Self-Force for a Scalar Particle in Schwarzschild Spacetime: Time-Domain Calculations with Adaptive Mesh Refinement

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Barack-Ori mode-sum regularization scheme

Lorenz gauge:

• metric perturbation from point particle is relatively simple (locally isotropic)

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Decompose metric perturbation into spherical harmonics (ℓ, m) :

- metric perturbations aren't separable
 - \Rightarrow compute modes via time-domain numerical integration

$$\Box \phi_{\ell m} + V_{\ell}(r)\phi_{\ell m} = S_{\ell m}(t)\,\delta\big(r(t) - r_{\mathsf{particle}}(t)\big)$$

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Compute self-force via regularized mode sum:

$$F_{\text{self}}^{c} \sim \sum_{\ell} \left[\left(\sum_{m} \left[\nabla^{c} \phi_{\ell m} \right]_{\text{particle}} \right) - (\ell + \frac{1}{2})A - B \right]$$

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Integrate one v = constantslice at a time via usual diamond-cell scheme (!)



4th Order Finite Differencing: Vacuum

Vacuum region:

- scheme described by Haas (2007):
- requires 3 past time levels, and 3 past points to start integrating a slice



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 $\delta\text{-function}$ source term $\Rightarrow\phi$ is C^{0} at the particle worldline

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 If evaluating at i = i_{particle}, compute 2nd order steps using cell sizes 1, 2, and 3 × normal, then Richardson-extrapolate to 4th order.





Adaptive Mesh Refinement: Motivation

 ϕ and its derivatives have very high dynamic range across the problem domain: phi(u) on relative v=190.0 slice



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- refine in both space and time
- fine grids overlay coarser grids
- integrate each grid independently



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- integrate each grid independently
- fine-grid initial & boundary data interpolated from coarse grids
- inject fine-grid results back into coarse grid when/where points coincide [keeps coarse-grid solution from drifting away from (accurate) fine-grid solution]



Adaptive Mesh Refinement: Characteristic Berger-Oliger

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Standard Cauchy Berger-Oliger also works in characteristic coordinates, with one modification: after injecting fine-grid data back into coarse grid, we need to re-integrate the tail of the coarse slice starting from the injected data.

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Tail Sum

Fit data points

$$\frac{1}{2}(\mathit{F}^{-}_{\ell,\mathsf{reg}} + \mathit{F}^{+}_{\ell,\mathsf{reg}}) \pm \frac{1}{2}|\mathit{F}^{+}_{\ell,\mathsf{reg}} - \mathit{F}^{-}_{\ell,\mathsf{reg}}|$$

to series expansion [Detweiler, Messaritaki, and Whiting (2003)]:

$$F_{\ell, \text{reg}} = \frac{\frac{a_2}{(\ell - \frac{1}{2})(\ell + \frac{3}{2})} + \frac{a_4}{(\ell - \frac{3}{2})(\ell - \frac{1}{2})(\ell + \frac{3}{2})(\ell + \frac{5}{2})} \\ + \frac{a_6}{(\ell - \frac{5}{2})(\ell - \frac{3}{2})(\ell - \frac{1}{2})(\ell + \frac{3}{2})(\ell + \frac{5}{2})(\ell + \frac{7}{2})} \\ + \mathcal{O}(\ell^{-8})$$

to determine coefficients $\{a_2, a_4, a_6\}$, then evaluate $\sum_{\ell=K+1}^{\infty} F_{\ell, \text{reg}}$ analytically (the sums telescope). [Better would be to use analytically-known value of a_2 .]

Sample Results: $(\ell, m) = (10, 10)$ Integration

Refinement-Level Map:



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Sample Frame \Rightarrow Movie:



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Sample Results: $(\ell, m) = (10, 10)$ Integration

4th Order Convergence with Grid Resolution:



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Test case:

- Particle in circular orbit at r = 10M
- Modes computed numerically for $\ell \leq$ 30 (256 modes).

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| • $\ell = 0$ | domain size | 30 000 <i>M</i> |
|---------------|-------------|-----------------|
| $\ell = 1$ | domain size | 5 000 <i>M</i> |
| $\ell = 2$ | domain size | 1000 <i>M</i> |
| $\ell=3,4$ | domain size | 500 <i>M</i> |
| $\ell \geq 5$ | domain size | 400 <i>M</i> |

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- $\ell = 0$ domain size $30\,000M$
 - $\ell = 1$ domain size $5\,000M$
 - $\ell = 2$ domain size $1\,000M$
 - $\ell = 3,4$ domain size 500*M*
 - $\ell \geq 5$ domain size 400M
- AMR error tolerance $\|\phi^{\text{normal}} \phi^{\frac{1}{2} \text{ resolution}}\| \le 10^{-15}$

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- AMR error tolerance $\|\phi^{\text{normal}} \phi^{\frac{1}{2} \text{ resolution}}\| \le 10^{-15}$
- Performance: $\ell = 0$: 108 minutes (1 proc), $\ell = 1-30$: ~5 hours (10 procs) median AMR speedup ≈ 25



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 F_{ℓ} convergence with ℓ :



rescaled F_{ℓ} convergence with ℓ :



rescaled F_{ℓ} convergence with ℓ for $r_* = 10M$ run:



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 $F[\ell = 0]$ time dependence starting at t = 1000M:



 $F[\ell = 0]$ time dependence starting at t = 10000M:



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F time-and inside/outside-dependence near end of evolution:



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F convergence with grid resolutions $\times 1$, $\times 2$, $\times 3$:



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F (outside) time-dependence near end of evolution:



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Conclusions

Current status:

- The characteristic Berger-Oliger algorithm works very well
- The combination of 4th order finite differencing and Berger-Oliger adaptive mesh refinement gives very high accuracy for self-force calculations

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Near-term plans:

• Clean up tail fitting!

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Near-term plans:

• Clean up tail fitting!

Longer-term plans:

• Extend finite differencing to handle eccentric orbits

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• Implement "self-consistent orbit correction"