h_{\mu\nu}^{\text{R,H}}$  : R part

Homogeneous solution

Tail force = R force

$$F_{\alpha} [h_{\mu\nu}^{\text{tail,H}}] = F_{\alpha} [h_{\mu\nu}^{\text{R,H}}]$$

Let's consider a **Finite Gauge Transformation**.

**Mino's idea**

$$\boxed{x_{\text{RW}}^\alpha = x_{\text{H}}^\alpha - \xi_{\text{H} \rightarrow \text{RW}}^\alpha [h_{\mu\nu}^{\text{R,H}}] .}$$

From harmonic to RW gauge

$$\begin{aligned} & \lim_{x \rightarrow z(\tau)} F_\alpha \left[ h_{\mu\nu}^{\text{R,H}} + 2 \xi_{(\mu;\nu)}^{\text{H} \rightarrow \text{RW}} [h_{\mu\nu}^{\text{R,H}}] \right] (x) \\ &= \lim_{x \rightarrow z(\tau)} F_\alpha \left[ h_{\mu\nu}^{\text{full,H}} - h_{\mu\nu}^{\text{S,H}} \right. \\ & \quad \left. + 2 \xi_{(\mu;\nu)}^{\text{H} \rightarrow \text{RW}} [h_{\mu\nu}^{\text{full,H}} - h_{\mu\nu}^{\text{S,H}}] \right] (x) \\ &= \lim_{x \rightarrow z(\tau)} F_\alpha \left[ h_{\mu\nu}^{\text{full,H}} + 2 \xi_{(\mu;\nu)}^{\text{H} \rightarrow \text{RW}} [h_{\mu\nu}^{\text{full,H}}] \right. \\ & \quad \left. - h_{\mu\nu}^{\text{S,H}} - 2 \xi_{(\mu;\nu)}^{\text{H} \rightarrow \text{RW}} [h_{\mu\nu}^{\text{S,H}}] \right] (x) \\ &= \lim_{x \rightarrow z(\tau)} F_\alpha \left[ h_{\mu\nu}^{\text{full,RW}} - h_{\mu\nu}^{\text{S,H}} - 2 \xi_{(\mu;\nu)}^{\text{H} \rightarrow \text{RW}} [h_{\mu\nu}^{\text{S,H}}] \right] (x) \\ &= \lim_{x \rightarrow z(\tau)} \left( F_\alpha [h_{\mu\nu}^{\text{full,RW}}] (x) \right. \\ & \quad \left. - F_\alpha [h_{\mu\nu}^{\text{S,H}} + 2 \xi_{(\mu;\nu)}^{\text{H} \rightarrow \text{RW}} [h_{\mu\nu}^{\text{S,H}}]] (x) \right) \\ &= F_\alpha^{\text{RW}}(\tau) \end{aligned}$$

This is a **Regge-Wheeler self-force**.

## §2. Regularization under RW gauge

★ First, we consider the **full metric perturbation**.

Metric perturbation in the **Schwarzschild B.G.**:

Using **tensor harmonics**,

$$\begin{aligned}
 \mathbf{h} = \sum_{\ell m} & \left[ f(r) H_{0\ell m}(t, r) \mathbf{a}_{\ell m}^{(0)} - i\sqrt{2} H_{1\ell m}(t, r) \mathbf{a}_{\ell m}^{(1)} \right. \\
 & + \frac{1}{f(r)} H_{2\ell m}(t, r) \mathbf{a}_{\ell m} \\
 & - \frac{i}{r} \sqrt{2\ell(\ell+1)} h_{0\ell m}^{(e)}(t, r) \mathbf{b}_{\ell m}^{(0)} + \frac{1}{r} \sqrt{2\ell(\ell+1)} h_{1\ell m}^{(e)}(t, r) \mathbf{b}_{\ell m} \\
 & + \sqrt{\frac{1}{2} \ell(\ell+1)(\ell-1)(\ell+2)} G_{\ell m}(t, r) \mathbf{f}_{\ell m} \\
 & + \left( \sqrt{2} K_{\ell m}(t, r) - \frac{\ell(\ell+1)}{\sqrt{2}} G_{\ell m}(t, r) \right) \mathbf{g}_{\ell m} \\
 & - \frac{\sqrt{2\ell(\ell+1)}}{r} h_{0\ell m}(t, r) \mathbf{c}_{\ell m}^{(0)} + \frac{i\sqrt{2\ell(\ell+1)}}{r} h_{1\ell m}(t, r) \mathbf{c}_{\ell m} \\
 & \left. + \frac{\sqrt{2\ell(\ell+1)(\ell-1)(\ell+2)}}{2r^2} h_{2\ell m}(t, r) \mathbf{d}_{\ell m} \right].
 \end{aligned}$$

Source:

$$\begin{aligned}
 \mathbf{T} = \sum_{\ell m} & \left[ A_{\ell m}^{(0)} \mathbf{a}_{\ell m}^{(0)} + A_{\ell m}^{(1)} \mathbf{a}_{\ell m}^{(1)} + A_{\ell m} \mathbf{a}_{\ell m} + B_{\ell m}^{(0)} \mathbf{b}_{\ell m}^{(0)} + B_{\ell m} \mathbf{b}_{\ell m} \right. \\
 & \left. + Q_{\ell m}^{(0)} \mathbf{c}_{\ell m}^{(0)} + Q_{\ell m} \mathbf{c}_{\ell m} + D_{\ell m} \mathbf{d}_{\ell m} + G_{\ell m}^{(s)} \mathbf{g}_{\ell m} + F_{\ell m} \mathbf{f}_{\ell m} \right],
 \end{aligned}$$

## Ten tensor harmonics:

$$\mathbf{a}_{\ell m}^{(0)} = \begin{pmatrix} Y_{\ell m} & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}, \quad \mathbf{a}_{\ell m}^{(1)} = (i/\sqrt{2}) \begin{pmatrix} 0 & Y_{\ell m} & 0 & 0 \\ Sym & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix},$$

$$\mathbf{a}_{\ell m} = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & Y_{\ell m} & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix},$$

$$\mathbf{b}_{\ell m}^{(0)} = ir[2\ell(\ell+1)]^{-1/2} \begin{pmatrix} 0 & 0 & (\partial/\partial\theta)Y_{\ell m} & (\partial/\partial\phi)Y_{\ell m} \\ 0 & 0 & 0 & 0 \\ Sym & 0 & 0 & 0 \\ Sym & 0 & 0 & 0 \end{pmatrix},$$

$$\mathbf{b}_{\ell m} = r[2\ell(\ell+1)]^{-1/2} \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & (\partial/\partial\theta)Y_{\ell m} & (\partial/\partial\phi)Y_{\ell m} \\ 0 & Sym & 0 & 0 \\ 0 & Sym & 0 & 0 \end{pmatrix},$$

$$\mathbf{c}_{\ell m}^{(0)} = r[2\ell(\ell+1)]^{-1/2} \times \begin{pmatrix} 0 & 0 & (1/\sin\theta)(\partial/\partial\phi)Y_{\ell m} & -\sin\theta(\partial/\partial\theta)Y_{\ell m} \\ 0 & 0 & 0 & 0 \\ Sym & 0 & 0 & 0 \\ Sym & 0 & 0 & 0 \end{pmatrix},$$

$$\mathbf{c}_{\ell m} = ir[2\ell(\ell + 1)]^{-1/2} \times \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & (1/\sin\theta)(\partial/\partial\phi)Y_{\ell m} - \sin\theta(\partial/\partial\theta)Y_{\ell m} & \\ 0 & Sym & 0 & 0 \\ 0 & Sym & 0 & 0 \end{pmatrix},$$

$$\mathbf{d}_{\ell m} = -ir^2[2\ell(\ell + 1)(\ell - 1)(\ell + 2)]^{-1/2} \times \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & -(1/\sin\theta)X_{\ell m} & \sin\theta W_{\ell m} \\ 0 & 0 & Sym & \sin\theta X_{\ell m} \end{pmatrix},$$

$$\mathbf{g}_{\ell m} = (r^2/\sqrt{2}) \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & Y_{\ell m} & 0 \\ 0 & 0 & 0 & \sin^2\theta Y_{\ell m} \end{pmatrix},$$

$$\mathbf{f}_{\ell m} = r^2[2\ell(\ell + 1)(\ell - 1)(\ell + 2)]^{-1/2} \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & W_{\ell m} & X_{\ell m} \\ 0 & 0 & Sym & -\sin^2\theta W_{\ell m} \end{pmatrix}.$$

$$X_{\ell m} = 2\frac{\partial}{\partial\phi} \left( \frac{\partial}{\partial\theta} - \cot\theta \right) Y_{\ell m},$$

$$W_{\ell m} = \left( \frac{\partial^2}{\partial\theta^2} - \cot\theta \frac{\partial}{\partial\theta} - \frac{1}{\sin^2\theta} \frac{\partial^2}{\partial\phi^2} \right) Y_{\ell m}.$$

**Odd parity:**  $(-1)^{\ell+1}$

**Even parity:**  $(-1)^\ell$

## Regge-Wheeler-Zerilli formalism [Zerilli ('70)]

$$\mathbf{h} = \sum_{\ell m} \left[ f(r) H_{0\ell m}(t, r) \mathbf{a}_{\ell m}^{(0)} - i\sqrt{2} H_{1\ell m}(t, r) \mathbf{a}_{\ell m}^{(1)} + \frac{1}{f(r)} H_{2\ell m}(t, r) \mathbf{a}_{\ell m} \cdots \right]$$

Linearized Einstein equation  $\rightarrow$

\* Introduce new radial functions:  $R_{\ell m \omega}^{(\text{odd/even})}$

Regge-Wheeler Equation:

(Using Chandrasekhar transformation)

$$\frac{d^2 R_{\ell m \omega}^{(\text{odd/even})}}{dr^{*2}} + [\omega^2 - V_\ell(r)] R_{\ell m \omega}^{(\text{odd/even})} = S_{\ell m \omega}^{(\text{odd/even})}$$

$V_\ell(r)$ : Regge-Wheeler potential

$S_{\ell m \omega}^{(\text{odd/even})}$ : Source

Radial function:  $R_{\ell m \omega}^{(\text{odd/even})}$

By a Green function method

Homogeneous solutions [Mano et al. ('96)]



Slow motion approximation:

$$z = \omega r \sim v$$

$$\epsilon = 2M\omega \sim v^3$$

\*  $O(v^n)$ :  $\binom{n}{2}$  post-Newtonian (PN) order

We can calculate arbitrarily higher order.

Regge-Wheeler gauge:

$$\boxed{h_{0lm\omega}^{(e)RW} = h_{1lm\omega}^{(e)RW} = G_{lm\omega}^{RW} = h_{2lm\omega}^{RW} = 0}$$

Reconstruction of metric perturbation:

$$h_{1lm\omega}^{RW} = \frac{r^2}{(r - 2M)} R_{lm\omega}^{(\text{odd})}$$

$$h_{0lm\omega} = \frac{i}{\omega} \frac{d}{dr^*} (r R_{lm\omega}^{(\text{odd})}) - \frac{8\pi r (r - 2M)}{\omega [\frac{1}{2}\ell(\ell + 1)(\ell - 1)(\ell + 2)]^{1/2}} D_{lm\omega}.$$

Regge-Wheeler-Zerilli formalism  $\rightarrow$  for  $\ell \geq 2$  modes

For  $\ell = 0$  and 1 modes

[Zerilli ('70), Detweiler and Poisson ('04)]

## S-Part of Metric Perturbation under H gauge

$$\begin{aligned} \bar{h}_{\mu\nu}^{\text{S,H}} = & 4\mu \left[ \frac{\bar{g}_{\mu\alpha}(x, z(\tau_{\text{ret}}))\bar{g}_{\nu\beta}(x, z(\tau_{\text{ret}}))u^\alpha(\tau_{\text{ret}})u^\beta(\tau_{\text{ret}})}{\sigma_{;\gamma}(x, z(\tau_{\text{ret}}))u^\gamma(\tau_{\text{ret}})} \right] \\ & + 2\mu \int_{\tau_{\text{ret}}}^{\tau_{\text{adv}}} \bar{g}_\mu{}^\alpha(x, z(\tau))\bar{g}_\nu{}^\beta(x, z(\tau))R_{\gamma\alpha\delta\beta}(z(\tau))u^\gamma(\tau)u^\delta(\tau) d\tau \\ & + O(y^2). \end{aligned}$$

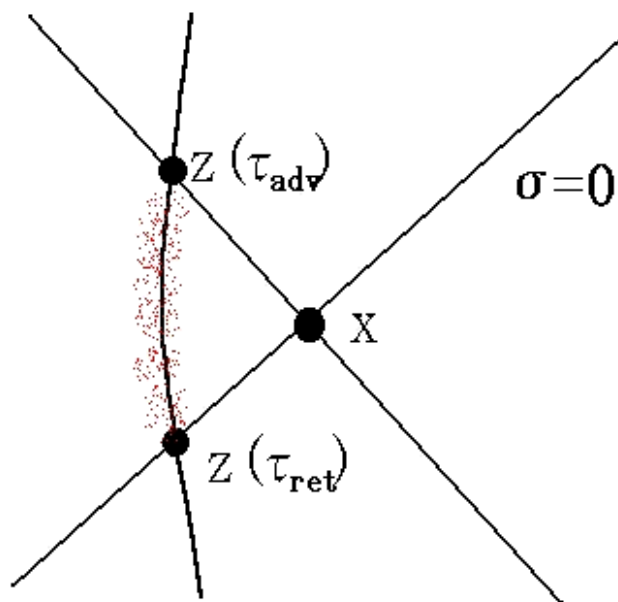
$\sigma(x, z)$ : **Bi-scalar**  
of half the squared geodesic distance

$\bar{g}_{\alpha\beta}(x, z)$ : **Parallel displacement bi-vector**

$\tau_{\text{ret}}(x)$ : **Retarded time for  $x$**

$y$ : **Coordinate difference between  $x$  and  $z_0$**

$z_0$ : **Location of the particle**



Metric components:

(Local coordinate expansion)

$$h_{\alpha\beta}^{\text{S,H}} = \mu \sum_{m,n,p,q,r} C_{\alpha\beta}^{m,n,p,q,r} \frac{T^m R^n \Theta^p \Phi^q}{\epsilon^r} + O(y^2)$$

$$\begin{aligned} \epsilon &:= (r_0^2 + r^2 - 2r_0 r \cos \Theta \cos \Phi)^{1/2}, \\ T &:= t - t_0, \quad R := r - r_0, \\ \Theta &:= \theta - \frac{\pi}{2}, \quad \Phi := \phi - \phi_0. \end{aligned}$$

When we use PN expansion,

→ the above quantities are useful.

$$\text{e.g.) } \frac{1}{\epsilon} = \sum_{\ell m} \frac{1}{r_{>}} \left( \frac{r_{<}}{r_{>}} \right) Y_{\ell m}^*(\Omega_0) Y_{\ell m}(\Omega)$$

Full relativistic treatment

**Mino, Nakano and Sasaki ('03)**

**Tensor harmonics expansion of local quantities**

Some **ambiguity** in spherical extension.

→ **Gauge transformation to RW gauge**

Standard form recovers under H gauge.

$$\begin{aligned}
F_{S,H}^{r(\text{even})}|_\ell = & \sum_m \left[ -\frac{1}{2} (2 \mathcal{E}^2 r_0^3 - i u_r \mathcal{L} m r_0^2 + 2 i r_0 u_r \mathcal{L} M m \right. \\
& \left. - 2 r_0 \mathcal{L}^2 + 4 \mathcal{L}^2 M) \mathcal{L}^2 m^2 G_{\ell m}^{\text{S,H}}(t_0, r_0) / r_0^6 \right. \\
& + \frac{1}{2} (2 \mathcal{E}^2 r_0^3 - i u_r \mathcal{L} m r_0^2 + 2 i r_0 u_r \mathcal{L} M m \\
& \left. - 2 r_0 \mathcal{L}^2 + 4 \mathcal{L}^2 M) \mathcal{L}^2 K_{\ell m}^{\text{S,H}}(t_0, r_0) / r_0^6 \right. \\
& + \frac{1}{2} \frac{\mathcal{E} \mathcal{L}^2 u_r m^2 \partial_t G_{\ell m}^{\text{S,H}}(t_0, r_0)}{r_0^2} \\
& - \frac{1}{2} \frac{\mathcal{E}^3 r_0 u_r \partial_t H_{0\ell m}^{\text{S,H}}(t_0, r_0)}{r_0 - 2M} \\
& - \frac{i (\mathcal{E}^2 r_0^3 - r_0 \mathcal{L}^2 + 2 \mathcal{L}^2 M) \mathcal{E} \mathcal{L} m \partial_t h_{1\ell m}^{(\text{e})\text{S,H}}(t_0, r_0)}{r_0^4 (r_0 - 2M)} \\
& \left. - \frac{i \mathcal{E}^2 u_r \mathcal{L} m \partial_t h_{0\ell m}^{(\text{e})\text{S,H}}(t_0, r_0)}{(r_0 - 2M) r_0} + \dots \right] Y_{\ell m}(\Omega_0).
\end{aligned}$$

$$\mathcal{E} = -g_{tt} \frac{dt}{d\tau}, \quad \mathcal{L} = g_{\phi\phi} \frac{d\phi}{d\tau}, \quad u_r = g_{rr} \frac{dr}{d\tau}.$$

Interruption factor of standard form:

$$\begin{aligned}
h_{0\ell m}^{(\text{e})\text{S,H}}(t, r) & \sim \frac{1}{\ell(\ell + 1)}, \\
h_{1\ell m}^{(\text{e})\text{S,H}}(t, r) & \sim \frac{1}{\ell(\ell + 1)}, \\
G_{\ell m}^{\text{S,H}}(t, r) & \sim \frac{1}{\ell(\ell + 1)(\ell - 1)(\ell + 2)}.
\end{aligned}$$

$$\begin{aligned}
h_{0\ell m}^{(e)S,H}(t, r) &\sim \frac{1}{\ell(\ell + 1)}, \\
h_{1\ell m}^{(e)S,H}(t, r) &\sim \frac{1}{\ell(\ell + 1)}, \\
G_{\ell m}^{S,H}(t, r) &\sim \frac{1}{\ell(\ell + 1)(\ell - 1)(\ell + 2)}.
\end{aligned}$$

<Standard form>

Regularization parameter:

$$F^{\mu(\pm)}|_{\ell} = \pm A^{\mu}L + B^{\mu} + \frac{C^{\mu}}{L} + D_{\ell}^{\mu}$$

$A^{\mu}$ -term: Quadratic divergence

$B^{\mu}$ -term: Linear divergence

$A^{\mu}$  and  $B^{\mu}$  are **independent of  $\ell$**  ( $L = \ell + 1/2$ )

$C^{\mu}$ -term: Logarithmic divergence

$$C^{\mu} = 0.$$

$D_{\ell}^{\mu}$ -term: Remaining finite contribution

$$\begin{aligned}
D_{\ell}^{\mu} &= \frac{d^{\mu}}{(2\ell - 1)(2\ell + 3)} \\
&\quad + \frac{e^{\mu}}{(2\ell - 1)(2\ell + 3)(2\ell - 3)(2\ell + 5)} + \dots \\
&= \frac{d^{\mu}}{4(L^2 - 1)} + \frac{e^{\mu}}{16(L^2 - 1)(L^2 - 4)} + \dots
\end{aligned}$$

$$\begin{aligned}
h_{0\ell m}^{(e)S,H}(t, r) &\sim \frac{1}{\ell(\ell+1)}, \\
h_{1\ell m}^{(e)S,H}(t, r) &\sim \frac{1}{\ell(\ell+1)}, \\
G_{\ell m}^{S,H}(t, r) &\sim \frac{1}{\ell(\ell+1)(\ell-1)(\ell+2)}.
\end{aligned}$$

**Summation over  $m$  modes:**

$$\begin{aligned}
\sum_m \left[ \partial_\phi^{2k} Y_{\ell m}^*(\Omega_0) \right] Y_{\ell m}(\Omega_0) &= \sum_m \left[ (-1)^k m^{2k} Y_{\ell m}^*(\Omega_0) \right] Y_{\ell m}(\Omega_0) \\
&\sim \ell(\ell+1) \times (\text{function of } \ell).
\end{aligned}$$

$$\begin{aligned}
\sum_m \left[ (\partial_\theta^2 - \partial_\phi^2) \partial_\phi^{2k} Y_{\ell m}^*(\Omega_0) \right] Y_{\ell m}(\Omega_0) \\
= \sum_m \left[ (-\ell(\ell+1) + 2m^2) (-1)^k m^{2k} Y_{\ell m}^*(\Omega_0) \right] Y_{\ell m}(\Omega_0) \\
\sim \ell(\ell+1)(\ell-1)(\ell+2) \times (\text{function of } \ell).
\end{aligned}$$

$$\sum_m |\partial_\theta Y_{\ell m}(\Omega_0)|^2 \sim \ell(\ell+1).$$

$$\begin{aligned}
\sum_m m^{2k} |\partial_\theta Y_{\ell m}(\Omega_0)|^2 &\sim \ell(\ell+1)(\ell-1)(\ell+2) \\
&\times (\text{function of } \ell).
\end{aligned}$$

**\* We are trying to prove “standard form”!**

## Generators of gauge transformation:

$$\begin{aligned}\xi_\mu^{(\text{odd})} &= \sum_{\ell m} \Lambda_{\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) \\ &\quad \times \left\{ 0, 0, \frac{-1}{\sin \theta} \partial_\phi Y_{\ell m}(\theta, \phi), \sin \theta \partial_\theta Y_{\ell m}(\theta, \phi) \right\}, \\ \xi_\mu^{(\text{even})} &= \sum_{\ell m} \left\{ M_{0\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) Y_{\ell m}(\theta, \phi), M_{1\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) Y_{\ell m}(\theta, \phi), \right. \\ &\quad \left. M_{2\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) \partial_\theta Y_{\ell m}(\theta, \phi), M_{2\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) \partial_\phi Y_{\ell m}(\theta, \phi) \right\}.\end{aligned}$$

## Regge-Wheeler gauge condition:

$$\boxed{h_{0\ell m}^{(\text{e})\text{S,RW}} = h_{1\ell m}^{(\text{e})\text{S,RW}} = G_{\ell m}^{\text{S,RW}} = h_{2\ell m}^{\text{S,RW}} = 0}$$

$$h_{2\ell m}^{\text{S,H}}(t, r) = -2i \Lambda_{\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r),$$

$$h_{0\ell m}^{(\text{e})\text{S,H}}(t, r) = -M_{0\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) - \partial_t M_{2\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r),$$

$$h_{1\ell m}^{(\text{e})\text{S,H}}(t, r) = -M_{1\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) - r^2 \partial_r \left( \frac{M_{2\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r)}{r^2} \right),$$

$$G_{\ell m}^{\text{S,H}}(t, r) = -\frac{2}{r^2} M_{2\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r).$$

## Gauge deterministic on sphere:

$$\Lambda_{\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) = \frac{i}{2} h_{2\ell m}^{\text{S,H}}(t, r),$$

$$M_{2\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) = -\frac{r^2}{2} G_{\ell m}^{\text{S,H}}(t, r),$$

$$M_{0\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) = -h_{0\ell m}^{(\text{e})\text{S,H}}(t, r) - \partial_t M_{2\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r),$$

$$M_{1\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) = -h_{1\ell m}^{(\text{e})\text{S,H}}(t, r) - r^2 \partial_r \left( \frac{M_{2\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r)}{r^2} \right).$$

**NOT** necessary to integrate w.r.t.  $t$  or  $r$ .

**Gauge functions are determined uniquely.**

(for  $\ell \geq 2$ )

**The odd parity components:**

$$h_{0\ell m}^{\text{S,RW}}(t, r) = h_{0\ell m}^{\text{S,H}}(t, r) + \partial_t \Lambda_{\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r),$$

$$h_{1\ell m}^{\text{S,RW}}(t, r) = h_{1\ell m}^{\text{S,H}}(t, r) + r^2 \partial_r \left( \frac{\Lambda_{\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r)}{r^2} \right),$$

**The even parity components:**

$$H_{0\ell m}^{\text{S,RW}}(t, r) = H_{0\ell m}^{\text{S,H}}(t, r) + \frac{2r}{r - 2M} [\partial_t M_{0\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) - \frac{M(r - 2M)}{r^3} M_{1\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r)],$$

$$H_{1\ell m}^{\text{S,RW}}(t, r) = H_{1\ell m}^{\text{S,H}}(t, r) + [\partial_t M_{1\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) + \partial_r M_{0\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) - \frac{2M}{r(r - 2M)} M_{0\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r)],$$

$$H_{2\ell m}^{\text{S,RW}}(t, r) = H_{2\ell m}^{\text{S,H}}(t, r) + \frac{2(r - 2M)}{r} [\partial_r M_{1\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r) + \frac{M}{r(r - 2M)} M_{1\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r)],$$

$$K_{\ell m}^{\text{S,RW}}(t, r) = K_{\ell m}^{\text{S,H}}(t, r) + \frac{2(r - 2M)}{r^2} M_{1\ell m}^{\text{S,H} \rightarrow \text{RW}}(t, r),$$

$h_{0\ell m \omega}^{(\text{e})\text{S,RW}} = h_{1\ell m \omega}^{(\text{e})\text{S,RW}} = G_{\ell m \omega}^{\text{S,RW}} = h_{2\ell m \omega}^{\text{S,RW}} = 0$
--



## S-force difference between H and RW for general orbit

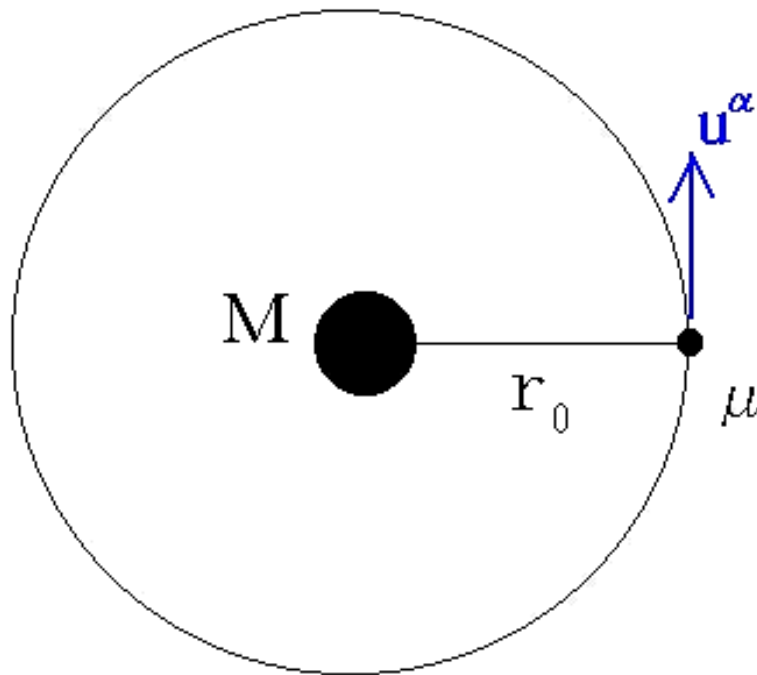
(In terms of the **M.P. under H gauge**)

$$\begin{aligned}
\delta F_{S,H \rightarrow RW}^{r(\text{even})}|_\ell = & \sum_m \left[ \frac{1}{2} (-i r_0 u_r \mathcal{L} (r_0 - 2M) m + 4 \mathcal{L}^2 M \right. \\
& + 2 \mathcal{E}^2 r_0^3 - 2 r_0 \mathcal{L}^2) \mathcal{L}^2 m^2 G_{\ell m}^{S,H}(t_0, r_0) / r_0^6 \\
& - \frac{1}{2} (\mathcal{L}^2 r_0 (r_0^3 \mathcal{E}^2 - 2 r_0^3 + 4 r_0^2 M \\
& - \mathcal{L}^2 r_0 + 2 \mathcal{L}^2 M) m^2 \\
& + i r_0 \mathcal{L} u_r (2 r_0^5 + 2 r_0^4 \mathcal{E}^2 - 7 r_0^3 M \\
& - 4 r_0^3 M \mathcal{E}^2 + 6 r_0^2 M^2 - r_0^2 \mathcal{L}^2 \\
& + 3 r_0^2 \mathcal{L}^2 M - 2 \mathcal{L}^2 M^2) m + 4 r_0 \mathcal{L}^4 M \\
& + 2 \mathcal{E}^2 r_0^3 \mathcal{L}^2 M - \mathcal{E}^2 r_0^4 \mathcal{L}^2 - 4 \mathcal{L}^4 M^2 \\
& - 2 r_0^6 \mathcal{E}^2 + 2 \mathcal{E}^4 r_0^6 - r_0^2 \mathcal{L}^4) \\
& \times \partial_r G_{\ell m}^{S,H}(t_0, r_0) / r_0^6 \\
& + \frac{r_0^2 \mathcal{E}^3 u_r \partial_t^2 h_{0\ell m}^{(e)S,H}(t_0, r_0)}{(r_0 - 2M)^2} \\
& - \frac{1}{2} (-2 r_0^3 + 3 \mathcal{E}^2 r_0^3 + 4 M r_0^2 - 3 r_0 \mathcal{L}^2 \\
& + 6 \mathcal{L}^2 M) r_0 \mathcal{E}^2 \\
& \times \partial_t^2 \partial_r G_{\ell m}^{S,H}(t_0, r_0) / (r_0 - 2M)^2 \\
& \left. + \dots \dots \dots \right] Y_{\ell m}(\Omega_0) .
\end{aligned}$$

Circular orbit:

$$\{z_0^\alpha\} = \left\{ t_0, r_0, \frac{\pi}{2}, \phi_0 \right\}$$

$$\{u^\alpha\} = \left\{ \frac{1}{\sqrt{1 - 3M/r_0}}, 0, 0, \sqrt{\frac{M}{r_0^2(r_0 - 3M)}} \right\}$$



$$F^\phi = \frac{r_0 - 2M}{r_0^3 \Omega} F^t,$$

$$\Omega = \frac{u^\phi}{u^t}.$$

## S-force difference between H and RW for circular orbit

$$\delta F_{S,H \rightarrow RW}^{t(\text{odd})} |_{\ell} = 0,$$

$$\delta F_{S,H \rightarrow RW}^{r(\text{odd})} |_{\ell} = 0,$$

$$\delta F_{S,H \rightarrow RW}^{t(\text{even})} |_{\ell} = 0,$$

$$\delta F_{S,H \rightarrow RW}^{r(\text{even})} |_{\ell} = \sum_m \left[ -3 \frac{M (r_0 - 2M)^2 h_{1\ell m}^{(e)S,H}(t_0, r_0)}{r_0^4 (r_0 - 3M)} + \frac{3M (r_0 - 2M)^2 \partial_r G_{\ell m}^{S,H}(t_0, r_0)}{2 r_0^2 (r_0 - 3M)} \right] Y_{\ell m}(\Omega_0).$$

Because

Source (frequency domain):  $\delta(\omega - m\Omega)$

$$\rightarrow \partial_t = -im\Omega$$

Gauge force: 2PN  $O(v^4)$

Because

$$h_{1\ell m}^{(e)S,H}(t, r) \sim \frac{M}{\ell(\ell + 1)},$$

$$G_{\ell m}^{S,H}(t, r) \sim \frac{M}{\ell(\ell + 1)(\ell - 1)(\ell + 2)}.$$

## 1PN Circular Example:

We consider only  $r$ -component. ( $\ell \geq 2$ )

$$\begin{aligned} F_{\text{RW}}^r|_{\ell} &= F_{\text{full,RW}}^r|_{\ell} - F_{\text{S,RW}}^r|_{\ell} \\ &= -\mu^2 \frac{45 M}{8(2\ell - 1)(2\ell + 3) r_0^3} \end{aligned}$$

After summing over  $\ell$  mode

$$F_{\text{RW}}^r(\ell \geq 2) = -\frac{3\mu^2 M}{4r_0^3}$$

For  $\ell = 0$  and 1 modes with appropriate boundary

[Detweiler and Poisson ('04)]

$$F_{\text{RW}}^r(\ell = 0, 1) = \frac{2\mu^2}{r_0^2} - \frac{57\mu^2 M}{4r_0^3}$$

Regularized gravitational self-force:

$$F_{\text{RW}}^r = \frac{2\mu^2}{r_0^2} - \frac{15\mu^2 M}{r_0^3}.$$

→ Correction to the radius of the orbit  
that deviates from the geodesic  
in the unperturbed background.

### §3. Discussion

Schwarzschild background / Regge-Wheeler gauge

**S-part of the self-force**

**Standard form?**

$$F^{\mu(\pm)}|_{\ell} = \pm A^{\mu}L + B^{\mu} + \frac{C^{\mu}}{L} + D_{\ell}^{\mu}$$

\* Plunging into a B.H.

Standard form recovers.

**Barack and Ori ('01), Barack and Lousto ('02)**

\* 1PN calculation for circular orbit

Standard form recovers.

**Higher PN order calculation?**

**Black Hole Perturbation Club**

<http://www2.yukawa.kyoto-u.ac.jp/~misao/BHPC/>